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Insights into the pion production mechanism and the symmetry energy at high density

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Abstract: The pion production mechanism is explored based on the ultrarelativistic quantum molecular dynamics model (UrQMD). By analyzing the time evolution of the pion production rate and the central region, we find that characteristic density of pion observable is in the region of 1–2.0 times normal density for Au+Au at $E_{beam}=0.4A$ GeV. The sensitivity of $\pi - /\pi +$ to symmetry energy is weakened after 4–5 $N-\Delta-\pi$ loops in the pion production path. The $\pi - /\pi +$ ratio in the reaction near the threshold energies retains its sensitivity to the symmetry energy. By comparing the UrQMD calculations to the FOPI and considering the constraint of symmetry energy from neutron star properties, the slope of symmetry energy L = 54-91.

Mean field and cross section in the UrQMD*

The isocalar part of potential energy density

 $u = \frac{\alpha}{2} \frac{\rho^2}{\rho_0} + \frac{\beta}{\eta+1} \frac{\rho^{\eta+1}}{\rho_0^{\eta}} + \frac{g_{\text{sur}}}{2\rho_0} (\nabla \rho)^2 + \frac{g_{\text{sur}, \text{ iso}}}{\rho_0} \left[\nabla \left(\rho_n - \rho_p \right) \right]^2 + u_{md}.$ The momentum dependence interaction: $v_{md} = t_4 \ln^2 (1 + t_5 (\mathbf{p}_1 - \mathbf{p}_2)^2) \delta(\mathbf{r}_1 - \mathbf{r}_2)$

The isovector part of potential energy density:

$$u_{\text{sym}=} \begin{cases} \left(A\left(\frac{\rho}{\rho_0}\right) + B\left(\frac{\rho}{\rho_0}\right)^{\gamma_s} + C\left(\frac{\rho}{\rho_0}\right)^{5/3} \right) \rho \delta^2 \\ \frac{C_5}{2} \left(\frac{\rho}{\rho_0}\right)^{\gamma_i} \rho \delta^2 \end{cases}$$

In this work, the Δ -mass dependent $N\Delta \rightarrow NN$ cross sections which were recently calculated based on the one-boson exchange model (OBEM) [1], i.e., $\sigma_{N\Delta \rightarrow NN}^{OBEM}(\sqrt{s}, m_{\Delta})$, are adopted in the UrQMD model At the given value of m_{Δ} , it is calculated as

 $\sigma_{N\Delta \to NN}^{\text{OBEM}} \left(\sqrt{s}, m_{\Delta} \right) \\= \frac{1}{1 + \delta_{N_1 N_2}} \frac{1}{64\pi^2} \int \frac{|\mathbf{p}'_{12}|}{\sqrt{s_{34}} \sqrt{s_{12}} |\mathbf{p}'_{34}(m_{\Delta})|} \\\times \left| \mathcal{M}_{N\Delta(m_{\Lambda}) \to NN} \right|^2 d\Omega.$

In which, $\frac{1}{|\mathcal{M}_{N\Delta(m_{\Delta})\to NN}|^2} = \frac{(2s_1+1)(2s_2+1)}{(2s_3+1)(2s_4+1)} \frac{1}{|\mathcal{M}(m_{\Delta})|^2}.$



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The cross section σ_{OBEM} are shown as red lines in Fig(1).b.

Pion production mechanism and its characteristic density



The pion freeze out after 4-5 N- Δ - π loops, which will reduce the sensitivity of π -/ π + to the symmetry energy. The characteristic density of pion has been calculated by eq.1, and results show that its characteristic density is in the range of 1–2.0 times normal density at Ebeam=0.4A GeV.

Effects of $\sigma_{NA \rightarrow NN}$, incompressibility K₀, and symmetry energy on M(π) and $\pi - /\pi +$



Fig.6. The excitation function of the $M(\pi)/A_{part}$ and $\pi/\pi+$ with different N Δ -NN cross section.



Fig.7. The excitation function of the $M(\pi)/A_{part}$ and $\pi/\pi+$ with different incompressibility K₀.



The $\pi - \pi / \pi + ratio$ in the reaction near the threshold energies retains its sensitivity to the symmetry energy, and it is insensitive to the nuclear incompressibility K_0 and effective mass when their values are selected in the commonly accepted range.

Constraints on symmetry energy at high density by $\pi - \pi / \pi +$ and neutron star prosperities.



By comparing the UrQMD calculations to the FOPI data at 0.4*A* GeV and considering the constraint of symmetry energy from neutron star properties, the slope of symmetry energy L = 54-91 MeV and the symmetry energy at two times normal density $S(2\rho_0) = 48-59$ MeV are deduced.