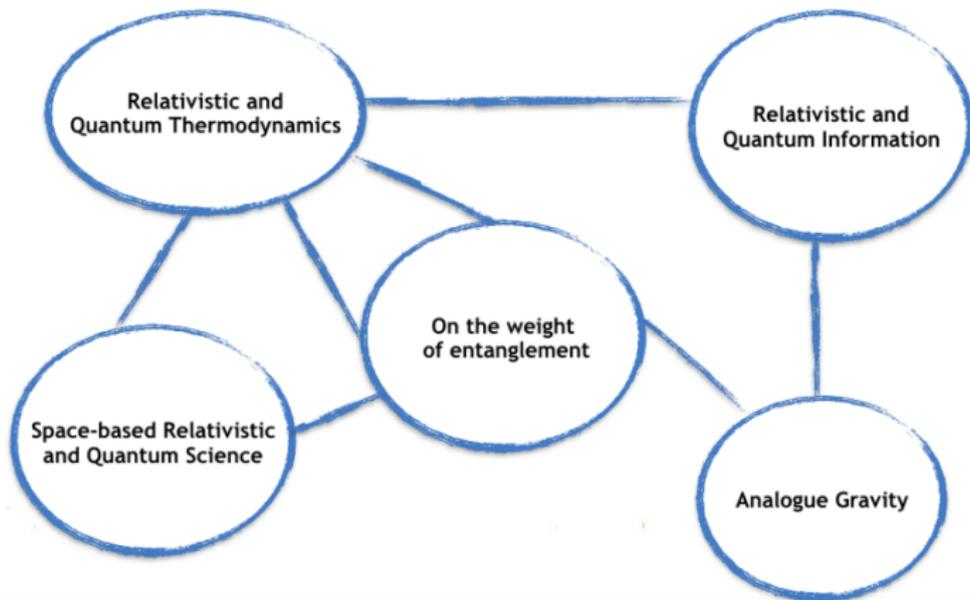


Self gravity affects quantum states

David Edward Bruschi

Institute for Quantum Computing Analytics (PGI-12)
Forschungszentrum Jülich
Germany

VIII.IX.MMXXI



Einstein gravity

We know that **ALL ENERGY GRAVITATES.**

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Einstein equations

$$G_{\mu\nu} = \kappa T_{\mu\nu}.$$

These equations have been **highly successful** in providing many predictions.

Successes

- Precession of orbits;
- Bending of light;
- Black Holes;
- Penrose process;
- Gravitational waves;
- Cosmology.

Difficulties

- Rotation curves of galaxies;
- Nonlinearity of equations;
- Gravitation of quantum objects;
- Quantum nature of gravity;
- Fundamental or emergent theory?

Einstein gravity extended

We know that **ALL ENERGY GRAVITATES**.

Semiclassical gravity

$$G_{\mu\nu} = \kappa \langle : \hat{T}_{\mu\nu} : \rangle$$

$\hat{T}_{\mu\nu}$: **stress-energy tensor** for quantum field. $: \cdot :$ is **normal ordering**.

Successes

- Takes into account (somehow) backreaction;
- ...

Problems

- Fluctuations of stress energy tensor big/huge;
- Curved spacetime: inconsistent renormalisation procedures;
- “Strange” predictions for gravitational fields of superpositions;
- ...

Experiments at the overlap of relativity/quantum phys.

Planning

There are plans to try to test the gravitational field of small quantum objects that can be found in quantum states.

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Experiments of interest

- Spontaneous WF collapse;
- Gravitational decoherence;
- Superposition of masses;
- Optomechanical systems;
- Space based tests;
- Atom Interferometry;
- More?

One setup

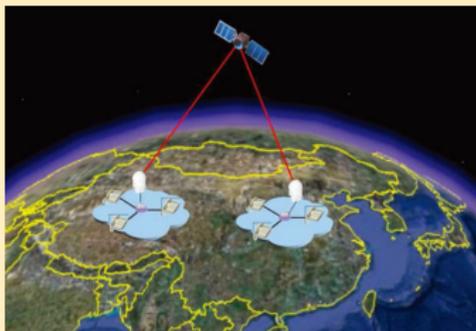


Figure: Micius satellite CAS

Quantum Information (QI) and Thermodynamics (QT)

Quantum Information

Allows us to connect concepts such as entropy and quantum correlations.

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QT investigates thermodynamics **far** from **thermodynamic limit**.

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- Small (quantum) constituents;
- Few (e.g **ONE**) systems;
- Concepts of energy and work not unique;
- Fluctuation relations;
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Applications

- Quantum chemistry;
- Quantum refrigerators;
- Fundamental physics;
- Connections to Information Theory;
- ...

Quantum Thermodynamics (QT)

Resources

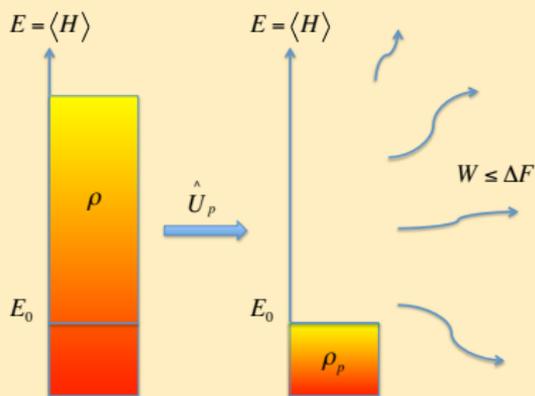
State $\hat{\rho}$. Unitaries \hat{U} . Then, exists \hat{U}_p : $\hat{\rho}_p = \hat{U}_p^\dagger \hat{\rho} \hat{U}_p$. And, $\hat{\rho}_p$ is “passive”.

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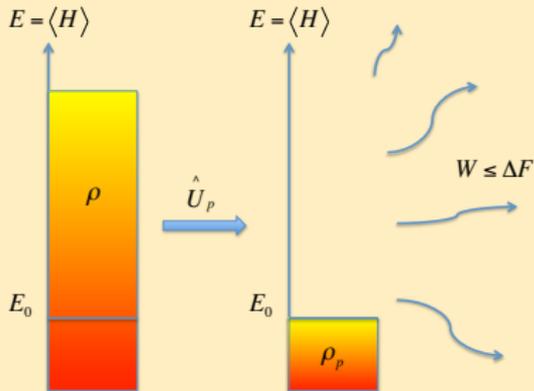


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Applications (PRE 91, 032118 (2015))

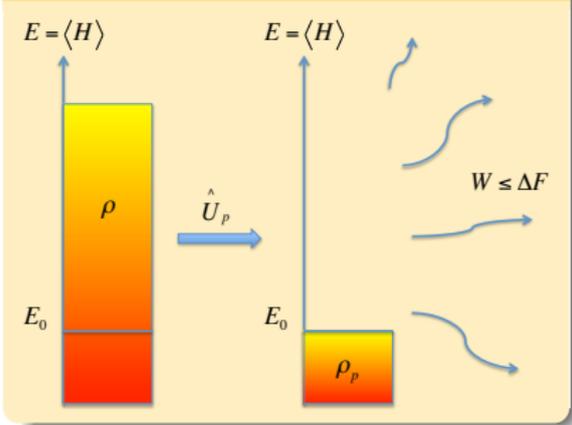
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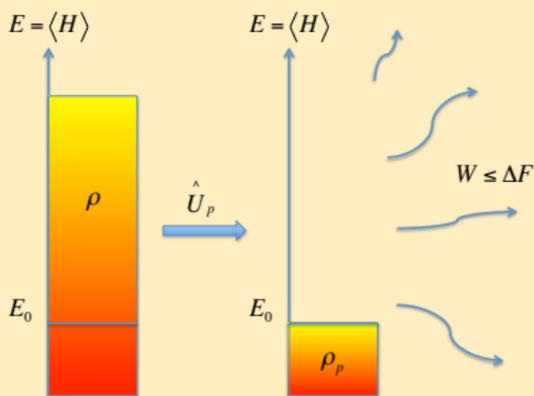
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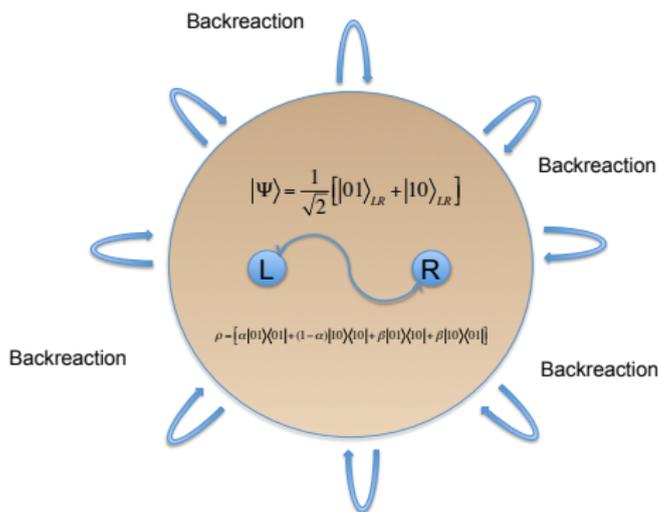
- Conclusion: Correlations must “carry energy”.

Role of correlations

On the weight of entanglement (Physics Letters B 54, 182-186 (2016))

Employing semiclassical gravity: "entanglement has a weight".

We find $G_{\beta}^{(1)} \propto \beta$, where $G^{(1)}$ is correction to flat Einstein tensor.



A gedankenexperiment

The weight of a passive state

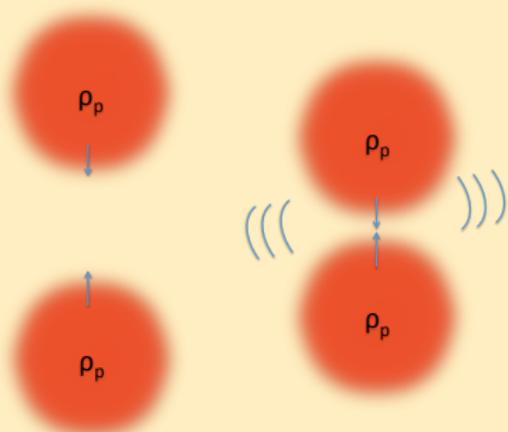
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Work from passive states

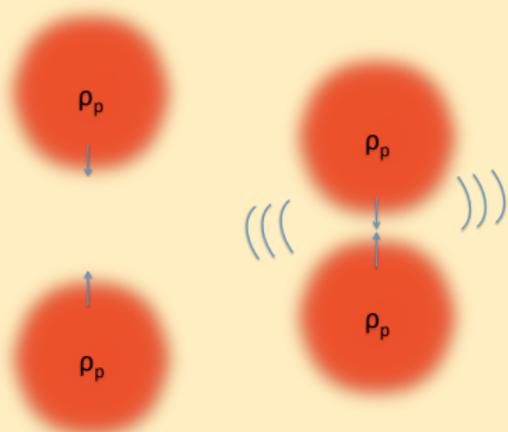


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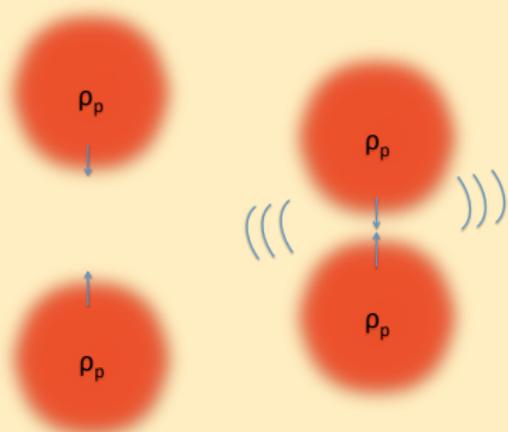
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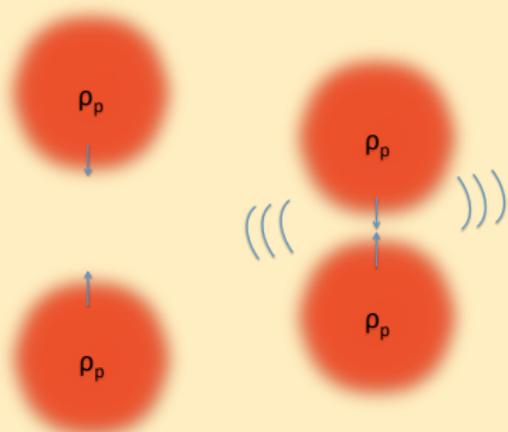
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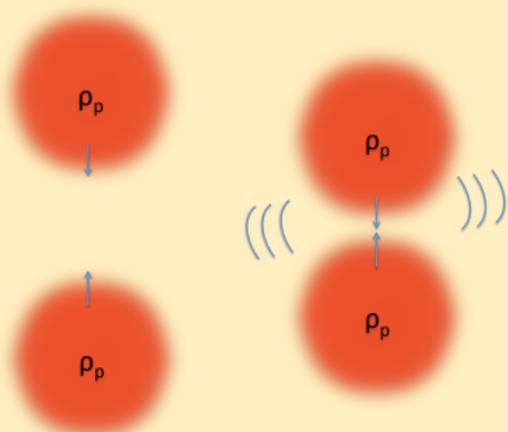
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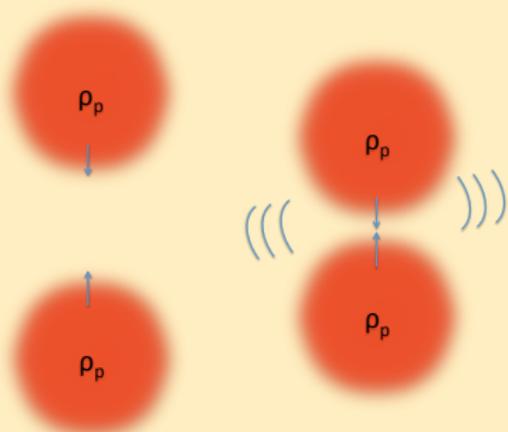
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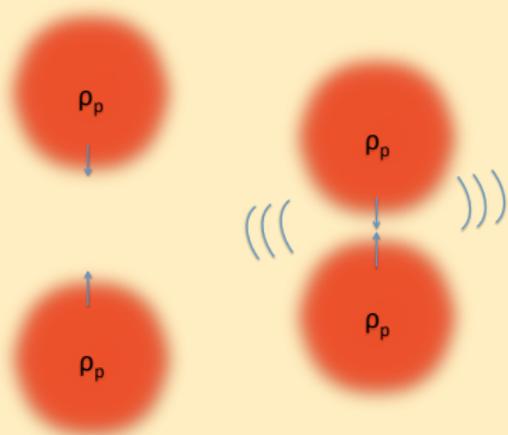
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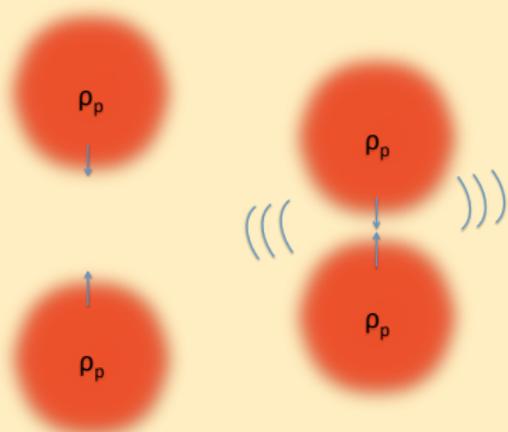
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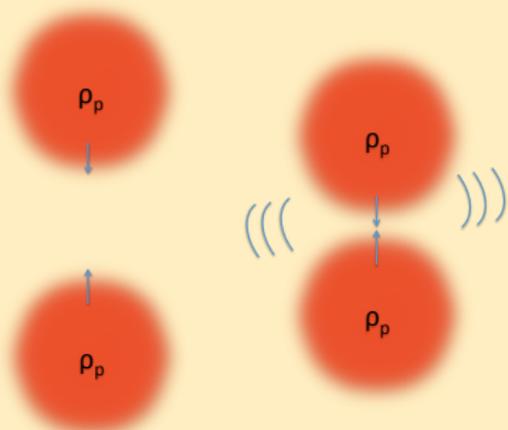
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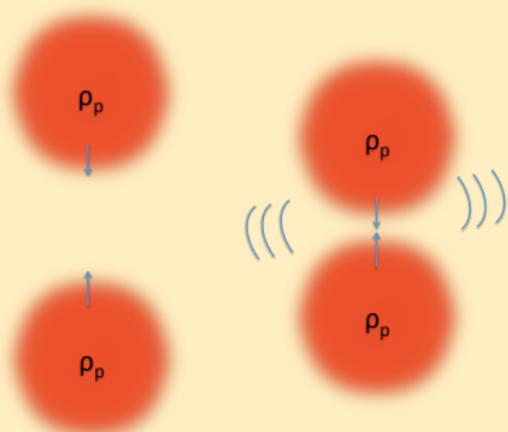
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A novel proposal (arXiv:1701.00699)

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Only **extractible energy** gravitates.

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New proposal

The field equations of Nature need to take into account the fact that it is extractible work that determines gravity, not the “whole amount of energy”.

However: this might require to change other fundamental equations (i.e., Heisenberg eq.).

Features of the proposal I

Clarification

Passive state $\hat{\rho}_p$ is vacuum state of **full theory** (including gravity part).

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N.B.

Classical states: little/vanishing zero point energy compared to total one.

Features of the proposal II

Considerations on known effects in QFT

- Unruh effect: theoretically requires infinite acceleration times (i.e., infinite fuel).
- Schwarzschild BH evaporation: evaporates forever as seen by somebody sitting at infinity (i.e., infinite energy).

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The proposal...

... predicts that these highly ideal scenarios should be “trivial” because the unitary transformation involved (“change of observer”) does **not** add energy to the observed system (which is in the vacuum state).

Time evolution by work (AoP 394, 155-161 (2018))

Time evolution

- Time evolution is typically “driven” by Hamiltonian.
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- $|\psi\rangle = \cos \theta |0\rangle + \sin \theta |1\rangle;$
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An experimental proposal

- Use highly entangled quantum state;
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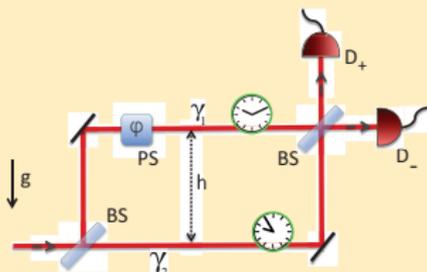


Figure: CQG 29, 224010 (2012)

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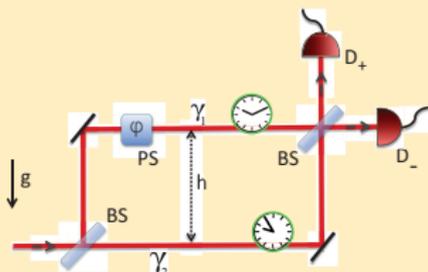


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Heisenberg evolution

- $|\psi\rangle = \frac{1}{\sqrt{2}} [|0N\rangle + |0N\rangle]$;
- $\dot{\rho} = \frac{i}{\hbar} [H, \rho] - \frac{i}{\hbar} [U_p^\dagger H U_p, \rho]$;
- $P|0N\rangle = \frac{1}{2} [1 + \frac{r_s}{r_E} \frac{h}{r_E} \sin^2(\frac{N\omega_0 t}{2})]$;
- $P|0N\rangle = \frac{1}{2} [1 - \frac{r_s}{r_E} \frac{h}{r_E} \sin^2(\frac{N\omega_0 t}{2})]$;
- Note: $\frac{r_s}{r_E} \frac{h}{r_E} \sim \frac{g h}{c^2}$;
- Note: $\frac{r_s}{r_E} \frac{h}{r_E} \sim h \times 10^{-16}$.

Conclusions

We have:

- Studied role of quantum correlations/entanglement in gravitating quantum systems;
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- Studied role of quantum correlations/entanglement in gravitating quantum systems;
- Proposed that only extractible energy gravitates;
- Modified time evolution to be compatible with new proposal

Then:

- Applications include localized scenarios, as well as global ones;
- Predictions are (very) different from the standard known theories;
- More advanced mathematical tools can be used to propose better models;
- Experiments can be potentially done with current technology;
- More work to come...

Acknowledgments

Grazie

PLB 54, 182-186 (2016) — arXiv:1701.00699 — AoP 394, 155-161 (2018)

