

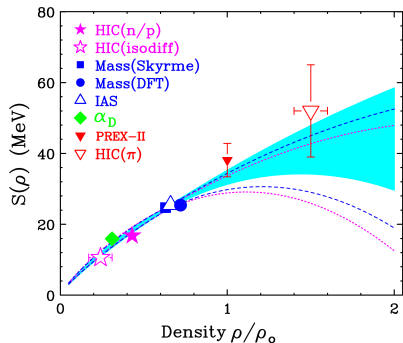
# Light clusters in dynamic evolution of heavy-ion collisions

Akira Ono

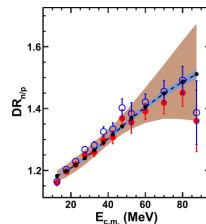
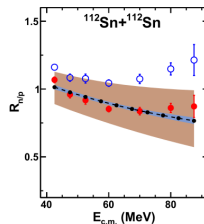
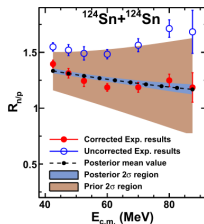
Tohoku University

NuSYM21: International Symposium on Nuclear Symmetry Energy,  
22–24 September & 13–15 October 2021, online

# Symmetry energy constraints and clusters



Lynch, Tsang, arXiv:2106.10119.



★ HIC(n/p) is a result from “coalescence-invariant (CI)” neutron and proton spectra from Sn + Sn central collisions at 120A MeV. [Morfouace et al., PLB 799 (2019) 135045.]

$$"n" = n + {}^2\text{H} + 2 \times {}^3\text{H} + {}^3\text{He} + 2 \times {}^4\text{He}$$

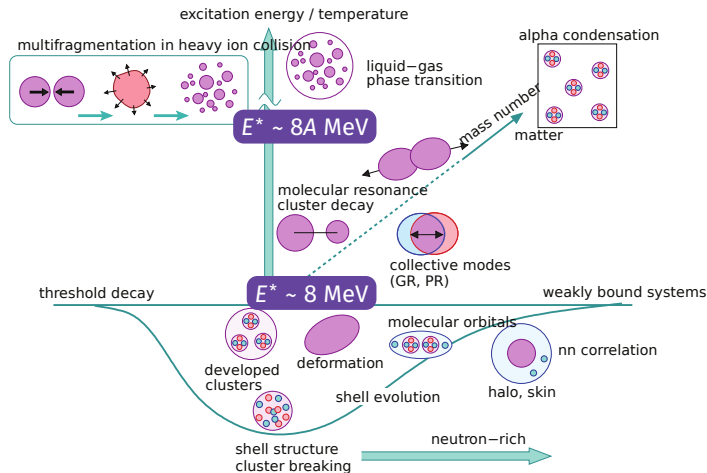
$$"p" = p + {}^2\text{H} + {}^3\text{H} + 2 \times {}^3\text{He} + 2 \times {}^4\text{He}$$

“CI” spectra may (or not?) be affected by cluster formation if many clusters are formed, e.g., through energy conservation.

We have to understand cluster formation.

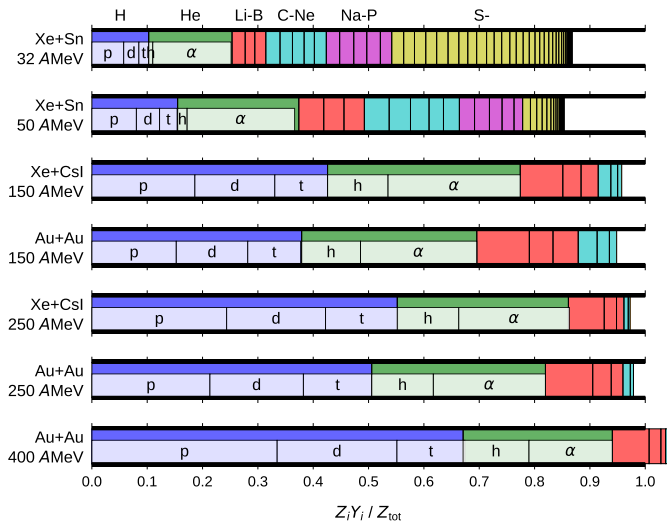
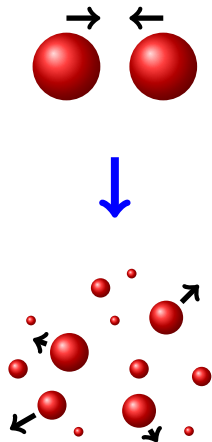
# Clustering phenomena in excited states of nuclear systems

$E^* \sim 80A \text{ MeV}$  Gas of clusters at higher energies



Kanada-En'yo, Kimura, Ono, Prog. Theor. Exp. Phys. 2012 01A202 (2012)

# Fraction of protons in clusters and fragments in heavy-ion collisions



INDRA: Hudan et al., PRC67 (2003) 064613. FOPI: Reisdorf et al., NPA 848 (2010) 366.

Figure in Ono, PPNP 105 (2019) 139.

This is a challenge to theorists and transport models.



- Calculate  $e^{-iHt/\hbar}|\psi(t=0)\rangle$ , and find clusters.
- Cluster correlation in dynamics — Find whether a nucleon moves together with some other nucleons for a while.
  - Clusters will not necessarily be emitted.
  - Correlations affect the time evolution.
- SACA, FRIGA [Le Fèvre et al., PRC 100 (2019) 034904.], for QMD.
- Coalescence prescription, to predict clusters in BUU.
- At large  $t$  (e.g.  $t = 200\text{--}1000$  fm/c), there is nothing controversial in finding clusters and fragments.
  - ..., if the state at this time is predicted by the transport model reasonably well.
  - The decay of excited fragments should be calculated by a statistical decay code.

## BUU with clusters

Danielewicz and Bertsch, NPA 533 (1991) 712.

Coupled equations for  $f_n(\mathbf{r}, \mathbf{p}, t)$ ,  $f_p(\mathbf{r}, \mathbf{p}, t)$ ,  $f_d(\mathbf{r}, \mathbf{p}, t)$ ,  $f_t(\mathbf{r}, \mathbf{p}, t)$ ,  $f_h(\mathbf{r}, \mathbf{p}, t)$  are solved by the test particle method.

$$\frac{\partial f_n}{\partial t} + \mathbf{v} \cdot \frac{\partial f_n}{\partial \mathbf{r}} - \frac{\partial U_n}{\partial \mathbf{r}} \cdot \frac{\partial f_n}{\partial \mathbf{p}} = I_n^{\text{coll}}[f_n, f_p, f_d, f_t, f_h]$$

$$\frac{\partial f_p}{\partial t} + \mathbf{v} \cdot \frac{\partial f_p}{\partial \mathbf{r}} - \frac{\partial U_p}{\partial \mathbf{r}} \cdot \frac{\partial f_p}{\partial \mathbf{p}} = I_p^{\text{coll}}[f_n, f_p, f_d, f_t, f_h]$$

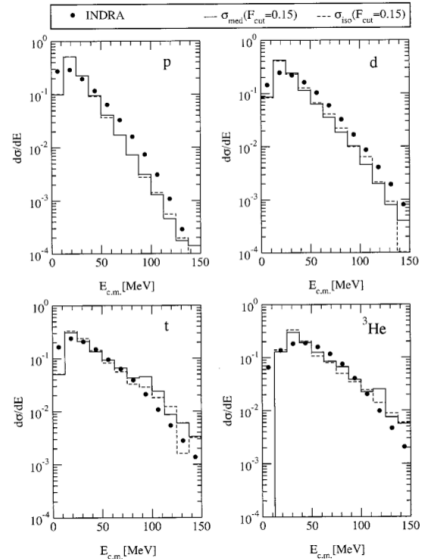
$$\frac{\partial f_d}{\partial t} + \mathbf{v} \cdot \frac{\partial f_d}{\partial \mathbf{r}} - \frac{\partial U_d}{\partial \mathbf{r}} \cdot \frac{\partial f_d}{\partial \mathbf{p}} = I_d^{\text{coll}}[f_n, f_p, f_d, f_t, f_h]$$

$$\frac{\partial f_t}{\partial t} + \mathbf{v} \cdot \frac{\partial f_t}{\partial \mathbf{r}} - \frac{\partial U_t}{\partial \mathbf{r}} \cdot \frac{\partial f_t}{\partial \mathbf{p}} = I_t^{\text{coll}}[f_n, f_p, f_d, f_t, f_h]$$

$$\frac{\partial f_h}{\partial t} + \mathbf{v} \cdot \frac{\partial f_h}{\partial \mathbf{r}} - \frac{\partial U_h}{\partial \mathbf{r}} \cdot \frac{\partial f_h}{\partial \mathbf{p}} = I_h^{\text{coll}}[f_n, f_p, f_d, f_t, f_h]$$

Similar transport models for relativistic collisions:

Oliinychenko et al, PRC 99 (2019) 044907. KJ Sun et al., arXiv:2106.12742 [nucl-th].



Renormalized cluster spectra

Kuhrts et al., PRC63(2001)034605.

## BUU with clusters

Danielewicz and Bertsch, NPA 533 (1991) 712.

Coupled equations for  $f_n(\mathbf{r}, \mathbf{p}, t)$ ,  $f_p(\mathbf{r}, \mathbf{p}, t)$ ,  $f_d(\mathbf{r}, \mathbf{p}, t)$ ,  $f_t(\mathbf{r}, \mathbf{p}, t)$ ,  $f_h(\mathbf{r}, \mathbf{p}, t)$  are solved by the test particle method.

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$$\frac{\partial f_p}{\partial t} + \mathbf{v} \cdot \frac{\partial f_p}{\partial \mathbf{r}} - \frac{\partial U_p}{\partial \mathbf{r}} \cdot \frac{\partial f_p}{\partial \mathbf{p}} = I_p^{\text{coll}}[f_n, f_p, f_d, f_t, f_h]$$

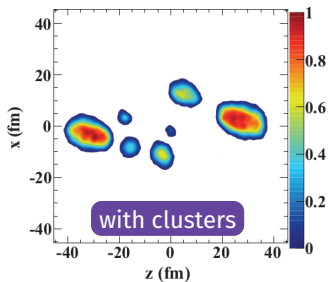
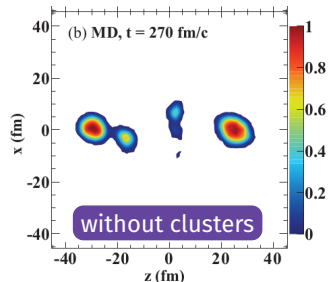
$$\frac{\partial f_d}{\partial t} + \mathbf{v} \cdot \frac{\partial f_d}{\partial \mathbf{r}} - \frac{\partial U_d}{\partial \mathbf{r}} \cdot \frac{\partial f_d}{\partial \mathbf{p}} = I_d^{\text{coll}}[f_n, f_p, f_d, f_t, f_h]$$

$$\frac{\partial f_t}{\partial t} + \mathbf{v} \cdot \frac{\partial f_t}{\partial \mathbf{r}} - \frac{\partial U_t}{\partial \mathbf{r}} \cdot \frac{\partial f_t}{\partial \mathbf{p}} = I_t^{\text{coll}}[f_n, f_p, f_d, f_t, f_h]$$

$$\frac{\partial f_h}{\partial t} + \mathbf{v} \cdot \frac{\partial f_h}{\partial \mathbf{r}} - \frac{\partial U_h}{\partial \mathbf{r}} \cdot \frac{\partial f_h}{\partial \mathbf{p}} = I_h^{\text{coll}}[f_n, f_p, f_d, f_t, f_h]$$

Similar transport models for relativistic collisions:

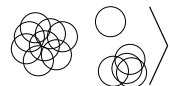
Oliinychenko et al, PRC 99 (2019) 044907. KJ Sun et al., arXiv:2106.12742 [nucl-th].



Isospin diffusion and fragmentation

Coupland et al., PRC 84 (2011) 054603.

# Antisymmetrized Molecular Dynamics (very basic version)



AMD wave function

$$|\Phi(Z)\rangle = \det_{ij} \left[ \exp\left\{-v\left(\mathbf{r}_j - \frac{\mathbf{Z}_i}{\sqrt{v}}\right)^2\right\} \chi_{\alpha_i}(j) \right]$$

$$\mathbf{Z}_i = \sqrt{v}\mathbf{D}_i + \frac{i}{2\hbar\sqrt{v}}\mathbf{K}_i$$

$v$  : Width parameter =  $(2.5 \text{ fm})^{-2}$

$\chi_{\alpha_i}$  : Spin-isospin states =  $p \uparrow, p \downarrow, n \uparrow, n \downarrow$

Equation of motion for the wave packet centroids  $Z$

$$\frac{d}{dt}\mathbf{Z}_i = \{\mathbf{Z}_i, H\}_{\text{PB}} + (\text{NN collisions})$$

$\{\mathbf{Z}_i, H\}_{\text{PB}}$ : Motion in the mean field

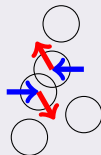
$$H = \frac{\langle \Phi(Z) | H | \Phi(Z) \rangle}{\langle \Phi(Z) | \Phi(Z) \rangle} + (\text{c.m. correction})$$

$H$ : Effective interaction (e.g. Skyrme force)

NN collisions

$$W_{i \rightarrow f} = \frac{2\pi}{\hbar} |\langle \psi_f | V | \psi_i \rangle|^2 \delta(E_f - E_i)$$

- $|V|^2$  or  $\sigma_{NN}$  (in medium)
- Pauli blocking



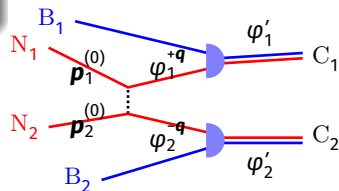
Ono, Horiuchi et al., Prog. Theor. Phys. 87 (1992) 1185.



# NN collisions with cluster correlations



- $N_1, N_2$  : Colliding nucleons
- $B_1, B_2$  : Spectator nucleons/clusters
- $C_1, C_2$  :  $N, (2N), (3N), (4N)$  (up to  $\alpha$  cluster)



## Transition probability

$$W(\text{NBNB} \rightarrow \text{CC}) = \frac{2\pi}{\hbar} |\langle \text{CC} | V | \text{NBNB} \rangle|^2 \delta(E_f - E_i)$$

$$vd\sigma \propto |\langle \varphi'_1 | \varphi_1^{+q} \rangle|^2 |\langle \varphi'_2 | \varphi_2^{-q} \rangle|^2 |M|^2 \delta(E_f - E_i) p_{\text{rel}}^2 dp_{\text{rel}} d\Omega$$

$|M|^2 = |\langle \text{NN} | V | \text{NN} \rangle|^2$ : Matrix elements of NN scattering

$\Leftarrow (d\sigma/d\Omega)_{\text{NN}}$  in free space (or in medium)

$$\mathbf{p}_{\text{rel}} = \frac{1}{2}(\mathbf{p}_1 - \mathbf{p}_2) = p_{\text{rel}} \hat{\Omega}$$

$$\mathbf{p}_1 = \mathbf{p}_1^{(0)} + \mathbf{q}$$

$$\mathbf{p}_2 = \mathbf{p}_2^{(0)} - \mathbf{q}$$

$$\varphi_1^{+q} = \exp(+i\mathbf{q} \cdot \mathbf{r}_{N_1}) \varphi_1^{(0)}$$

$$\varphi_2^{-q} = \exp(-i\mathbf{q} \cdot \mathbf{r}_{N_2}) \varphi_2^{(0)}$$

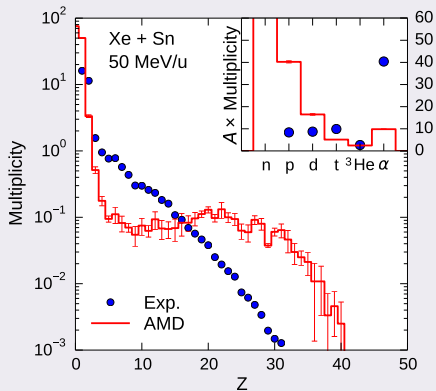
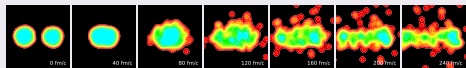
Ono, J. Phys. Conf. Ser. 420 (2013) 012103.

Ikeno, Ono et al., PRC 93 (2016) 044612.

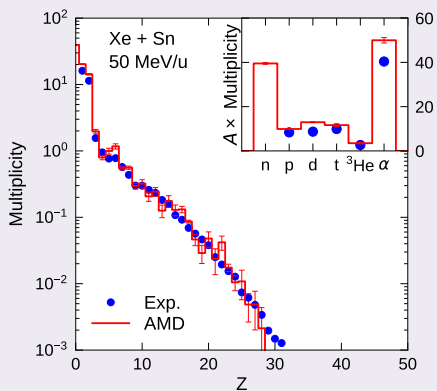
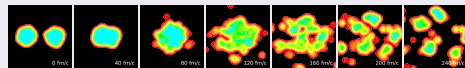
Ono, JPS Conference Proceedings 32 (2020) 010076.

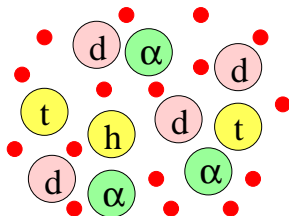
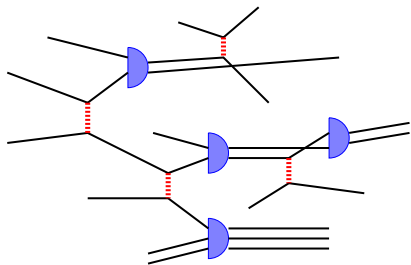
# Effect of cluster correlations: central Xe + Sn at 50 MeV/u

## Without clusters



## With clusters





- A cluster in AMD is still composed of nucleon wave packets. The many-body state is a Slater determinant of nucleons.
- Clusters are not only created but also broken.
- Clusters are in medium, so the existence and the properties may be modified.
- For a collision, there are many possible configurations ( $C_1, C_2$ ) of cluster formation. Non-orthogonality is treated suitably.
- Correlations to bind several clusters (like  $\alpha + t = {}^7\text{Li}$ ) are also important and taken into account.

- $n + p + X \leftrightarrow d + X'$
- $d + n + X \leftrightarrow t + X'$
- $d + p + X \leftrightarrow h + X'$
- $t + p + X \leftrightarrow \alpha + X'$
- $h + n + X \leftrightarrow \alpha + X'$
- $d + d + X \leftrightarrow \alpha + X'$
- $2n + p + X \leftrightarrow t + X'$
- $n + 2p + X \leftrightarrow h + X'$
- $d + n + p + X \leftrightarrow \alpha + X'$
- $2n + 2p + X \leftrightarrow \alpha + X'$
- $d + d \leftrightarrow p + t$
- $d + d \leftrightarrow n + h$
- $p + t \leftrightarrow n + h$
- $d + t \leftrightarrow n + \alpha$
- $d + h \leftrightarrow p + \alpha$
- $d + t \leftrightarrow 2n + h$
- $d + h \leftrightarrow 2p + t$
- $d + \alpha \leftrightarrow t + h$

## Finding a cluster in a many-body state (in general)

One proton( $\uparrow$ ) and one neutron( $\uparrow$ ) (both described by Gaussian wave packets)

$$\left| \begin{array}{c} \text{p} \\ \text{n} \end{array} \right\rangle = c_d \left| \text{d} \right\rangle + c' \left| \begin{array}{c} \text{p} \\ \text{n} \end{array} \right\rangle + \dots$$

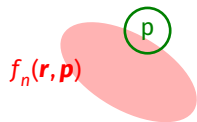
$$\varphi_p(\mathbf{r}_1) \varphi_n(\mathbf{r}_2) = \left[ c_d \psi_d(\mathbf{r}) + (\text{continuum states}) \right] \varphi_{\text{cm}}(\mathbf{R}), \quad \mathbf{r} = \mathbf{r}_1 - \mathbf{r}_2, \quad \mathbf{R} = \frac{1}{2}(\mathbf{r}_1 + \mathbf{r}_2)$$

$$P_d(\mathbf{P}) = \left| \langle e^{i\mathbf{P}\cdot\mathbf{R}/\hbar} \psi_d | \varphi_p \varphi_n \rangle \right|^2 = \int \frac{d\mathbf{r}_1 d\mathbf{r}_2 d\mathbf{p}}{(2\pi\hbar)^3} \rho_d^W(\mathbf{r}_1 - \mathbf{r}_2, \mathbf{p}) f_p(\mathbf{r}_1, \frac{1}{2}\mathbf{P} + \mathbf{p}) f_n(\mathbf{r}_2, \frac{1}{2}\mathbf{P} - \mathbf{p})$$

One proton( $\uparrow$ ) and many neutrons( $\uparrow$ )

$$P_d(\mathbf{P}) \stackrel{???}{=} \int \frac{d\mathbf{r}_1 d\mathbf{r}_2 d\mathbf{p}}{(2\pi\hbar)^3} \rho_d^W(\mathbf{r}_1 - \mathbf{r}_2, \mathbf{p}) f_p(\mathbf{r}_1, \frac{1}{2}\mathbf{P} + \mathbf{p}) f_n(\mathbf{r}_2, \frac{1}{2}\mathbf{P} - \mathbf{p})$$

LW Chen, CM Ko, BA Li, NPA 729 (2003) 809. Mattiello et al., PRC 55 (1997) 1443.



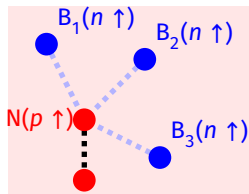
This is valid only in the dilute limit. In general, e.g, the “probability”

$$N_d = \int \frac{d\mathbf{P}}{(2\pi\hbar)^3} P_d(\mathbf{P}) \quad \text{can be} > 1.$$

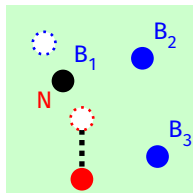
- Identifying clusters in a many-body system is a fundamental problem.
- Another question is how the identified clusters are propagated.

# Construction of Final States in AMD/Cluster

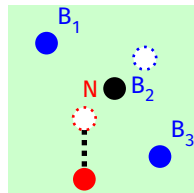
Clusters (in the final states) are assumed to have  $(0s)^N$  configuration.



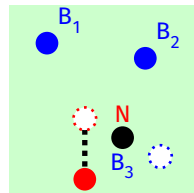
$|\Phi^q\rangle$   
After  $\mathbf{p}^{(0)} \rightarrow \mathbf{p}^{(0)} + \mathbf{q}$



$|\Phi'_1\rangle$   
 $N + B_1 \rightarrow C_1$



$|\Phi'_2\rangle$   
 $N + B_2 \rightarrow C_2$



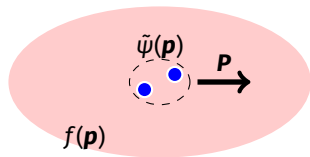
$|\Phi'_3\rangle$   
 $N + B_3 \rightarrow C_3$

Final states are not orthogonal:  $N_{ij} \equiv \langle \Phi'_i | \Phi'_j \rangle \neq \delta_{ij}$

The probability of cluster formation with one of B's:

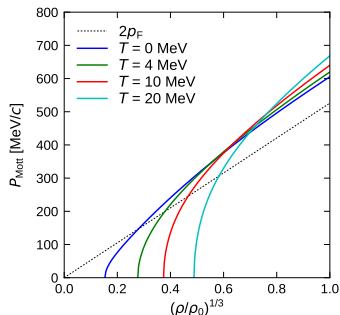
$$\hat{P} = \sum_{ij} |\Phi'_i\rangle N_{ij}^{-1} \langle \Phi'_j|, \quad P = \langle \Phi^q | \hat{P} | \Phi^q \rangle \neq \sum_i |\langle \Phi'_i | \Phi^q \rangle|^2$$

- $\left\{ \begin{array}{l} P \\ 1 - P \end{array} \right. \Rightarrow$  Choose one of the candidates and make a cluster.
- $\left\{ \begin{array}{l} P \\ 1 - P \end{array} \right. \Rightarrow$  Don't make a cluster (with any  $n \uparrow$ ).



## Equation for a deuteron in uncorrelated medium

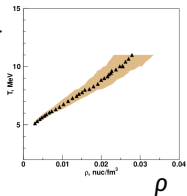
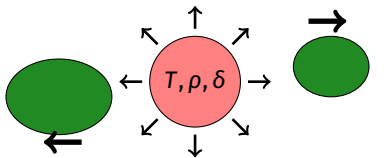
$$\begin{aligned} & \left[ e(\frac{1}{2}\mathbf{P} + \mathbf{p}) + e(\frac{1}{2}\mathbf{P} - \mathbf{p}) \right] \tilde{\psi}(\mathbf{p}) \\ & + \left[ 1 - f(\frac{1}{2}\mathbf{P} + \mathbf{p}) - f(\frac{1}{2}\mathbf{P} - \mathbf{p}) \right] \int \frac{d\mathbf{p}'}{(2\pi)^3} \langle \mathbf{p} | v | \mathbf{p}' \rangle \tilde{\psi}(\mathbf{p}') \\ & = E \tilde{\psi}(\mathbf{p}) \end{aligned}$$



Formula from Röpke, NPA867 (2011) 66.

- A bound deuteron cannot exist inside the Fermi sphere, except at very low densities.
- A deuteron can exist if its momentum is high enough.
- When there is no bound solution, is it still possible that nucleons are correlated in the continuum like a resonance?
- What happens when the medium is correlated?

# Clusters in low-density medium (experiments)



Qin et al.,  
PRL108(2012)172701  
Ar, Zn + Sn @ 47A MeV

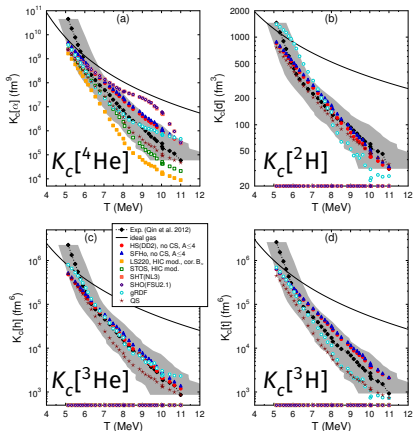
## Assumption

The particles emitted at the same velocity  $v_{surf}$  were emitted at the same time from the same source characterized by  $(T, \rho, \delta)$ .

## Equilibrium Constants

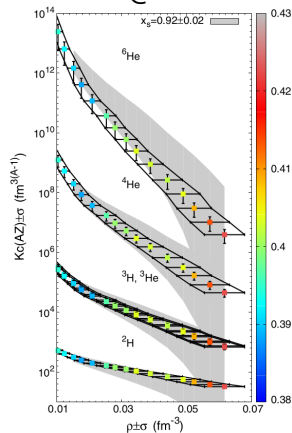
$$K_c(N, Z) = \frac{\rho(N, Z)}{\rho_p^Z \rho_n^N} \quad \text{for cluster } (N, Z)$$

Hempel et al., PRC 91 (2015) 045805.



Paris, Bougault et al.,  
PRL 125 (2020) 012701.

Xe + Sn @ 32A MeV



## Condition to switch on/off clusters

### With or without clusters

$$N_1 + N_2 + B_1 + B_2 \rightarrow C_1 + C_2 \quad \text{or} \quad N_1 + N_2 \rightarrow N_1 + N_2$$

The condition to switch on clusters

$$\rho' < \rho_c, \quad \rho_c = 0.125 \text{ fm}^{-3} \text{ or } 0.060 \text{ fm}^{-3} \text{ etc.}$$

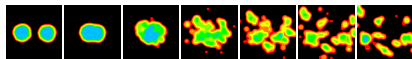
Density with a momentum cut for the nucleon  $N_i$  ( $i = 1, 2$ )

$$\begin{aligned}\rho_i'^{(ini)} &= \left(\frac{2V}{\pi}\right)^{\frac{3}{2}} \sum_{k(\neq i)} \theta(p_{\text{cut}} > |\mathbf{P}_i - \mathbf{P}_k|) e^{-2\nu(\mathbf{R}_i - \mathbf{R}_k)^2} \\ \rho_i'^{(\text{fin})} &= \left(\frac{2V}{\pi}\right)^{\frac{3}{2}} \sum_{k(\neq i)} \theta(p_{\text{cut}} > |\mathbf{P}_i^{(\text{fin})} - \mathbf{P}_k|) e^{-2\nu(\mathbf{R}_i - \mathbf{R}_k)^2} \\ \rho' &= (\rho_1'^{(ini)} \rho_1'^{(\text{fin})} \rho_2'^{(ini)} \rho_2'^{(\text{fin})})^{\frac{1}{4}}\end{aligned}$$

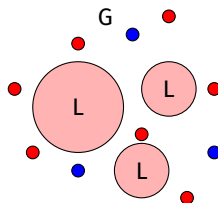
An energy-dependent momentum cut was chosen,  $p_{\text{cut}} = (375 \text{ MeV}/c) e^{-\epsilon/(225 \text{ MeV})}$ , where  $\epsilon$  is the collision energy (i.e. the sum of the kinetic energies of  $N_1$  and  $N_2$  in their c.m. frame).



# Liquid-Gas separation in fragmentation reactions



In a late stage of reaction



## Fractionation/Distillation

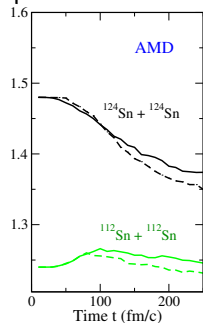
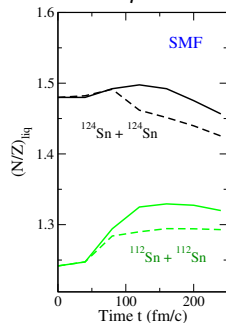
$$\frac{N}{Z} (\text{Liquid}) < \frac{N}{Z} (\text{Gas})$$

- Gas =  $\sum(A \leq 4 \text{ particles})$
- Liquid =  $\sum(A > 4 \text{ fragments})$
- Total = Gas + Liquid

$^{124}\text{Sn} + ^{124}\text{Sn}$  and  $^{112}\text{Sn} + ^{112}\text{Sn}$ ,  $E/A = 50 \text{ MeV}$ ,  $b \approx 0$ .

Colonna and Ono, PRC 82 (2010) 054613.

## N/Z of the Liquid Part

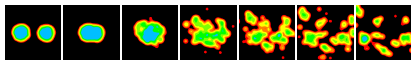


Stiff  $S(\rho)$   
Soft  $S(\rho)$

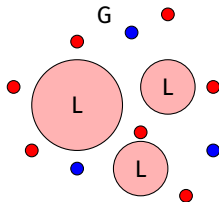
## Symmetry energy effects in Fermi energy domain

- Fractionation and related dynamical effects
- Isospin drift
- Isospin diffusion

# Liquid-Gas separation in fragmentation reactions



In a late stage of reaction



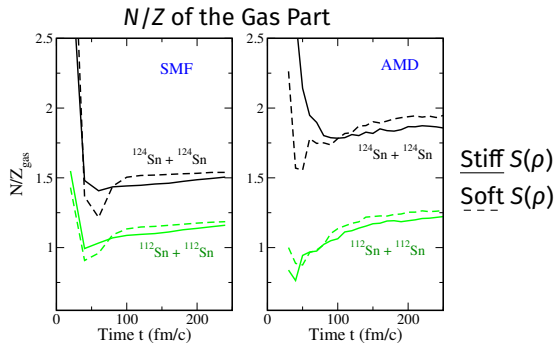
## Fractionation/Distillation

$$\frac{N}{Z} (\text{Liquid}) < \frac{N}{Z} (\text{Gas})$$

- Gas =  $\sum(A \leq 4 \text{ particles})$
- Liquid =  $\sum(A > 4 \text{ fragments})$
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$^{124}\text{Sn} + ^{124}\text{Sn}$  and  $^{112}\text{Sn} + ^{112}\text{Sn}$ ,  $E/A = 50 \text{ MeV}$ ,  $b \approx 0$ .

Colonna and Ono, PRC 82 (2010) 054613.

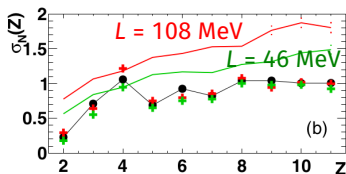
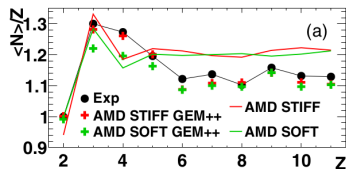


## Symmetry energy effects in Fermi energy domain

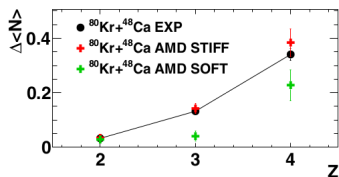
- Fractionation and related dynamical effects
- Isospin drift
- Isospin diffusion

# Isospin drift (+ isospin diffusion + isospin fractionation)

Fragments in  $v_{cm} < v < v_{QP}$

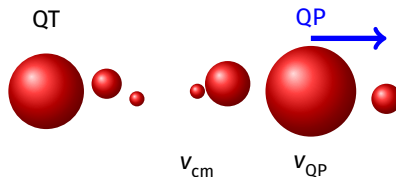


$(v_{cm} < v < v_{QP}) - (v > v_{QP})$



Piantelli et al. (FAZIA Collaboration), PRC 103 (2021) 014603.

$^{80}\text{Kr} + ^{48}\text{Ca}$  at 35 MeV/nucleon (INFN-LNS)



- In the width of the isotope distribution  $\sigma_N(Z)$ , clear dependence on the stiffness of the symmetry energy is found, for the excited fragments produced in AMD.
  - After statistical decays calculated by GEMINI++, the sensitivity in  $\sigma_N(Z)$  becomes weaker.
- For the final He, Li and Be fragments (emitted backward in the QP frame), the yields of neutron-rich isotopes are sensitive to the stiffness of the symmetry energy.

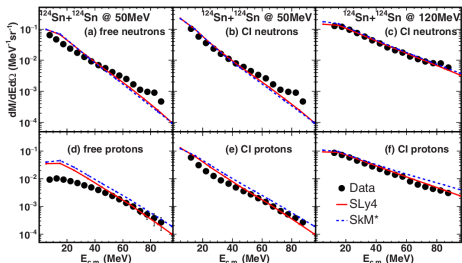
$L = 108$  MeV (strong isospin drift) is more favored than  $L = 46$  MeV.

# Spectra and ratios of emitted nucleons and clusters

MSU data:

Coupland et al., PRC 96 (2016) 011601(R).

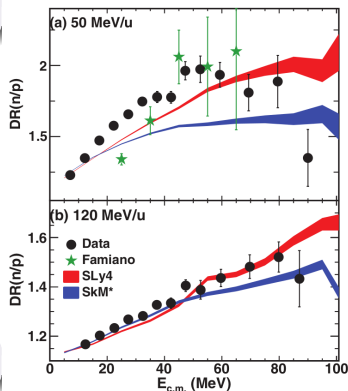
Sn + Sn,  $E/A = 50$  and  $120$  MeV,  $b \leq 3$  fm



CI = (free nucleons) + (nucleons in clusters)

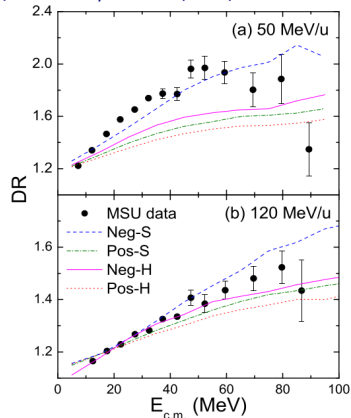
$$DR = \frac{“n/p” \text{ in } ^{124}\text{Sn} + ^{124}\text{Sn}}{“n/p” \text{ in } ^{112}\text{Sn} + ^{112}\text{Sn}}$$

compared with ImQMD



DR compared with an IQMD

J. Su et al., PRC 94 (2016) 034619.



Similar result in

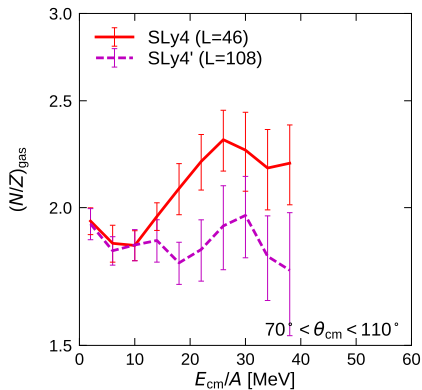
YX Zhang et al., PLB 732 (2014) 186.

Bayesian analyses at  $E/A = 120$  MeV by Morfouace et al., PLB 799 (2019) 135045.

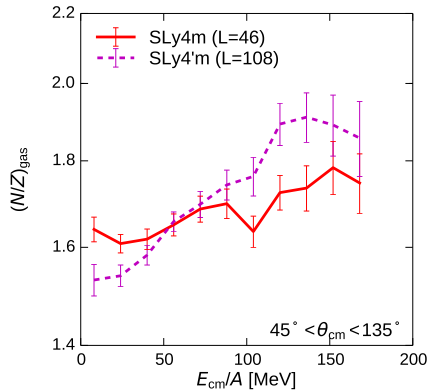
- $(m_n^* - m_p^*)/m_N = -0.05_{-0.09}^{+0.09} (\rho_n - \rho_p)/\rho$
- $S(\rho_s) = 16.8_{-1.2}^{+1.2}$  MeV at  $\rho_s/\rho_0 = 0.43_{-0.05}^{+0.05}$

# N/Z Ratio at 50 and 300 MeV/u calculated by AMD with clusters

$^{124}\text{Sn} + ^{124}\text{Sn}$ ,  $E/A = 50$  MeV,  $b \approx 0$



$^{132}\text{Sn} + ^{124}\text{Sn}$ ,  $E/A = 300$  MeV,  $b \approx 0$



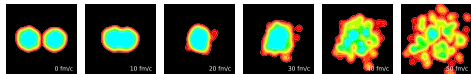
Low density effect  $\iff$  High density effect

$$\left(\frac{N}{Z}\right)_{\text{gas}} = \frac{Y_n(\epsilon) + Y_d(\epsilon) + 2Y_t(\epsilon) + Y_h(\epsilon) + 2Y_\alpha(\epsilon)}{Y_p(\epsilon) + Y_d(\epsilon) + Y_t(\epsilon) + 2Y_h(\epsilon) + 2Y_\alpha(\epsilon)}, \quad \epsilon = E_{\text{cm}}/A$$

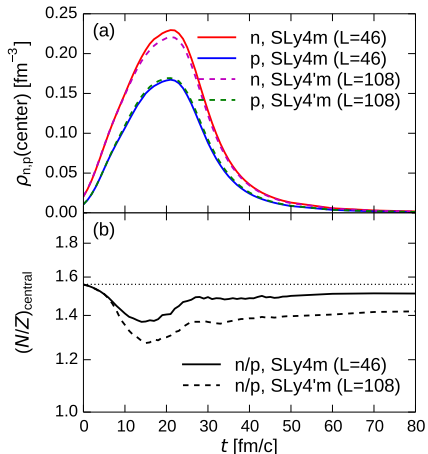
For 50A MeV, the  $(N/Z)_{\text{gas}}$  ratio in the AMD calculation seems to be higher than in QMD calculations.

However, the nucleon spectra (before taking ratio) need to be understood better in calculations to draw a definite conclusion.

# Compression and expansion in collisions at 300 MeV/nucleon



$^{132}\text{Sn} + ^{124}\text{Sn}$ ,  $E/A = 300$  MeV,  $b \sim 0$



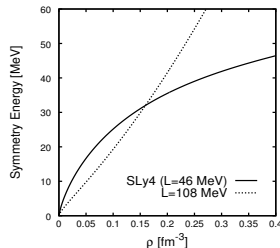
Momentum dependence of Skyrme (SLy4) interaction has been corrected for high energy collisions, in a similar way to [Gale, Bertsch, Das Gupta, PRC 35 \(1987\) 1666](#).

## Nuclear EOS (at $T = 0$ )

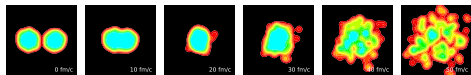
$$(E/A)(\rho_p, \rho_n) = (E/A)_0(\rho) + S(\rho)\delta^2 + \dots$$

$$\rho = \rho_p + \rho_n, \quad \delta = \frac{\rho_n - \rho_p}{\rho_n + \rho_p}$$

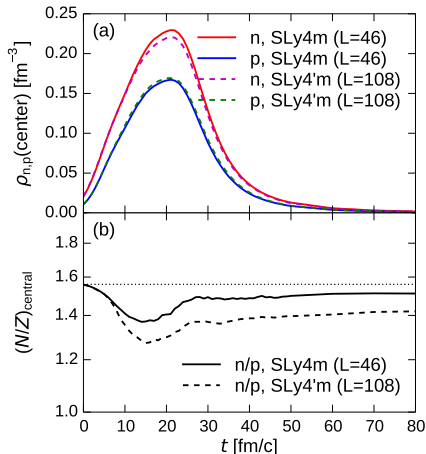
- $S_0 = S(\rho_0)$
- $L = 3\rho_0(dS/d\rho)_{\rho=\rho_0}$



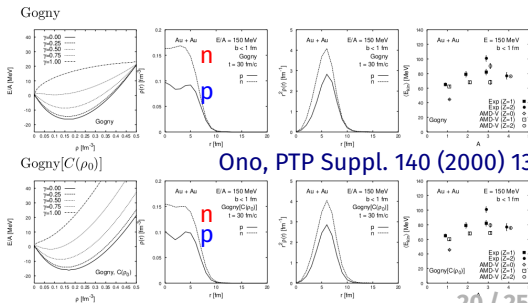
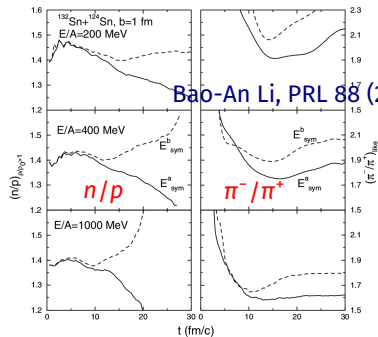
# Compression and expansion in collisions at 300 MeV/nucleon



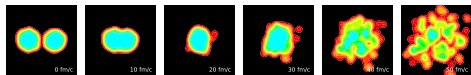
$^{132}\text{Sn} + ^{124}\text{Sn}, E/A = 300 \text{ MeV}, b \sim 0$



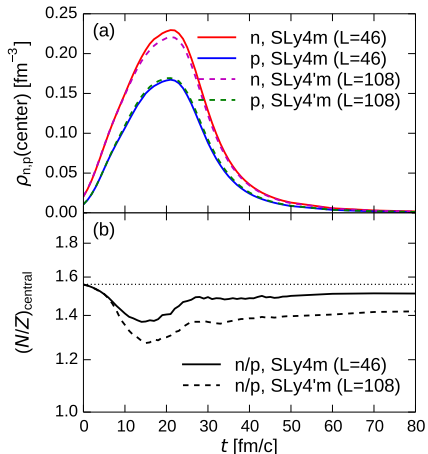
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# Compression and expansion in collisions at 300 MeV/nucleon



$^{132}\text{Sn} + ^{124}\text{Sn}$ ,  $E/A = 300$  MeV,  $b \sim 0$

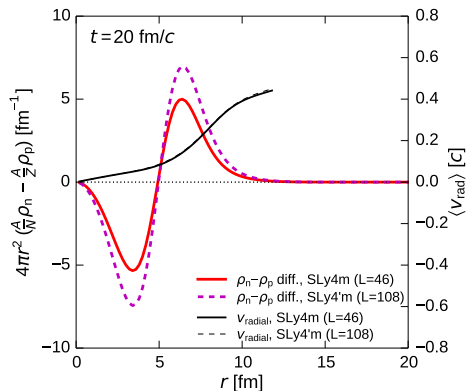


Momentum dependence of Skyrme (SLy4) interaction has been corrected for high energy collisions, in a similar way to [Gale, Bertsch, Das Gupta, PRC 35 \(1987\) 1666](#).

- Neutron-proton density diff. (fn of  $r$ )

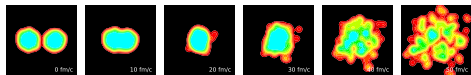
$$4\pi r^2 \left[ \frac{A}{N} \rho_n(r) - \frac{A}{Z} \rho_p(r) \right]$$

- Radial expansion velocity  $v_{\text{rad}}(r)$

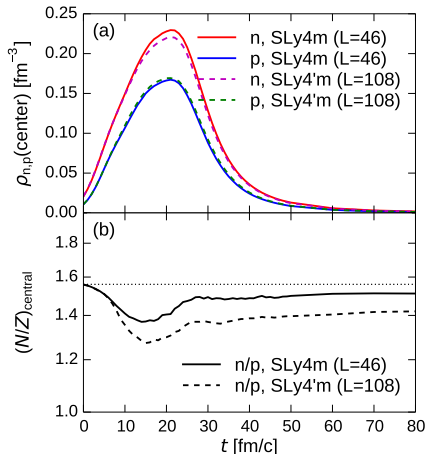




# Compression and expansion in collisions at 300 MeV/nucleon



$^{132}\text{Sn} + ^{124}\text{Sn}$ ,  $E/A = 300$  MeV,  $b \sim 0$

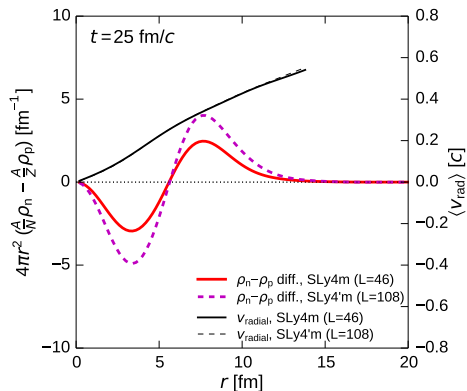


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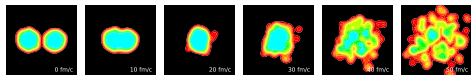
- Neutron-proton density diff. (fn of  $r$ )

$$4\pi r^2 \left[ \frac{A}{N} \rho_n(r) - \frac{A}{Z} \rho_p(r) \right]$$

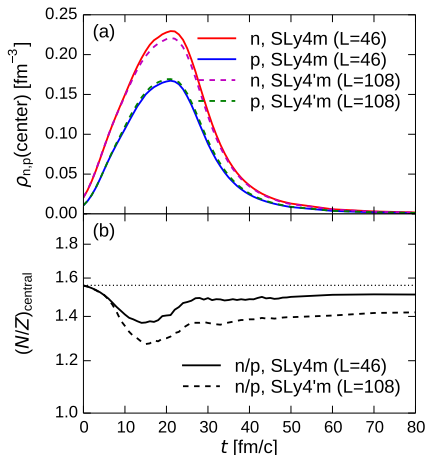
- Radial expansion velocity  $v_{\text{rad}}(r)$



# Compression and expansion in collisions at 300 MeV/nucleon



$^{132}\text{Sn} + ^{124}\text{Sn}$ ,  $E/A = 300$  MeV,  $b \sim 0$

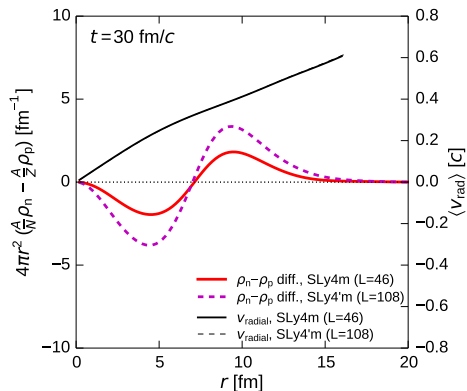


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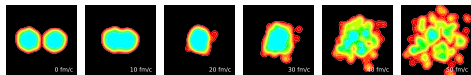
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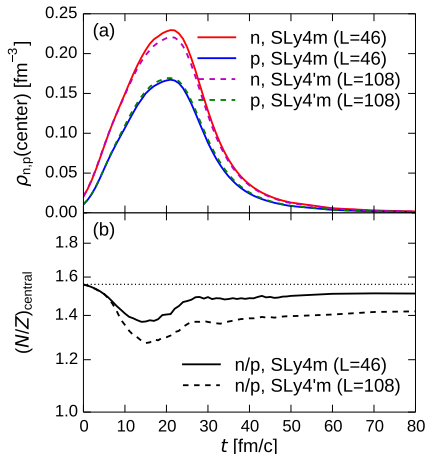
- Radial expansion velocity  $v_{\text{rad}}(r)$



# Compression and expansion in collisions at 300 MeV/nucleon



$^{132}\text{Sn} + ^{124}\text{Sn}$ ,  $E/A = 300$  MeV,  $b \sim 0$

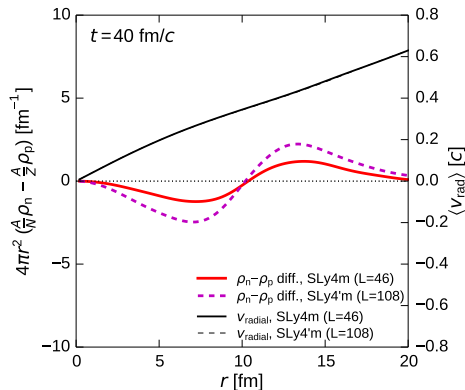


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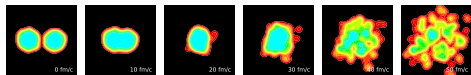
- Neutron-proton density diff. (fn of  $r$ )

$$4\pi r^2 \left[ \frac{A}{N} \rho_n(r) - \frac{A}{Z} \rho_p(r) \right]$$

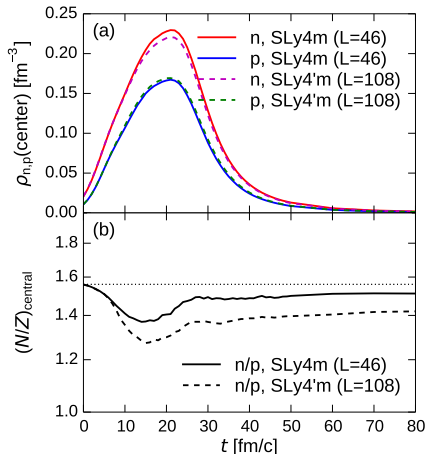
- Radial expansion velocity  $v_{\text{rad}}(r)$



# Compression and expansion in collisions at 300 MeV/nucleon



$^{132}\text{Sn} + ^{124}\text{Sn}$ ,  $E/A = 300$  MeV,  $b \sim 0$

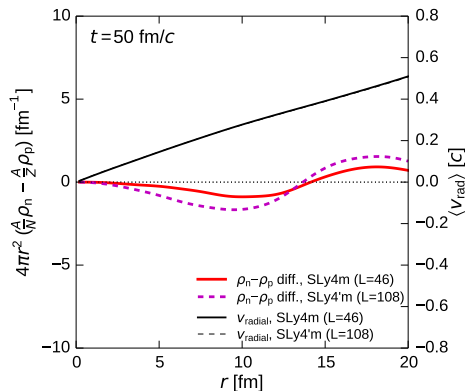


Momentum dependence of Skyrme (SLy4) interaction has been corrected for high energy collisions, in a similar way to [Gale, Bertsch, Das Gupta, PRC 35 \(1987\) 1666](#).

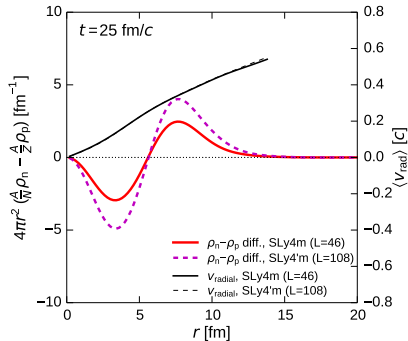
- Neutron-proton density diff. (fn of  $r$ )

$$4\pi r^2 \left[ \frac{A}{N} \rho_n(r) - \frac{A}{Z} \rho_p(r) \right]$$

- Radial expansion velocity  $v_{\text{rad}}(r)$

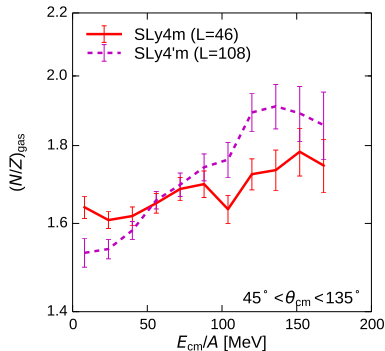


$\rho_n - \rho_p$  at the compression stage



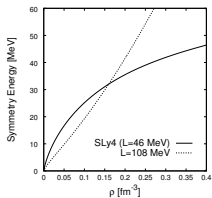
↔  
similar

N/Z of the spectrum of emitted particles



$$\left(\frac{N}{Z}\right)_{\text{gas}} = \frac{Y_n(\epsilon) + Y_d(\epsilon) + 2Y_t(\epsilon) + Y_h(\epsilon) + 2Y_\alpha(\epsilon)}{Y_p(\epsilon) + Y_d(\epsilon) + Y_t(\epsilon) + 2Y_h(\epsilon) + 2Y_\alpha(\epsilon)}$$

$$\epsilon = E_{\text{cm}}/A$$



M. Kaneko, Murakami, Isobe, Kurata-Nishimura, Ono, Ikeno et al. ( $S\pi$ RIT), PLB 822 (2021) 136681.

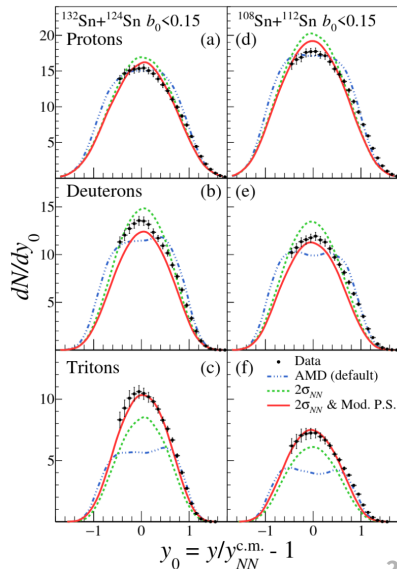
## Rapidity distributions

for  $p$ ,  $d$  and  $t$ , in the central collisions of

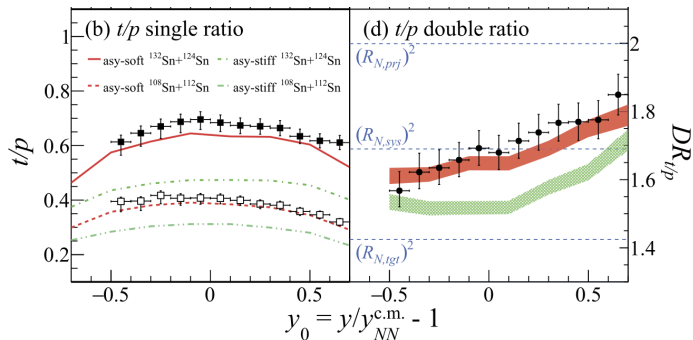
- $^{132}\text{Sn} + ^{124}\text{Sn}$  at 270 MeV/nucleon
- $^{108}\text{Sn} + ^{112}\text{Sn}$  at 270 MeV/nucleon

- Black points:  $S\pi$ RIT data
- Lines: AMD calculations (the asy-soft symmetry energy  $L = 46$  MeV) with different model parameters.
  - The results in the next page don't depend much of the model parameters.

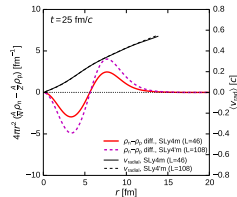
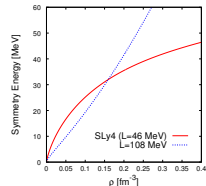
neutron rich vs. neutron deficient



# $t/p$ ratio and its implication on the symmetry energy



$$\frac{t/p \text{ in } ^{132}\text{Sn} + ^{124}\text{Sn}}{t/p \text{ in } ^{108}\text{Sn} + ^{112}\text{Sn}} = (t/p \text{ double ratio})$$



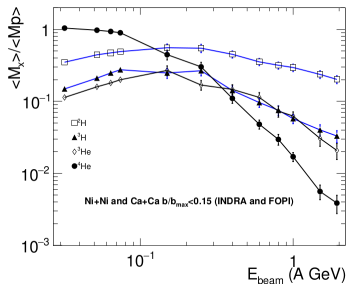
- The  $\pi$ RIT data (black points) favor **the asy-soft symmetry energy (L=46 MeV)** rather than **the asy-stiff symmetry energy (L=108 MeV)**.
- The moderate rapidity dependence of the  $t/p$  double ratio implies a partial isospin mixing.
- $t/{}^3\text{He}$  should be a better probe.
- Stopping needs to be understood better systematically.
- Comparisons of other observables (e.g. the collective flow) are in progress.

# Cluster yields at various incident energies

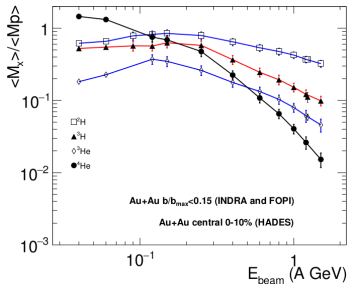
Bougault et al., *Symmetry* 2021, 13, 1406.

Data sets from INDRA and FOPI

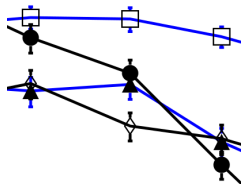
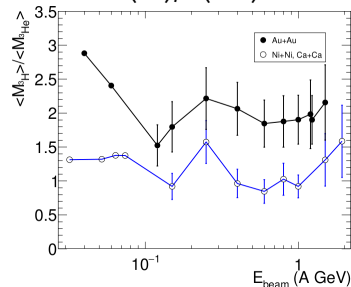
$M(x)/M(p)$  in Ni + Ni, Ca + Ca



$M(x)/M(p)$  in Au + Au

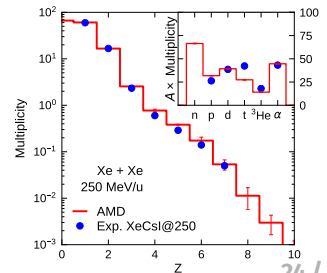


$M(^3H)/M(^3He)$



Is there any special mechanism for the triton production around 250A MeV (beyond the current transport model descriptions)?

Does it depend on  $N/Z$  of the system?





- Light clusters and heavier fragments cannot be ignored in understanding the global dynamics of heavy-ion collisions in the full range of energies.
- Many symmetry energy observables are related to clusters.
- In Fermi energy domain, clusters can be regarded as constituents of the gas part which has larger  $N/Z$  ratio than the liquid part in neutron-rich systems.
  - $S(\rho < \rho_0)$  from the  $N/Z$  ratio of the gas part (i.e., CI spectrum ratio).
  - Some hint on  $\nabla S(\rho < \rho_0)$  from isospin drift ( $L = 108$  MeV rather than  $L = 46$  MeV), from a comparison of FAZIA data and AMD calculation.
  - ...
- At higher energies (e.g.,  $E/A = 270$  MeV), clusters are mainly created in the inner part of the expanding system, and therefore they carry residual information on the  $N/Z$  of the compressed matter.
  - The  $t/p$  ratio of the rapidity distributions of the  $S\pi$ RIT data for Sn + Sn at 270A MeV favors  $L = 46$  MeV rather than  $L = 108$  MeV, in a comparison with an AMD calculation.
  - ...
- It is a highly nontrivial problem to handle cluster correlations in transport models, though there are some models which are practically working (e.g., AMD with an extension for cluster correlations).
  - How can we identify clusters in general situations?
  - How should clusters be suppressed in medium?
  - ...
- Anything special beyond the current AMD, to produce extra tritons and neutron-rich fragments?

## More about cluster production in NN collisions



$$vd\sigma \propto |\langle \varphi'_1 | \varphi_1^{+q} \rangle|^2 |\langle \varphi'_2 | \varphi_2^{-q} \rangle|^2 |M|^2 \delta(E_f - E_i) p_{\text{rel}}^2 dp_{\text{rel}} d\Omega \quad E_i, E_f = \frac{\langle \Phi(Z) | H | \Phi(Z) \rangle}{\langle \Phi(Z) | \Phi(Z) \rangle}$$

$$\Rightarrow P(C_1, C_2, p_{\text{rel}}, \Omega) \times \left| M(p_{\text{rel}}^{(0)}, p_{\text{rel}}, \Omega) \right|^2 \times \frac{p_{\text{rel}}^2 d\Omega}{\partial E_f / \partial p_{\text{rel}}}$$

- Gaussian width  $v_{\text{cl}} = 0.24 \text{ fm}^{-2}$  for the overlap factors.
- There are a huge number of final cluster configurations  $(C_1, C_2)$ .

$$\sum_{C_1 C_2} P(C_1, C_2, p_{\text{rel}}, \Omega) = 1 \quad \text{for any fixed } (p_{\text{rel}}, \Omega)$$

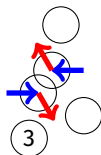
- The energy-conserving final momentum depends on the cluster configuration

$$p_{\text{rel}} = p_{\text{rel}}(C_1, C_2, \Omega)$$

When clusters are formed,  $p_{\text{rel}}$  tends to be large, and the effect of collisions will increase.

- the phase space factor  $\uparrow$
- Pauli blocking  $\downarrow$  (collision probability  $\uparrow$ )
- momentum transfer  $\uparrow$

$$W_{i \rightarrow f} = \frac{2\pi}{\hbar} |\langle \Psi_f | V | \Psi_i \rangle|^2 \delta(E_f - E_i)$$

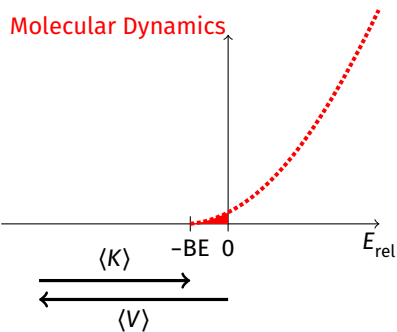


In the usual way of NN collision, only the two wave packets are changed.

$$\{ |\Psi_f\rangle \} = \{ | \varphi_{k_1}(1) \varphi_{k_2}(2) \Psi(3, 4, \dots) \rangle \}$$

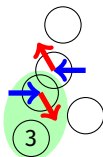
(ignoring antisymmetrization for simplicity of presentation.)

Phase space or the density of states for two nucleon system



# NN collisions without or with cluster correlations

$$W_{i \rightarrow f} = \frac{2\pi}{\hbar} |\langle \Psi_f | V | \Psi_i \rangle|^2 \delta(E_f - E_i)$$



In the usual way of NN collision, only the two wave packets are changed.

$$\{ |\Psi_f\rangle \} = \{ |\varphi_{k_1}(1)\varphi_{k_2}(2)\psi(3, 4, \dots)\rangle \}$$

(ignoring antisymmetrization for simplicity of presentation.)

## Extension for cluster correlations

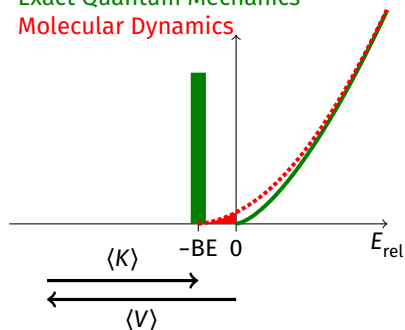
Include correlated states in the set of the final states of each NN collision.

$$\{ |\Psi_f\rangle \} \ni |\varphi_{k_1}(1)\psi_d(2, 3)\psi(4, \dots)\rangle, \dots$$

Phase space or the density of states for two nucleon system

Exact Quantum Mechanics

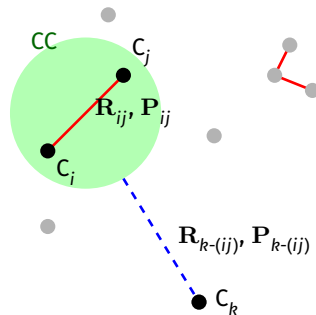
Molecular Dynamics



Several clusters may form a loosely bound state.

e.g.,  ${}^7\text{Li} = \alpha + t - 2.5 \text{ MeV}$

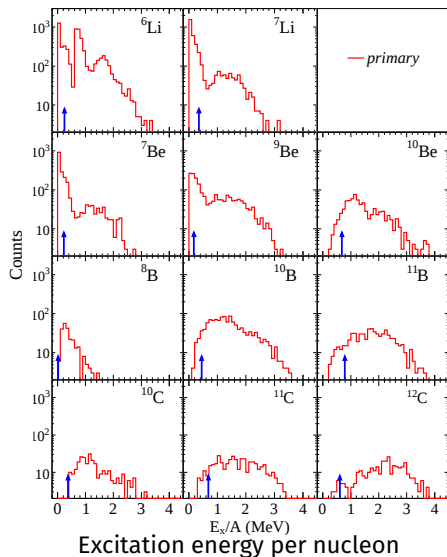
Need more probability of  $|\alpha + t\rangle \rightarrow |{}^7\text{Li}\rangle$  etc.



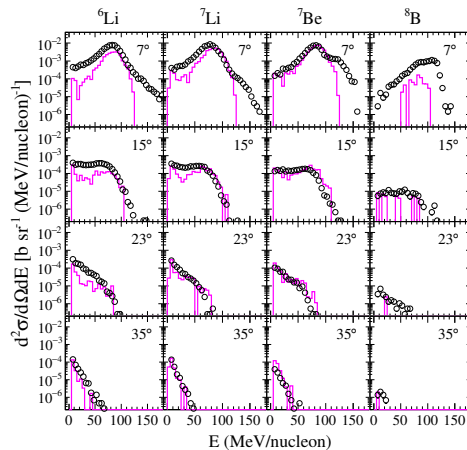
# Production of light nuclei

$^{12}\text{C} + ^{12}\text{C}$  at 95 MeV/nucleon

Tian et al., PRC 97 (2018) 034610.

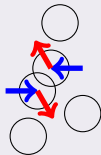


Some light nuclei are emitted at large angles ( $\theta_{\text{lab}} > 20^\circ$ ) almost in their ground states, at  $t = 300$  fm/c.



## Without cluster correlation

$$W_{i \rightarrow f} = \frac{2\pi}{\hbar} |\langle \psi_f | V | \psi_i \rangle|^2 \delta(E_f - E_i)$$



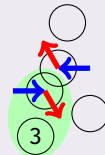
A collision of particles 1 and 2 will change only the two wave packets.

$$\{ |\psi_f\rangle \} = \{ | \varphi_{k_1}(1) \varphi_{k_2}(2) \psi(3, 4, \dots) \rangle \}$$

(ignoring antisymmetrization for simplicity of presentation.)

## With cluster correlation

$$W_{i \rightarrow f} = \frac{2\pi}{\hbar} |\langle \psi_f | V | \psi_i \rangle|^2 \delta(E_f - E_i)$$

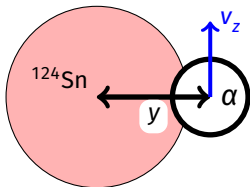


Include correlated states in the set of the final states of each NN collision.

$$\{ |\psi_f\rangle \} \ni | \varphi_{k_1}(1) \psi_d(2, 3) \psi(4, \dots) \rangle, \dots$$

(ignoring antisymmetrization for simplicity of presentation.)

# A cluster put into a nucleus in AMD



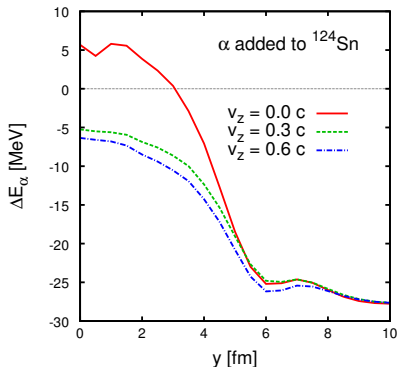
$\alpha$  cluster  $|\alpha, \mathbf{Z}\rangle$ : Four wave packets with different spins and isospins at the same phase space point  $\mathbf{Z}$ .

$$E_\alpha : A |\alpha, \mathbf{Z}\rangle |^{124}\text{Sn}\rangle$$

$$E_N : A |\mathbf{Z}\rangle |^{124}\text{Sn}\rangle \quad (N = p \uparrow, p \downarrow, n \uparrow, n \downarrow)$$

$$-B_\alpha = \Delta E_\alpha = E_\alpha - (E_{p\uparrow} + E_{p\downarrow} + E_{n\uparrow} + E_{n\downarrow})$$

(Energies are defined relative to  $|^{124}\text{Sn}\rangle$ .)



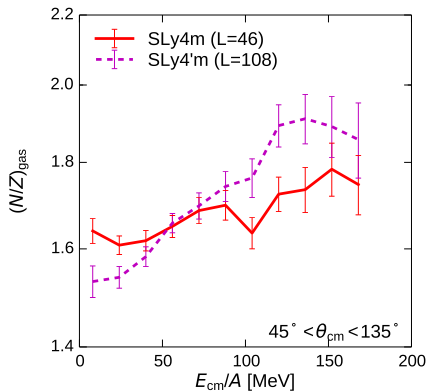
$$\frac{\text{Re } \mathbf{Z}}{\sqrt{v}} = (0, y, 0), \quad \frac{2\hbar\sqrt{v} \text{Im } \mathbf{Z}}{M} = (0, 0, v_z)$$

- Distance from the center:  $y$   
 $\approx$  Dependence on density
- Dependence on  $P_\alpha = M_\alpha v_z$
- Due to the density dependence of the Skyrme force, the interaction between nucleons in the  $\alpha$  cluster is weakened in the nucleus.

Energy is OK, but dynamics is ...

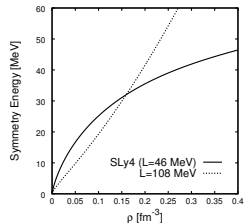
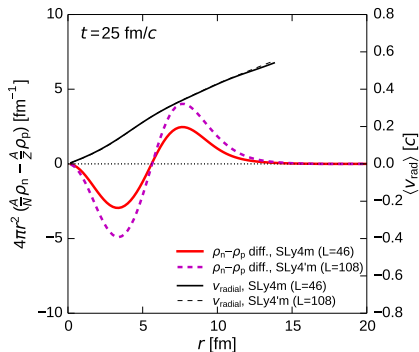


# N/Z Spectrum Ratio (AMD with clusters)

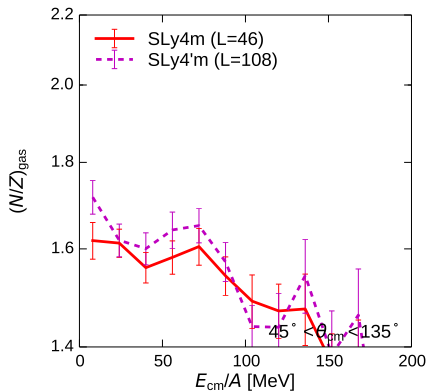


$$\left(\frac{N}{Z}\right)_{\text{gas}} = \frac{Y_n(v) + Y_d(v) + 2Y_t(v) + Y_h(v) + 2Y_\alpha(v)}{Y_p(v) + Y_d(v) + Y_t(v) + 2Y_h(v) + 2Y_\alpha(v)}$$

N/Z of spectrum of emitted particles is similar to the neutron-proton density difference at the compression stage.

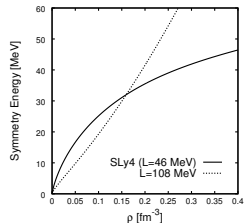
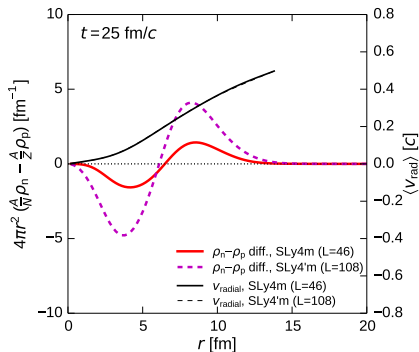


# N/Z Spectrum Ratio (AMD withOUT clusters)



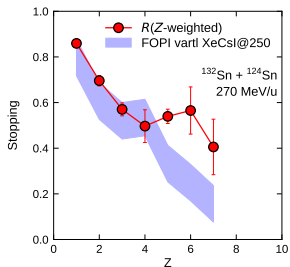
$$\left(\frac{N}{Z}\right)_{\text{gas}} = \frac{Y_n(v) + Y_d(v) + 2Y_t(v) + Y_h(v) + 2Y_\alpha(v)}{Y_p(v) + Y_d(v) + Y_t(v) + 2Y_h(v) + 2Y_\alpha(v)}$$

N/Z of spectrum of emitted particles is NOT similar to the neutron-proton density difference at the compression stage.

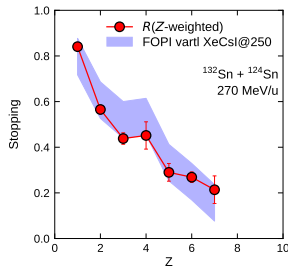


# Results from different choices of cluster and $\sigma_{NN}$ in medium

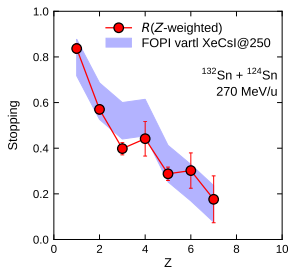
Cluster(full)  
 $\sigma_{NN}$ (free)



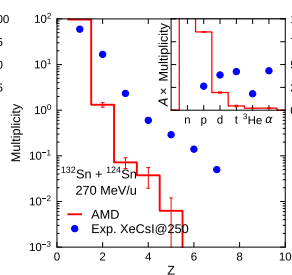
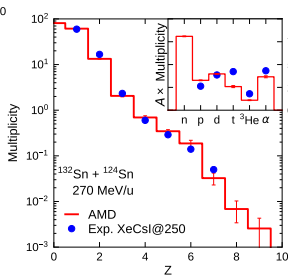
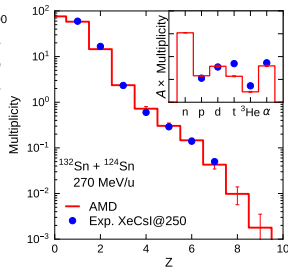
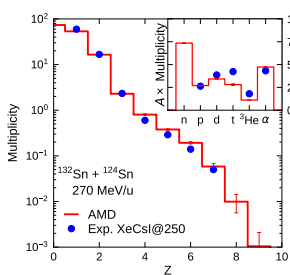
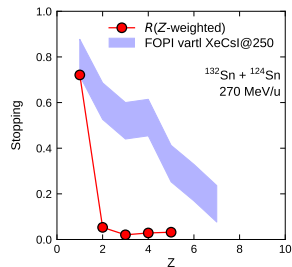
Cluster(0.125)  
 $\sigma_{NN}$ (in-medium)



Cluster(0.060)  
 $\sigma_{NN}$ (free)

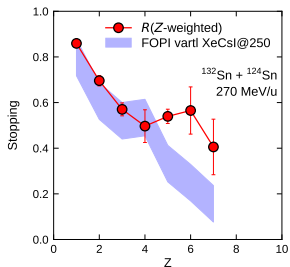


Cluster(0)  
 $\sigma_{NN}$ (in-medium)

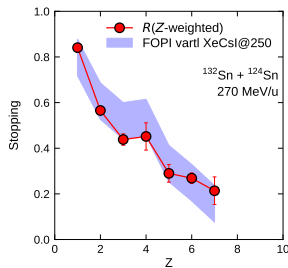


# Results from different choices of cluster and $\sigma_{NN}$ in medium

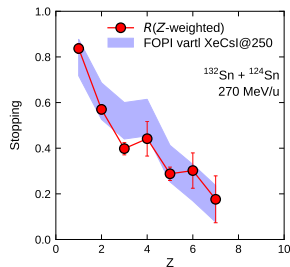
Cluster(full)  
 $\sigma_{NN}$ (free)



Cluster(0.125)  
 $\sigma_{NN}$ (in-medium)



Cluster(0.060)  
 $\sigma_{NN}$ (free)



Cluster(0)  
 $\sigma_{NN}$ (in-medium)

