

# TORSION-PENDULUM TESTING OF LISA CHARGE MANAGEMENT WITH A REPLICA LPF GRAVITATIONAL REFERENCE SENSOR

SECOND EPS CONFERENCE ON GRAVITATION

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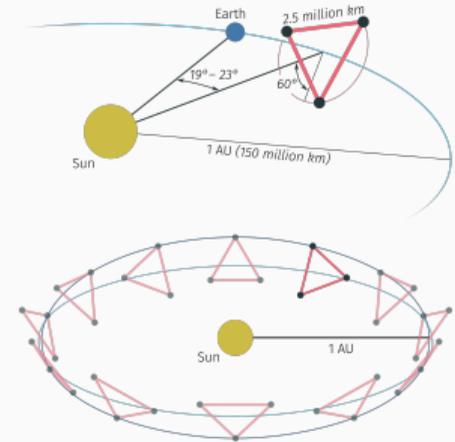
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Università di Trento & INFN/TIFPA

# WHAT IS LISA AND HOW DOES IT WORK?

LISA is a **space mission** to detect *gravitational waves*:

- Free-falling test masses are **tidally accelerated** by curvature.
- Relative acceleration is detected as a **frequency modulation** of the received laser beam.
- Spacecrafts have the goal to **protect** the purity of the free fall of test masses.

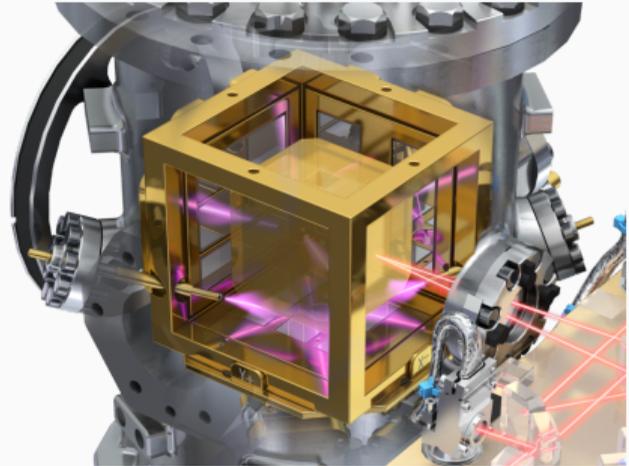


# NEED FOR CHARGE MANAGEMENT AND LPF HERITAGE

## Problem

Cosmic rays and SEP charge **positively** the TMs. The net charge rate for CRs is expected to be between  $+10 e/s$  and  $+100 e/s$ .

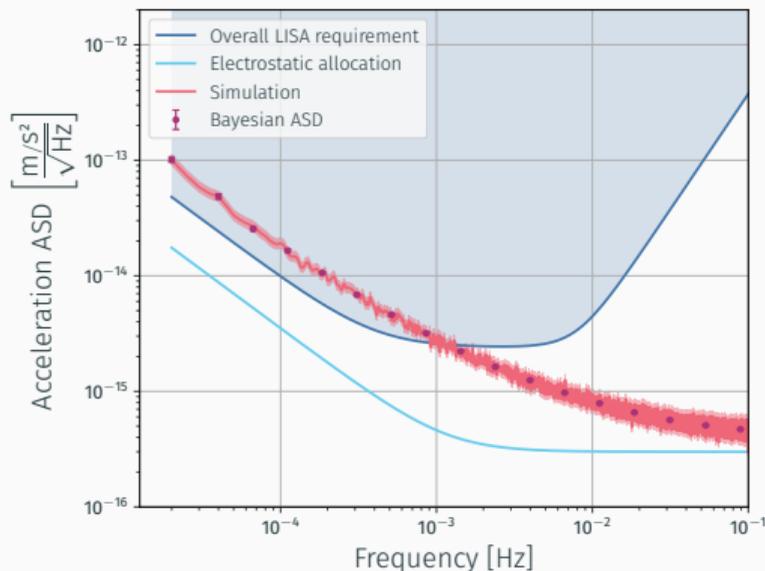
- TM charge is required to be kept in the range  $\pm 1.5 \cdot 10^7 e$  (i.e.  $\pm 70 mV$ ) to ensure the full LISA sensitivity.
- Need for a **charge management system** (CMS) to keep the TM voltage under control.
- LPF successfully proved that this could be achieved with Hg UV-lamps.



# BUT, WHY?

The **electrostatic force** on the TM (at *linear order*) is

$$F_x = \frac{Q_{TM}}{C_{tot}} \frac{dC_x}{dx} \Delta_x$$



The *two* main sources of **electrostatic force noise** are:

- Interaction of a charged TM with **noisy stray potentials** (*dominant*)

$$S_{F_x(\delta\Delta_x)} = \left[ \frac{Q_{TM}}{C_{tot}} \frac{dC_x}{dx} \right]^2 S_{\Delta_x}$$

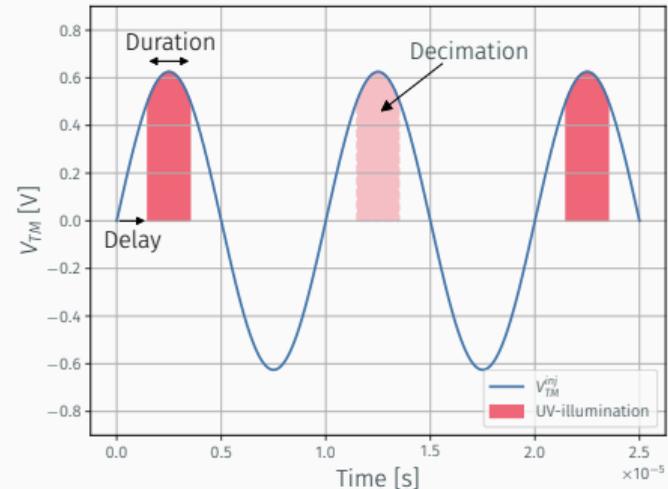
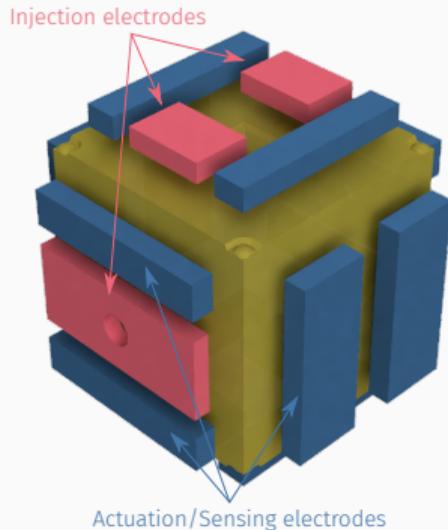
- Interaction of a **fluctuating TM charge** with DC stray potentials (*relevant at low-frequencies*)

$$S_{F_x(\delta Q_{TM})} = S_{Q_{TM}} \left[ \frac{1}{C_{tot}} \frac{dC_x}{dx} \Delta_x \right]^2$$

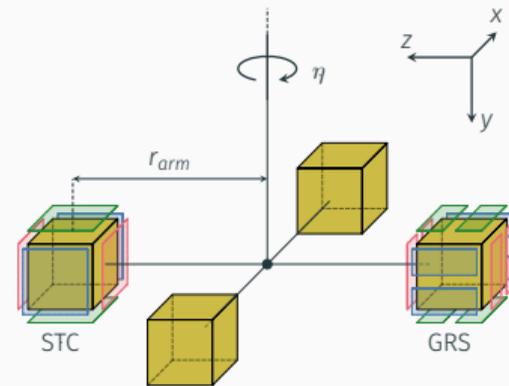
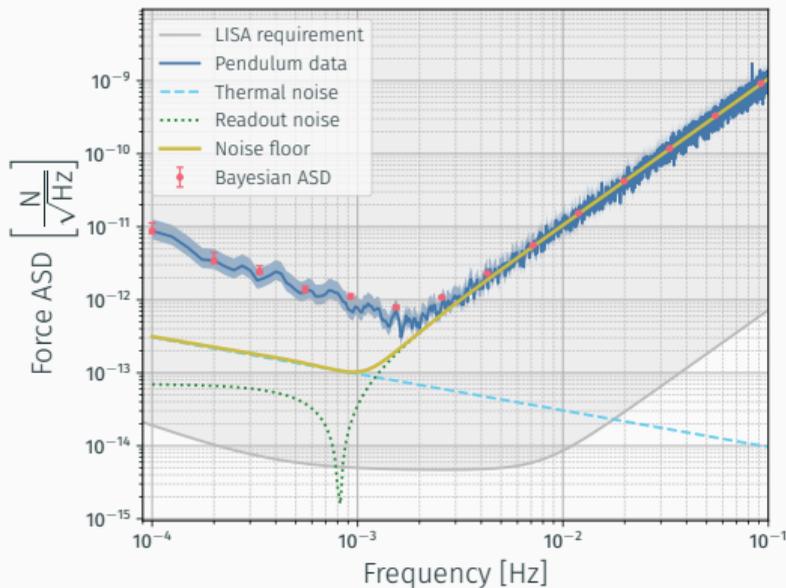
# CHARGE MANAGEMENT IN LISA

For LISA the current design aims to use **UV-LEDs**, with the following advantages:

- Can be **synchronized** with the 100 kHz TM bias (also improved redundancy).
- Weight and volume saving.
- More reliable and robust (longer lifetime).

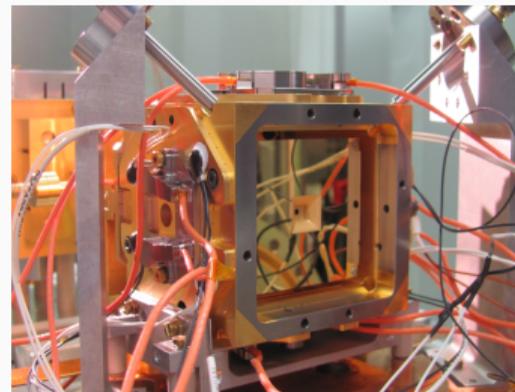


# 4-MASS TORSION PENDULUM AT UTN



EH ISUK

TM ISUK



## Main features of our facility

Flight model replica of LPF GRS, copy of LPF FEE for actuation/sensing, ISUK positioned as in LPF.

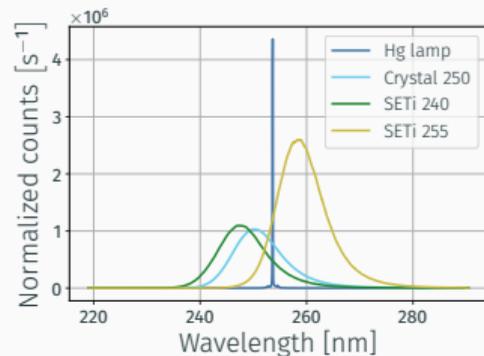
# UV LIGHT SOURCES AND PROTOTYPE CMD

## Light sources

- LPF Hg-lamp replica
- SETi 255
- SETi 240
- Crystal 250

## UV power measurement

- Si-Photodiode (*S1337-1010BQ*)
- PMT (*H6780-03*)
- Spectrometer (*AvaSpec-ULS2048XL-EVO*)

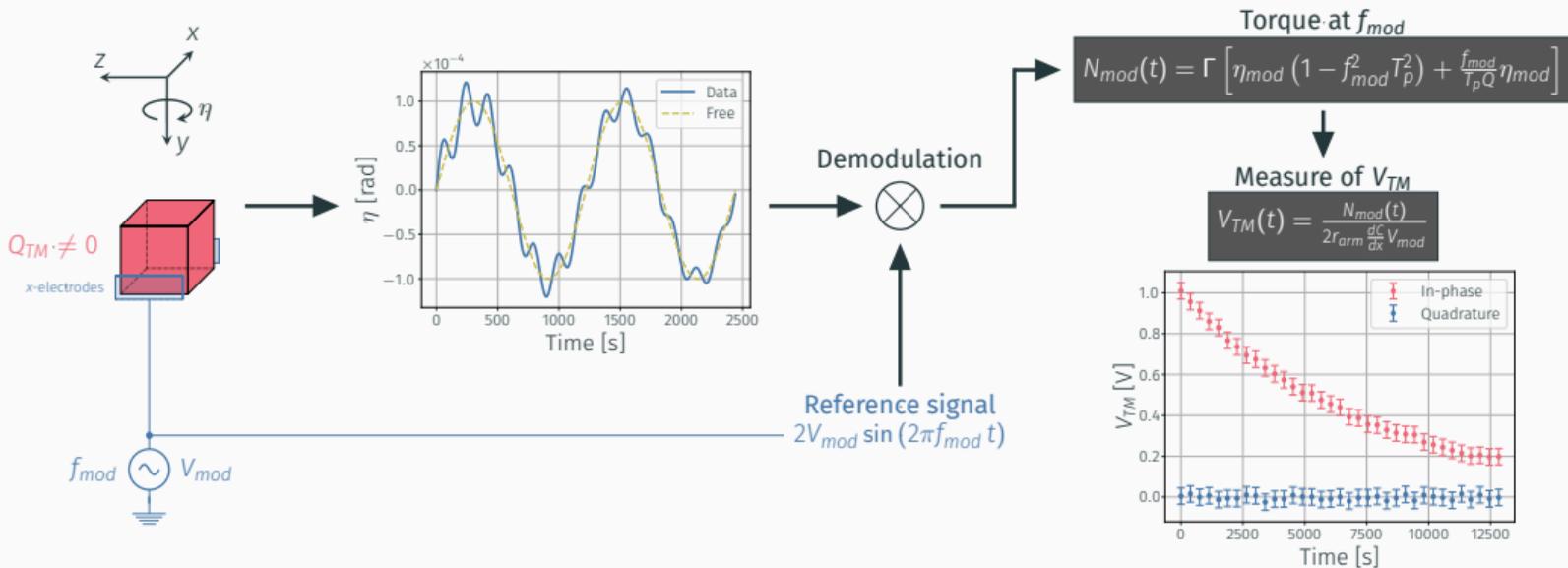


## Charge Management Device

Prototype CMD provided by the INFN/Roma Tor Vergata group.



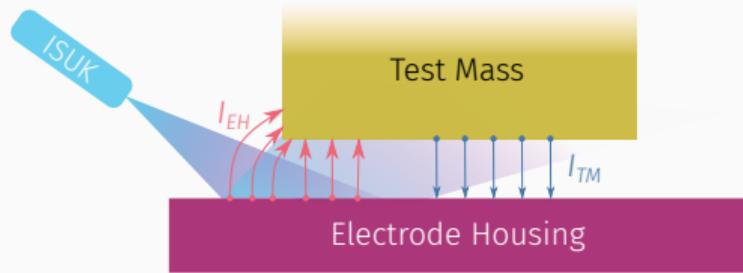
# APPARENT YIELD MEASUREMENTS



Apparent yield: 
$$AY \equiv \frac{\# \text{ charges/s}}{\# \text{ photons/s}} = \frac{\frac{1}{e} \frac{dQ_{TM}}{dt}}{\frac{P_{UV}}{hf_{UV}}} = \frac{hf_{UV}}{eP_{UV}} C_{tot} \frac{dV_{TM}}{dt}$$

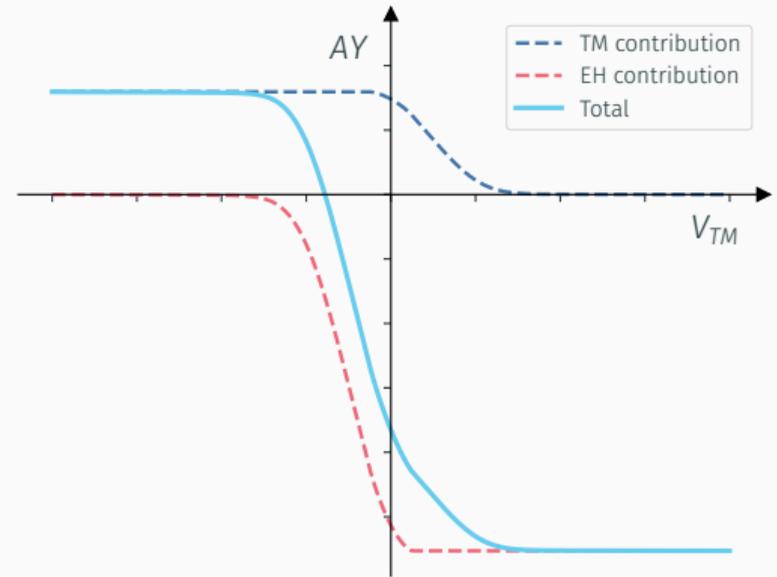
## EXPECTED QUALITATIVE BEHAVIOR

As the inner surfaces of the gravitational reference sensor (GRS) are *reflective*, a fraction of the incident UV-light is **absorbed** by the opposing surface.

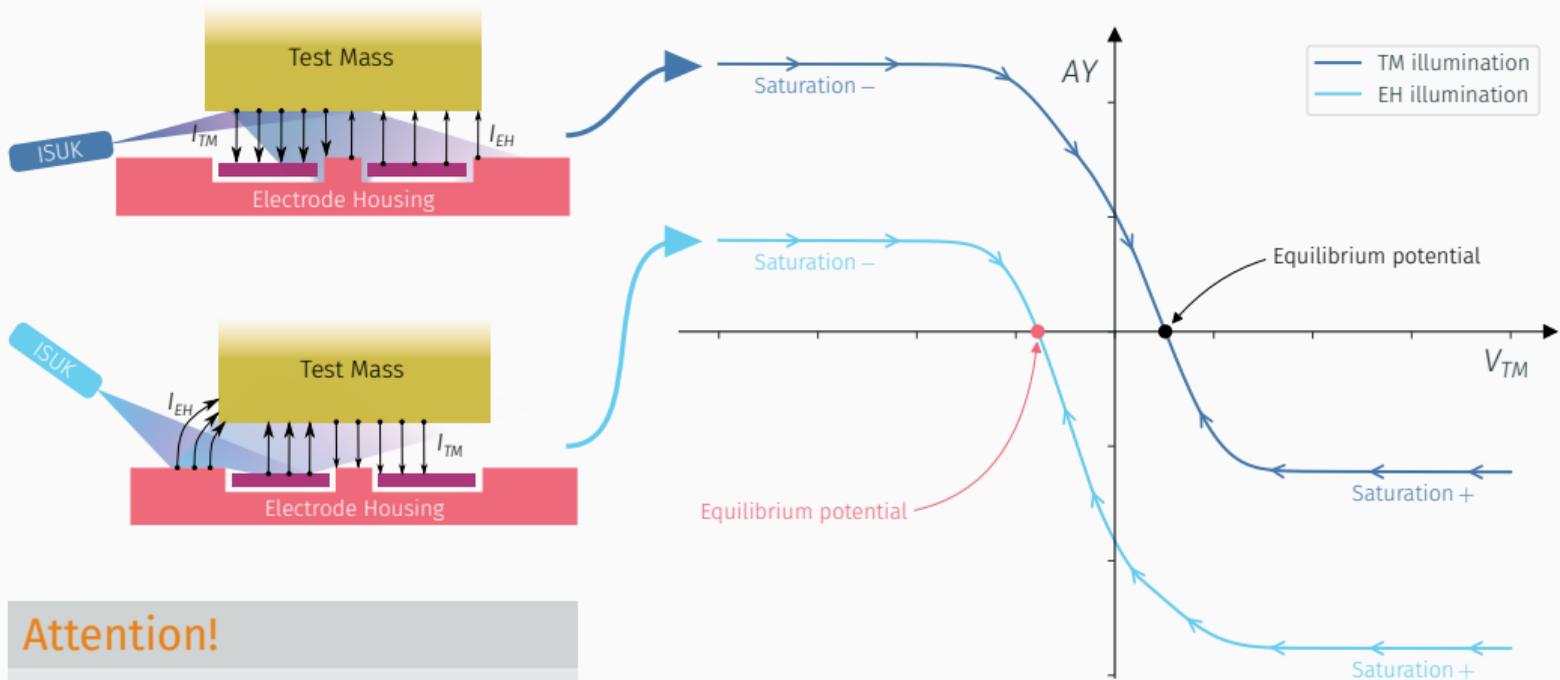


### Attention!

The arrows indicate the direction of  $e^-$  motion!



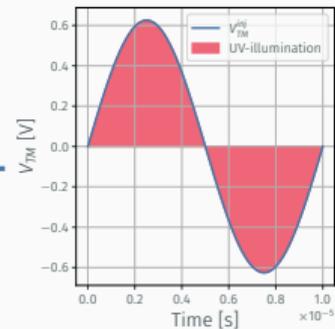
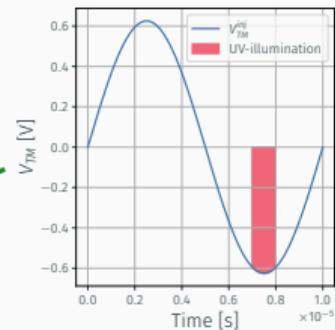
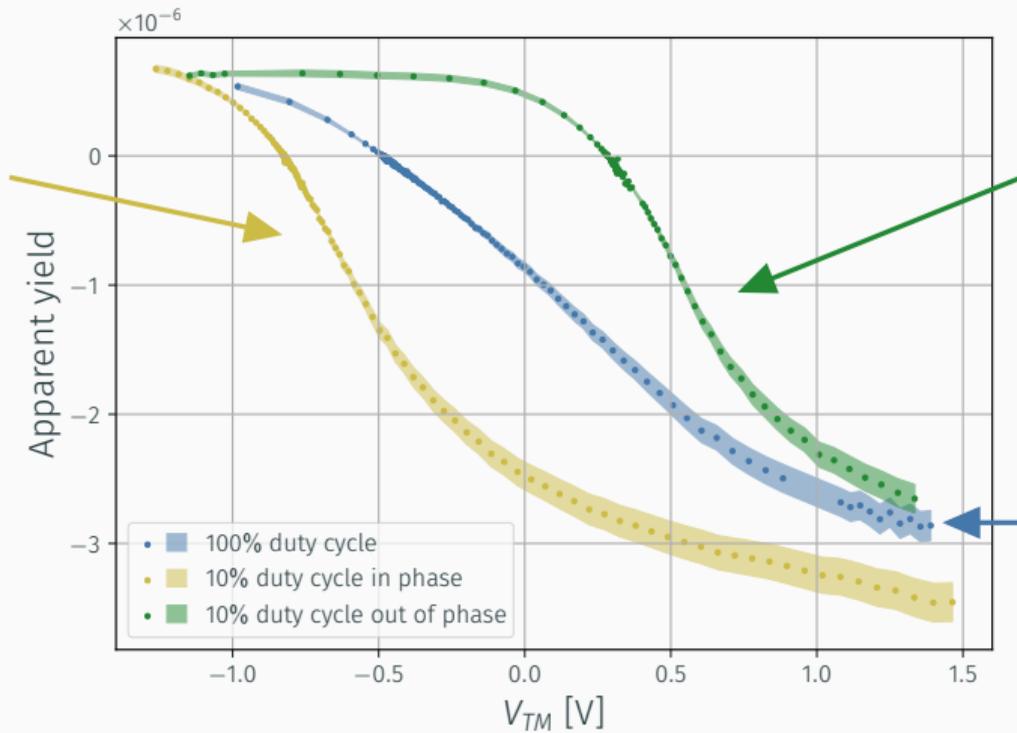
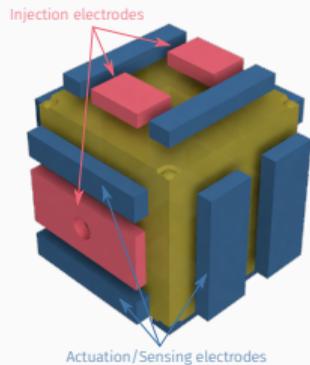
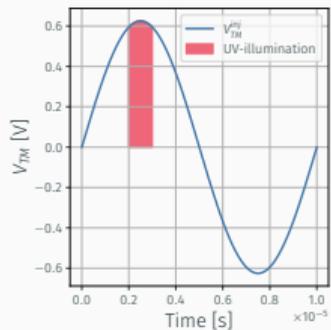
# DIFFERENCES FOR TM AND EH ILLUMINATIONS



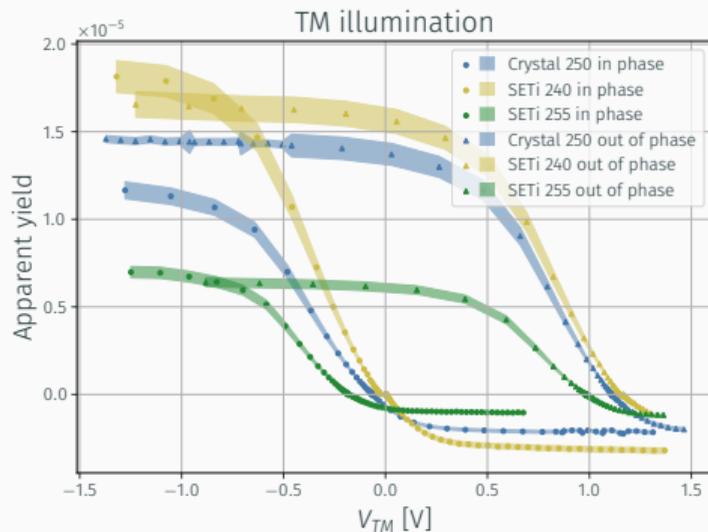
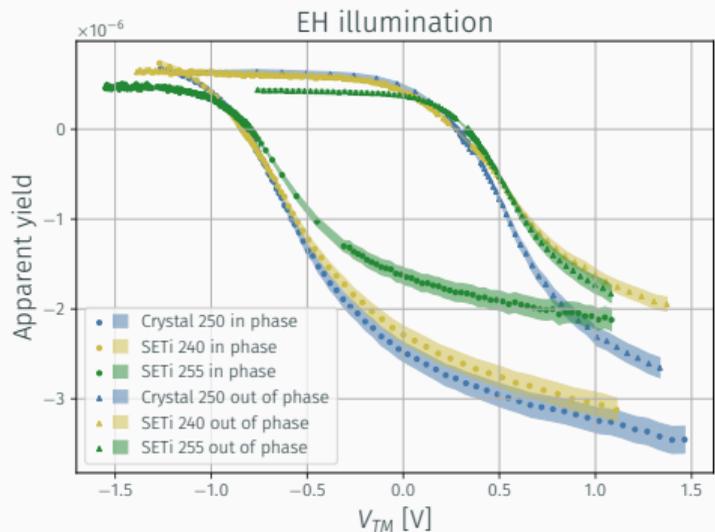
## Attention!

The arrows indicate the direction of  $e^-$  motion!

# MEASURED APPARENT YIELD (CRYSTAL 250, EH ILLUMINATION)



# APPARENT YIELD FOR EH AND TM ILLUMINATIONS



**NB!**

We had some issues regarding the power normalization!

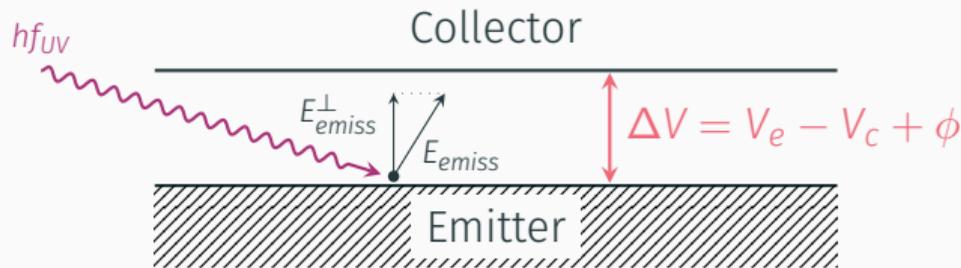
## INTERPRETATION OF THE DATA: A SIMPLE MODEL

The model is based on the following **assumptions**:

- **Idealized geometry**: EH and TM surfaces are plane parallel sheets.
- The electrons in the emitter are modeled as a free *Fermi gas*. The absorbed photon brings their normal kinetic energy to  $E_0^\perp + hf_{UV}$ .
- If  $E_0^\perp + hf_{UV} > E_{F,e} + W_e$ , there is a non-zero probability of being emitted with

$$E_{emiss}^\perp = E_0^\perp + hf_{UV} - W_e - E_{F,e}$$

- All emitted  $e^-$  with energy larger than the *potential barrier*  $\Delta V$  contribute to the photocurrent (where  $\phi$  is the contact potential).



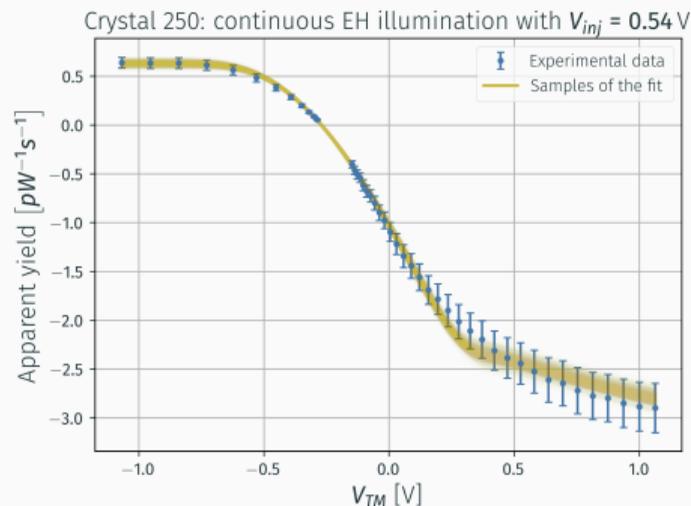
## FIT OF THE DATA

The Bayesian fit obtained by sampling the following *Gaussian log-likelihood* with a **Hamiltonian Monte Carlo** algorithm (NUTS by the API PyMC3)

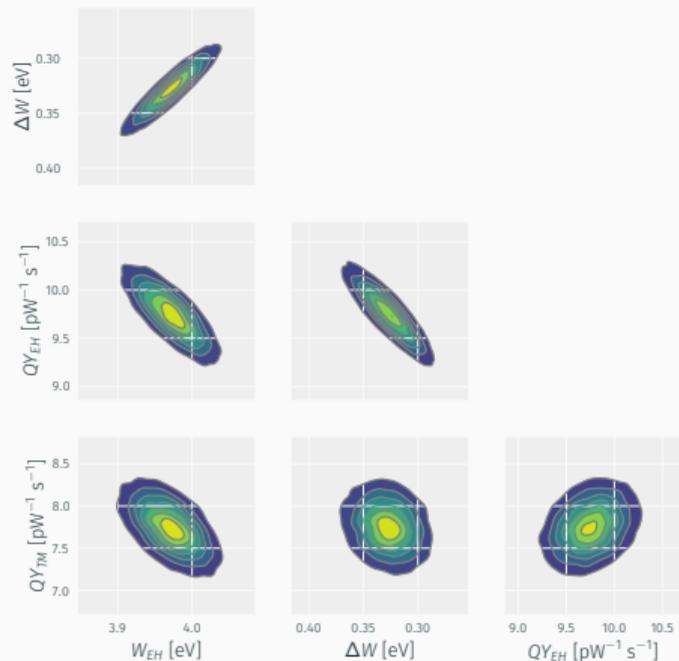
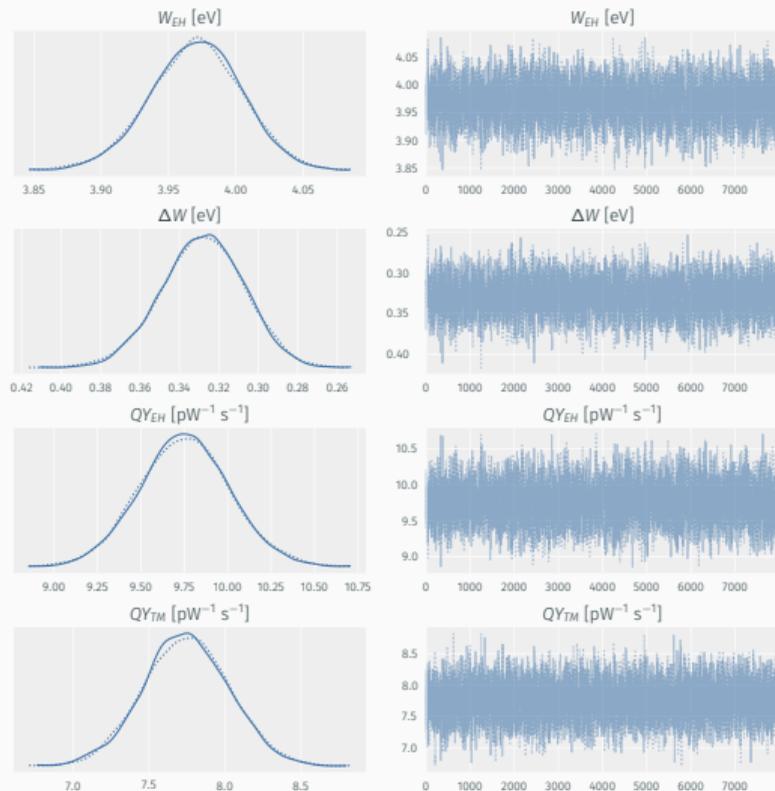
$$LL(\mathbf{y}|\boldsymbol{\theta}) = -\frac{1}{2} \sum_i \frac{[y_i - f(\boldsymbol{\theta}, x_i)]^2}{\sigma_i^2}$$

The fit parameters  $\boldsymbol{\theta}$  are:

- The work function of EH ( $W_{EH}$ )
- The difference of TM and EH work function ( $\Delta W = W_{TM} - W_{EH}$ )
- The quantum yield of EH ( $QY_{EH}$ )
- The quantum yield of TM ( $QY_{TM}$ )



# POSTERIOR DISTRIBUTION OF FIT PARAMETERS



## CONCLUSIONS AND FUTURE WORK

Except for some anomalies, we experimentally proved the **advantages** of using UV-LEDs synchronized with the injection bias, such as of *tuning the TM voltage* and the *discharge rate* with the phase of the illumination.

We have a simple but effective model to interpret the data and extract relevant surface parameters (e.g.  $W_{EH} = 3.97 \pm 0.03$  eV and  $W_{TM} = 3.64 \pm 0.05$  eV), which would be hard to measure otherwise.

### Future steps

- Apply the fitting procedure to all the AY data.
- Improve our models to interpret the data (*FEM model* for photoelectrons trajectories and fields, accurate *mapping light absorption*).

THANK YOU FOR THE ATTENTION!

