

Summary of reconstruction and analysis

E. Di Marco

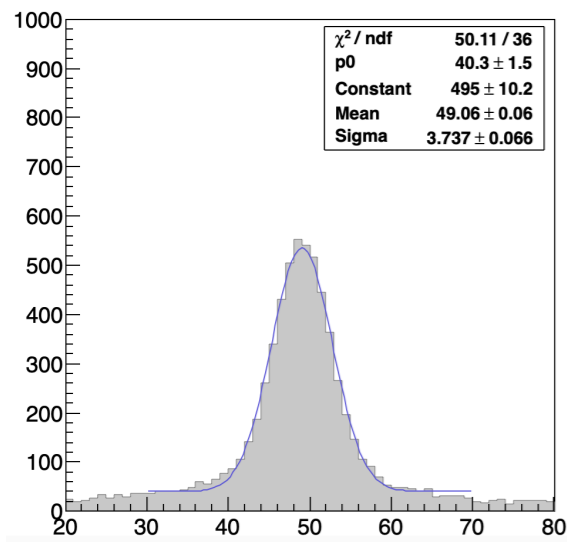
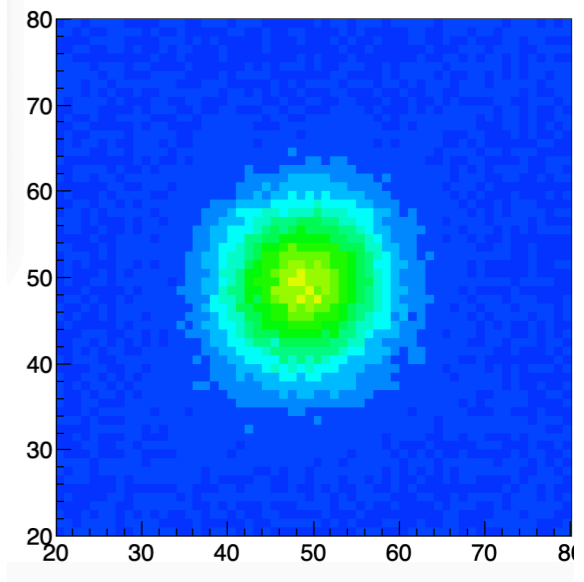
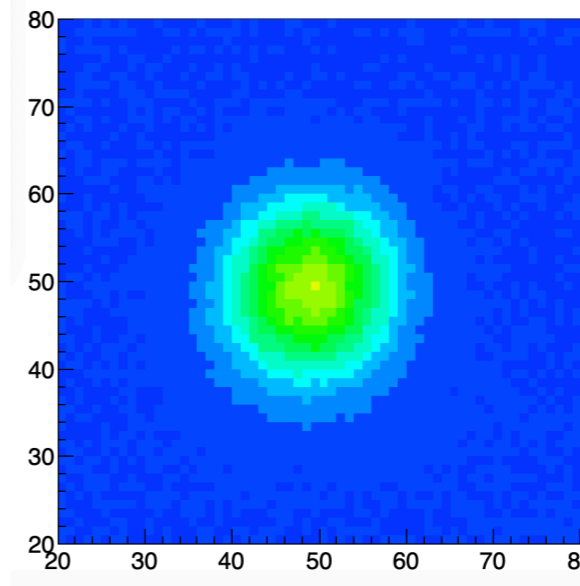
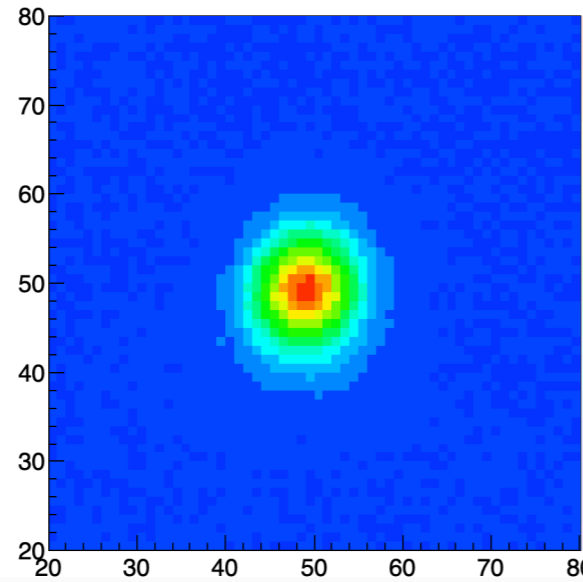
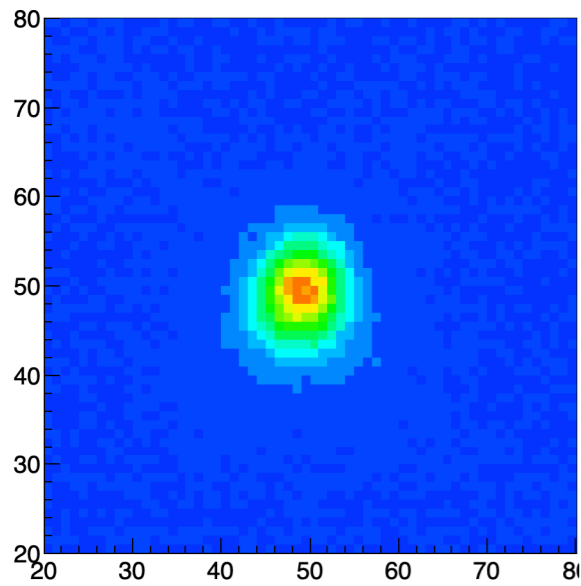
CYGN0 meeting,

10 December 2020

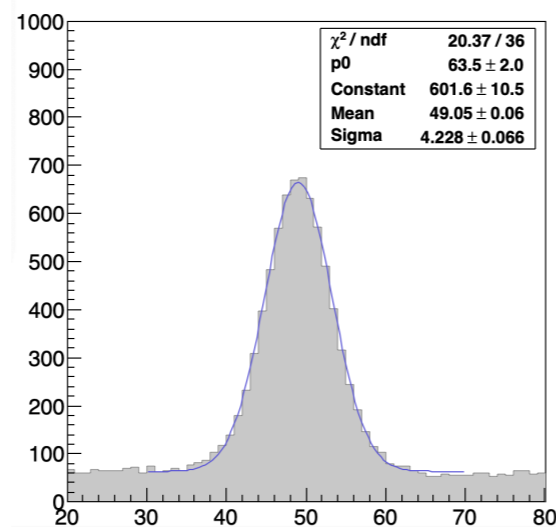
saturation studies

^{55}Fe spots at increasing distance from GEMs: transverse, but also longitudinal diffusion helps against saturation [the spot is a sphere in 3D]

D. Pinci

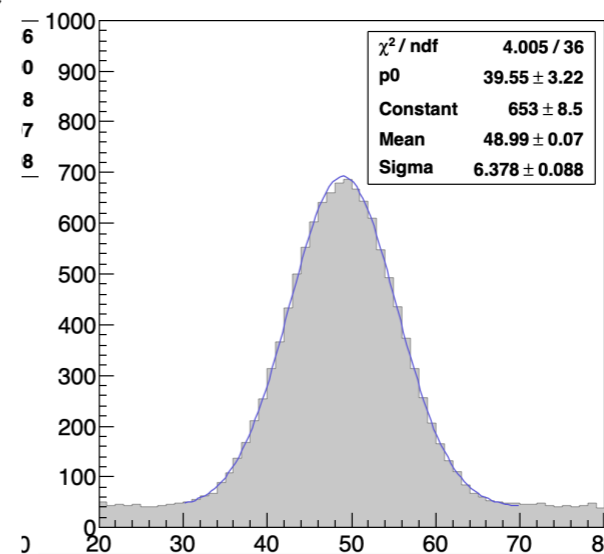


6 cm - light = 4635

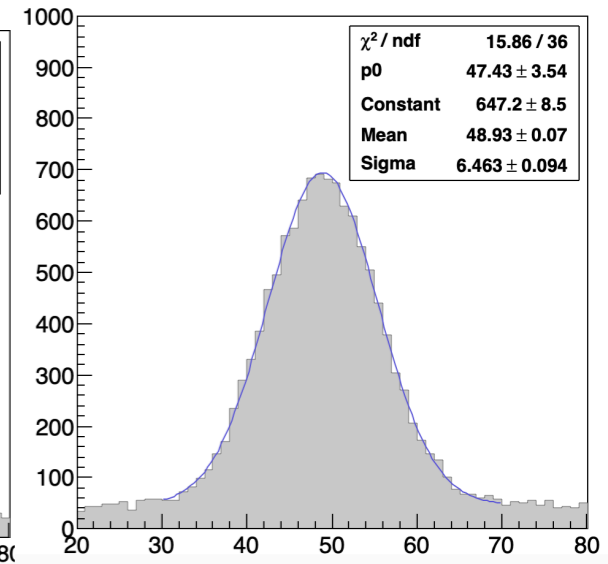


11 cm - light = 6373

...



41 cm - 10436

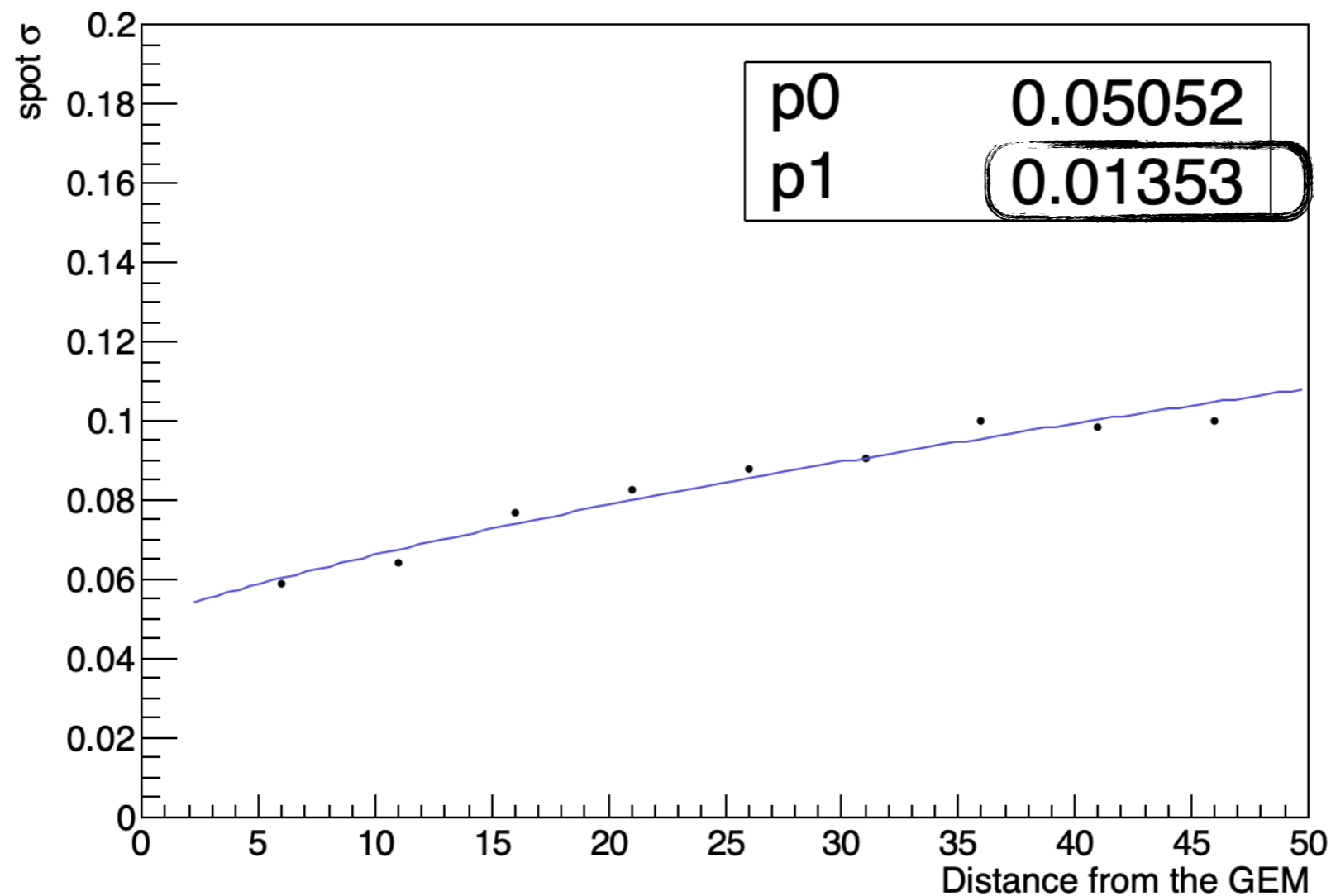


46 cm - 10480

comparison with sim

transverse diffusion 0.14 mm/cm similar to value predicted from simulation (0.12mm/cm)

D. Pinci

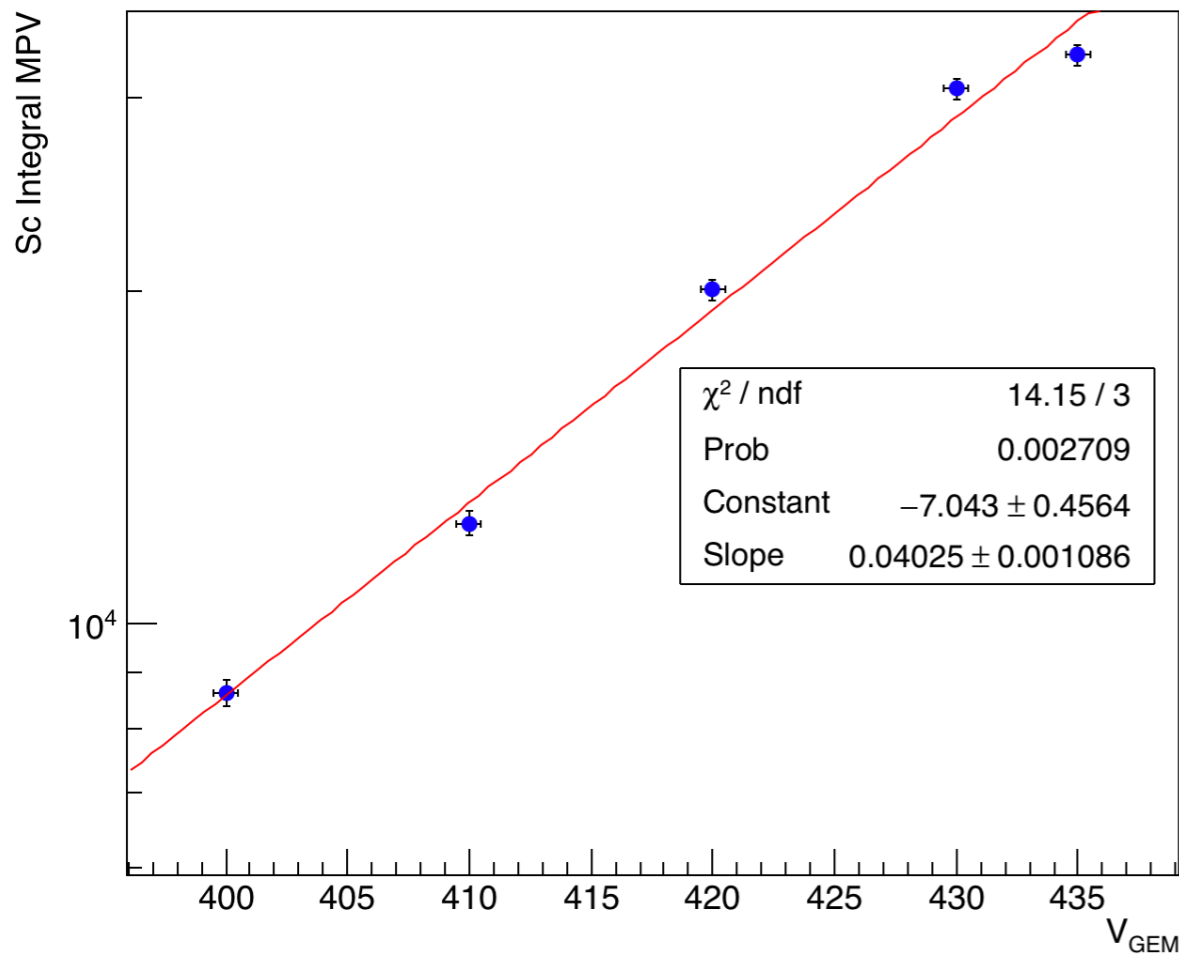


V_{GEM} SCAN WITH MANGO

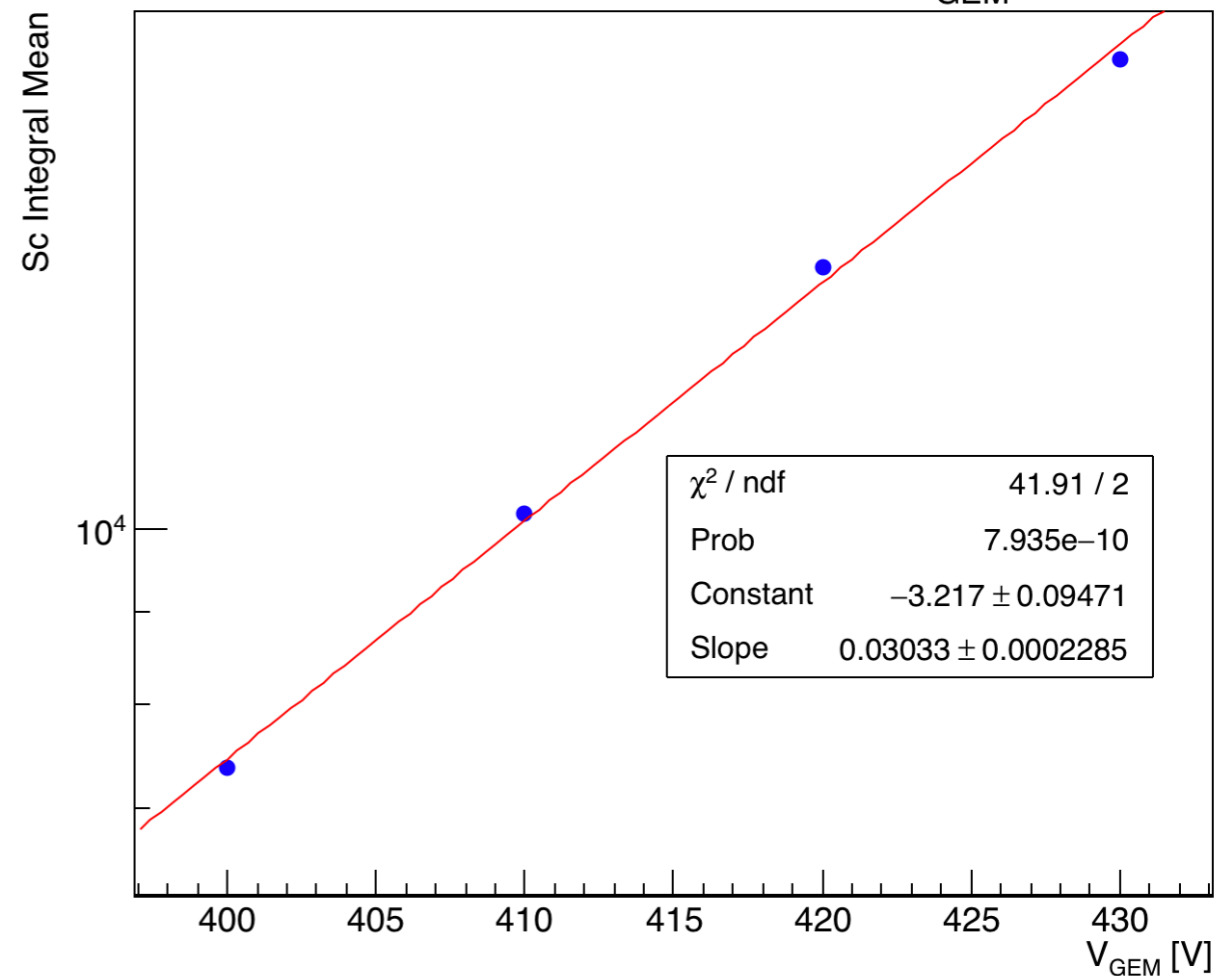
Initially used ¹³³Ba source, then got ⁵⁵Fe source and performed a V_{GEM} scan.

E. Baracchini et al.

¹³³Ba Sc Integral MPV vs V_{GEM}



⁵⁵Fe Sc Integral Mean vs V_{GEM}

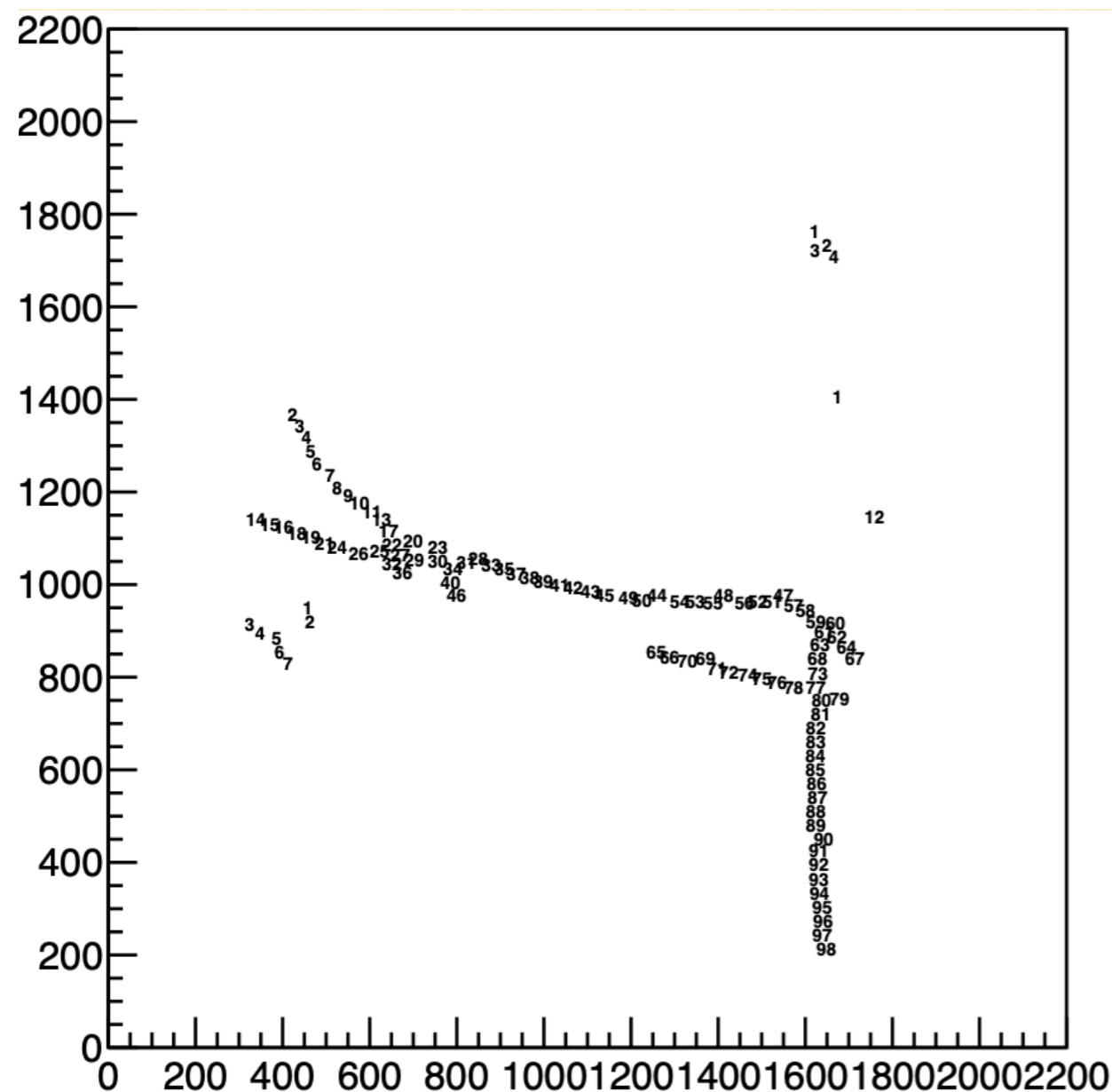


Expected behavior observed

Head-Tail studies

Ongoing preparatory development for head-tail studies:
improving the algorithm to follow the pattern of curly /
branched tracks

S. Torelli



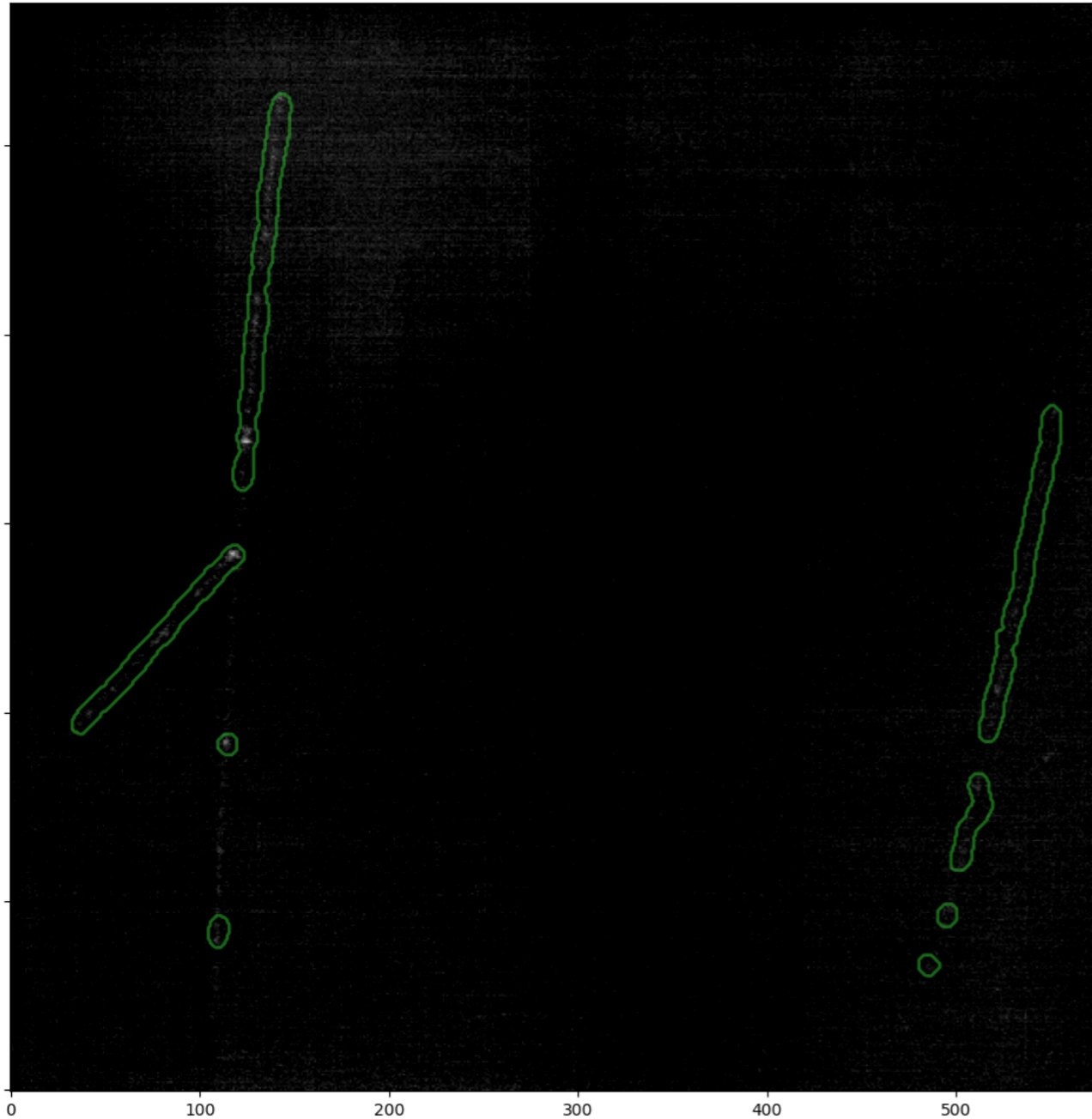
For long tracks, developing
an algorithm to better
handle branches in the
track / overlapping tracks

For the time being, use
what is there in the reco.
Focus on the simple tracks,
reducing the slices sizes.

AmBe @ LIME update

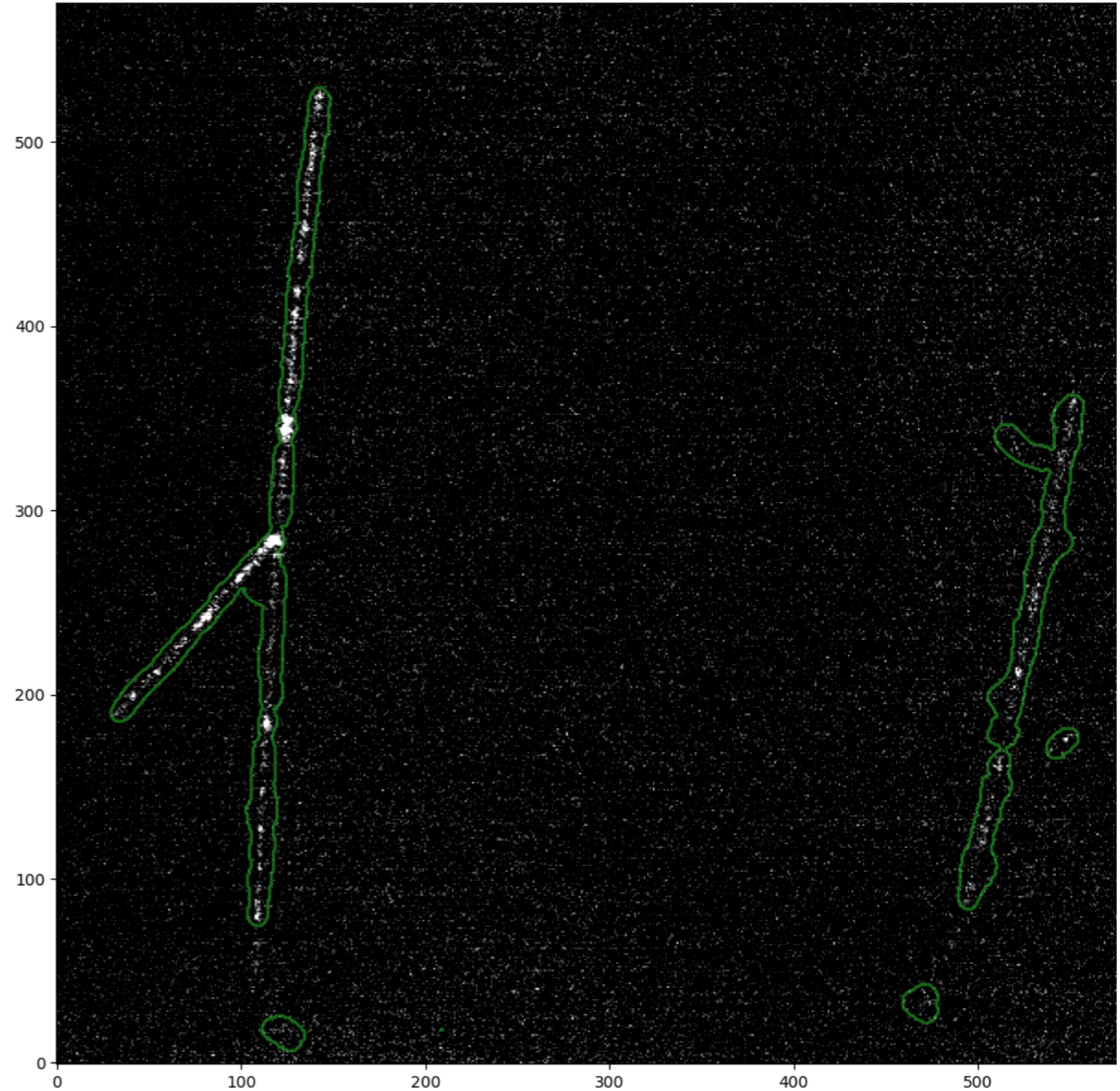
Improved noise filtering applying it to full-resolution images

Superclusters found



old noise filtering

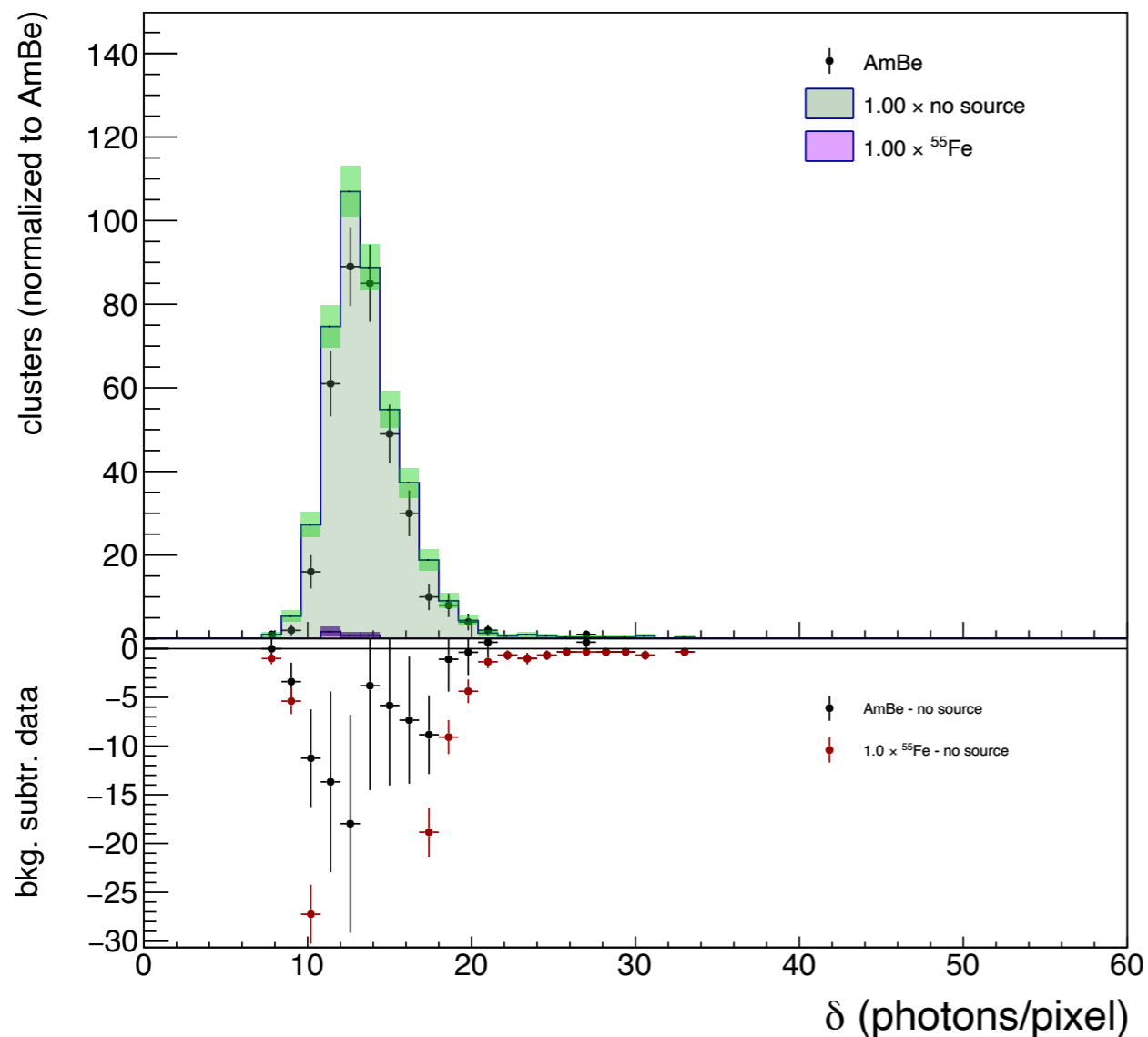
Superclusters found



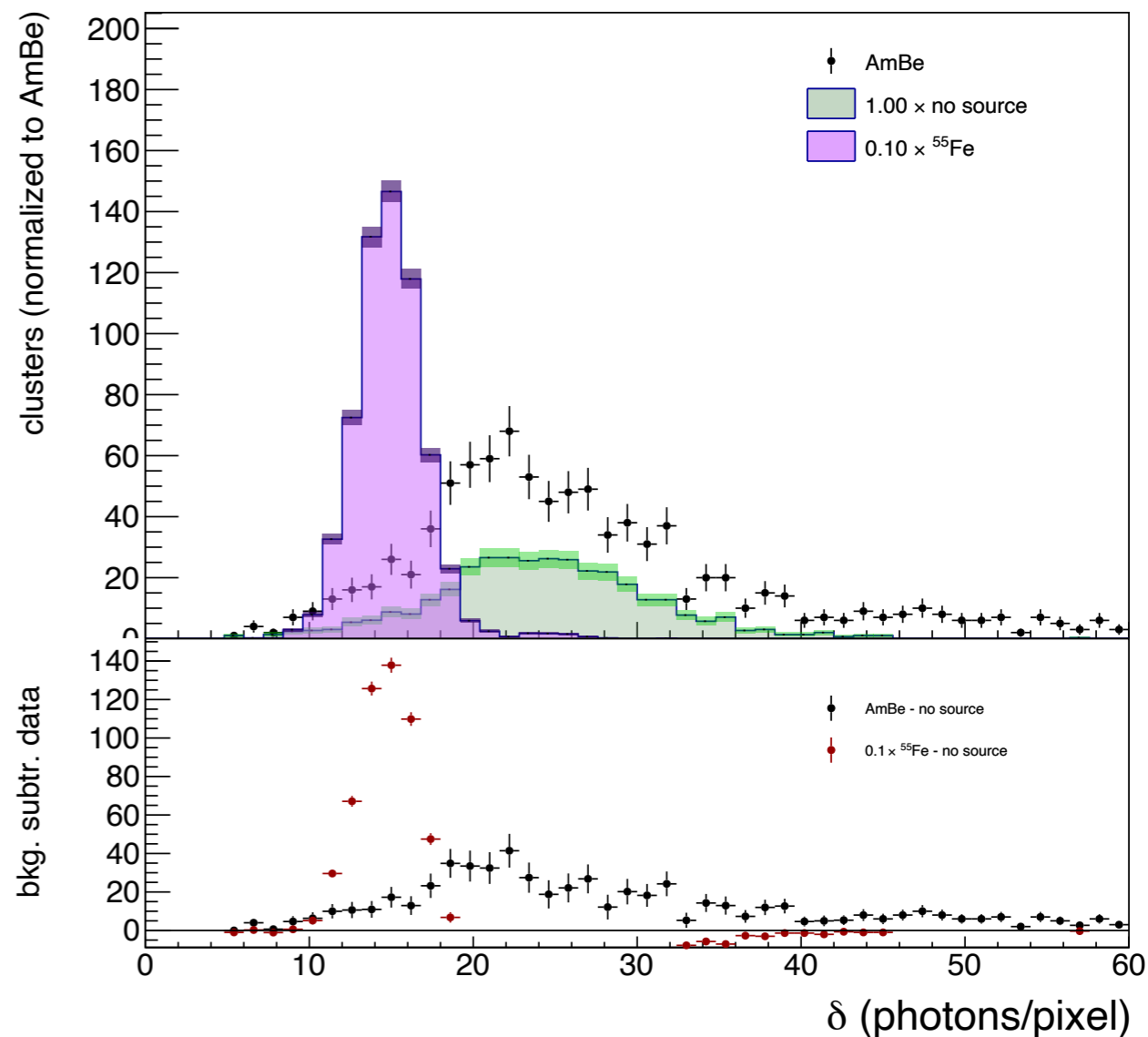
new noise filtering

applied to AmBe / ^{55}Fe

Solved partially the high density of long tracks, due to added noise, but...



long (>12 cm) slim tracks



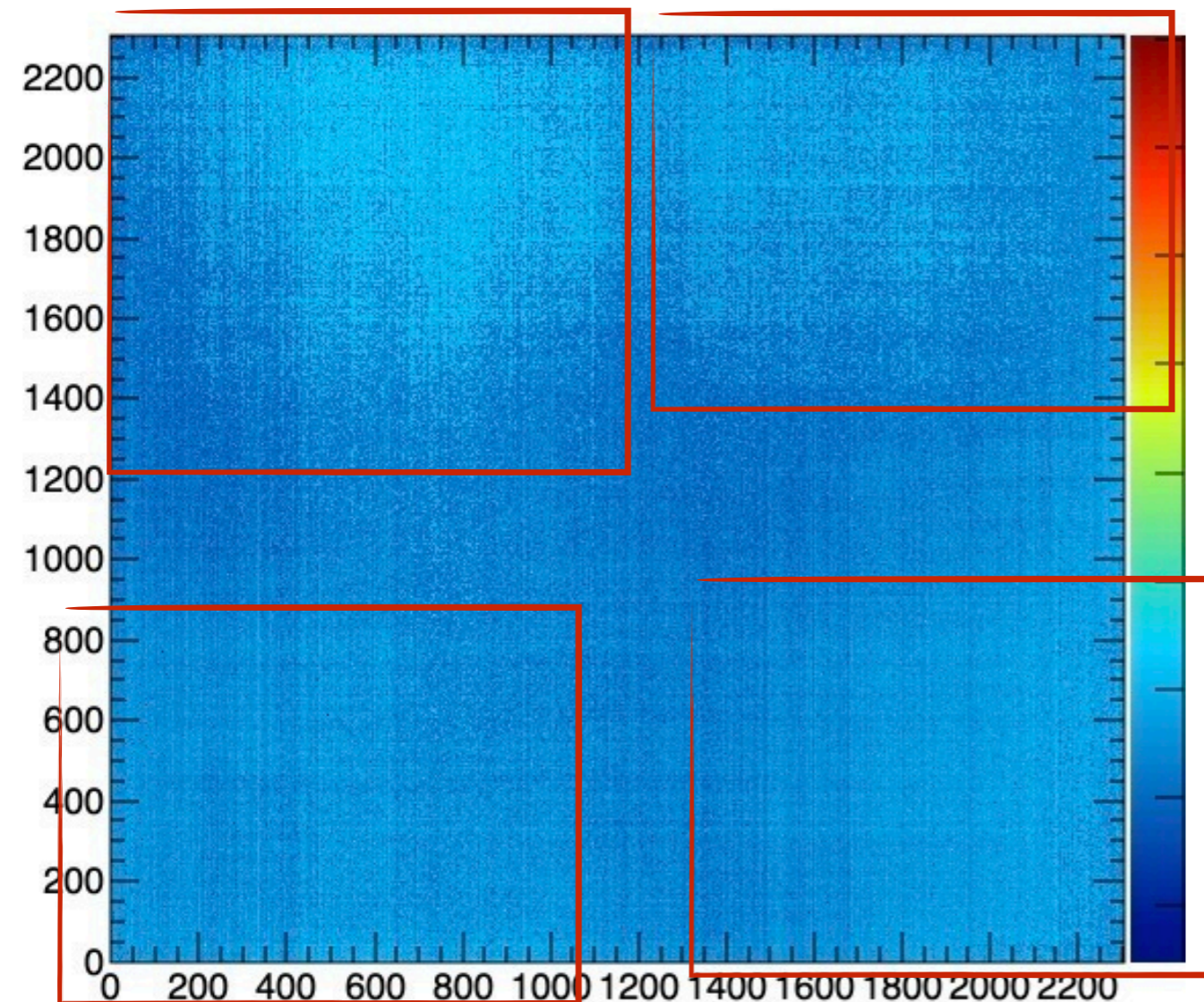
short tracks (<1cm)

δ of cosmics now < Fe. Still high...

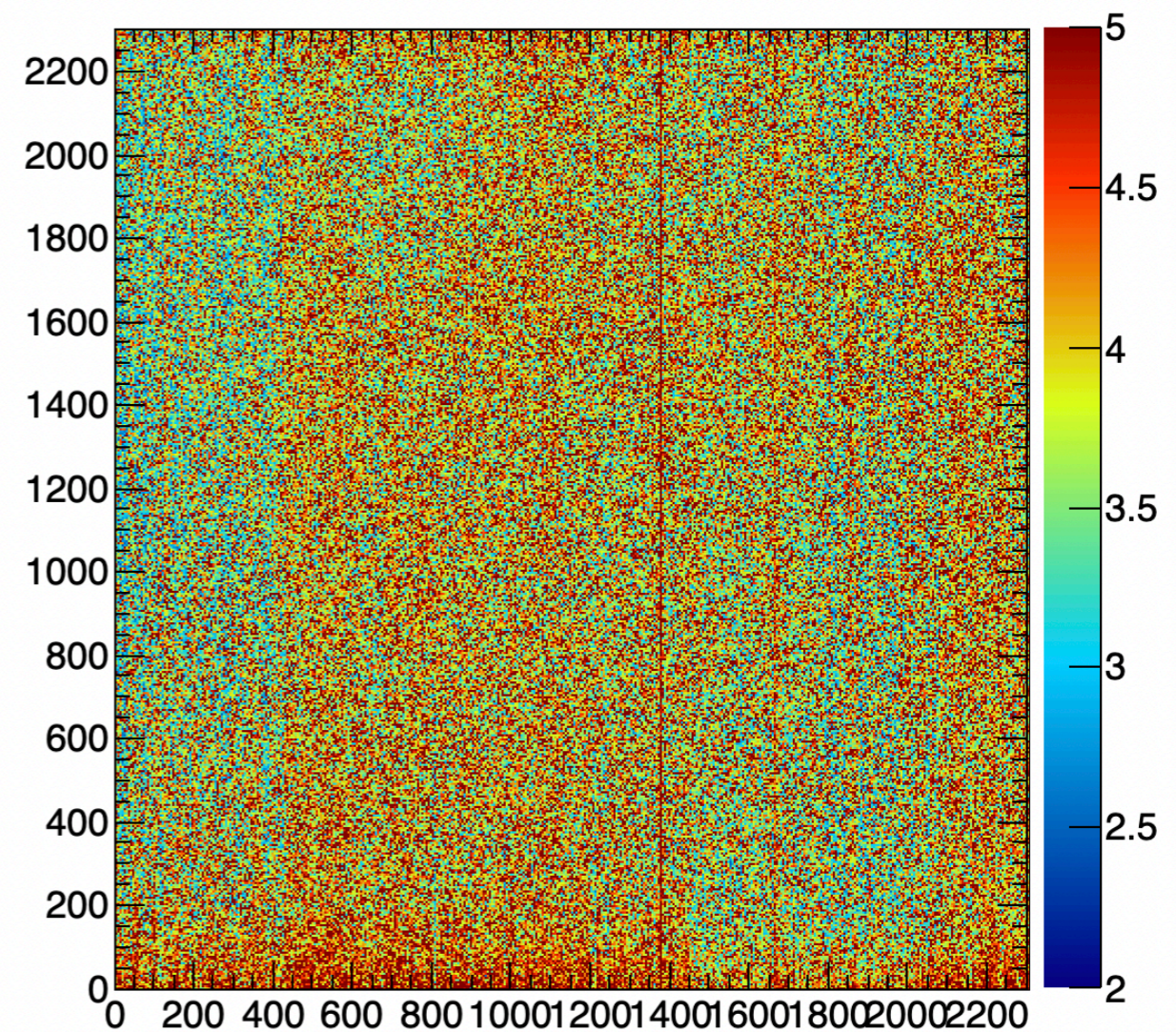
pattern sensor noise

Pedestal value, but also noise of the sensor is not random. It goes along strips and is different in 4 areas.

pedestal run with camera on LIME and HV OFF



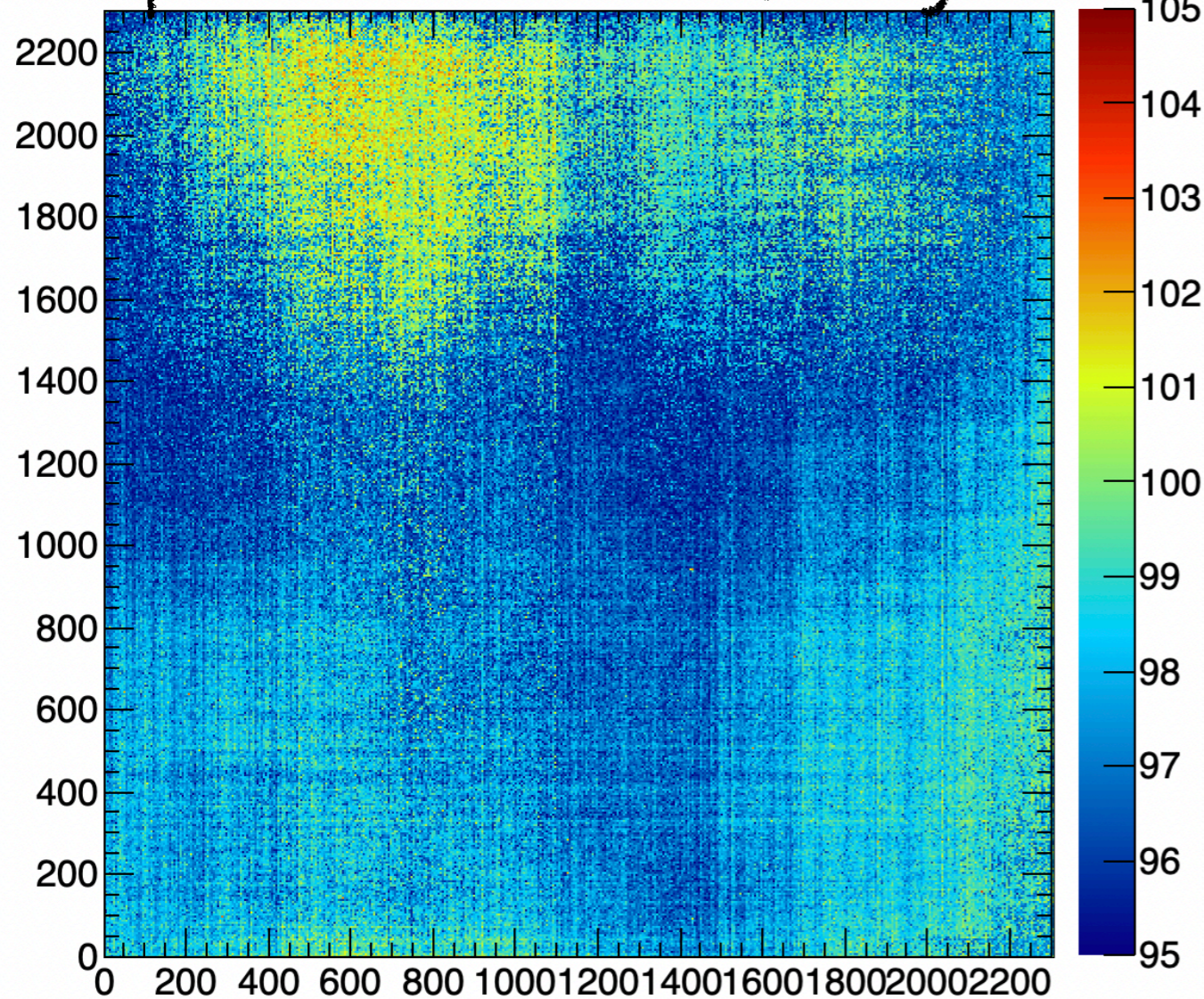
pedestal mean



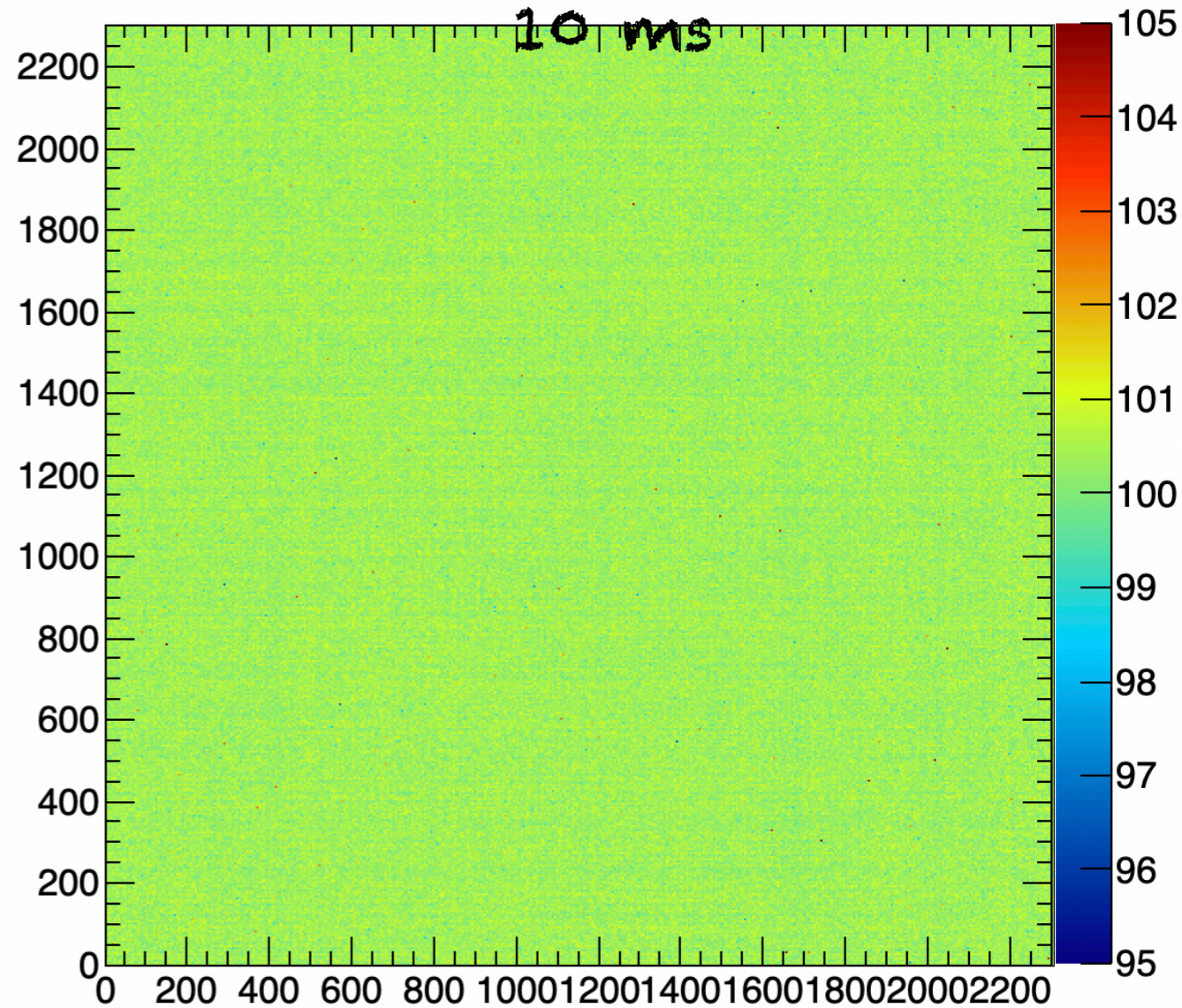
pedestal RMS

Pedestals?

pedestal on LIME, July 17



pedestal with cap, 27 Nov
10 ms



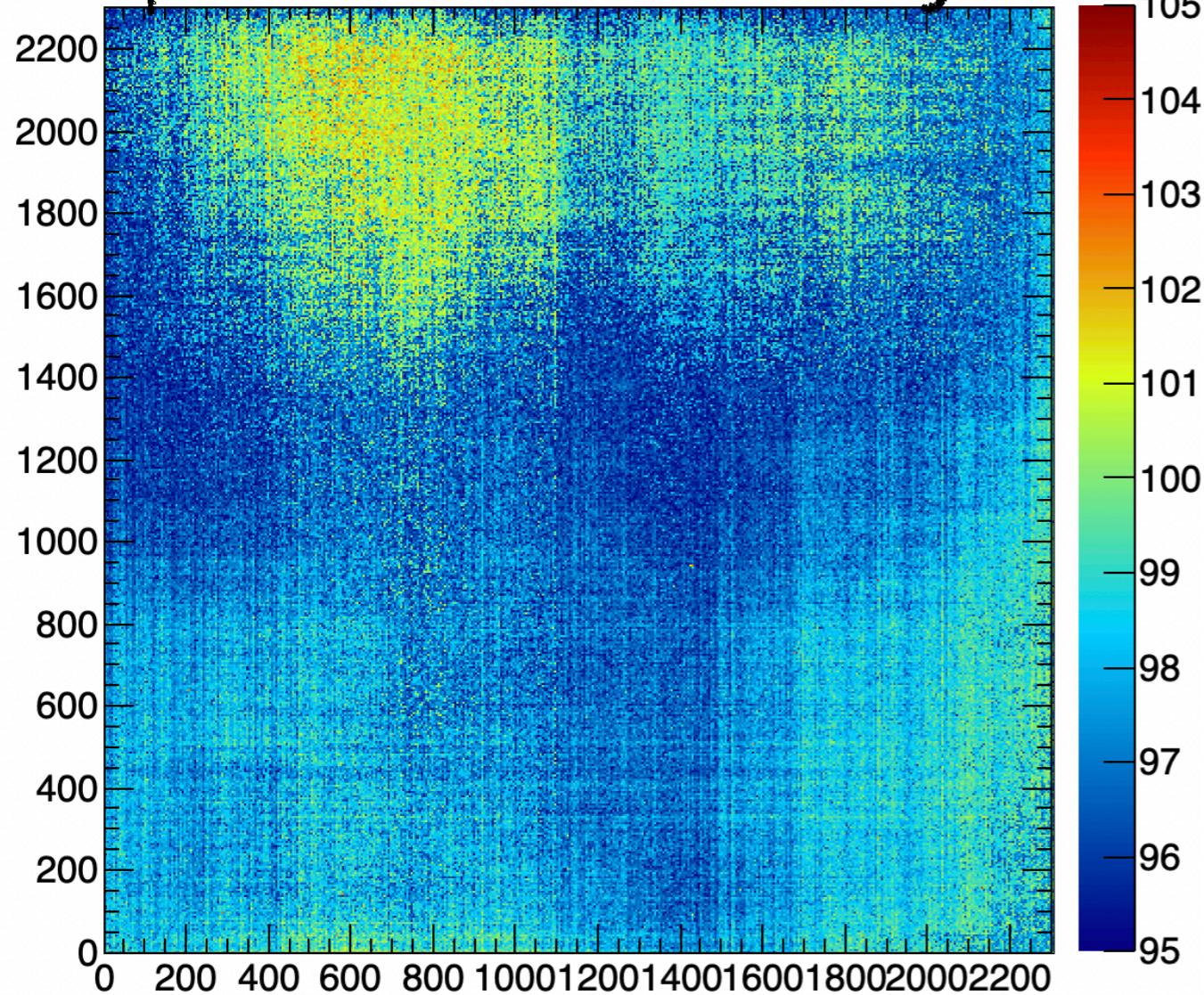
patterns at the borders
up to 10% local variations

no patterns at all

LIGHT entering LIME during AmBe and Cs runs !? !?
This adds large fake energy in different zones

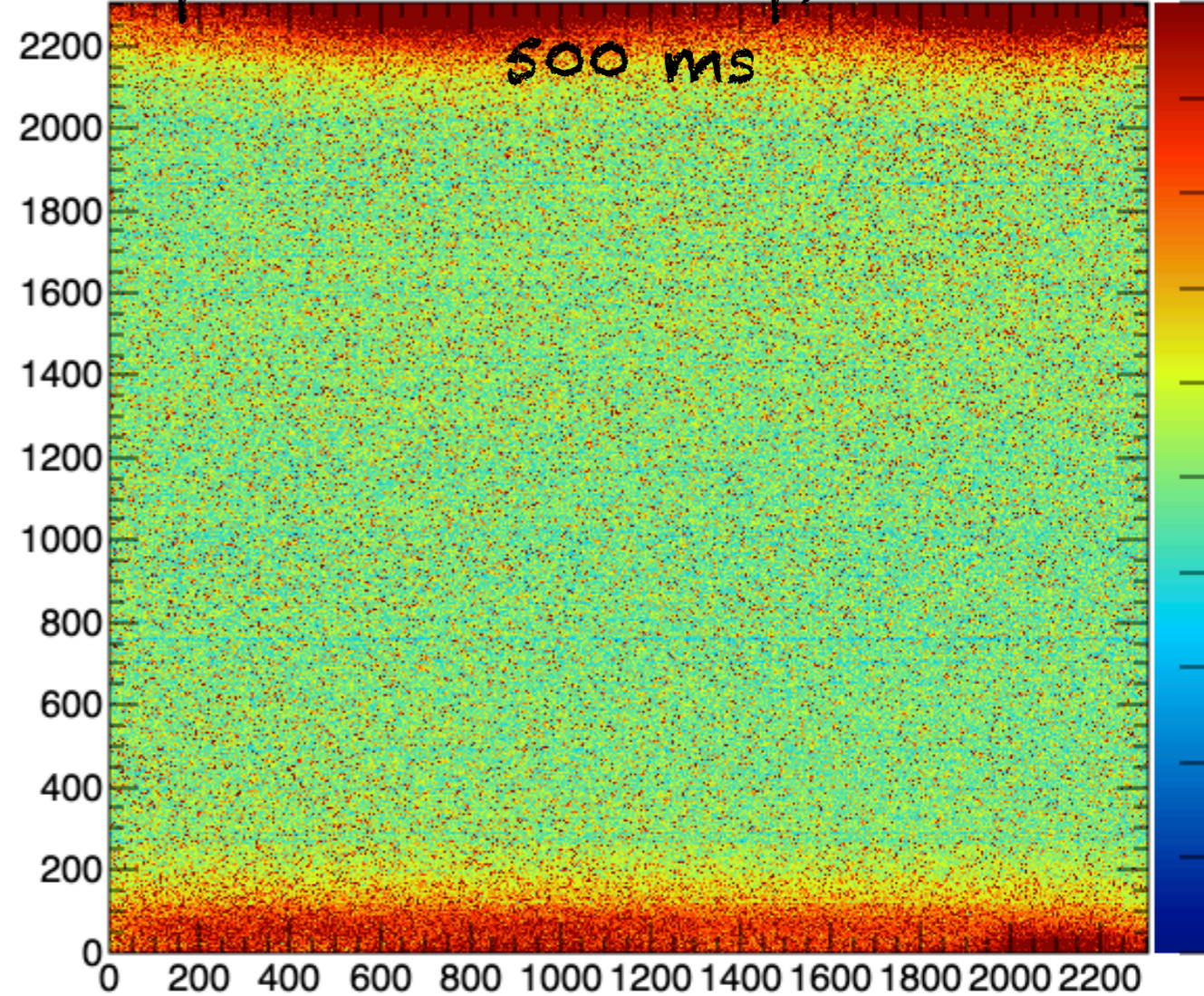
Pedestals?

pedestal on LIME, July 17



patterns at the borders
up to 10% local variations

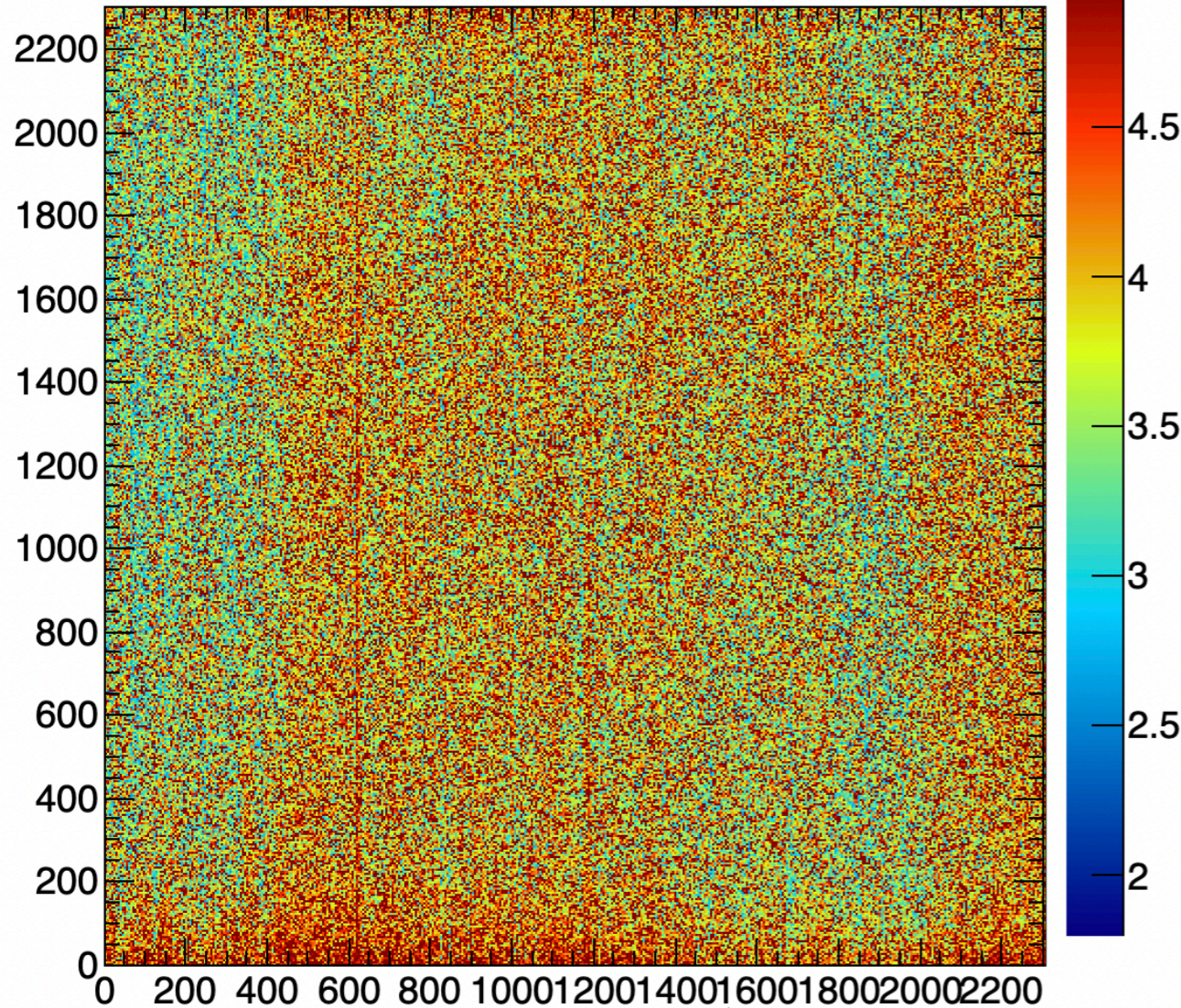
pedestal with cap, 27 Nov



no patterns at all,
apart noise arising for aperture
longer than 500 ms

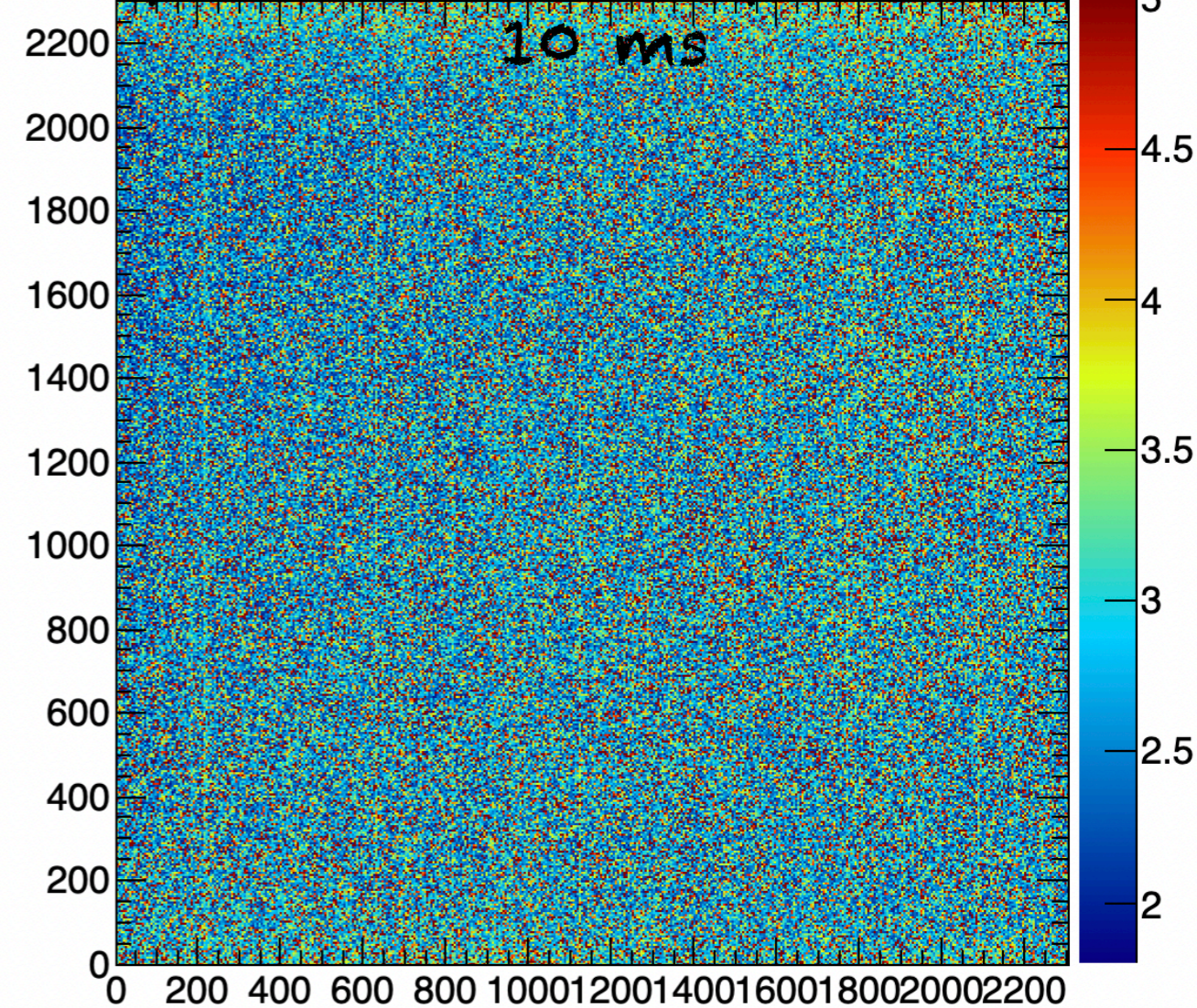
Pedestals RMS

pedestal on LIME, July 17



higher at the bottom
can be light variations?

pedestal with cap, 27 Nov

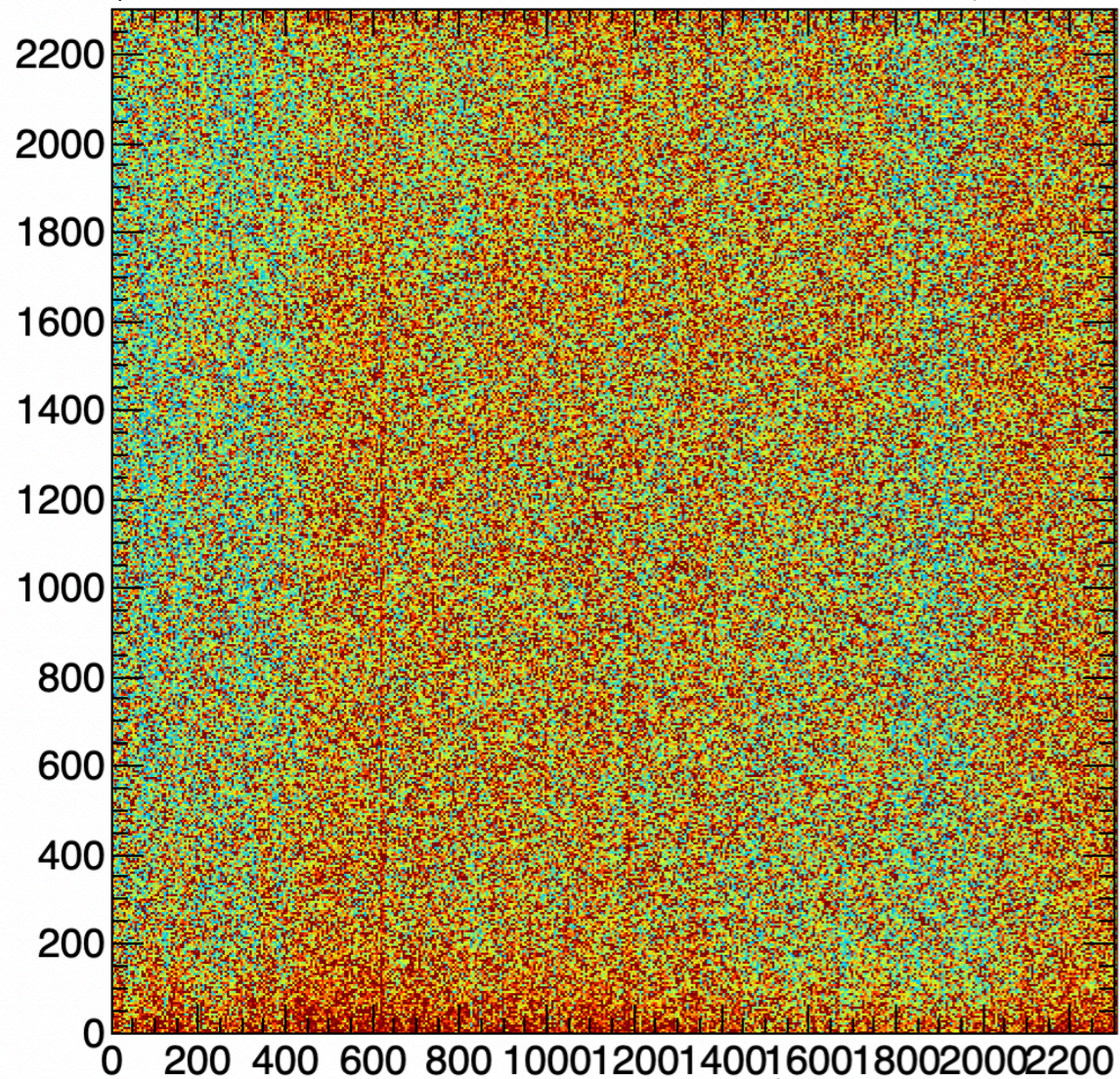


only a bit higher at the top/bottom

it means that 1.50 cut of the ZS
is in reality 20 cut

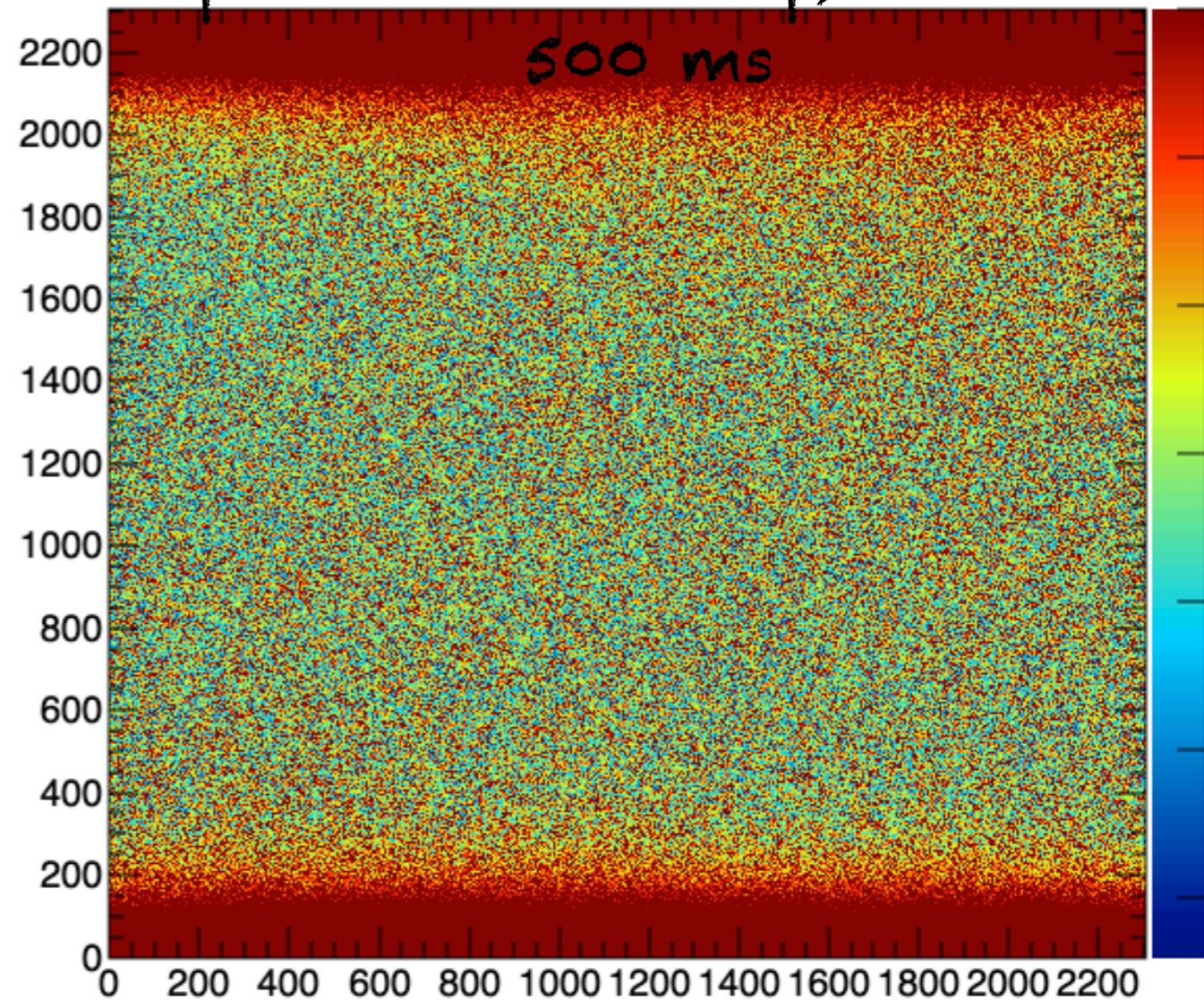
Pedestals RMS

pedestal on LIME, July 17



higher at the bottom
can be light variations?

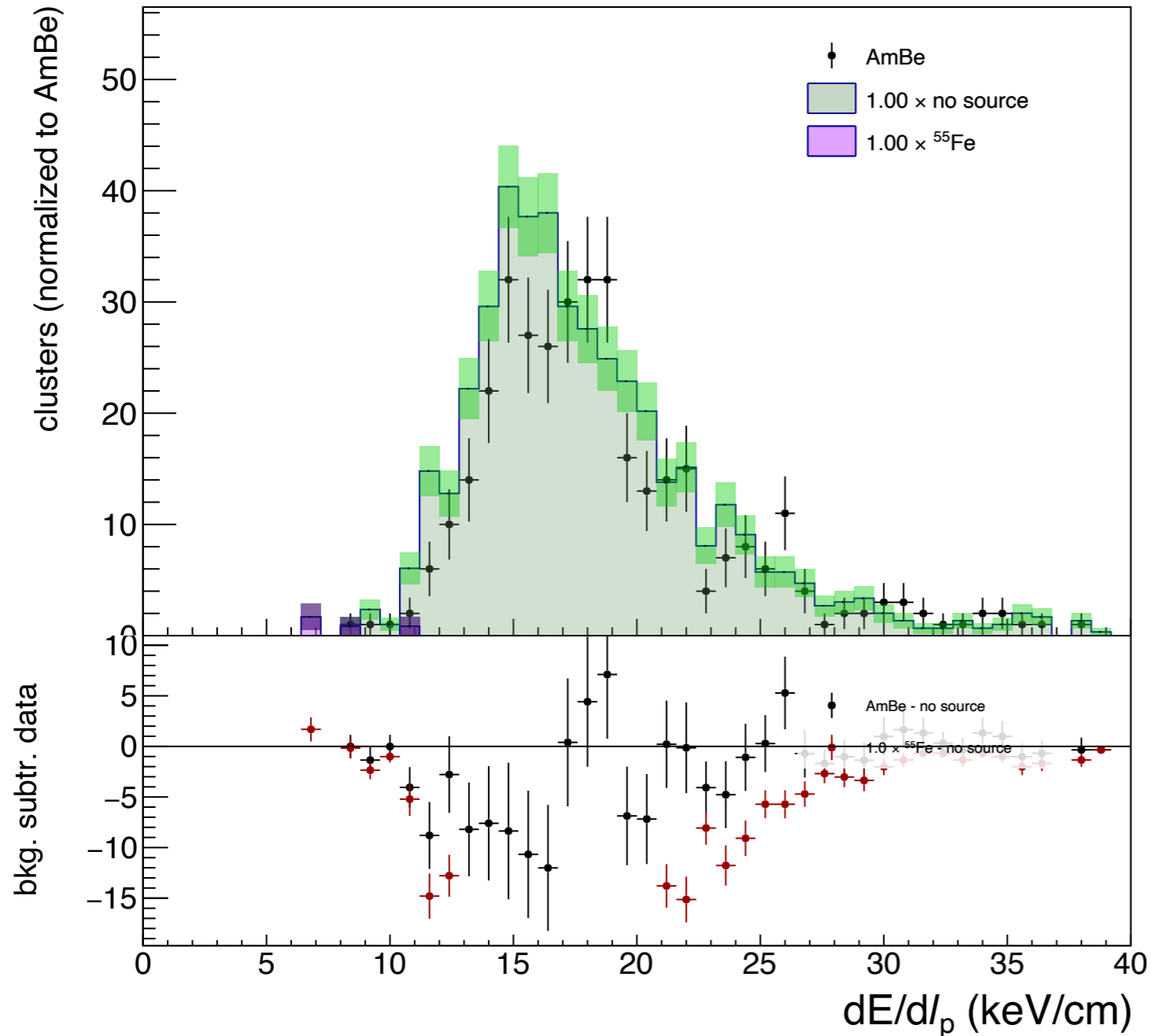
pedestal with cap, 27 Nov



500 ms
only a bit higher at the top/bottom

it means that 1.50 cut of the ZS
is in reality 20 cut

possible effect on energy

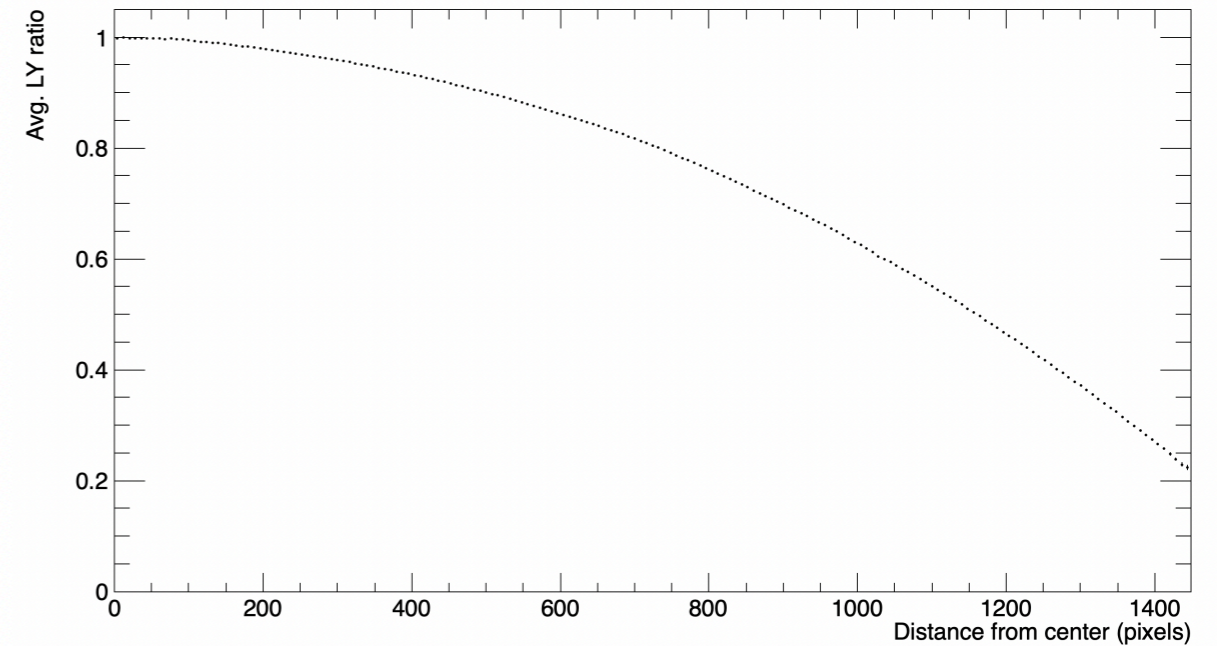
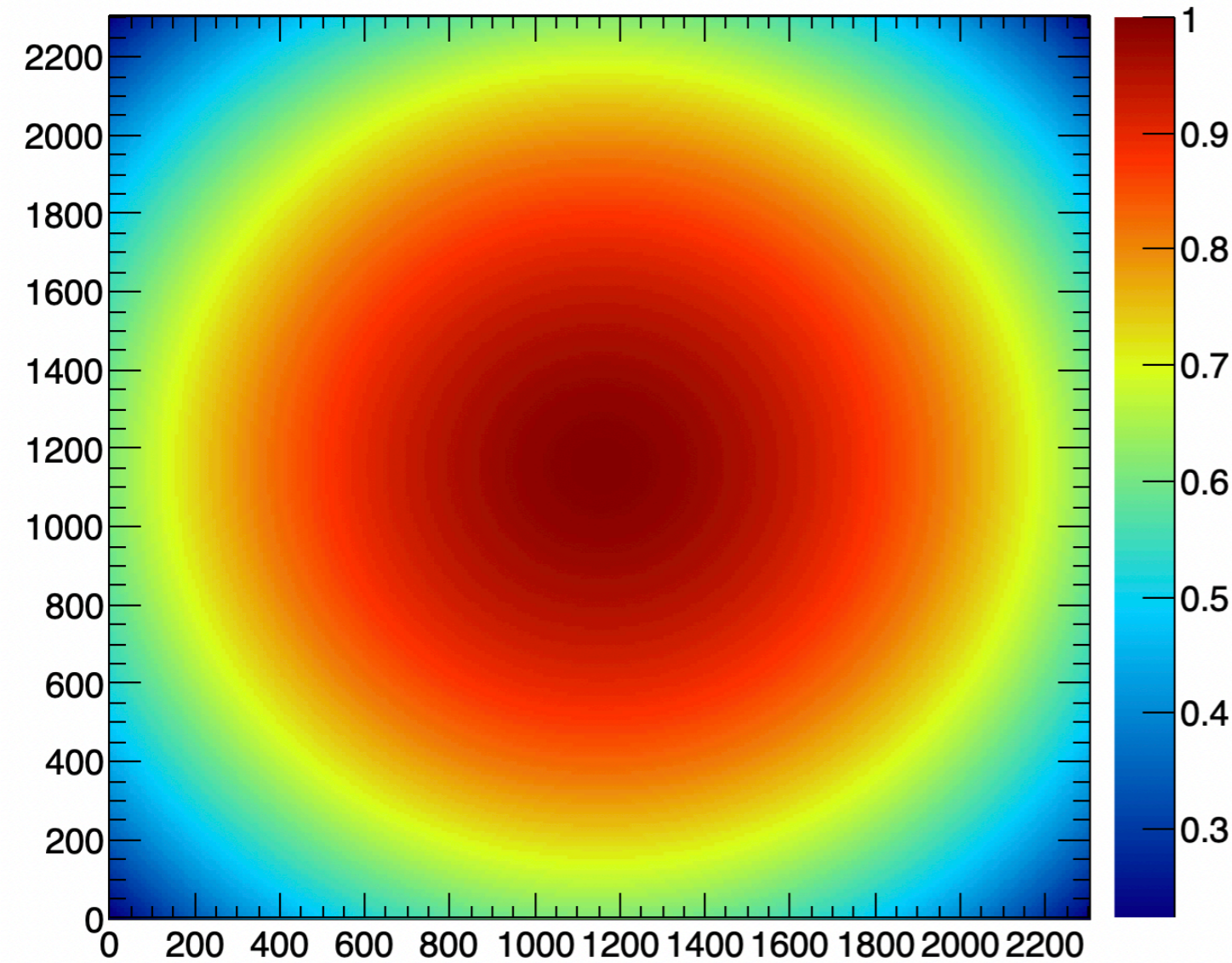


The extra light can be the reason of large dE/dx in cosmic tracks, which always traverse a poisoned region.

N.B. Even subtracting a pedestal with "light in" doesn't remove the issue, since ambient light variations during data-taking can be significant

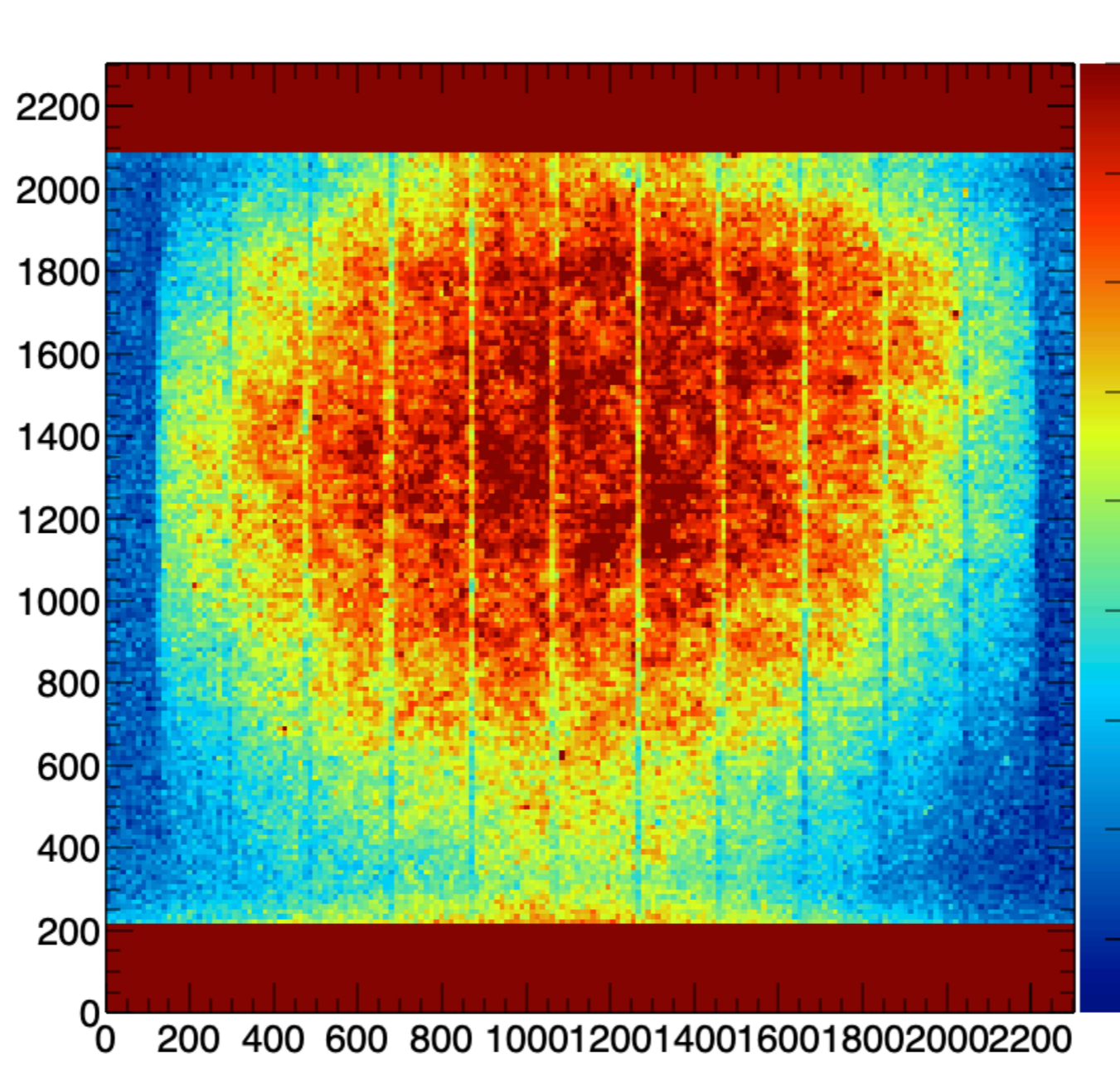
Pure optical vignetting

From pictures of white surface, averaged in 3 directions with ORCA flash



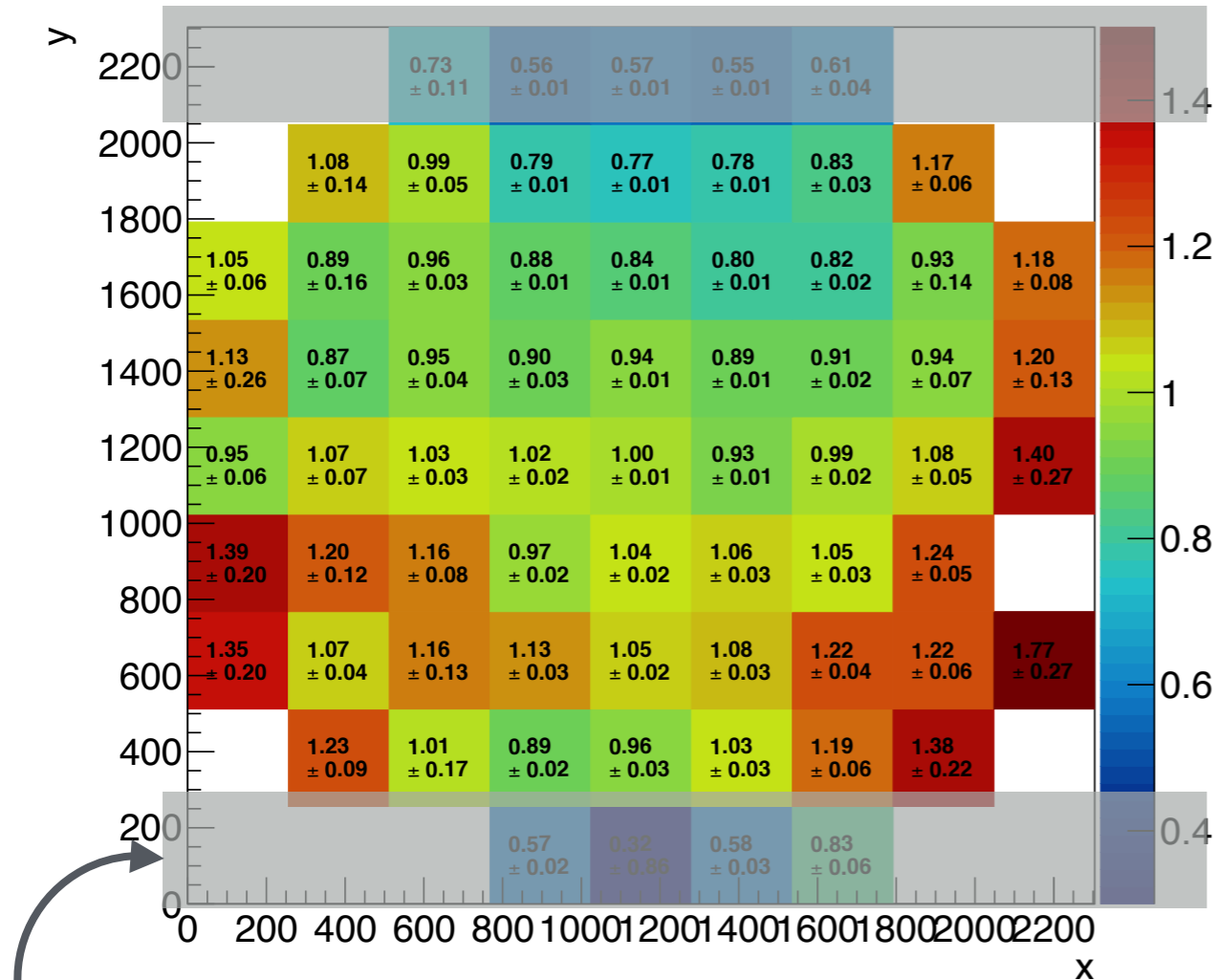
Full "vignetting"

Using average response from cosmics run 3806



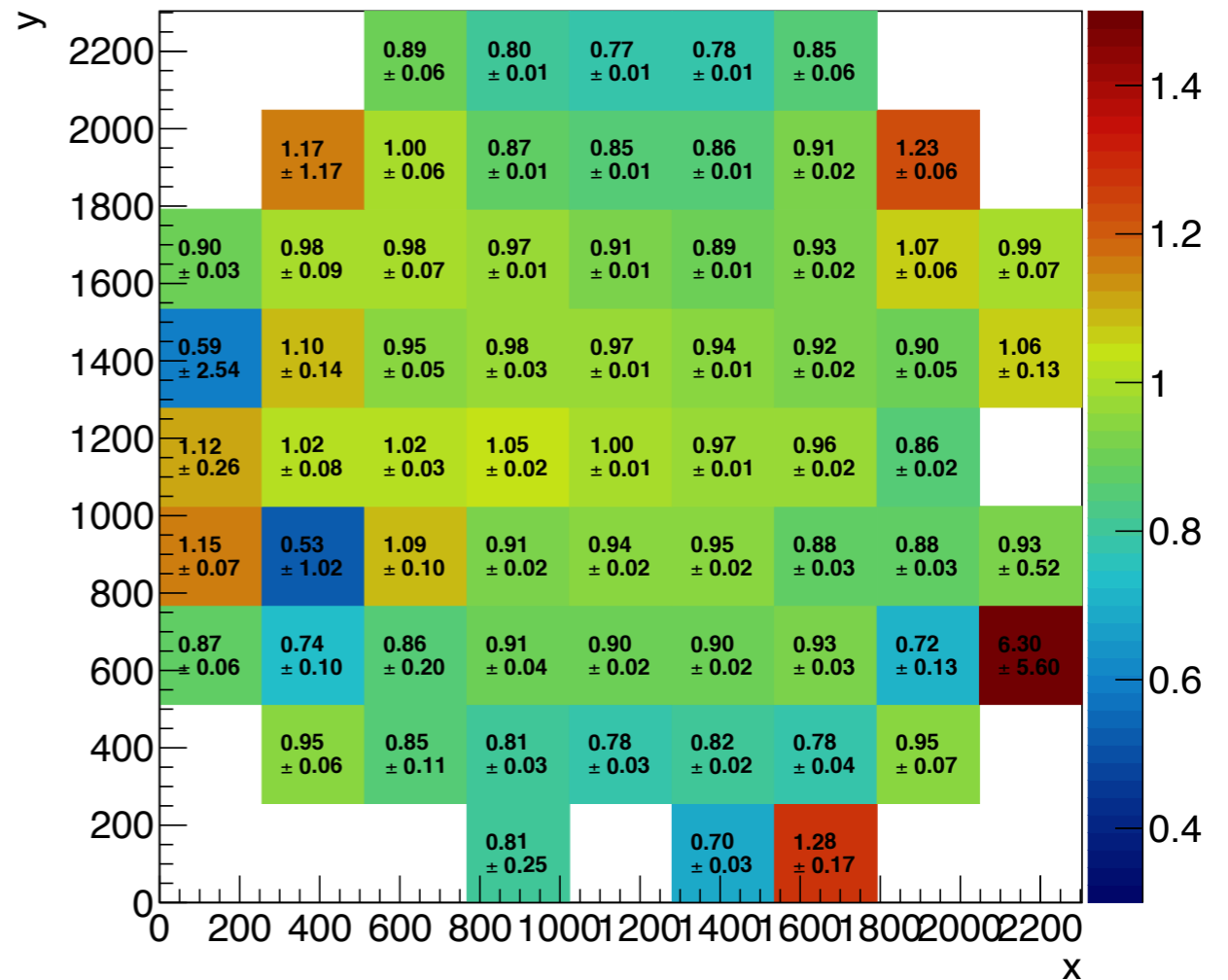
E scale with SSFe

Applied to SSFe spots. Energy scale a bit more uniform with optical vignetting correction. Still a 10-20% residual effect?



noisy regions not corrected

Correction with run 3806

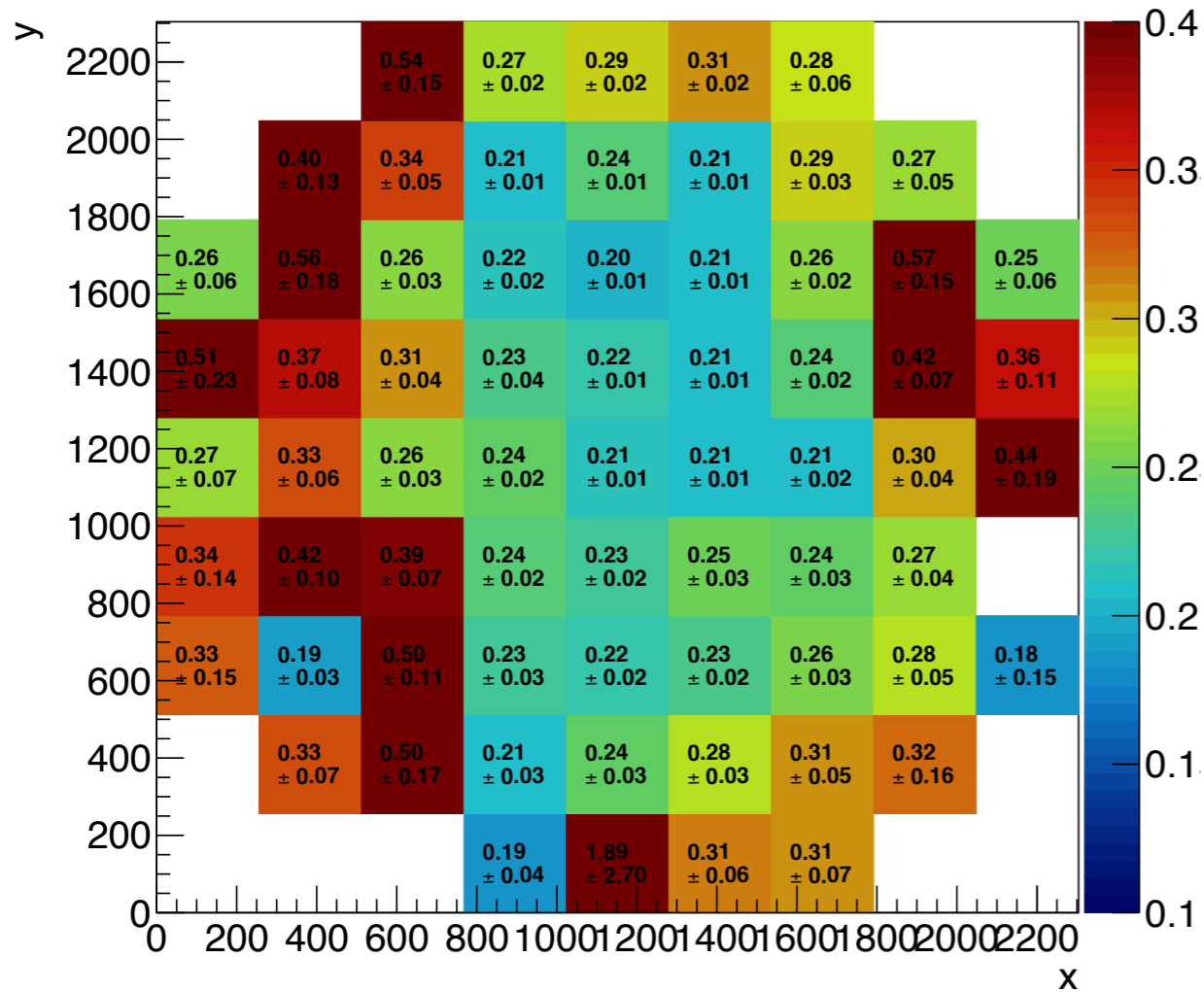


Optical-only correction

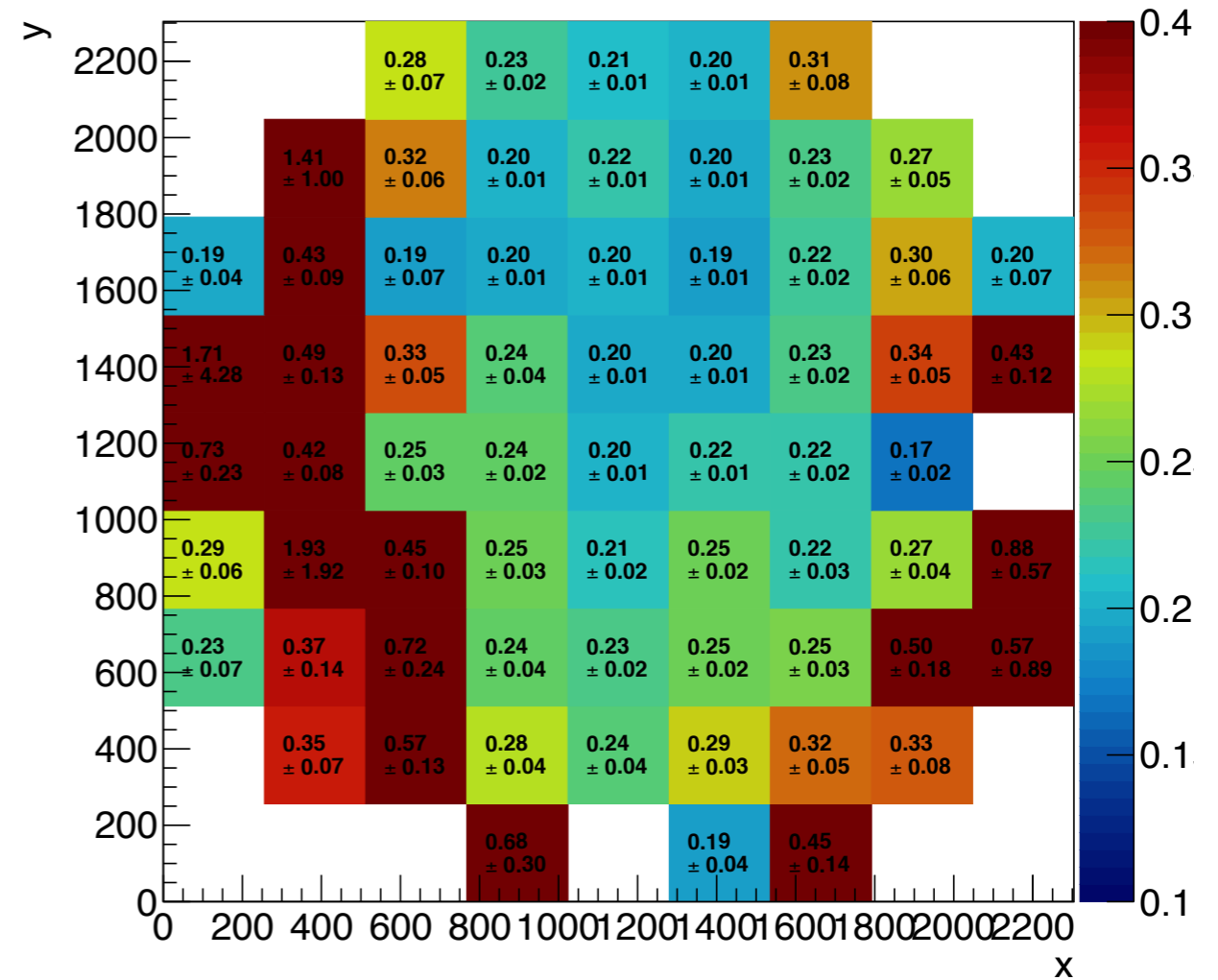
E resolution with SSFE



Similar resolution with both corrections (~20%)



Correction with run 3806



Optical-only correction