



AGENCE NATIONALE DE LA RECHERCHE

ANR



Neutrino mass ordering determination through a combined JUNO and KM3NeT/ORCA analysis

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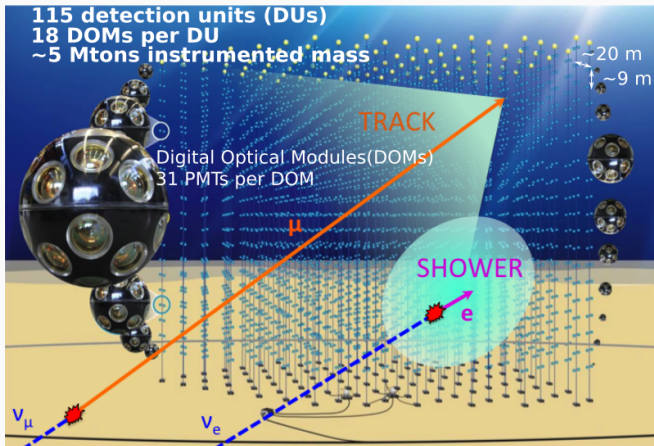
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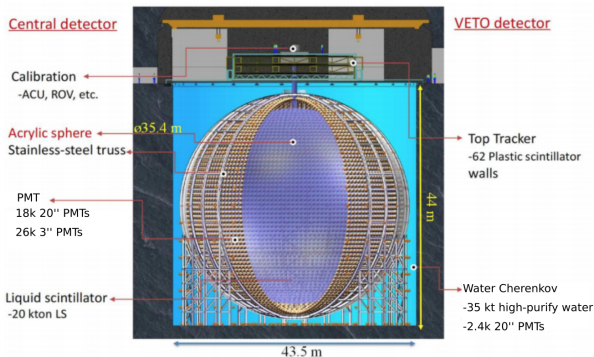
KM3NeT/ORCA

- **ORCA**: Oscillation Research with Cosmics in the Abyss.
- Study of **atmospheric neutrino oscillations above 1 GeV**.
- **Matter effect** allows the determination of the NMO.



JUNO

- **JUNO**: The Jiangmen Underground Neutrino Observatory.
- **Reactor neutrinos at medium baseline**(53 km).
- **Determine NMO by interference effects** between fast oscillations in $\bar{\nu}_e$ spectrum.
- The **Yangjiang NPP** is already fully operational, with **6 reactors** and The **Taishan NPP** has already **2 reactors** in operation.



*Model following F. An et al., J. Phys. G 43 (2016) no.3, 030401 [arXiv:1507.05613].

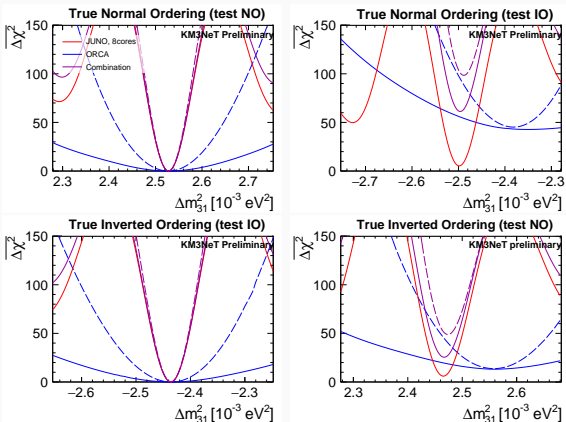
Combination strategy

- $\Delta\chi^2$ minimization of Asimov dataset.
- Combination on Δm_{31}^2 and θ_{13} using a scanned grid:

$$\chi^2(\Delta m_{31}^2, \theta_{13}) = \chi_{\text{JUNO}}^2(\Delta m_{31}^2, \theta_{13}) + \chi_{\text{ORCA}}^2(\Delta m_{31}^2, \theta_{13}) + \frac{(\sin^2 \theta_{13} - \sin^2 \theta_{13}^{\text{GF}})^2}{\sigma_{\sin^2 \theta_{13}^{\text{GF}}}}$$

| Osc parameters | JUNO | ORCA |
|-------------------|--------------|--------|
| θ_{13} | scan in grid | |
| Δm_{31}^2 | scan in grid | |
| θ_{23} | × | fitted |
| θ_{12} | fitted | fixed |
| Δm_{21}^2 | fixed | |
| δ_{CP} | × | fitted |

Synergy effect, 6 years data taking



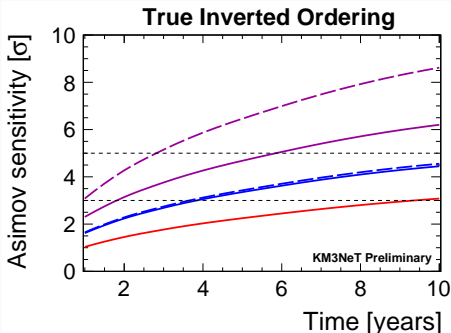
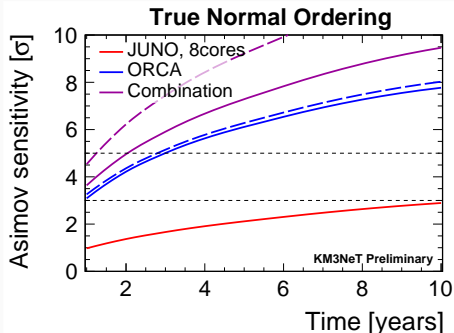
2 systematic approaches for ORCA:

- Dashed - Optimistic approach: with similar systematic set in Ref [1]
10.1103/PhysRevD.101.032006
- Solid - Conservative approach: with detector energy scale systematic leading to worse precision on Δm_{31}^2 .

| True NMO | JUNO (8 cores) | ORCA | Simple Sum | Combination |
|----------|----------------|--------------|--------------|--------------|
| NO | 2.3 σ | 6.5 σ | 6.9 σ | 7.8 σ |
| IO | 2.4 σ | 3.6 σ | 4.3 σ | 5.1 σ |

¹ Combined sensitivity to the neutrino mass ordering with JUNO, the IceCube Upgrade, and PINGU. 10.1103/PhysRevD.101.032006

Sensitivity in time

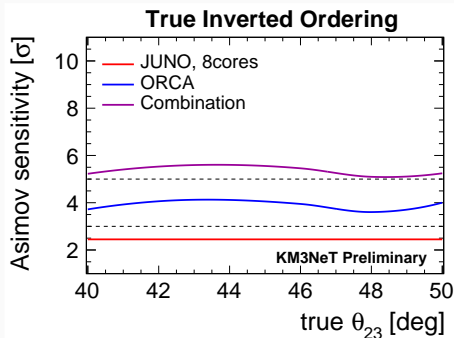
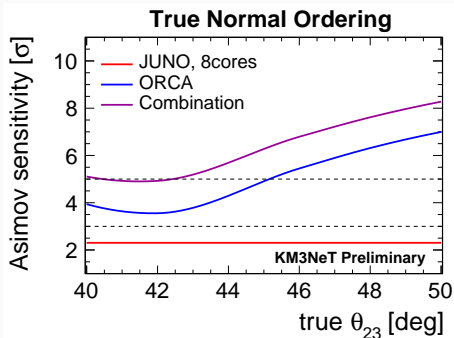


Dashed: Optimistic, Solid: Conservative

With the combined analysis

- Time needed to reach 5σ is reduced by at least one year compared to ORCA alone.
- A 5σ significance can be obtained within 2/6 years in case of true NO/IO, respectively, and independently of the systematic approach.

θ_{23} dependence, 6 years data taking



- θ_{23} dependence driven by ORCA sensitivity
- The combination ensures 5σ after 6 years regardless of the true value of θ_{23} and the true NMO.

- **The tension in the best fit values of Δm_{31}^2 boosts the NMO sensitivity in a joint fit** between JUNO and KM3NeT/ORCA.
- 5σ can be reached after **6 years of combination** of JUNO and ORCA.
- **Detector energy scale for ORCA** can have strong impact on the combination.

Thank you for your listening!

Back up

Table 1: Systematics parameters in ORCA analysis

| Parameter | ORCA standard systematic | PINGU-like systematic |
|-----------------------------------|--------------------------|-----------------------|
| PID-class norm. factors | free | × |
| Effective area scale | × | 10% prior |
| Detector energy scale | 5% prior | × |
| Flux energy scale | × | 10% prior |
| Flux $\nu_e/\bar{\nu}_e$ skew | | 7% prior |
| Flux $\nu_\mu/\bar{\nu}_\mu$ skew | | 5% prior |
| Flux $\nu_e/\bar{\nu}_\mu$ skew | | 2% prior |
| Flux E-tilt | | free |
| NC norm | | 10% prior |

Table 2: Systematic parameters in JUNO analysis

| Error source | Value |
|----------------------------|-------|
| Correlated reactor error | 2.0% |
| Uncorrelated reactor error | 0.8% |
| Reactor spectrum | 1.0% |
| Detector response | 1.0% |

Background rate and shape, together with all the associated uncertainties were taken from Ref. [2]

²J. Phys. G 43 (2016) 030401

Impact of JUNO energy resolution

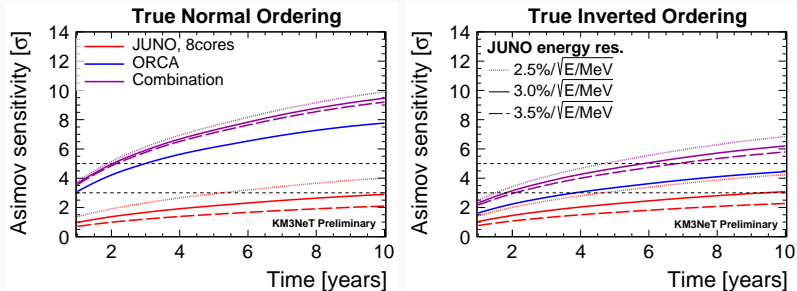


Figure 1: NMO sensitivity as a function of time, considering a better (dotted) and worse (dashed) energy resolution for JUNO than the nominal one (solid) by $\pm 0.5\%/\sqrt{E/\text{MeV}}$.

Δm_{31}^2 dependence

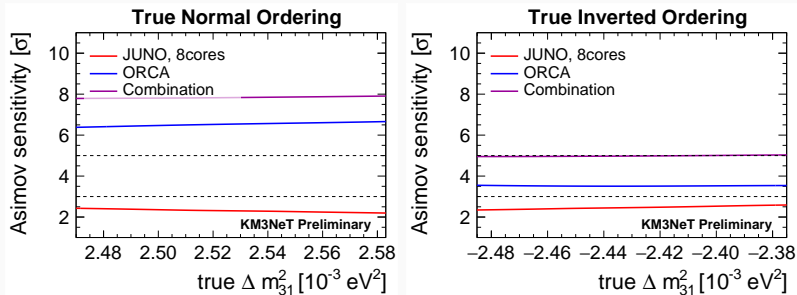


Figure 3: NMO sensitivity as a function of the true Δm_{31}^2 value for 6 years of data.