Cosmogenic backgrounds in Plan C

Shawn Westerdale (presenting work from Zoe, Teena, Sagar, and others) Backgrounds call 28 May 2021

Muon-induced backgrounds



Two dangerous types of background

Material activation

- Muon passing through detector components creates unstable isotopes with µs—hrs half-life
- Particular danger from β-delayed neutrons produced near TPC

Prompt neutrons

- Produce by spallation, shower, etc. induced by muon while it traverses rock, lab, and detector
- High energy, may penetrate shielding
- Muon doesn't need to come near the detector for a neutron to reach TPC
 - Neutrons may even emerge from rock with the muon

Some facts about prompt neutrons



Fish can't live in a LAr tank, so we can't expect Poisson stats



Mei and Hime. Phys Rev D 73, 053004 (2006) A. Empl et al JCAP08(2014)064

Some facts about prompt neutrons

'n

Neutrons can **stray** even farther from the muon track

Significant neutron production **displaced** from muon track And the neutrons travel **far** once they are produced



To catch a lot of muons, you need a big net



Muon caught – background vetoed :)⁵

Muon missed – background!

Plan C FLUKA simulations

- 30.6 years of muons + showers simulated in Hall C by Toni Empl
- Hall C muons propagated into Plan C geometry by Sagar, further development underway by Teena
- Disclaimers:
 - FLUKA output is cumbersome and complicated; currently existing FLUKA simulations are limited in what variables are stored for each event
 - Existing FLUKA output records
 - Total energy deposited in TPC, neutron veto, and LAr bath
 - ID of all particles entering the TPC
 - Kinetic energy of all neutrons that entered the TPC
 - Cannot apply fiducial and multiple scatter cuts with this maybe they would improve things by a factor of ~a
 few
 - Some very rough approximations can be explored by cutting on what/how many particles enter TPC
 - Timescale for additional MC with higher stats and more variables stored > 1 month

Plan C FLUKA simulations



There are likely some minor differences in geometry (e.g. Cu vs. Ti barriers), but covers main elements

Plan C FLUKA simulations, 30.6 yrs: No µ veto



Plan C FLUKA simulations, 30.6 yrs: Adding µ veto

No µ veto

No NVeto, 1 visible TPC neutron : 234 NVeto thresh 500 keV, 1 visible TPC neut : 38 NVeto thresh 500 keV, precisely 1 TPC neut: 28 NVeto thresh 100 keV, 1 visible TPC neut : 15 NVeto thresh 0 keV, 1 visible TPC neut : 9 NVeto thresh 0 keV, precisely 1 TPC neut : 6 **µ veto threshold 1 GeV**

No NVeto, 1 visible TPC neutron : 40 NVeto thresh 500 keV, 1 visible TPC neut : 7 NVeto thresh 500 keV, precisely 1 TPC neut: 6 NVeto thresh 100 keV, 1 visible TPC neut : 1 NVeto thresh 0 keV, 1 visible TPC neut : 0 NVeto thresh 0 keV, precisely 1 TPC neut : 0

μ veto threshold 100 MeV

No NVeto, 1 visible TPC neutron : 12 NVeto thresh 500 keV, 1 visible TPC neut : 4 NVeto thresh 500 keV, precisely 1 TPC neut: 3 NVeto thresh 100 keV, 1 visible TPC neut : 0 NVeto thresh 0 keV, 1 visible TPC neut : 0 NVeto thresh 0 keV, precisely 1 TPC neut : 0

μ veto threshold 50 MeV

No NVeto, 1 visible TPC neutron : 5 NVeto thresh 500 keV, 1 visible TPC neut : 1 NVeto thresh 500 keV, precisely 1 TPC neut : 1 NVeto thresh 100 keV, 1 visible TPC neut : 0 NVeto thresh 0 keV, 1 visible TPC neut : 0 NVeto thresh 0 keV, precisely 1 TPC neut : 0 This is ~ the nominal design, modulo fiducial+multiscatter cuts

Note: DS-50 saw 2 cosmo neutrons (vetoed) DS-20k has a cross sectional area 136x larger

Neglecting vetoes, a simple scaling of background rate would predict 272 events → these simulations predict 234!

One event that's hard to kill...

- Edep TPC = 158 keV
- Edep NVeto = 459 keV
- Edep LAr Bath = 616 keV
- n_{neutrons} = 1, at 3.3 MeV

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Can we achieve this?

- Zoe has been running g4ds simulations of Plan C geometry with PMTs added
 - Studying PE detected vs. energy deposited by muons for different configurations
- One realistic scenario: Reuse MiniCLEAN PMTs
 - Of course, not the only option (maybe we want PDMs?), but it is ~free and sets a realistically achievable baseline for argument's sake
 - 92 PMTs (80 are working perfect, they think remaining 12 are just a problem with the base and can be revived)
 - For these simulations, Zoe is using the DS-50 LSV PMTs (R5912, 8" diameter), spaced ~uniformly around the cryostat walls
- Disclaimers:
 - Muon simulations are not currently producing anything heavier than electrons. Debugging from Igor indicates that this may be fixed by upgrading to latest Geant4-10 version
 - Muon veto background model from ³⁹Ar and γ's are still being set. Realistically, this will set the achievable threshold

Optical muon veto simulations

Three configurations considered

- 92 PMTs, no reflector, no TPB
- 92 PMTs, Tyvek reflector, no TPB
- 92 PMTs, Tyvek reflector, TPB on PMT faces

Scintillation+Cherenkov modes
 ³⁹Ar pileup + γ backgrounds, but significantly enhanced μ signal

No scintillation: Cherenkov only modes
 Negligible ³⁹Ar background, mostly γ's



Cherenkov-only modes



⁽Note the different x axes)

Scintillation + Cherenkov mode



From scaling arguments we naively expect to find a LY ~ 40–400 pe/MeV

Dead time

Full simulations are being developed. For now, the following are some preliminary considerations

- ³⁹Ar: 700 tonnes \rightarrow 700 kBq!
 - Negligible contribution to Chernekov (expect something like 10 PE/³⁹Ar decay at endpoint)
 - If we need a 5 (10) μs coincidence window to detect scintillation light, then we expect an average of 3.5 (7)
 ³⁹Ar decays to pile up
 - With a mean β energy of 219 keV, ³⁹Ar pileup will produce a wall below ~800 (1600) keV for Scintillation configuration
- γ-rays from cryostat (from Vicente's spreadsheet): Average 1.9 MeV at 77 kBq



- Assume half of these γ 's scatter in LAr, then we have 38.5 kBq background of 1.9 MeV
- This gives a 20 (40)% pile-up rate below 2 MeV, assuming 100% of γ energy lost in LAr
- γ-rays from rock/lab surroundings: Need reference

Dead time: ³⁹Ar toy MC and lab γ 's

10 μ s window

For the scintillation case, we will have <10% dead time from ³⁹Ar pileup above ~2.5 MeV

Lacking information of ambient lab radioactivity...

- DS-50 LSV saw ~300 kHz trigger rate prior to filling WCV.
- Assuming γ's have same energy as cryostat gammas, expect ~1.5 (3) pileup events per 5 (10) µs window, summing to 2-5 MeV total.

Conclusions

- More FLUKA statistics are needed to pin down final background expectation, along with better models of multiple scatter and fiducial cuts
- However, from the available simulations, the (unrealistically) best case scenario leaves us with two cosmogenic neutron backgrounds in 10 years we need a muon veto
- It looks like achieving an expectation << 1 will require a muon veto threshold as low as ~50 MeV and decreasing neutron veto threshold below 500 keV (a 100 keV threshold works)
- Optical simulations of the muon veto are still under development, but it looks like these goals are imminently achievable with modest instrumentation
 - A 2 MeV threshold on muon tagging seems achievable given backgrounds
- Bonus note: Potentially some interesting neutrino physics through the neutrino absorption channel (ER signal at E_v-1.5 MeV with a taggable delayed coincidence and a σ) with 700 tonnes of instrumented LAr...

