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Lightsaber: A simulator of the angular sensing and control system in LIGO

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- Importance and requirements
- Ongoing activities and motivation
- Optomechanical system



Importance and requirements

- LIGO limited by ASC noise
- No straight-forward solution
- ET issue needs to be addressed already with its design
- Low light power, increased mass of TM (reduce the angular optomechanical instabilities)
- Requirements:
 - 1. 1 nrad RMS
 - 2. Lowest possible noise above 10 Hz







Frequency [Hz]





Ongoing activities and motivation

- Time domain simulations of the ASC A2L coupling and strain noise
- Currently feedback is done by linear filter designed by commissioners
- > ML to improve our feedback control
- Test nonlinear feedback control such as reinforcement learning – high demands (robust, stable and optimal) – will be developed and tested with the time-domain simulations
- Caltech 40m prototype first experimental test
- DeepMind and Caltech







Overview

The linear couplings of the simulation are based on state-space models, which means that also the zpk specifications of control filters are internally converted into state-space models



Optomechanical system

- Radiation pressure exerts torque on the suspended mirrors, adding to the fixed restoring torque of the suspension
- > ASC allows us to operate IFO with angular mechanics dominated by RP
- Normal mode basis which decouples the effects of RP in 2 independent modes

Control feed

Sidles-Sigg effect



L_{ISI}[m]





Optomechanical system



Bode plot of Sidles-Sigg feedback transfer function changing the arm cavity power > The most significant coupling of angular motion to cavity length occurs when the beam spot is offcenter from the mirror's axis of rotation Results in an increase in the sensed longitudinal motion





Some results







8





Thank you for your attention

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BACKUP SLIDES

External noises and overall noise









Frequency [Hz]

Transfer functions









Bode plots and PSD of power fluctuations

