Supported by U.S. DOE: Award DE-SC0015903

## **Most Recent Results** from T2K: Is the Mirror Cracking? Les Rencontres de Physique de la Vallée d'Aoste 2021-03-12 Luke Pickering for the T2K Pronouns: He/Him/His MICHIGAN STATE Collaboration



## Big Picture Neutrino Oscillation Questions<sup>2</sup>

What is the mass ordering of the neutrino mass states?

What are the precise values of the neutrino oscillation parameters? Is there significant CP violation in the neutrino sector?

Could neutrino sector CP violation explain the matter/antimatter asymmetry?





## Big Picture Neutrino Oscillation Questions

What is the mass ordering of the neutrino mass states?

What are the precise values of the neutrino oscillation parameters?

#### Experiment

Is there significant CP violation in the neutrino sector?

Could neutrino sector CP violation explain the matter/antimatter asymmetry?

Theory





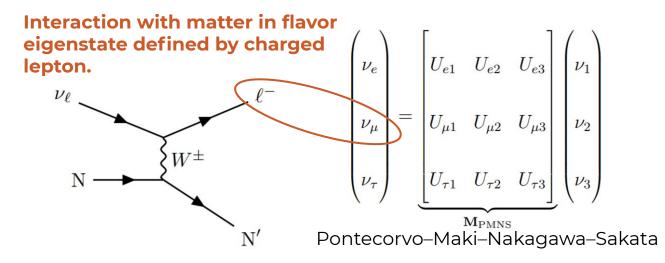


## Neutrino Oscillations: A Brief Introduction





#### **Neutrino Oscillation: PMNS**

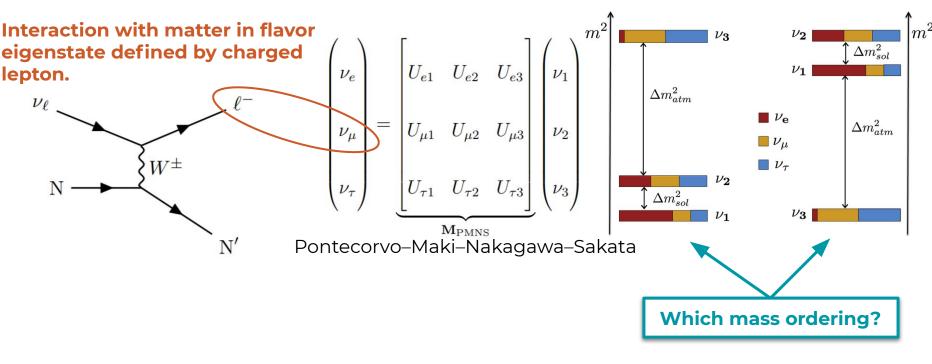






Journal of Physics G: Nuclear and Particle Physics. 43. 10.1088/0954-3899/43/8/084001

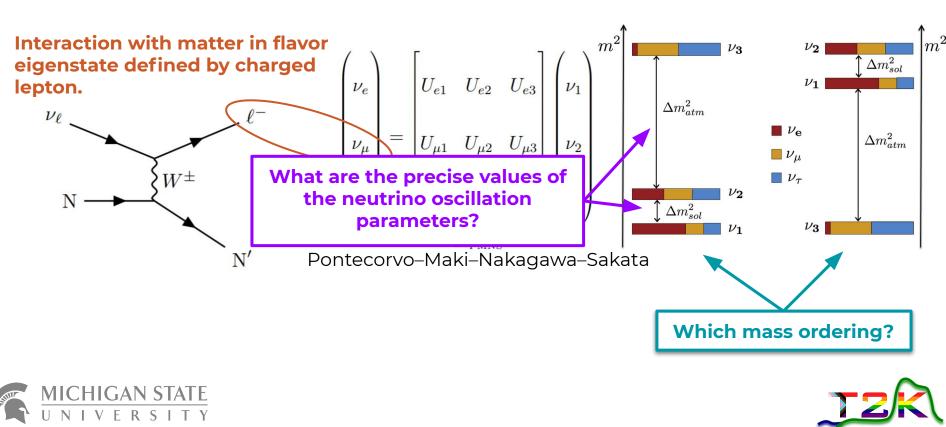
#### **Neutrino Oscillation: PMNS**



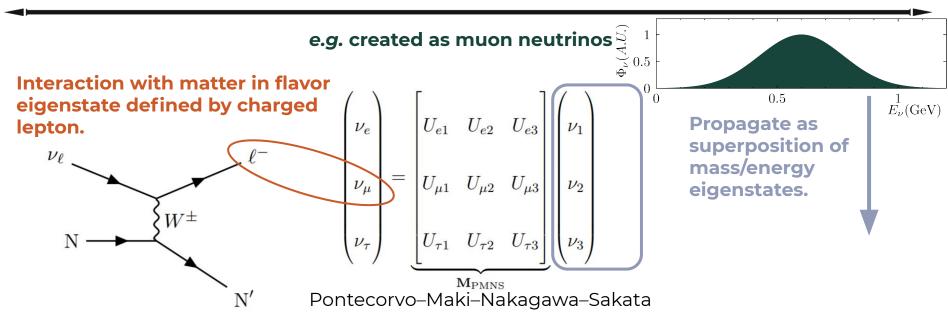
TZK

#### **Neutrino Oscillation: PMNS**

Journal of Physics G: Nuclear and Particle Physics. 43. 10.1088/0954-3899/43/8/084001



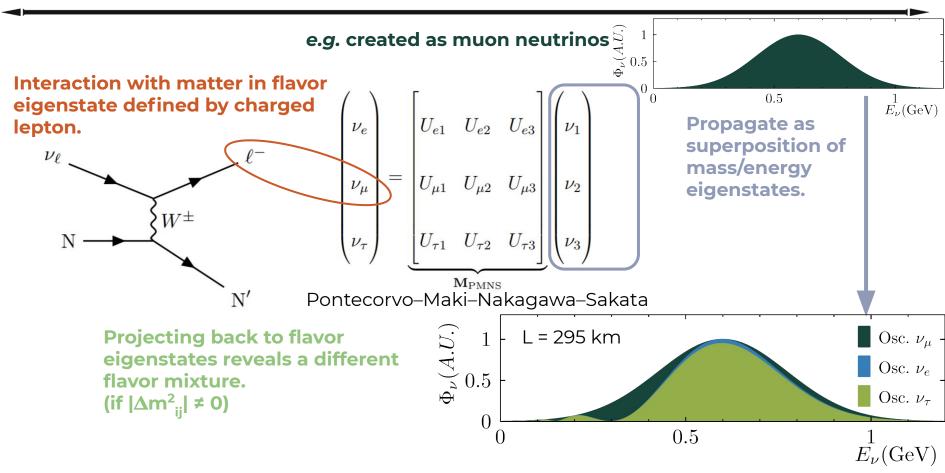
#### **Neutrino Oscillation: PMNS**







#### **Neutrino Oscillation: PMNS**



#### **Re-parameterizing the PMNS**

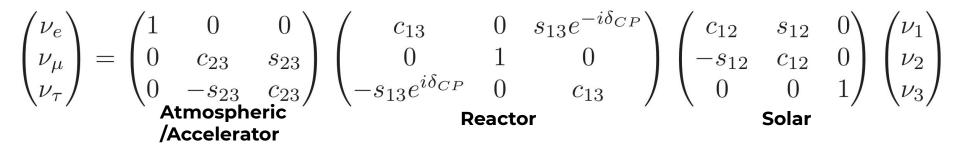
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$
Atmospheric Accelerator Reactor Solar

- Unitarity lets us re-parameterize PMNS matrix in terms of:
  - Three mixing angles:  $C_{ij} = cos(\theta_{ij})$
  - CP violating phase:  $0 < \delta_{CP}^{"} < 2\pi$ , or  $-\pi < \delta_{CP} < \pi$

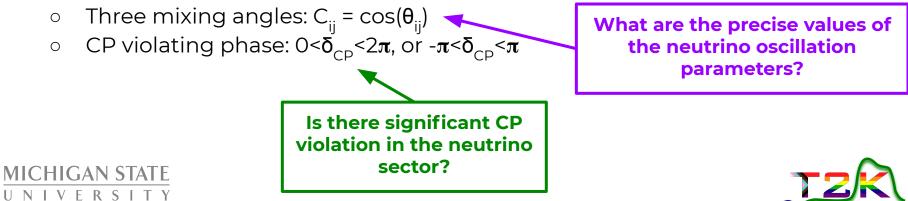




#### **Re-parameterizing the PMNS**



• Unitarity lets us re-parameterize PMNS matrix in terms of:







muon: neutrino

#### **Measuring Oscillations**

• Long baseline experiments study two oscillation channels:

#### Muon neutrino disappearance

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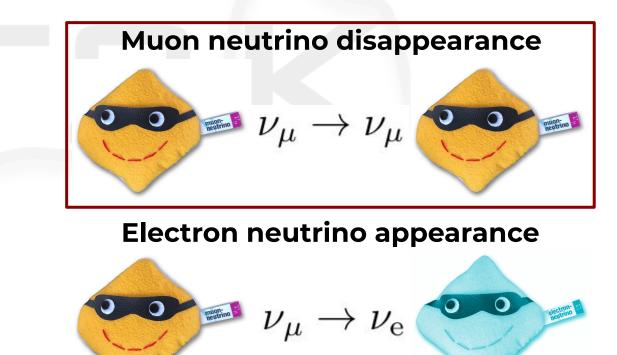
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#### **Electron neutrino appearance**



#### Measuring Oscillations

• Long baseline experiments study two oscillation channels:







## **Muon Neutrino Disappearance**

 To leading order, muon neutrino survival probability depends on mixing angles, and mass-squared splittings.

$$P(\nu_{\mu} \rightarrow \nu_{\mu}) \simeq 1 - 4\cos^{2} \theta_{13} \sin^{2} \theta_{23}$$
$$\times [1 - \cos^{2} \theta_{13} \sin^{2} \theta_{23}] \sin^{2} \frac{\Delta m_{32}^{2} L}{4E}$$
$$+ (\text{solar, matter effect terms})$$





## **Muon Neutrino Disappearance**

0.

0.

- To leading order, muon neutrino survival probability depends on mixing angles, and mass-squared splittings.
- Choose L/E for maximum effect:  $\sin^2\left(\Delta m_{23}^2 L/4E\right) \simeq 1$

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ttings.  

$$P(\nu_{\mu} \rightarrow \nu_{\mu}) \simeq 1 - 4\cos^{2} \theta_{13} \sin^{2} \theta_{23}$$

$$\times [1 - \cos^{2} \theta_{13} \sin^{2} \theta_{23}] \sin^{2} \frac{\Delta m_{32}^{2} L}{4E}$$

$$+ (\text{solar, matter effect terms})$$

$$\int_{1}^{3} \uparrow_{1}^{0} 0.8$$

$$\int_{0}^{3} 0.6$$

$$0.4$$

$$0.2$$

$$0$$

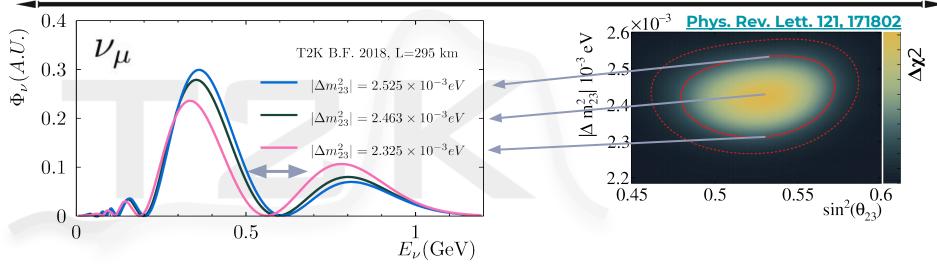
$$0.5$$

$$I$$

$$E_{\nu}(\text{GeV})$$

## **Disappearance Sensitivity**

L. Pickering **17 68% Confidence Level 90%** 



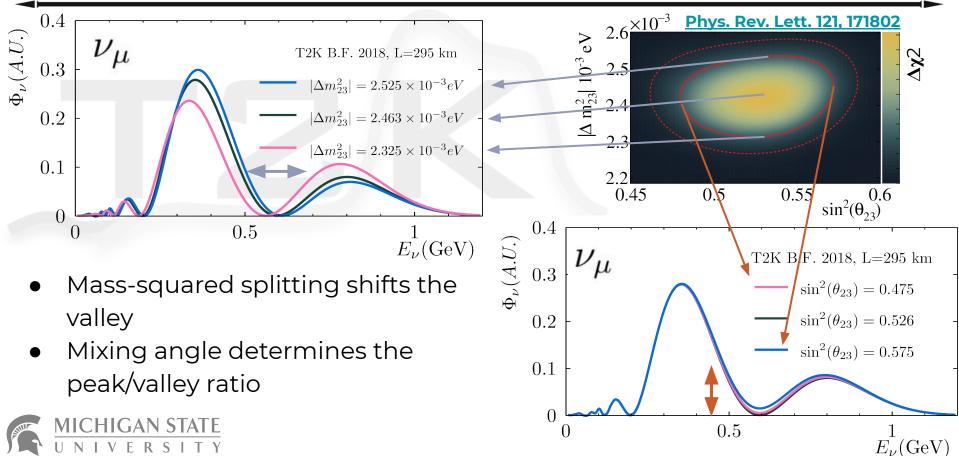
• Mass-squared splitting shifts the valley





## **Disappearance Sensitivity**

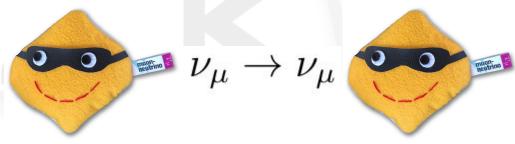
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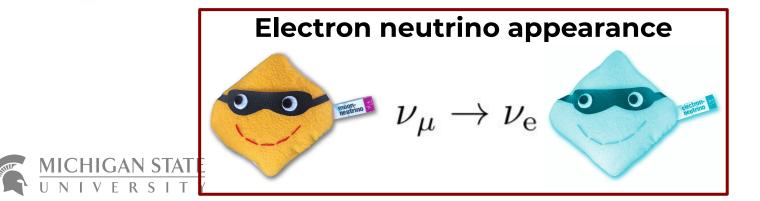


#### Measuring Oscillations

• Long baseline experiments study two oscillation channels:

#### Muon neutrino disappearance





## **Electron Neutrino Appearance**

- Electron neutrino appearance probability has 'CP odd' term.
  - Sign flip between matter and antimatter.

 $P(\stackrel{(\leftarrow)}{\nu_{\mu}} \rightarrow \stackrel{(\leftarrow)}{\nu_{e}}) \simeq \frac{\sin^{2} \theta_{23} \sin^{2} 2\theta_{13} \sin^{2} \frac{\Delta m_{32}^{2} L}{4E}}{(+) - \left[ \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \right] \times \sin \frac{\Delta m_{21}^{2} L}{4E} \sin^{2} \frac{\Delta m_{32}^{2} L}{4E} \sin \delta_{CP} \right] + (CP-even, solar, matter effect terms)$ 

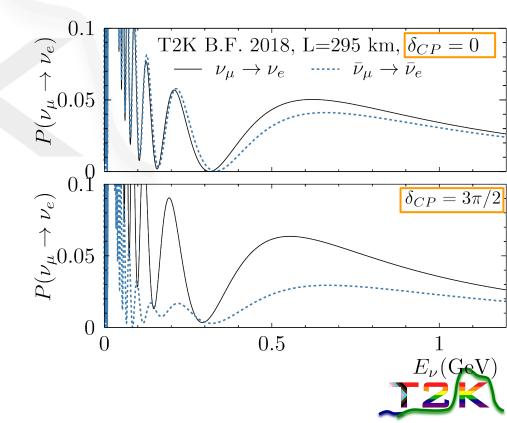




## **Electron Neutrino Appearance**

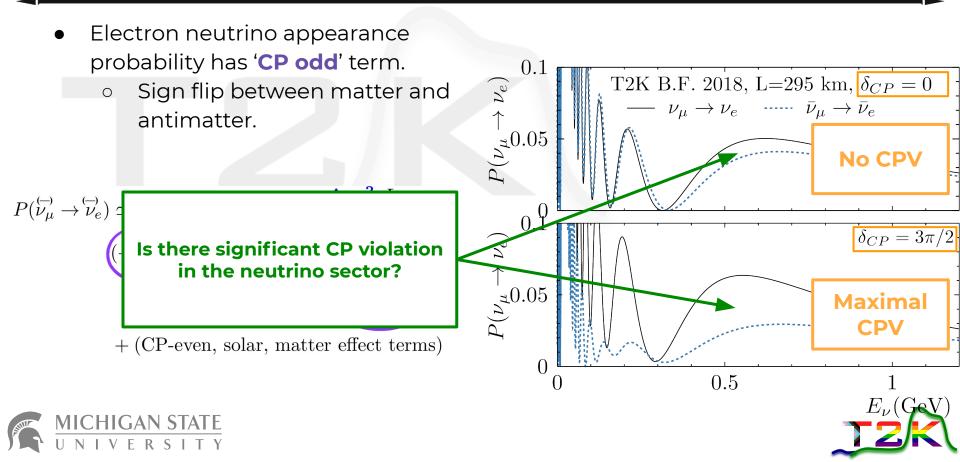
- Electron neutrino appearance probability has 'CP odd' term.
  - Sign flip between matter and antimatter.

$$\begin{split} P(\stackrel{(\rightarrow)}{\nu_{\mu}} \rightarrow \stackrel{(\rightarrow)}{\nu_{e}}) \simeq \sin^{2}\theta_{23}\sin^{2}2\theta_{13}\sin^{2}\frac{\Delta m_{32}^{2}L}{4E} \\ (+)-\left[\sin 2\theta_{12}\sin 2\theta_{23}\sin 2\theta_{13}\cos \theta_{13}\right] \\ \times \sin \frac{\Delta m_{21}^{2}L}{4E}\sin^{2}\frac{\Delta m_{32}^{2}L}{4E}\sin \delta_{CP}\right] \\ + (\text{CP-even, solar, matter effect terms}) \end{split}$$

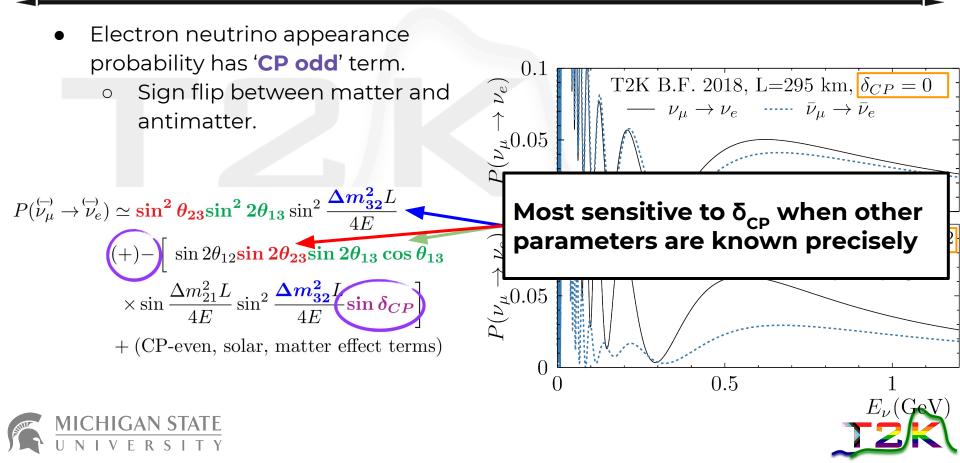


22

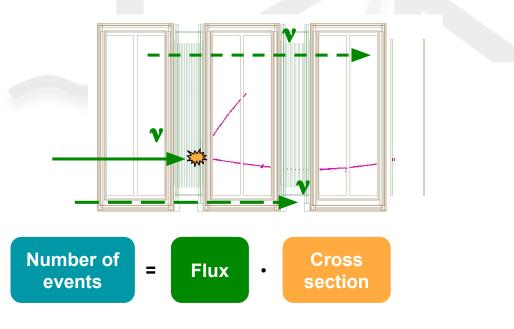
## **Electron Neutrino Appearance**



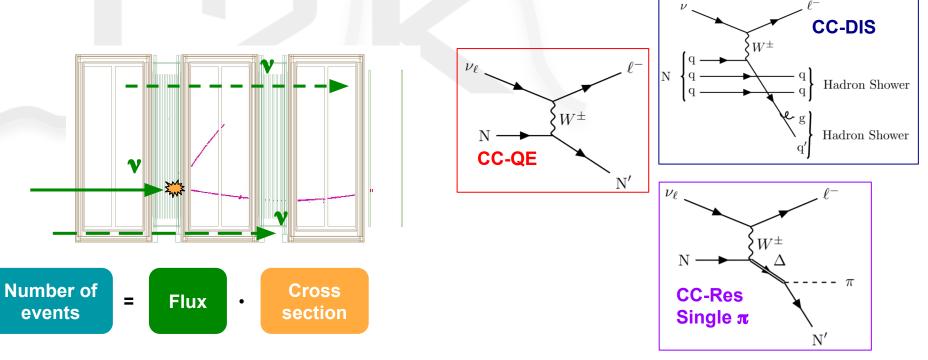
## **Electron Neutrino Appearance**

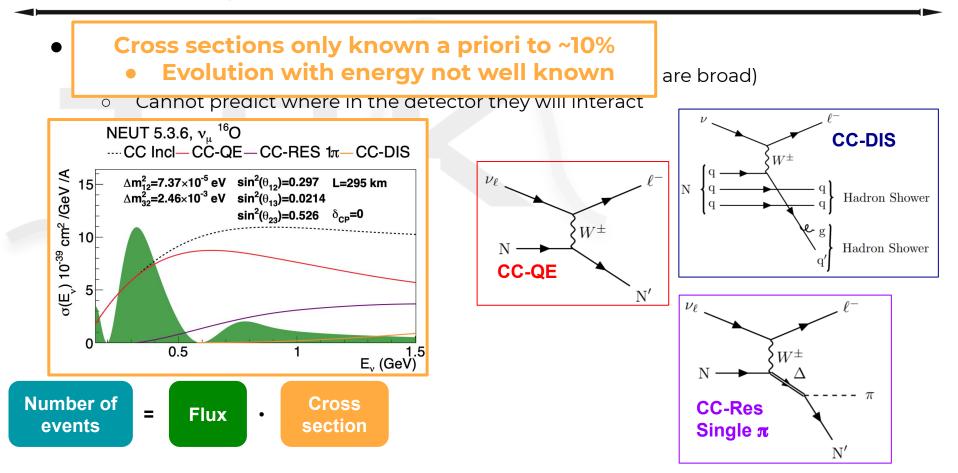


- Measuring oscillation is hard because:
  - Do not know incoming particle energy (neutrino beams are broad)
  - Cannot predict where in the detector they will interact

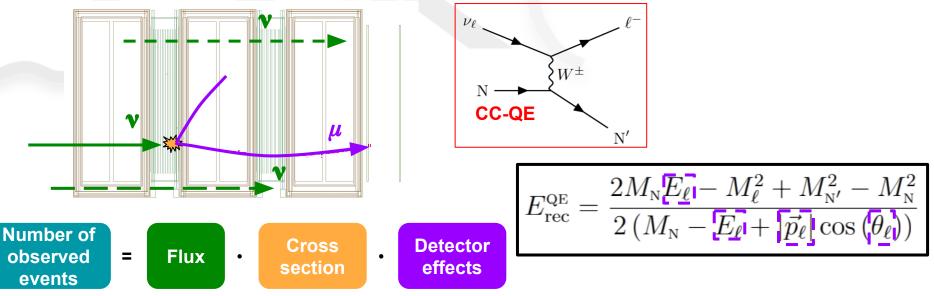


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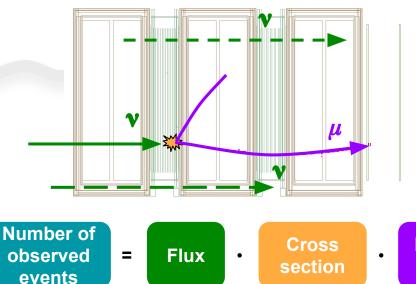


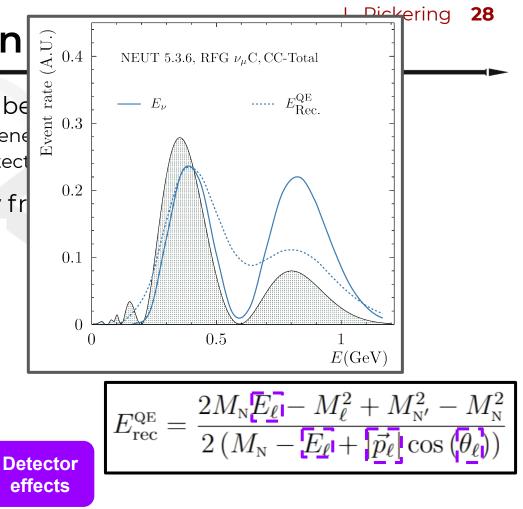


- Measuring oscillation is hard because:
  - Do not know incoming particle energy (neutrino beams are broad)
  - Cannot predict where in the detector they will interact
- Reconstruct neutrino energy from visible final state particles



- Measuring oscillation is hard be
  - Do not know incoming particle ene
  - Cannot predict where in the detect
- Reconstruct neutrino energy fr





# The T2R Collaboration L. Pickering 29







# The T2R Collaboration <sup>L. Pickering 30</sup>

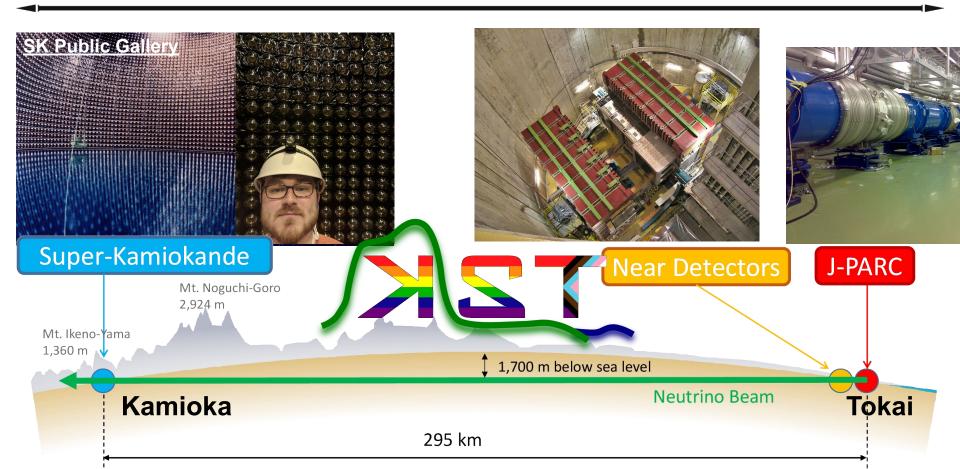


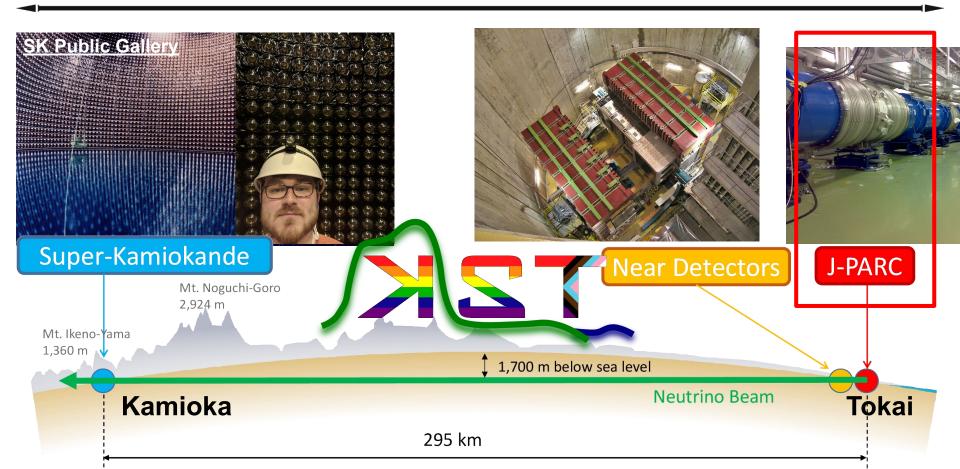










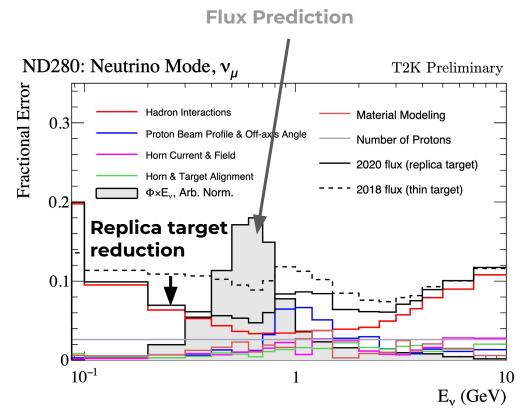


## **J-PARC Beam: Uncertainties**

- Can run the beam in neutrino mode or antineutrino mode.
- Flux uncertainties dominated by

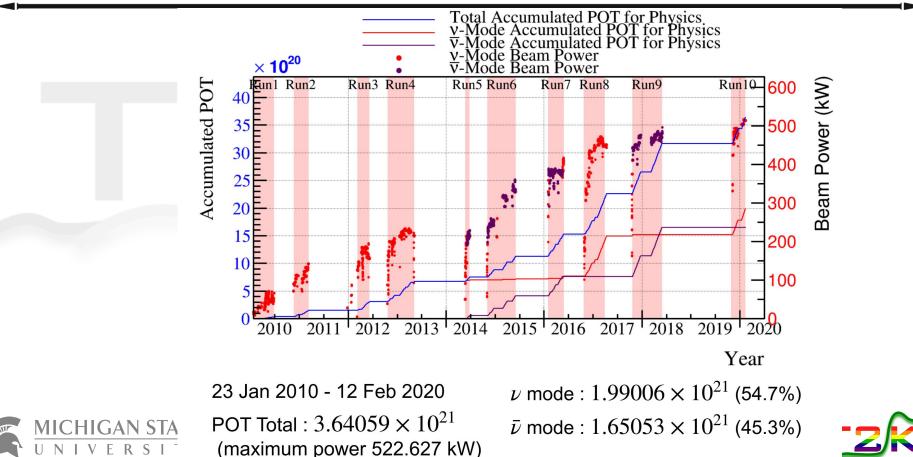
#### hadron-production:

- Simulation tuned to NA61/SHINE hadron-production data.
- Incorporation of NA61/SHINE data taken with a **replica** T2K target reduced uncertainties by factor ~2.



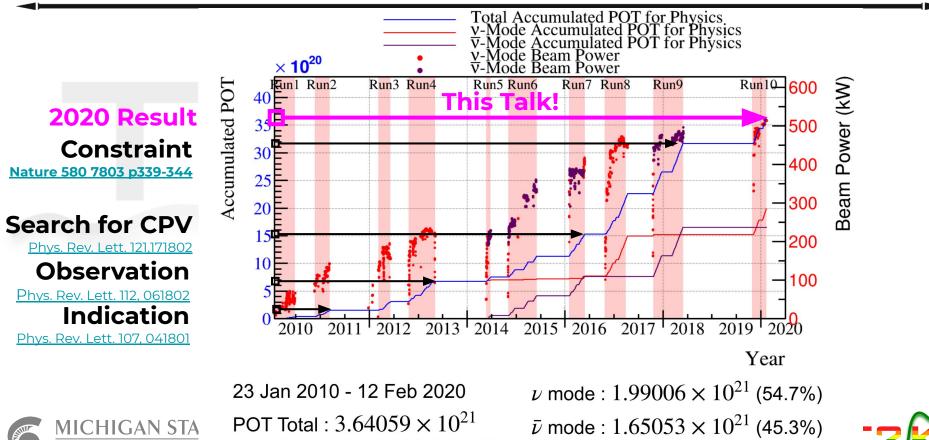
#### T2K Exposure





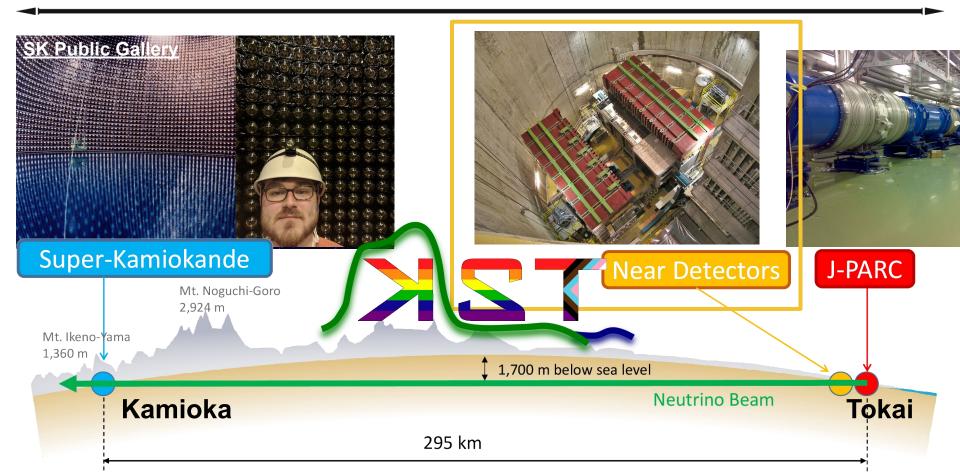
#### T2K Exposure





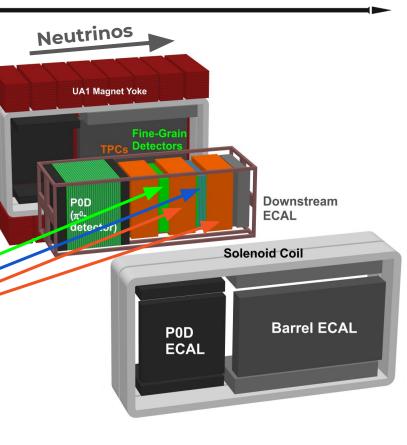
(maximum power 522.627 kW)

## Tokai To Kamioka



### **T2K Near Detector**

- Magnetized: Charge and momentum measurements
- Constrain 'wrong sign' backgrounds
   μ<sup>+</sup> in neutrino mode, μ<sup>-</sup> in antineutrino mode
- FGD used as the neutrino target:
  - Active **CH** target + passive **water** target.
- Time Projection Chambers:
  - Good momentum/PID for charged final state particles.



# T2K Near Detector Measurement Programme

#### **Charged-Current 0 Pions**

- v<sub>µ</sub> CCQE C<sup>12</sup> (2014)
- ν<sub>u</sub> CCOπ C<sup>12</sup>H (2016)
- $v_{\mu} CC0\pi H_2 O^{16}$  (2017)
- **ν** CCOπ C<sup>12</sup>H (2018)
- NC 1γ C<sup>12</sup>H (2019)
- $\bar{v}_{\mu}$  CCO $\pi$  H<sub>2</sub>O<sup>16</sup> (2019)
- $v_{\mu}^{\prime}/v_{\mu}^{-}$  CC0 $\pi$  C<sup>12</sup>H (2020)
- $v_{\mu} CC0\pi C^{12}H/H_2O^{16}$  (2020)

#### **Charged-Current 1 Pion**

- $v_{\mu} CC1\pi H_2O^{16}$  (2016)
- **ν** CC1π C<sup>12</sup>H (2019)
- $\dot{C}C1\pi$  Transverse Imbalance (2021)

New since La Thuile 2019

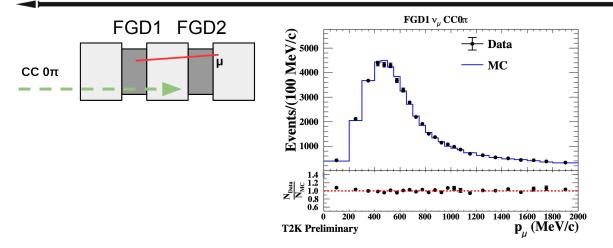
#### **Other Cross-sections**

- $v_{\mu}$  CCInc C<sup>12</sup> (2013)
- **v**NCQE O<sup>16</sup> (2014)
- v<sub>e</sub>CCInc C<sup>12</sup> (2014)
- $v_{\mu}$  CC Coherent  $1\pi$  C<sup>12</sup> (2017)
- $v_{\mu}$  CCInc C<sup>12</sup> (2018)
- v CCInc POD (2018)
- $v_e / v_e$  CCInc C<sup>12</sup> (2020)

#### Exotic/BSM

• Search for Heavy Neutrinos (2019)

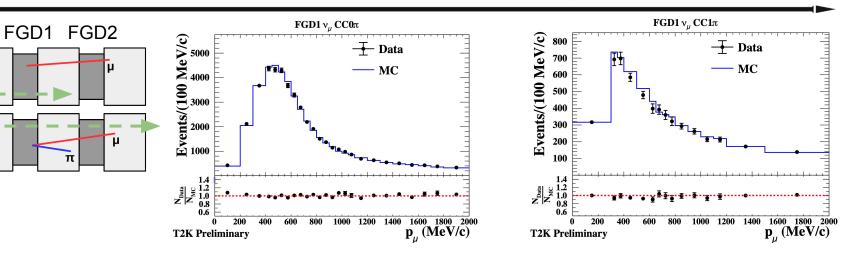
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L. Pickering

41

- Near detector samples separated by:
  - Reconstructed pion multiplicity: N=0



L. Pickering

42

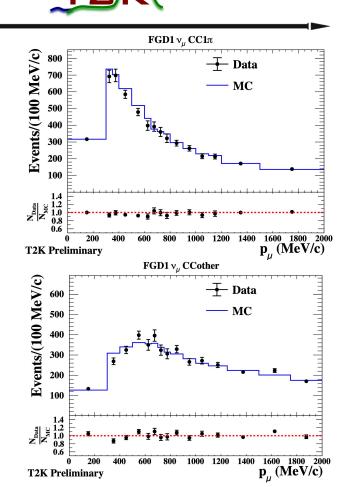
• Near detector samples separated by:

CC 0π

CC 1π

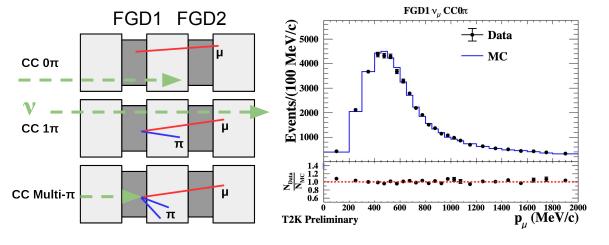
ν

Reconstructed pion multiplicity: N=0, 1



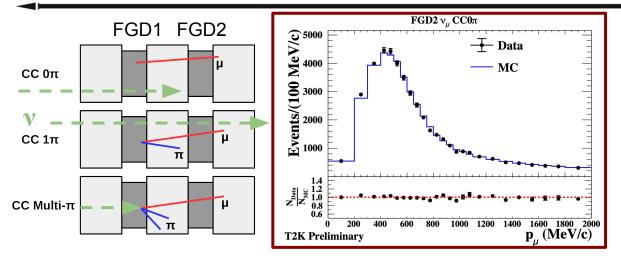
L. Pickering

43



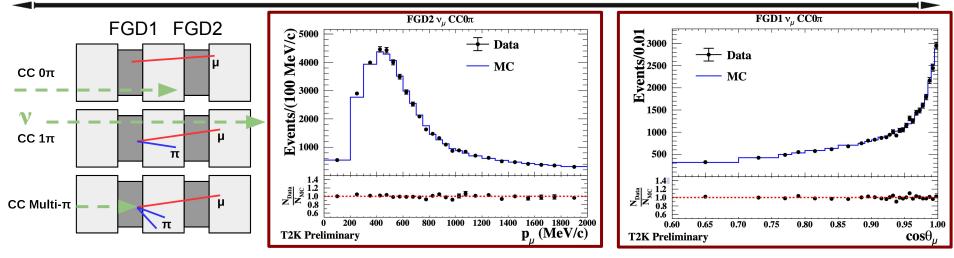
- Near detector samples separated by:
  - Reconstructed pion multiplicity: N=0, 1, >1





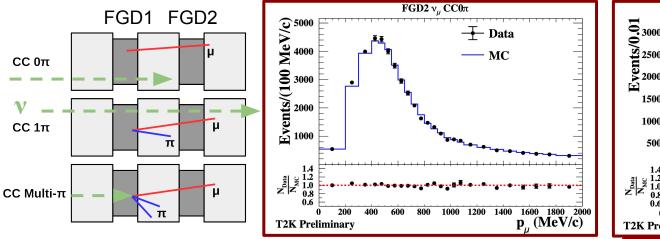
- Near detector samples separated by:
  - Reconstructed pion multiplicity: N=0, 1, >1
  - Detector material: CH (FGD) or CH+H2O (FGD2)



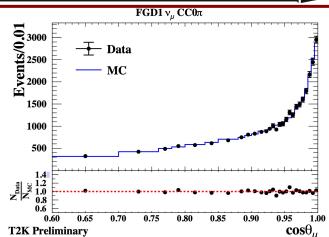


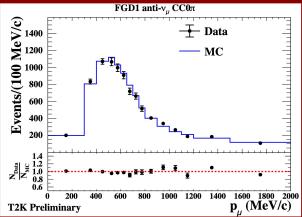
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- Binned in observed lepton kinematics only.



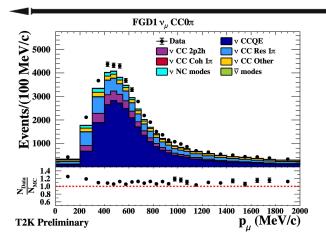


- Near detector samples separated by:
  - Reconstructed pion multiplicity: N=0, 1, >1
  - Detector material: CH (FGD) or CH+H2O (FGD2)
- Binned in **observed lepton kinematics** only.
- Both neutrino and antineutrino beam modes.



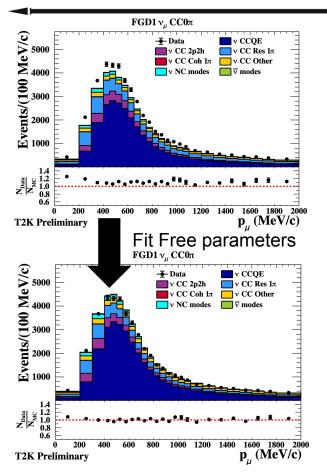


#### Near Detector Fit



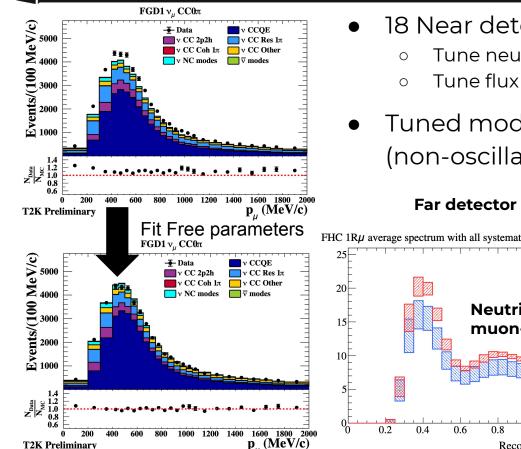
• 18 Near detector samples:

## Near Detector Fit



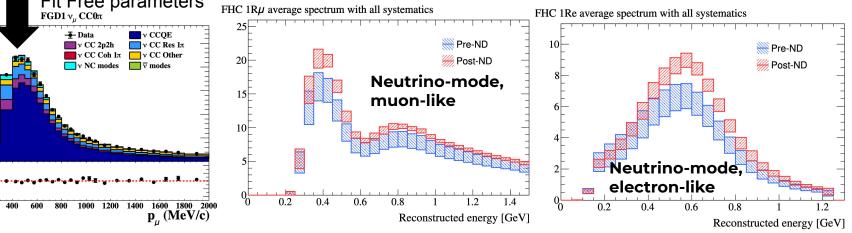
- 18 Near detector samples:
  - Tune neutrino interaction model
  - Tune flux prediction

## Near Detector Fit



- 18 Near detector samples:
  - Tune neutrino interaction model
  - Tune flux prediction
- Tuned model constrains event rate (non-oscillation) uncertainties for Far detector

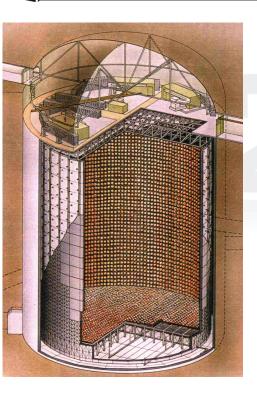
#### Far detector predicted event rates with oscillations



## Tokai To Kamioka



#### Super-Kamiokande

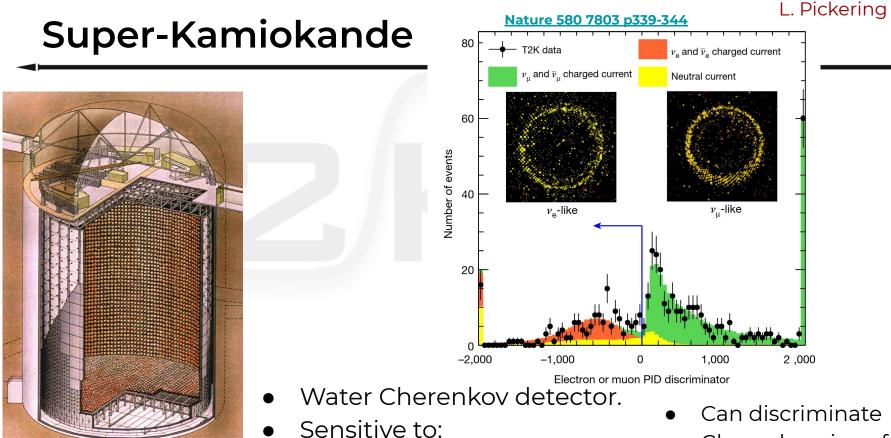






Material:Ultrapure WaterTotal Water Mass:50 kTDetection technique:CherenkovBaseline:295 kmPeak neutrino energy:0.6 GeVLocation:Mozumi Mine, Gifu, Japan





MICHIGAN STATE

- Electrons, muons, pions Ο
  - Decaying pions via Michel Ο electron

- Cherenkov rings from:
  - electrons ('fuzzy') -Ο

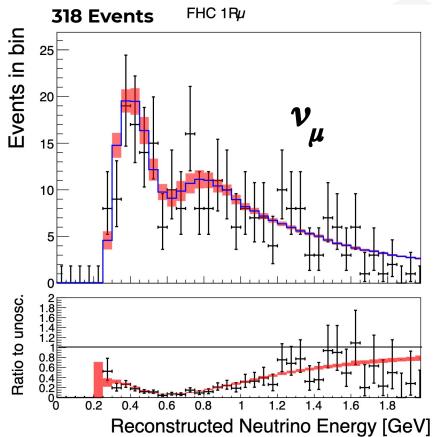
52

muons ('sharp') Ο

# Oscillation Results



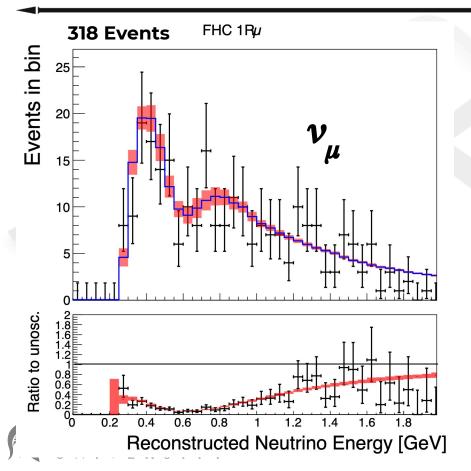




- Posterior predictive error on measured spectra. Shows remaining post-fit uncertainty in:
  - Neutrino flux prediction
  - Neutrino interaction modelling
  - Oscillation parameters

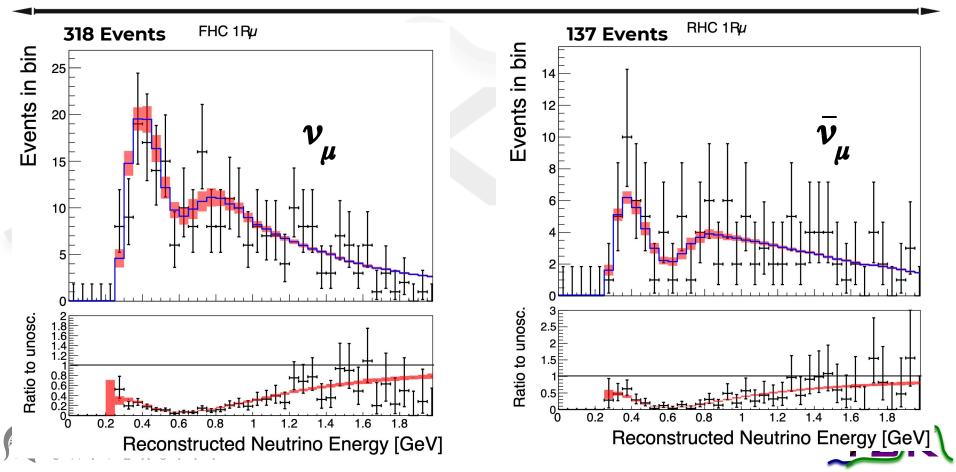
	$\ $ 1R $\mu$		$1 \mathrm{R}e$		
Error source (units: %)	FHC	RHC $\parallel$ FHC	RHC	FHC CC1 $\pi^+$	FHC/RHC
Flux Xsec (ND constr)	$ \begin{array}{c c} 2.9 \\ 3.1 \end{array} $	$\begin{array}{c c} 2.8 \\ 3.0 \\ 3.2 \end{array}$	$2.9 \\ 3.1$	2.8 $4.2$	$\begin{array}{c} 1.4 \\ 1.5 \end{array}$
Flux+Xsec (ND constr) Xsec (ND unconstrained) SK+SI+PN	$ \begin{array}{c c} 2.1 \\ 0.6 \\ 2.1 \end{array} $	$\begin{array}{c c} 2.3 \\ 2.5 \\ 1.9 \\ 3.1 \end{array}$	$2.3 \\ 3.6 \\ 3.9$	$4.1 \\ 2.8 \\ 13.4$	$ \begin{array}{c c} 1.7 \\ 3.8 \\ 1.2 \end{array} $
Total	3.0	$4.0 \parallel 4.7$	5.9	14.3	4.3

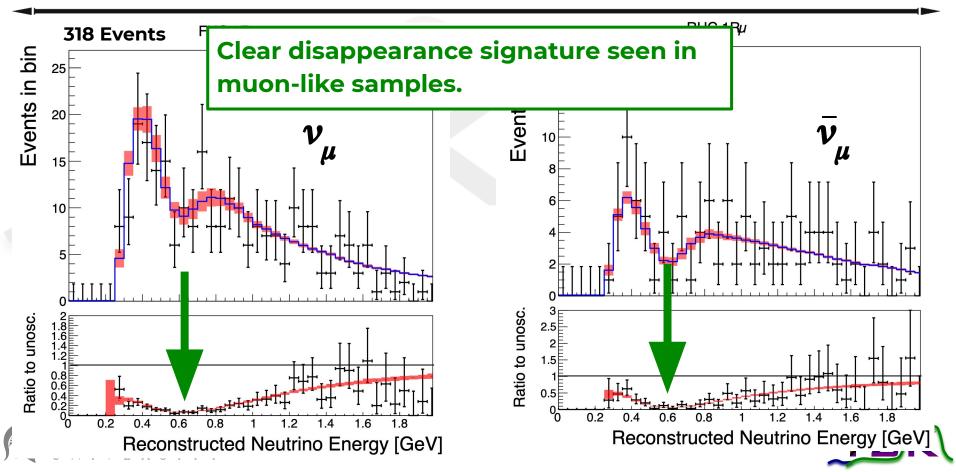


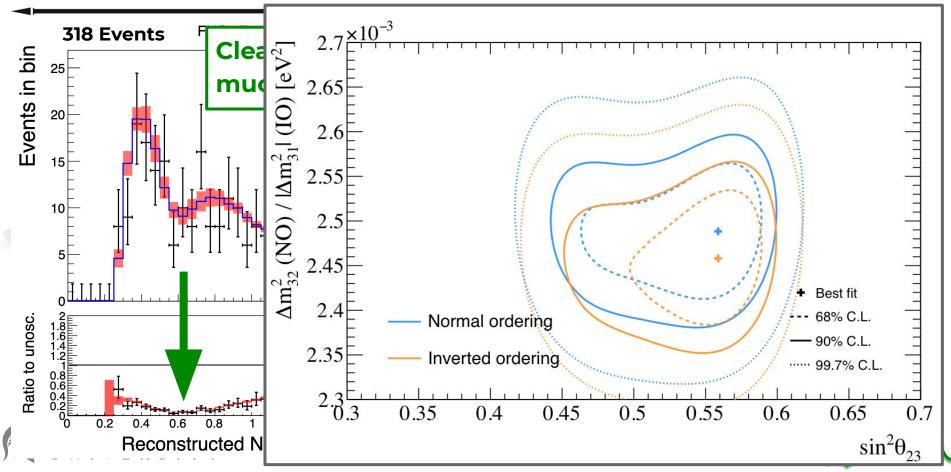


- Posterior predictive error on measured spectra. Shows remaining post-fit uncertainty in:
  - Neutrino flux prediction
  - Neutrino interaction modelling
  - Oscillation parameters
- Simultaneous fit to five samples:
  - muon-like and electron-like samples
  - neutrino-mode and anti-neutrino mode samples



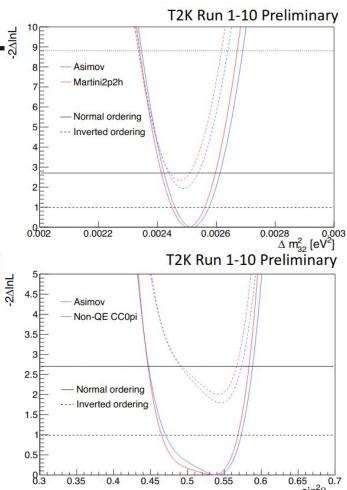






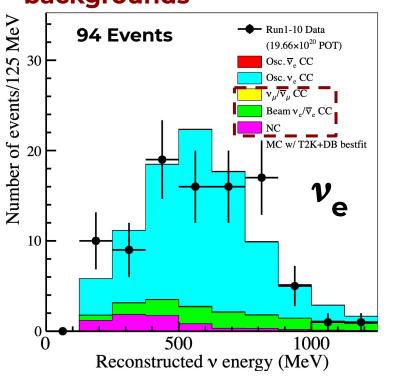
## **Robustness Studies**

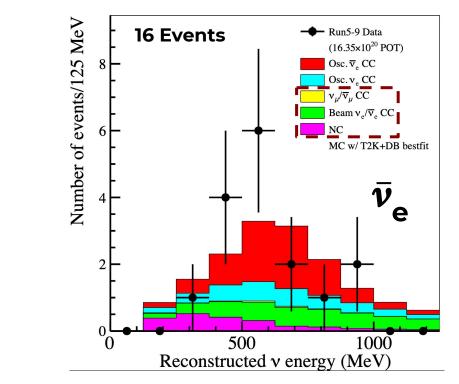
- Systematic uncertainty in neutrino interaction models are large.
- Full simulation of multiple different models is computationally infeasible.
- Develop reweighting schemes to make 'mock data' predictions from various alternate models.
  - Test if analysis is robust to such variations and inflate uncertainties as necessary.
- No significant bias seen on  $\delta_{CP}$ ,  $\theta_{23}$ , or  $\theta_{13}$  from any alternate model study:
  - $\circ$  Small bias on  $\Delta m^2_{~32}$  accounted for by inflating error by 1.4x10^{-5} eV^2



## **Electron-like samples**

• Significant electron appearance seen above predicted backgrounds

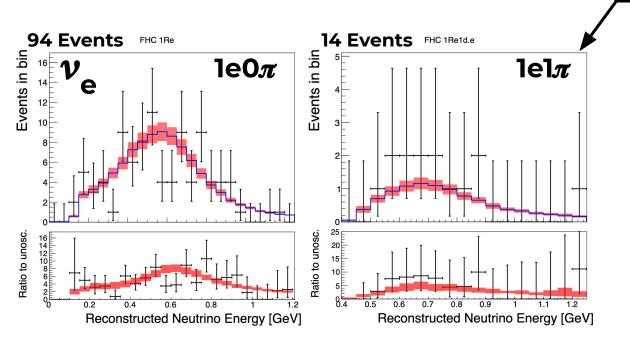




## **Electron-like samples**

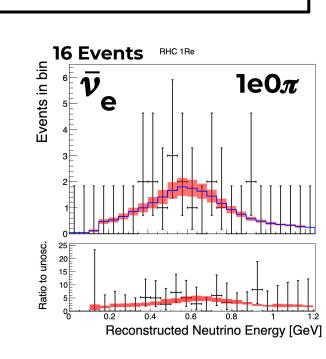
Ο

Select electron-neutrino single-pion events by tagging delayed michel electron from pion decay



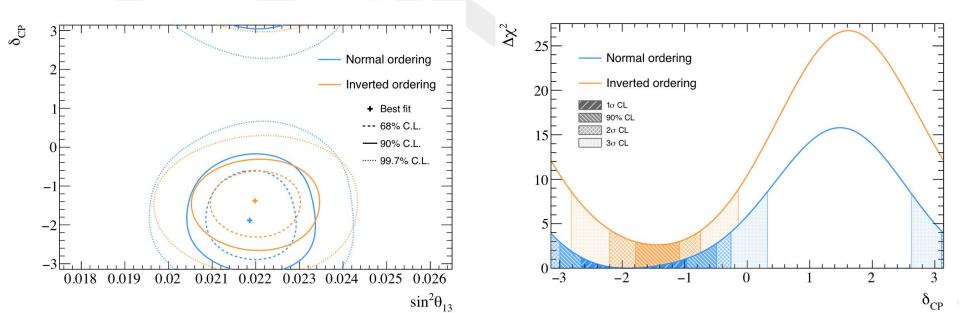
Significant electron appearance seen:

Both neutrino and antineutrino beam modes



## **Appearance Constraints**

- Includes PDG constraint on  $\theta_{13}$  from reactor neutrino measurements.
- Disfavor CP conservation ( $\delta_{CP}$ =0,± $\pi$ ) at 90%
- Bayesian analysis favors Normal mass ordering at 80% confidence.
  - Largest bias seen in robustness studies moves edge of 90% contour by **0.08**.



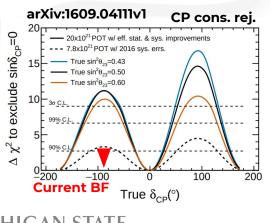
## **The Future**

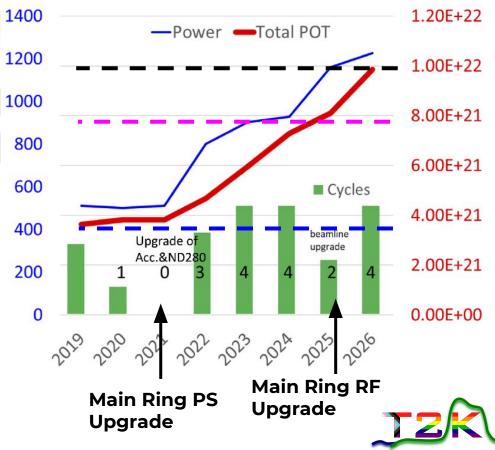




## T2K Extension and J-PARC Beam Upgrade

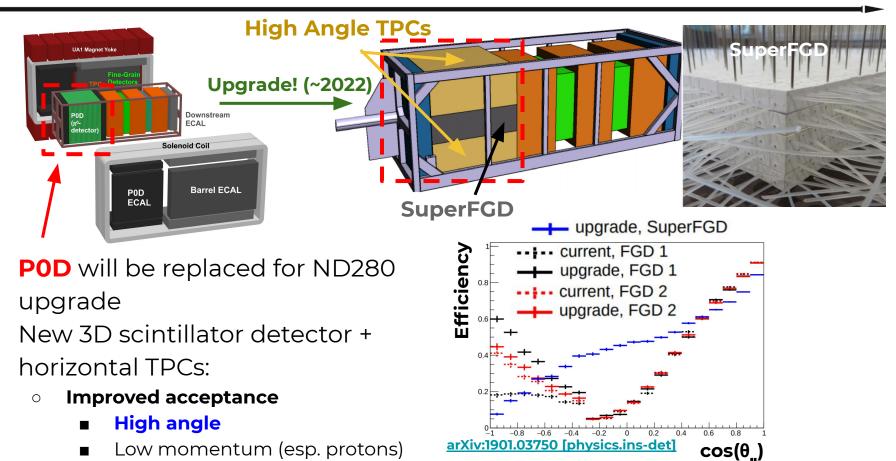
- T2K has recorded **3.64x10<sup>21</sup> POT** 1400
- KEK now budget for: 10x10<sup>21</sup>
- Continued rich physics program <sup>1</sup> and improved oscillation sensitivity until Hyper-K





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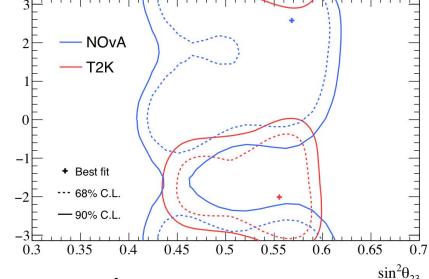
## ND280 Upgrade





- Ongoing joint oscillation fit effort with the NOvA collaboration.
- Probe the same oscillation physics (Same L/E)
  - Different L
    - T2K: 295 km
    - NOvA: 810 km
  - Different <E>
    - T2K: 0.6 GeV
    - NOvA: 2.3 GeV
  - Different dominant systematic uncertainties
- Complementary constraints!



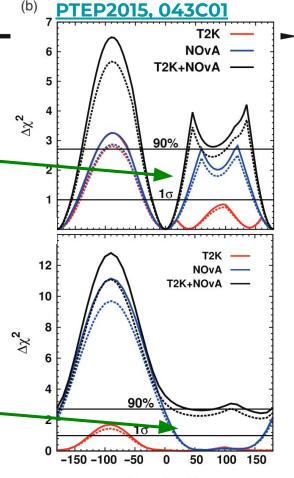


#### Joint T2K-NOvA Meeting, J-PARC 2018





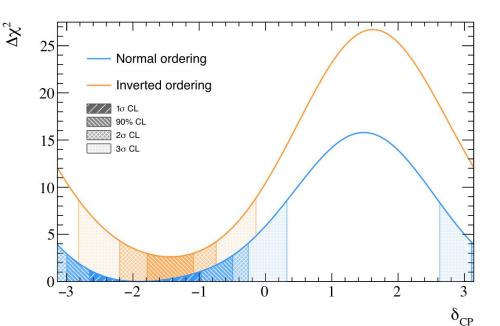
CP conservation rejection: Degenerate for T2K but not T2K+NOvA



True δ<sub>CP</sub>(°)

## Summary

- It's an exciting time in long baseline neutrino physics!
- World-leading measurements of neutrino mixing parameters.
- Seeing sensitivity to lepton-sector CPV.
- Future joint fit with NOvA expected to increase sensitivity to CPV!







#### Dawn from the summit of Fuji-san

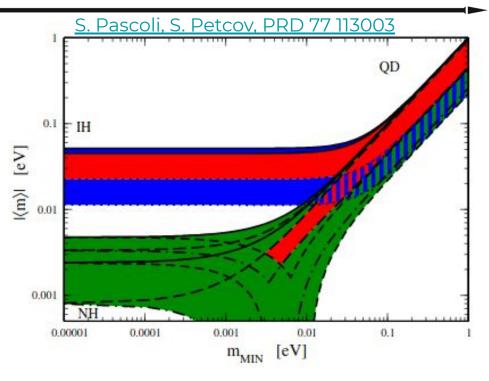
## **Thanks for listening**





## **Impact of Mass Ordering**

- If inverted ordering, why?
  - Would break a structure seen in other areas of the standard model.
- If inverted ordering (IH), then there is a lower limit on required sensitivity for neutrinoless double beta decay searches for Majorana nature of the neutrino.







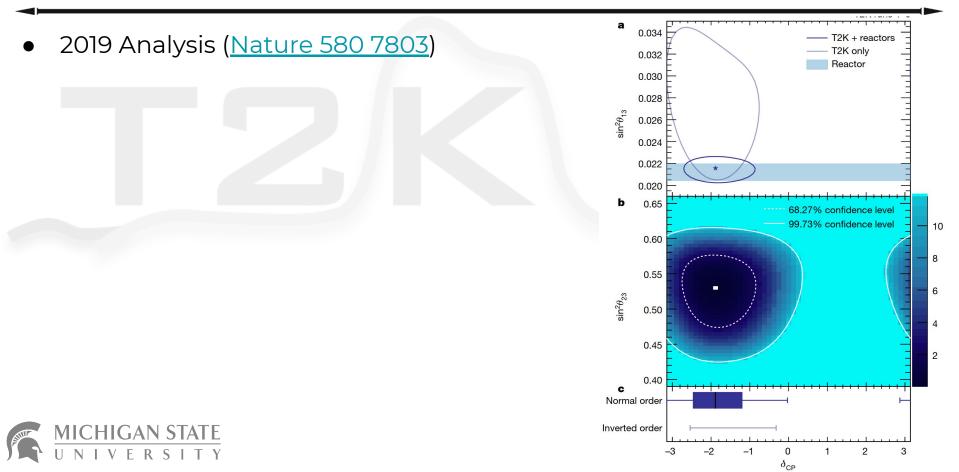
# What Changed Since 2019?

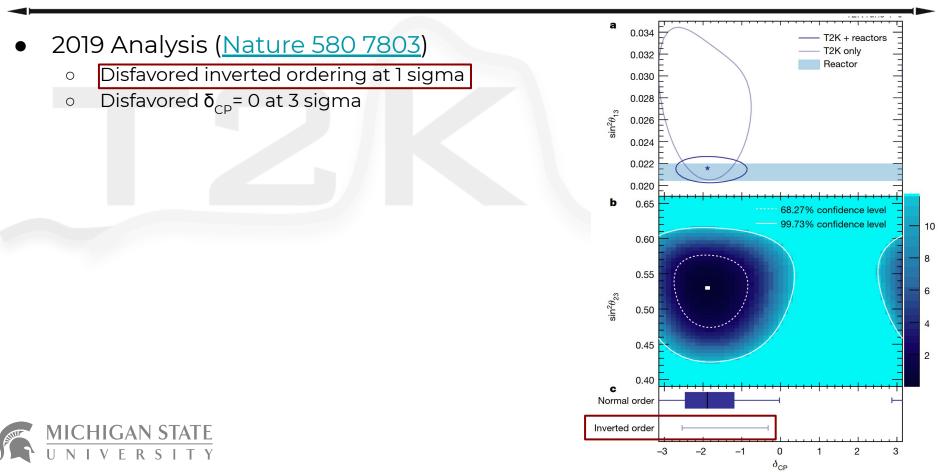


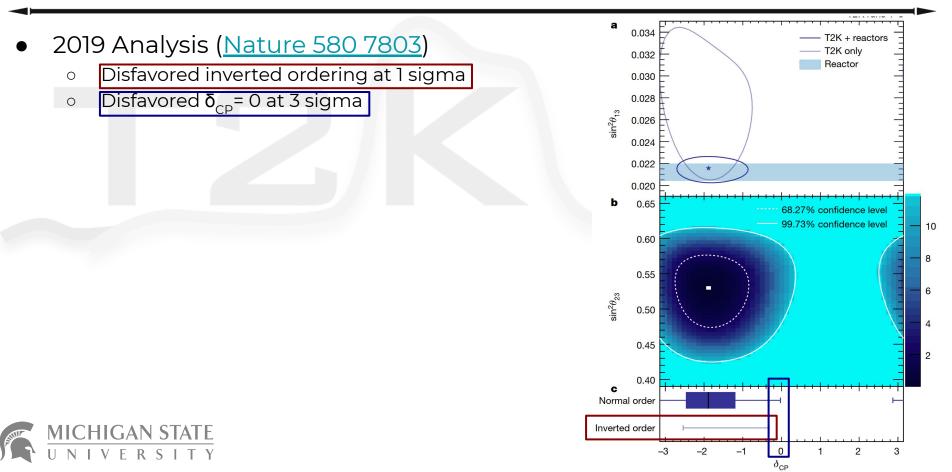




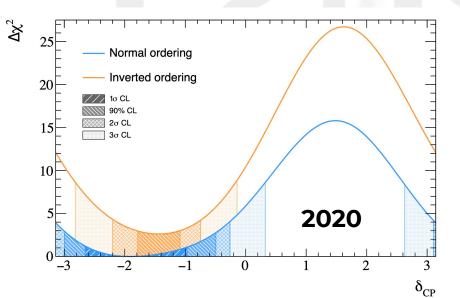


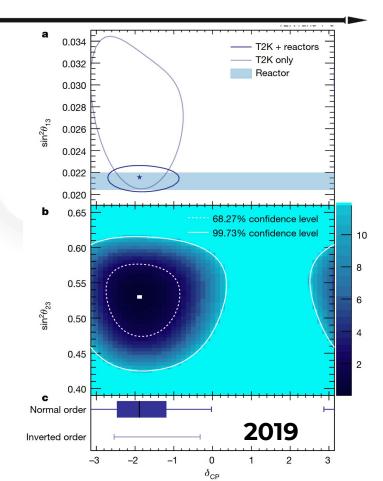




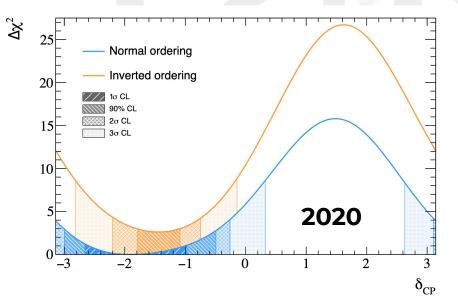


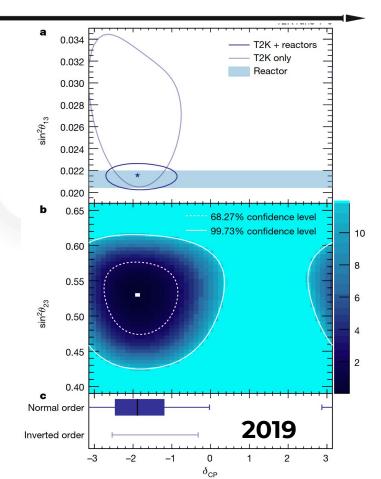
- 2019 Analysis (<u>Nature 580 7803</u>)
  - Disfavored inverted ordering at 1 sigma
  - Disfavored  $\delta_{CP}$  = 0 at 3 sigma
- 2020 Analysis presents slightly looser constraint.





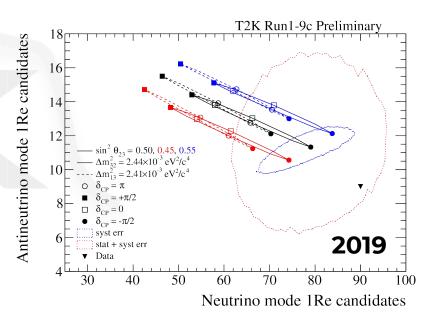
- 2019 Analysis (<u>Nature 580 7803</u>)
  - Disfavored inverted ordering at 1 sigma
  - Disfavored  $\delta_{CP}$  = 0 at 3 sigma
- 2020 Analysis presents slightly looser constraint. **Why?**





## What Changed?

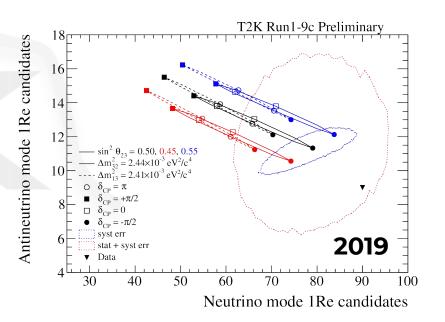
 Previous result exceeded projected CPV sensitivity through 1-sigma tension with PNMS prediction.





## What Changed?

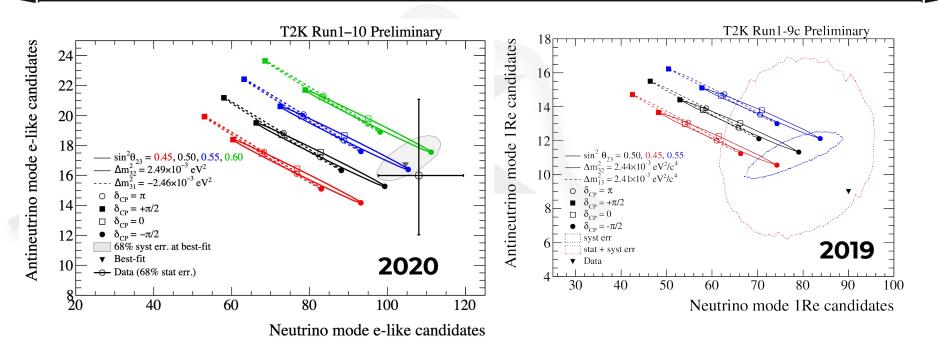
- Previous result exceeded projected CPV sensitivity through 1-sigma tension with PNMS prediction.
  - Expected at the time to be a statistical fluctuation
  - 'Too many' neutrinos and 'too few' anti-neutrinos.







#### What Changed?



2020 result is 'more consistent' with PMNS predictions, more inline with projected sensitivity.





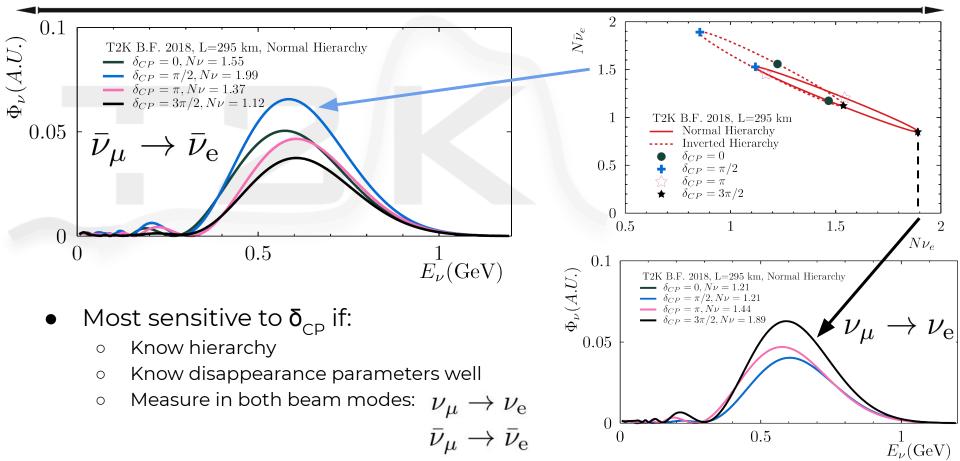




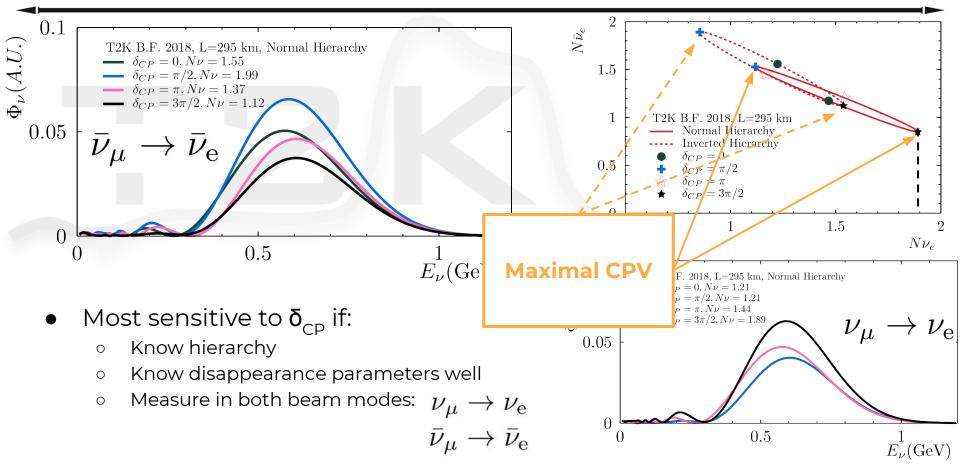




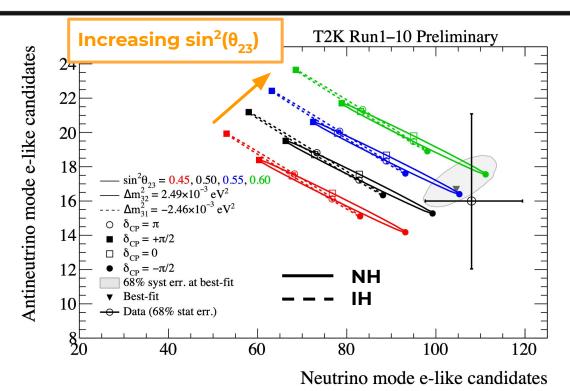
## Neutrino and Antineutrino Appearance



## Neutrino and Antineutrino Appearance



#### Matter vs. Antimatter rates

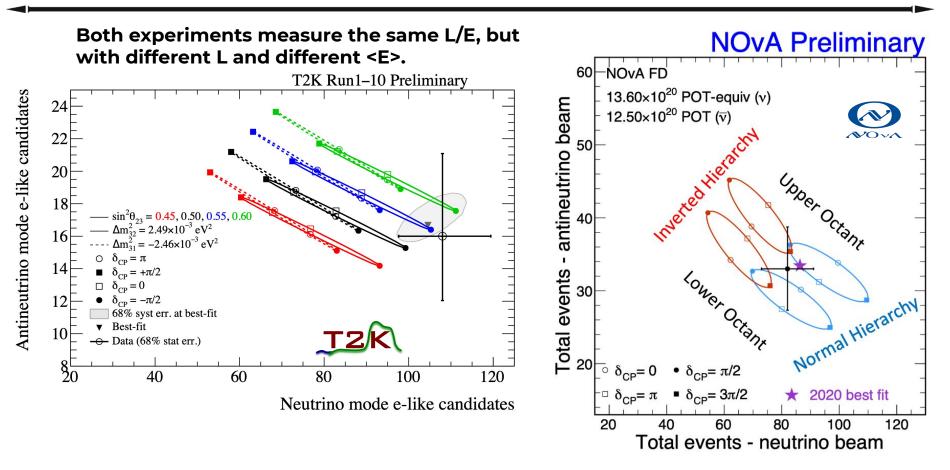


• Compare expected and observed  $v_{a}$  and  $\overline{v}_{a}$  rates:

 $\circ$  ~ Sensitivity to  $\square^{}_{\rm CP}$  and Mass hierarchy

## T2K and NOvA

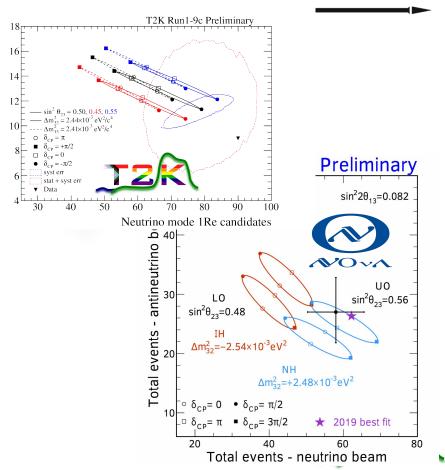
A. Himmel, Neutrino2020



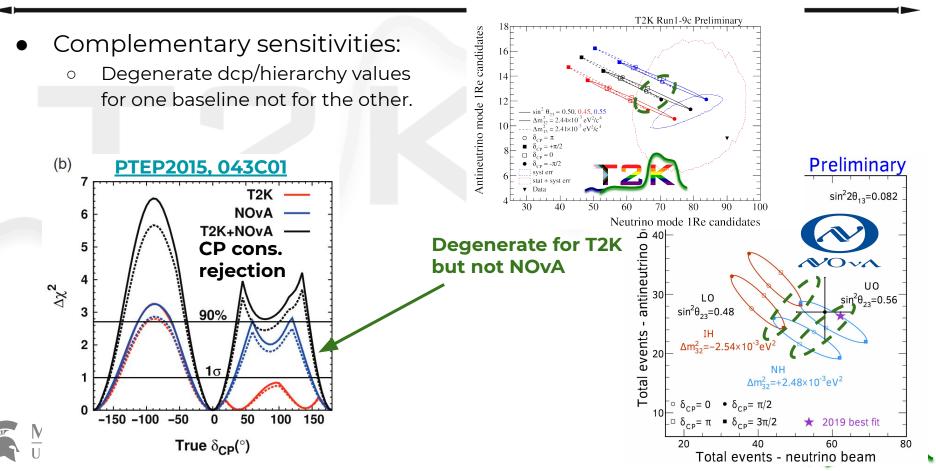
## T2K-NOvA Complementarity

Antineutrino mode 1Re candidates

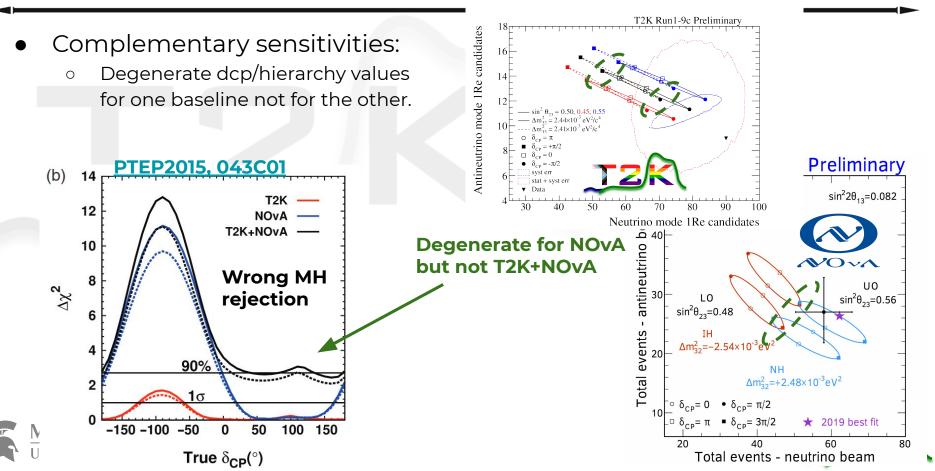
- Complementary sensitivities:
  - Degenerate dcp/hierarchy values for one baseline not for the other.



## T2K-NOvA Complementarity



## T2K-NOvA Complementarity



0

#### A. Himmel, Neutrino2020

L. Pickering

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T2K Run1-10 Preliminary Antineutrino mode e-like candidates Complementary sensitivities 24 22 from both experiments: 20F Degenerate dcp/hierarchy values 18 = 0.45, 0.50, 0.55, 0.  $= 2.49 \times 10^{-3}$  $= -2.46 \times 10^{-3} \text{ eV}^{-3}$ for one baseline not for the other.  $= +\pi/2$  $\Box = \delta_{CD} = 0$  $\delta_{CP} = -\pi/2$ 68% syst err. at best-fit **NOvA** Preliminary ▼ Best-fit Data (68% stat err. 20 40 60 80 100 120 -equiv (v) Inverted Hierarchy Upper Octant ï Total events -8 8 lower Octant Normal Hierarchy  $\delta_{\rm CP} = 0$  •  $\delta_{\rm CP} = \pi/2$  $\delta_{CP} = \pi = \delta_{CP} = 3\pi/2$ 2020 best fit 20 40 60 80 100 120 Total events - neutrino beam

#### A. Himmel, Neutrino2020

L. Pickering

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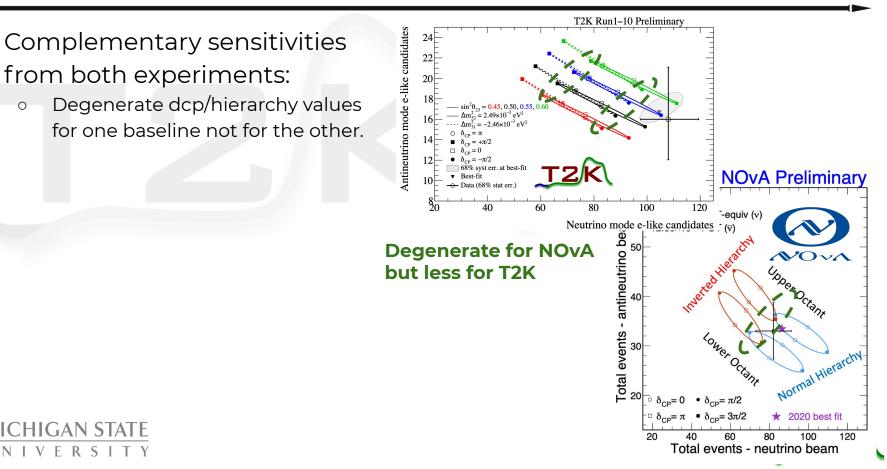
T2K Run1-10 Preliminary Antineutrino mode e-like candidate Complementary sensitivities 24 22 from both experiments: 20 Degenerate dcp/hierarchy values 18 0 = 0.45, 0.50, 0.55, 0.60 $= 2.49 \times 10^{-3} \text{ eV}^2$  $= -2.46 \times 10^{-3} \text{ eV}^{-3}$ for one baseline not for the other.  $= +\pi/2$  $\Box \delta_{cn} = 0$  $\delta_{cp} = -\pi/2$ 68% syst err. at best-fit **NOvA** Preliminary Best-fit Data (68% stat err 20 40 60 80 100 120 -equiv (v) Neutrino mode e-like candidates  $\overline{(v)}$ be Inverted Hierarchy **Degenerate for T2K** 50 antineutrino but less for NOvA 40 Total events -8 8 lowero Normal Hierarchy  $\circ \delta_{CP} = 0 \quad \bullet \delta_{CP} = \pi/2$  $\delta_{CP} = \pi = \delta_{CP} = 3\pi/2$ 2020 best fit 20 40 60 80 100 120 Total events - neutrino beam

•

#### A. Himmel, Neutrino2020

L. Pickering

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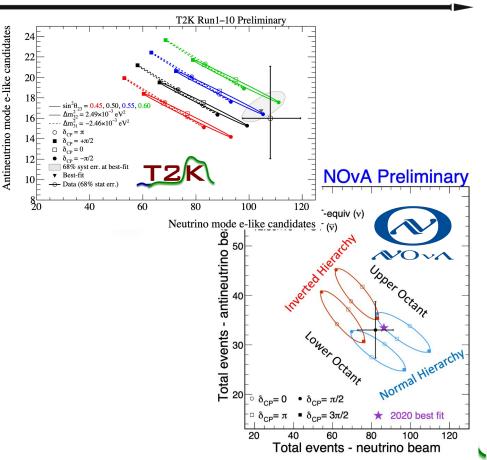


#### A. Himmel, Neutrino2020

L. Pickering

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- Complementary sensitivities from both experiments:
  - Degenerate dcp/hierarchy values for one baseline not for the other.
- Joint fit with access to full likelihood allows for:
  - more robust statistical treatment,
  - correlations between important systematic parameters.

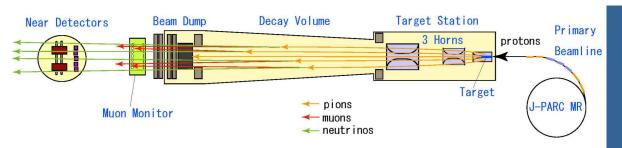


## **Oscillation Details**

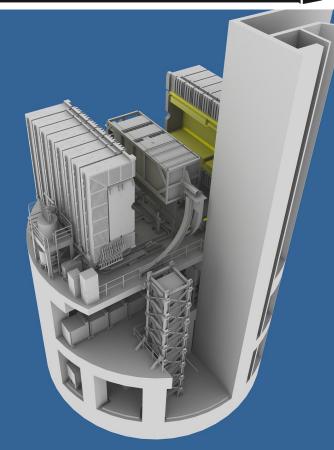




#### **Near Detector Complex**

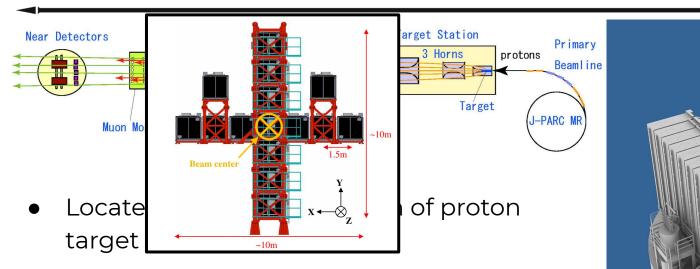


• Located 280 m downstream of proton target station.



**NGRID** 

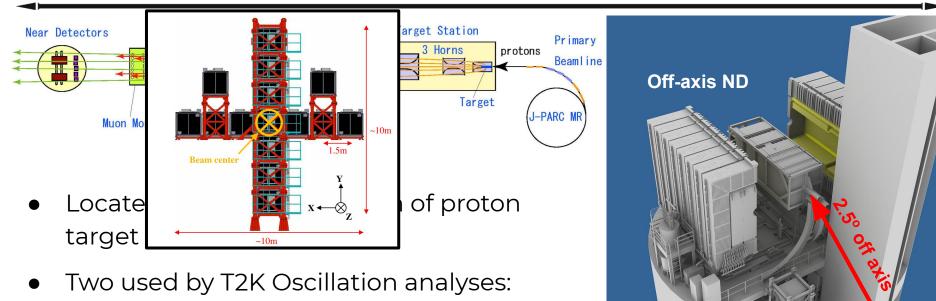
#### **Near Detector Complex**



- Two used by T2K Oscillation analyses:
  - INGRID: On-axis, ensures beam alignment

**NGRID** 

#### **Near Detector Complex**



- INGRID: On-axis, ensures beam alignment
- Off-axis near detector

#### **Uncerts table**

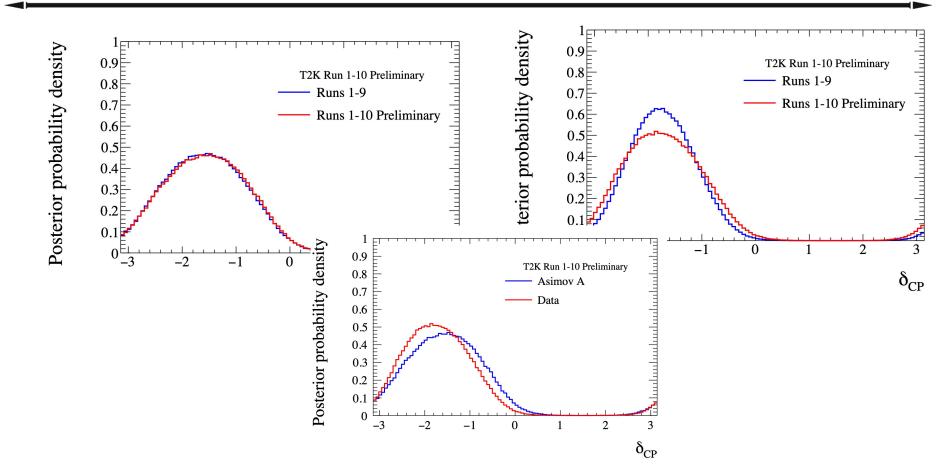
Table 20: Uncertainty on the number of event in each SK sample broken by error source after the BANFF fit. To obtain error rates comparable with the "Flux+Xsec (ND constrained)" presented by MaCh3 [22], square sum the "Flux+Xsec (ND constr)", " $\sigma(\nu_e)$ ,  $\sigma(\bar{\nu}_e)$ ", "NC  $\gamma$ ".

	$\ $ 1R $\mu$		1Re			
Error source	FHC	RHC	FHC	RHC	FHC CC1 $\pi^+$	FHC/RHC
Flux	2.9	2.8	2.8	2.9	2.8	1.4
Xsec (ND constr)	3.1	3.0	3.2	3.1	4.2	1.5
Flux+Xsec (ND constr)	$\parallel 2.1$	2.3	2.0	2.3	4.1	1.7
2p2h Edep	0.4	0.4	0.2	0.2	0.0	0.2
$\mathrm{BG}_A^{\mathrm{RES}}$ low- $p_{\pi}$	0.4	2.5	0.1	2.2	0.1	2.1
$\sigma( u_e),\sigma(ar u_e)$	0.0	0.0	2.6	1.5	2.7	3.0
NC $\gamma$	0.0	0.0	1.4	2.4	0.0	1.0
NC Other	0.2	0.2	0.2	0.4	0.8	0.2
SK	2.1	1.9	3.1	3.9	13.4	1.2
Total	3.0	4.0	4.7	5.9	14.3	4.3



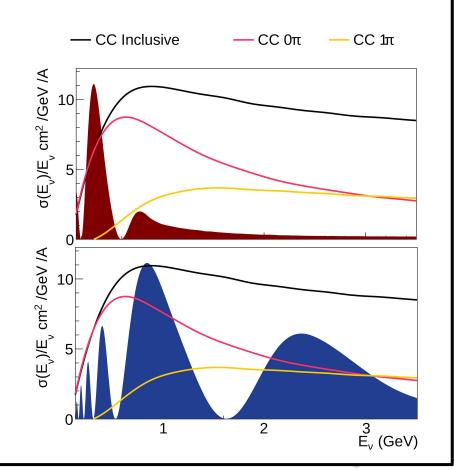


#### **Expected Sensitivity**



- Complementary sensitivities from both experiments:
  - Degenerate dcp/hierarchy values for one baseline not for the other.
- Joint fit with access to full likelihood allows for:
  - more robust statistical treatment,
  - correlations between important systematic parameters.
- Studies of which systematic parameters should be correlated are underway!

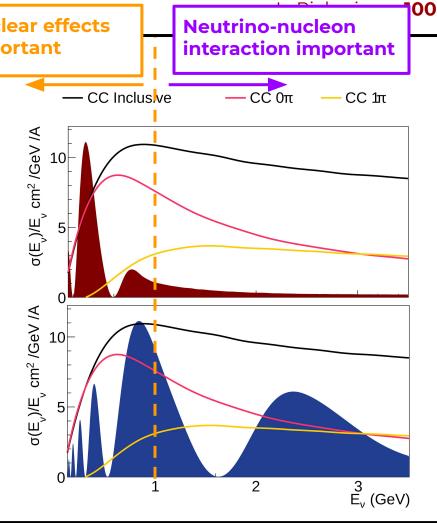




## T2K-NOvA Compleme

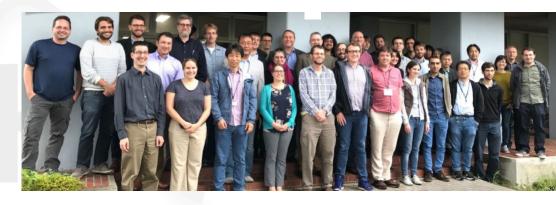
- Complementary sensitivities from both experiments:
  - Degenerate dcp/hierarchy values for one baseline not for the other.
- Joint fit with access to full likelihood allows for:
  - more robust statistical treatment,
  - correlations between important systematic parameters.
- Studies of which systematic parameters should be correlated are underway!





### T2K-NOvA

- Joint analysis workshops on-going:
  - Four successful meetings since 2017 J-PARC and FNAL
  - Strong US-Japan support!
- Challenging joint analysis:
  - Different experimental setups
  - Different peak energy
  - Different analysis methodology
- But T2K-NOvA sensitivity is worth the challenge!



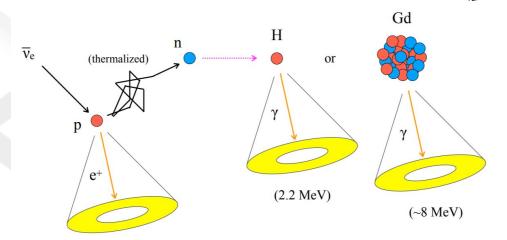






## SK-Gd

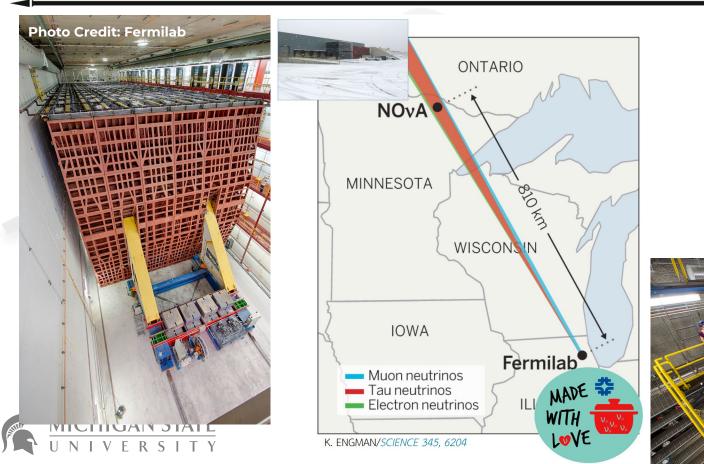
- Super-K deep cleaned in preparation for Gadolinium doping.
  - Gd loading is underway!
- Much improved efficiency for neutron capture:
  - Statistical separation of neutrino/antineutrino rate
- Extra constraint on 'wrong sign' backgrounds at the far detector



#### The NOvA Experiment

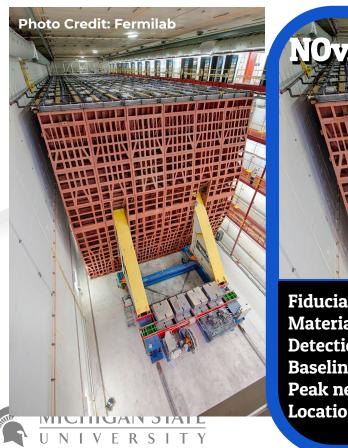


Photo Credit: Fermilab



#### **The NOvA Experiment**

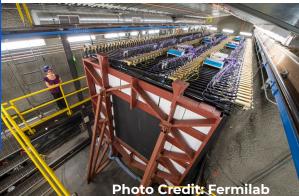


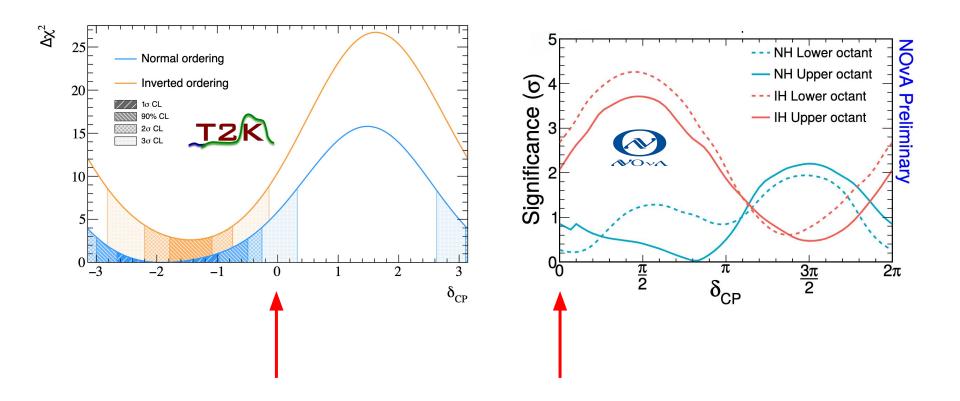


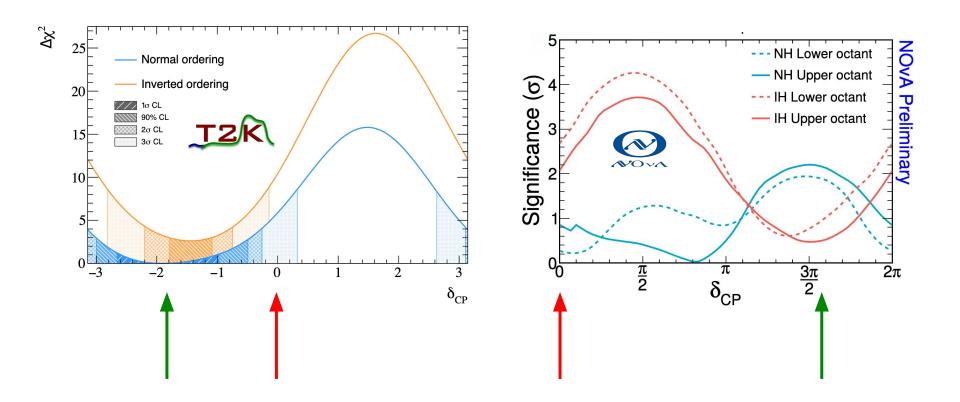


Fiducial Mass: Material: Detection technique: Baseline: Peak neutrino energy: Location:

14 kT Liquid scintillator Scintillation 810 km 1.9 GeV Ash River, MN





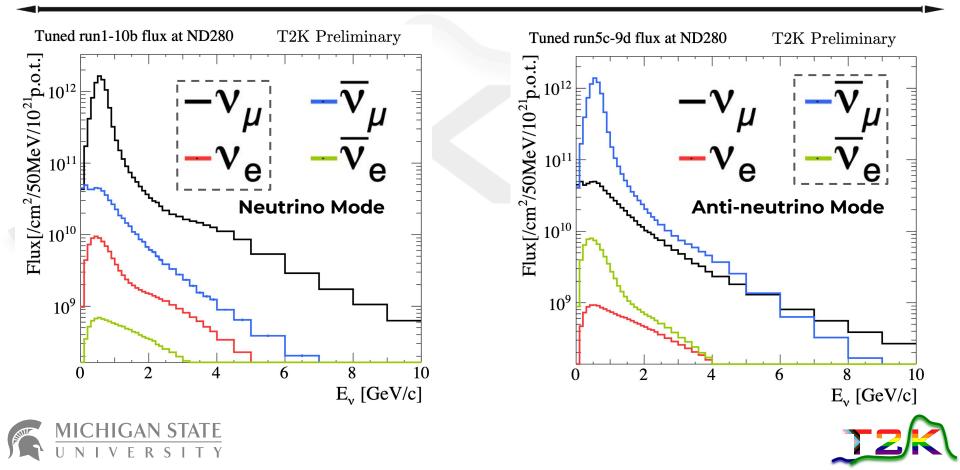


# Flux Uncertainties





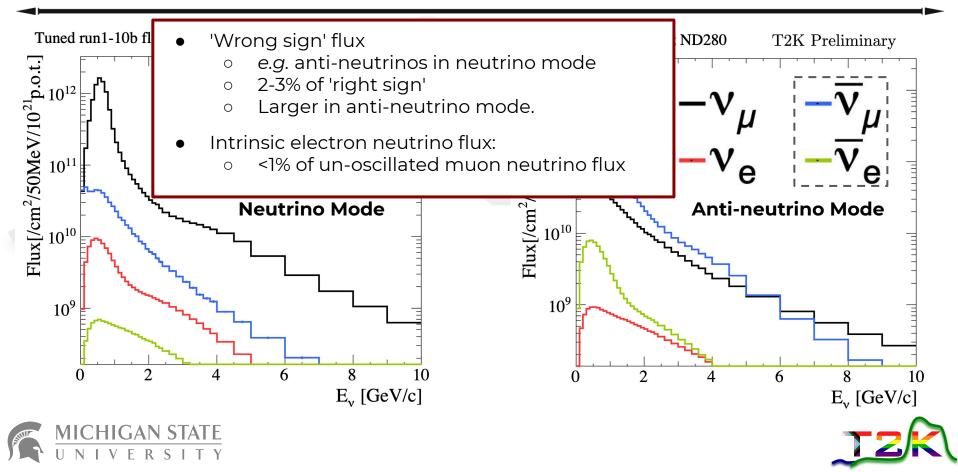
### J-PARC Beam: Neutrino Species

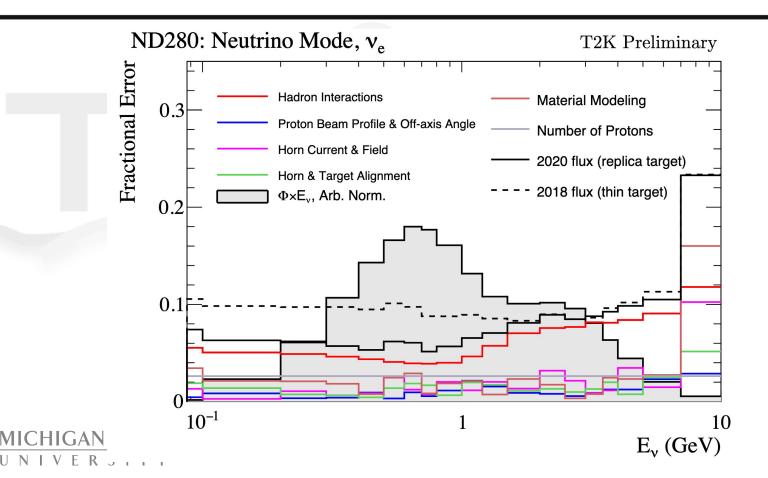


L. Pickering

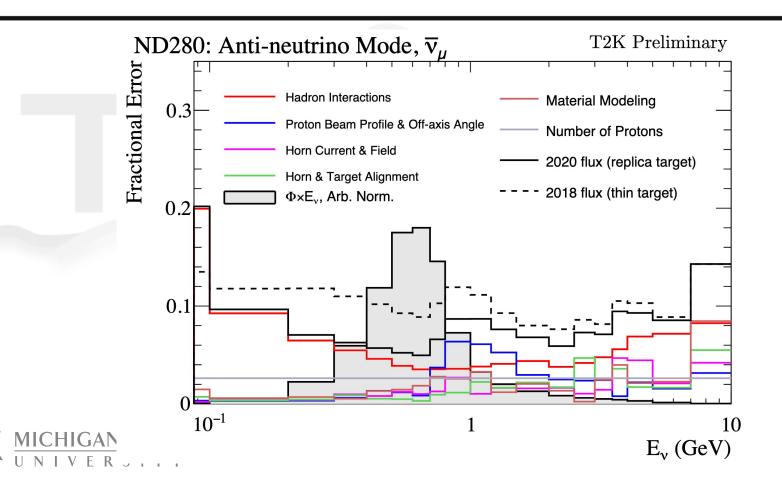
108

#### **J-PARC Beam: Neutrino Species**

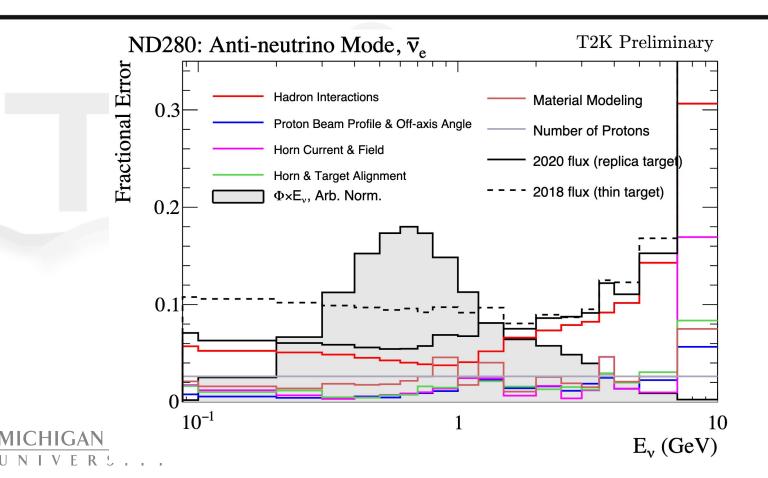












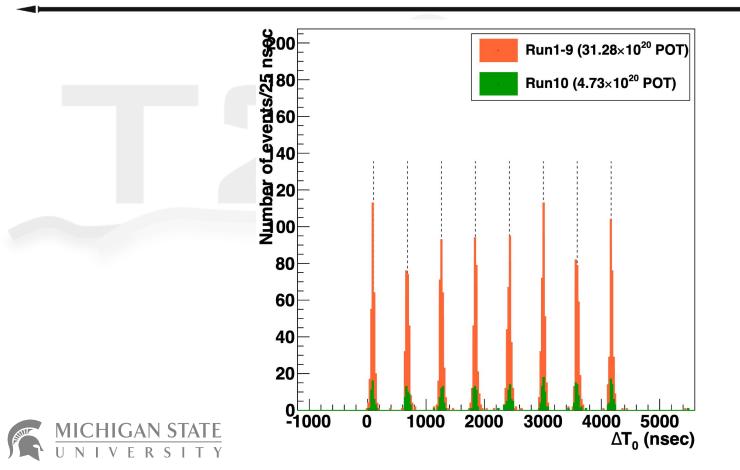








### **Beam Bunch Timing at SK**



## **Neutrino–Matter Interactions**



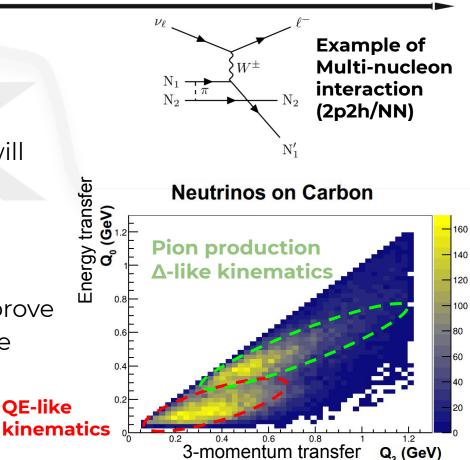
NIWG



# Multi-nucleon Interactions

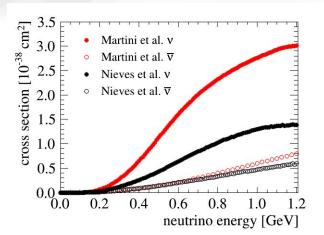
- Scattering from bound nucleon-nucleon pairs within the nucleus: different E,↔E<sub>Rec.</sub>
- Not possible to study in isolation, will always also have:
  - True CCQE
  - CC1pi with missed pion
  - Other nuclear effects
- Current multi-nucleon models improve experimental agreement, but some way still to go.

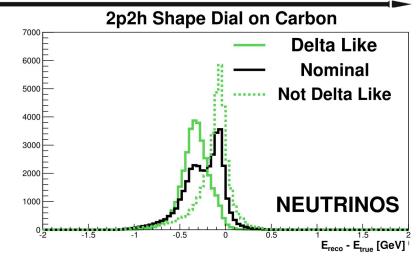




# Effect on Oscillation Analysis

- Want to check how biased the results might be if the wrong multi-nucleon model was chosen:
  - Assign uncertainty to QE-like/Δ-like nature of multi-nucleon interaction.
  - Run oscillation analysis with 'fake data' generated with an alternate model.



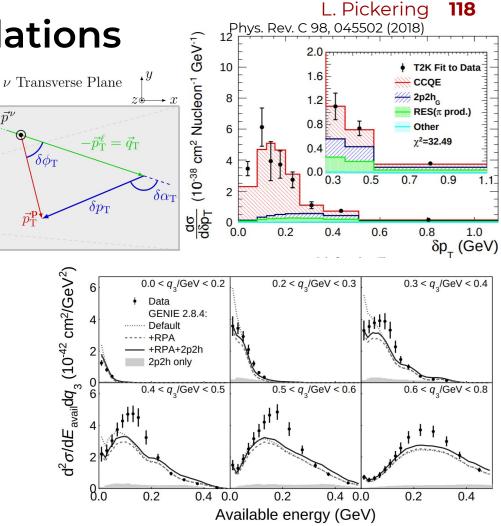


Near detector fit prefers between nominal and  $\Delta$ -like

#### 

- Investigate lepton-hadron correlations.
- Two recent approaches:
  - Transverse imbalance
  - q0/q3 reconstruction
- Hard to use directly in OA:
  - Existing models can't be bent to fit with current freedom...
  - Build 'fake data' informed by these results and use to test OA robustness.





#### 

- Energy associated with liberating struck nucleon from nuclear potential
- A. Bodek's re-analysis found that the default NEUT value was poor [arXiv:1801.0797]
- For 2018 T2K OA, a fit to mock-data with a large shift in Eb was used to assess uncertainty
  - Largest single source of error.
- In the future, a smaller prior from A. Bodek's analysis will be used.

