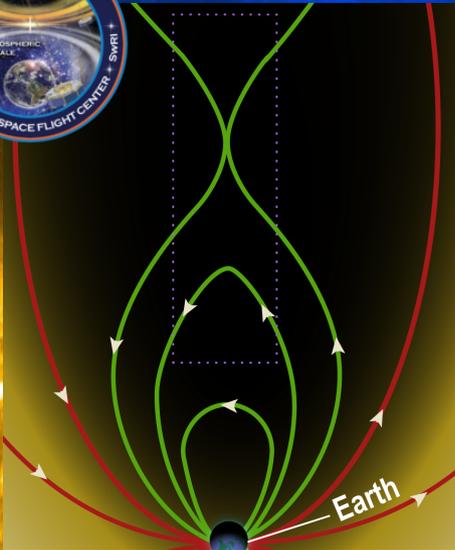
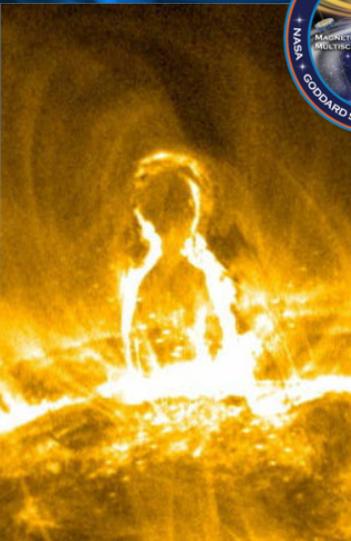


Suprathermal electrons in Earth's magnetotail: What can they teach us about flare electron energization?



***Li-Jen Chen, Shan Wang, James Drake,
Spiro Antiochos, Tai Phan***



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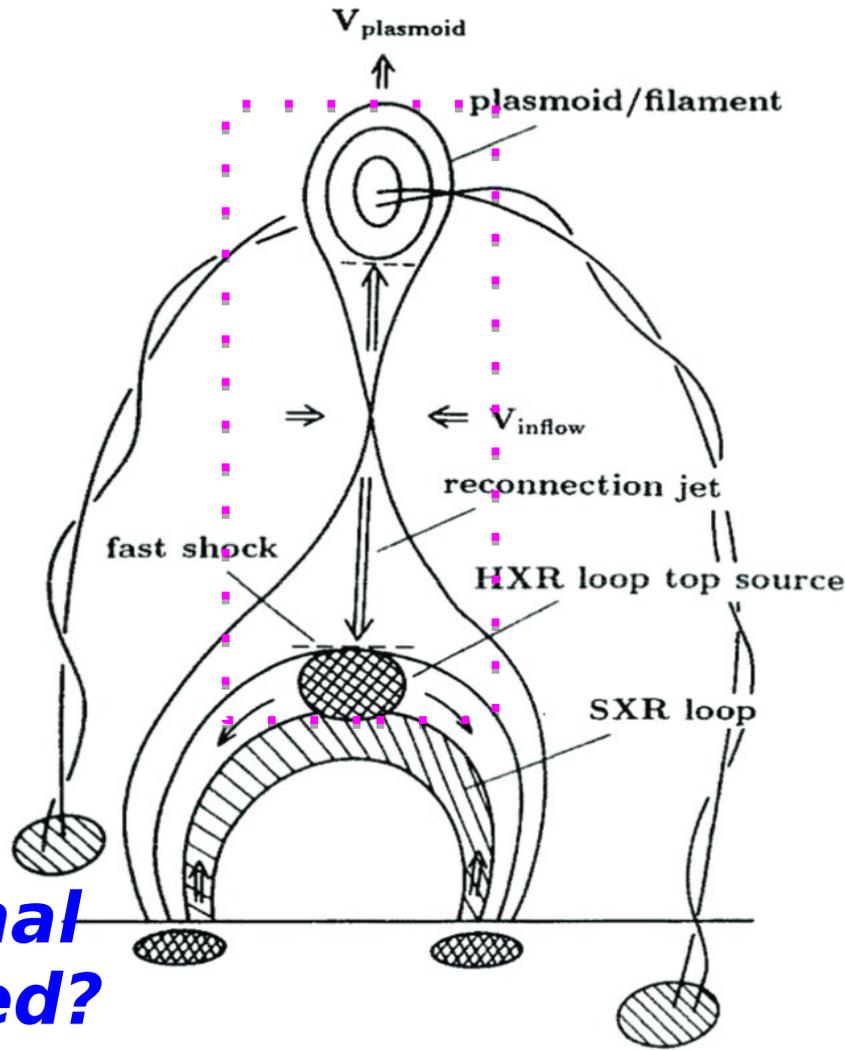
+

other MMS & CLUSTER team members

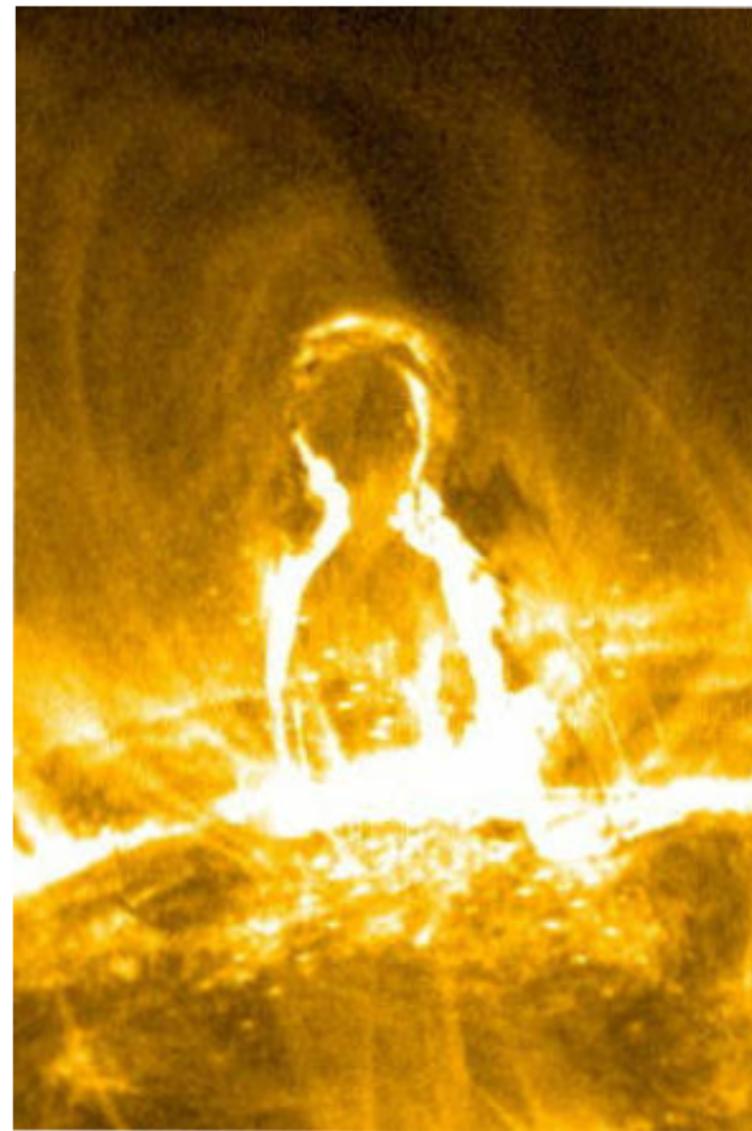
Up to 50% of flare released energy goes to suprathermal electrons

[e.g., Lin+, 2003
Holman+, 2003;
Krucker+, 2010;
Krucker & Battaglia, 2014]

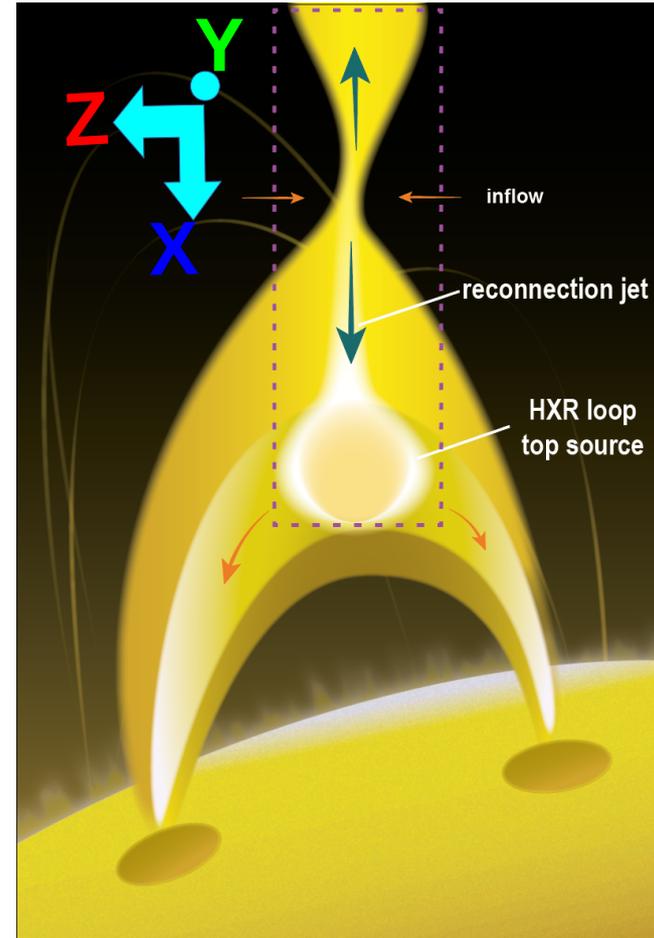
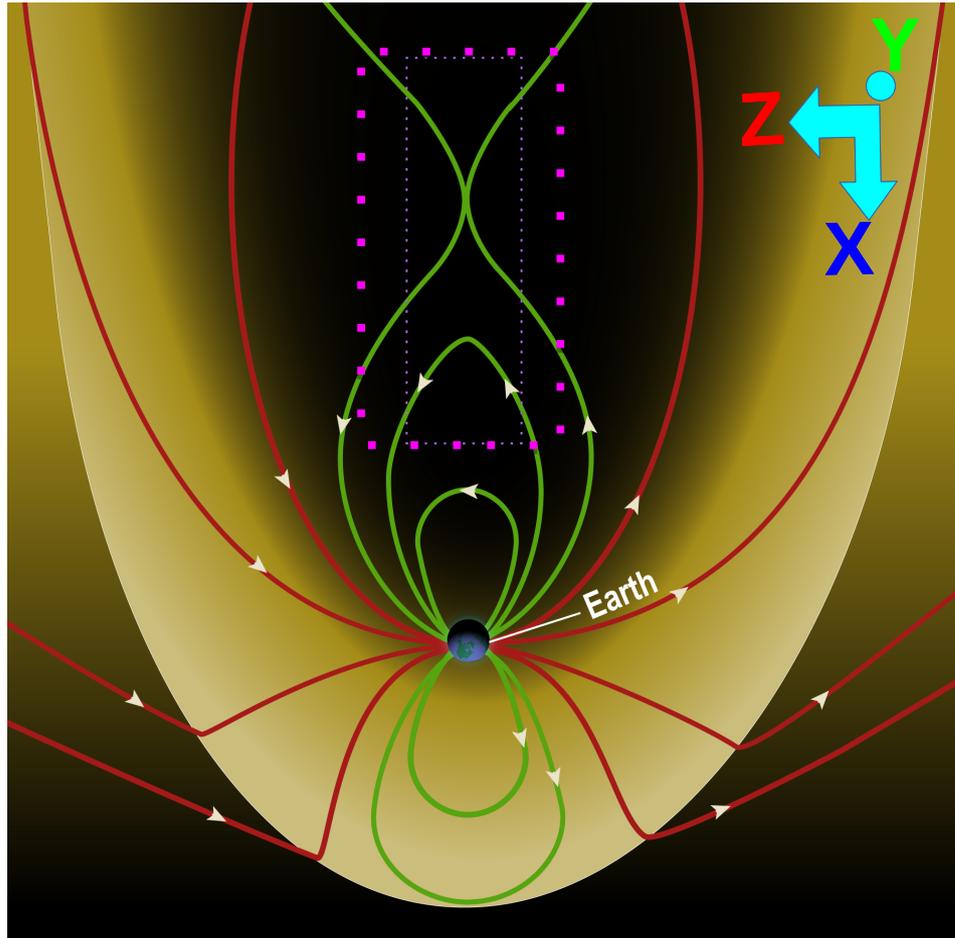
How are suprathermal e's produced?



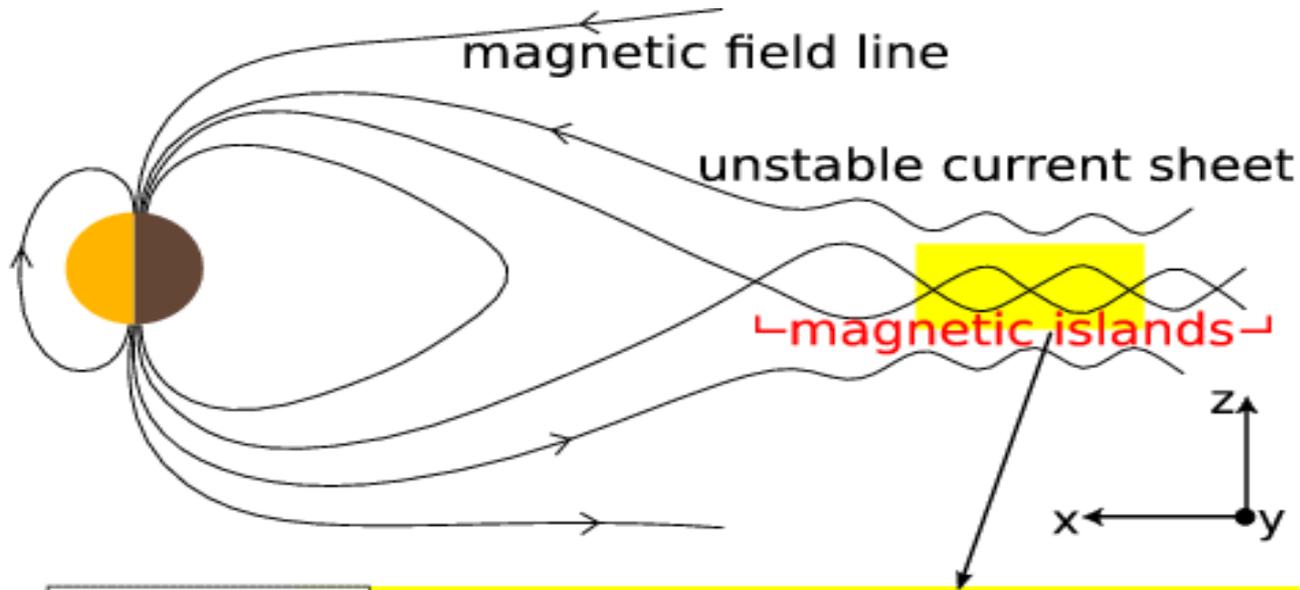
[Shibata et al., 1995]



Magnetic configuration & beta in solar flares are similar to those in the magnetotail

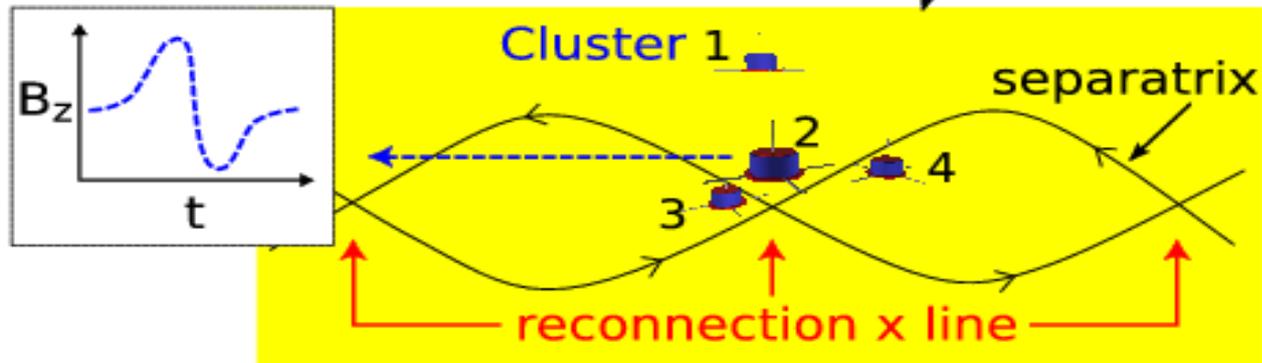


Volume-filling magnetic islands?



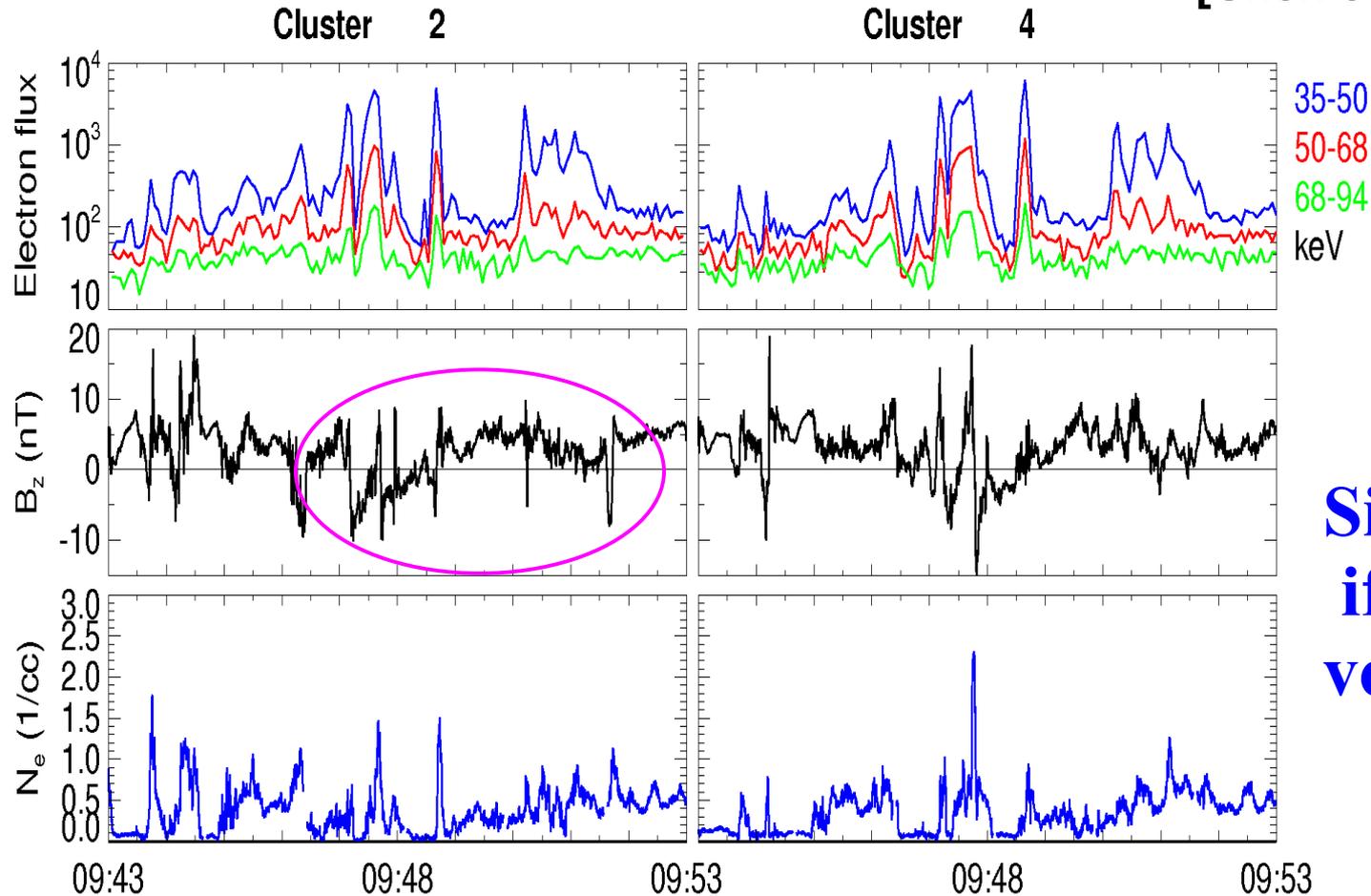
Motivated by
[Drake et al.,
Nature, 2006]

Later work:
[Dahlin et al., PP,
2014, 2015;
Arnold et al., PRL,
2021]



Energetic electrons in magnetic islands: magnetotail

[Chen et al., Nature Phys., 2008]



**Islands are di scale,
grew out of the
electron current layer**
[Chen et al., PP, 2012]

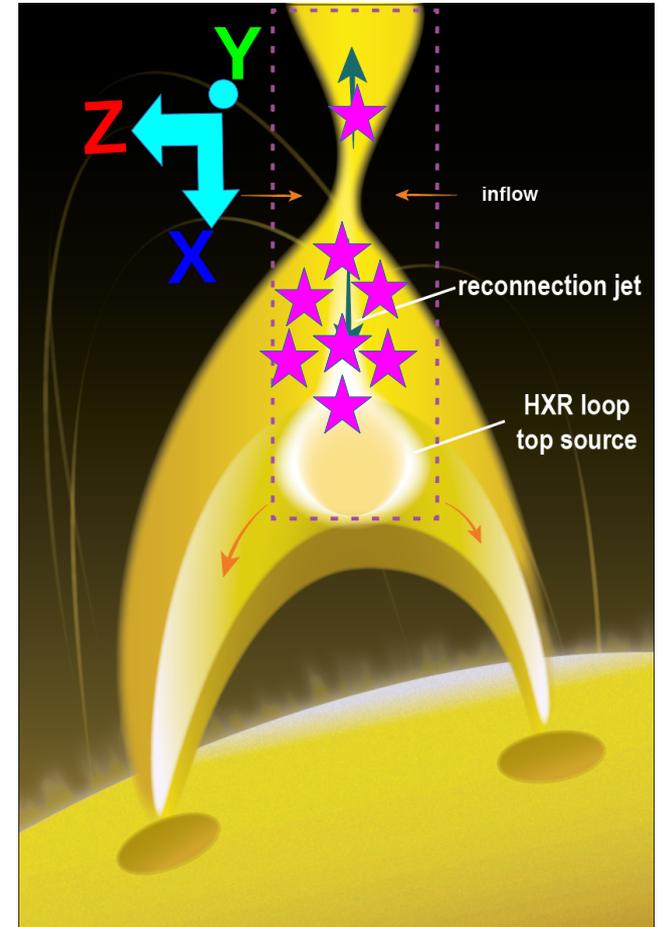
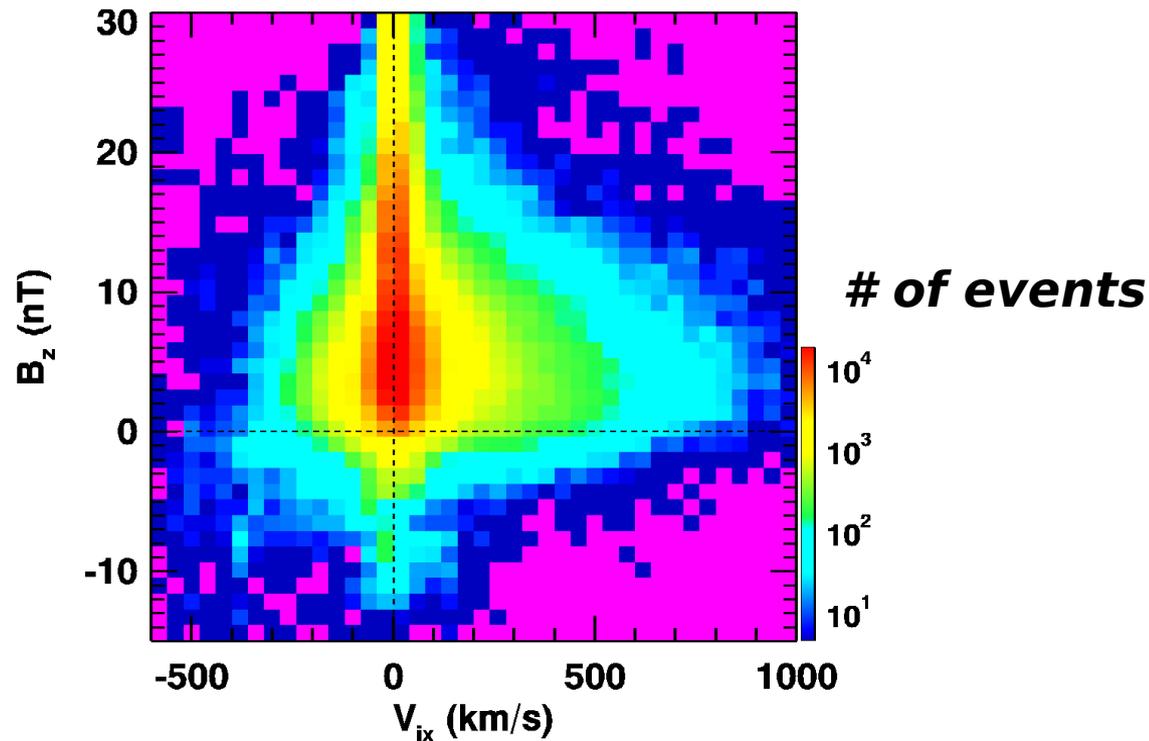
**Signatures =?
if islands are
volume-filling**

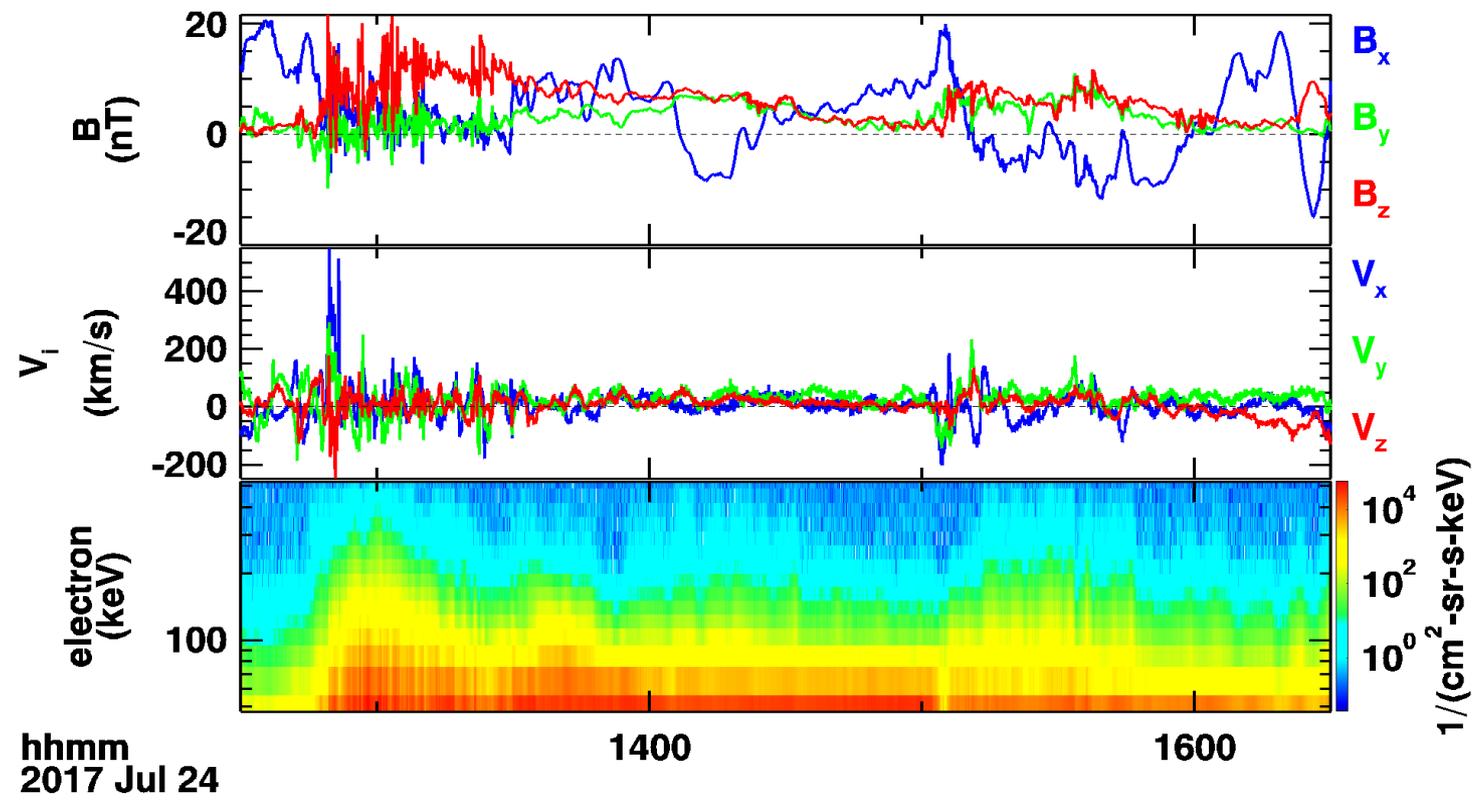
***Use all MMS magnetotail passes
(127 orbits in 4 years)
to test out***

- Whether volume-filling islands are dominant accelerators
- If primary energization occurs at the X line or downstream

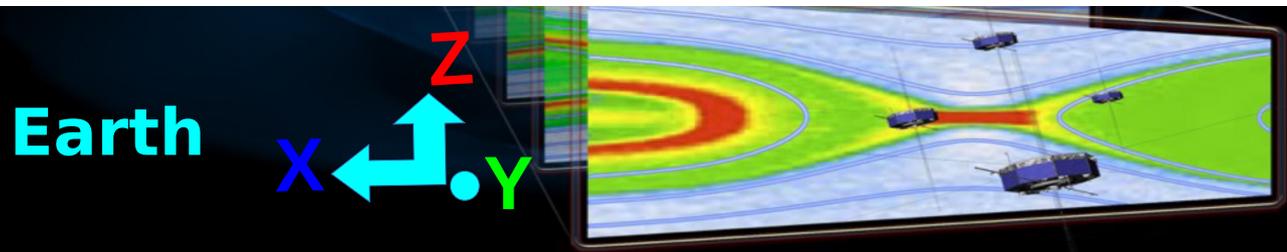
90 keV e's tend to be detected in regions with $B_z > 0$ and $V_{ix} \sim 0$

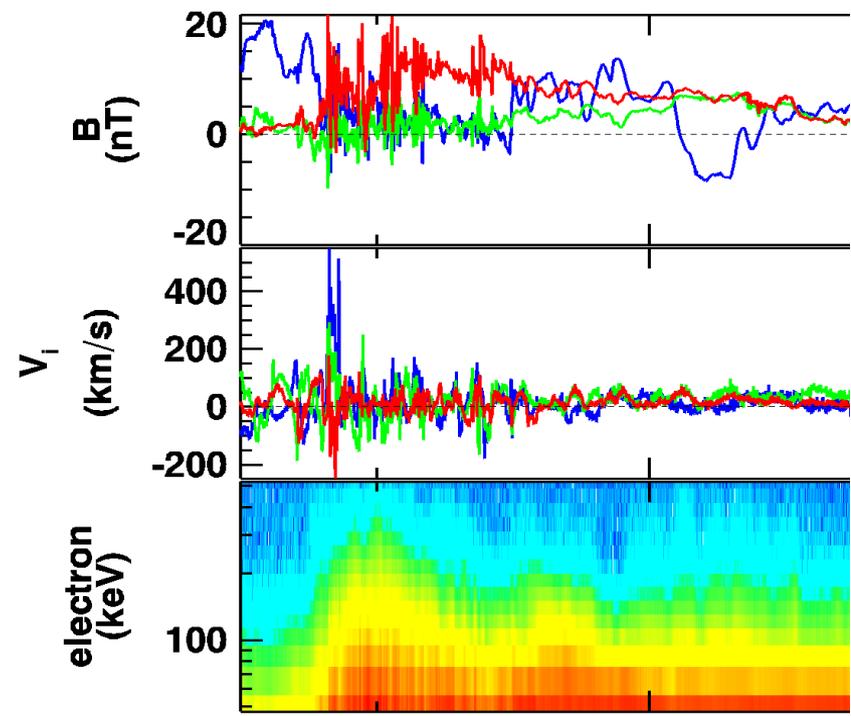
Events: flux > threshold





hhmm
2017 Jul 24

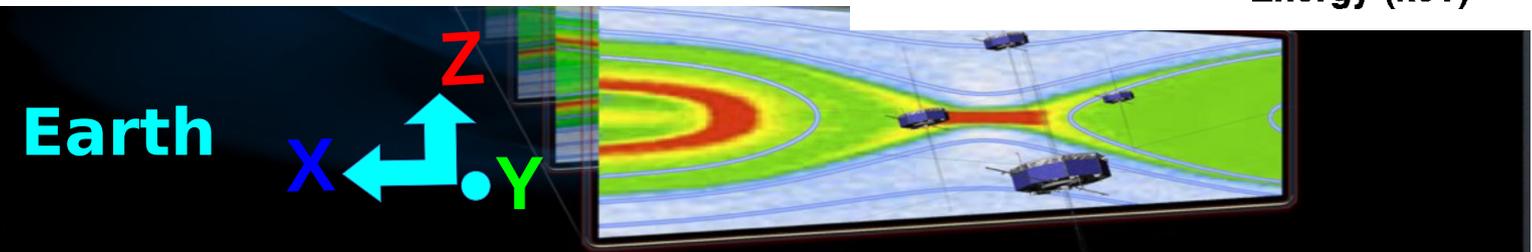
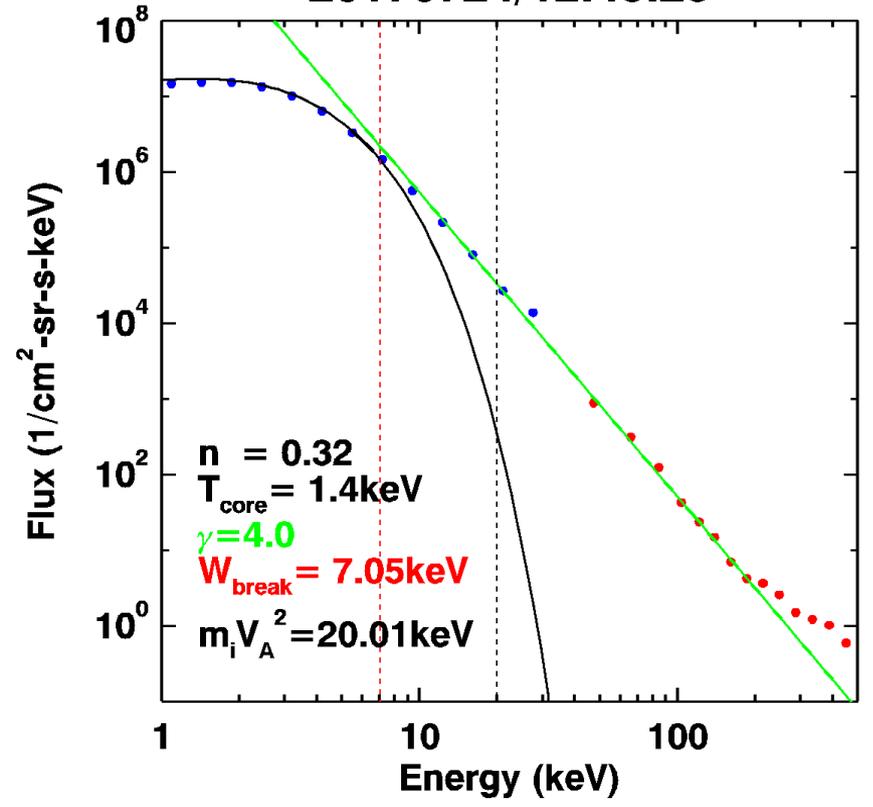


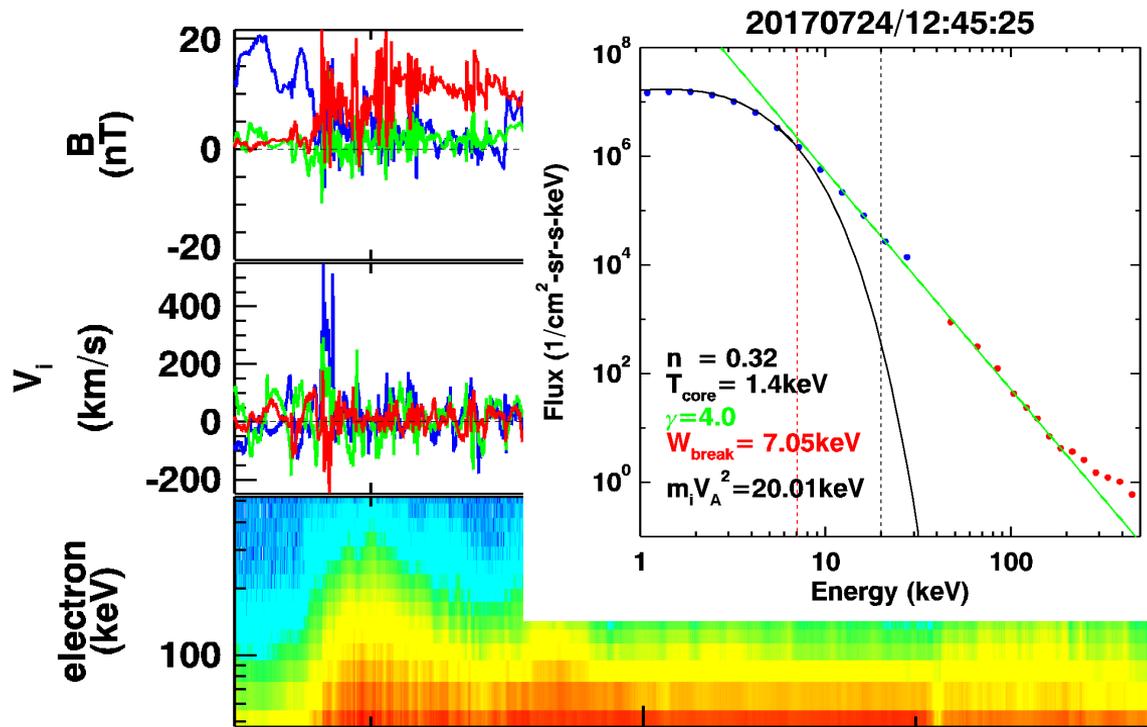


hhmm
2017 Jul 24

1400

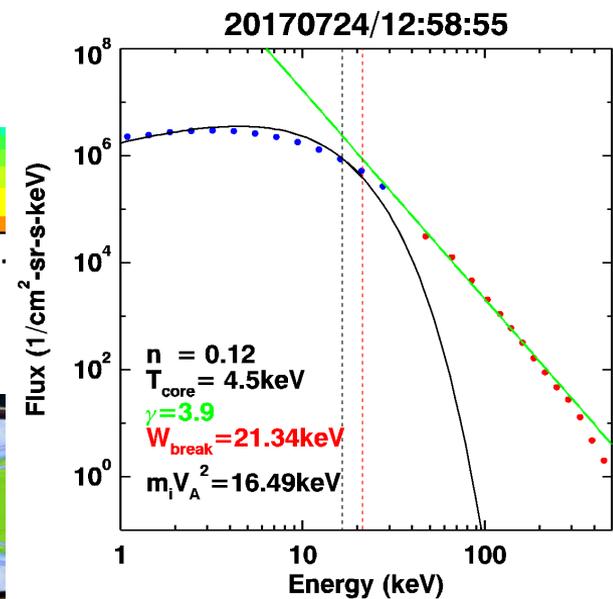
20170724/12:45:25



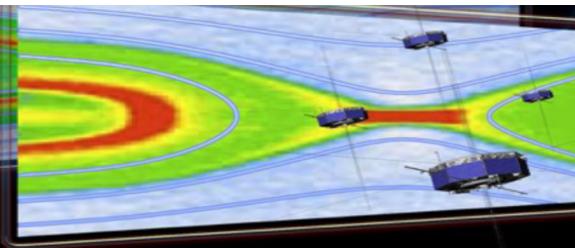


hhmm
2017 Jul 24

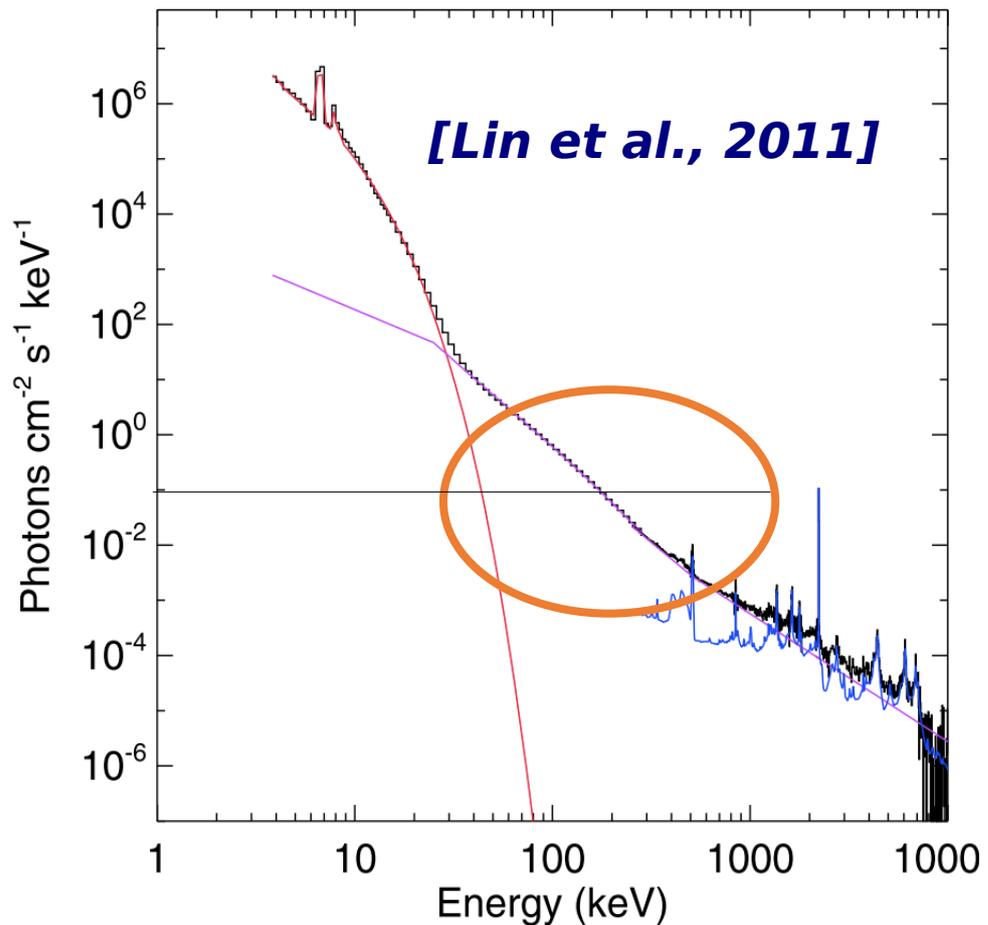
1400



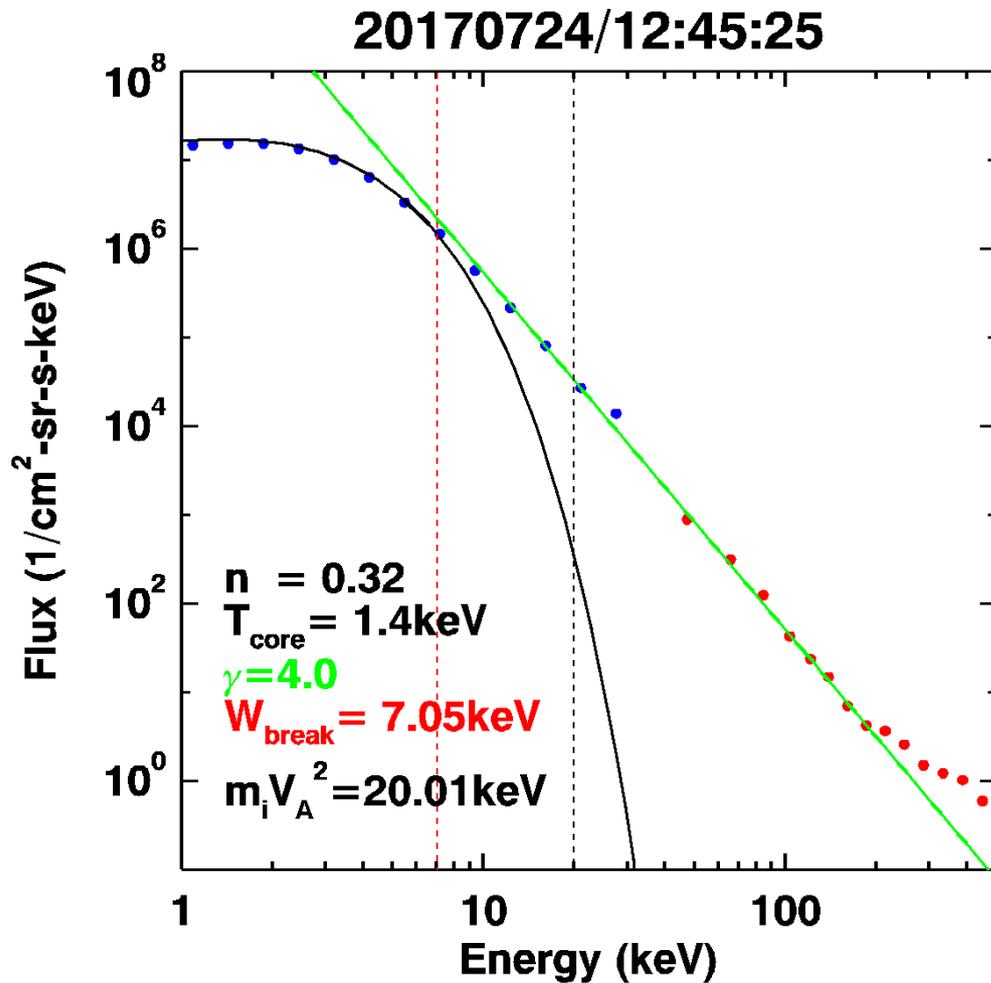
Earth



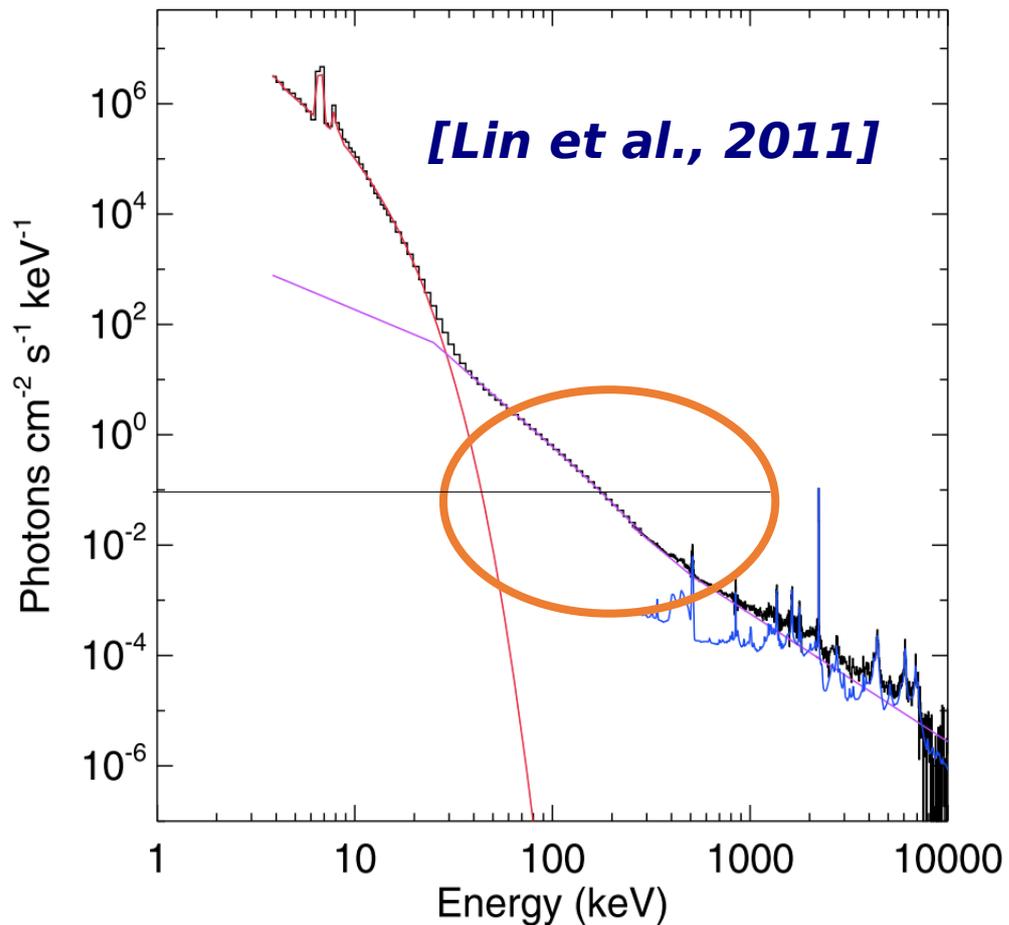
Solar flare



Magnetotail

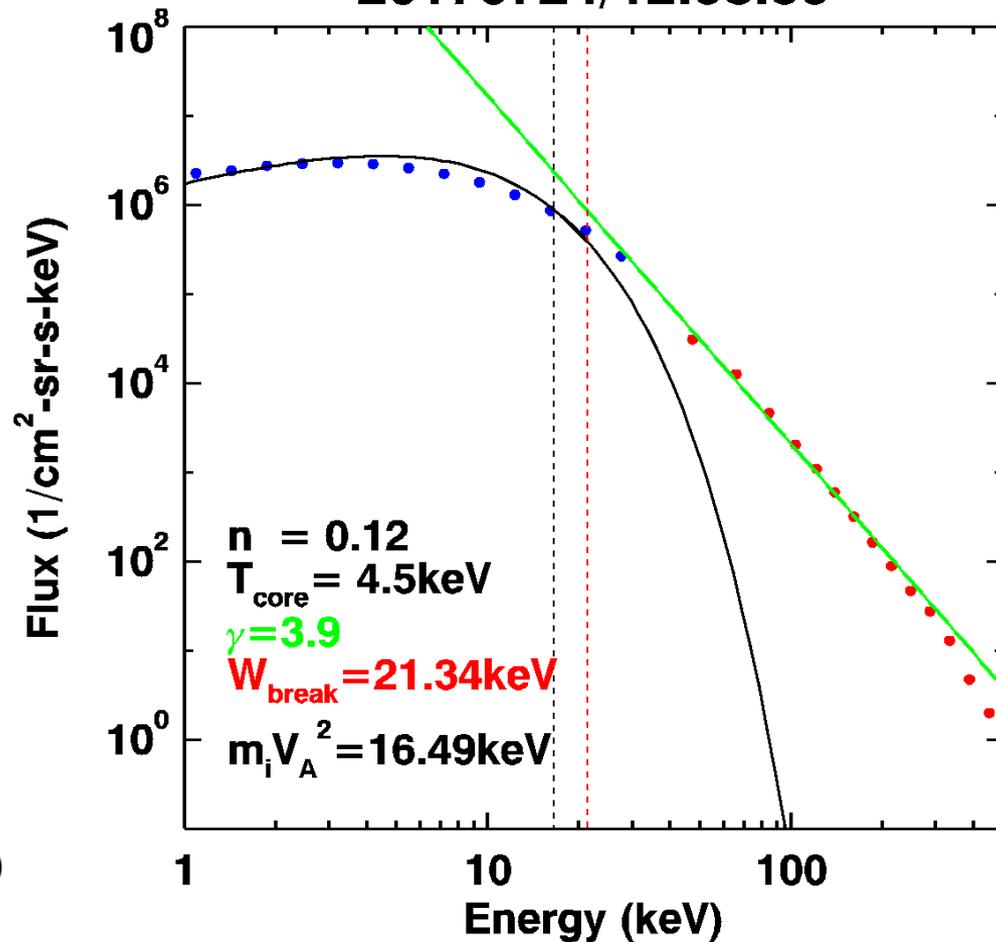


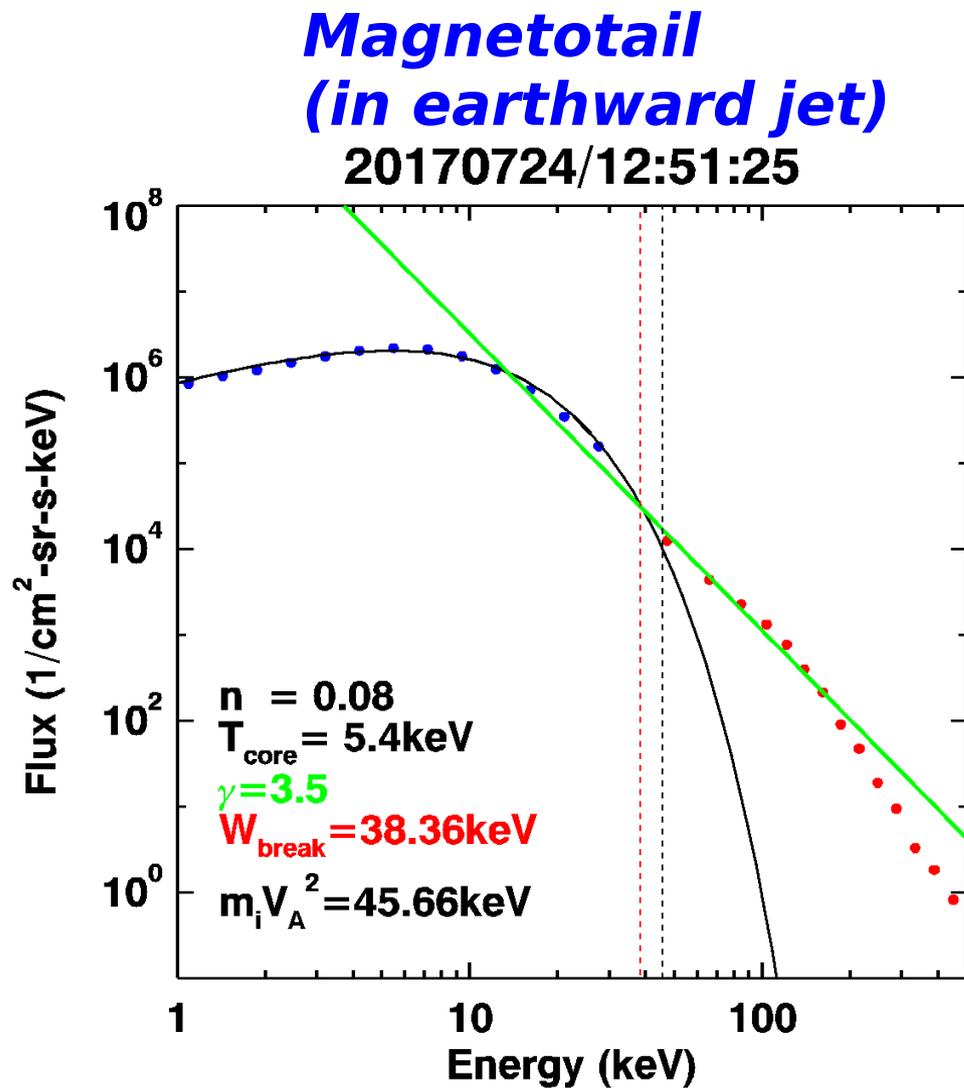
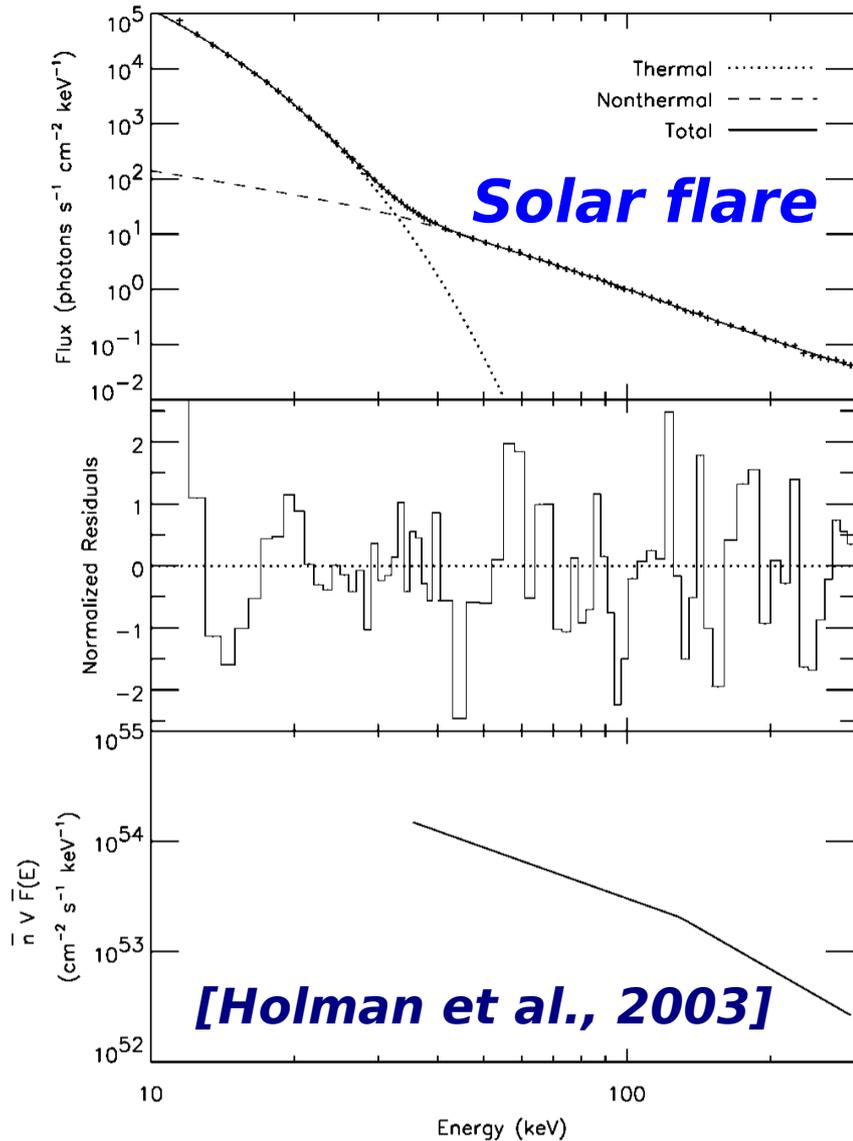
Solar flare



Magnetotail

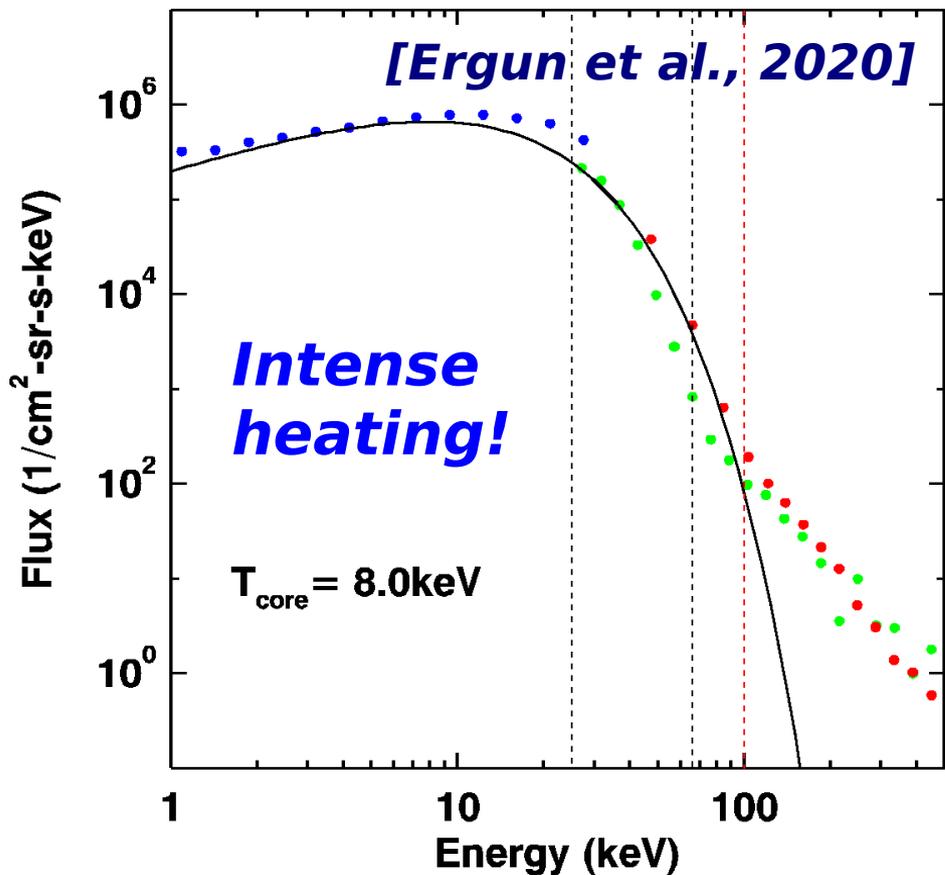
20170724/12:58:55





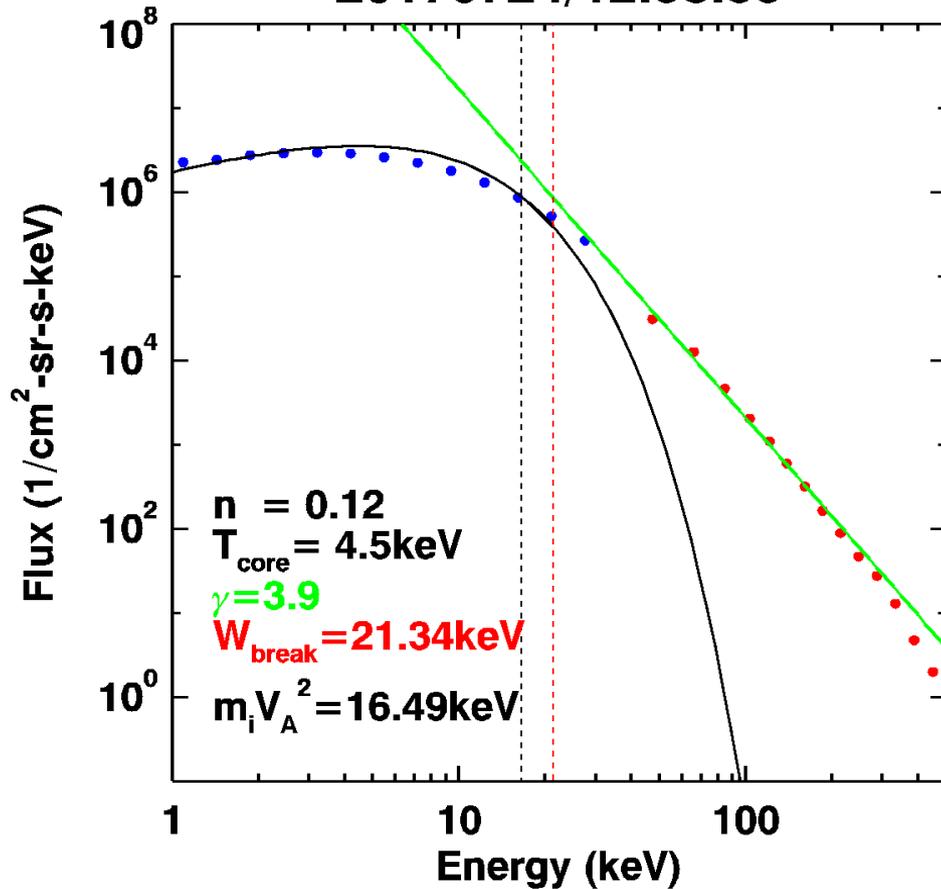
Magnetotail/Turbulence

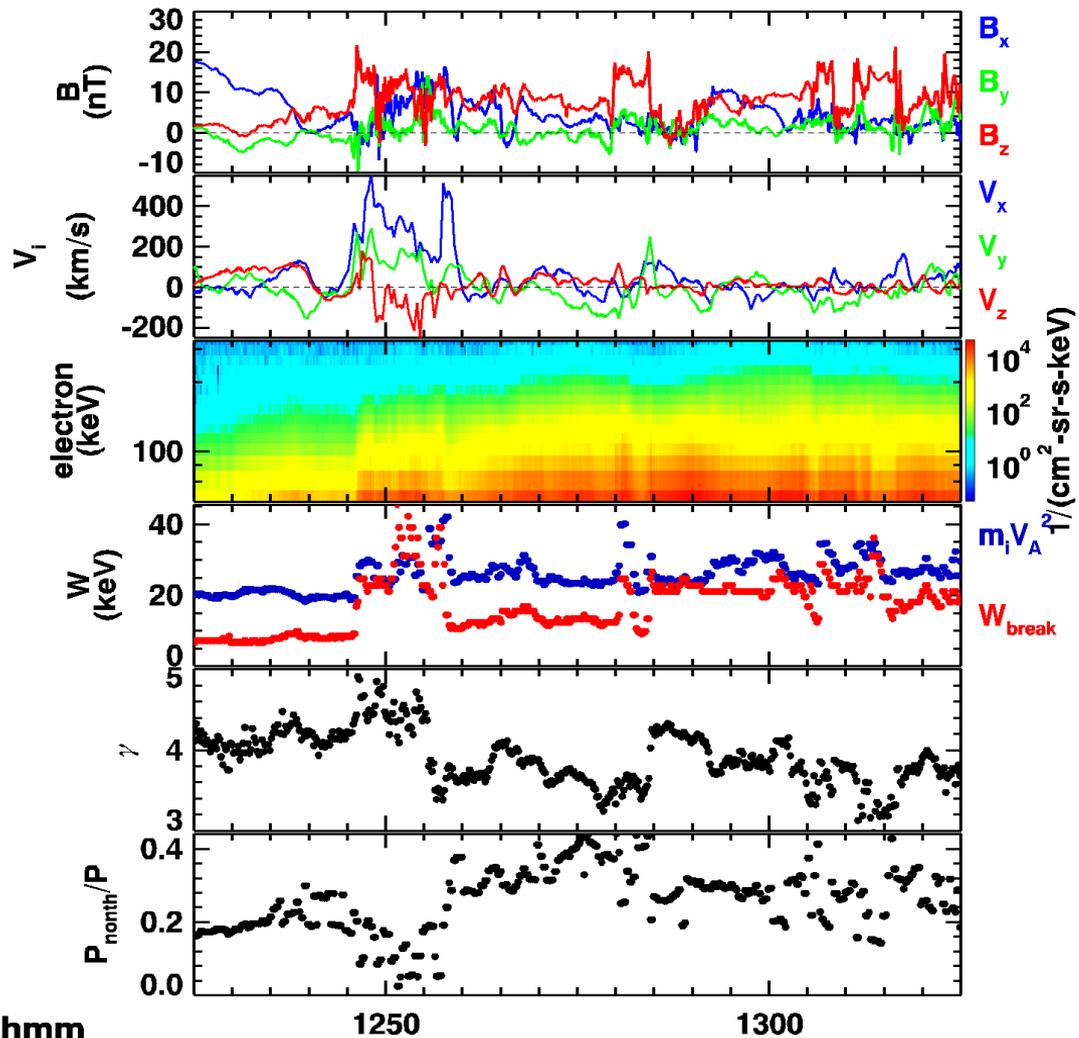
20170726/07:37:20-07:37:23



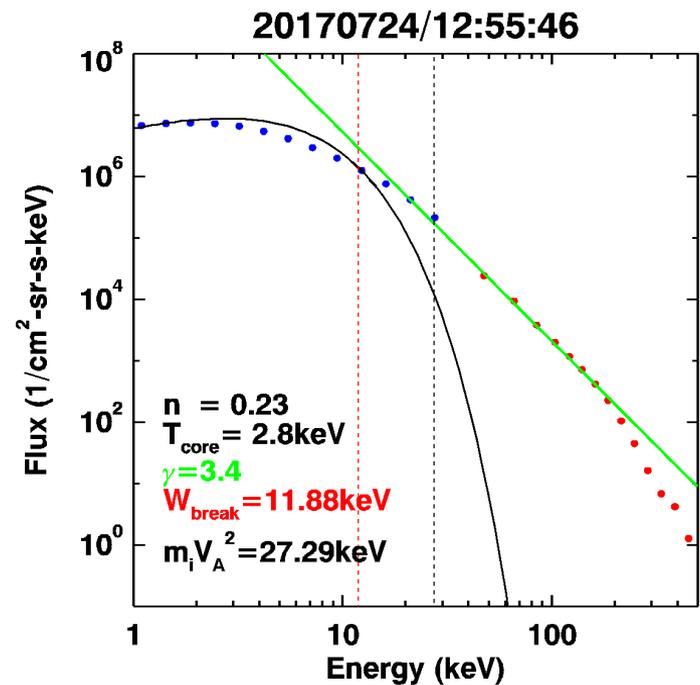
Magnetotail/flux pileup

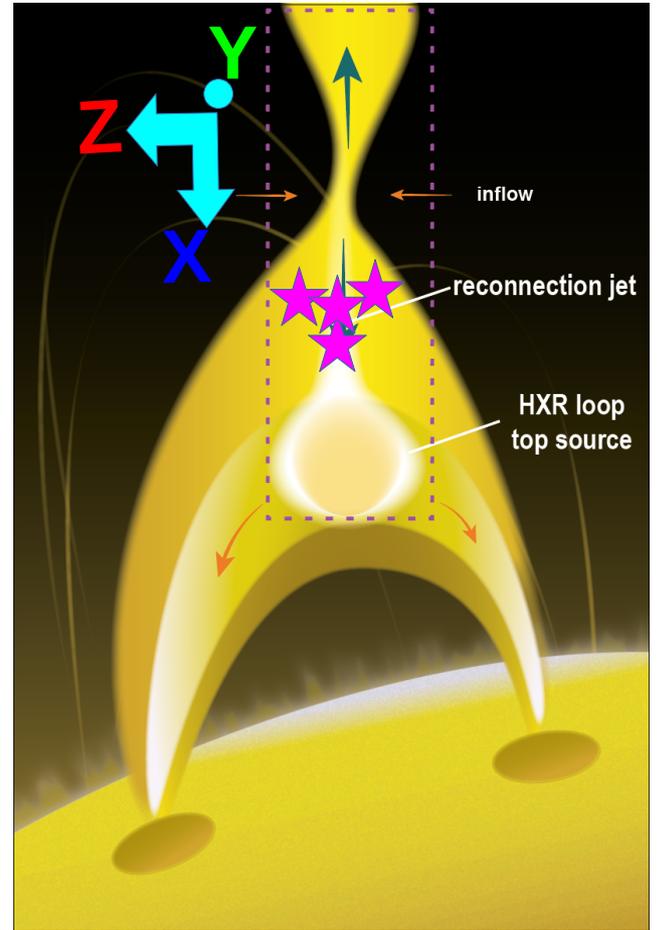
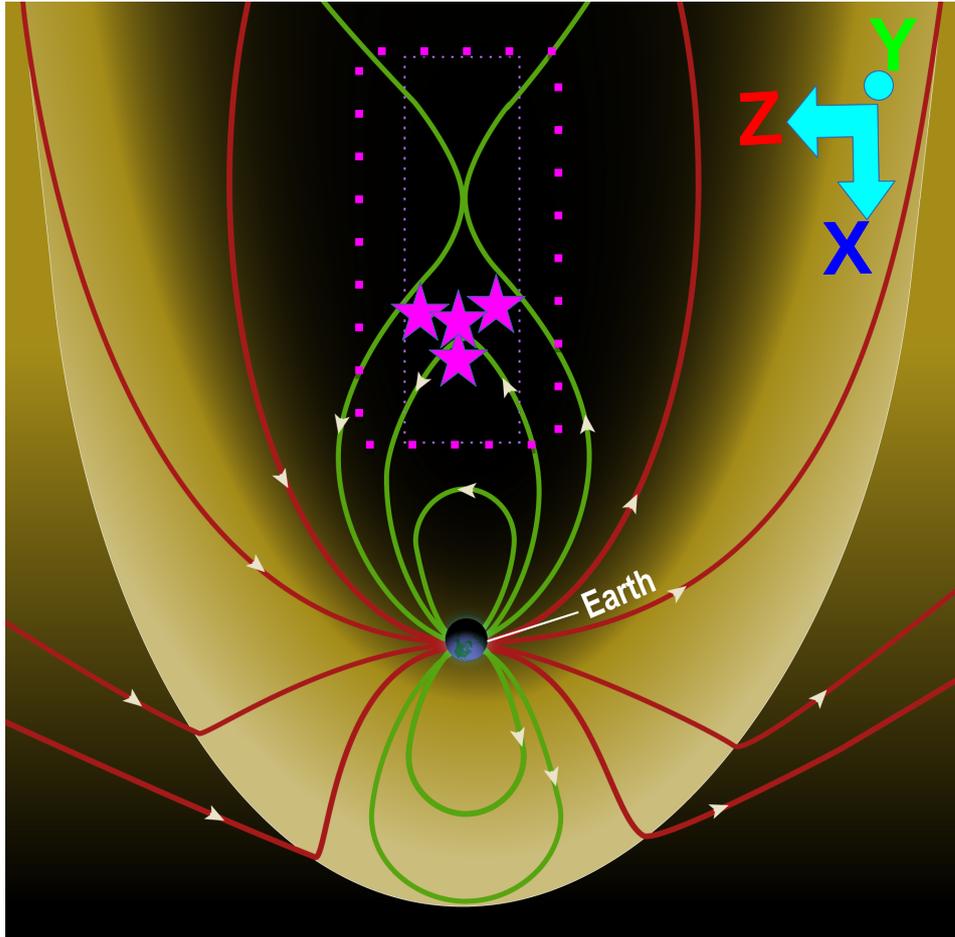
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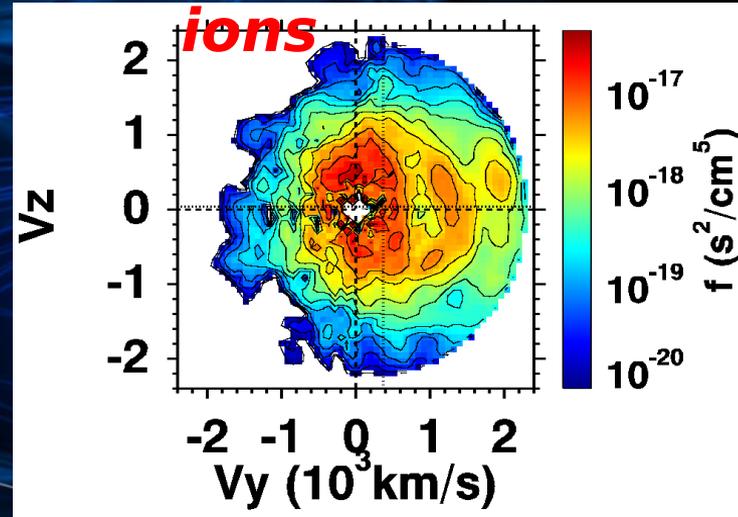
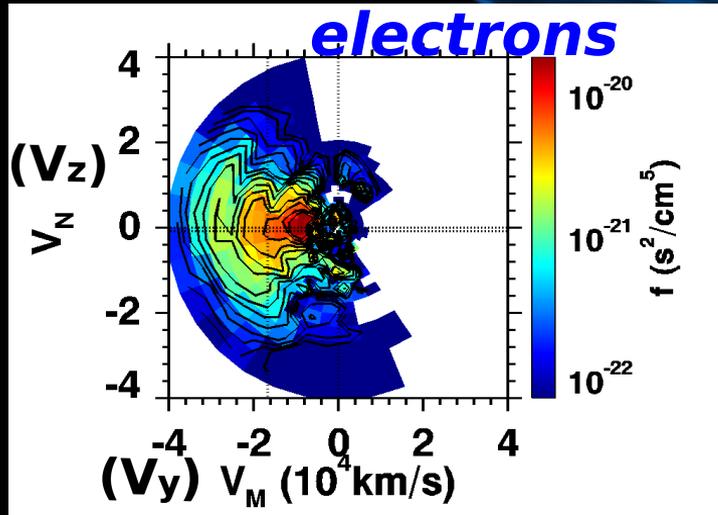


hhmm
2017 Jul 24





***In the diffusion region, acceleration
by E_{rec} can achieve: a few keV for e's ;
a few tens of keV for ions***

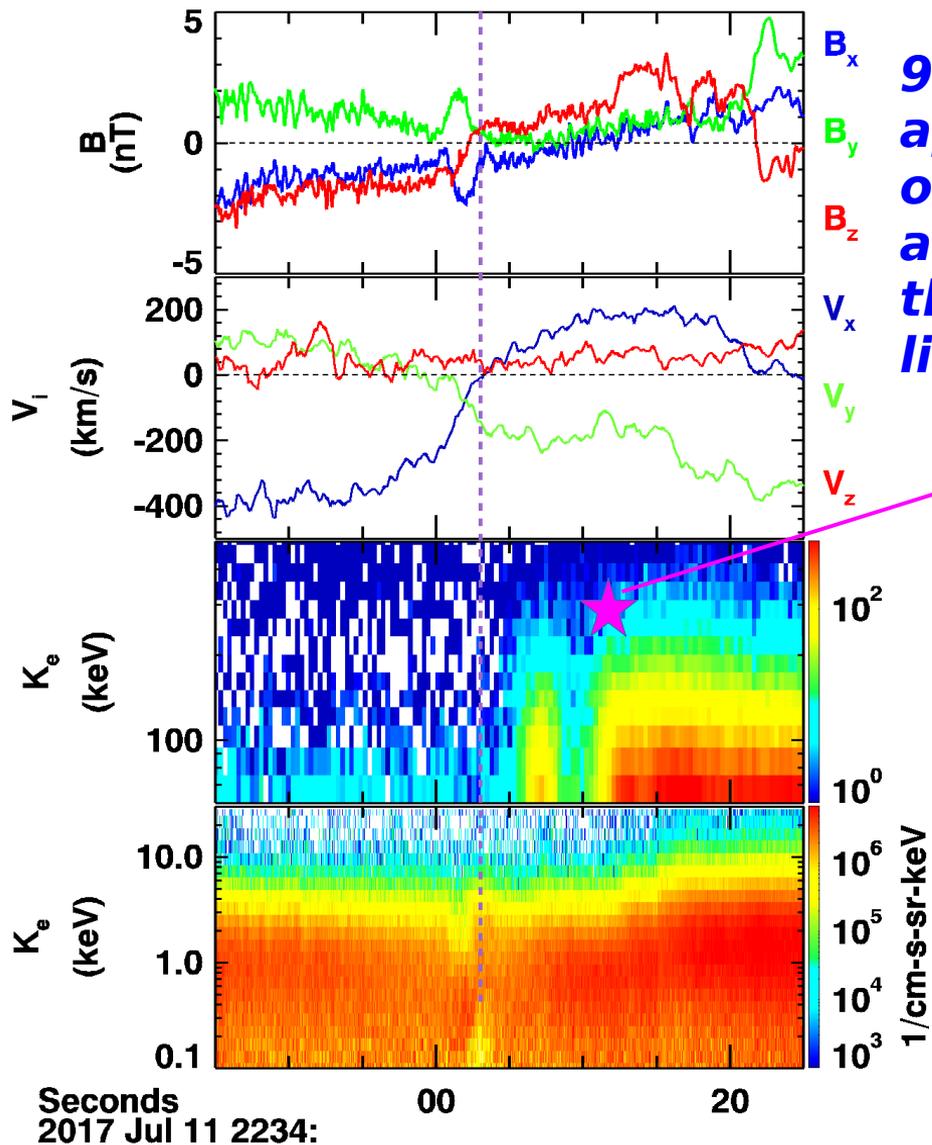


[Torbert et al., Science, 2018;
Bessho et al., GRL, 2018]

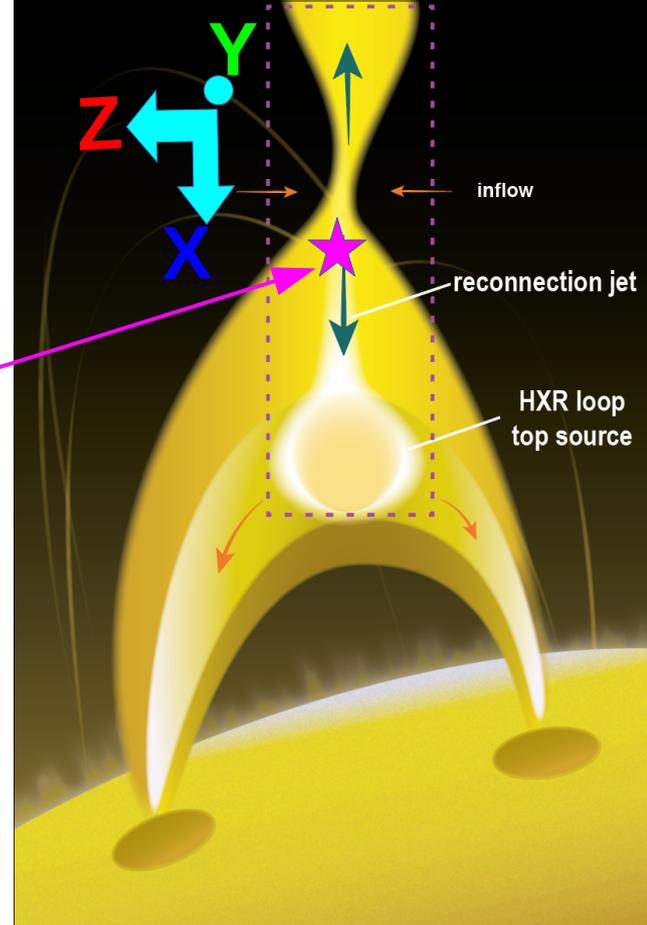
[Giles et al., GRL, 2021]

Sun/Earth





90 keV e's only appear Earthward of the X-line: key acc. physics is in the closed field line region

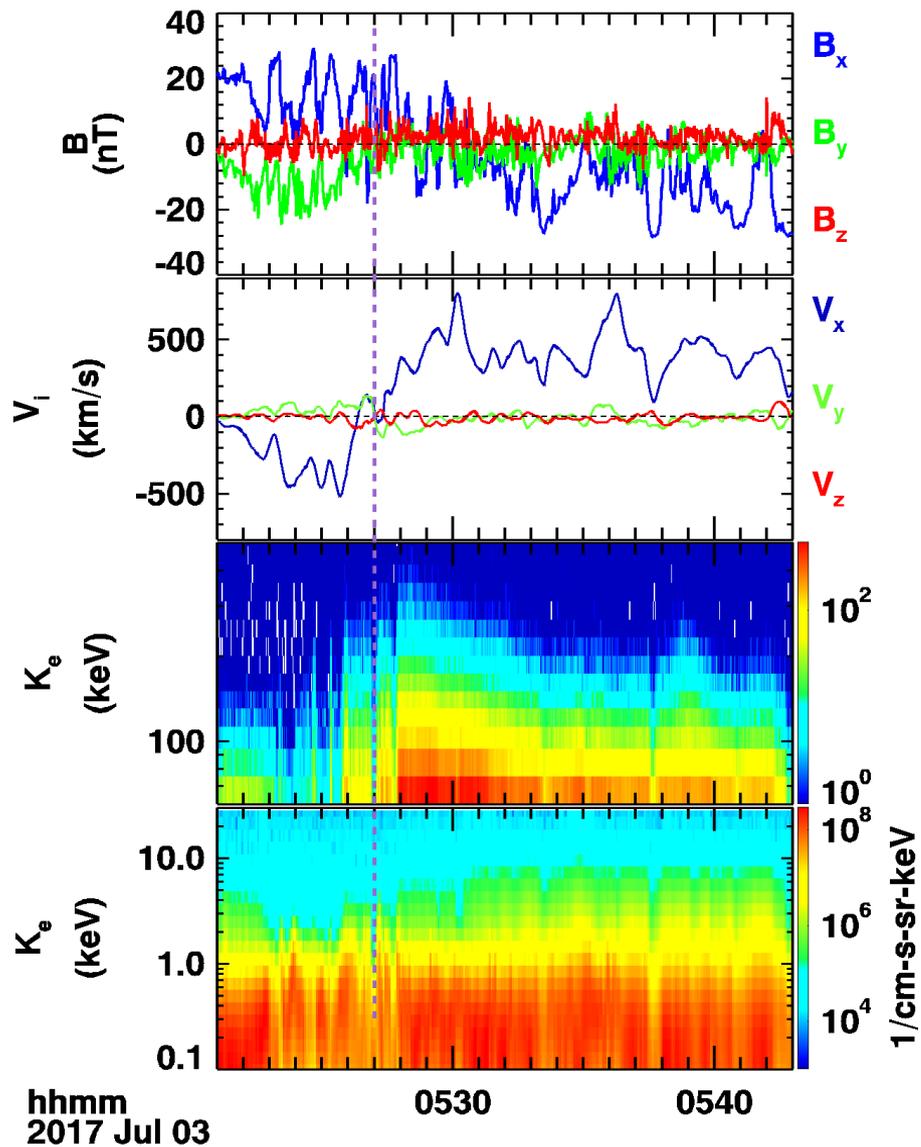


X-line (EDR) crossing in [Torbert et al., Science, 2018]

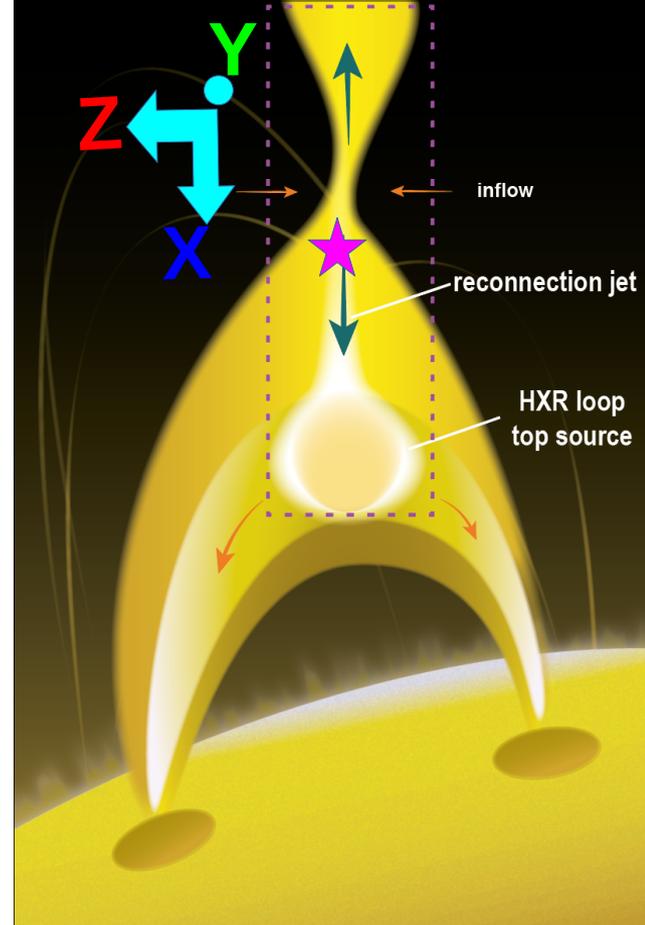
Seconds
2017 Jul 11 2234:

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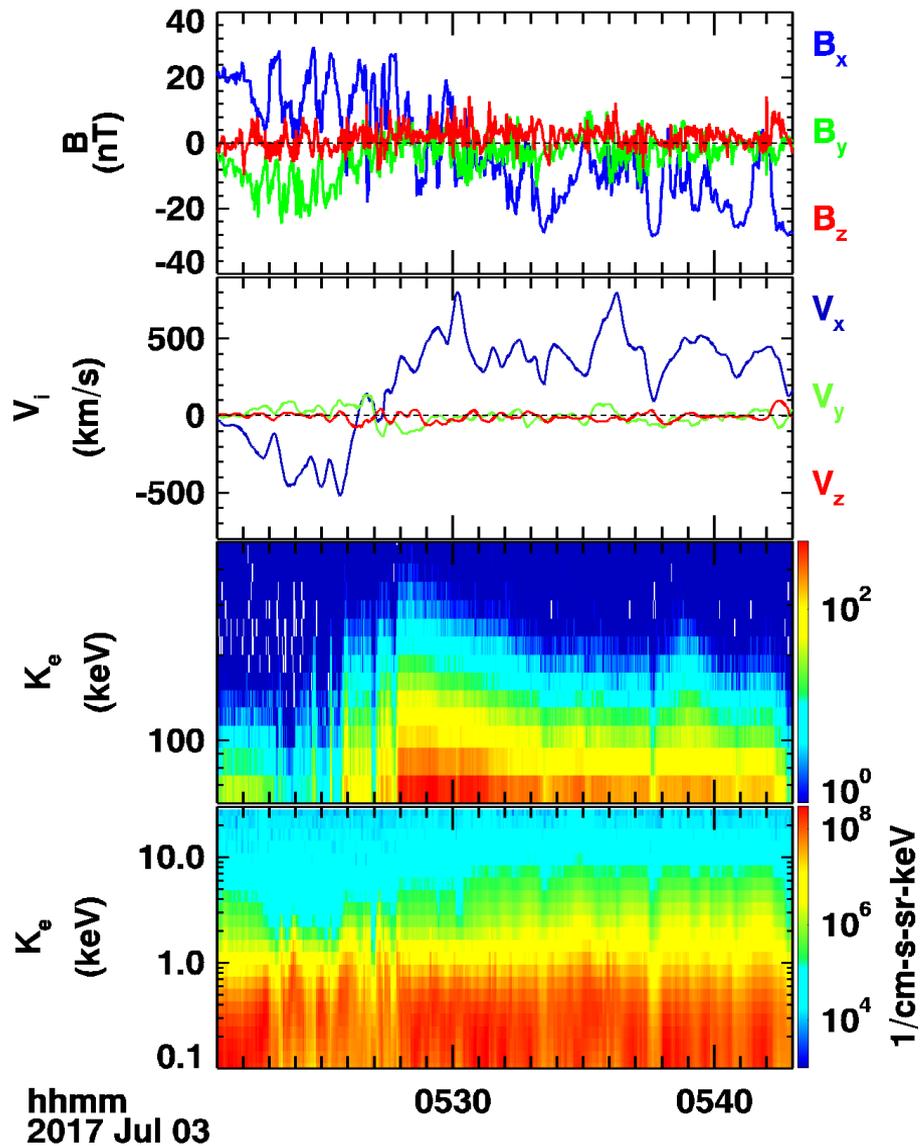
20



**90 keV e
Flux more
intense
Earthward
of the X-line:
a higher
density case
for spectra
analysis**

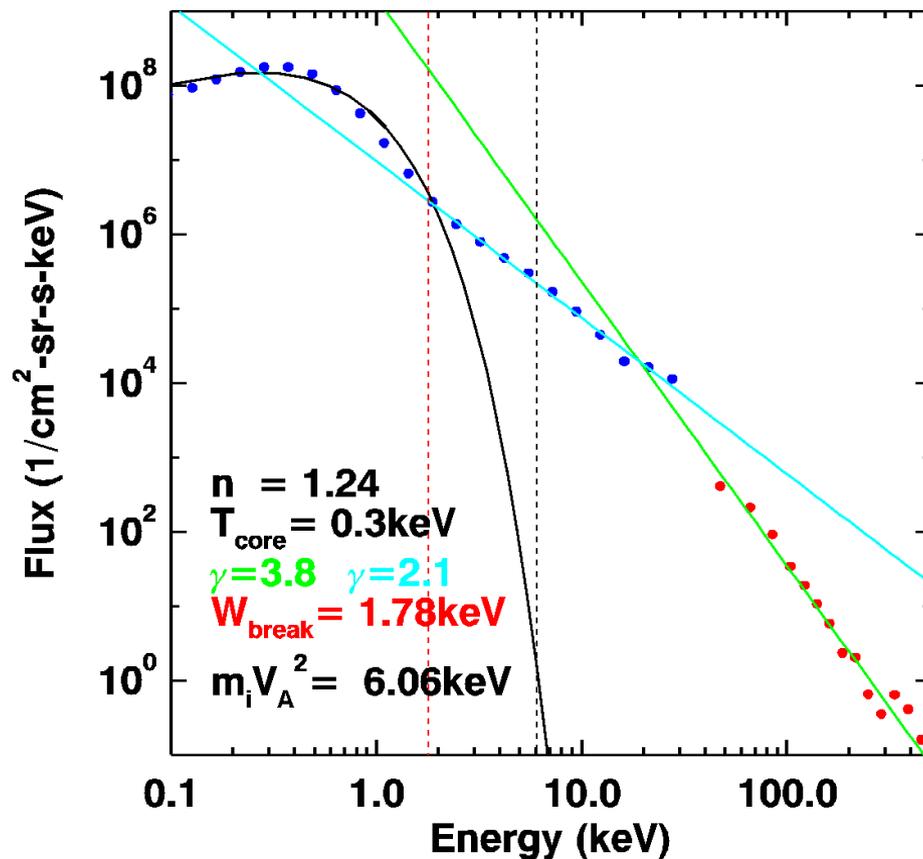


**X-line (3 EDRs) crossing in
[Chen et al., GRL, 2019]**

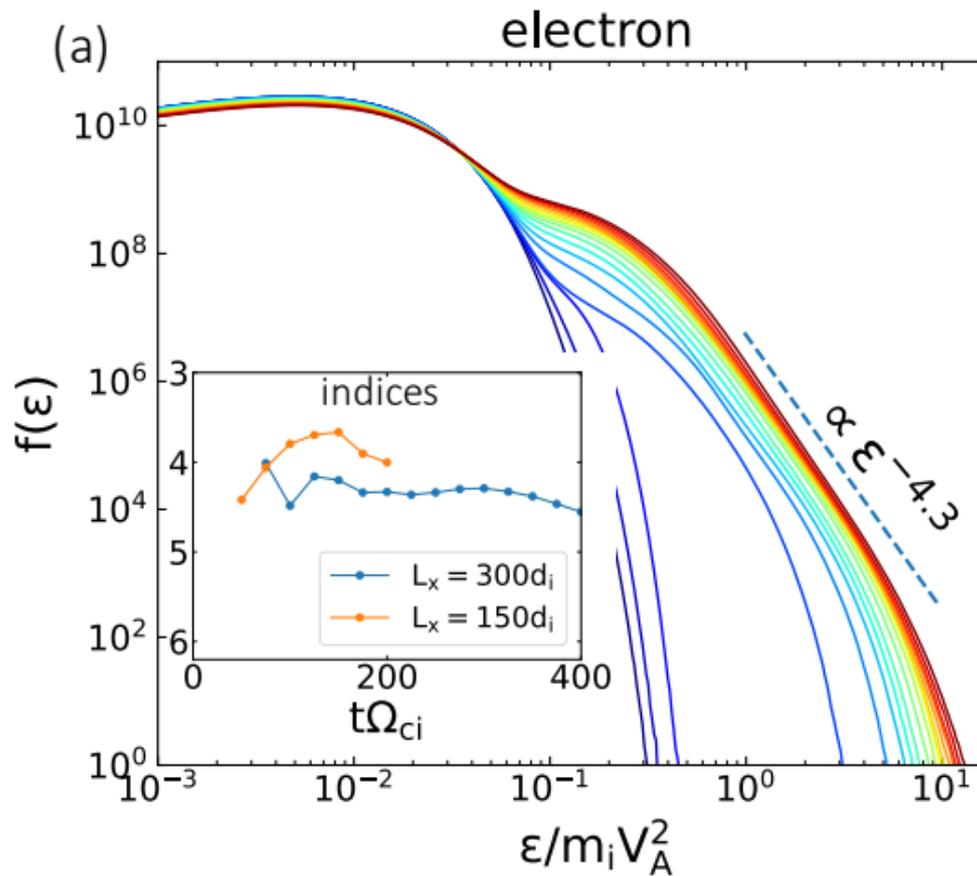


thermal+double power laws

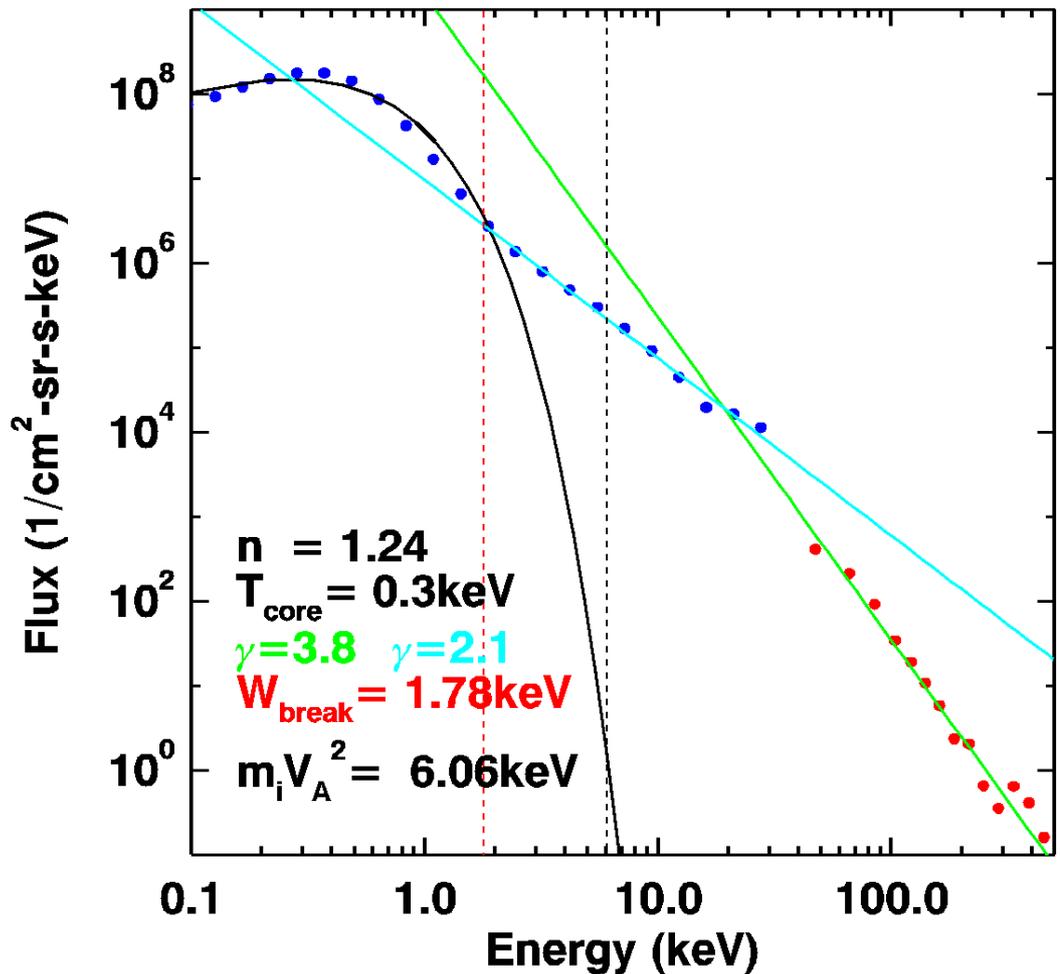
20170703/05:35:27



PIC with flux rope kinks [Qile Zhang, 2021]



Magnetotail reconnection outflow 20170703/05:35:27



Summary

- Suprathermal e's are primarily found in the earthward B pileup regions.
- Across the X line, earthward exhaust is favored.
- Statistical results do not support volume-filling islands as the dominant accelerators in the magnetotail.

If...

