#### **Optical-parametric signal-amplification for a high-frequency gravitational-wave detector**

Department of physics, Tokyo Institute of Technology, Japan

Ken-ichi Harada

Collaborators

Department of physics, Tokyo Institute of Technology, Japan Sotatsu Otabe, Kaido Suzuki, Kentaro Somiya

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#### Outline

Introduction

**Optical spring Optical parametric amplification (OPA)** 

Experimental setup at Tokyo Tech.

Signal recycling Michelson interferometer with OPA

Result

Summary

## Introduction

The gravitational wave (GW) sources in a few kHz bands

binary neutron star merger, supernova, ....

The high-frequency signal cannot be detected by current GW detectors because of sensitivity degradation due to shot noise.

By improving the sensitivity in the kHz band, we significantly boost our understanding of the universe.



Time / ms

For improving the detection sensitivity,

- Squeezed vacuum: Reduce the noise of the sensitivity.
- Optical spring and Optical parametric amplification: Amplify the signal

# **Optical spring**

For improving the detection sensitivity in a high-frequency band,

Optical spring (OS) : Induced by the interaction between electromagnetic radiation and mechanical motion.



# **Optical parametric amplification**



#### Intracavity OPA

The optical spring becomes stiff.

The resonance frequency of OS is adjusted by detuned phase and pump power.

# **Sensitivity estimation**

Sensitivity of signal recycling Michelson interferometer (SRMI) with OPA



Laser sources



Fiber Amp.



#### Output power after the fiber amp.

λ/2: Half-wave plateEOM: Electro-optic modulatorFC: Fiber couplerPBS: Polarizing beam splitter



Michelson interferometer (MI)

Piezo1: Stabilize the MI system



λ/2: Half-wave plate
EOM: Electro-optic modulator
FC: Fiber coupler
PBS: Polarizing beam splitter
BS: Beam splitter
PD: Photodetector

#### Suspended mirror

Diameter: 6 mm Weight: 0.2 g Resonant frequency: 11 Hz Mount made of polyester



Signal recycling cavity (SRC)

Piezo1: Stabilize the MI system

Piezo2: Stabilize SRC by using the subcarrier 1 light



Subcarrier 1 ( $\omega + \omega_{sub1}$ )

The frequency-shifted by the double-pass configuration of AOM.

Carrier frequency

 $\omega_{sub1} = 2\omega_{AOM}$ 

 $\omega_{AOM} = 40 \text{ MHz}$ 

The frequency-shifted light is leaked from AS port because of the asymmetry of arm lengths of MI.

> EOM: Electro-optic modulator AOM: Acousto-optic modulator FC: Fiber coupler  $\lambda$ /2: Half-wave plate  $\lambda$ /4: Quarter-wave plate BS: Beam splitter PBS: Polarizing beam splitter PD: Photodetector SRM: Signal recycling mirror

Second harmonic generation (SHG)

Bow-tie cavity: Generate 532 nm light by SHG.

Stabilized by the Pound-Drever-Hall (PDH) method.





Observation of the signal amplification of the 1064 nm light



We confirmed the signal amplification by OPA.

#### **Subcarrier 2**

Subcarrier 2: Stabilize the relative phase between the carrier and pump. Employ the coherent control (CC) method reported in Phys. Rev. Lett. 97, 011101 (2006).



Subcarrier 2 ( $\omega + \omega_{sub2}$ )

The frequency is identical to subcarrier 1.

 $\omega_{sub2} = \omega_{sub1} = 2\omega_{AOM}$ 

The polarization is the identical to the carrier light.

The frequency-shifted light is leaked from AS port because of the asymmetry of arm lengths of MI.

The beat signal between  $\omega + \omega_{sub2}$ and  $\omega - \omega_{sub2}$  generates by OPA.

The error signal  $I_{err}$  can be obtained by demodulating  $2 \times \omega_{sub2}$ .

 $I_{\rm err} \propto \sin \theta \quad \theta$ : relative phase

## Subcarrier 2

#### Subcarrier 2 intensity detected by PD3





Time / s

13

#### Summary

Observed OPA in SRMI.

Obtained the error signal using subcarrier 2.

## **Future plans**

Improvement of the OPA signal Observation of the frequency shift of OS by OPA

Thank you for your attention!