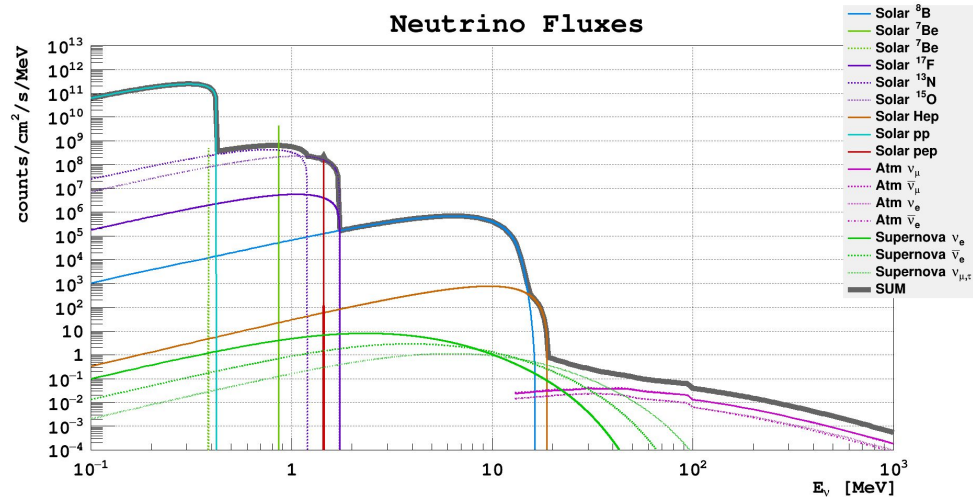


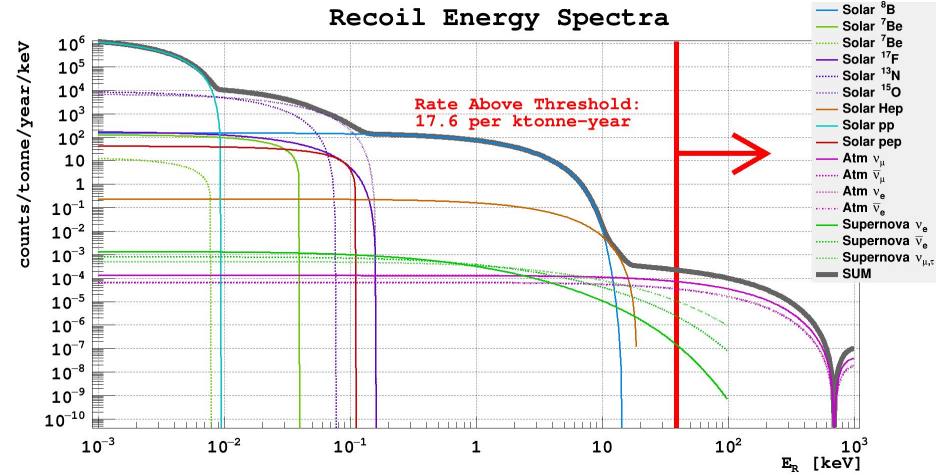
CEvNS Backgrounds

- **Problem:** CEvNS from atmospheric neutrinos present an irreducible background
 - Fortunately, they will be rare in DS-20k, but nonzero
- **People:**
 - Shawn, Matteo, Emmanuele, Ellen?
- **Procedure:**
 - The CEvNS background rate is calculable from the the neutrino flux and cross section
- **Mitigation strategy:**
 - It's irreducible, so we can't mitigate it
 - Background subtraction, PLR, and energy cuts are all possible avenues for mitigating their impact on sensitivity



Current status and results

- Above $40 \text{ keV}_{\text{nr}}$, we expect 17.6 events/ktonne/year
 - 1.7 events expected in 100 tonne year exposure
- Systematic uncertainty of $\sim 20\%$ from uncertainty on the atmospheric neutrino flux



Prospects and plans

- Problems:
 - Irreducible background
- Uncertainties:
 - There's a ~20% uncertainty on the atmospheric neutrino flux
- Plans:
 - The calculation is largely done, module cross-checks
 - It's worth further developing infrastructure for studying and mitigating their impact on sensitivity, like improving PLR code, energy range of interest