

Particle Dark Matter Lecture 3

Useful relations:

$$\Omega = \frac{\rho}{\rho_c} \quad \rho_c = \frac{3H^2}{8\pi G_N} = 11,23 h^2 m_p / m^3$$

here $H_0 = h \text{ } 100 \text{ km/s/Mpc}$

$m_p = \text{proton mass} \sim 1 \text{ GeV}$

$h \sim 0.67 - 0.72$

independent
of h !

$$\Omega h^2 = \frac{\rho h^2}{\rho_c} = \frac{\rho}{11,23 m_p / m^3}$$

Unitarity bound: $\sigma_{\text{non-el}, J} = \frac{4\pi (2J+1)}{(2s_p+1)^2} \frac{(1 - \eta_J^2)}{p_i^2}$

i.e. maximal σ for $J = s_p = 0$

$$\sigma_{\text{non-el}}^{\text{max}} = \frac{4\pi}{p_i^2} = \frac{4\pi}{m_{\text{DM}}^2 v^2} \quad \text{i.e. } \sigma_{\text{non-el}}^{\text{max}} \sim \frac{4\pi}{m_{\text{DM}}^2 v}$$

giving $\langle \sigma_{\text{non-el}}^{\text{max}} v \rangle_X \sim \frac{4\pi}{m_{\text{DM}}^2} \langle \frac{1}{v} \rangle_X \sim \frac{4\pi}{m_{\text{DM}}^2} X^{-1/2}$

So we have

$$\Omega_{DM} h^2 = \frac{m_{DM} Y_{DM} s(T_x)}{8\pi/h^2} = \frac{m_{DM}}{8\pi/h^2} \frac{H(m_{DM}) s(T_{now})}{s(m_{DM})} \frac{\sqrt{6} m_{DM}^2}{4\pi} \frac{5}{2} x_f^{3/2}$$

$$= 0.04 \times 10^{-10} \left(\frac{m_{DM}}{\text{GeV}} \right)^2 x_f^{3/2}$$

$$\Rightarrow m_{DM} < 10^5 \text{ GeV} x_f^{-3/4} \left(\frac{\Omega_{DM} h^2}{0.04} \right)^{1/2}$$

$$\Rightarrow m_{DM} < 12 \text{ TeV for } x_f \sim 30$$

More complex picture:

* coannihilation: particle j decoupling at the same time as DM

Coupled Boltzmann eqs:

$$\frac{dn_i}{dt} + 3Hn_i = - \sum_j \langle \sigma_{ij} v_{ij} \rangle (n_i n_j - n_i^{eq} n_j^{eq})$$



$$- \sum_{j \neq i} \langle \sigma'_{ij} v_{ij} \rangle (n_i n_{SM} - n_i^{eq} n_{SM}^{eq})$$

+ decays / inverse decays

But consider the sum: $n_j + n_{DM} = n$

$$\Rightarrow \frac{dn}{dt} + 3Hn = -\langle \sigma_{\text{eff}} v \rangle (n^2 - (n^{\text{eq}})^2)$$

where $\langle \sigma_{\text{eff}} v \rangle = \sum_{i,j} \langle \sigma_{ij} v_{ij} \rangle \frac{n_i^{\text{eq}} n_j^{\text{eq}}}{(n^{\text{eq}})^2}$

assuming conversion in equilibrium

