# UNCOVERING THE HIDDEN POPULATION OF 

## SYMBIOTIC STARS

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## SYMBIOTIC STARS


$\star$ Interacting binaries
$\star$ Red giant + white dwarf

## Why are they important?

- Promising Type la supernova progenitor
- Enrichment of the interstellar medium with lithium


## SYMBIOTIC STARS

- Two phases of symbiotic stars: accreting-only and burning-type

Optical spectrum is dominated by the red giant and there are no or only very weak emission lines present

How can we distinguish them from single giants?

Strong nebular continuum and a rich emission line spectrum

- Observed excess light in the UV and X-rays
- Faint emission lines like $\mathrm{H} \alpha$ and $\mathrm{H} \beta$
- Flickering
- The IR excess


## GOALS AND METHODS

## RANDOM FOREST

- To characterize the entire population and quantify the number of burning versus accreting-only systems
- Separating symbiotic stars which contain AGB or RGB giant as a primary star.
- Supervised machine learning technique
- Builds decision trees on different samples and takes their majority vote for classification


## ADVANTAGES

$\rightarrow$ Highly accurate, robust, does not suffer the overfitting problem
$\rightarrow$ Can be used in classification problems
$\rightarrow$ Provides feature importance

## DATA AND SAMPLES

## Two groups: AGB/RGB and AGB/RGB/SYST

Color indices
(WISE, 2MASS)

- W1-W4
- Ks-W3
- H-W2
- W1-W2
- W3-W4
- J-W1
- J-H
- $\mathrm{H}-\mathrm{Ks}$

Color indices + lightcurves

- W1-W4
- Ks-W3
- H-W2
- W1-W2
- W3-W4
- J-W1
- J-H
- H-Ks
- Period
- Amplitude

Lightcurves
(ASAS-SN)

- Amplitude
- Period
- Kurtosis
- Skewness
- HI_ampl_ratio
- Shapiro_w
- Weighted_std


## FIRST RESULTS: AGB/RGB

## AGB/RGB CLASSIFICATION

- Best results when using color indices, amplitude and period
- Accuracy around 0.91
- Most important features:
- Ks-W3
- Period
- H-Ks
- W1-W2
- J-H



## FIRST RESULTS: AGB/RGB/SYST

- Best results when using color indices, amplitude and period
- Accuracy around 0.897
- Most important features:
- Ks-W3
- W3-W4
- W1-W4
- Period
- W1-W2



## FUTURE

We are exploring ways of connecting photometric detection of AGB stars with their spectra. If we are successful we will be able to search for AGB and RGB distinction based on their spectra from large surveys such as GALAH, Gaia-ESO and in the future 4MOST.


