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Probability Theory Part 2

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Outline

Basic Definitions

2 Discrete Distributions

- Binomial
- Poisson
- Multinomial
- 3 Continuous Distributions
 - Gaussian
 - χ²
 - Cauchy



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1. Probability Function

A function f(x) that gives a rule for assigning a probability P(x) to outcome x is called a probability function.

So far, we have talked about outcomes that can be modeled with *n*-tuples, (z_1, \dots, z_n) , whose elements are drawn from the set of natural numbers $\mathbb{N} = \{0, 1, \dots, \aleph_0\}$. But outcomes can also be modeled using *n*-tuples with elements drawn from the set of real numbers $\mathbb{R} = (-c, c)^a$.

^aGeorg Cantor (1845 - 1918), inventor of set theory, proved the astonishing theorem $c = 2^{\aleph_0}$, that the cardinalities c and \aleph_0 of sets \mathbb{R} and \mathbb{N} , respectively, are related in this amazing way. We typically use the symbol ∞ instead of c.

2. Probability Mass Function

If x is from \mathbb{N} , then the probability function f(x) is called a probability mass function (pmf). But we physicists hardly ever use this jargon. Notice that f(x) is a probability.

3. Probability Density Function

If x is from \mathbb{R} , then the probability function f(x) is called a probability density function (pdf) and is often written with a lower case letter, e.g., p. Notice that f(x) is not a probability.

To get a probability, we must integrate the pdf over an interval whose size is at least as large as an infinitesimal dx^a . More usefully, we compute

$$P=\int_{x_1}^{x_2}f(X)\,dX.$$

^aAn infinitely small non-zero number!

Random Variables

Formal books on statistics distinguish between a random variable X, denoted with an upper case letter, and its outcomes x, denoted by lower case letters.

However, most physicists typically do not make this distinction

More Definitions

Several standard numbers are used to characterize a probability distribution. Here are a few.

4. Moments

The r^{th} moment $\mu_r(a)$ about *a* of a probability distribution with probability function f(x) is defined by^{*a*}

$$\mu_r(a) = \int_{S_x} (x-a)^r f(x) \, dx,$$

where S_x is the domain^b of f(x).

 $\mu = \mu_1(0)$ is called the mean and is a measure of the location of the function f(x); $V(x) = \mu_2(\mu)$ is called the variance and $\sigma = \sqrt{V}$ is the standard deviation, which is one measure of the width of f(x).

^aFor discrete distributions, we replace the integral by a sum.

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^bThe domain of a function is the set of its "input" values. The range is the set of its "outputs".

Yet More Definitions

5. Quantile Function

The function

$$D(x) = \int_{X \le x} f(X) \, dX$$

is called the cumulative distribution function (cdf) of f(x). (Here distinguishing between X and x turns out to be helpful!) The function x = Q(P) that returns x given P = D(x) is called the quantile function and x is called the P-quantile of f(x).

Sometimes it is convenient to distinguish between the left cdf $D_L(x) \equiv D(x)$ and the right cdf defined by

$$D_R(x) = \int_{X \ge x} f(X) \, dX.$$

And More Definitions ...!

6. Covariance, Correlation, Independence

The covariance of random variables x and y with probability function f(x, y) is defined by

$$\operatorname{Cov}(x,y) = \int_{\mathcal{S}_x} \int_{\mathcal{S}_y} (x - \mu_x) (y - \mu_y) f(x,y) \, dx \, dy.$$

It is a measure of the correlation between the variables x and y.

If the probability function f(x, y) can be written as f(x, y) = f(x) f(y)then variables x and y are said to be independent in which case Cov(x, y) = 0.

However, in general, Cov(x, y) = 0 does not imply independence.

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Example (2.1 The Binomial and the LHC)

Consider *m* proton-proton collisions at the LHC and suppose we have *r* successes, say the creation of a Higgs boson. However, we are able to record only n < m collision events of which $k \le n$ are successes.

Problem What is the probability P(k, n|r, m) to get k successes and n - k failures in n trials given that they are drawn at random from a "box" called the LHC containing r unknown successful collisions and m - r unknown failures?

Assumptions

- The order of proton-proton collisions is irrelevant.
- 2 Every sample of collisions of size n is equally probable.

This problem is exactly the same as drawing k red balls and n - k blue balls from an urn with r red balls and m - r blue balls.

And, like the LHC, the drawing of red and blue balls, that is, Higgs boson and non-Higgs boson events, is done without replacement.



Example (2.1 The Binomial and the LHC) Solution Plan:

- Determine the number of ways T to get n collisions from m collisions regardless of whether a collision is a success or a failure.
- Obtermine the number of ways S to get exactly k successful collisions from r successful collisions.
- Solution Determine the number of ways F to get exactly n k failed collisions from m r failed collisions.
- Since successes and failures are assumed to be independent, the number of samples of size *n* with *k* successes and *n* − *k* failures is N = S × F.
- Therefore, $P(k, n|r, m) = S \times F/T$.

• How many samples T of *n* collisions can be drawn from *m* collisions?

 $\binom{m}{n}$

2 How many samples S of k successes can be drawn from r successes?

• How many samples F of n - k failures can be drawn from m - r failures?

$$\binom{m-r}{n-k}$$

$$P(k, n|r, m) = \binom{r}{k} \binom{m-r}{n-k} / \binom{m}{n}$$

This probability can be rewritten as

$$P(k,n|r,m) = \binom{n}{k} f(k,n,r,m),$$

where
$$f(k, n, r, m) = \frac{r!}{(r-k)!} \frac{(m-r)!}{(m-r-n+k)!} / \frac{m!}{(m-n)!}$$
.

We now ask what is the probability of k, n irrespective of r, m?

This requires that we consider all values of r and m that are possible a priori and sum the probability P(k, n|r, m) weighted by the probability P(r, m) of r and m.

That is, we need to compute the sum

$$P(k,n) = \sum_{r,m} P(k,n|r,m) P(r,m).$$

The elimination of quantities like r and m that are not of current interest is an example of a common procedure in probability theory called marginalization.

We can already see a potential problem. It is far from clear what we should put for P(r, m). But let's nevertheless continue and see where this leads.

$$P(k,n) = \sum_{r,m} P(k,n|r,m) P(r,m).$$

Let's rewrite the expression above in terms of the unknown relative frequency of success, z = r/m:

$$P(k,n) = \sum_{z,m} P(k,n|zm,m) P(zm,m),$$
$$= {n \choose k} \sum_{z,m} f(k,n,z,m) P(zm,m).$$

At the LHC, *m*, the number of proton-proton collisions, is huge.

Therefore, let's consider the idealization $m \to \infty$ while keeping k and n fixed. It's as if we've used up our disk space quota at CERN even as the LHC continues to run!

$$P(k,n) = \binom{n}{k} \sum_{z,m} f(k,n,z,m) P(zm,m),$$

Exercise 2.1

show that $f(k, n, z, m) \rightarrow z^k (1-z)^{n-k}$ as $m \rightarrow \infty$.

What about the probabilities P(zm, m)?

To see what happens, write P(k, n) as

$$P(k,n) \rightarrow \sum_{z} \sum_{m} {n \choose k} z^{k} (1-z)^{n-k} P(zm,m) \text{ with } z = r/m,$$
$$= \sum_{z} {n \choose k} z^{k} (1-z)^{n-k} \sum_{m} P(zm,m),$$

 $P(k,n) = \sum_{z} {n \choose k} z^k (1-z)^{n-k} \sum_{m} P(zm,m),$

As $m \to \infty$, the sum converges to an integral and we obtain:

Bruno de Finetti's Representation Theorem

$$P(k, n) = \int_0^1 \operatorname{binomial}(k, n, z) \pi(z) \, dz, \text{ where}$$

binomial(k, n, z) = $\binom{n}{k} z^k (1 - z)^{n-k}$ and
 $\pi(z) = \lim_{m \to \infty} \sum_m P(zm, m).$
 $\pi(z)$ is an example of a prior density.

The Binomial Distribution

What are we to make of the prior density $\pi(z)$?

We ask our friendly theorist for a prediction of the relative frequency of Higgs boson production at the LHC. She predicts that it is *p*.

We might consider modeling that prediction by setting $\pi(z) = \delta(z - p)$ in de Finetti's theorem. If we do so, we obtain the binomial distribution

$$P(k,n) = \mathrm{binomial}(k,n,p)$$



"Ohhhhhhh . . . Look at that, Schuster . . . Dags are so cute when they try to comprehend quantum mechanics."

The Poisson Distribution

There are many situations in which the count n in the binomial distribution

binomial
$$(k, n, p) = \binom{n}{k} p^k (1-p)^{n-k},$$

is large and the probability p is small. For example, at 13 TeV only one in 10^{10} proton-proton collisions leads to a Higgs boson event. Consider, therefore, the limit $n \to \infty$ and $p \to 0$ while a = pn and k remain fixed.

Exercise 2.2

Show that in this limit the binomial distribution becomes the Poisson distribution, $Poisson(k, a) = a^k \exp(-a)/k!$

The Poisson Distribution

The more fundamental definition of a Poisson distribution is via a stochastic model.

Suppose that at time t + dt we have recorded k counts and that in the time interval (t, t + dt) only two things can happen:

- no event occurred during (t, t + dt) or
- one event occurred during (t, t + dt).

We further suppose that the probability to get an event during the time interval (t, t + dt) is proportional to its size dt.

We can now assign probabilities.

The Poisson Distribution

Here are the transition probabilities that define the Poisson model:

 $P_k(t + dt) = ext{probability that the count is } k ext{ at time } t + dt$ $P_k(t) = ext{probability that the count is } k ext{ at time } t$ $P_{k-1}(t) = ext{probability that the count is } k - 1 ext{ at time } t$ $qdt = ext{probability to record 1 event during } t + dt$ $1 - qdt = ext{probability to record 0 events during } t + dt$

In principle, *q* could depend on time. Using the probability rules, we can write

$$P_k(t+dt) = (1-qdt) P_k(t) + qdt P_{k-1}(t),$$

or noting that $dP_k(t)/dt = [P_k(t+dt) - P_k(t)]/dt$,

$$\frac{dP_k}{dt} = -q P_k + q P_{k-1}.$$

The Poisson Distribution Equations such as

$$\frac{dP_k}{dt} = -q \, P_k + q \, P_{k-1},$$

can be solved recursively.

Exercise 2.3

Show that

$$P_k(t) = ext{Poisson}(k, a) = rac{e^{-a}a^k}{k!},$$

where the mean count is a = qt. Also, show that $\operatorname{Var}_k = a$, an important fact about the Poisson distribution that justifies the statement that for a mean count a we would expect counts k to fluctuate by roughly $\pm \sqrt{a}$.

A widely used model in particle physics, astronomy, and cosmology is the multi-Poisson model defined by

$$P(\mathbf{k}|\mathbf{a}) = \prod_{m=1}^{M} rac{a_m^{k_m} e^{-a_m}}{k_m!}.$$

This is the standard statistical model for binned data when the counts are conditionally independent.

In particle physics, analyses that use this model are sometimes referred to as shape analyses.

Exercise 2.4

Show that

$$P(\mathbf{k}|\mathbf{a}) = \text{multinomial}(k_1, \cdots, k_m, p_1, \cdots, p_m) Poisson(k, a), \text{ where}$$

$$k = \sum_{m=1}^{\infty} k_m, \quad a = \sum_{m=1}^{\infty} a_m, \quad p_m = \frac{a_m}{a}, \quad \sum_{m=1}^{\infty} p_m = 1,$$

and the multinomial distribution is given by

multinomial
$$(k_1, \cdots, k_m, p_1, \cdots, p_m) \equiv \binom{k}{k_1, \cdots, k_m} \prod_{m=1}^M p_m^{k_m}$$

When bin counts are large, or when they have a large dynamic range, typical of steeply falling spectra, it is sometimes convenient to drop the Poisson term and rely solely on the multinomial.

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Gaussian Distribution

The probability density function of the Gaussian distribution is

Gauss
$$(x, \mu, \sigma) = \frac{e^{-\frac{1}{2}(x-\mu)^2/\sigma^2}}{\sigma\sqrt{2\pi}}$$

It has mean μ and variance σ^2 . The other oft-used properties are the probability contents of various intervals. Let $z = (x - \mu)/\sigma$. Then

$$P(z \in [-1.00, 1.00]) = 0.683$$

$$P(z \in [-1.64, 1.64]) = 0.900$$

$$P(z \in [-1.96, 1.96]) = 0.950$$

$$P(z \in [-2.58, 2.58]) = 0.990$$

$$P(z \in [-3.29, 3.29]) = 0.999$$

$$P(z \in [5.00, \infty)) = 2.7 \times 10^{-7}$$

The Gaussian, or normal, distribution is the most important distribution in statistics.

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Gaussian Distribution

A bumper sticker: All sensible probability distributions approach a Gaussian in some limit. The precise statement is the central limit theorem.

Example (2.2 The Central Limit Theorem)

Consider the average $t = \frac{1}{n} \sum_{i=1}^{n} x_i$, where $x_i \sim p(\mu, \sigma)$ and $p(\mu, \sigma)$ is any probability density with finite mean μ and standard deviation σ .

Define the standardized variable $z = \sqrt{n(t - \mu)}/\sigma$. The mean of the probability density of z, p(z), is 0 and its standard deviation is 1. The central limit theorem states

$$\lim_{n\to\infty}p(z< x)=\int_{-\infty}^x \operatorname{Gauss}(X,0,1)\,dX\,.$$

When measurement errors can be modeled as the sum of a large number of random contributions, we expect, and this is borne out in practice, the probability density of these errors to be roughly Gaussian. χ^2 Distribution Write $z = (x - \mu)/\sigma$, where $x \sim \text{Gaussian}(\mu, \sigma)$ (~ means "is sampled from") and consider the sum

$$t=\sum_{i=1}^n z_i^2.$$

What is the probability density function (pdf) of t? For any well-behaved probability density function, $p(z_1, \dots, z_n)$, the pdf of t, p(t), is given by the random variable theorem¹

$$p(t) = \int dz_1 \cdots \int dz_n \,\delta\left(t - g(z_1, \cdots, z_n)\right) \,p(z_1, \cdots, z_n)$$

where $g(z_1, \dots, z_n)$ is the function, such as the sum above, that maps z_1 to z_n to t. The δ -function imposes the constraint $t = g(z_1, \dots, z_n)$.

¹A theorem for physicists in the theory of random variables, D. Gillespie, Am. J. of Phys. **51**, 520 (1983).

χ^2 Distribution

First note that $p(z_1, \dots, z_n) = p(z_1)p(z_2)\cdots p(z_n)$ and

$$\delta(x) = \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{i\omega x} \, d\omega.$$

(Burn this formula into your brain...it's one of the most useful in physics and statistics!) Putting together the pieces and shuffling the order of integration, we get

$$egin{aligned} p(t) &= rac{1}{2\pi} \int_{-\infty}^\infty d\omega \, e^{i\omega t} \prod_{j=1}^n \int_{-\infty}^\infty e^{-i\omega z_j^2} \, p(z_j) \, dz_j, \ &= rac{1}{2\pi} \int_{-\infty}^\infty d\omega \, rac{e^{i\omega t}}{(2i)^{n/2}} \, rac{1}{(\omega-i/2)^{n/2}}. \end{aligned}$$

χ^2 Distribution

Writing m = n/2, and performing the integral for integer m, we find

$$\frac{1}{2\pi i} \int_{-\infty}^{\infty} d\omega \, \frac{i e^{i\omega t}}{(2i)^m} \, \frac{1}{(\omega - i/2)^m} = \frac{1}{\Gamma(m)} \frac{t^{m-1} \, e^{-t/2}}{2^m}.$$

This result remains valid for non-integral values of *m*. Therefore, the pdf of the sum of the square of *n* standardized Gaussian variates is $(t = \chi^2)$

$$p(t) = \frac{1}{\Gamma(n/2)} \frac{t^{n/2-1} e^{-t/2}}{2^{n/2}}$$
, mean *n*, variance $2n$

Cauchy Distribution

Let $x, y \sim \text{Gaussian}(0, 1) \equiv g(x)$. What is the pdf of t = y/x? It is given by

$$p(t) = \int_{-\infty}^{\infty} dx \int_{-\infty}^{\infty} dy \,\delta(t - y/x) g(x) g(y),$$

= $\frac{1}{2\pi} \int_{-\infty}^{\infty} dx \int_{-\infty}^{\infty} dy \,\delta(t - y/x) e^{-\frac{1}{2}(x^2 + y^2)}.$

This integral is begging us to use polar coordinates, $y = r \sin \theta$, $x = r \cos \theta$ and $dx dy \rightarrow r dr d\theta$, so that we can write

$$p(t) = \frac{1}{\pi} \left(\int_0^\infty e^{-\frac{1}{2}r^2} r \, dr/2 \right) \int_0^{2\pi} \, \delta(t - \tan \theta) \, d\theta,$$
$$= \frac{1}{\pi} \int_0^{2\pi} \, \delta(t - \tan \theta) \, d\theta.$$

At first glance, the odd looking beast

$$p(t) = rac{1}{\pi} \int_0^{2\pi} \, \delta(t - an heta) \, d heta,$$

looks tricky! But, recall that $\delta(h(\theta)) = \delta(\theta - \theta_0)/|dh/d\theta|$, where θ_0 is the root of $h(\theta)$.

For this problem, $h(\theta) = t - \tan \theta = 0$ and $1/|dh/d\theta| = \cos^2 \theta$. Therefore,

$$egin{aligned} p(t) &= rac{1}{\pi} \int_{0}^{2\pi} \, \delta(heta - heta_0) \, \cos^2 heta \, d heta, \ &= rac{1}{\pi} \, \cos^2 heta_0. \end{aligned}$$

But, $\tan \theta = t \implies \cos \theta = 1/\sqrt{1+t^2}$. Therefore,

$$p(t)=rac{1}{\pi(1+t^2)}$$

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Summary

- According to Kolmogorov, probabilities are functions defined on suitable sets, have range [0, 1], and follow simple rules.
- The two most common interpretations are: relative frequency and degree of belief.
- If it is possible to decompose experimental outcomes (basically, a set of *n*-tuples) into outcomes considered equally likely, then the probability of an outcome may be taken to be the ratio of the number of favorable outcomes to that of all possible outcomes.
- More generally, we use probability functions; probability mass functions for discrete distributions and probability densities for continuous ones.
- And **no**, the probability distributions we use do not come from Nature! We create them through mathematical reasoning. Our hope, however, is that the distributions we create are adequate models of the data generation mechanisms.