

Gamma rays production via intracavity Compton back-scattering with FEL photons

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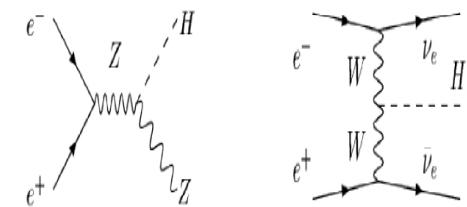
Applications

- *Non conventional source of positron for future colliders.*

- *Production of polarized e^+ beams*

Beam polarization for Higgs searches

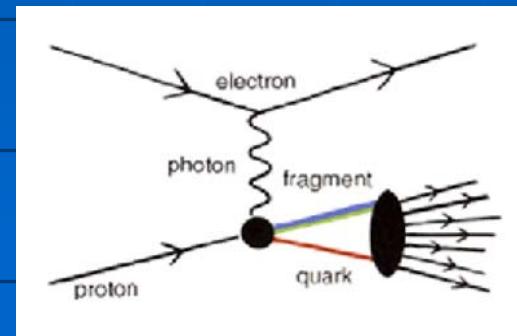
■ Light Higgs, e.g. $m_H=130$ GeV: HZ and $H\gamma\gamma$ similar rates



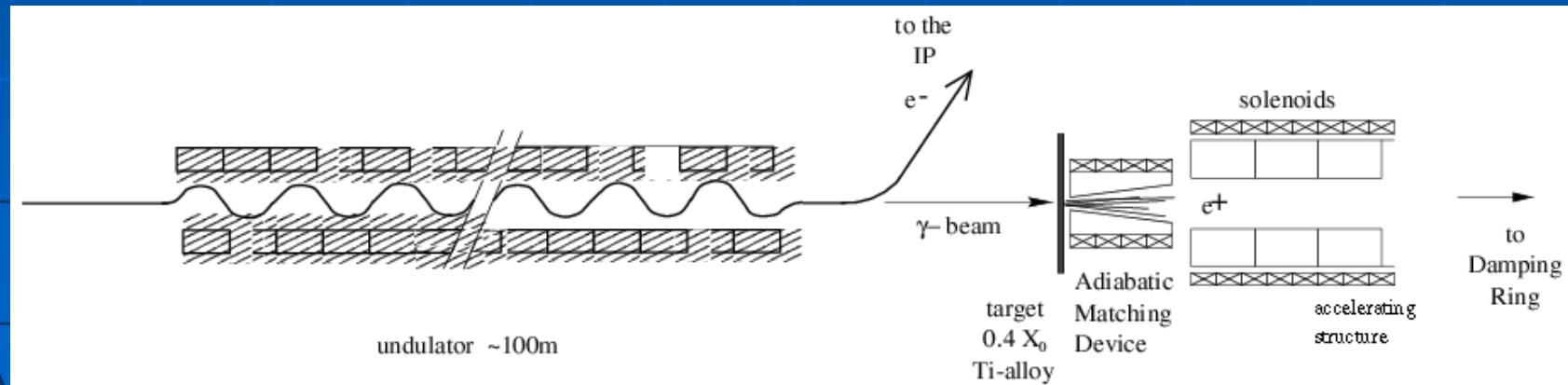
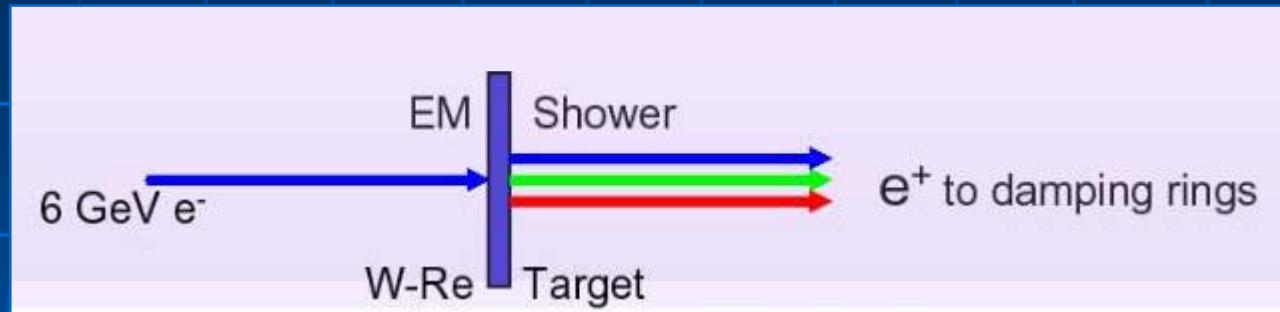
- *Nuclear Physics,*
 - *Non perturbative QCD:*
 - *Proton polarizability*

Hadron EM polarizabilities

G. Dattoli - Lett. N. Cim. (1977)



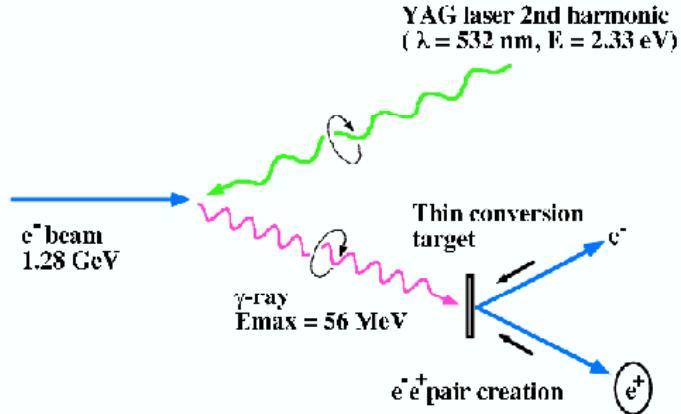
Positron source:



Compton back-scattering

■ *Proof of principle*

ATF-Compton Experiment@KEK

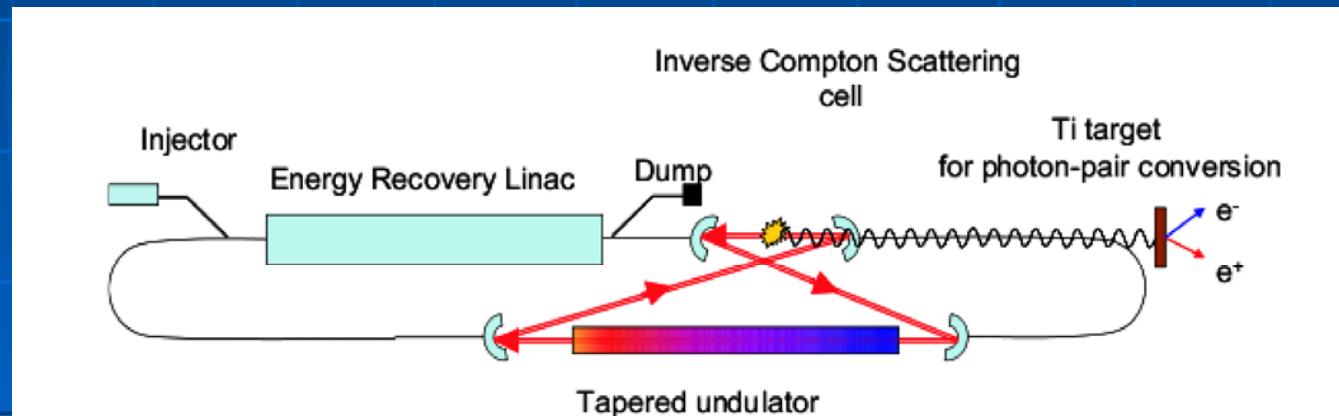


- i) proof-of-principle demonstration
- ii) accumulate technical informations:
polarimetry, beam diagnosis, ...

No Optical Cavity at Collision Point

ICS+FEL idea is not new...

- *C. Pagani, E.L Saldin, E.A. Schneidmiller and M.V. Yurkov NIM A 429 (1999) 476-480.*
- *V.N. Litvinenko and J.M.J. Madey, NIM A 375 (1996) 580-583*

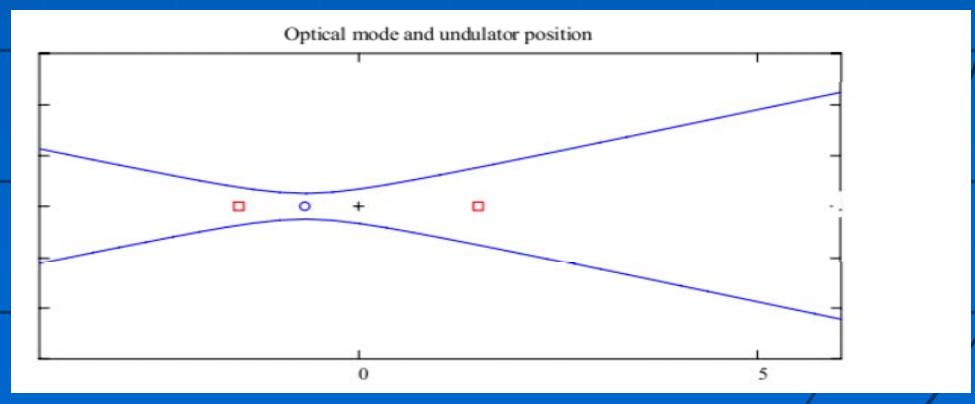


- *P. Musumeci [http://www.astec.ac.uk/id_mag/ID-Mag_Helical
ILC_Positron_Production_Workshop.htm](http://www.astec.ac.uk/id_mag/ID-Mag_Helical_ILC_Positron_Production_Workshop.htm) (broken link?)*

Lack an estimation of the γ yield.

Proposed scheme

- *FEL oscillator, Optical Klystron @ $\lambda \sim 260\text{nm}$*
- *Intracavity back-scattering, yielding $\gamma @ +70\text{ MeV}$*
- $E \sim 1\text{ GeV}$



Beam parameters

. $Q = 300 \text{ pC}$

. $I_{peak} = 80 \text{ A}$ $\sigma_t = 1.5 \text{ ps}$

. $\varepsilon_n = 10 \mu\text{m}$

. $E = 1 \text{ GeV}$ $\Delta E = 20 \text{ KeV}$

. $v_b = 2.986 \text{ GHz}$ (Rep rate in macropulse)

. $rr = 100 \text{ Hz}$

. $\Delta T = 5 \mu\text{s}$ (Macropulse)

. $n_b = \Delta T \cdot v_b = 1.5 \cdot 10^4$ (bunches in macrop.)

Undulator

N_p	60
$L_u = N_p \cdot \lambda_u$	3m

Optical Cavity

$$N_{br} = 200$$

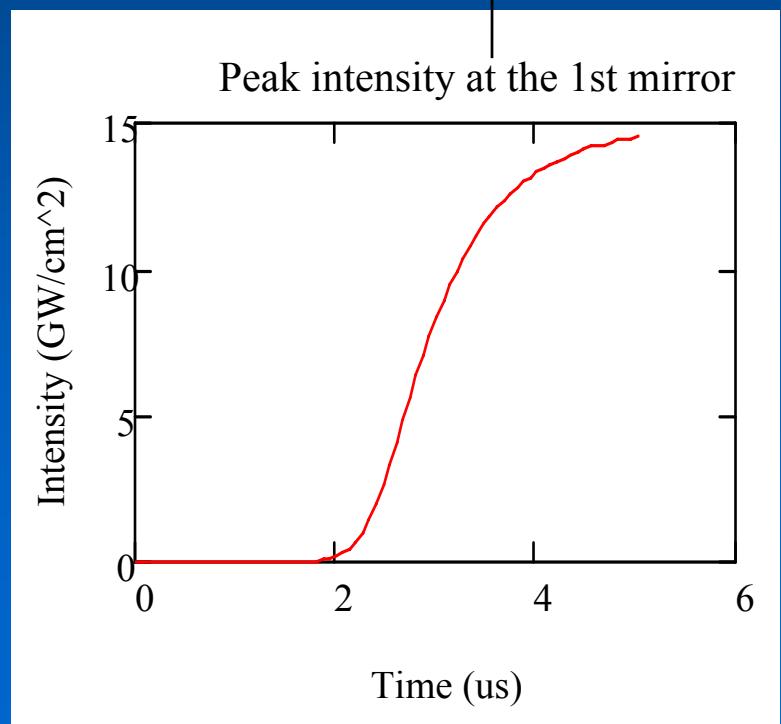
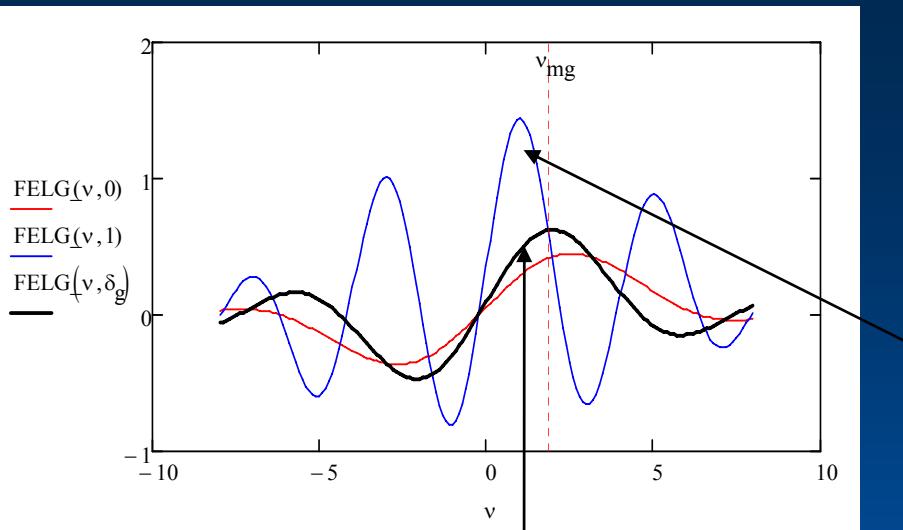
$$L_c = N_{br} \cdot c/2 \cdot v_b = 10.04 \text{ m}$$

$$R_1 = 3.5 \text{ m} \quad R_2 = 6.8 \text{ m}$$

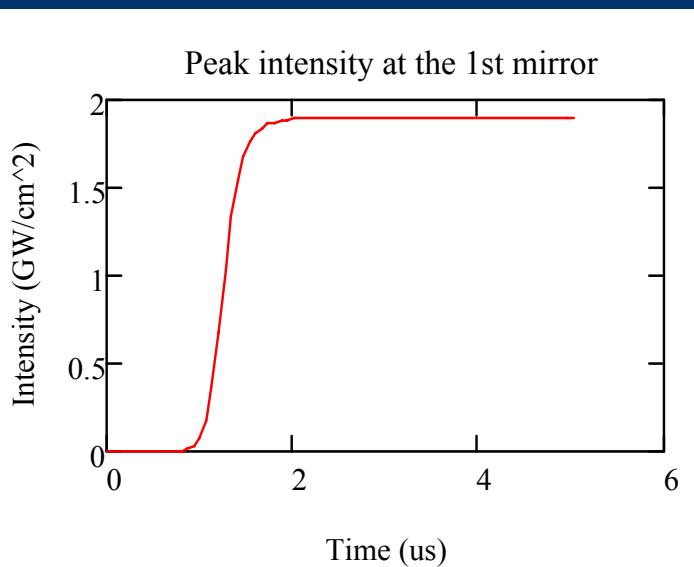
$$w_1 = 1.126 \text{ mm} \quad w_2 = 2.229 \text{ mm}$$

$$w_o = 250 \mu\text{m} \quad z_r = 761 \mu\text{m}$$

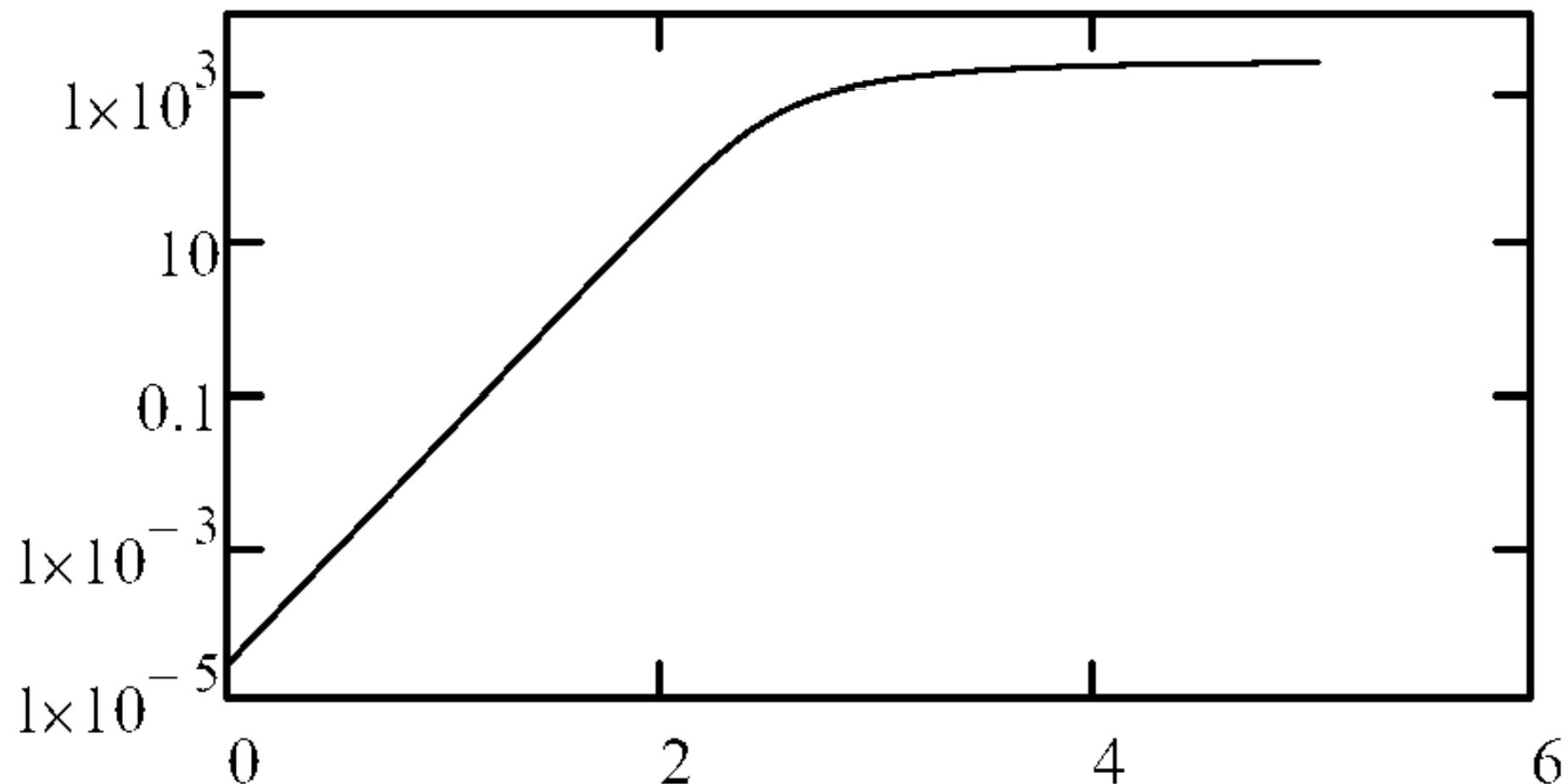
$$\text{Reflectivity} = 99\% \quad \eta = 1 - R^2 = 2\%$$



Gain optimization



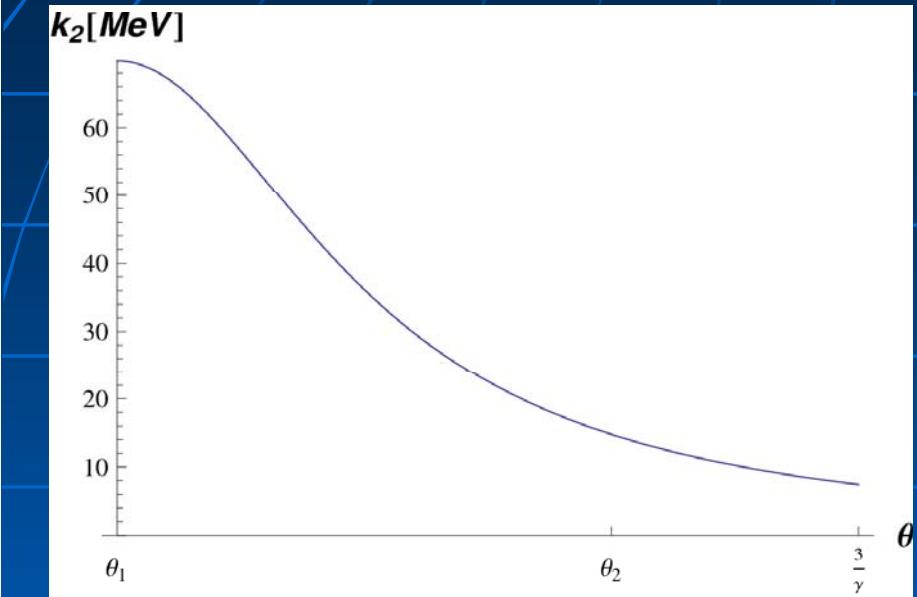
Intracavity peak power



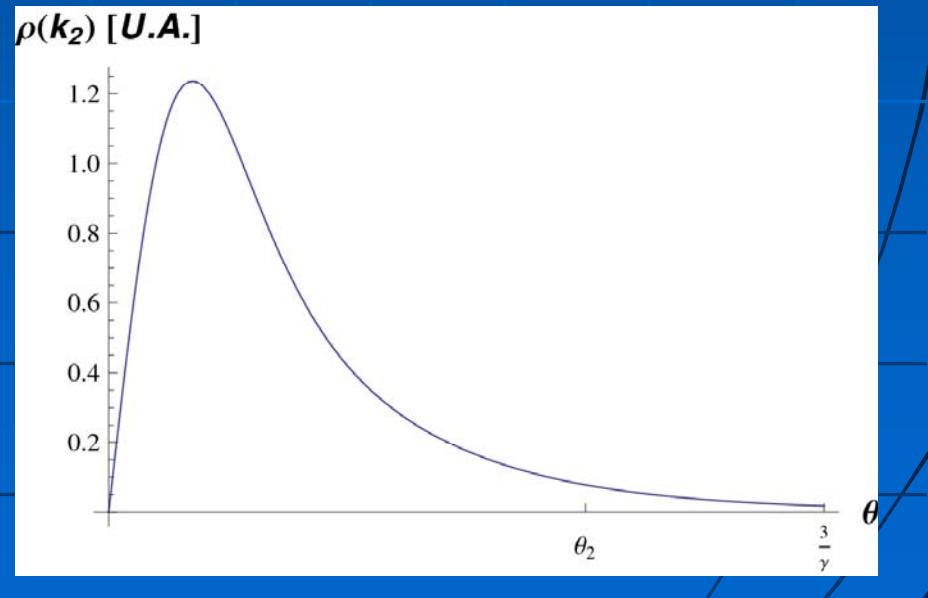
$$2^{nd} \text{ mirror} \rightarrow 3.311 \frac{\text{J}}{\text{cm}^2}$$

Time (μs)

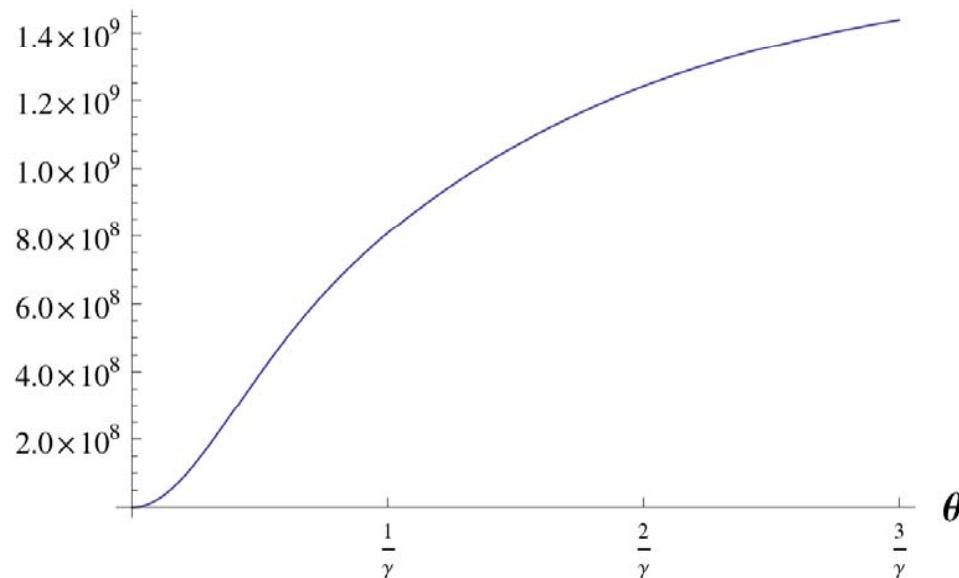
$k_2(\theta)$



$\rho_{k_2}(\theta)$

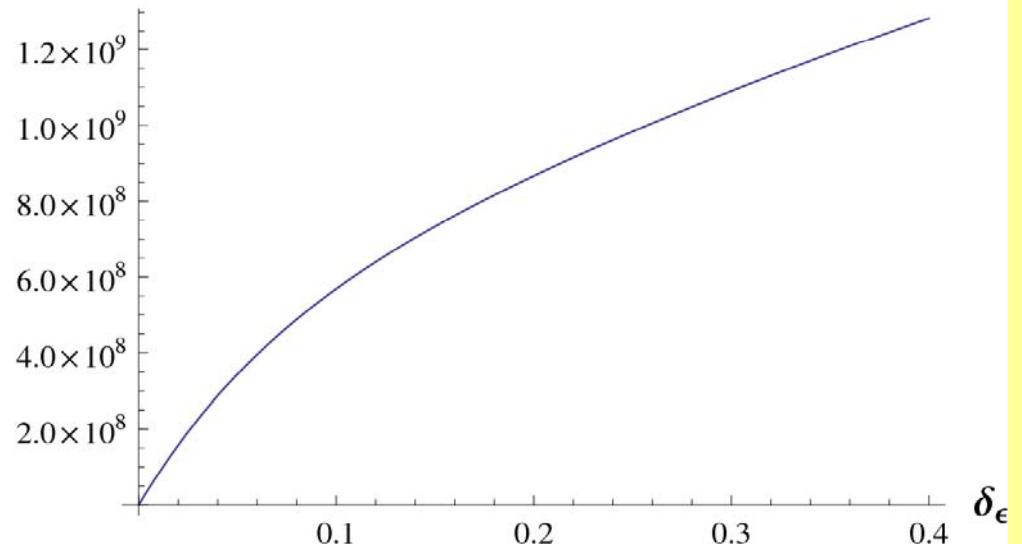


$$\frac{dN_\gamma}{dt} [\text{photons/sec}]$$



$$\frac{dN_\gamma}{d\Omega} = rrN_{br} \frac{1}{2\pi} \frac{d\sigma}{d\Omega} \frac{N_{FEL} N_e}{\sigma_{ph}^2 + \sigma_e^2}$$

$$\frac{dN_\gamma}{dt} [\text{photons/sec}]$$



Conclusions

- The hard γ yield from a practical Optical Klystron FEL has been estimated in terms of collection angle and energy spread;
- Monte Carlo simulations of electromagnetic showers are needed to characterize the interaction with target for e^+ production to assess achievable polarization;