

SarWorS 2021
7 September 2021

TMDs & EIC impact studies

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MAP Collaboration

(Multi-dimensional Analyses of Partonic distributions)



Unpolarized quark TMDs

TMD PDF

$$\text{scales } \mu = Q$$
$$\zeta = Q^2$$

$$\mu_b = \frac{2e^{-\gamma_E}}{b_*}$$

$$f_1^q(x, b; \mu, \zeta) = \sum_j \left(C_{q/j} \otimes f_1^j \right) (x, b_*; \mu_b) e^{S(b_*; \mu_b, \mu)} e^{g_K(b) \ln \frac{\mu}{\mu_0}} f_{\text{NP}}^q(x, b, \zeta)$$

Unpolarized quark TMDs

TMD PDF

scales $\mu_b = \frac{2e^{-\gamma_E}}{b_*}$
 $\mu = Q$
 $\zeta = Q^2$

collinear PDFs

Sudakov factor
 perturbative evolution

$$f_1^q(x, b; \mu, \zeta) = \sum_j \left(C_{q/j} \otimes f_1^j \right) (x, b_*; \mu_b) e^{S(b_*; \mu_b, \mu)} e^{g_K(b) \ln \frac{\mu}{\mu_0}} f_{\text{NP}}^q(x, b, \zeta)$$

matching coefficients

non perturbative
 evolution

intrinsic non perturbative
 transverse content

Unpolarized quark TMDs

TMD PDF

collinear PDFs

Sudakov factor
perturbative evolution

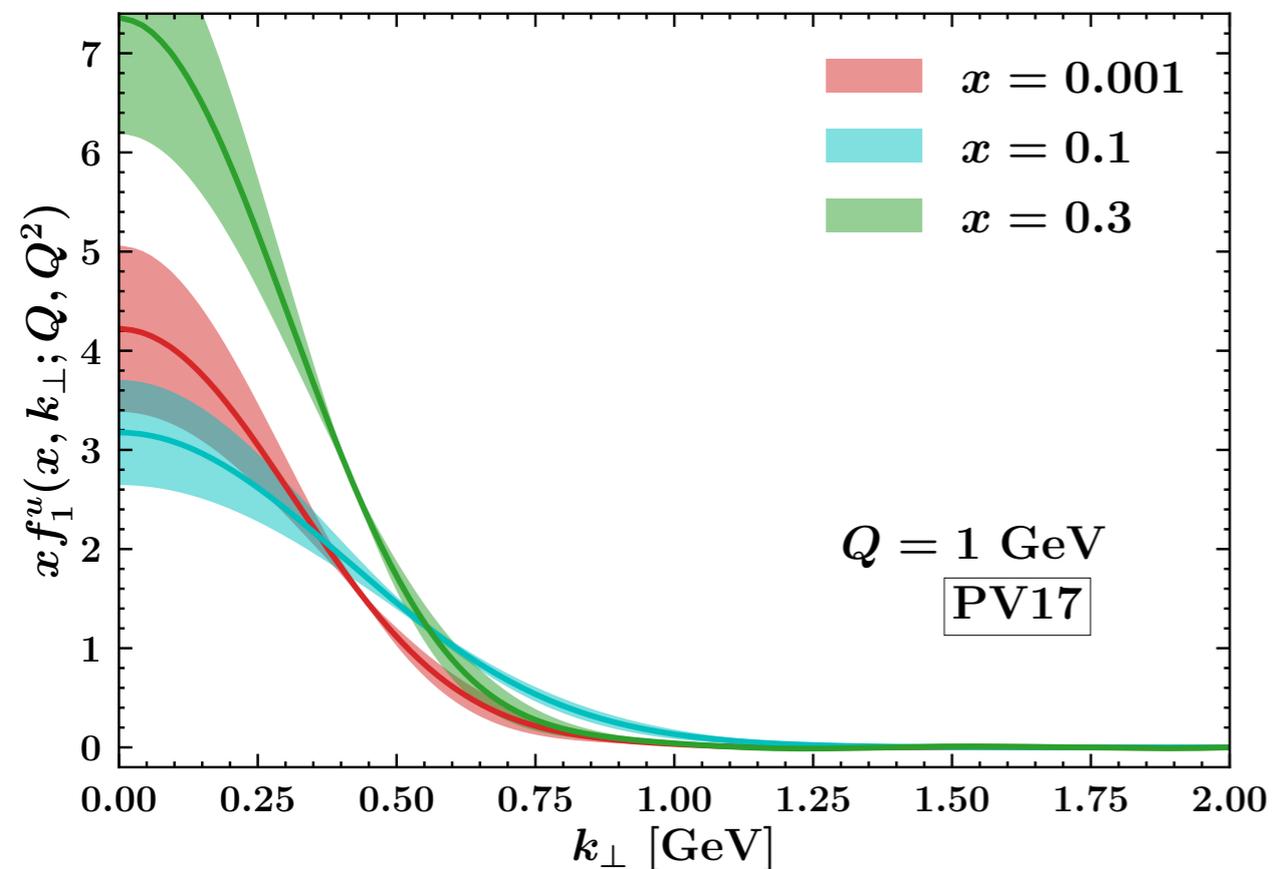
$$f_1^q(x, b; \mu, \zeta) = \sum_j \left(C_{q/j} \otimes f_1^j \right) (x, b_*; \mu_b) e^{S(b_*; \mu_b, \mu)} e^{g_K(b) \ln \frac{\mu}{\mu_0}} f_{\text{NP}}^q(x, b, \zeta)$$

matching coefficients

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transverse content

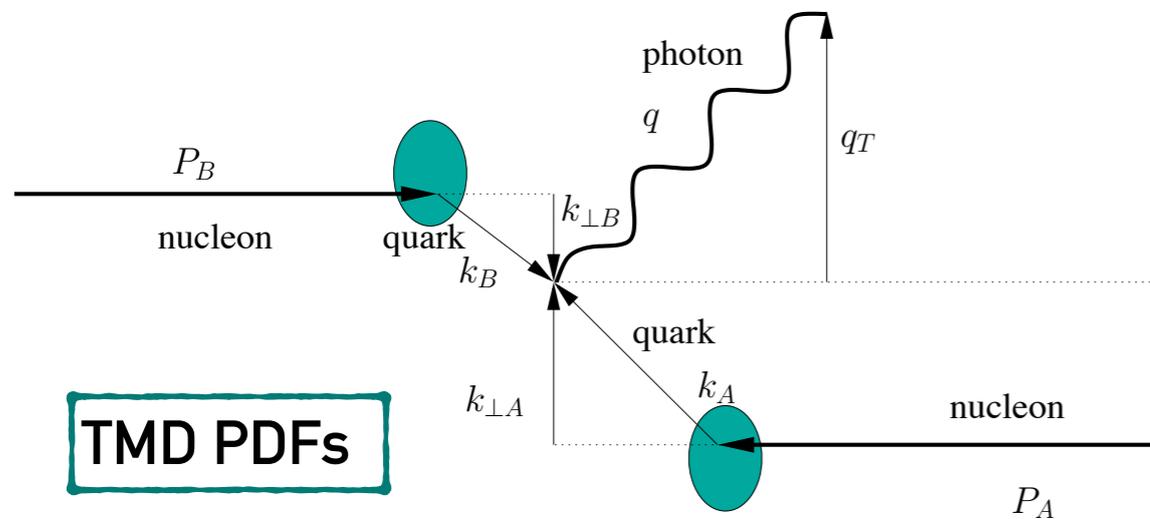
non perturbative
parametrized and fitted to data



A. Bacchetta, F. Delcarro, C. Pisano, M. Radici, A. Signori
JHEP06 (2017) 081, arXiv:1703.10157

Drell-Yan and SIDIS

$$N(P_A) + N(P_B) \rightarrow \gamma^*/Z \rightarrow l^+l^-$$



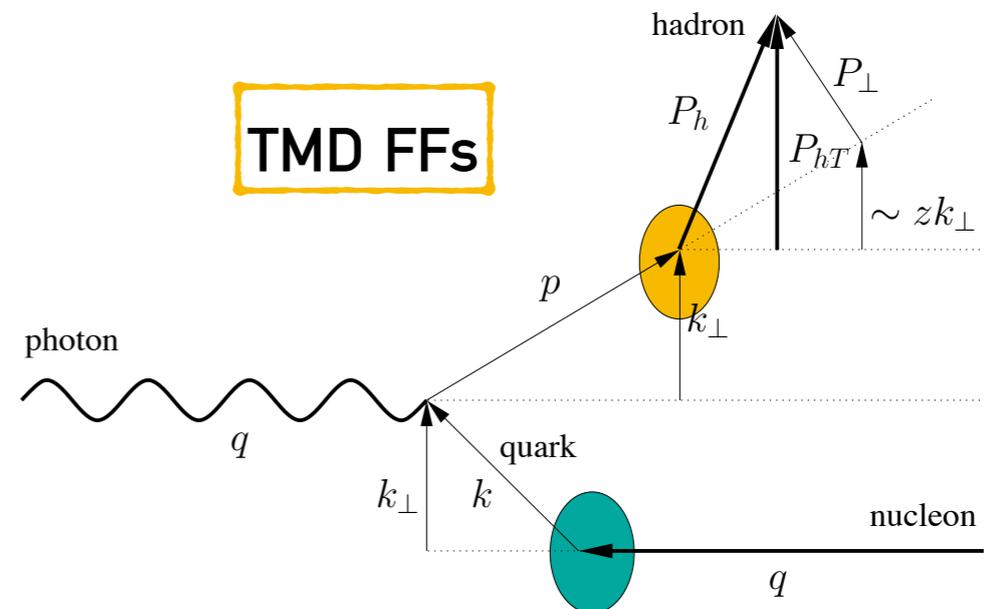
$$q_T \ll Q$$

TMD factorization

$$\left(\frac{d\sigma}{dq_T} \right) \propto$$

$$\int \frac{d^2\mathbf{b}}{4\pi} e^{i\mathbf{b}\cdot\mathbf{q}_T} x_1 \boxed{f_1^q(x_1, \mathbf{b})} x_2 \boxed{f_1^{\bar{q}}(x_2, \mathbf{b})}$$

$$\ell(l) + N(p) \rightarrow \ell(l') + h(P_h) + X$$



$$M^2 \ll Q^2 \quad P_{hT}^2 \ll Q^2$$

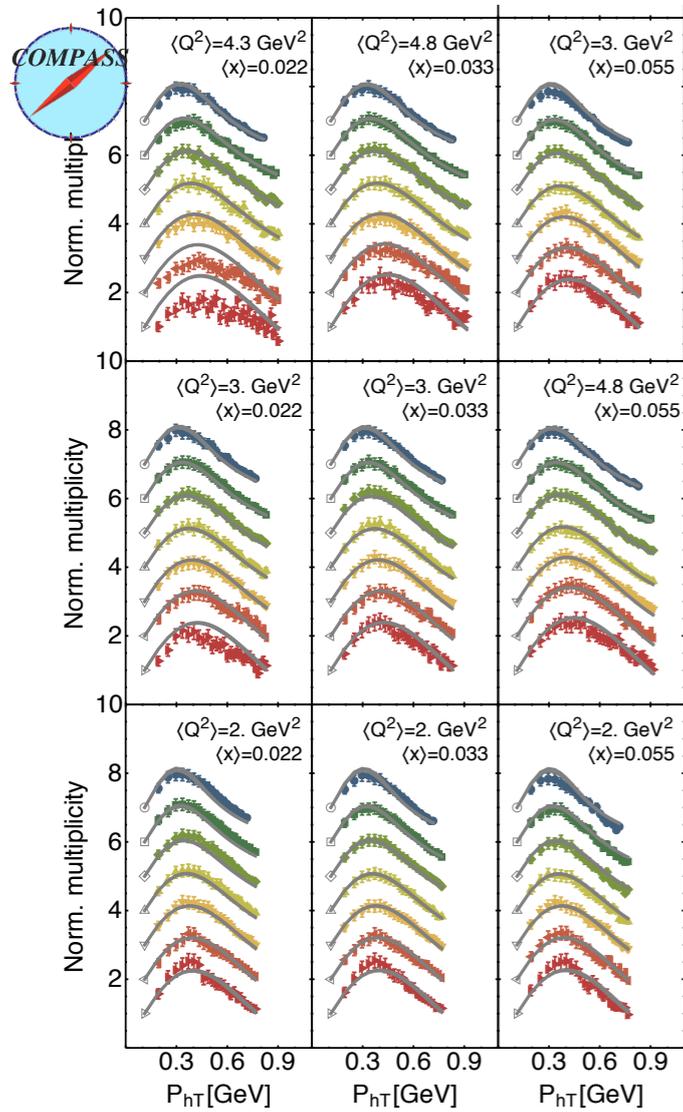
$$\left(\frac{d\sigma}{dq_T} \right) \propto$$

$$\int \frac{d^2\mathbf{b}}{4\pi} e^{i\mathbf{b}\cdot\mathbf{q}_T} \boxed{f_1^q(x, \mathbf{b})} \boxed{D_1^{q \rightarrow h}(z, \mathbf{b})}$$

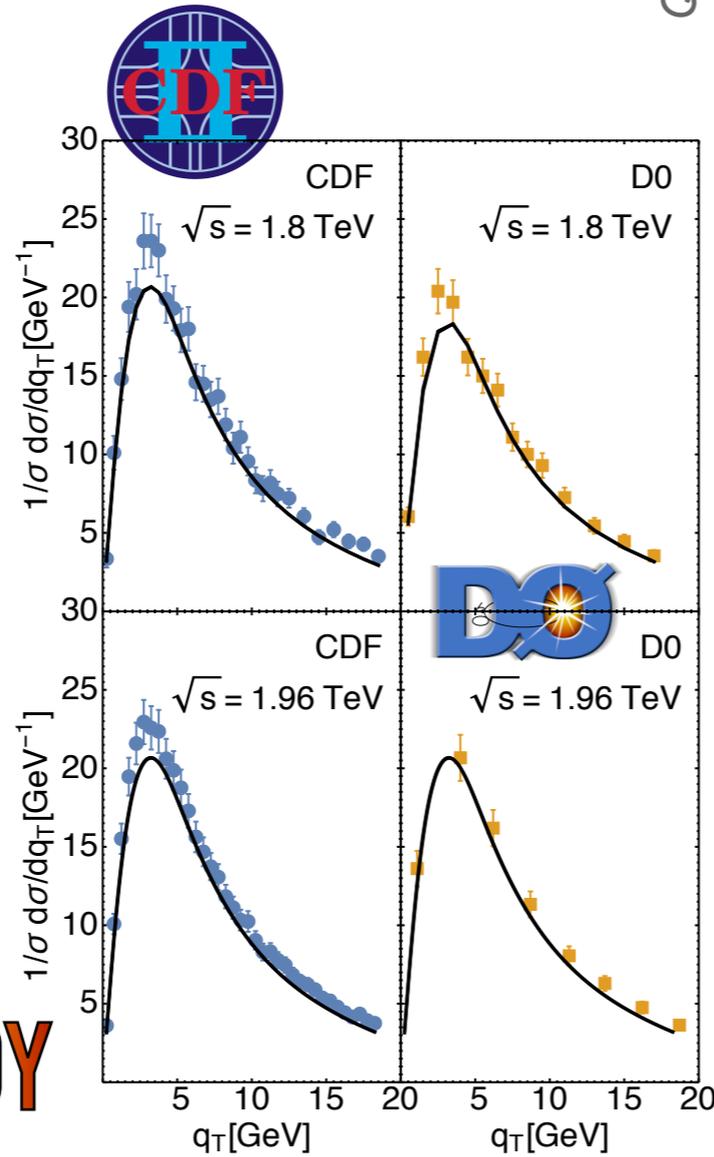
PV17 fit

A. Bacchetta, F. Delcarro, C. Pisano, M. Radici, A. Signori
 JHEP06 (2017) 081, arXiv:1703.10157

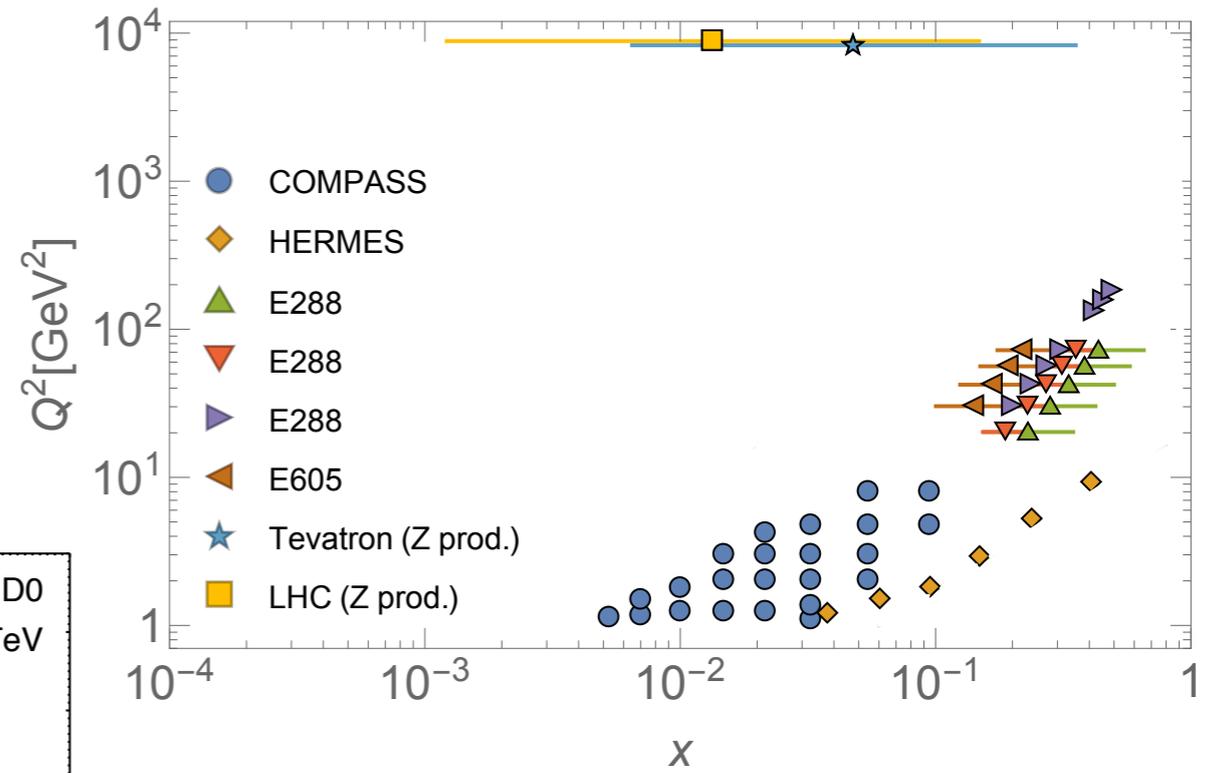
SIDIS



8059
 data points



DY



**with
 normalization
 coefficients**

NLL

global

$\chi^2 = 1.55$

PV17 non perturbative functions

A. Bacchetta, F. Delcarro, C. Pisano, M. Radici, A. Signori

JHEP06 (2017) 081, arXiv:1703.10157

$$f_{1\text{NP}}^a(x, \mathbf{k}_\perp^2) = \frac{1}{\pi} \frac{(1 + \lambda \mathbf{k}_\perp^2)}{g_{1a} + \lambda g_{1a}^2} e^{-\frac{\mathbf{k}_\perp^2}{g_{1a}}}$$

$$D_{1\text{NP}}^{a \rightarrow h}(z, \mathbf{P}_\perp^2) = \frac{1}{\pi} \frac{1}{g_{3a \rightarrow h} + (\lambda_F/z^2) g_{4a \rightarrow h}^2} \left(e^{-\frac{\mathbf{P}_\perp^2}{g_{3a \rightarrow h}}} + \lambda_F \frac{\mathbf{P}_\perp^2}{z^2} e^{-\frac{\mathbf{P}_\perp^2}{g_{4a \rightarrow h}}} \right)$$

x-dependence

$$g_1(x) = N_1 \frac{(1-x)^\alpha x^\sigma}{(1-\hat{x})^\alpha \hat{x}^\sigma}$$

$$g_{3,4}(z) = N_{3,4} \frac{(z^\beta + \delta) (1-z)^\gamma}{(\hat{z}^\beta + \delta) (1-\hat{z})^\gamma}$$

non-perturbative Sudakov factor

$$g_K(b_T) = -g_2 b_T^2 / 2$$

total of **11 parameters**

PV17 fit

A. Bacchetta, F. Delcarro, C. Pisano, M. Radici, A. Signori
JHEP06 (2017) 081, arXiv:1703.10157

8059 points
11 parameters
SIDIS + DY
NLL

$$\chi_{\text{d.o.f.}}^2 = 1.55 \pm 0.05$$

all standard deviations at 68% c.l.

g_2	N_1 [GeV ²]	β	α	λ_F [GeV ⁻²]
0.13 ± 0.01	0.28 ± 0.06	1.65 ± 0.49	2.95 ± 0.05	5.50 ± 1.23

$$\langle x \rangle \pm \Delta x$$

fairly well determined

not very constrained

$g_2 \rightarrow$ non-perturbative evolution

$\beta \rightarrow$ low-z width of TMD FF

$N_1 \rightarrow$ mid-x width of TMD PDF

$\alpha \rightarrow$ high-x width of TMD PDF

$\lambda_F \rightarrow$ weight of second Gaussian in TMD FF

PV17 fit

A. Bacchetta, F. Delcarro, C. Pisano, M. Radici, A. Signori
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8059 points
11 parameters
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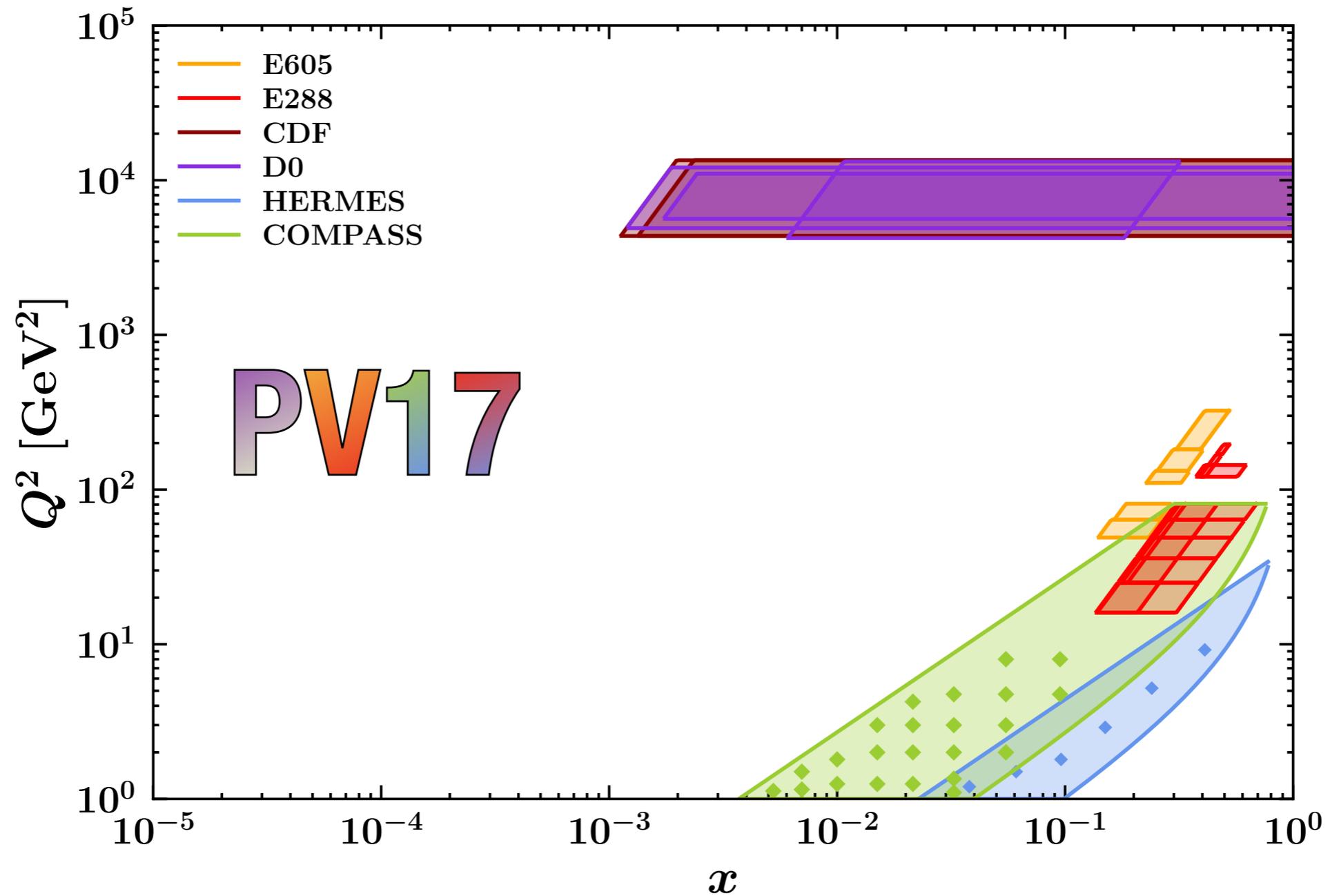
$$\langle x \rangle \pm \Delta x$$

fairly well determined

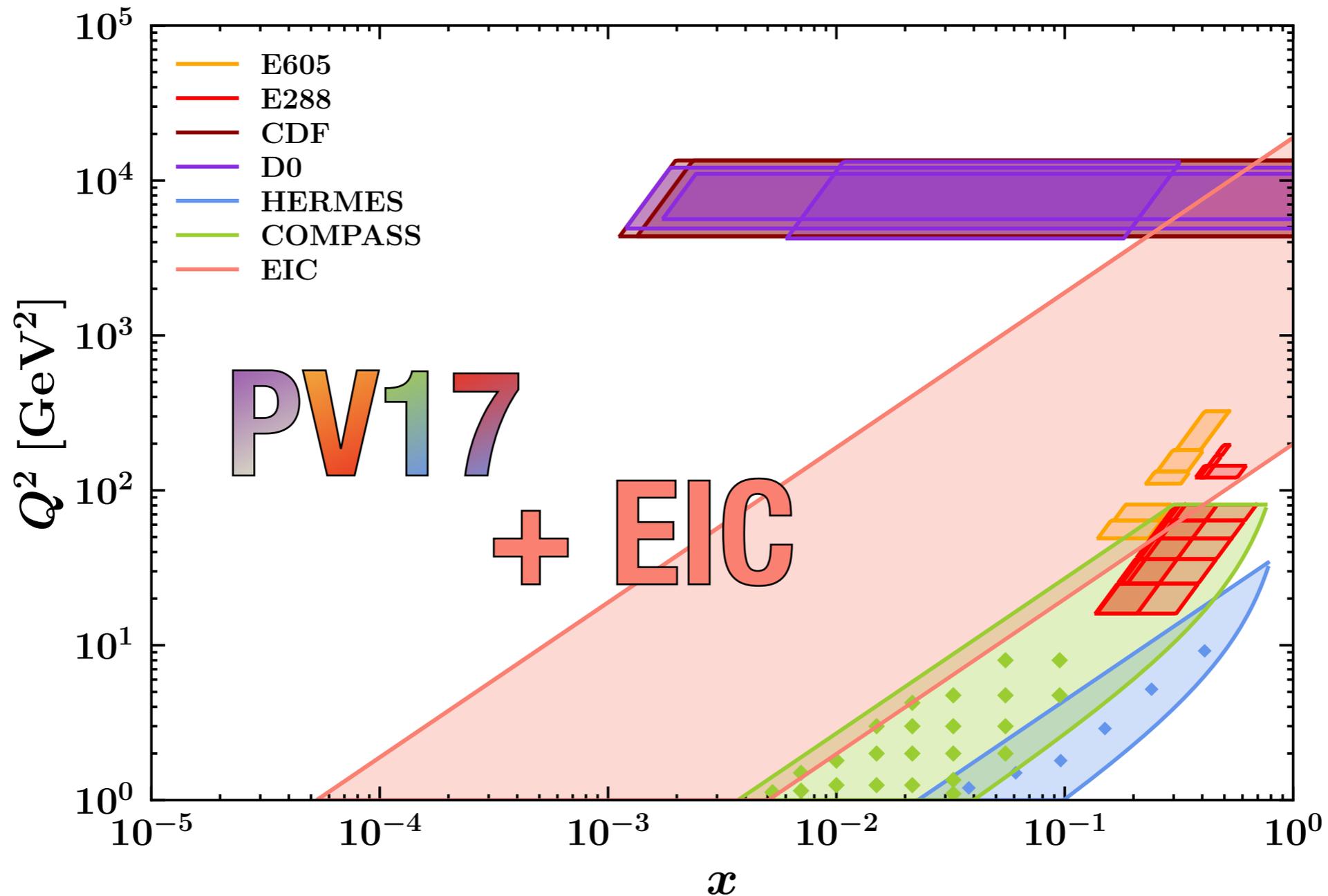
not very constrained

what impact will the EIC have on those uncertainties?

What happens to PV17 TMDs ...

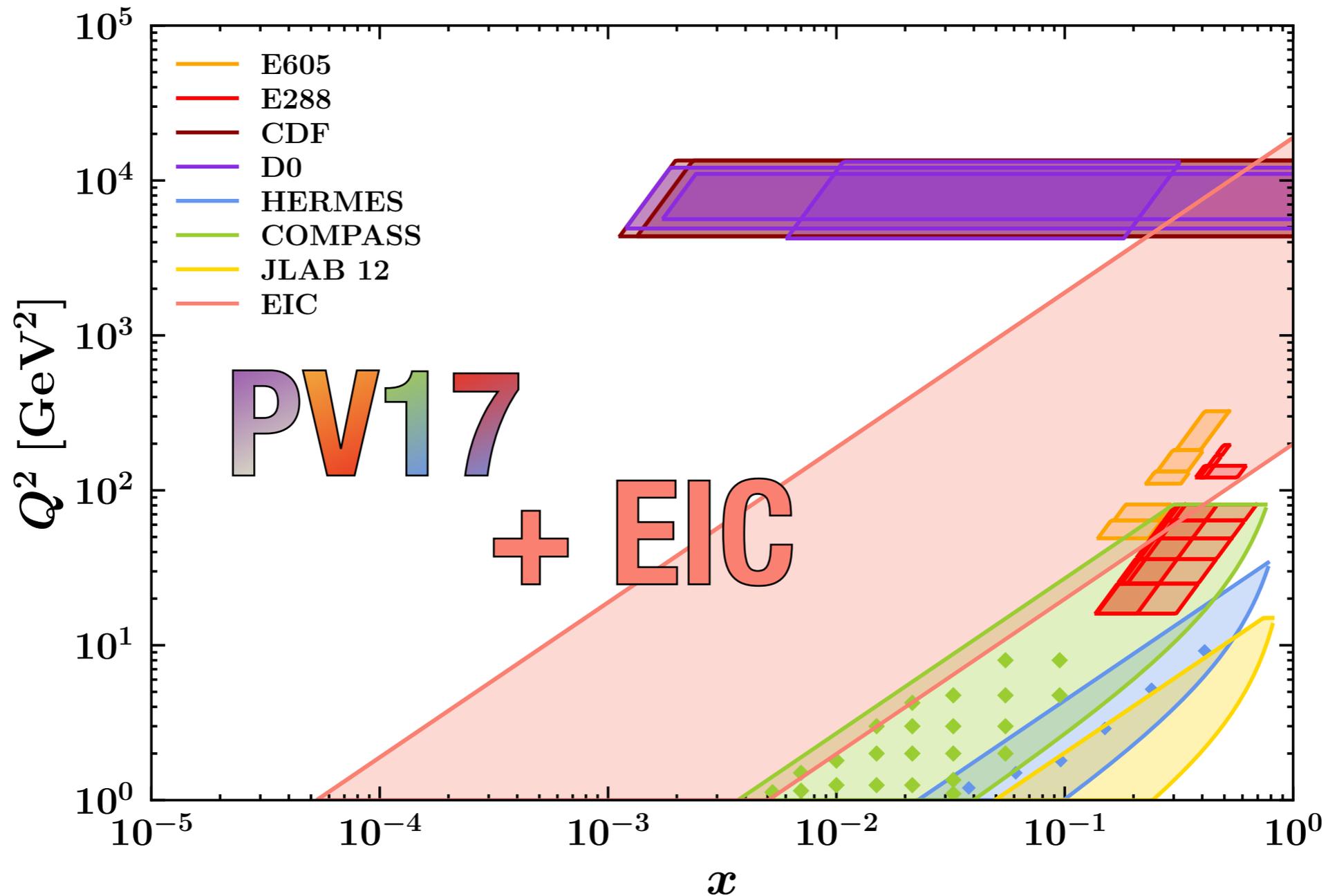


What happens to PV17 TMDs ...



... once we include EIC data?

Electron-Ion Collider



little side-note:

impact studies can be also done for **JLAB 12**

EIC pseudodata

generated by Ralf Seidl and available on https://github.com/VladimirovAlexey/EIC_YR_TMD

π^+
 π^- final state hadrons
 K^+
 K^-

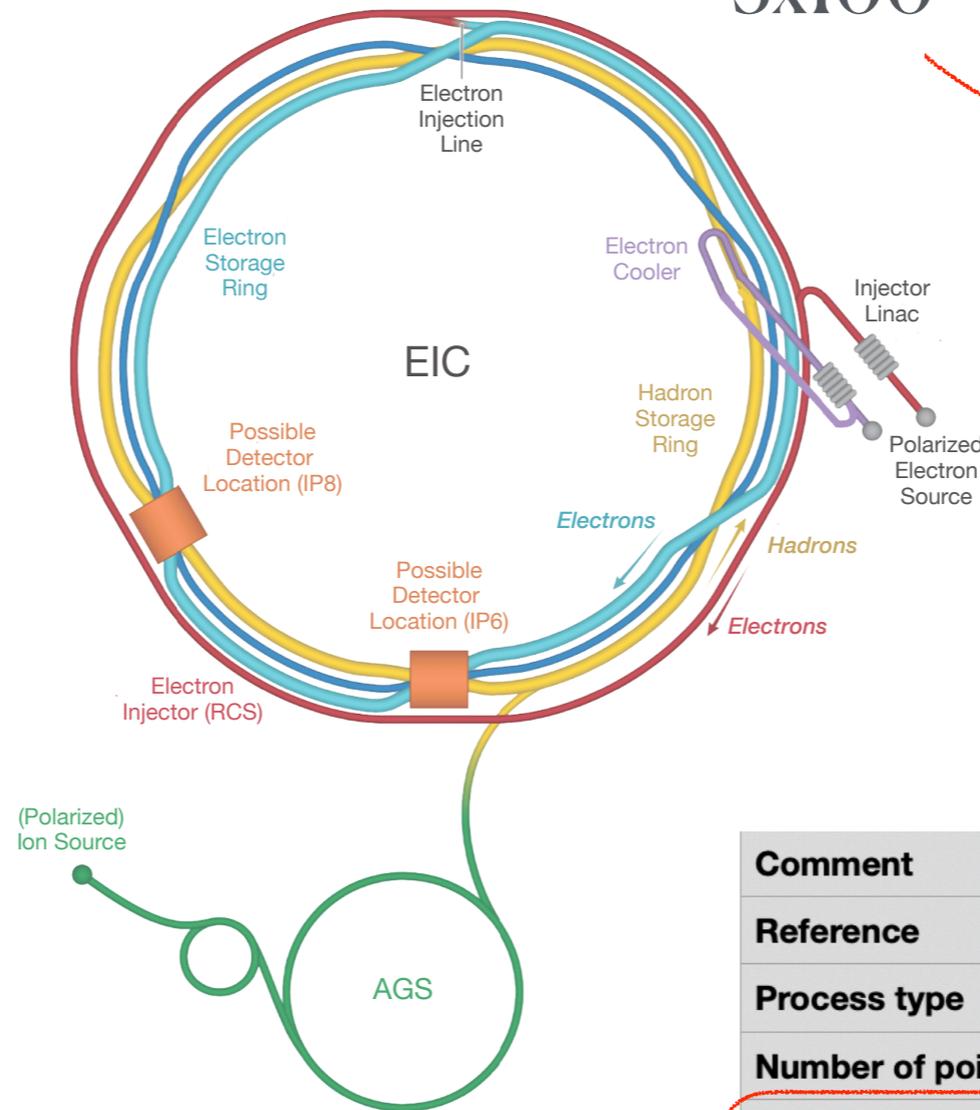


image from EIC YR
arXiv:2103.05419

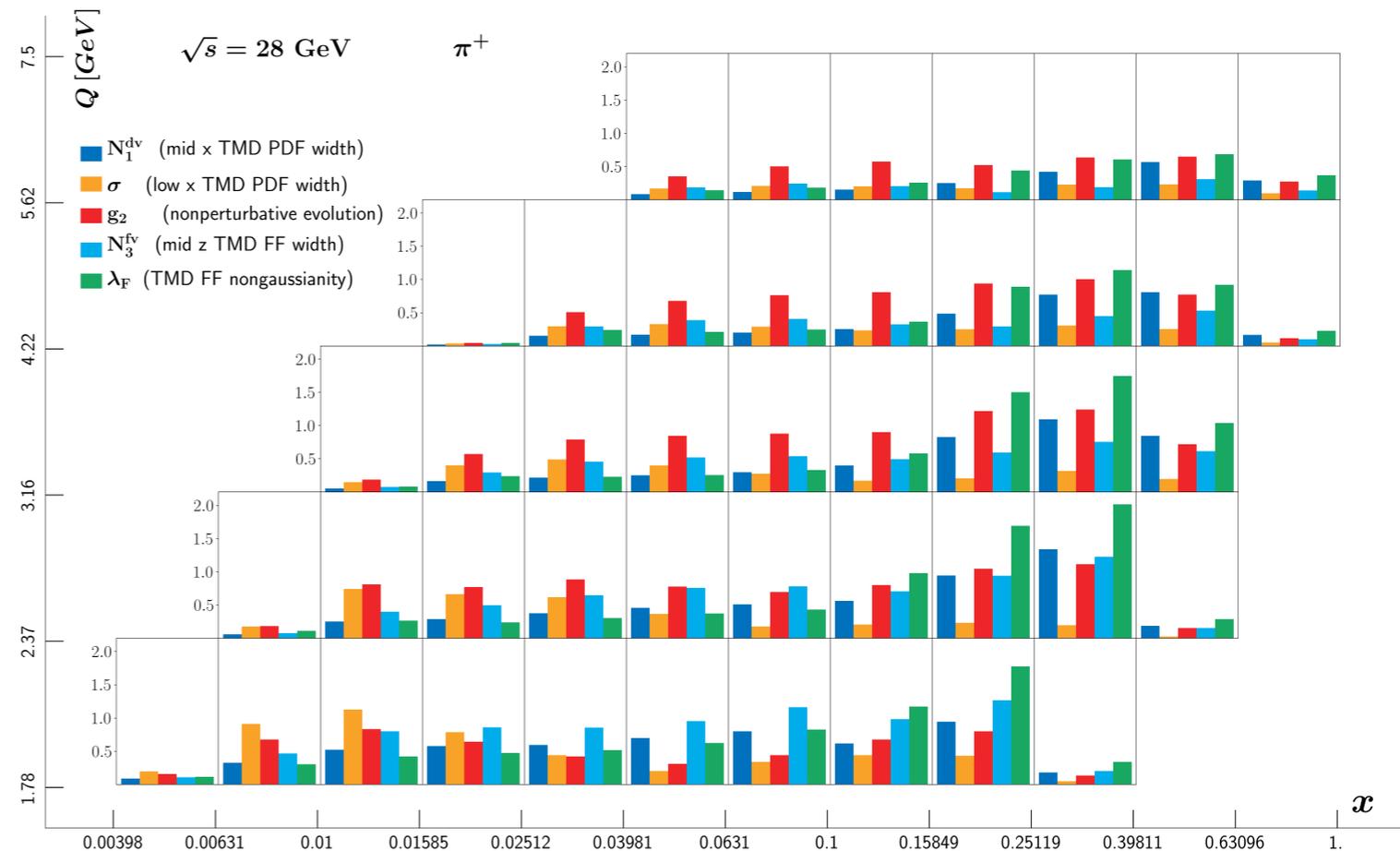
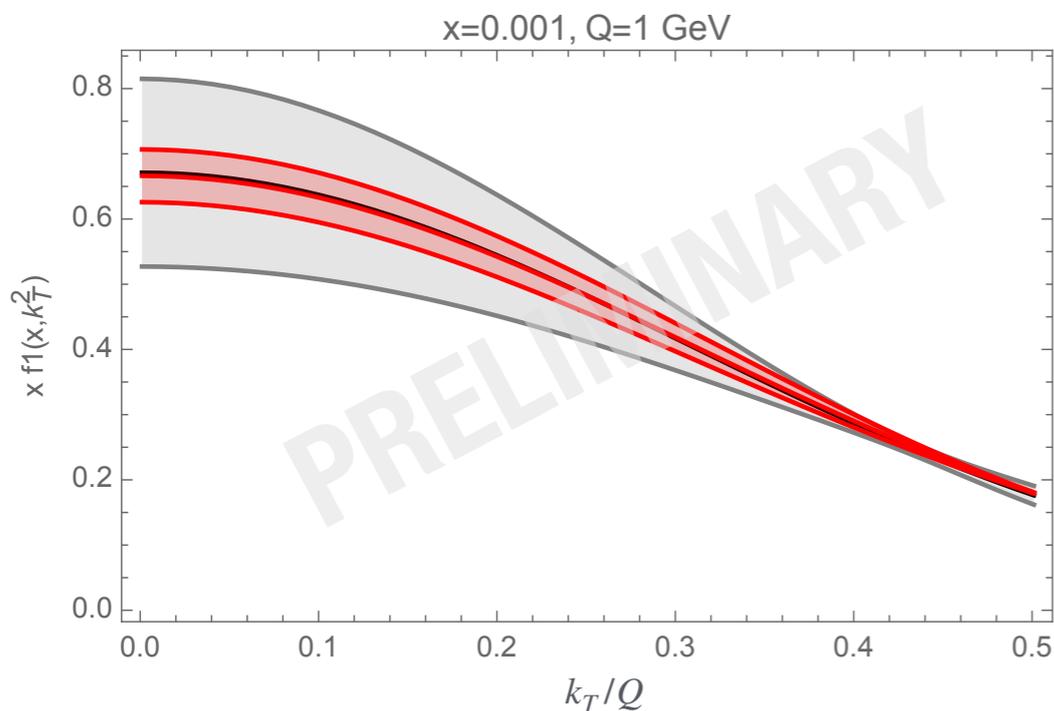
5x41 ————— 18x275
 configurations
 5x100 ————— 18x100
 10x100

eight options for EIC settings:
we choose **option 8**

Comment	Ralf's pseudo data for EIC.
Reference	Ralf
Process type	SIDIS
Number of points	2958
Number of uncorr.errors	2
Number of corr.errors	uncertainties 0
Number of norm.errors	1
List of norm.errors (relative)	0.03
Total cross-section normalized	False

A few tools to estimate EIC impact

reweighting



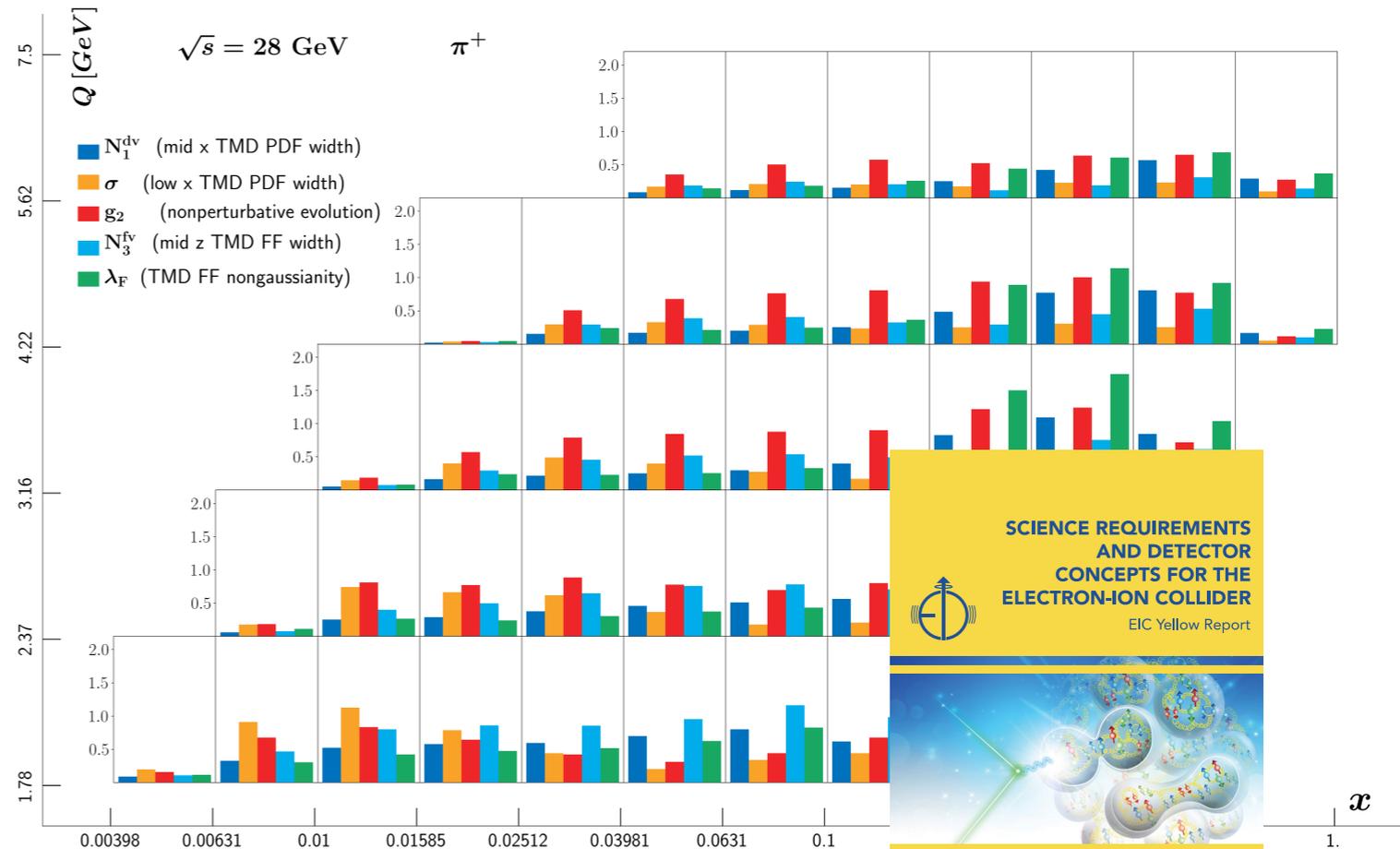
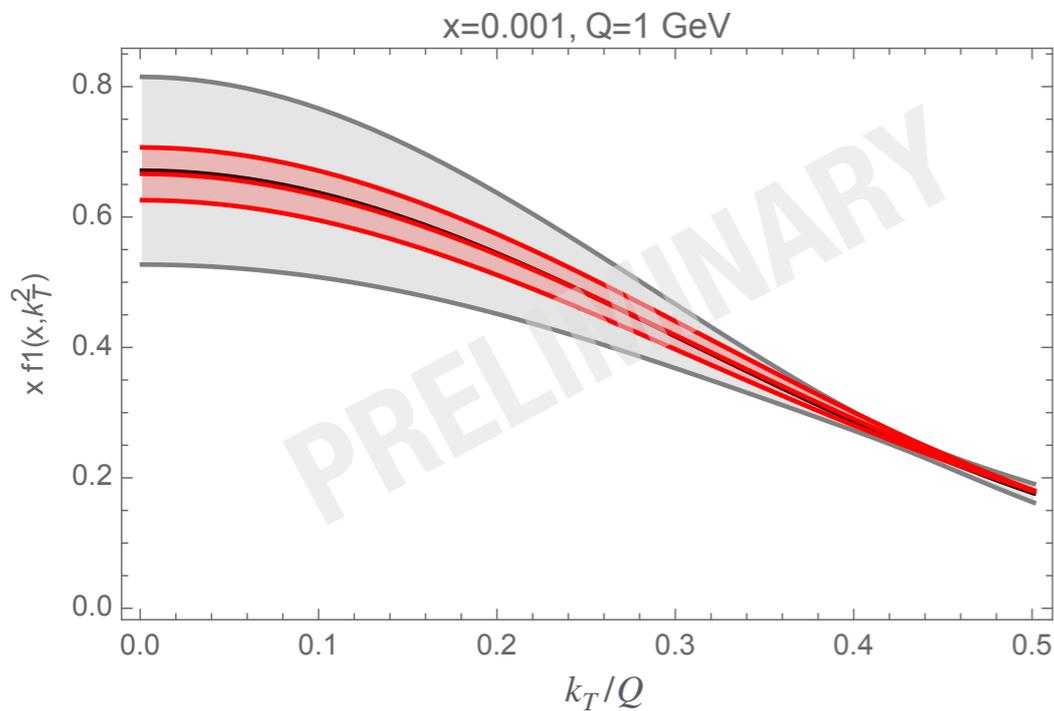
sensitivity coefficients

$$S[f_i, \mathcal{O}] = \frac{\langle \mathcal{O} \cdot f_i \rangle - \langle \mathcal{O} \rangle \langle f_i \rangle}{\delta \mathcal{O} \Delta f_i}$$

A few tools to estimate EIC impact

reweighing

not satisfactory results



sensitivity coefficients

$$S[f_i, \mathcal{O}] = \frac{\langle \mathcal{O} \cdot f_i \rangle - \langle \mathcal{O} \rangle \langle f_i \rangle}{\delta \mathcal{O} \Delta f_i}$$

NECESSARY

new fit that includes EIC pseudodata

Baseline fit

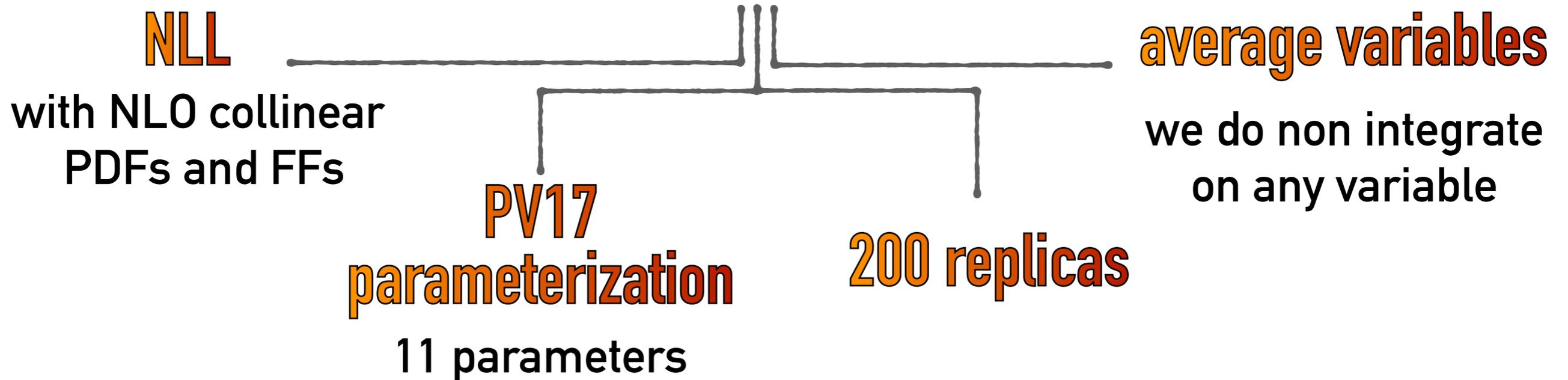
<https://github.com/MapCollaboration/NangaParbat>



with **NangaParbat**
is not possible to replicate
exactly **PV17** results

APFEL++

NEW FIT
to use as baseline



Baseline fit - datasets

NEW FIT

(almost) same datasets as in PV17

global fit



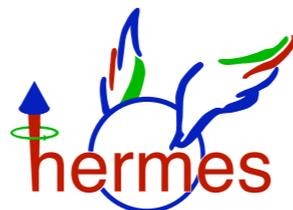
SIDIS



Drell-Yan



2017 COMPASS



Fermilab
low energy

remarks:

- ☼ normalization coeff. for Tevatron
- ☼ NO LHC data
- ☼ **uncertainties correlations taken into account**

cut in q_T/Q

more restrictive than PV17

$$q_T/Q < \min \left[\min[0.2/z, 0.5] + 0.3/zQ, 1 \right]$$

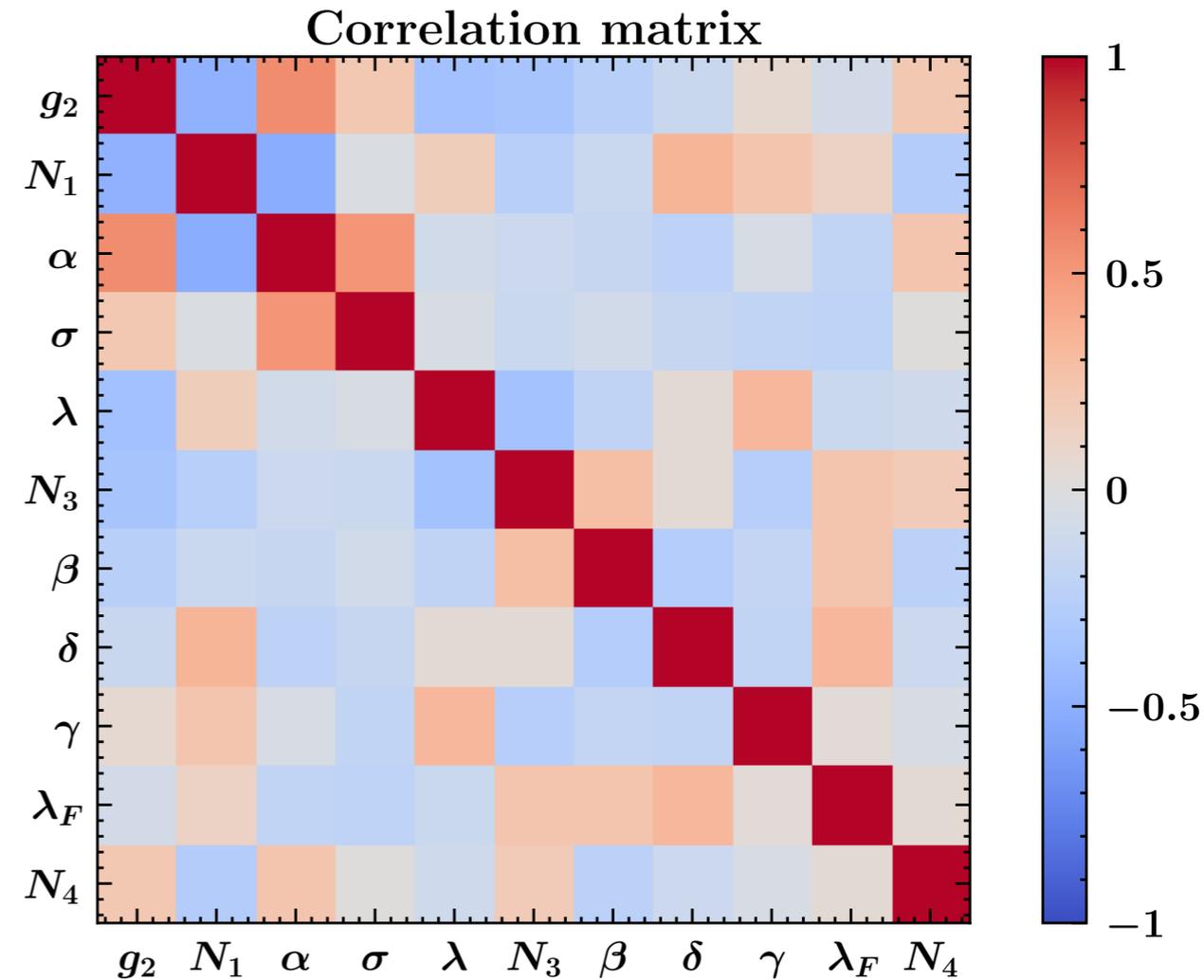
2516

data points

Baseline fit - results

parameters

Parameter	Average over replicas à la PV17	
g_2	0.118	± 0.013
N_1	0.288	± 0.030
α	2.473	± 1.538
σ	-0.142	± 0.110
λ	0.069	± 0.285
N_3	0.220	± 0.021
β	2.948	± 0.966
δ	0.115	± 0.0316
γ	2.454	± 0.151
λ_F	7.022	± 3.298
N_4	0.0317	± 0.006



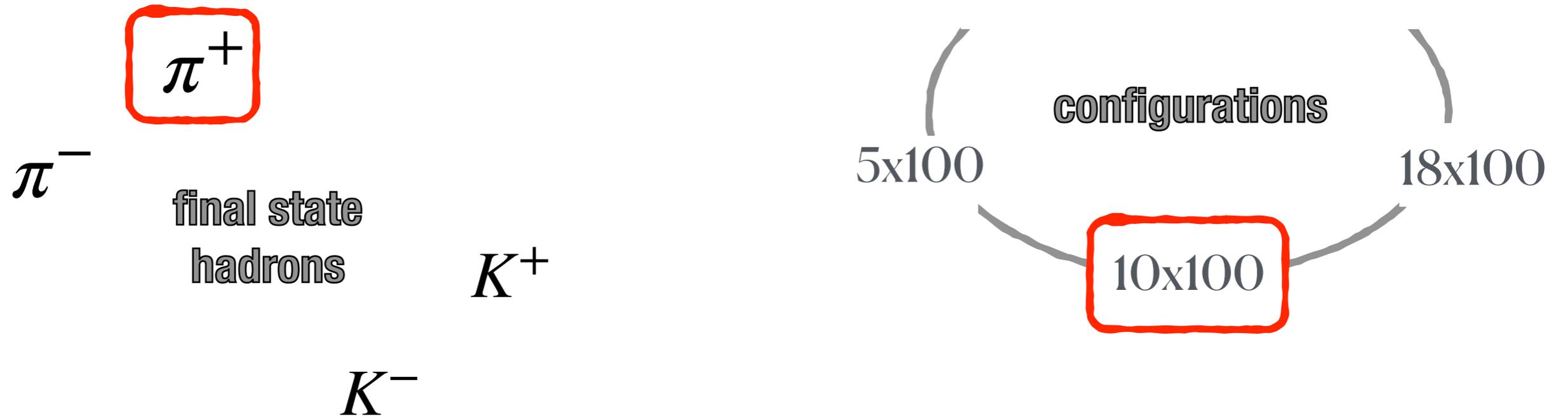
with **NangaParbat**

a **new fit** with uncertainties
similar to **PV17**

$$\chi_{\text{d.o.f.}}^2 = 1.14 \pm 0.06$$

Step towards PV17 + EIC fit

generation of pseudo data



central value of pseudo data obtained
using average parameters of the PV17 baseline fit

uncertainties of pseudo data
are given by simulations

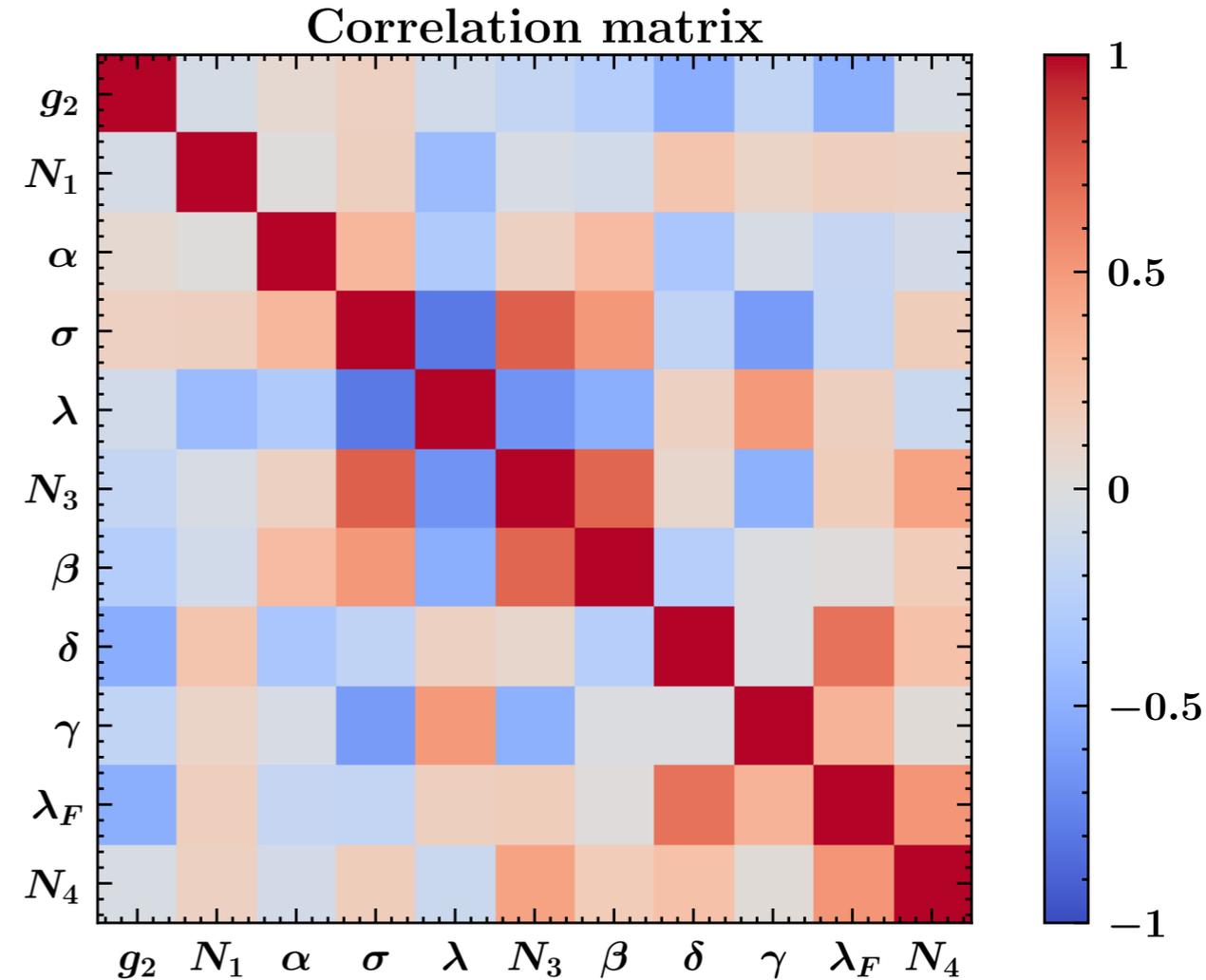
PV17 baseline + EIC fit

impact on parameters

$$g_K(b_T) = -g_2 b_T^2 / 2$$

$$g_1(x) = N_1 \frac{(1-x)^\alpha x^\sigma}{(1-\hat{x})^\alpha \hat{x}^\sigma}$$

$$g_{3,4}(z) = N_{3,4} \frac{(z^\beta + \delta)(1-z)^\gamma}{(\hat{z}^\beta + \delta)(1-\hat{z})^\gamma}$$



$$f_{1\text{NP}}^a(x, \mathbf{k}_\perp^2) = \frac{1}{\pi} \frac{(1 + \lambda \mathbf{k}_\perp^2)}{g_{1a} + \lambda g_{1a}^2} e^{-\frac{\mathbf{k}_\perp^2}{g_{1a}}}$$

$$\chi_{\text{d.o.f.}}^2 = 0.88 \pm 0.32$$

$$D_{1\text{NP}}^{a \rightarrow h}(z, \mathbf{P}_\perp^2) = \frac{1}{\pi} \frac{1}{g_{3a \rightarrow h} + (\lambda_F / z^2) g_{4a \rightarrow h}^2} \left(e^{-\frac{\mathbf{P}_\perp^2}{g_{3a \rightarrow h}}} + \lambda_F \frac{\mathbf{P}_\perp^2}{z^2} e^{-\frac{\mathbf{P}_\perp^2}{g_{4a \rightarrow h}}} \right)$$

Impact of EIC

PV17 baseline

reduction by factor 10

PV17 baseline + EIC

Average over replicas à la PV17		Parameter	Average over replicas à la PV17	
0.118	± 0.013	g_2	0.129	± 0.002
0.288	± 0.030		0.300	± 0.036
2.473	± 1.538		3.409	± 1.079
-0.142	± 0.110		0.2147	± 0.073
0.069	± 0.285	λ	0.136	± 0.551
0.220	± 0.021	N_3	0.213	± 0.0149
2.948	± 0.966	β	2.039	± 0.238
0.115	± 0.0316	δ	0.090	± 0.0181
2.454	± 0.151	γ	2.439	± 0.1114
7.022	± 3.298	λ_F	4.497	± 1.0423
0.0317	± 0.006	N_4	0.0331	± 0.002

non - perturbative evolution

PRELIMINARY

Impact of EIC

PV17 baseline

Average over replicas à la PV17	
0.118	± 0.013
0.288	± 0.030
2.473	± 1.538
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0.069	± 0.285
0.220	± 0.021
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7.022	± 3.298
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reduction by
factor 10

Parameter
g_2
N_1
α
σ
λ
N_3
β
δ
γ
λ_F
N_4

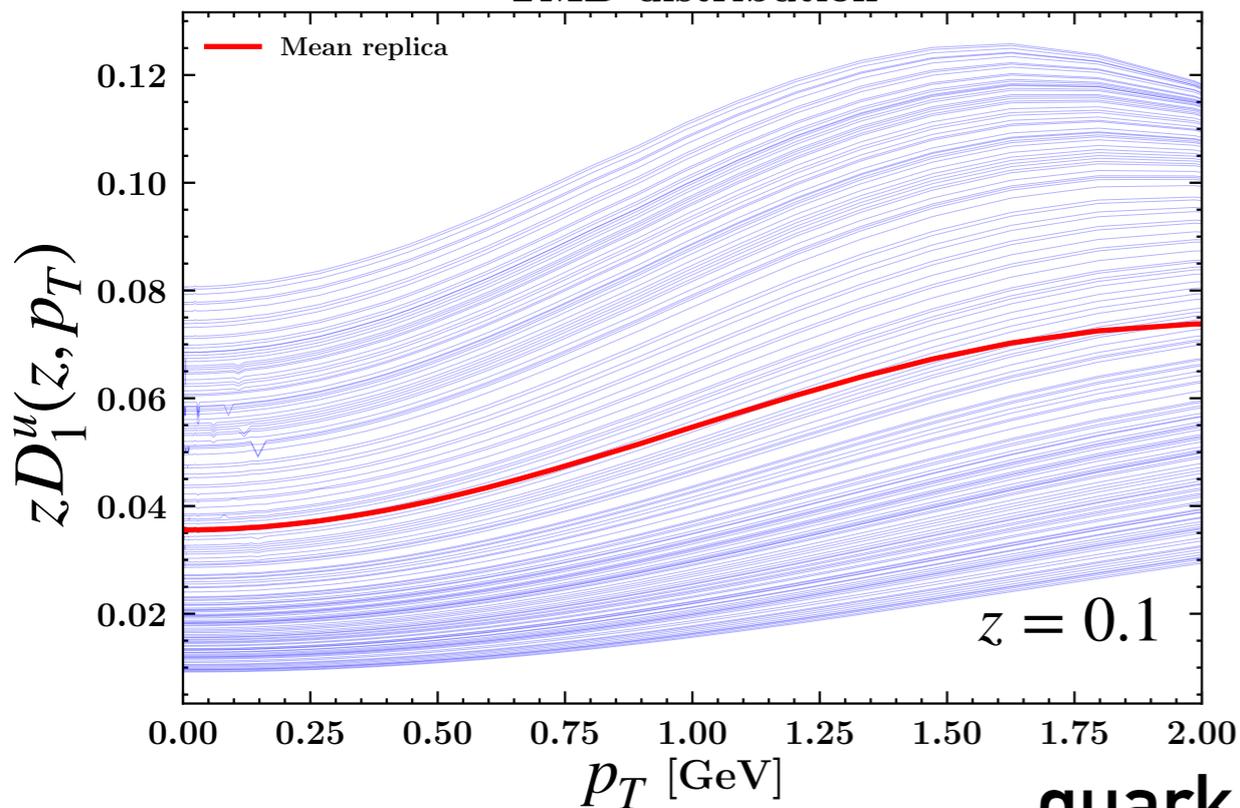
PV17 baseline
+ EIC

Average over replicas à la PV17	
0.129	± 0.002
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2.439	± 0.1114
4.497	± 1.0423
0.0331	± 0.002

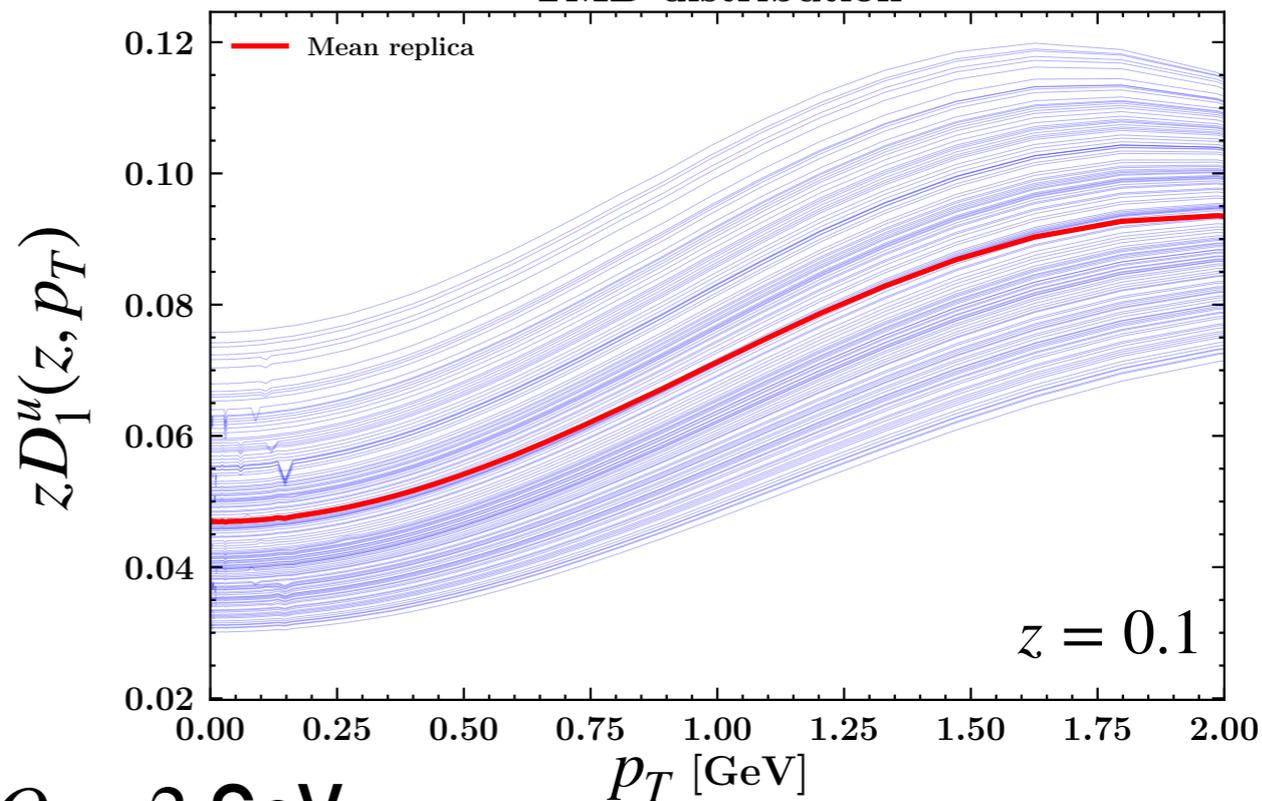
general improvement of uncertainties

Impact on unpolarized TMD FFs

TMD distribution

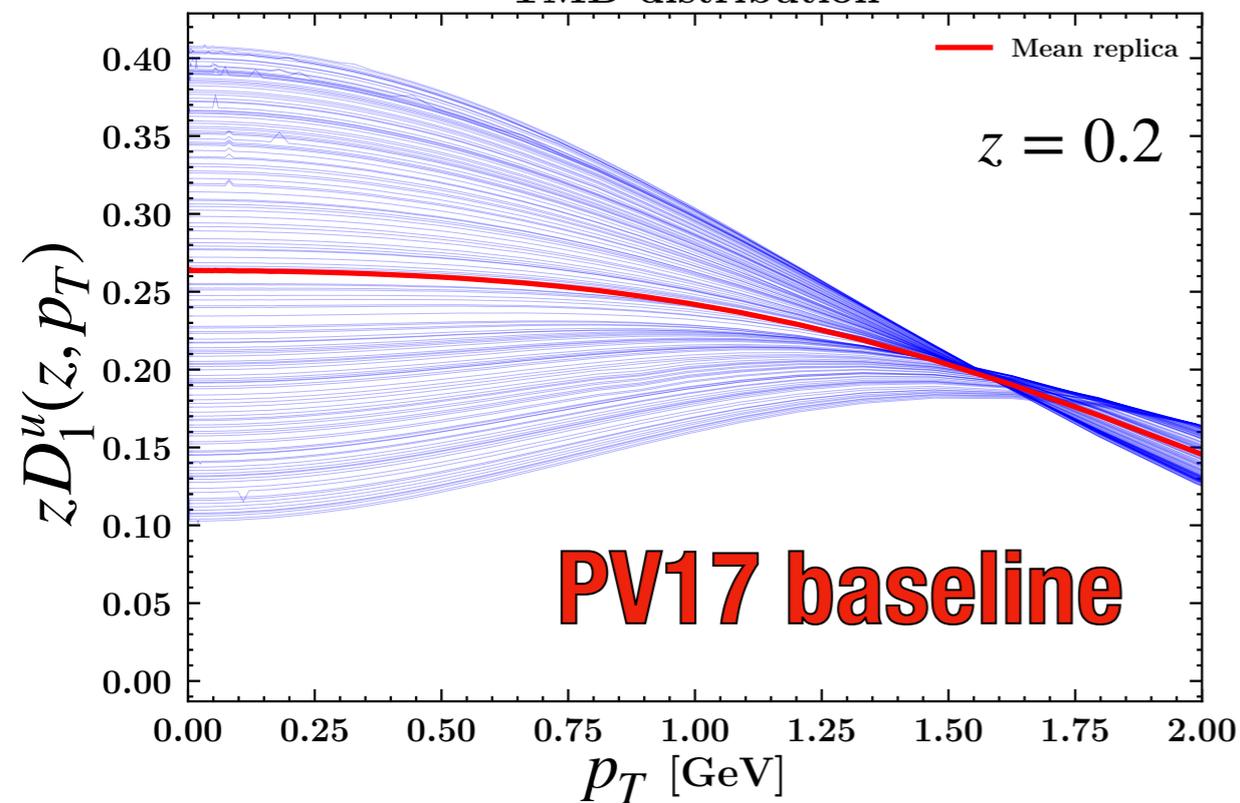


TMD distribution

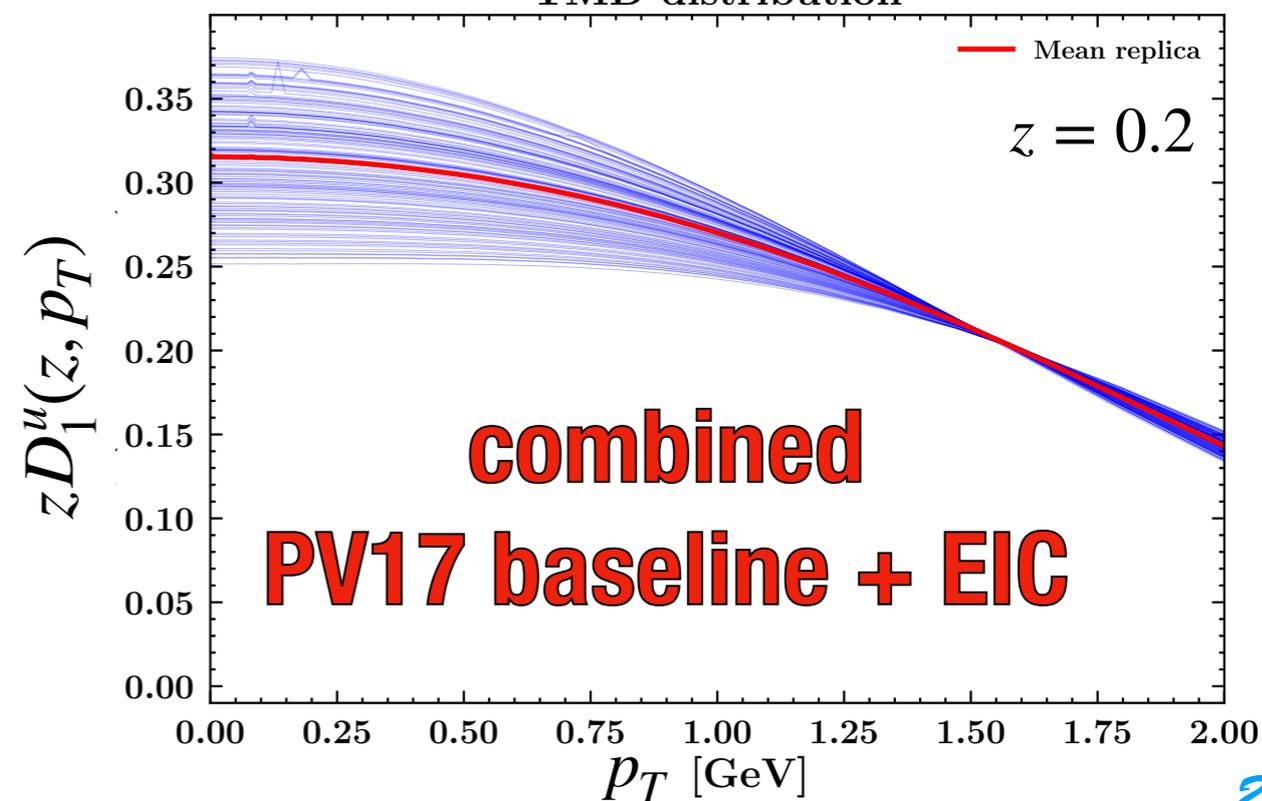


quark up, $Q = 2$ GeV

TMD distribution



TMD distribution



Conclusions

EIC will have a very big **impact** on **TMDs**

will cover a large region
not covered by present data

from **impact studies**,

and in particular from **EIC PSEUDO DATA FITS**

we have **encouraging results** on **uncertainties reduction**

CAVEAT

results depend on
the chosen parameterization

preliminary work

need to include also
other EIC energy configurations

Backup

EIC impact studies

sensitivity coefficients

E. Aschenauer, I. Borsa, G. Lucero, A. S. Nunes, R. Sassot

arXiv:2007.08300

$F_{UU,T}(x, z, q_T; Q^2)$ — **observable** — **distribution** — **TMD parameters**

$$S[f_i, \mathcal{O}] = \frac{\langle \mathcal{O} \cdot f_i \rangle - \langle \mathcal{O} \rangle \langle f_i \rangle}{\xi \Delta \mathcal{O} \Delta f_i}$$

experimental uncertainty
(from pseudodata)

$$\xi \equiv \frac{\delta \mathcal{O}}{\Delta \mathcal{O}}$$

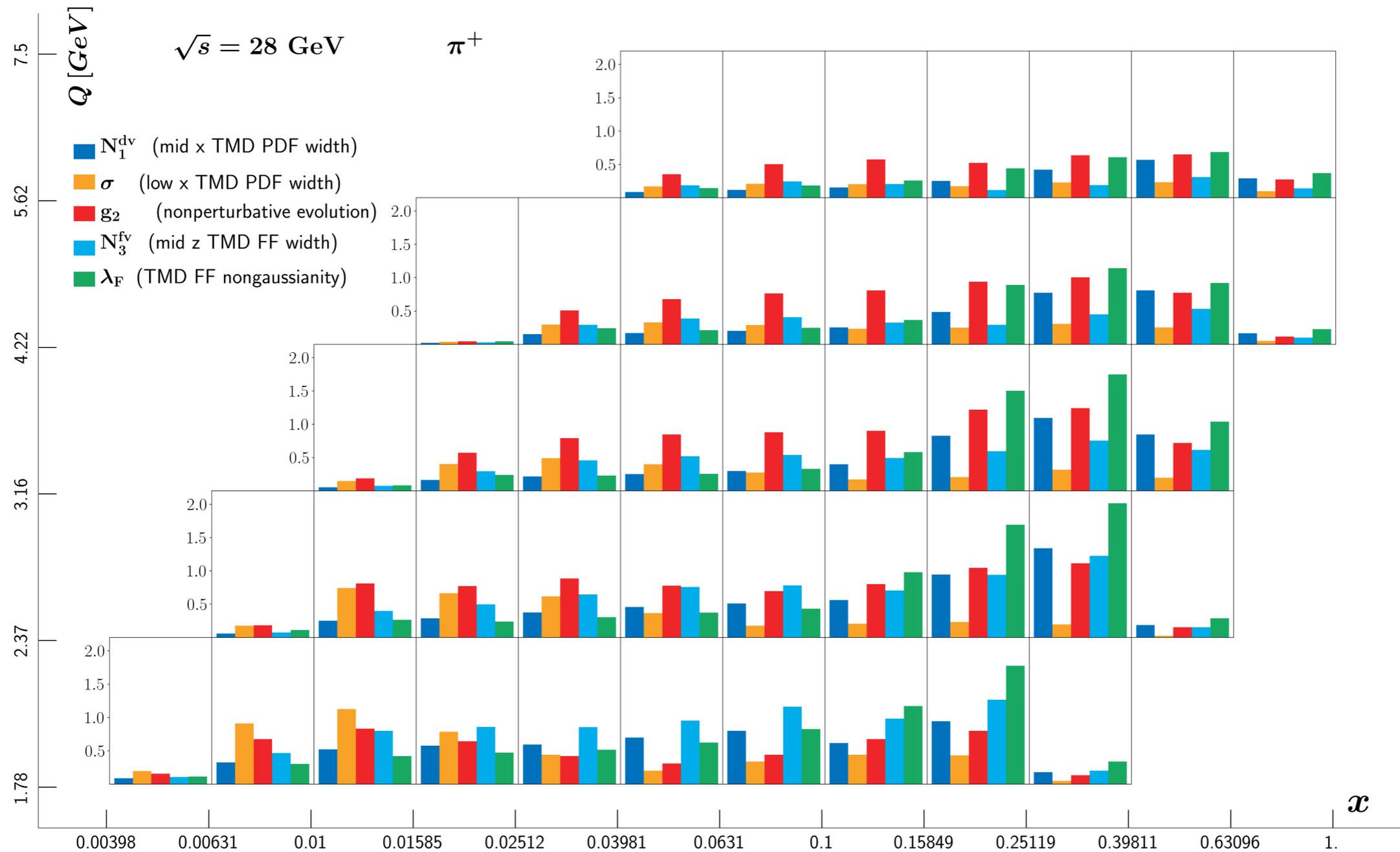
~~theoretical uncertainty~~

$$\langle \mathcal{O} \rangle = \frac{1}{N} \sum_{k=1}^N \mathcal{O}[f_i^{(k)}]$$

ELC impact studies

sensitivity coefficients

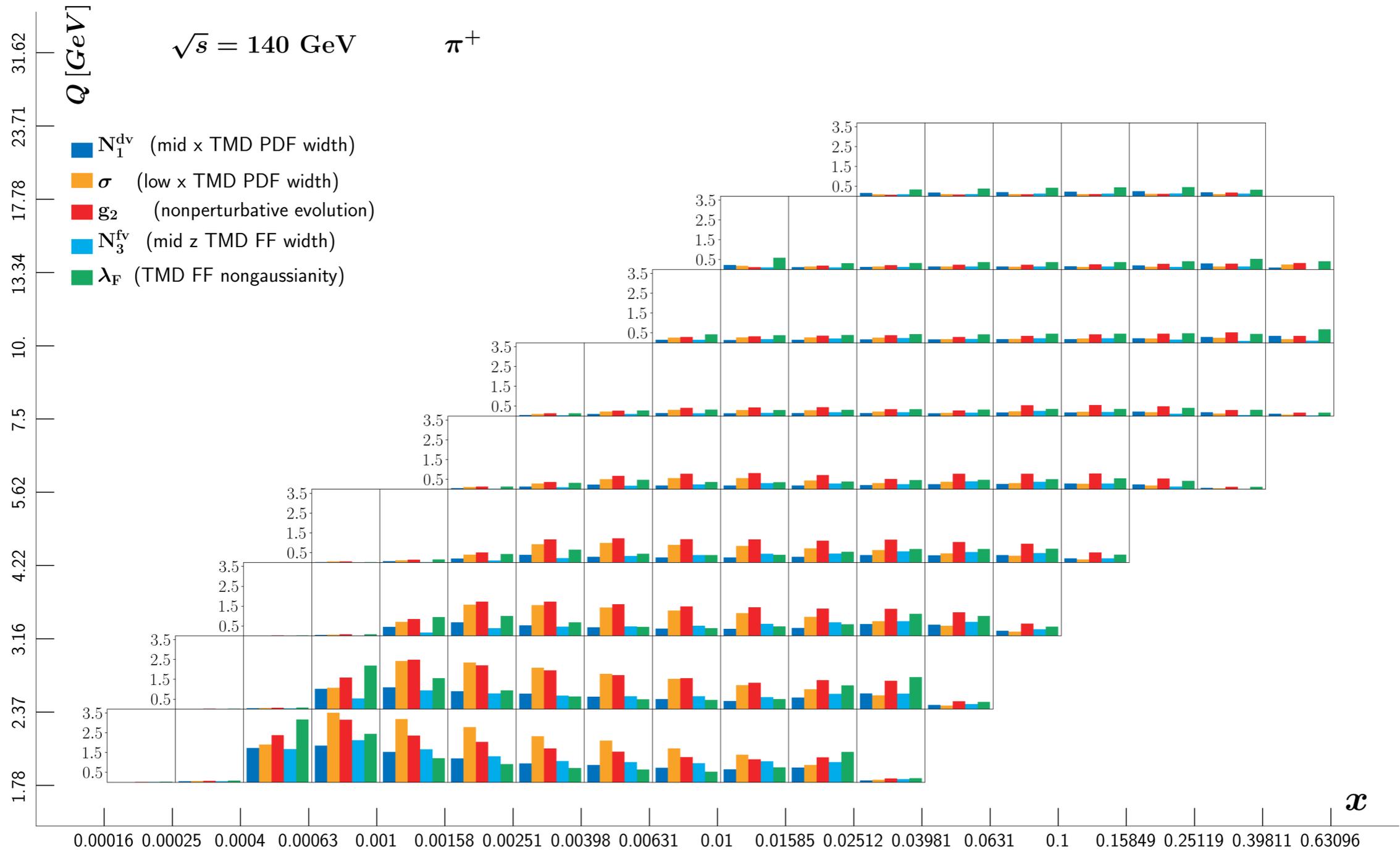
$$S[f_i, \mathcal{O}] = \frac{\langle \mathcal{O} \cdot f_i \rangle - \langle \mathcal{O} \rangle \langle f_i \rangle}{\delta \mathcal{O} \Delta f_i}$$



ELC impact studies

sensitivity coefficients

$$S[f_i, \mathcal{O}] = \frac{\langle \mathcal{O} \cdot f_i \rangle - \langle \mathcal{O} \rangle \langle f_i \rangle}{\delta \mathcal{O} \Delta f_i}$$



Sensitivity coefficient and standard deviation

how to estimate uncertainties reduction?

$$S[\mathcal{O}, f] = \frac{\langle \mathcal{O} f \rangle - \langle \mathcal{O} \rangle \langle f \rangle}{\delta \mathcal{O} \Delta f}$$

from PV17

we know a parameter A with error ΔA

☀ if we perform a new measurement that produces on A an error equal to its initial standard deviation, $\delta A = \Delta A$

→ this corresponds to $S(A) = 1$



in fact, if A can be ideally considered as parameter and observable, then

☀ the error on A scales as
 $1/\sqrt{2} = 1/\sqrt{1 + (S = 1)}$

$$S(A, A) = \frac{\langle A A \rangle - \langle A \rangle \langle A \rangle}{\delta A \Delta A} = \frac{(\Delta A)^2}{\Delta A \Delta A} = 1$$

if the new measurement is more precise, then $S > 1$ and the error is further reduced

for n measurements, the error on A should scale as

$$1/\sqrt{1 + S_1 + \dots + S_n}$$

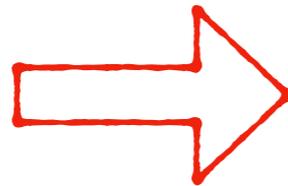
Non perturbative evolution g_2

which energy configuration has the highest impact?

summing over all (x, Q^2) bins

PV17 fit

$$\Delta g_2 = 0.01$$



$$R(A) = \frac{(\Delta A)_{\text{prev}}}{(\delta A)_{\text{EIC}}}$$

run at $\sqrt{s} = 28$ GeV π^+	→	0.00155
run at $\sqrt{s} = 44$ GeV π^+	→	0.00120
run at $\sqrt{s} = 63$ GeV π^+	→	0.00108
run at $\sqrt{s} = 84$ GeV π^+	→	0.00105
run at $\sqrt{s} = 140$ GeV π^+	→	0.00096

$$R(g_2) = 6.45$$

$$R(g_2) = 8.33$$

$$R(g_2) = 9.26$$

$$R(g_2) = 9.52$$

$$R(g_2) = 10.36$$

(no correlation between measurements in different bins)

consistent trend:
 evolution parameter better constrained by larger covered (x, Q^2)
 → larger \sqrt{s} configuration

Attempts at reweighting

200 replicas are compared
with pseudodata

$$\chi_k^2 = \chi_{k,\text{EIC}}^2 + \chi_{k,\text{PV17}}^2$$

'original' χ^2
with respect to PV17 data

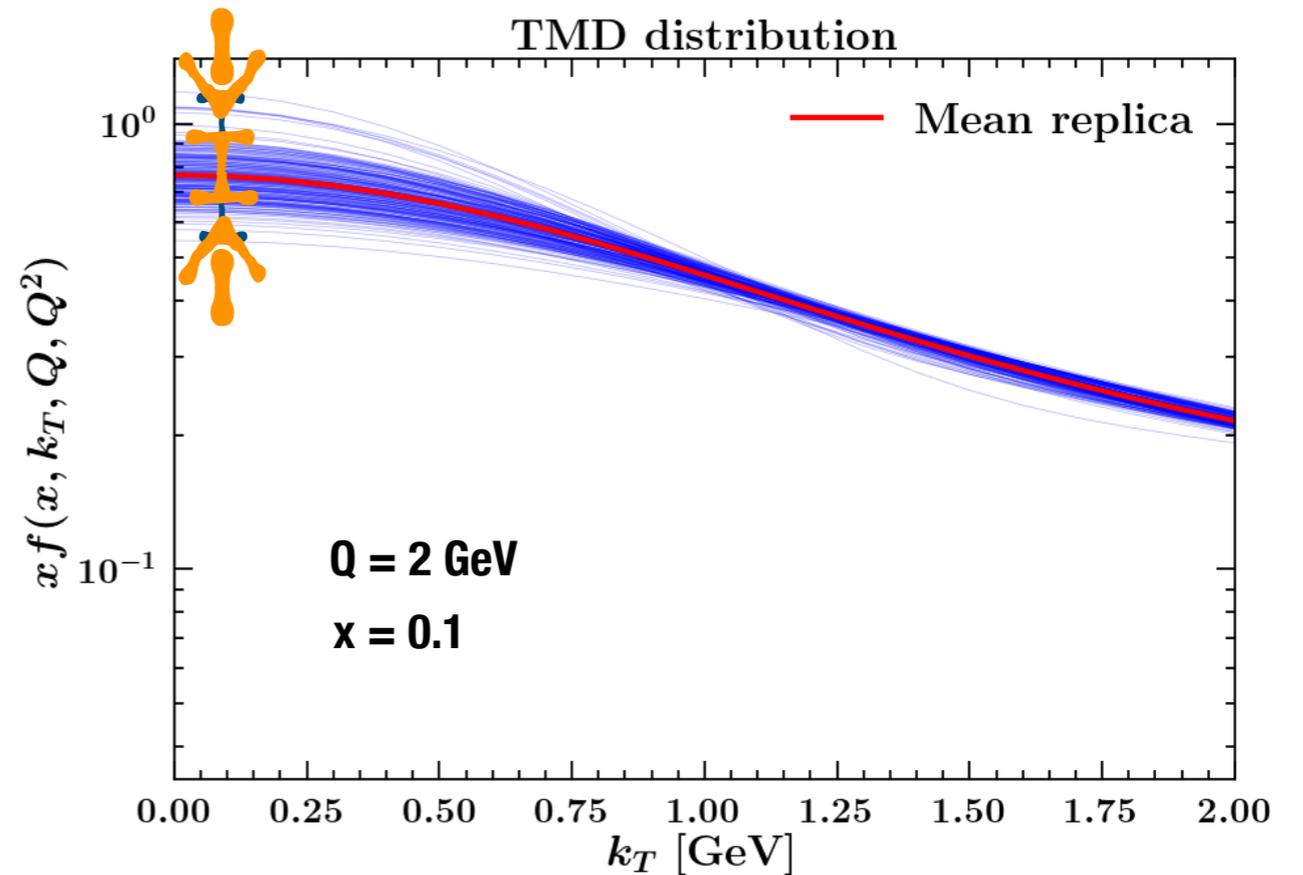
weights

$$w_k \propto \mathcal{P}(f_k | \chi_k) \propto \chi_k^{n-1} e^{-\frac{1}{2}\chi_k^2}$$

used to select replicas

from NNPDF Collaboration
[arXiv:1108.1758](https://arxiv.org/abs/1108.1758)

reflect the impact of EIC data on extracted TMDs



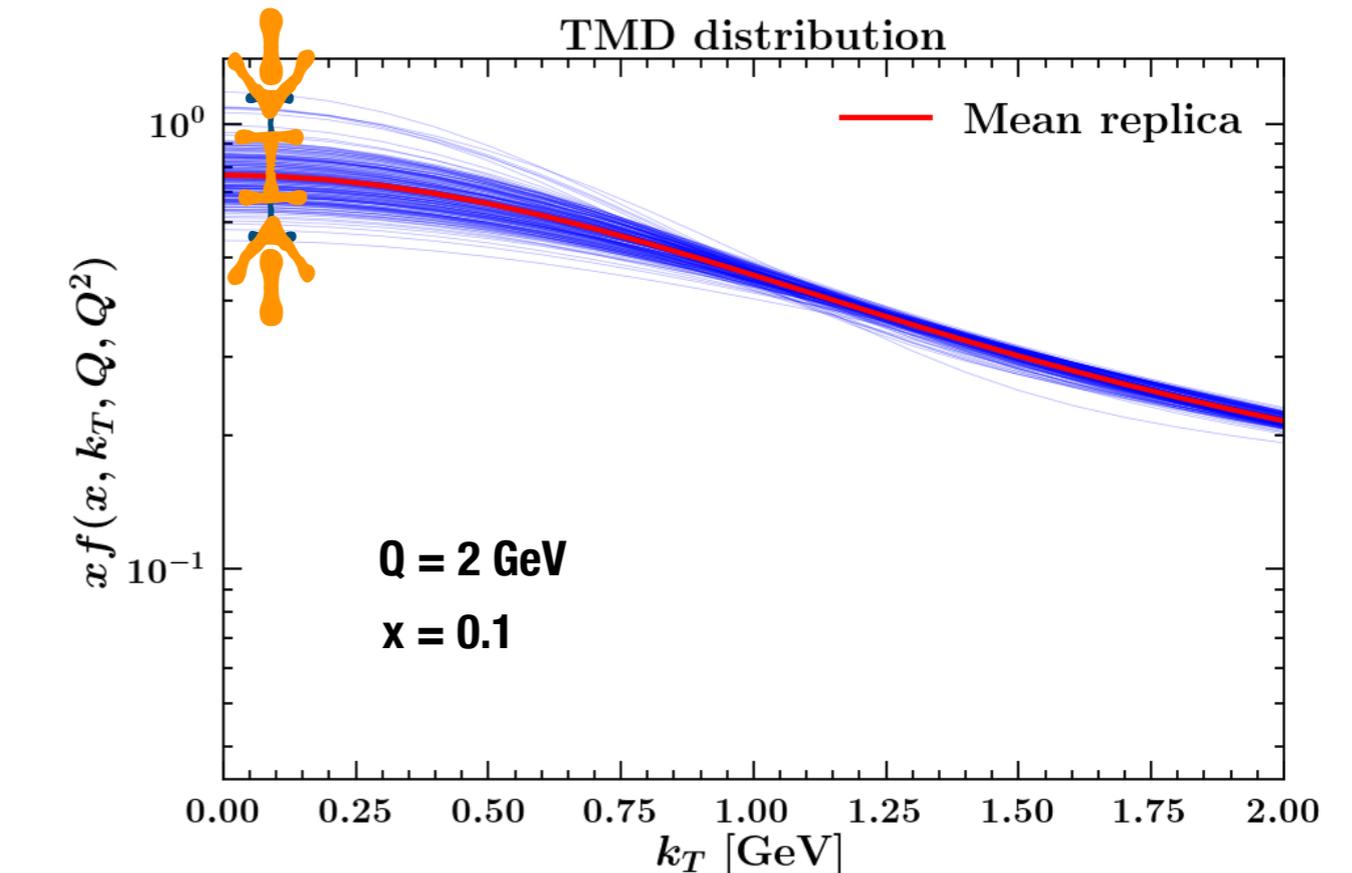
Attempts at reweighting

different mathematical formulas
to compute the weights

N. Sato, J. Owens, H. Prosper, PRD 89 (2014) 114020;
H. Paukkunen, P. Zurita, JHEP 12 (2014) 100

$$w_k \propto \mathcal{P}(f_k | \chi_k) \propto e^{-\frac{1}{2} \chi_k^2}$$

selects replicas with very low χ^2



NNPDF Collaboration
[arXiv:1108.1758](https://arxiv.org/abs/1108.1758)

total number
of points

$$w_k \propto \mathcal{P}(f_k | \chi_k) \propto \chi_k^{n-1} e^{-\frac{1}{2} \chi_k^2}$$

suppresses replicas with very
high AND very low χ^2