29 September 2021 to 1 October 2021 Gran Sasso National Laboratory (LNGS)

# Survey on other next generation Double Beta experiments



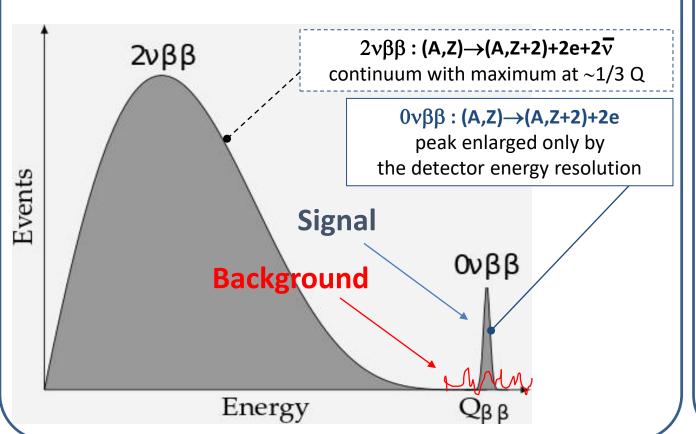
Andrea Giuliani
IJCLab, Orsay, France

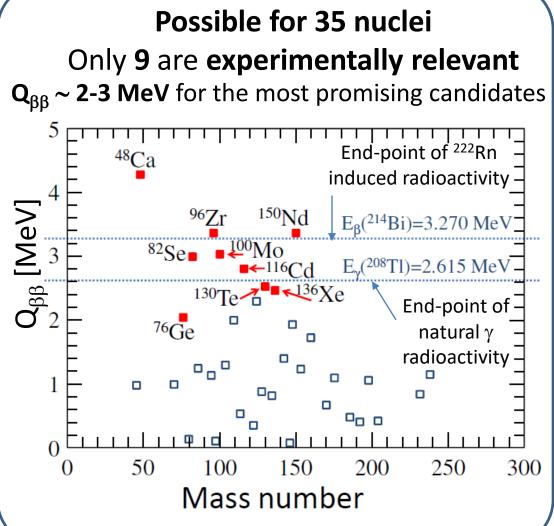


# Which signature and which nuclei?

### Sum energy spectrum of the two electrons

 $\mathbf{Q}_{\beta\beta}$ : energy available for the products

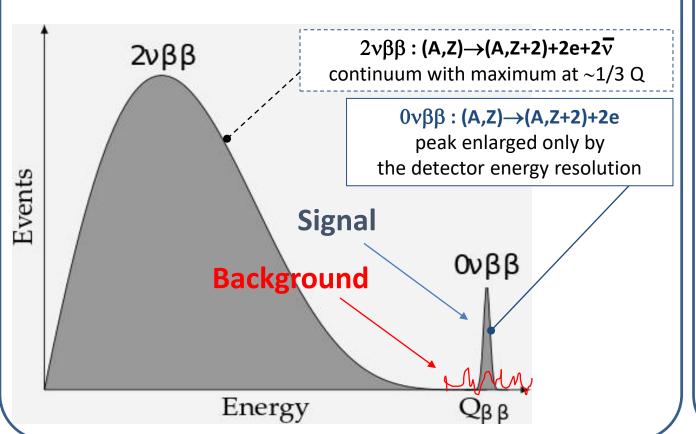


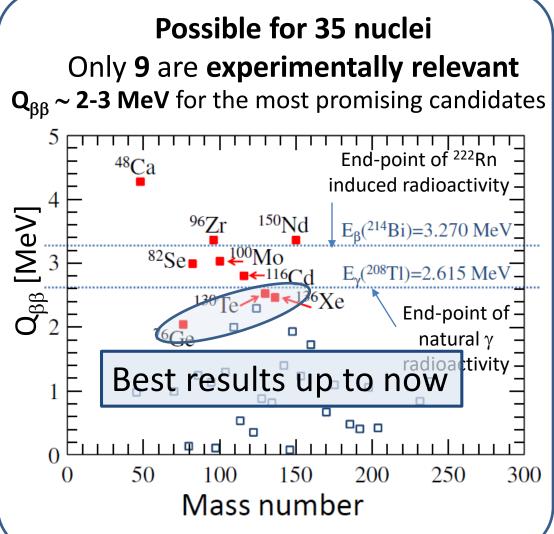


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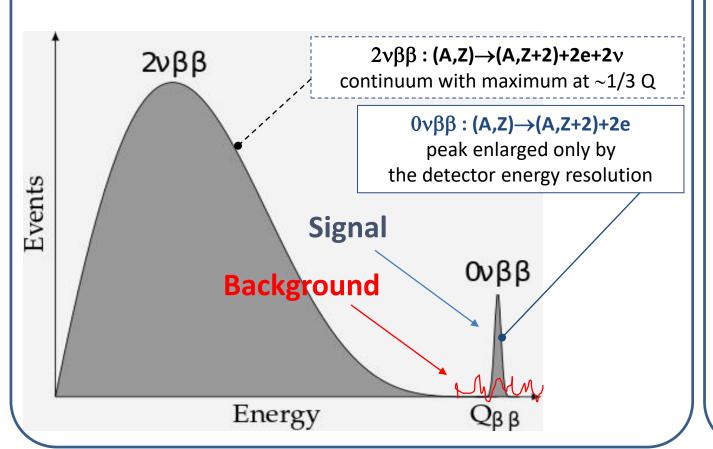




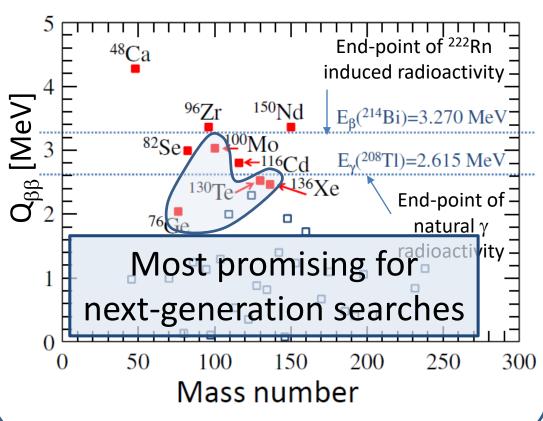
# Which signature and which nuclei?

### Sum energy spectrum of the two electrons

 $\mathbf{Q}_{\mathbf{B}\mathbf{B}}$ : energy available for the products

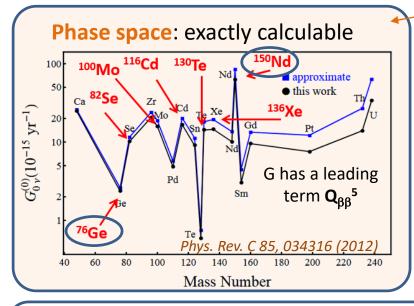


# Possible for 35 nuclei Only 9 are experimentally relevant



# Which half-lives?

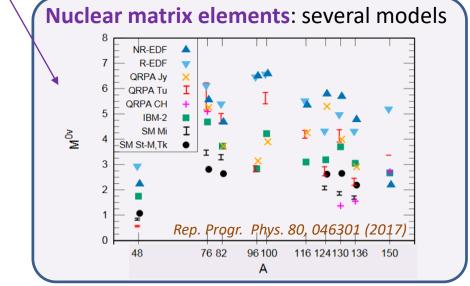
Mass mechanism  $1/\tau = G(Q_{\beta\beta}, Z) g_A^4 |M_{nucl}|^2 m_{\beta\beta}^2$ 



 $g_A = \begin{bmatrix} 1.269 \text{ Free nucleon} \\ 1 & \text{Quark} \end{bmatrix}$ 

 $g_{A,eff} \sim 0.6 - 0.8$  to be taken (« quenching ») to describe  $\beta$  and  $2\nu\beta\beta$  rates with current nuclear models

- Controversial
- Ab-initio calculation with unquenched g<sub>A</sub> are required
- Progress ongoing

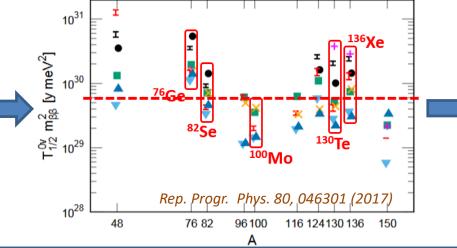


### 0νββ rate

The  $0\nu\beta\beta$  community still assumes  $\mathbf{g}_{\mathbf{A}}\approx$  **1.27** (no quenching) with «traditional models» for  $\mathbf{M}_{\text{nucl}}$ 

This point should be revised in the future, after an expected maturation of ab-initio calculations

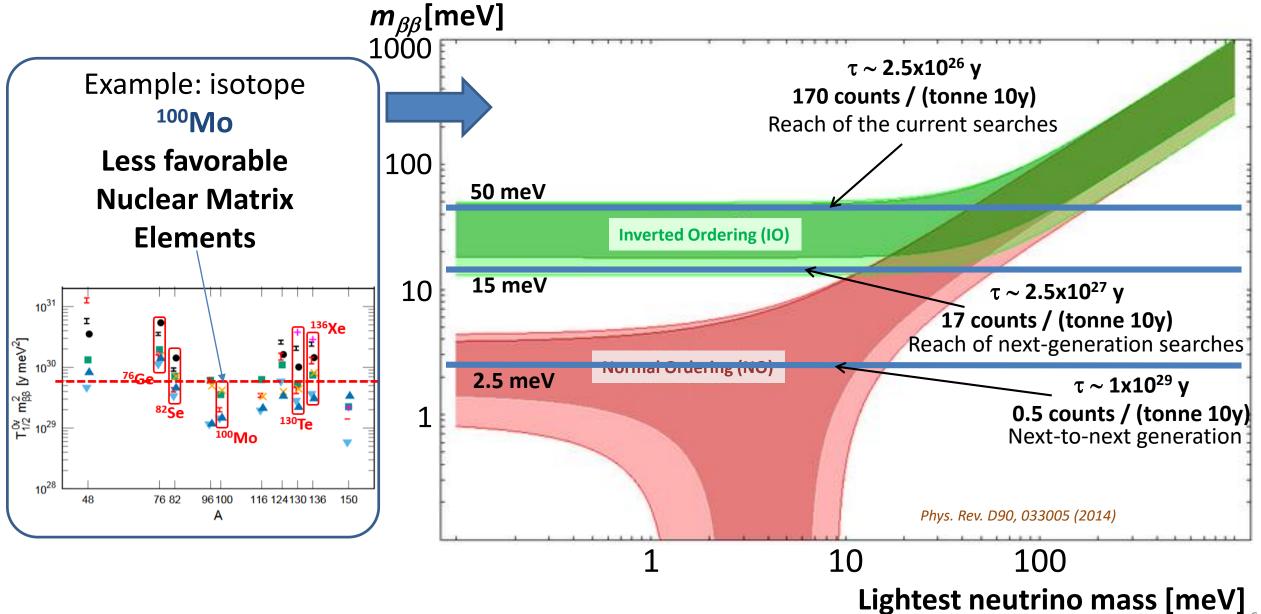
arXiv:2108.11805



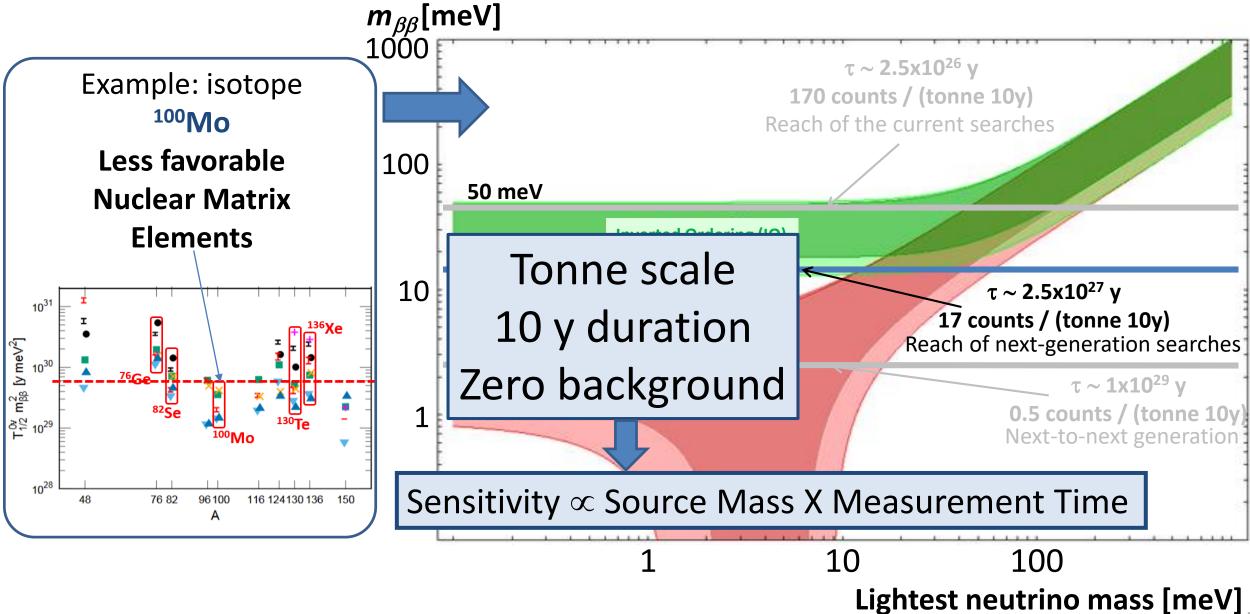
 $T_{1/2}^{0\nu} \simeq 10^{27-28} \left( \frac{0.01 \,\text{eV}}{\langle m_{\beta\beta} \rangle} \right)^2 \text{y}$ 

Working formula for general experiment design

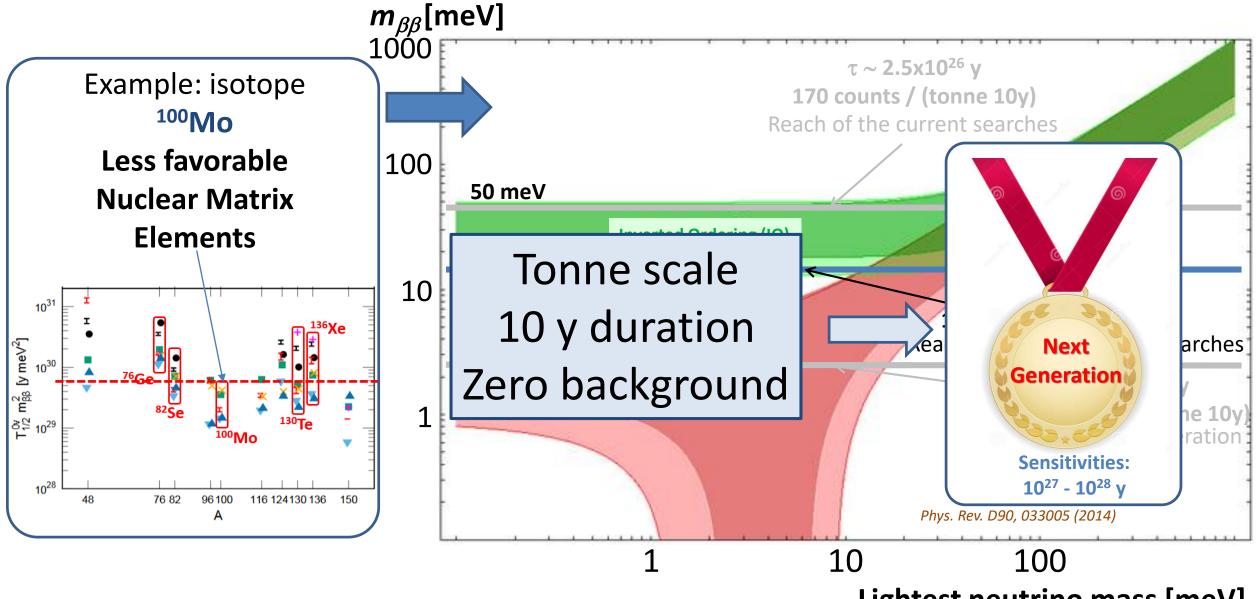
# How many counts?



# How many counts?



# How many counts?



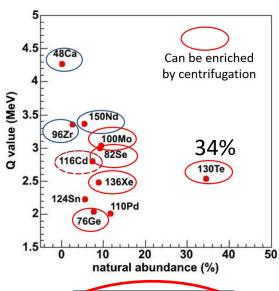
# Expand the source, abate the background

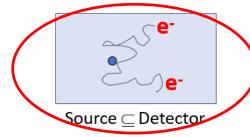
### **Source**

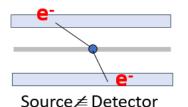
- 1 Large source  $\rightarrow$  tonne scale  $\rightarrow$  > 10<sup>27</sup> nuclei
- 2 Isotopic enrichment
- → the isotopic abundance is artificially increased to > 80%
  - $\rightarrow$  Isotope selection

## 3 Maximize efficiency

→ The option in which the source is separated from the detector is abandoned for next-generation experiments







### **Background**

- → Standard common actions
- **1** Natural radioactivity ( $\alpha$ ,  $\beta$ ,  $\gamma$  radiation)

Levels  $< 1 \mu Bq/kg$  are required  $\longleftrightarrow$  Ordinary material  $\sim 1-100 Bq/kg$ 

2 Cosmic muons

**Underground laboratory**  $\rightarrow$  Muon flux reduction by > 10<sup>6</sup>

3 Neutrons

Generated by rock radioactivity and muons

Quality and depth of the underground laboratory

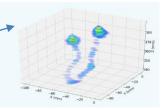
Dedicated shieldings are often required

**Dedicated shieldings** are often required

4 Cosmogenic induced activity (long living)
Delayed effect of the cosmic radiation (activation)
Choice of detector materials – Storage of material underground

### → Specific actions depending on the technology

- High energy resolution
- Particle identification
- Tracking / Event topology
- Multi-site vs. single-site events
- Surface vs. bulk events
- Fiducial volume / Active shielding
- Final-state nucleus identification



background index b

background counts @Q<sub>ββ</sub>

 $M \times \Delta E \times T$ 

# Deployment of an arsenal of technologies

# Two main classes of experiment

The source is a fluid or is diluted in a fluid

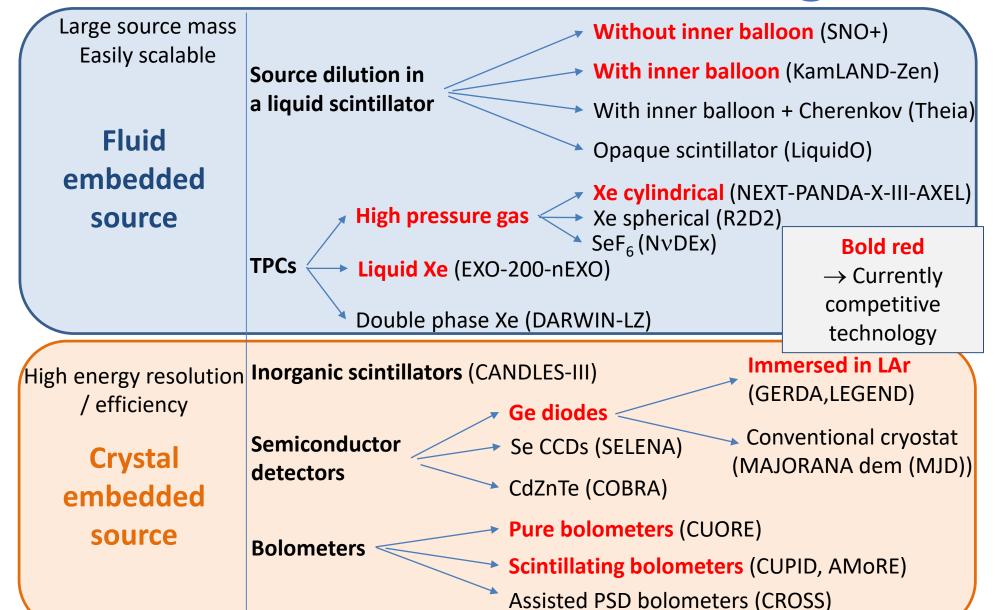
Scalability – Increase:

- Vessel size
- Source concentration

The source is contained in a crystal

Scalability – Increase:

- Crystal size (marginal)
- Numbers of crystals



# Implementation in tens of experiments

Legenda (color code)



### **Completed**

Data taking

Construction / Commissioning

Advanced R&D sometimes at CDR/TDR level

R&D

TGV-2 EC/EC  $\beta$ +/EC  $^{106}$ Cd

Large source mass Easily scalable

Fluid embedded source

NvDEx	High pressure TPC	<sup>82</sup> Se
ZICOS	Dilution in liquid scintillator+Cherenkov	<sup>96</sup> Zr
SNO+	Dilution in liquid scintillator	<sup>130</sup> Te
SNO+-Phase II	Dilution in liquid scintillator	<sup>130</sup> Te
Theia	Dilution in liquid scintillator+Cherenkov	<sup>130</sup> Te- <sup>136</sup> Xe
KamLAND-Zen 400	Dilution in liquid scintillator	<sup>136</sup> Xe
KamLAND-Zen 800	Dilution in liquid scintillator	<sup>136</sup> Xe
KamLAND2-Zen 800	Dilution in liquid scintillator	<sup>136</sup> Xe
EXO-200	Liquid TPC	<sup>136</sup> Xe
nEXO	Liquid TPC	<sup>136</sup> Xe
NEXT-White	High pressure TPC	<sup>136</sup> Xe
NEXT-100	High pressure TPC	<sup>136</sup> Xe
NEXT-HD / NEXT-BOLD	High pressure TPC	<sup>136</sup> Xe
PANDAX-III	High pressure TPC	<sup>136</sup> Xe
AXEL	High pressure TPC	<sup>136</sup> Xe
DARWIN	Double-phase TPC	<sup>136</sup> Xe
LZ	Double-phase TPC	<sup>136</sup> Xe
R2D2	High pressure TPC	<sup>136</sup> Xe

High energy resolution
/ efficiency

Crystal embedded source

CANDLES-III CANDLES-IV Scintillating bolometers MAJORANA DEM. Semiconductor detectors GERDA LEGEND-200 Semiconductor detectors LEGEND-1000 Semiconductor detectors  CDEX-300 / CDEX-1000 Semiconductor detectors  SELENA Semiconductor detectors  SELENA Semiconductor detectors  CUPID-0 Scintillating bolometers CUPID-MO Scintillating bolometers AMORE-I Scintillating bolometers CUPID CUPID Scintillating bolometers  CUPID Scintillating bolometers  CUPID CUPID Scintillating bolometers  100 Mo CUPID COBRA Semiconductor detectors  TIN-TIN Bolometers  CUORE Bolometers  Scintillating bolometers  100 Mo-130Te CROSS Scintillating bolometers  100 Mo-130Te COBORO CIPID Scintillating bolometers  CROSS Scintillating bolometers  100 Mo-130Te COBORO	LiquidO	Dilution in opaque liquid scintillator	multi
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TIN-TIN Bolometers 124Sn CUORE Bolometers 130Te CROSS Scintillating bolometers 100Mo-130Te	CUPID Reach / CUPID-1T	Scintillating bolometers	<sup>100</sup> Mo
CUORE Bolometers 130Te CROSS Scintillating bolometers 100Mo-130Te	COBRA	Semiconductor detectors	<sup>116</sup> Cd
CROSS Scintillating bolometers 100 Mo-130 Te	TIN-TIN	Bolometers	<sup>124</sup> Sn
	CUORE	Bolometers	<sup>130</sup> Te
DINCO Scintillating / Charankov balamatars 100Ma 130Ta	CROSS	Scintillating bolometers	<sup>100</sup> Mo- <sup>130</sup> Te
BINGO SCITULIALING / CHERENKOV DOTOMETERS 100 MIO-100 TE	BINGO	Scintillating / Cherenkov bolometers	<sup>100</sup> Mo- <sup>130</sup> Te

# **Current situation**

$$T_{1/2} > 10^{24}$$
 y 90% C.I. restricted club

**GERDA** T<sub>1/2</sub> > 1.8×10<sup>26</sup> y

Phys. Rev. Lett. 125, 252502 (2020)

 $T_{1/2} > 1.07 \times 10^{26} \text{ y}$ KamLAND-Zen 400 Phys. Rev. Lett. 117, 082503 (2016)

EXO-200  $T_{1/2} > 3.5 \times 10^{25} \text{ y}$ Phys. Rev. Lett. 123, 161802 (2019)

 $T_{1/2} > 2.7 \times 10^{25} \text{ y}$ MAJORANA dem. Phys. Rev. C 100, 025501

 $T_{1/2} > 2.2 \times 10^{25} \text{ y}$ **CUORE** *arXiv:1907.09376* 

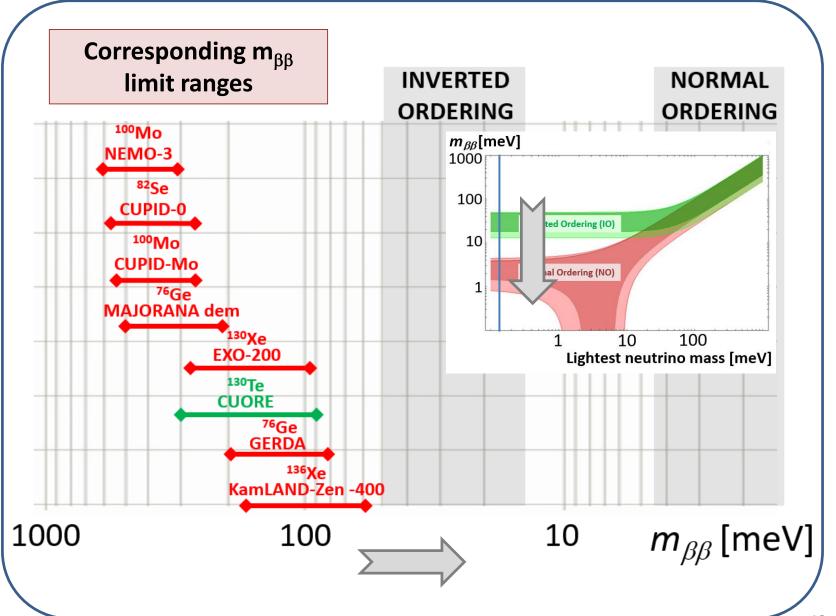
 $T_{1/2} > 4.7 \times 10^{24} \text{ y}$ **CUPID-0** 

L. Pagnanini, TAUP 2021

 $T_{1/2} > 1.8 \times 10^{24} \text{ y}$ **CUPID-Mo** B. Welliver, TAUP 2021

T<sub>1/2</sub> > 1.1×10<sup>24</sup> y

Phys. Rev. D 92, 072011 (2015)



# Most promising next-generation experiments

7 research lines / experiments are more mature:

- 4 fluid embedded
- 3 crystal embedded

In my view, these experiments are in the best position for actual construction and data taking on a-few-year time scale, for several reasons:

- Technology maturity
- Solid collaboration
- Funding prospects

Large source mass Easily scalable

Fluid embedded source

- $\bigcirc$  **NEXT-White**  $\rightarrow$  **NEXT-100**  $\rightarrow$  NEXT-HD / NEXT-BOLD

Data taking
Construction /
Commissioning

**Completed** 

Advanced R&D

High energy resolution
/ efficiency

Crystal embedded source

- GERDA MAJORANA dem. → LEGEND-200 → LEGEND-1000
- CUPID-Mo
  CUPID-O
  CUPID → CUPID Reach / CUPID 1t
  CUORE
- **7** AMORE-I → AMORE-II

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**Fluid** embedded

- KamLAND-Zen 400 → KamLAND-Zen 800 → KamLAND2-Zen
- $EXO-200 \rightarrow nEXO$
- NEXT-White ightarrow NEXT-HD / NEXT-BOLD

Completed

**Data taking** Construction / Commissioning

Advanced R&D

R&D

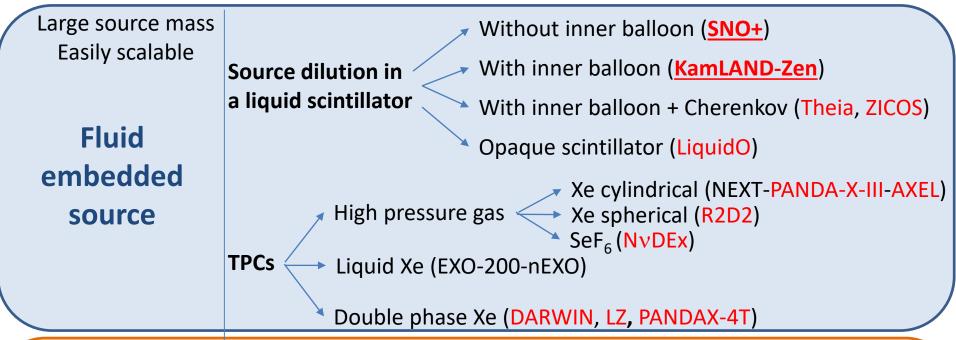
Crystal source

→ **CUPID** → CUPID Reach / CUPID 1t

**AMORE-I** → **AMORE-II** 

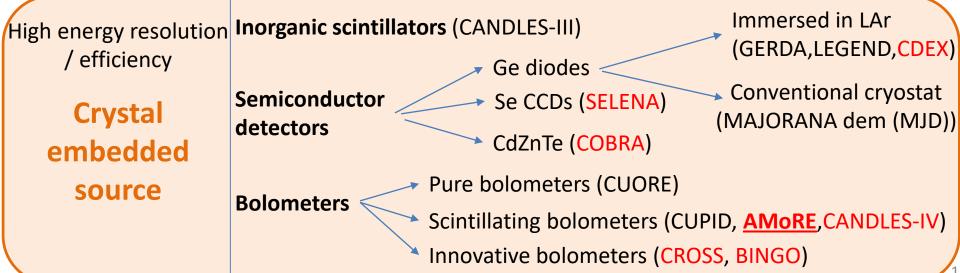
# Other promising developments

**SuperNEMO** (Tracking+Calorimetry) Source ≠ Detector



/ efficiency

Crystal embedded source



# KamLAND-Zen (400 and 800)

KamLAND-Zen 400 → KamLAND-Zen 800 → KamLAND2-Zen

KamLAND-Zen 400 – Kamioka, Japan  $T_{1/2} > 1.07 \times 10^{26}$  y 350 kg of <sup>136</sup>Xe – Leading experiment  $m_{\beta\beta} < 60 - 160$  meV

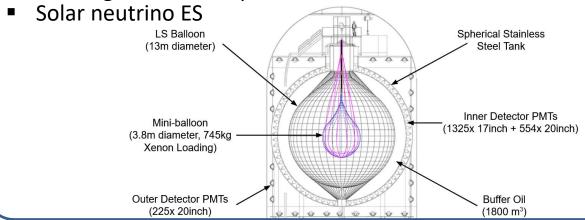
### Concept

Enriched Xenon diluted (3 wt%) in liquid scintillator exploiting the existing KamLAND detector with the addition of a nylon balloon

- Scalability increase diameter of nylon inner balloon (IB)
- <sup>136</sup>Xe On-off
- Energy resolution:  $\Delta E(\sigma) \sim 7\%/VE(MeV) 4.5\%@Q_{BB}$
- Single event position Vertex resolution 15 cm/ vE(MeV)

### **Background:**

- $2\nu\beta\beta$  decay of <sup>136</sup>Xe
- Xe-LS, IB and outer-LS radioactive impuritities
- Cosmogenic: muon-spallation



KamLAND-800 (started Jan 2019)

### Major new points with respect to KamL-400

- More isotope 745 kg of <sup>136</sup>Xe
- New balloon (2X larger, more radiopure)
- Reduction of <sup>12</sup>C-spallation by analysis
- Characterization of <sup>136</sup>Xe spallation
- Improve KamL-400 results by ~4X in 5 y

$$\rightarrow$$
 m <sub>$\beta\beta$</sub>  < 30 – 80 meV  
J. Phys.: Conf. Ser. 1468 012142 (2020)

# KamLAND2-Zen

Data taking Advanced R&D

Completed

KamLAND-Zen 400  $\rightarrow$  KamLAND-Zen 800  $\rightarrow$  KamLAND2-Zen

3 neutrinos and beyond, K. Ueshima, 2019

### KamLAND2-Zen

- Larger source 5X brighter  $\rightarrow$  2X better  $\Delta$ E
- $m_{\beta\beta}$  < 20 meV

### **Directions of improvement**

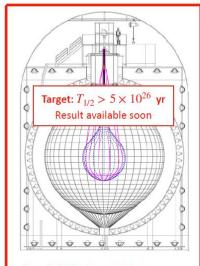
- $> 1000 \text{ kg of } ^{136}\text{Xe}$
- Reduce <sup>214</sup>Bi background
- Reduce 2v2β background



### 5y sensitivity

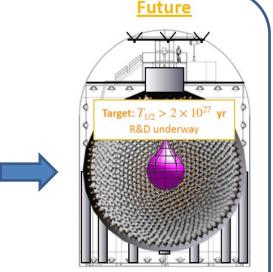
 $T_{1/2} > 2 \times 10^{27} \text{ y}$  $m_{\beta\beta}$  < 14 – 37 meV

### Present



### KamLAND-Zen 800:

- · Mini-balloon Radius = 1.90 m
- Xenon mass = 745 kg
- Data taking starts Jan. 2019



### KamLAND2-Zen:

- Aiming at 100% Photocoverage
- · PEN scintillation balloon film

### Reduce <sup>214</sup>Bi background



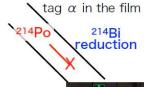
Identify BiPo events in balloon



Tag  $\alpha$ 's with a film on balloon



scintillator film



Reduce  $2v2\beta$  background



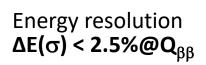
Improve energy resolution



Increase light yield



- Brighther scintillator (x 1.4)
- High Q.E. 20" PMT (22% $\rightarrow$ 30%) (x 1.9)
- Winstone cone for light collection (1.8)



Ultimate background: elastic scattering of solar v's on e

Further evolutions



Possibility to include scintillating inorganic crystals embedding other  $2\beta$  candidates Ambitious long-term developments: Super-KamLAND-Zen – a few tens of tons of Xe in a 20 kton detector

# SNO+

### SNO+ → SNO+-phase II

### Concept

*arXiv:2104.11687v2* 

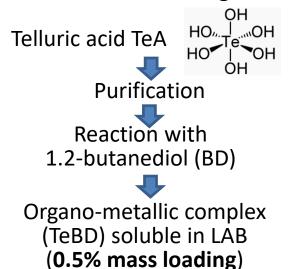
Reuse the acrylic vessel, the PMT array and the electronics of the SNO detector at SNOLAB with a new target:

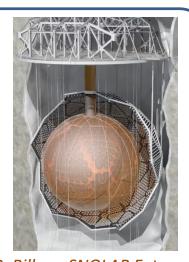
natural-Te-loaded liquid scintillator (LAB + 2g/l PPO "fluor")

- 780 tons of scintillator
- 3.9 tons of natural tellurium
- $\rightarrow$  **1.3** tons of <sup>130</sup>Te (34% I.A.)

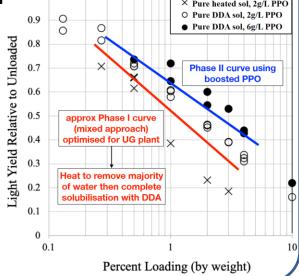
→ Scintillator purification system

→ Novel metal loading technique 1





S.B. Biller – SNOLAB Future Project Workshop – May 2021

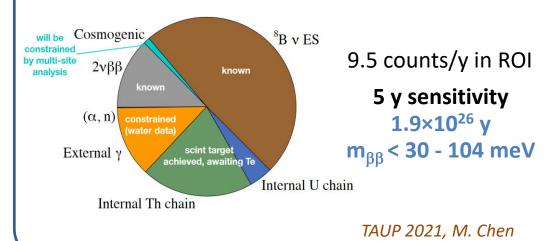


### **SNO+** consists of **three phases**

- Pure-water phase (from May 2017)
  - → measurement of the external background
  - $\rightarrow$  physics results (8B v's, invisible nucleon decays)
- Liquid scintillator phase without Te (ongoing)
  - → measurement of scintillator background
  - $\rightarrow$  U, Th concentration  $\sim 5 \times 10^{-17}$  g/g
  - $\rightarrow$  Background level low enough for  $0\nu\beta\beta$
- **Te phase** (from 2022) Study of  $2\nu\beta\beta$  and  $0\nu\beta\beta$

 $\Delta E = 190 \text{ keV FWHM } @Q_{\beta\beta}$ 

**Background budget and sensitivity** 



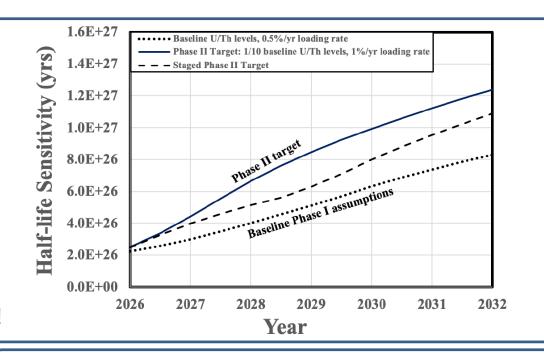
# **SNO+** and evolutions

**SNO+**  $\rightarrow$  **SNO+-phase II**  $\rightarrow$  THEIA, ZICOS

### **SNO+-phase II** (start in 2026)

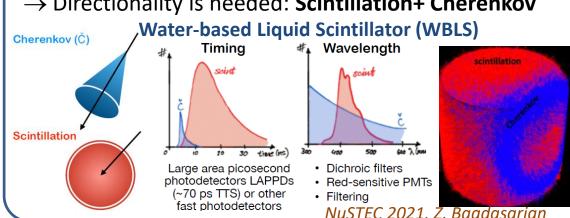
### 5 y sensitivity

- $1 \times 10^{27} \text{ y}$  $0.5\% \rightarrow$  **3%** Te concentration
- $m_{\beta\beta} < 13 45 \text{ meV}$ Improve transparency
  - No changes to underground set-up
  - No changes to loading method
  - Increase PPO to ~6g/L
  - Adiabatically increase Te loading towards ~ 3%
  - Loading rate depends on staging and system throughput (project 0.5% per 6-9 months)
- 3% Te concentration  $\rightarrow$  24 t of Te / 8 t of <sup>130</sup>Te
- → 70% "thrown away" by fiducial cuts but still cost effective!



### Scaling up towards NO is limited by solar v's background

→ Directionality is needed: **Scintillation+ Cherenkov** 



### THEIA – multipurpose neutrino detector

**DUNE** cavern

KAMIOKA

- 50 kt WBLS 16 m radius balloon with isotope
- 5% natural Te loading 10 y sensitivity: 1.1×10<sup>28</sup> y
- Another option: 3% enriched <sup>136</sup>Xe loading

TAUP 2021, Y. Fukuda

### ZICOS – **Zr-loaded liquid scintillator detector**

- 180 t total 3.5 m radius inner balloon with isotope
- 10 wt.% Zr(iprac)4 loaded in liquid scintillator 50% <sup>96</sup>Zr
- 1.7 t of loaded LS  $\rightarrow$  865 kg of  $^{96}$ Zr
- Cherenkov+Scintillation demonstrated in protoypes
  - $\rightarrow$  topological background rejection Sensitivity: 1×10<sup>27</sup> y

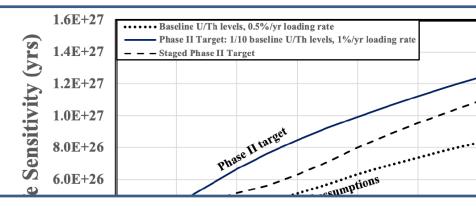
# **SNO+** and evolutions

 $SNO+ \rightarrow SNO+-phase II \rightarrow THEIA, ZICOS$ 



- $0.5\% \rightarrow 3\%$  Te concentration  $1\times10^{27}$  y
- Improve transparency  $m_{\beta\beta}$  < 13 45 meV No changes to underground set-up No changes to loading method Increase PPO to ~6g/L

Adiabatically increase Te loading towards ~ 3%



### LiquidO

R&D activity based on the well-known liquid scintillator technology

Radical change of paradigm: Opaque scintillator + tight array of fibres

conventional paradigm of transparency is abandoned

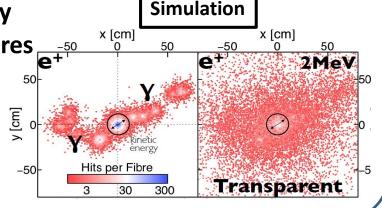
scintillation light is confined and collected near its creation point

Target loading fractions: 5-30 % (vs. current 1-3 %)

Transparency constraint is relaxed

~10 ton isotope (130Te, 82Se, 100Mo, 150Nd)

Prototype at Bordeaux

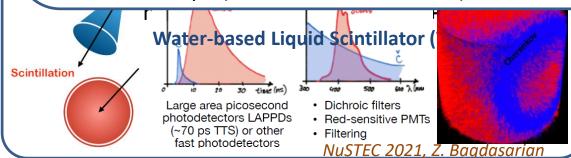


r detector

ZICOS - Zr-loaded liquid scintillator detector

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- Cherenkov+Scintillation demonstrated in protoypes

  → topological background rejection Sensitivity: 1×10<sup>27</sup> y

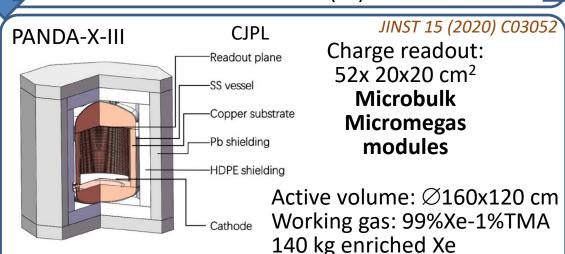


# Other gas TPC experiments

### High pressure (10-15 bar) enriched Xe gas TPC

NEXT technology, with variants, is adopted by

- PANDA-X-III electron collection
- AXEL electroluminescence (EL)



 $\Delta E = 3\%$  FWHM  $@Q_{\beta\beta}$  – Signal efficency = 35% Background index ~10<sup>-4</sup> c/(keV kg yr) after topological cuts

3 y sensitivity:  $8.5 \times 10^{25}$  y  $m_{BB} < 68 - 180$  meV

**Future extension:** 

PANDA-X-III 1t Ton scale detector Several modules Special readout for EL 25
Electroluminescence
Light Collection Cell
(ELCC)

10mm-pitch

Drift anode electrode

PTFE w/ holes

Mesh electrode

MPPC (pholon sensor) array

X-axis

2.5kV/cm/bar ~0.1kV/cm/bar

Proof of principle in 10L prototype Extrapolated energy resolution:  $\Delta E = 0.82-1.74 \%$  FWHM  $@Q_{\beta\beta}$  180L set-up under development

### NvDEX High pressure 82SeF<sub>6</sub> TPC

**CJPL** 

- Same gas used to enrich Se  $\rightarrow$  in principle extendable to Te, Mo
- SeF<sub>6</sub> does not admit free electrons → Ion TPC
- Collect and read-out ions without multiplication with CMOS technology
- Topology and good energy resolution
- 100 kg set-up under development

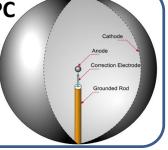
### R2D2 Spherical high pressure (40 bar) Xe gas TPC

Simple readout and mechanics Light readout Primary scintillation

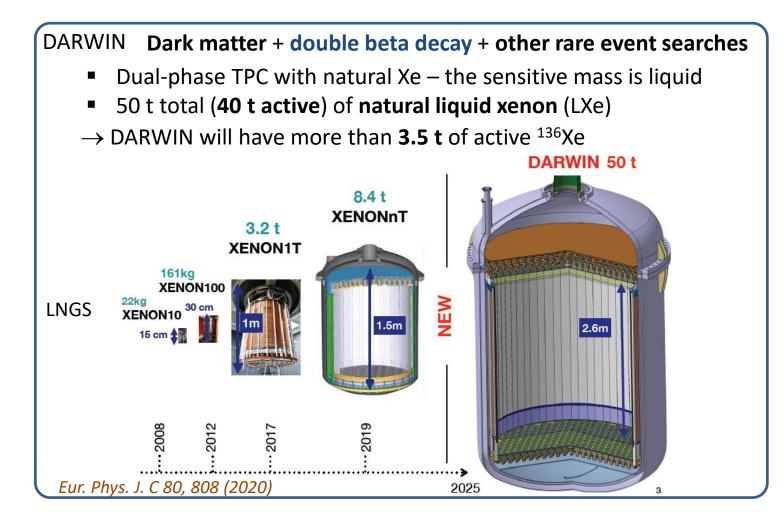
+ EL during avalanche

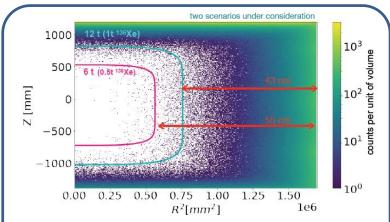
 $\emptyset$ 40 cm prototype under test

Prove zero background with  $\emptyset$ 74 cm  $\rightarrow$  50 kg Xe  $\emptyset$ 40 bar



# **Dual-phase Xe TPC experiments**





### Main background sources

- <sup>222</sup>Rn in LXe
- <sup>137</sup>Xe from μ-induced neutrons
- 8B Solar neutrinos

Fiducial volume: 5 t

Background: 0.2 c/(ton y) in ROI

10 y sensitivity:  $T_{1/2} > 2.4 \times 10^{27}$  y  $m_{\beta\beta} < 11 - 35$  meV

Similar calculations from PANDAX-4T (3.7 t) and LZ (10 t) with sensitivities two orders of magnitude lower

Ke Han, TAUP 2021 arXiv:2104.13374

# **AMORE**

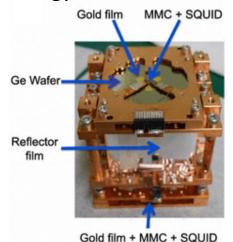
JINST 15 C08010 (2020) J. Phys.: Conf. Ser. 1468, 012130 (2020)

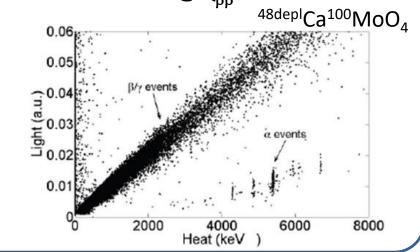
### $AMORE-I \rightarrow AMORE-II$

### AMoRE – Y2L Lab (AMORE-I), Yemilab (AMORE-II), Korea Concept

- <sup>100</sup>**Mo**-containing scintillating bolometers
- Initially chosen compound (AMoRE pilot R&D): 48deplCa<sup>100</sup>MoO<sub>4</sub>
  - $\rightarrow$  high light yield, excellent  $\alpha/\beta$  separation by PSD and light yield
  - → challenging internal contamination (<sup>238</sup>U chain)
- <sup>48depl</sup>Ca<sup>100</sup>MoO<sub>4</sub> has been accompanied by Li<sub>2</sub><sup>100</sup>MoO<sub>4</sub> in AMORE-I
- Li<sub>2</sub><sup>100</sup>MoO<sub>4</sub> is currently the only compound foreseen in **AMORE-II**
- Heat readout based on MMC sensors (faster than CUORE/CUPID)
  - $\rightarrow 2\nu\beta\beta$  random coincidences provide negligible background

Energy resolution  $\Delta E \sim 10-15$  keV FWHM @Q<sub>68</sub>

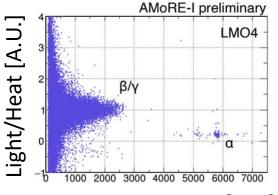




AMORE-I – started in Aug 2020 - stop in 2022

 $13x^{48\text{depl}}\text{Ca}^{100}\text{MoO}_{4}$  (CMO, 4.6 kg) 5x Li<sub>2</sub><sup>100</sup>MoO<sub>4</sub> (LMO, 1.6 kg)

3 kg of <sup>100</sup>Mo





Target BI:  $< 10^{-2}$  counts/(keV kg y)

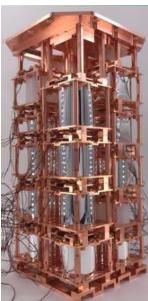
Projected sensitivity:  $7 \times 10^{24}$  y  $m_{\beta\beta} < 130 - 250$  meV

**AMORE-II** - 2022 - 2027

Secured **110 kg of**  $^{100}$ Mo – 596x Li<sub>2</sub> $^{100}$ MoO<sub>4</sub> crystals New cryostat and underground lab – work in progress

Target BI:  $< 10^{-4}$  counts/(keV kg y)

Projected sensitivity:  $8 \times 10^{26}$  y  $m_{\beta\beta} < 13 - 25$  meV



# Other bolometric experiments

### CANDLES

CANDLES-III Pure scintillation experiment with CaF<sub>2</sub>(Eu) crystals

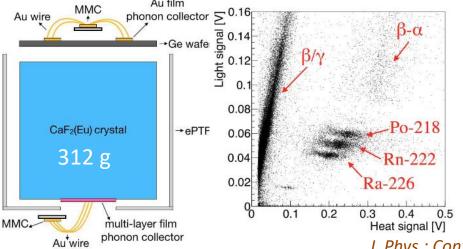
Natural crystals – 96x 3.2 kg  $\rightarrow$  only **350 g of** <sup>48</sup>Ca

**KAMIOKA** 

New phase of the experiment  $\rightarrow$  CANDLES-IV

- Study the possibility of enrichment with Laser Isotopic Separation
- Move to scintillating bolometers (as CUPID, AMORE)
  - high energy resolution
  - $\alpha/\beta$  rejection

Preliminary encouraging results with large crystals (MMC technology)

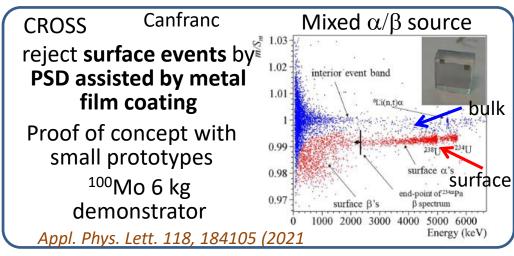


Modest energy resolution  $\sigma = 3.2\% @ 4.9 \text{ MeV}$ 

Position dependence

J. Phys.: Conf. Ser. 1468 (2020) 012116

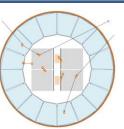
Techniques for background rejection in future TeO<sub>2</sub> / Li<sub>2</sub>MoO<sub>4</sub> based experiments



BINGO Modane Luminescent bolometers

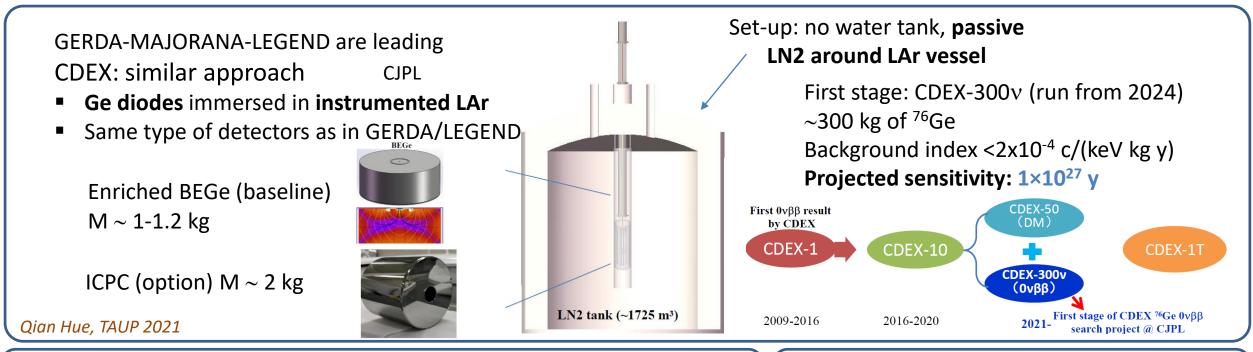
- Internal active shield (ultrapure ZnWO<sub>4</sub> scintillators)
   → mitigate γ background in TeO<sub>2</sub>
- Revolutionary assembly to reject surface background
- Enhanced-sensitivity light detectors

C, Nones, TAUP 2021





# Other semiconductor-based experiments



SELENA Innovative approach

Amorphous 82Se x-ray detectors (0.2 mm thick) +HVReadout: CMOS pixel array

Stack to achieve high density, high mass array

- 5 μm pixel size gives full track reconstruction
- Industrial production
- Very promising for background control ~0.001 c/(FWHM t y)

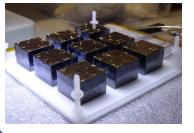
JINST 12 (2017) P03022

# Integrated electronics CMOS pixel array 100 cm² x 5 µm E field Amorphous 82Se (200 µm thick) (30 nm)

### COBRA CdZnTe detector technology

Several  $\beta\beta$  isotopes – <sup>116</sup>Cd most promising

- Demonstrator: 64x 1 cm³ detectors
- Recent upgrade: 9x 2x2x1.5 cm<sup>3</sup> detectors



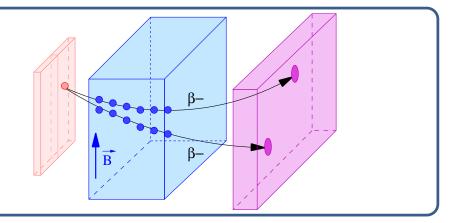
Improvements in energy resolution and background

NIMA 1010 (2021) 165524

# SuperNEMO

### **Tracker-Calorimeter Technique**

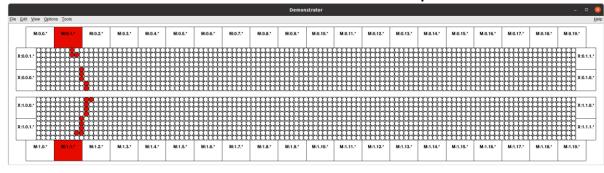
- Foils can be made of any solid  $\beta\beta$  isotope (SuperNEMO uses <sup>82</sup>Se)
- Identification of  $e^-$ ,  $e^+$ ,  $\gamma$  and  $\alpha \rightarrow$  leads to excellent background rejection.
- Event topology reconstruction (energies, angles).
- $e-\gamma$  separation can probe decays to excited states



### **SuperNEMO demonstrator status**

Modane

- Final commissioning.
- First tracker-calorimeter data September 2021



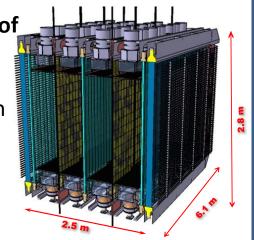
First  $\beta\beta$ candidate event,
9/9/21

# Current role of SuperNEMO Providing Supporting Measurements

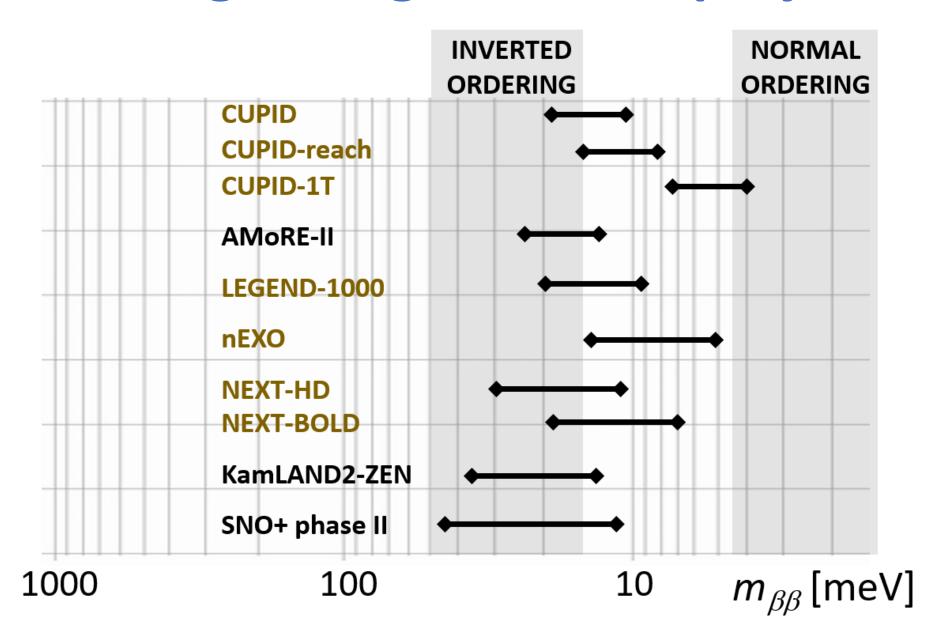
- Precision measurements of 2vββ
  - g<sub>A</sub> quenching constraints (NEMO-3 analysis in preparation)

Understanding the Ultimate Reach of the Tracker-Calorimeter Technique

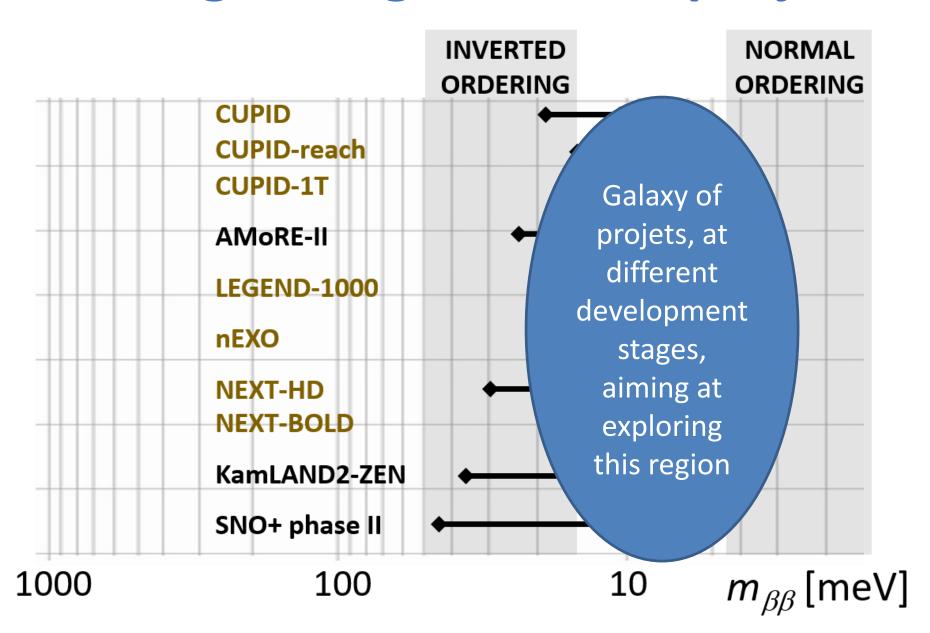
- Can the technique be used to confirm & probe a signal found in the next generation of 0vββ experiments?
- Explore different detector technologies & isotopes



# Promising next-generation projects



# Promising next-generation projects



# BACK UP

# Neutrinoless double beta decay in a nutshell

$$0\nu\beta\beta: (A,Z) \rightarrow (A,Z+2) + 2e^{-}$$

$$2n \rightarrow 2p + 2e^{-}$$

Creation of matter without antimatter partners

**Beyond Standard Model** 

Never observed – Best limits  $\tau > 10^{24}$  -  $10^{26}$  y

- ① Standard mechanism: neutrino physics  $0v2\beta$  is mediated by light massive Majorana neutrinos (exactly those which oscillate) Sometimes defined "mass mechanism"
- 2 Non-standard mechanisms: Sterile v, LNV Not necessarily neutrino physics

The only currently viable experimental approach to probe the Majorana nature of neutrino

Francesco Vissani, this workshop

### Experiments measure / constrain $\tau$ Assuming mass mechanism, this translates into information on the effective Majorana mass m<sub>BB</sub> $m_{\beta\beta}$ [meV] 1000 $\mathbf{m}_{\beta\beta} = |U_{e1}|^2 m_1 + e^{i\alpha 1} |U_{e2}|^2 m_2 + e^{i\alpha 2} |U_{e3}|^2 m_3$ $_{100}$ 1/ $\tau \propto m_{BR}^2$ Inverted Ordering (IO) 10 Normal Ordering (NO)

Phys. Rev. D90, 033005 (2014)

100

Lightest neutrino mass [meV]

# Neutrinoless double beta decay in a nutshell

$$0\nu\beta\beta: (A,Z) \rightarrow (A,Z+2) + 2e^{-}$$

$$2n \rightarrow 2p + 2e^{-}$$

Creation of matter without antimatter partners

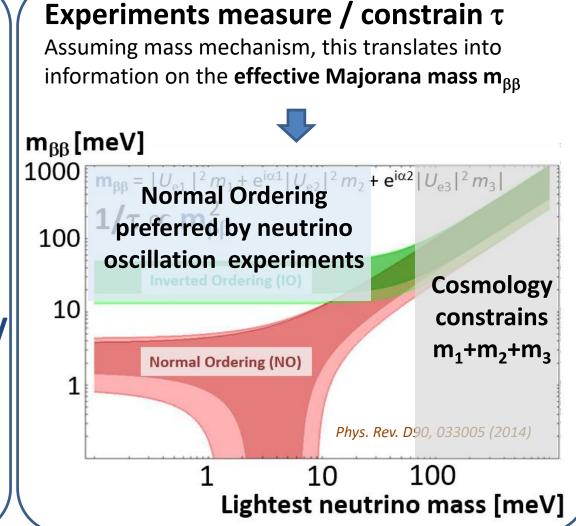
**Beyond Standard Model** 

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The only currently viable experimental approach to probe the Majorana nature of neutrino

Francesco Vissani, this workshop



# **nEXO**

### $EXO-200 \rightarrow nEXO$

nEXO is built on the successful **EXO-200** – **WIPP, US** 

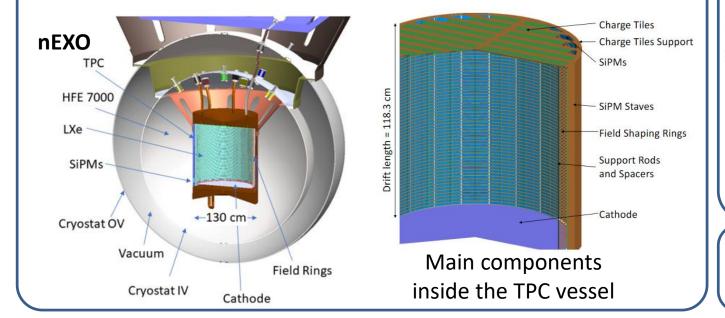
150 kg of 
$$^{136}$$
Xe  $-T_{1/2} > 3.5 \times 10^{25}$  y  $-m_{\beta\beta} < 93 - 286$  meV

First observation of  $2\nu\beta\beta$  of <sup>136</sup>Xe (2011) –  $T_{1/2}$  = 2.165×10<sup>21</sup> y

### Concept

### Single phase enriched LXe TPC

- Energy resolution  $\Delta E(\sigma) \sim 0.8\%@Q_{\beta\beta}$
- Measurement of both charge and scintillation
- Single site (including signal) vs. multi site events (background)
- Multi-dimensional analysis using energy, 3D position and topology



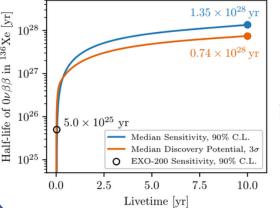
# **nEXO** (under DoE Portfolio Review) – **SNOLab Major upgrades with respect to EXO-200**

- More isotope  $\sim 5000 \text{ kg of }^{136}\text{Xe}$
- Improvement in light sensors (LAAPDs→SiPM)
- Increased light collection
- Improvement in radiopurity (electroformed Cu)
- Cold electronics

	EXO-200	nEXO
Fiducial Mass [kg]	74.7	3281
Energy resolution σ/Q <sub>ββ</sub> [%]	1.2%	0.8%

LXe self shielding

preCDR - arXiv:1805.11142v2 arXiv:2106.16243



Background dominated by Rn outgassing and intrinsic radioactivity

Equivalent background index: 7×10<sup>-5</sup> c/keV kg y)

### 10 y sensitivity

$$1.35 \times 10^{28} \text{ y}$$
  $m_{\beta\beta} < 5 - 15 \text{ meV}$ 

Tagging of individual <sup>136</sup>Ba daughter

Demonstrated by

 $^{136}$ Xe  $\rightarrow$   $^{136}$ Ba + 2e<sup>-</sup>

fluorescence in solid Xenon Nature 569, 203–207 (2019)

# **GERDA** → **LEGEND**

GERDA
MAJORANA dem.

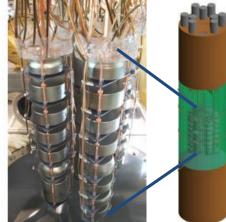
 $\rightarrow$  LEGEND-200  $\rightarrow$  LEGEND-1000

GERDA - LNGS, Italy  $T_{1/2} > 1.8 \times 10^{26} \text{ y} - m_{\beta\beta} < 79 - 180 \text{ meV}$ 35 kg of <sup>76</sup>Ge – Leading experiment in terms of half-life

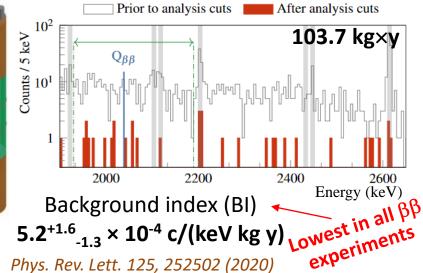
### Concept

### High purity naked Ge detectors immersed in instrumented LAr

- Energy resolution  $\Delta E \sim 3 \text{ keV FWHM } @Q_{BB}$
- Pulse shape discrimination: multi site vs. single site events
- Anticoincidence with LAr active shield, instrumented with
  - Wavelength shifting fiber shroud coupled to SiPMs
  - PMTs on top and bottom of the setup



37 HP Ge detectors



### **LEGEND-200** combines the best of GERDA and MJD

- Adopt GERDA detector configuration
- Reuse GERDA infrastructure at LNGS (after upgrade)
- Follow MJD selection of radiopure parts
- MJD electronics and low threshold
- <sup>76</sup>Ge: 35 kg from GERDA, 30 kg from MJD
   140 kg are new material
- New detector type, already tested in GERDA ICPC detector, > 2 kg vs. previous 0.7-0.9 kg
- → same energy resolution and PSD capability **Commissioning:**

Detector deployment starts in Sep 2021 Data taking: end 2021 / beginning 2022

AIP Conference Proceedings 1894, 020027 (2017)

### LEGEND-1000 (under DoE Portfolio Review) Discovery

- Same technology, new larger infrastructure
- Phased approach, up to 1000 kg of <sup>76</sup>Ge Background
- Site to be decided baseline: SNOLAB free approach

LEGEND-200	LEGEND-1000
BI: 2× 10 <sup>-4</sup> c/(keV kg y)	BI: 10 <sup>-5</sup> c/(keV kg y)
$T_{1/2} > 10^{27} y - 5 y live time$	$T_{1/2} > 1.3 \times 10^{28} - 10 \text{ y live time}$
$m_{\beta\beta}$ < 34 – 78 meV	$m_{\beta\beta}$ < 9 – 21 meV

*arXiv:2107.11462v1* 

sensitivity

# CUORE $\rightarrow$ CUPID

**CUPID-Mo CUPID-0** 

→ CUPID → CUPID Reach / CUPID-1T

**CUORE** 

CUORE - LNGS, Italy  $T_{1/2} > 2.2 \times 10^{25} \text{ y} - m_{\beta\beta} < 90 - 305 \text{ meV}$ 

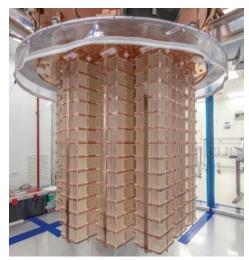
Exposure: 1038.4 kg  $\times$  y – Record for a bolometric experiment

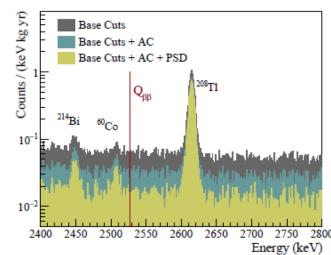
### Concept

*arXiv:1907.09376* 

Array of natural TeO<sub>2</sub> crystals operated as bolometers at 10 mK

- Built on the precursor CUORICINO experiment
- 988 TeO<sub>2</sub> crystals, arranged in 19 towers 206 kg of <sup>130</sup>Te
- Energy resolution  $\Delta E \sim 7.8$  keV FWHM  $@Q_{\beta\beta} Q_{\beta\beta} = 2527$  keV Background index **1.49** × **10**<sup>-2</sup> **c/(keV·kg·y)**
- - Dominated by energy-degraded surface  $\alpha$ 's





Target half-life sensitivity: 9×10<sup>25</sup> y

CUPID-Mo - LSM, France

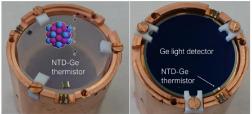
Exposure:  $2.71 \text{ kg} \times \text{y}$ 

 $T_{1/2} > 1.8 \times 10^{24} \text{ y}$  $m_{\beta\beta} < 280 - 490 \text{ meV}$ 

Concept ← LUMINEU R&D

**2 changes** wrt CUORE:

(1) Pure bolometers  $\rightarrow$ **Scintillating bolometers**  Phys. Rev. Lett. 126, 181802 (2021)



 $(2)^{130}$ Te (TeO<sub>2</sub>)  $\rightarrow$   $^{100}$ Mo (enriched Li<sub>2</sub>MoO<sub>4</sub>)  $Q_{BB}$ =3034 keV > 2.6 MeV (reject external  $\gamma$  background)

20  $\text{Li}_2\text{MoO}_4$  crystals – 2.26 kg of <sup>100</sup>Mo

Energy resolution  $\Delta E \sim 7.8 \text{ keV FWHM } @Q_{BB}$ Reject  $2\nu\beta\beta$  decay of <sup>100</sup>Mo Counts/keV background background free after light cut Heat Energy [keV]

CUPID-0 - LNGS, Italy Zn82Se First scintillating

bolometer demonstrator

🌊. Pagnanini, this conference  $T_{1/2} > 4.7 \times 10^{24} \text{ y}$ 

Small scale demonstrators

# CUORE → CUPID

CUPID-Mo CUPID-0 CUORE

→ **CUPID** → CUPID Reach / CUPID-1T

**CUPID** (under DoE Portfolio Review) – **LNGS, Italy** 

### **Concept**

- Single module: Li<sub>2</sub><sup>100</sup>MoO<sub>4</sub>
   45×45x45 mm − ~ 280 g
- 57 towers of 14 floors with2 crystals each 1596 crystals
- ~240 kg of <sup>100</sup>Mo with >95% enrichment
- ~1.6×10<sup>27</sup> 100 Mo atoms
- Bolometric Ge light detectors as in CUPID-Mo, CUPID-0

Copper str

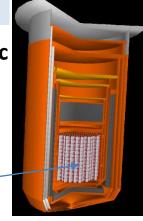
arXiv:1907.09376

### CUPID is built on successful CUPID-Mo + CUORE

 $\text{Li}_{2}\text{MoO}_{4}$  scintillating bolometer technology, with demonstration of energy resolution, crystal radiopurity and  $\alpha$  rejection

Ton-scale bolometric experiment is possible Electronics and data analysis tools Reuse CUORE

infrastructure



### **CUPID** sensitivity

### Data driven background model

- Information from CUPID-Mo, CUPID-0
- CUORE background model (same infrastructure!)

Projected background index: 1×10<sup>-4</sup> c/(keV kg y)

Critical background component: random coincidence of  $2\nu\beta\beta$  events (100 Mo fastest  $2\nu\beta\beta$  emitter:  $T_{1/2} = 7.1 \times 10^{18}$  y)

10 y discovery sensitivity  $m_{\beta\beta} < 12 - 20 \text{ meV}$ 

### Possible follow-up of CUPID

CUPID-reach - Same sensitive mass and cryostat as CUPID Background improvement by factor 5 Criticalities:

 $2.3 \times 10^{27} \text{ y} \rightarrow \text{m}_{ee} < 7.9 - 14 \text{ meV}$ 

- 2νββ

CUPID-1T - 1 ton isotope → new cryostat Background improvement by factor 20

Surface events

 $9.2 \times 10^{27} \text{ y} \rightarrow \text{m}_{ee} < 4.0 - 6.9 \text{ meV}$ 

Intense R&D to improve background in Li<sub>2</sub>MoO<sub>4</sub> and TeO<sub>2</sub> based bolometric experiments

CROSS → reject surface events by PSD assisted by metal film coating

BINGO →

- Internal active shield (ZnWO<sub>4</sub> scintillators)
- Enhanced-sensitivity light detectors

Revolutionary detector assembly