## Machine Learning for BSM Searches A New Perspective

#### **Charanjit Kaur Khosa**



Theory Seminar, 4th November 2020

## Outline

- A. Why to use Machine Learning for BSM searches?
- B. Basics of Machine Learning methods
- C. Anomaly detection
- D. SMEFT analysis
- E. Summary

## Physics Beyond the SM

#### Standard Model: the mysteries

- Dark matter
- Neutrino masses
- Baryon Asymmetry of the Universe

#### Standard Model: the unexplained

- Flavour puzzle
- Hierarchy problem
- Fourth fundamental force (gravity) is not included

## Current Status of BSM

We have not seen any signal for New Physics (NP) so far!

#### Nature of BSM

- New Physics signature is beyond the reach of the current colliders
- Known BSM models do not include the "correct" model

#### How do we search for it?

- Traditional analysis strategies are not suitable
- Model dependent searches (pre-bias)

#### Need to go beyond these limitations

- 1. Large volumes of data
- 2. High dimensionality of the data sets
- 3. Large number of model parameters

"Hidden" correlations in the data can be explored using "powerful" ML techniques



#### ARTICLE

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## Searching for exotic particles in high-energy physics with deep learning

P. Baldi<sup>1</sup>, P. Sadowski<sup>1</sup> & D. Whiteson<sup>2</sup>

Collisions at high-energy particle colliders are a traditionally fruitful source of exotic particle discoveries. Finding these rare particles requires solving difficult signal-versus-background classification problems, hence machine-learning approaches are often used. Standard approaches have relied on 'shallow' machine-learning models that have a limited capacity to learn complex nonlinear functions of the inputs, and rely on a painstaking search through manually constructed nonlinear features. Progress on this problem has slowed, as a variety of techniques have shown equivalent performance. Recent advances in the field of deep learning make it possible to learn more complex functions and better discriminate between signal and background classes. Here, using benchmark data sets, we show that deep-learning methods need no manually constructed inputs and yet improve the classification metric by as much as 8% over the best current approaches. This demonstrates that deep-learning approaches can improve the power of collider searches for exotic particles.

#### Low-level features: $p_T^{l_1}, p_T^{l_2},$

 $p_{T}^{l_{1}}, p_{T}^{l_{2}}, \sum p_{T}^{j}, MET, N_{j}$ 

#### High-level features: Axial MET, $M_{T_2}$ , razor quantities



#### SUSY benchmark: chargino production (lepton+MET final state)

Technique	Low-level	High-level	Complete
AUC			
BDT	0.850 (0.003)	0.835 (0.003)	0.863 (0.003)
NN	0.867 (0.002)	0.863 (0.001)	0.875 (<0.001)
NN <sub>dropout</sub>	0.856 (<0.001)	0.859 (<0.001)	0.873 (<0.001)
DN	0.872 (0.001)	0.865 (0.001)	0.876 (<0.001)
DN <sub>dropout</sub>	0.876 (<0.001)	0.869 (<0.001)	0.879 (<0.001)
Discovery sigr	nificance		
NN	$6.5\sigma$	6.2 <i>o</i>	6.9 <i>o</i>
DN	$7.5\sigma$	7.3σ	7.6σ

BDT, boosted decision tree; DN, deep neural network; NN, shallow neural network; SUSY, supersymmetry particle.

#### Neural Networks in a Nutshell

## Neural Network (NN) Basic Structure

Training sample validation sample test (real) sample

- A. Input layer nodes: set of observables(kinematical features)/images
- B. Number of hidden layers (shallow or deep NN)
- C. Output layer: predictions



Schematic of a Neural Network

Train the network using training sample and make predictions for the test (real) dataset

### How To Train Your NN?



#### **NN Parameters and Concepts**



#### **Bias-Variance Trade-Off**

![](_page_9_Figure_1.jpeg)

### **Broad Categories of Machine Learning**

![](_page_10_Figure_1.jpeg)

## **Convolutional Neural Networks**

![](_page_11_Figure_1.jpeg)

![](_page_11_Figure_2.jpeg)

n3 units

#### Autoencoders

![](_page_12_Figure_1.jpeg)

#### Could be used as anomaly detector:

- 1. Train with the background sample.
- 2. Compare how the reconstructed output is different from the input (reconstructed error).

#### Reconstructed error will be more for the anomalous event.

T.Heimel, G.Kasieczka, T.Plehn and J.M.Thompson, SciPost Phys.6 (2019), 030 M. Farina, Y. Nakai and D. Shih, arXiv: 1808.08992

![](_page_13_Picture_0.jpeg)

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## Searching for exotic particles in high-energy physics with deep learning

# P. Baldi<sup>1</sup>, P. Sadowsk Deep Learning methods improve

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Background

BDT, boosted decision tree; DN, deep neural network; NN, shallow neural network; SUSY, supersymmetry particle.

## **More Recent Review**

nttps://doi.org/10.1038/s41586-018-0361-2

## Machine learning at the energy and intensity frontiers of particle physics

Alexander Radovic<sup>1</sup>\*, Mike Williams<sup>2</sup>\*, David Rousseau<sup>3</sup>, Michael Kagan<sup>4</sup>, Daniele Bonacorsi<sup>5,6</sup>, Alexander Himmel<sup>7</sup>,

## *"Machine Learning techniques"*

#### increase the discovery potential of

Table 1   the Higg	the	expe	rime	nts	study of
				1.9	

## REVIEW

## Machine learning at the energy and intensity frontiers of particle physics

Alexander Radovic<sup>1</sup>\*, Mike Williams<sup>2</sup>\*, David Rousseau<sup>3</sup>, Michael Kagan<sup>4</sup>, Daniele Bonacorsi<sup>5,6</sup>, Alexander Himmel<sup>7</sup>, Adam Aurisano<sup>8</sup>, Kazuhiro Terao<sup>4</sup> & Taritree Wongjirad<sup>9</sup>

Our knowledge of the fundamental particles of nature and their interactions is summarized by the standard model of particle physics. Advancing our understanding in this field has required experiments that operate at ever higher energies and intensities, which produce extremely large and information-rich data samples. The use of machine-learning techniques is revolutionizing how we interpret these data samples, greatly increasing the discovery potential of present and future experiments. Here we summarize the challenges and opportunities that come with the use of machine learning at the frontiers of particle physics.

Table 1   Effect of machine learning on the discovery and study of the Higgs boson					
Analysis	Years of data collection	Sensitivity without machine learning	Sensitivity with machine learning	Ratio of <i>P</i> values	Additional data required
$\frac{1}{CMS^{24}}$ $H \rightarrow \gamma \gamma$	2011–2012	2.2 $\sigma$ , $P = 0.014$	2.7 <i>σ</i> , <i>P</i> = 0.0035	4.0	51%
$\begin{array}{c} {\rm ATLAS^{43}} \\ {\rm H} \rightarrow \tau^+ \tau^- \end{array}$	2011–2012	2.5 $\sigma$ , P = 0.0062	3.4 $\sigma$ , $P = 0.00034$	18	85%
$ATLAS^{99}$ $VH \rightarrow bb$	2011–2012	$1.9\sigma, P = 0.029$	2.5 <i>σ</i> , <i>P</i> = 0.0062	4.7	73%
$ATLAS^{41}$ $VH \rightarrow bb$	2015–2016	2.8 <i>σ</i> , <i>P</i> = 0.0026	3.0 <i>о</i> , <i>P</i> = 0.00135	1.9	15%
$CMS^{100}$ $VH \rightarrow bb$	2011–2012	1.4 <i>σ</i> , <i>P</i> = 0.081	2.1 <i>σ</i> , <i>P</i> = 0.018	4.5	125%

#### A. Radovic et al., Nature 560(2018) no. 7716,41

## ML Techniques at Various Experimental Analysis Steps

- A. Decision-making for data storage
- B. Jet reconstruction and heavy flavor tagging
- C. Track reconstruction
- D. Signal and background classification
- E. Event generation

## New Potential for BSM searches

#### **Incomplete list of references**

- SMEFT (new physics deformations)
- Top tagging
- Anomaly detection
- Dark matter searches
- DNN likelihood
- Decoding black box
- Many more..

- J. Brehmer et al., Phys. Rev. Lett. 121 (2018) 111801, Phys. Rev. D 98 (2018) 052004, F.F. Freitas, CKK, V. Sanz, Phys. Rev. D 100 (2019) no.3, 035040 [arXiv:1902.05803 [hep-ph]].
- G.Kasieczka, T.Plehn, M.Russell and T.Schell, JHEP 1705 (2017) 006.
- arXiv: 1807.10261, Shih et al. 1808.08992[hep-ph]. Heimel et al., SciPost Phys.6 (2019), 030. **CKK** and Veronica Sanz, arXiv:2007.14462 [cs.LG] and others
- CKK, L. Mars, J. Richards and V. Sanz, J. Phys. G 47 (2020) no.9, 095201. CKK, V. Sanz and M. Soughton [arXiv:1910.06058 [hep-ph]].
- A.Coccaro, M.Pierini, L.Silvestrini and R. Torre, Eur. Phys. J. C 80 (2020) no.7, 664 [arXiv:1911.03305 [hep-ph]].
- G. Kasieczka, S. Marzani, G. Soyez and G.Stagnitto, JHEP 09 (2020), 195 [arXiv:2007.04319 [hep-ph]].
   T. Faucett, J. Thaler and D. Whiteson, [arXiv:2010.11998 [hep-ph]].

(Deep)Neural Networks, CNNs, (V)Autoencoders, Clustering, GNNs,..

Classification, Jet tagging, Anomalous Jet, Anomalous Events, Limit setting, Resonance,..

> Assisting BSM detection

### **Model Independent Searches**

![](_page_18_Figure_1.jpeg)

Image taken from 2001.04990[arxiv: hep-ph]

## Classification Without Labels (CWoLa)

![](_page_19_Figure_1.jpeg)

E.M.Metodiev, B.Nachman and J.Thaler, JHEP 10 (2017), 174, [arXiv:1708.02949 [hep-ph]]

## **Other Recent Developments**

- Learning New Physics from a Machine: R.T.D'Agnolo and A.Wulzer, Phys. Rev. D 99 (2019) no.1, 015014
- Guiding New Physics Searches with Unsupervised Learning: A. De Simone et al., Eur. Phys. J. C 79 (2019) no.4, 289, [arXiv:1807.06038].
- Using Variational Autoencoders: Cerri et al., JHEP 05 (2019), 036, [arXiv:1811.10276 [hep-ex]]. Cheng et al.,[arXiv:2007.01850].
- Uncovering latent jet substructure: B.M.Dillon et al., Phys. Rev. D 100 (2019) no.5, 056002)[arXiv:1904.04200].
- Tag N' Train: O.Amram and C.M.Suarez, [arXiv:2002.12376].
- Anti QCD tagger: J. A.Aguilar-Saavedra, J. H. Collins and R. K. Mishra, JHEP 11 (2017), 163, [arXiv:1709.01087].
- Anomaly Detection with Density Estimation: B. Nachman and D.Shih, Phys. Rev. D 101 (2020), 075042, [arXiv:2001.04990]
- Simulation Assisted Likelihood-free Anomaly Detection: A.Andreassen et al., Phys. Rev. D 101 (2020) no.9, 095004 [arXiv:2001.05001].

For other proposals, see recent review by Nachman: Anomaly Detection for Physics Analysis and Less than Supervised Learning[arXiv:2010.14554 [hep-ph]].

### **Creating Anomaly Aware Methods**

Anomaly Awareness (AA) a new algorithm for anomaly detection (C. K. Khosa & V. Sanz. 2020)

Which type of BSM we consider?

**Boosted Regime** 

What type of Input data?

Jet Images

Which model do we use?

CNN

Our algorithm learns about normal events (SM) while being made aware of an array of anomalies (BSM) in a way that it becomes sensitive to unseen BSM anomalies.

 Lets see how AA works in a well-known topology for new physics searches: Fat Jets

- Demonstrate its use against an array of BSM scenarios:
- 1.EFT Higgs,
- 2.Resonances  $\longrightarrow$  leading

jet with 2, 3 or 4 subsets

## The "Algorithm"

Algorithm 1 Anomaly Awareness (AA).
Important parameters are $\lambda_{AA}$ , $p_{An}^{min}$ , $p_{An}^{max}$ .
Prior Run
Initialize test: train splitting of Normal $(N)$ dataset
Initialize hyper parameters
Initialize Model (CNN architecture)
for Training over the epochs do
Cross entropy loss
Update model parameters.
end for
Get accuracy for $D_{test}$ and $D_{train}$
This run sets the hyper-parameters for the AA run
Anomaly Detection Run
Load the Anomaly (An) dataset
Initialize amount of data w.r.t. the Normal dataset
Initialize $\lambda_{AA}$
for Training over the epochs do
$l_1$ = Cross entropy loss (Normal dataset)
$l_2$ = Cross entropy loss (Anomaly dataset with Uniform
Distribution)
$Loss = l_1 + \lambda_{AA} l_2$
end for
Get softmax probabilities for all the datasets,
$p_i, i = N, An$
Select datapoints in a range $[p_{An}^{min}, p_{An}^{max}]$ ,
range optimized to select anomaly over normal events

## Top and QCD Jets

#### the input dataset

SM  $t\bar{t}$  and QCD diJet production,  $\sqrt{s} = 13 \ TeV$ 

Madgraph + Pythia

Leading jet with  $p_t > 750 \text{ GeV}$ , R=1 Anti-kt jet

 $\Delta\eta=0.087, \Delta\phi=0.087$ 

![](_page_23_Figure_6.jpeg)

![](_page_23_Figure_7.jpeg)

**Top Quark Decay** 

![](_page_23_Figure_8.jpeg)

Averaged over 50,000 events

#### CNN for Top vs QCD Classification (2C Baseline Classification)

![](_page_24_Figure_1.jpeg)

#### **Related work**

G.Kasieczka, T.Plehn, M.Russell and T.Schell, JHEP 05 (2017), 006 S.Macaluso and D.Shih, JHEP 10(2018), 121

#### **BSM Benchmarks**

![](_page_25_Figure_1.jpeg)

![](_page_26_Figure_0.jpeg)

## Anomaly Awareness for 2C Data

We see the effect of adding awareness to the classification task

![](_page_27_Figure_2.jpeg)

As we add more types of BSM examples, ALL BSMs gather in the centre Uniform Distribution over the baseline classes for all the BSM events

### **Robust Anomaly Detector**

![](_page_28_Figure_1.jpeg)

### Baseline vs AA Comparison

![](_page_29_Figure_1.jpeg)

The addition of the AA term does not degrade the baseline classification but adds the ability to use its output for anomaly detection

## Signal Cross-section Reach

![](_page_30_Picture_1.jpeg)

We scan on windows of the classifier output Cutting a small window around 0.5 anomaly detection is enhanced

We use  $S/\sqrt{B}$  as an example of quantity to maximise (S=BSM, B=SM)

![](_page_30_Figure_4.jpeg)

![](_page_30_Figure_5.jpeg)

![](_page_30_Figure_6.jpeg)

#### Three-class Example

This procedure can be generalized beyond binary classification

Top Jet, QCD jet, W-jet

150 K images (balanced data set), training:test=70:30%

**Prior Run** 

With AA

![](_page_31_Figure_6.jpeg)

**Unseen Data Set: EFT** 

### Key Message

- All methods have their "advantages" and "limitations".
- No "Generic" method: Each method operates in a limited region of parameter space.
- "None" of the current methods are suitable for whole phase space and can not target the anomalous events of all types.
- Simulated benchmarks are also different, so it is "non-trivial" task to compare one with other.

![](_page_32_Picture_5.jpeg)

LHC Olympics 2020 was organised to address this issue https://lhco2020.github.io/homepage/

#### SMEFTs

Model independent framework to parametrize the new physics Higher dimensional operators respecting SM symmetries and involving SM fields

$$\mathscr{L}^{d=6} = \mathscr{L}_{SM} + \sum_{i} \frac{C_i}{\Lambda^2} \mathcal{O}_i$$

SMEFT fit to Higgs, diboson and EW data

$$\begin{split} \mathcal{L}_{\mathrm{SMEFT}}^{\mathrm{SILH}} &\supset \frac{\bar{c}_W}{m_W^2} \frac{ig}{2} \left( H^{\dagger} \sigma^a \overset{\leftrightarrow}{D^{\mu}} H \right) D^{\nu} W_{\mu\nu}^a + \frac{\bar{c}_B}{m_W^2} \frac{ig'}{2} \left( H^{\dagger} \overset{\leftrightarrow}{D^{\mu}} H \right) \partial^{\nu} B_{\mu\nu} + \frac{\bar{c}_T}{v^2} \frac{1}{2} \left( H^{\dagger} \overset{\leftrightarrow}{D_{\mu}} H \right)^2 \\ &\quad + \frac{\bar{c}_U}{v^2} (\bar{L} \gamma_{\mu} L) (\bar{L} \gamma^{\mu} L) + \frac{\bar{c}_{He}}{v^2} (iH^{\dagger} \overset{\leftrightarrow}{D_{\mu}} H) (\bar{e}_R \gamma^{\mu} e_R) + \frac{\bar{c}_{Hu}}{v^2} (iH^{\dagger} \overset{\leftrightarrow}{D_{\mu}} H) (\bar{u}_R \gamma^{\mu} u_R) \\ &\quad + \frac{\bar{c}_{Hd}}{v^2} (iH^{\dagger} \overset{\leftrightarrow}{D_{\mu}} H) (\bar{d}_R \gamma^{\mu} d_R) + \frac{\bar{c}'_{Hq}}{v^2} (iH^{\dagger} \sigma^a \overset{\leftrightarrow}{D_{\mu}} H) (\bar{Q}_L \sigma^a \gamma^{\mu} Q_L) \\ &\quad + \frac{\bar{c}_{Hq}}{v^2} (iH^{\dagger} \overset{\leftrightarrow}{D_{\mu}} H) (\bar{Q}_L \gamma^{\mu} Q_L) + \frac{\bar{c}_{HW}}{m_W^2} ig (D^{\mu} H)^{\dagger} \sigma^a (D^{\nu} H) W_{\mu\nu}^a + \frac{\bar{c}_{HB}}{m_W^2} ig' (D^{\mu} H)^{\dagger} (D^{\nu} H) B_{\mu\nu} \\ &\quad + \frac{\bar{c}_{3W}}{m_W^2} g^3 \epsilon_{abc} W_{\mu}^{a\nu} W_{\nu\rho}^b W^{c\,\rho\mu} + \frac{\bar{c}_g}{m_W^2} g_s^2 |H|^2 G_{\mu\nu}^A G^{A\mu\nu} + \frac{\bar{c}_\gamma}{m_W^2} g'^2 |H|^2 B_{\mu\nu} B^{\mu\nu} \\ &\quad + \frac{\bar{c}_H}{v^2} \frac{1}{2} (\partial^{\mu} |H|^2)^2 + \sum_{f=e,u,d} \frac{\bar{c}_f}{v^2} y_f |H|^2 \bar{F}_L H^{(c)} f_R \\ &\quad + \frac{\bar{c}_{3G}}{m_W^2} g_s^3 f_{ABC} G_{\mu}^{A\nu} G_{\nu}^{A\rho} G_{\rho}^{C\mu} + \frac{\bar{c}_{uG}}{m_W^2} g_s y_u \bar{Q}_L H^{(c)} \sigma^{\mu\nu} \lambda_A u_R G_{\mu\nu}^A \end{split}$$

J.Ellis, C.W.Murphy, V.Sanz and T.You JHEP06(2018)146

## SMEFT : Global Analysis

 Precision electroweak data, LHC Run 1 & 2 data (Higgs production, pair of gauge bosons)

Observable	Measurement	Ref.	SM Prediction	Ref.
$\Gamma_Z \; [\text{GeV}]$	$2.4952 \pm 0.0023$	[41]	$2.4943 \pm 0.0005$	[40]
$\sigma_{\rm had}^0 \; [{\rm nb}]$	$41.540 \pm 0.037$	[41]	$41.488 \pm 0.006$	[40]
$R^0_\ell$	$20.767 \pm 0.025$	[41]	$20.752 \pm 0.005$	[40]
$A_{ m FB}^{0,\ell}$	$0.0171 \pm 0.0010$	[41]	$0.01622 \pm 0.00009$	[120]
$\mathcal{A}_{\ell}\left(P_{\tau}\right)$	$0.1465 \pm 0.0033$	[41]	$0.1470 \pm 0.0004$	[120]
$\mathcal{A}_{\ell}\left(\mathrm{SLD}\right)$	$0.1513 \pm 0.0021$	[41]	$0.1470 \pm 0.0004$	[120]
$R_b^0$	$0.021629 \pm 0.00066$	[41]	$0.2158 \pm 0.00015$	[40]
$R_c^0$	$0.1721 \pm 0.0030$	[41]	$0.17223 \pm 0.00005$	[40]
$A_{ m FB}^{0,b}$	$0.0992 \pm 0.0016$	[41]	$0.1031 \pm 0.0003$	[120
$A_{ m FB}^{0,c}$	$0.0707 \pm 0.0035$	[41]	$0.0736 \pm 0.0002$	[120
$\mathcal{A}_b$	$0.923 \pm 0.020$	[41]	0.9347	[120]
$\mathcal{A}_{c}$	$0.670\pm0.027$	[41]	$0.6678 \pm 0.0002$	[120
$M_W$ [GeV]	$80.387 \pm 0.016$	[42]	$80.361 \pm 0.006$	[120]
$M_W$ [GeV]	$80.370 \pm 0.019$	[100]	$80.361 \pm 0.006$	[120

LEP data + WW LEP2 data + Mw Tevatron

J.Ellis, C.W.Murphy, V.Sanz and T.You JHEP06(2018)146

## LHC data

Production	Decay	Signal Strength	Production	Decay	Signal Strength
$gg\mathrm{F}$	$\gamma\gamma$	$1.10\substack{+0.23\\-0.22}$	Wh	au au	$-1.4 \pm 1.4$
$gg\mathrm{F}$	ZZ	$1.13\substack{+0.34\\-0.31}$	Wh	bb	$1.0 \pm 0.5$
ggF	WW	$0.84\pm0.17$	Zh	$\gamma\gamma$	$0.5^{+3.0}_{-2.5}$
ggF	au au	$1.0 \pm 0.6$	Zh	WW	$5.9^{+2.6}_{-2.2}$
VBF	$\gamma\gamma$	$1.3 \pm 0.5$	Zh	au au	$2.2^{+2.2}_{-1.8}$
VBF	ZZ	$0.1^{+1.1}_{-0.6}$	Zh	bb	$0.4 \pm 0.4$
VBF	WW	$1.2 \pm 0.4$	tth	$\gamma\gamma$	$2.2^{+1.6}_{-1.3}$
VBF	au au	$1.3 \pm 0.4$	tth	WW	$5.0^{+1.8}_{-1.7}$
Wh	$\gamma\gamma$	$0.5^{+1.3}_{-1.2}$	tth	au au	$-1.9^{+3.7}_{-3.3}$
Wh	WW	$1.6^{+1.2}_{-1.0}$	tth	bb	$1.1 \pm 1.0$
pp	$Z\gamma$	$2.7^{+4.6}_{-4.5}$	pp	$\mu\mu$	$0.1 \pm 2.5$

#### Run 1 data

	Production	Decay	Sig. Stren.		Production	Decay	Sig. Stren.
[102]	1-jet, $p_T > 450$	$b\overline{b}$	$2.3^{+1.8}_{-1.6}$	[110]	pp	$\mu\mu$	$-0.1\pm1.5$
[103]	Zh	$b\overline{b}$	$0.9\pm0.5$	[111]	Zh	$b\overline{b}$	$1.12^{+0.50}_{-0.45}$
[103]	Wh	$b\overline{b}$	$1.7\pm0.7$	[111]	Wh	$b\overline{b}$	$1.35^{+0.68}_{-0.59}$
[104]	$t\bar{t}h, \ge 1\ell$	$b\overline{b}$	$0.72\pm0.45$	[112]	$t\bar{t}h$	$b\overline{b}$	$0.84^{+0.64}_{-0.61}$
[105]	$t\bar{t}h$	$1\ell + 2\tau_h$	$-1.52^{+1.76}_{-1.72}$	[113]	$t\bar{t}h$	$2\ell os + 1\tau_h$	$1.7^{+2.1}_{-1.9}$
[105]	$t\bar{t}h$	$2\ell ss + 1\tau_h$	$0.94\substack{+0.80\\-0.67}$	[113]	$t\bar{t}h$	$1\ell + 2\tau_h$	$-0.6^{+1.6}_{-1.5}$
[105]	$t\bar{t}h$	$3\ell + 1\tau_h$	$1.34^{+1.42}_{-1.07}$	[113]	$t\bar{t}h$	$3\ell + 1\tau_h$	$1.6^{+1.8}_{-1.3}$
[ <b>105</b> ]	$t\bar{t}h$	$2\ell ss$	$1.61\substack{+0.58\\-0.51}$	[113]	$t\bar{t}h$	$2\ell ss + 1\tau_h$	$3.5^{+1.7}_{-1.3}$
[105]	$t\bar{t}h$	$3\ell$	$0.82^{+0.77}_{-0.71}$	[113]	$t\bar{t}h$	$3\ell$	$1.8^{+0.9}_{-0.7}$
[105]	$t\bar{t}h$	$4\ell$	$0.9^{+2.3}_{-1.6}$	[113]	$t\bar{t}h$	$2\ell ss$	$1.5^{+0.7}_{-0.6}$
[106]	0-jet DF	WW	$1.30\substack{+0.24\\-0.23}$	[114]	$gg\mathrm{F}$	WW	$1.21\substack{+0.22\\-0.21}$
[106]	1-jet DF	WW	$1.29^{+0.32}_{-0.27}$	[114]	VBF	WW	$0.62^{+0.37}_{-0.36}$
[106]	2-jet DF	WW	$0.82^{+0.54}_{-0.50}$	[115]	${ m B}(h  o \gamma \gamma)/ \ { m B}(h$	$\rightarrow 4\ell$ )	$0.69\substack{+0.15\\-0.13}$
[ <b>106</b> ]	VBF 2-jet	WW	$0.72\substack{+0.44 \\ -0.41}$	[115]	0-jet	$4\ell$	$1.07\substack{+0.27 \\ -0.25}$
[ <b>106</b> ]	Vh 2-jet	WW	$3.92^{+1.32}_{-1.17}$	[115]	1-jet, $p_T < 60$	$4\ell$	$0.67\substack{+0.72 \\ -0.68}$
[ <b>106</b> ]	Wh 3-lep	WW	$2.23^{+1.76}_{-1.53}$	[115]	1-jet, $p_T \in (60, 120)$	$4\ell$	$1.00\substack{+0.63\\-0.55}$
[ <b>107</b> ]	$gg\mathrm{F}$	$\gamma\gamma$	$1.10\substack{+0.20 \\ -0.18}$	[115]	1-jet, $p_T \in (120, 200)$	$4\ell$	$2.1^{+1.5}_{-1.3}$
[107]	VBF	$\gamma\gamma$	$0.8^{+0.6}_{-0.5}$	[115]	2-jet	$4\ell$	$2.2^{+1.1}_{-1.0}$
[ <b>107</b> ]	$t\bar{t}h$	$\gamma\gamma$	$2.2^{+0.9}_{-0.8}$	[115]	"BSM-like"	$4\ell$	$2.3^{+1.2}_{-1.0}$
[ <b>107</b> ]	Vh	$\gamma\gamma$	$2.4^{+1.1}_{-1.0}$	[115]	VBF, $p_T < 200$	$4\ell$	$2.14_{-0.77}^{+0.94}$
[ <b>108</b> ]	$gg\mathrm{F}$	$4\ell$	$1.20\substack{+0.22\\-0.21}$	[115]	$Vh \ \mathrm{lep}$	$4\ell$	$0.3^{+1.3}_{-1.2}$
[109]	0-jet	au au	$0.84\pm0.89$	[115]	$t\bar{t}h$	$4\ell$	$0.51^{+0.86}_{-0.70}$
[109]	boosted	au au	$1.17\substack{+0.47\\-0.40}$	[116]	Wh	WW	$3.2^{+4.4}_{-4.2}$
[109]	VBF	au au	$1.11\substack{+0.34\\-0.35}$				
[106]	Zh 4-lep	WW	$0.77^{+1.49}_{-1.20}$				

#### (Early) Run 2 data + STXS

![](_page_36_Figure_0.jpeg)

J.Ellis, C.W.Murphy, V.Sanz and T.You JHEP06(2018)146

A.Biekötter, T.Corbett and T.Plehn, arXiv:1812.07587 [hep-ph] E.da Silva Almeida et al.,Phys.Rev.D 99(2019)

#### Why VH channel?

VBF channel

J. Brehmer, K. Cranmer, G. Louppe and J Pavez, Phys. Rev. Lett. 121 (2018) 111801, Phys. Rev. D 98 (2018) 052004

#### VH channel (higher statistics)

Felipe F. Freitas, CKK, Veronica Sanz, arXiv: 1902.05803 [hep-ph]

#### SMEFT via VH channel

![](_page_38_Figure_1.jpeg)

Specific operator which produces it

Felipe F. Freitas, CKK, Veronica Sanz, arXiv: 1902.05803 [hep-ph]

NN.

 $V_{\nu}$ 

#### Analysis set-up (VH Channel)

 $p_T^{b_1}, p_T^{b_2}, p_T^{VH}, M_T^{VH}, p_T^{W/Z}, p_T^H, \eta^H, \phi^H$ 

0-lepton	$pp \rightarrow HZ, (H \rightarrow b\bar{b}, Z \rightarrow \nu\bar{\nu})$	$MET, \Delta \phi_{b_1 MET}$
1-lepton	$pp \rightarrow HW, (H \rightarrow b\bar{b}, W \rightarrow lv_l)$	$M_T^W, p_T^l, MET, \Delta R_{wl}, \Delta \phi_{b_1l}, \Delta \phi_{lMET}$
2-lepton	$pp \to HZ, (H \to b\bar{b}, Z \to l^+l^-)$	$p_T^{l_1}, p_T^{l_2}, \Delta R_{ll}, \Delta \phi_{b_1 l_1}, \Delta \phi_{b_2 l_1}$

Feynrules Model : Higgs Effective Lagrangian arXiv:1310.5150

 $\sqrt{s} = 14$  TeV, 100K events for both SMEFT and SM using MC@NLO Madgraph

LO, parton level analysis

Channel	Inclusive
0L	$E_T > 150 \text{ GeV}$
1L	$p_T^l > 25 \text{ GeV},  \eta_l  < 2.7$
	$E_T > 30 \text{ GeV}, p_T^V > 150 \text{ GeV}$
2L	$ p_T^l > 7 \text{ GeV},  \eta_l  < 2.7, p_T^V > 75 \text{ GeV}$
	Leading lepton $p_T > 27 \text{ GeV}$
0L, 1L, 2L	$p_T^b > 20 \text{ GeV},  \eta_b  < 2.5,$
	Leading b-jet $p_T > 45 \text{ GeV}$

M. Aaboud et al., Phys. Lett. B 786(2018) 59

Felipe F. Freitas, CKK, Veronica Sanz, arXiv: 1902.05803 [hep-ph]

#### 1D and 2D features

![](_page_40_Figure_1.jpeg)

#### Neural Network

![](_page_41_Figure_1.jpeg)

#### **Neural Network Architecture**

- Training set: Test set = 70 %: 30% (data scaling)
- Hidden layers: 1 (optimised)
- Activation function: ReLu
- Dropouts: 0.2
- Loss function: Asimov loss function (pre-training with MSE)

# ROC curve (SMEFT vs SM Higgs background)

![](_page_43_Figure_1.jpeg)

AUC: area under ROC

# Appropriate performance measure for HEP analysis

Asimov Significance

$$Z_A = \left[ 2 \left( (s+b) \ln \left[ \frac{(s+b)(b+\sigma_b^2)}{b^2+(s+b)\sigma_b^2} \right] - \frac{b^2}{\sigma_b^2} \ln \left[ 1 + \frac{\sigma_b^2 s}{b(b+\sigma_b^2)} \right] \right) \right]^{1/2}$$
$$s = W_s \sum_{i}^{N_{batch}} y_i^{pred} \times y_i^{true} \qquad b = W_b \sum_{i}^{N_{batch}} y_i^{pred} \times (1 - y_i^{true})$$

Specific(Asimov) Loss function

$$\ell_{Asimov} = 1/Z_A$$

Adam Elwood and Dirk Krücker, arXiv: 1806.00322[hep-ex]

Glen Cowan, Kyle Cranmer, Eilam Gross, Ofer Vitells, arXiv:1007.1727[physics.data-an]

## Classifier Output : OL channel $(C_{HW} = 0.03)$

![](_page_45_Figure_1.jpeg)

#### Current limit at 95% CL

![](_page_46_Figure_1.jpeg)

Irreducible background only

# Combining 0L+1L+2L for a limiting case

![](_page_47_Figure_1.jpeg)

Irreducible background only

For a realistic analysis combining different channels may help

## Summary and Outlook

- ML techniques are emerging as a competitive tool to look for new phenomena in the complex data
- HEP community is adapting these techniques for various tasks: trigger, heavy flavour tag, quark gluon discrimination, jet tagging etc.
- There is lot of activity to build methods for anomaly detection.
- We present a new algorithm for anomaly detection. It is based on the procedure of classifying 'normal' (SM) events, While the algorithm is made aware of the presence of anomalies (BSM) through a modification of the learning function.
- We used supervised learning techniques to exploit kinematic information in VH channel for SMEFT framework. This approach may provide a significant stronger bounds on EFT coefficients (scalability for more operators, realistic simulation).
- Finally, we need to test/use these approaches for LHC data.

#### Thanks

![](_page_49_Figure_0.jpeg)

![](_page_49_Figure_1.jpeg)

![](_page_49_Figure_2.jpeg)

## Generative Adversarial Networks (GANs)

Can interpolate the phase space for the event generation

![](_page_50_Figure_2.jpeg)

Aim is to see if GANs could be used for fast simulations (dijetGAN, arxiv:1903.02433)