### A New Multiwavelength Census of Blazars

**A. Paggi** <sup>1,2,3</sup> M. Bonato <sup>4</sup>, C. M. Raiteri<sup>3</sup>, M. <u>Villata<sup>3</sup></u>, G. De Zotti<sup>5</sup> and M. I. Carnerero<sup>3</sup>

<sup>1</sup>Dipartimento di Fisica, Università degli Studi di Torino <sup>2</sup>INFN, Sezione di Torino
 <sup>3</sup>INAF, Osservatorio Astrofisico di Torino <sup>4</sup>INAF, Istituto di Radioastronomia
 <sup>5</sup>INAF, Osservatorio Astronomico di Padova

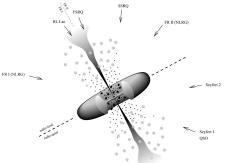




### Blazars

Radio-loud sources with emission across whole e.m. spectrum, high and variable polarization, superluminal motion, high luminosities, intense/rapid variability, flat radio spectrum.

Jet aligned with l.o.s. within a few degrees (e.g., Blandford & Rees 1978; Begelman et al. 1984; Königl 1986).



Dominate the  $\gamma$ -ray sky: 80% of extragalactic  $\gamma$ -ray sources, 75% of extragalactic TeV sources.

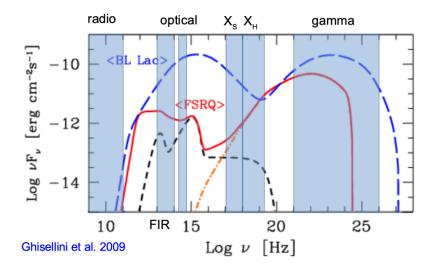
### BL Lacs:

- $L_{\rm iso} \lesssim 10^{46}$
- Spectra with weak em. lines (EW < 5 Å)</li>
- Pure non-thermal radiation
- $z \lesssim 1.3$  (peak at  $z \sim 0.3$ )

### Flat Spectrum Radio Quasars:

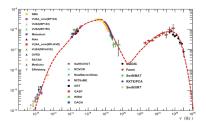
- $L_{\rm iso} \lesssim 10^{48}$
- Spectra with broad em. lines (EW > 5 Å)
- Thermal emission (BBB)
- $z \lesssim 5.5$  (peak at  $z \sim 1.5$ )

But recently,  $\gamma$ -ray **BL Lacs** with z up to 2.5 (4LAC, Ajello et al. 2020) and even 3.6 (Paliya et al. 2020).



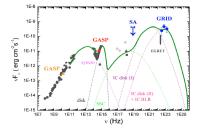
### BL Lacs vs. FSRQs

# **BL Lac**: Mrk 501 (Spurio et al. 2018)

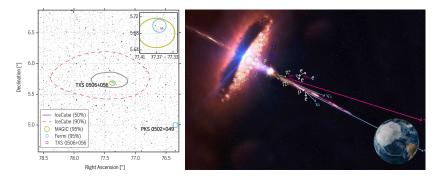


Pure non thermal continuum  $\rightarrow$ SSC: synch. by electrons in jets + IC on synch. photons (Marscher et al. 2010). Developed accretion disk (BBB + lines),  $\gamma$ -ray dominance  $\rightarrow$  **EC**: seed photons from disk, torus, BLR (Maraschi et al. 2001).





### TXS 0506+065 (IceCube Collaboration et al. 2018, Garrappa et al. 2019).



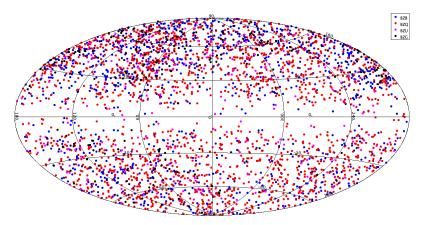
### Blazars as multi-messenger sources.

A New Multiwavelength Census of Blazars

INFN, Sezione di Torino November 17 2020 The 5<sup>th</sup> edition of the Roma-*BZCAT* Multifrequency Catalogue of Blazars (Massaro et al. 2005) represents the most comprehensive list of blazars confirmed by means of available spectra.

- 3561 sources
- Multi-wavelength informations (optical, radio flux, microwave, soft/hard X-ray flux, γ-ray) and redshift
- 1151 BZBs (1059 BL Lac objects + 92 BL Lac candidates)
- 1909 BZQs (Flat Spectrum Radio Quasars)
- 227 BZUs (blazars of uncertain type)
- 274 BZGs (BL Lacs with dominant galactic emission)

# Mind the Gap



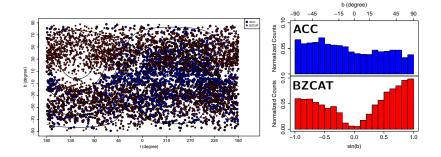
Blazar catalogs as complete as possible are necessary to provide candidate counterparts to unassociated  $\gamma$ -ray sources and to sources of high-energy neutrino emission or UHECRs.

Combining the ALMACAL (Bonato et al. 2018) with the ALMA Calibrator Source Catalogue, Bonato et al. (2019) built the ALMA Calibrator Catalogue (ACC): 3364 bright, compact radio sources with ALMA flux density measurements,  $\alpha_{low}$  between 1 (NVSS, Condon et al. 1998; SUMMS, Mauch et al. 2003) and 5 GHz (GB6, Gregory et al. 1996; PMN, Griffith & Wright 1993).

■ 3037 (~ 90%) blazar candidates:  $\alpha_{\rm low} <$  0.5, evidence of variability,  $\gamma$ -ray emission

A. Paggi

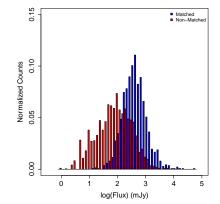
# BZCAT vs ACC



■ 
$$BZCAT$$
:  $|b| < 10^{\circ} \rightarrow 3\%$   
■ ACC (Dec  $\leq +60^{\circ}$ ):  $|b| < 10^{\circ} \rightarrow 15\%$ 

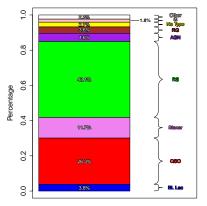
A New Multiwavelength Census of Blazars

INFN, Sezione di Torino November 17 2020 3340 BZCAT sources with Dec  $<60^\circ \rightarrow$  1391/1949 with/without match in ACC: 1.4/0.843 GHz radio flux from BZCAT.



We select the brightest radio end of the blazar population.

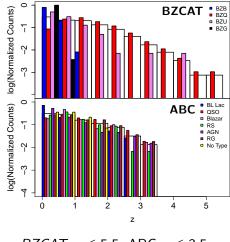
ACC 3037 blazar candidates - 1391 in  $BZCAT \rightarrow$  1646 sources. Literature (SIMBAD) classification.



Radio sources (**RS**, 43%), quasars (**QSO**, 26%), and **Blazars** (12%) are the main contributors. We exclude galaxies (**G**), galaxies in clusters, stars, etc. (Other)  $\rightarrow$  1580 sources (96%).

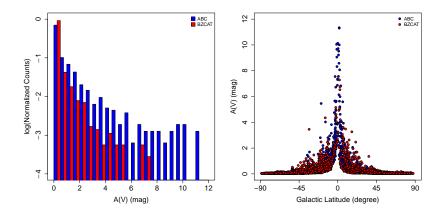
ALMA Blazar Candidates (ABC).

### **Redshift Distribution**



BZCAT  $z \leq 5.5$ , ABC  $z \leq 3.5$ .

# Absorption Distribution



ABC extends to lower galactic latitudes  $\rightarrow$  stronger absorption.

4<sup>th</sup> *Fermi* Large Area Telescope Source Catalog (4FGL, Abdollahi et al. 2020):

- 5064 sources 50 MeV 1 TeV
- 3474 associated (spatial coincidence) + 359 identified (periodic/correlated variability)
- 3137 blazars (62 % of 4FGL, 82 % of associated/identified):
  BLL (BL Lacs, 36 %), FSRQ (22 %), BCU (blazar of uncertain type, 42 %)
- 1231 unassociated (UGS)

259 ABC matches with associated/identified 4FGL sources (no UGSs).

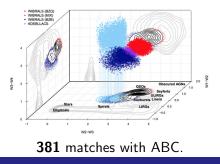
- 3<sup>rd</sup> catalogue of extreme and High-Synchrotron Peaked blazars (3HSP, Chang et al. 2019):
  - 2013 high-synchrotron peaked blazars (HSPs),  $\nu_p > 10^{15}$  Hz (selected with multi-wavelength analysis)
  - 657 in *BZCAT*, similar *b* distribution
    - 9 matches in ABC (7 BL Lacs, 1 RG and 1 No Type).

A. Paggi

# Catalog Comparison: WIBRaLS2

WISE Blazar-like Radio-Loud Sources (WIBRaLS2) catalog (D'Abrusco et al. 2019), candidate  $\gamma$ -ray blazars:

- Radio-loud sources from NVSS and SUMSS (22 µm/radio)
- 9541 sources with WISE colors compatible with γ-ray BZCAT blazars (detection in all 4 WISE bands, D'Abrusco et al. 2012)
- Classified depending on how compatible they are with BZB, BZQ, and MIXED regions



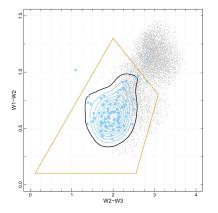
A. Paggi

A New Multiwavelength Census of Blazars

INFN, Sezione di Torino November 17 2020

# Catalog Comparison: KDEBLLACS

KDEBLLACS (D'Abrusco et al. 2019), candidate  $\gamma$ -ray **BL Lacs** selected using the similar criteria to WIBRaLS2:

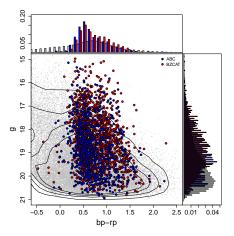


- radio-loud (12 µm/radio)
- 5579 sources selected to be compatible with 90%
   KDE level of BZCAT γ-ray BL Lacs (detection in first 3 WISE bands)

 $\mid b \mid > 10^{\circ}$ 

1 match with ABC (Blazar).

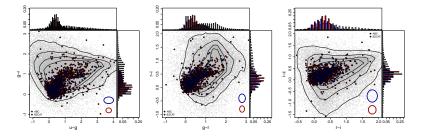
*Gaia* DR2 (Gaia Collaboration, 2018): **805** matches (null parallaxes and proper motions).



- ABC sources are less bright than BZCAT sources (p = 3 × 10<sup>-11</sup>) although spanning similar ranges of magnitudes
- ABC sources are redder than BZCAT sources ( $p = 1 \times 10^{-5}$ )

#### A. Paggi

### SDSS DR12 (Alam et al., 2015): 295 matches.



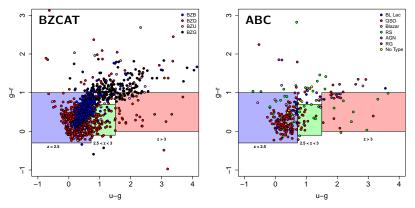
ABC sources are bluer than *BZCAT* sources, although with low significance (u-g:  $p = 6 \times 10^{-3}$ , g-r:  $p = 6 \times 10^{-3}$ , r-i:  $p = 4 \times 10^{-3}$ , i-z:  $p = 9 \times 10^{-5}$ ).

A AL A A 1.1 1

A. Paggi

# SDSS Cross-match

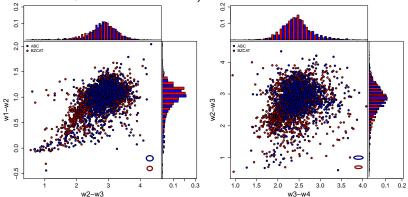
Butler & Bloom (2011): three regions to select low (z < 2.5), intermediate (2.5 < z < 3), and high (z > 3) redshift quasars.



Most most **BZQs/QSOs** and **BZBs/BL Lacs** fall into the region of low z quasars, BZGs outside.

## WISE Cross-match

AllWISE (Cutri et al., 2013): 1311 matches (906 detected in all four WISE bands, 739 with SNR > 3).



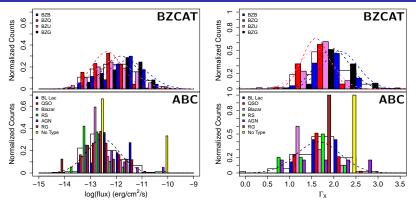
ABC sources have similar w1-w2 colors to *BZCAT* sources ( $p = 4 \times 10^{-2}$ ), while they appear bluer in the w2-w3 ( $p = 4 \times 10^{-4}$ ) and w3-w4 ( $p = 9 \times 10^{-11}$ ) colors.

Swift/XRT, Chandra/ACIS and XMM-Newton/EPIC observations:

- X-ray counterparts for 173 ABC sources
- 140 spectra for 92 sources
- 109 power-law spectra  $\rightarrow$  spectral slope  $\Gamma_X$
- 31 "complex" spectra for 24 sources (1 Blazar, 11 AGNs, 9 QSOs, 1 RS, 1 RG, 1 No Type)
- Fluxes: from spectra or from count-rates assuming  $\Gamma_X = 1.8$

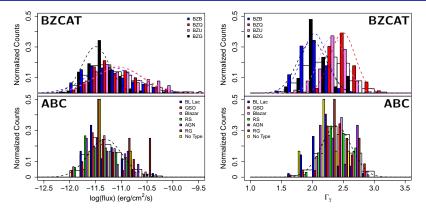
A. Paggi

# X-rays: Fluxes and Slopes



- ABC sources are dimmer (p = 3 × 10<sup>-16</sup>) in X-rays compared to BZCAT sources
- ABC sources show similar  $\Gamma_X$  to *BZCAT* sources (p = 0.1)and differ significantly from **BZBs**  $(p = 2 \times 10^{-13})$  and BZGs  $(p = 6 \times 10^{-5})$

### Gamma-rays: Fluxes and Slopes

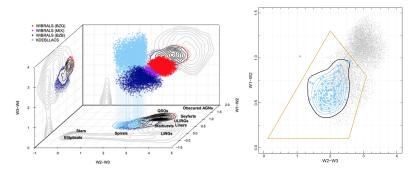


- ABC sources are dimmer (p = 1 × 10<sup>-5</sup>) in γ-rays compared to BZCAT sources
- ABC are softer than *BZCAT* sources ( $p = 4 \times 10^{-13}$ ) and show  $\Gamma_{\gamma}$  similar to BZUs ( $p = 4 \times 10^{-2}$ )

A. Paggi

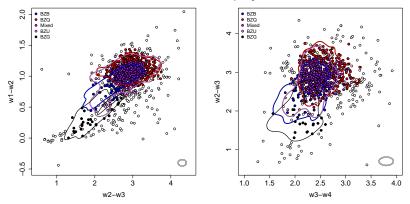
WIBRaLS2: WISE counterparts to  $\gamma$ -ray *BZCAT* blazars detected in all 4 bands, 3D locus.

KDEBLLACS: WISE counterparts to  $\gamma$ -ray *BZCAT* **BZB** detected in first 3 bands, 2D KDE.



### WISE Selection of $\gamma$ -ray Blazar Candidates

WISE counterparts to  $\gamma$ -ray *BZCAT* blazars detected in all 4 bands, 2D KDE contours in *both* color-color projection.

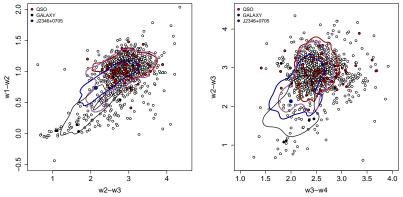


**697**  $\gamma$ -ray blazar candidates.

A. Paggi

### WISE Selection of $\gamma$ -ray Blazar Candidates

Optical spectra: 98 SDSS DR12 + 28 LAMOST DR5 (15 already in SDSS)  $\rightarrow$  111 optical spectra, 99 QSO + 12 GALAXY.

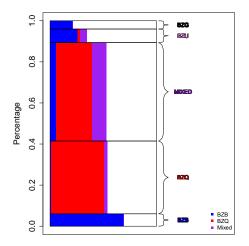


Most spectroscopic QSO are classified as BZQ and MIXED. Spectroscopic GALAXY are classified as BZB, BZU and BZG. J2346+0705 (z=2.62): GALAXY + BZB: EW < 1 Å.

INFN, Sezione di Torino

### Comparison with WIBRaLS2

ABC  $\gamma$ -ray blazar candidates as WIBRaLS2 candidates:



Not equivalent and complementary selections.

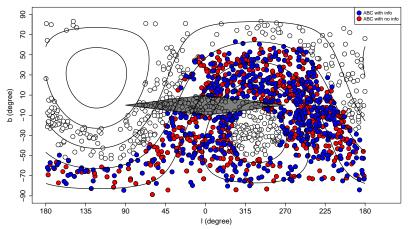
A. Paggi

# Conclusions

- ABC contains 1580 blazar candidates, filling the lack of  $|b|<10^\circ$  blazars in current catalogs
- Multi-wavelength informations Gaia DR2, SDSS DR12, LAMOST DR5, AllWISE, and 4FGL catalogs. X-ray data from Swift, Chandra and XMM-Newton archives
- ABC sources are brighter in the radio than BZCAT blazars by selection, while they appear dimmer in Gaia, X-rays and γ-rays
- ABC sources appear bluer than *BZCAT* blazars in SDSS and WISE
- ABC sources show similar X-ray spectral slopes to *BZCAT* blazars, while they appear softer in  $\gamma$ -rays
- $\blacksquare$  We classified 697 ABC WISE-detected sources as candidate  $\gamma\text{-ray}$  blazar
- Future work: investigate multi-wavelength correlations, optical spectral observations

### http://cdsarc.u-strasbg.fr/viz-bin/cat/J/A+A/641/A62

### Future Prospective: LSST @ Vera C. Rubin Observatory



WFD survey area: 556/529 ABC sources with/without additional info. 125 ABC sources in WFD survey area with SDSS DR12 counterpart: 114 (> 90%) with 24.16 > r > 16 (planned WFD median single-visit 5 $\sigma$  point sources depth and nominal LSST saturation limit for 15 s exposures).