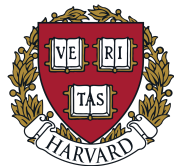

Search for Solar Atmospheric Neutrinos with 9 Years of IceCube Data

Josh Villarreal

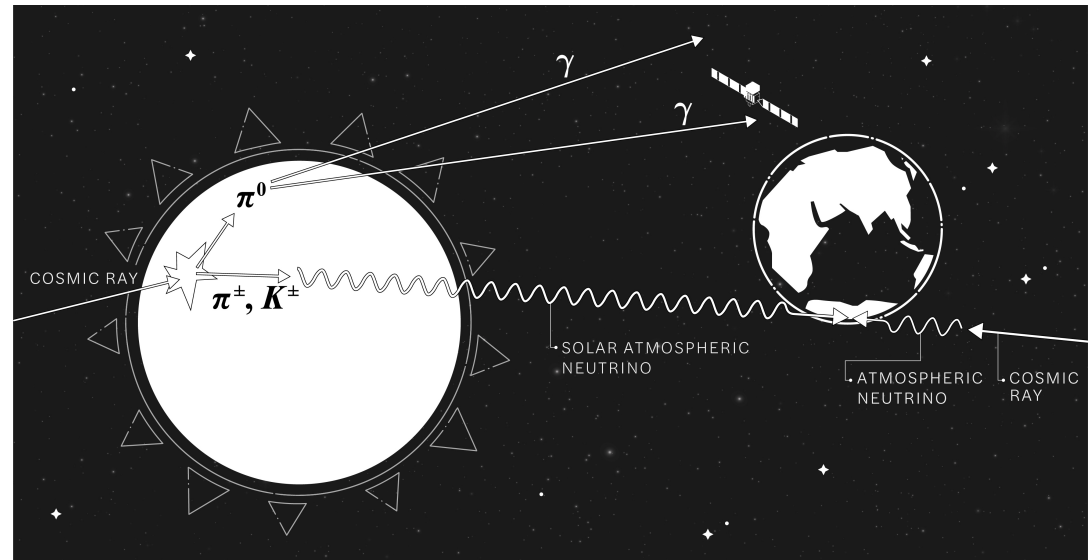
for the IceCube Collaboration

XIX International Workshop on Neutrino Telescopes, 25 Feb 2021

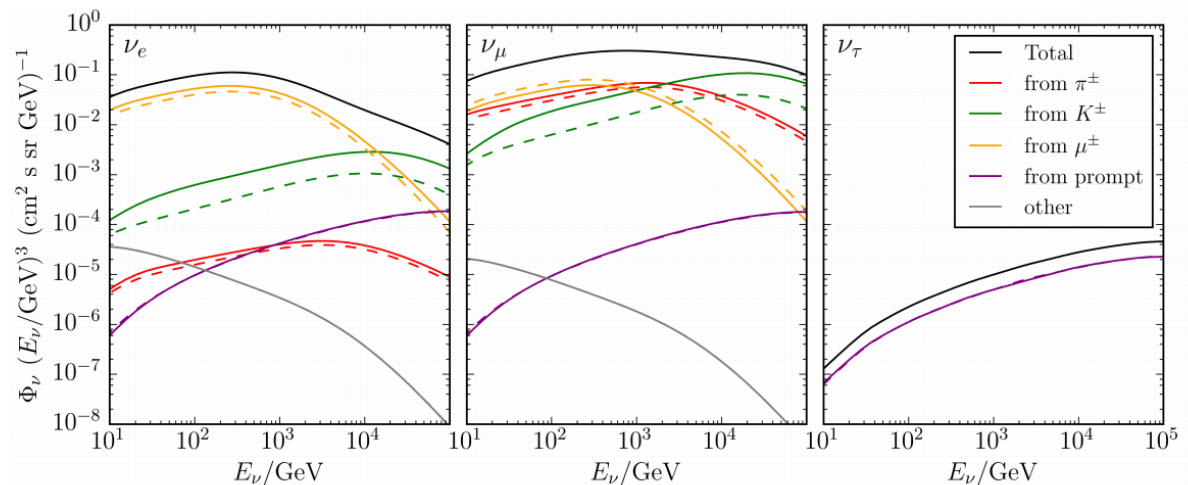


Motivation

- Solar atmospheric neutrinos are produced by cosmic-rays interacting with solar matter
- Energy spectrum falls off with $\sim E^{-3}$
- Previous IceCube analysis (<https://arxiv.org/abs/1912.13135>) set limits on normalization of nominal spectrum at $12.6 \times$ nominal value



• CARTOON OF SOLAR ATMOSPHERIC NEUTRINO PRODUCTION •

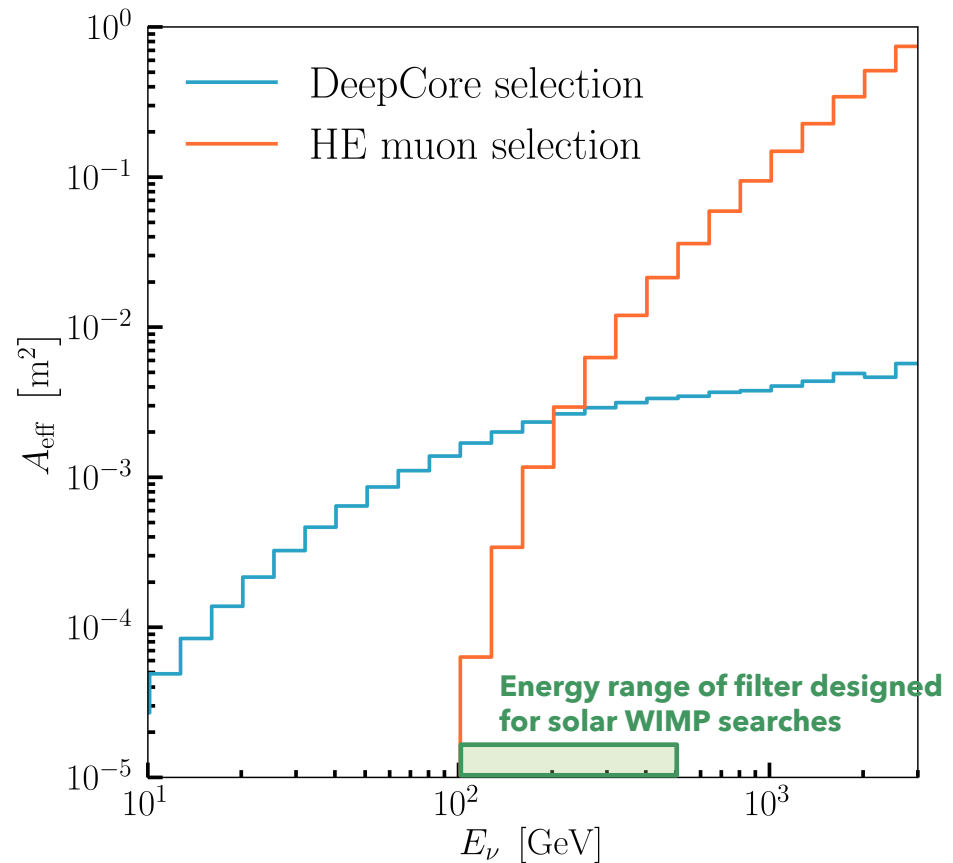


Flux of solar atmospheric neutrinos by flavor by production channel.

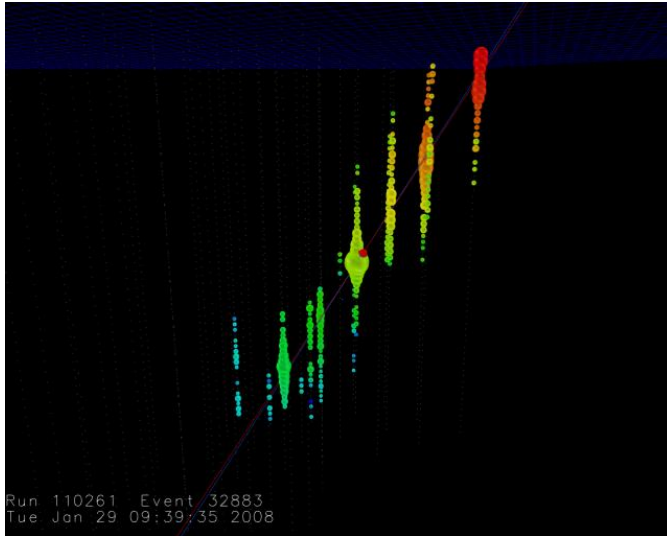
<https://arxiv.org/pdf/1703.07798.pdf>.

Optimized event selection

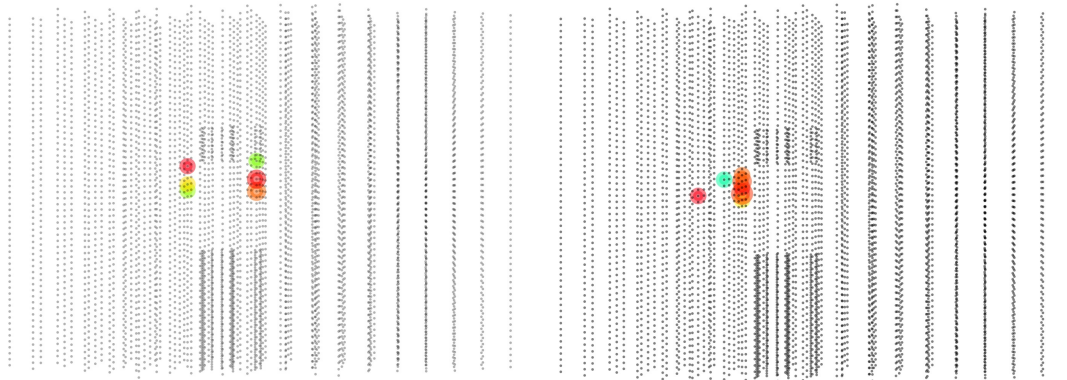
- Soft spectrum of solar atmospheric neutrinos (a background for solar WIMP searches) means event selection must account for this energy range
- **Our new event selection allows us to probe this mid-energy range, an improvement on existing event selections**



Challenges of low-energy optimization



Example of a high-energy track.
Source: [IceCube Gallery](#).



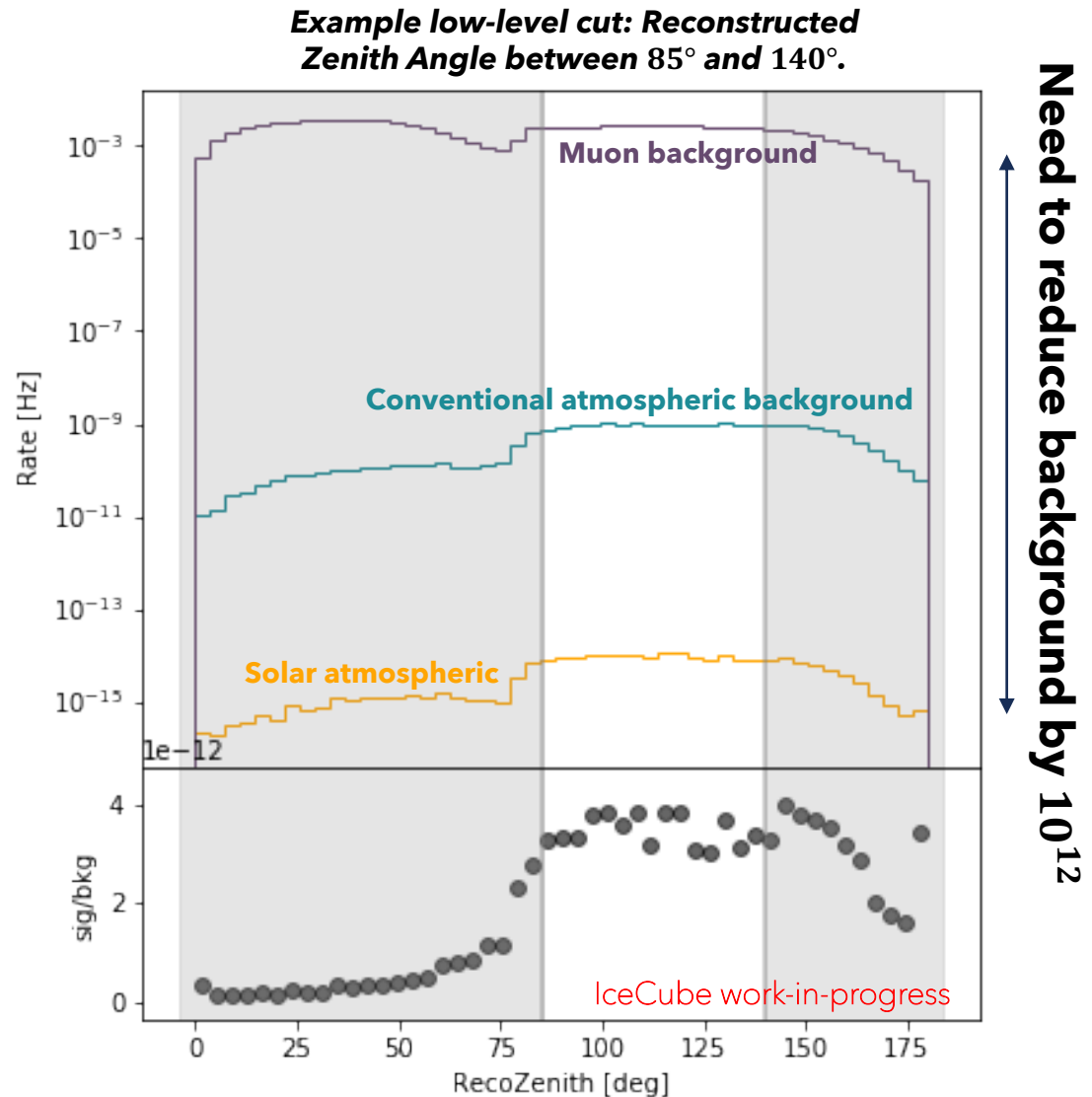
Example of a 200 GeV muon track.

Example of a 200 GeV electron cascade.

- Previous analysis only used high-energy tracks (muon neutrino dominated)
- Our MC simulations include:
 - High-energy tracks (~ 1 TeV) and are track-like
 - Low- and mid-range energy tracks/low energy cascades **which include all flavors**
 - Angular resolution is poorer anyway, so might as well include all flavors
 - Cascade background 10x lower than tracks

Status of event selection

- We've combined low- and high-energy Monte Carlo samples from previously existing simulation toolkits
- Currently verifying low-level background exclusion cuts by investigating cuts from previous analysis



Next steps and acknowledgements

- Low-level cuts:
 - Tune cuts to numbers optimized for this new event selection
 - Investigate how cuts effect signal purity
- Beyond:
 - Build a Boosted Decision Tree capable of filtering out additional background to further purify solar atmospheric signal
 - Optimize BDT through appropriate choices of boosting methods/hyperparameters

Special thank you to Jeff Lazar, Carlos Argüelles-Delgado, Harvard IceCube team, IceCube's cosmic ray working group, and Jack Pairin