

# NEWS

20<sup>th</sup> of May 2021

# Meeting with referees

Status report with referees last week:

This project, has now almost **50 collaborators**, from **8 Institutions** in **4 Countries**

Recent results were presented:

- electroluminescence, hydrocarbon and resistive foil FC R&D;
- saturation and signal simulations;
- studies on directionality sensitivity;
- progress on DAQ and trigger systems;
- test of GAS system and development of filtration stage;
- preliminary results AmBe-LIME analysis;

# Meeting with referees

An important part was devoted to LIME situation:

- performance with  $^{55}\text{Fe}$ ;
- hint about PMT issues and studies of fast signals;
- status and plans about the underground lab civil works;
- ongoing commissioning and plans for LIME characterisation before the installation at LNGS;

All activities on these different items were appreciated and the report had a positive feedback;

# Meeting with referees: documentation

INFN is trying to standardise the communication with experiments;

Collaborations will be required to use standard templates to submit CDR, TDR and periodical status reports to referees;

Our referees agreed that our CDR (<https://doi.org/10.15161/oar.it/73356>) is fine for this phase;

Two steps are needed:

1. submit a progress report by July;
2. submit a complete TDR describing the final demonstrator when final decisions will be taken about all details;

# Other documentation

In order to being able to write the TDR, we should know what will be the room available at LNGS;

To make a proper request:

- we should decide what we really want and what is financially possible;
- write a CDR++ (or TDR- -) to submit (exploiting the the experience gained with LIME);
- **tentative schedule:** CDR++: submit Feb-22, Approval June-22, TDR submit Sept/Dec-22

In the meanwhile we are finalising “the CYGNO paper” (thanks to Elisabetta and Giorgio!), soon to start the internal review; it will be useful as a basis for the general part of the CDR++;

# Simulation: LIME shield

Flaminia is working at LIME Shield simulation: can we use Opera Lead to make 5 cm Cu + 5 cm Pb?

## Background rates

PRELIMINARY

- 40cm water + 10cm copper

External gammas ( $0.56 \text{ cm}^{-2} \text{ s}^{-1}$ )

(ER):  $(2.05 \pm 0.10) \times 10^5$  events/yr (total)

(ER):  $(4.0 \pm 0.4) \times 10^4$  events/yr (0-20 keV)

External neutrons ( $2.7 \times 10^{-6} \text{ cm}^{-2} \text{ s}^{-1}$ )

(ER+NR):  $(23.6 \pm 0.6)$  events/yr (total)

(ER+NR):  $(9.9 \pm 0.3)$  events/yr (0-20 keV)

(Only NR):  $(4.7 \pm 0.1)$  events/yr (total)

(Only NR):  $(3.2 \pm 0.1)$  events/yr (0-20 keV)

With Cu:  $4.0 \times 10^4$  ER/yr (0-20keV)

With Pb:  $4.5 \times 10^4$  ER/yr (0-20keV)

Internal background:

$0.5\text{-}1.0 \times 10^5$  ER/yr (0-20keV)

- 40cm water + 5cm lead + 5cm copper

External gammas ( $0.56 \text{ cm}^{-2} \text{ s}^{-1}$ )

(ER):  $(1.8 \pm 0.5) \times 10^5$  events/yr (total)

(ER):  $(3.9 \pm 2.2) \times 10^4$  events/yr (0-20 keV)

External neutrons ( $2.7 \times 10^{-6} \text{ cm}^{-2} \text{ s}^{-1}$ )

(ER+NR):  $(30.1 \pm 0.3)$  events/yr (total)

(ER+NR):  $(12.1 \pm 0.2)$  events/yr (0-20 keV)

(Only NR):  $(4.38 \pm 0.09)$  events/yr (total)

(Only NR):  $(2.66 \pm 0.07)$  events/yr (0-20 keV)

Radioactivity from  $^{210}\text{Bi}$ \*:

(ER):  $(2.47 \pm 0.05) \times 10^4$  events/yr (total)

(ER):  $(6.2 \pm 0.3) \times 10^3$  events/yr (0-20 keV)

\*( $80 \text{ Bq/kg} \times 5658.7 \text{ kg}$  of lead = 452693 Bq)

If confirmed difference in background rate between copper and lead configurations is small – using lead might be a good choice

To be cross-checked

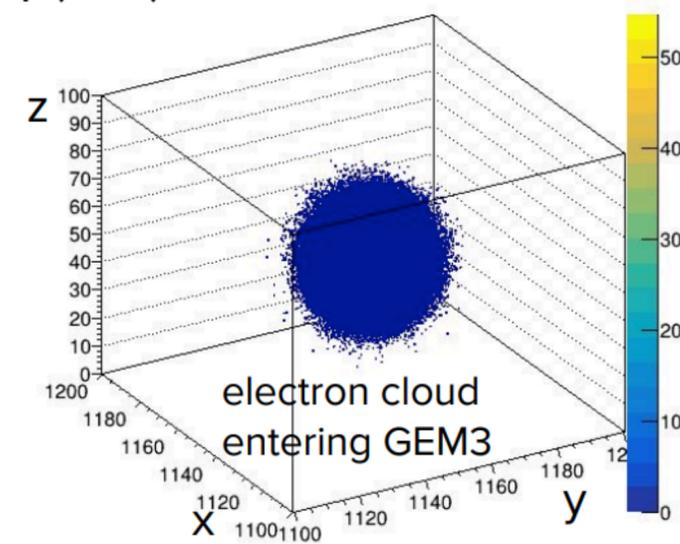
# Simulation: Saturation

Method:

- Add absorption length parameter for e- in the gas  $\lambda$ :  $n = n_0 \exp(-z/\lambda)$
- Only GEM3 saturated, G1 and G2 simulated as before
- Simulate the 3D cloud of electrons entering GEM3:
  - spatial smearing given by  $\sigma_{OT}, \sigma_T$  and  $\sigma_{OL}, \sigma_L$  and drift distance  $z$
  - divide electron cloud in voxels  $160(x) \times 160(y) \times 100(z) \mu\text{m}^3$
  - apply formula of saturated gain in each voxel

$$G = A \frac{g}{1 + \frac{n}{n_h}(g - 1)}$$

- Conversion to number of photons as before



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Giulia introduced:

- electron absorption in gas;
- gain saturation on the last GEM;
- behavior of GEM gain and  $\epsilon = \epsilon_{extr} \times \epsilon_{coll}$

