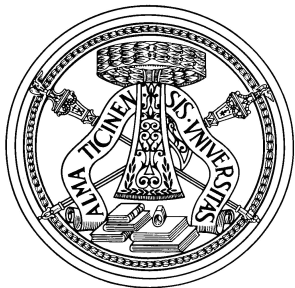


Light dark forces at flavour factories - [1007.4984]



Luca Barzè

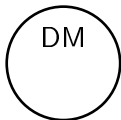
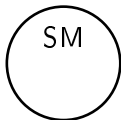
University of Pavia
INFN

Discrete 2010
10th December 2010

with Balossini, Bignamini, Carloni
Calame, Montagna, Nicosini, Piccinini

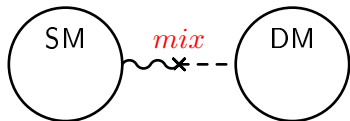
Axiom 1: Dark Matter exists

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{DM}$$



Axiom 1: Dark Matter exists and interacts with SM

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{DM} + \mathcal{L}_{mix}$$



$$\mathcal{L}_{mix} = \sum_{ij} k_{ij} \Theta_{SM}^i \Theta_{DM}^j$$

A simple way

$$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$$

$$\mathcal{L}_{SM} = \mathcal{L}_{SM}^F + \mathcal{L}_{SM}^B + \mathcal{L}_{SM}^H$$

$$\mathcal{L}_{DM} = ?$$

A simple way: a New Symmetry

$$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \otimes U(1)_{DM} \otimes \dots$$

$$\mathcal{L}_{SM} = \mathcal{L}_{SM}^F + \mathcal{L}_{SM}^B + \mathcal{L}_{SM}^H$$

$$\begin{aligned} \mathcal{L}_{DM} &= \mathcal{L}_{DM}^F(\chi) && \Rightarrow M_\chi \sim 10 - 100 \text{ GeV (WIMP)} \\ &+ \mathcal{L}_{DM}^B(U) && \Rightarrow m_U \sim ? \\ &+ \dots \end{aligned}$$

A simple way: a New Symmetry

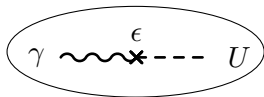
$$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \otimes U(1)_{DM} \otimes \dots$$

$$\mathcal{L}_{SM} = \mathcal{L}_{SM}^F + \mathcal{L}_{SM}^B + \mathcal{L}_{SM}^H$$

$$\begin{aligned} \mathcal{L}_{DM} &= \mathcal{L}_{DM}^F(\chi) && \Rightarrow M_\chi \sim 10 - 100 \text{ GeV (WIMP)} \\ &+ \mathcal{L}_{DM}^B(U) && \Rightarrow m_U \sim ? \\ &+ \dots \end{aligned}$$

$$\mathcal{L}_{mix} = \frac{\epsilon_Y}{2} F^{DM\mu\nu} F_{\mu\nu}^Y, \quad \epsilon \equiv \epsilon_Y c_W$$

$$\mathcal{L}_{mix} = \frac{\epsilon}{2} F^{DM\mu\nu} F_{\mu\nu}^{EM} \quad \leftarrow \text{SSB}$$



Small effects at low energies.

A simple way: a New Symmetry

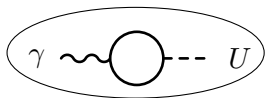
$$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \otimes U(1)_{DM} \otimes \dots$$

$$\mathcal{L}_{SM} = \mathcal{L}_{SM}^F + \mathcal{L}_{SM}^B + \mathcal{L}_{SM}^H$$

$$\begin{aligned} \mathcal{L}_{DM} &= \mathcal{L}_{DM}^F(\chi) && \Rightarrow M_\chi \sim 10 - 100 \text{ GeV (WIMP)} \\ &+ \mathcal{L}_{DM}^B(U) && \Rightarrow m_U \sim ? \\ &+ \dots \end{aligned}$$

$$\mathcal{L}_{mix} = \frac{\epsilon_Y}{2} F^{DM\mu\nu} F_{\mu\nu}^Y, \quad \epsilon \equiv \epsilon_Y c_W$$

$$\mathcal{L}_{mix} = \frac{\epsilon}{2} F^{DM\mu\nu} F_{\mu\nu}^{EM} \quad \leftarrow \text{SSB}$$



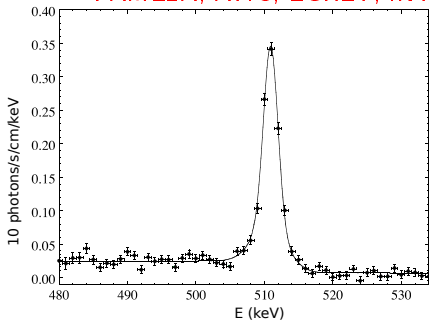
Small effects at low energies.

Motivation

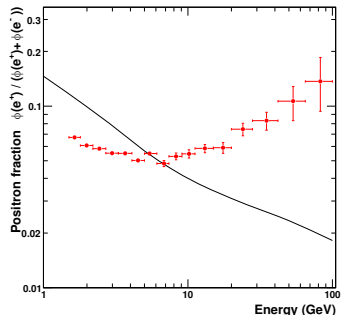
- Compatible with a number of theories;

Motivation

- Compatible with a number of theories;
- it would be possible to explain some experimental data:
 - PAMELA, ATIC, EGRET, INTEGRAL, DAMA, FERMI ...



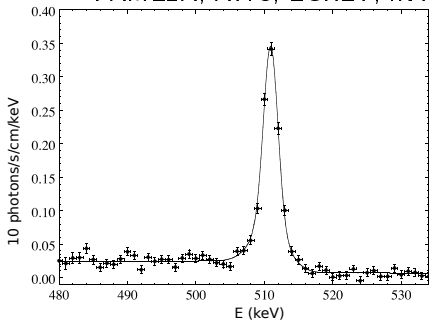
511 keV line - INTEGRAL



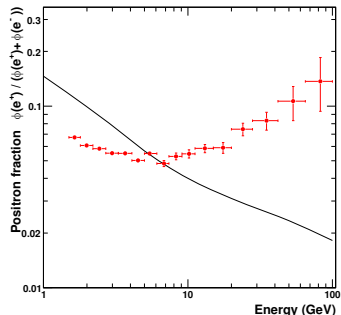
Excess of positrons - PAMELA

Motivation

- Compatible with a number of theories;
- it would be possible to explain some experimental data:
 - PAMELA, ATIC, EGRET, INTEGRAL, DAMA, FERMI ...



511 keV line - INTEGRAL

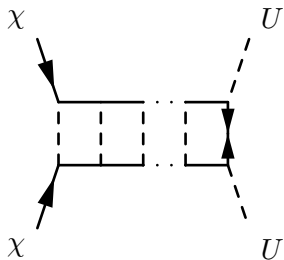


Excess of positrons - PAMELA

Astrophysical sources → difficulties

Axiom 2: Data due to DM

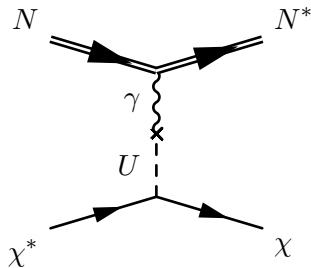
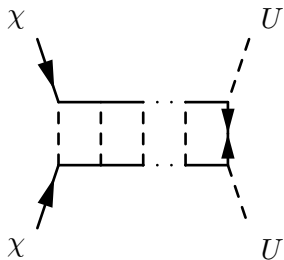
An excess of e^+ without \bar{p}



No \bar{p} excess $\rightarrow U$ must be light (\sim MeV - GeV) $\Rightarrow \epsilon \lesssim 10^{-2,-3}$

*hep-ph[0810.0713] - Arkani-Hamed, Finkbeiner, Slatyer, Weiner
A Theory of Dark matter*

An excess of e^+ without \bar{p}



DAMA/CoGeNT signals

No \bar{p} excess $\rightarrow U$ must be light (\sim MeV - GeV) $\Rightarrow \epsilon \lesssim 10^{-2,-3}$

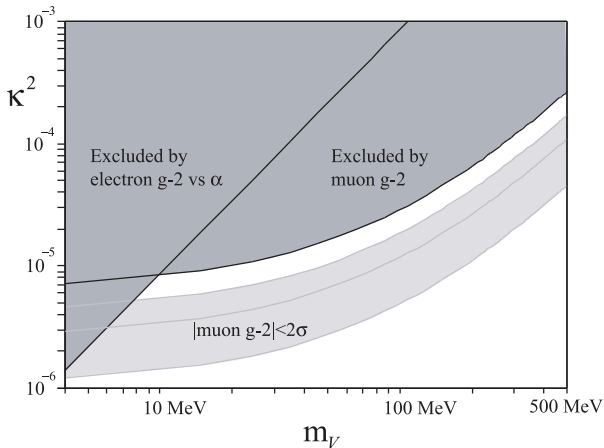
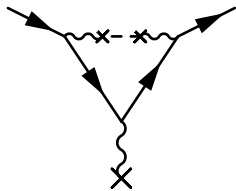
*hep-ph[0810.0713] - Arkani-Hamed, Finkbeiner, Slatyer, Weiner
A Theory of Dark matter*

Constraints:

From Particle Physics... ... From Astrophysics

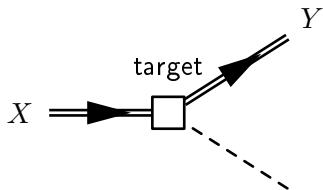
There is quite a wide window not excluded by any obvious laboratory measurement or astrophysical argument, while the INTEGRAL or observation could be easily accounted for

Predictions are testable: anomalous magnetic moment



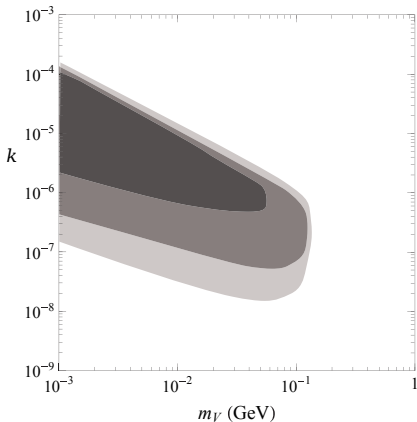
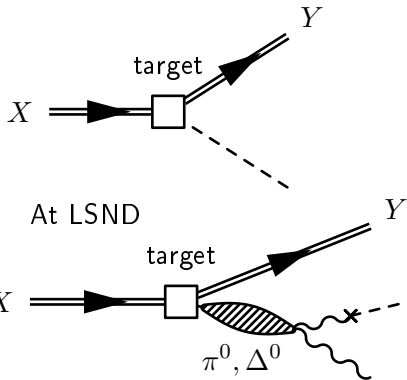
*hep-ph[0811.1030] - Pospelov
Secluded U(1) below the weak scale*

Predictions are testable: beam dump



*hep-ph[0906.5614] - Batell, Pospelov, Ritz
Exploring portals to a hidden sector through fixed targets*

Predictions are testable: beam dump



*hep-ph[0906.5614] - Batell, Pospelov, Ritz
Exploring portals to a hidden sector through fixed targets*

Flavour factories: An Ideal Environment

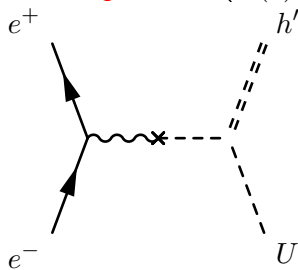
- Low energy \sim GeV ($\sigma \propto E^{-2}$);
- high luminosity (up to ab^{-1} at BaBar/Belle);
- clear signatures:

Flavour factories: An Ideal Environment

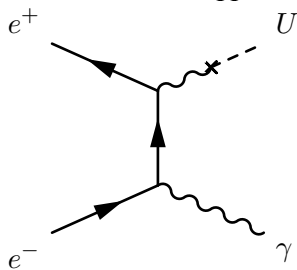
- Low energy \sim GeV ($\sigma \propto E^{-2}$);
- high luminosity (up to ab^{-1} at BaBar/Belle);
- clear signatures:

Flavour factories: An Ideal Environment

- Low energy $\sim \text{GeV}$ ($\sigma \propto E^{-2}$);
- high luminosity (up to ab^{-1} at BaBar/Belle);
- **clear signatures:** ($U(1)_{DM}$ broken $\sim \text{GeV} \Rightarrow$ dark Higgs $\sim \text{GeV}$)

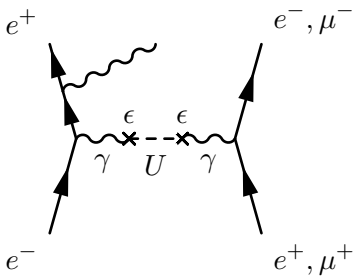


resonances ($\Upsilon, \Phi \rightarrow X + \cancel{E}_T$)
model dependent



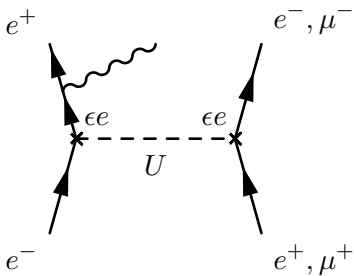
$l^+l^-\gamma$
model independent

A really difficult channel



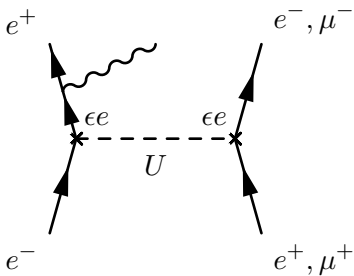
- Resonant channel:
 - particular signal shape \neq BG;
- radiative return:
 - energy scan;

A really difficult channel



- Resonant channel:
 - particular signal shape \neq BG;
- radiative return:
 - energy scan;

A really difficult channel



- Resonant channel:
 - particular signal shape \neq BG;
- radiative return:
 - energy scan;

- 2nd order process,
- 2 ϵ :
 - really small signal! ($\sigma_U \sim 10^{-7} \sigma_{BG}$)
($\sigma_Z(1 \text{ GeV}) \sim 10^{-3} \sigma_{BG}$)
- An accurate estimate of the background is mandatory.

Necessity of a very accurate event generator

$$| \text{[Feynman diagrams]} + \dots + \dots |^2 \quad 14 \text{ terms for } e^\pm, 6 \text{ for } \mu^\pm$$

ALPHA

BabaYaga

- Exact tree level calculation;
- very well tested generator.

hep-ph[0607181] - Balossini, Carloni Calame, Montagna, Nicosini, Piccinini
Matching perturbative and Parton Shower corrections to Bhabha process at flavour factories

hep-ph[9507237v1] - Caravaglios, M. Moretti
An algorithm to compute Born scattering amplitudes without Feynman graphs

BabaYaga

A MCEG for $e^+e^- \rightarrow e^+e^-, \mu^+\mu^-, \gamma\gamma$ processes at flavour factories.

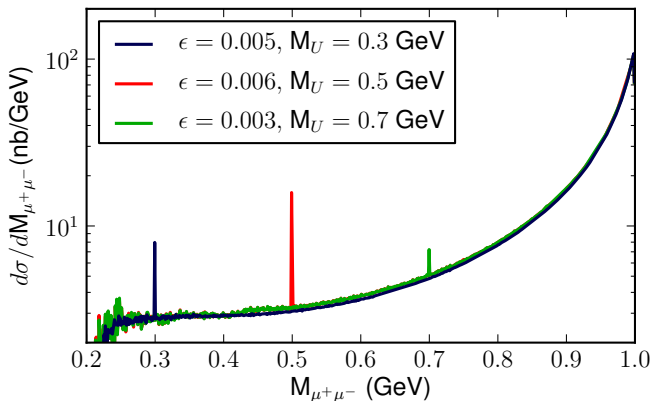
- Commonly used for the determination of flav. factories luminosity:
 - $(g-2)_\mu, R, \Delta\alpha_{had}$;
- theoretical error $\sim 1\%$ ($\mathcal{O}(\alpha^2)$) for first order processes.

<http://www.pv.infn.it/hepcomplex/babayaga.html>

A MCEG for Light Dark Matter at Leptonic Colliders

- Exact tree level calculation for the process
 $e^+e^- \rightarrow U, Z, \gamma \rightarrow l^+l^-\gamma$;
- exact three body kinematics;
- vacuum polarization (hadronic contribution \rightarrow [HADR5N09](#) from Jegerlehner
 \rightarrow [HMNT](#) from Teubner et al.);
- radiative corrections \rightarrow structure functions of the electron;
- theoretical error $\mathcal{O}(\alpha)$ (second order processes).

A possible signal

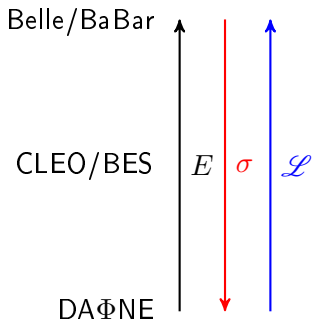


Final state invariant mass for different ϵ and m_U .

- Huge background on e^+e^-

Statistical significance

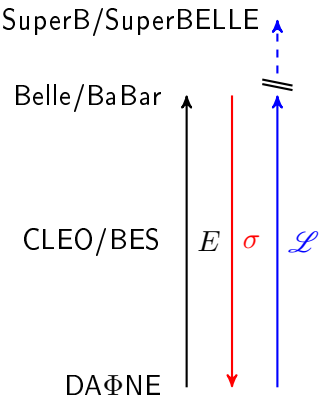
$$\frac{\#S}{\sqrt{\#B}} = \frac{\mathcal{L}(\sigma_{SM+U} - \sigma_{SM})}{\sqrt{\mathcal{L}\sigma_{SM}}} \equiv \sqrt{\mathcal{L}} \frac{\sigma_S}{\sqrt{\sigma_{SM}}} > 5 \text{ for discovery}$$



*hep-ph[0904.1743] - Reece, Wang
Searching for the light dark gauge boson in
GeV-scale experiments*

Statistical significance

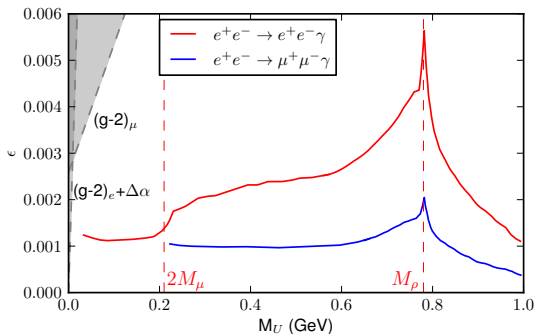
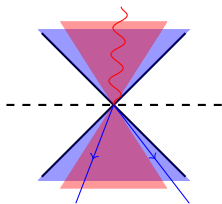
$$\frac{\#S}{\sqrt{\#B}} = \frac{\mathcal{L}(\sigma_{SM+U} - \sigma_{SM})}{\sqrt{\mathcal{L}\sigma_{SM}}} \equiv \sqrt{\mathcal{L}} \frac{\sigma_S}{\sqrt{\sigma_{SM}}} > 5 \text{ for discovery}$$



*hep-ph[0904.1743] - Reece, Wang
Searching for the light dark gauge boson in
GeV-scale experiments*

Simulation's results

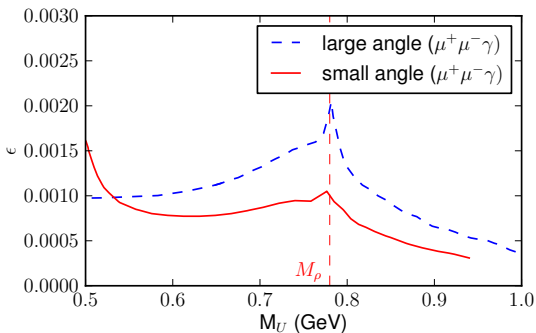
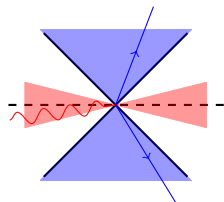
Large Angle Selection



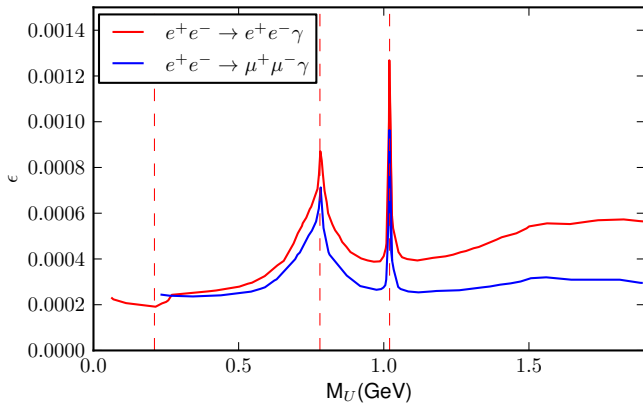
5 σ reach at KLOE+KLOE2 (5 fb^{-1} - 1.02 GeV)

Simulation's results

Small Angle Selection

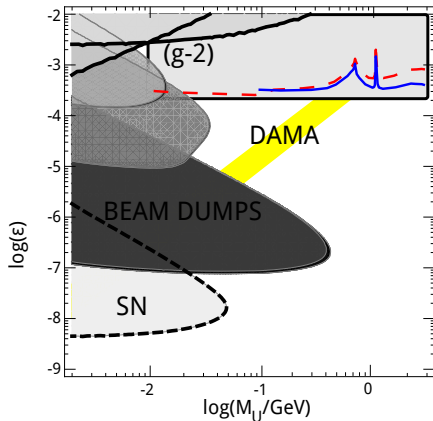


5σ reach at KLOE+KLOE2 (5 fb^{-1} - 1.02 GeV)



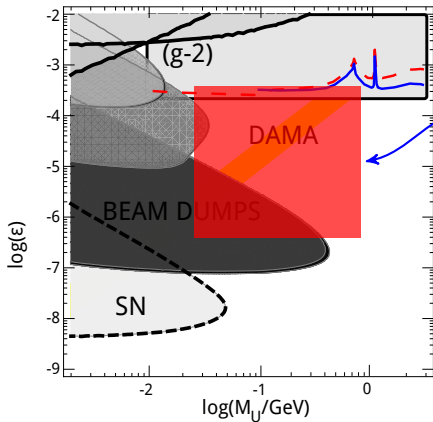
5 σ reach at Possible SuperB (100 ab^{-1} - 10.56 GeV)

Not satisfied?

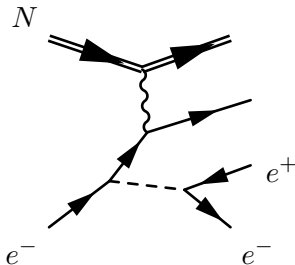


[0904.1743] - Bjorken, Essig, Schuster, Toro
New Fixed-Target Experiments to Search for Dark Gauge Forces

Not satisfied?



Fixed target experiment
really high luminosity!



[0904.1743] - Bjorken, Essig, Schuster, Toro
New Fixed-Target Experiments to Search for Dark Gauge Forces

A lot of work to do!

- **Experimental:**
 - analyze existing data;
 - produce new data (flavour factories, beam dump);
- model builders:
 - explain the experimental data;
 - explain ALL the data at the same time;
- phenomenologists;
 - describe other possible signatures;
 - prepare accurate event generators.

A lot of work to do!

- Experimental:
 - analyze existing data;
 - produce new data (flavour factories, beam dump);
- **model builders:**
 - **explain the experimental data;**
 - **explain ALL the data at the same time;**
- phenomenologists;
 - describe other possible signatures;
 - prepare accurate event generators.

A lot of work to do!

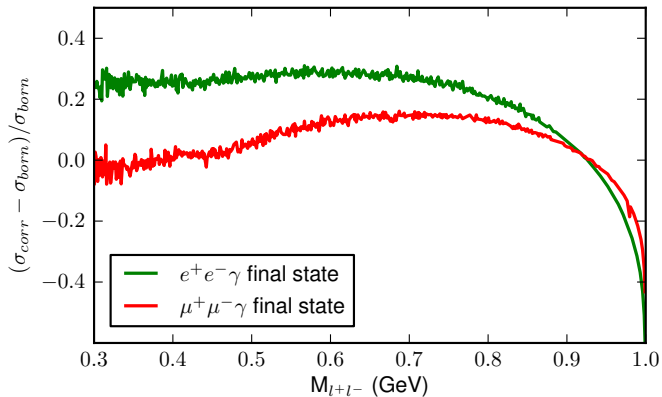
- Experimental:
 - analyze existing data;
 - produce new data (flavour factories, beam dump);
- model builders:
 - explain the experimental data;
 - explain ALL the data at the same time;
- phenomenologists;
 - describe other possible signatures;
 - prepare accurate event generators.



to Gauss & Kandinsky by G Martinelli

THANK YOU!

Effects of radiative correction



Width of U boson

