



KamLAND-Zen 800

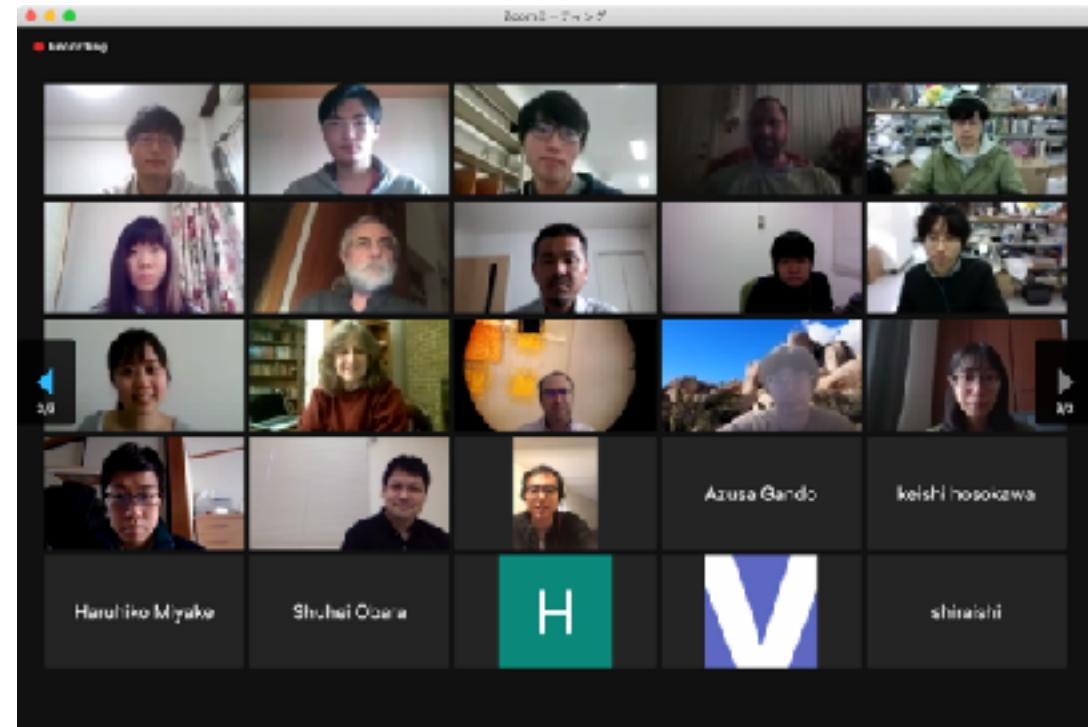
Hideyoshi Ozaki

RCNS, Tohoku University

for the KamLAND-Zen Collaboration

KamLAND-Zen Collaboration

KamLAND collaboration meeting in October 2020



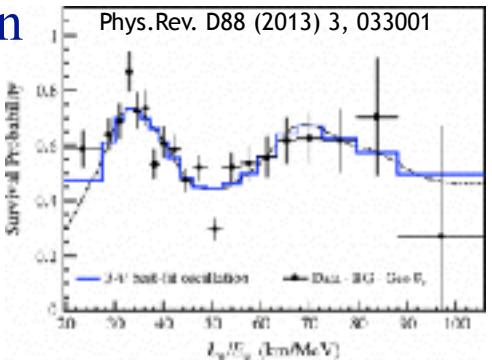
KamLAND-Zen collaboration
~60 researchers
~20 institutions

Majorana neutrinos

Neutrino oscillation

$$\underline{mv \neq 0}$$

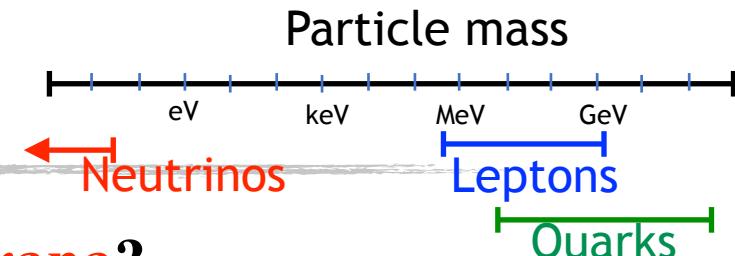
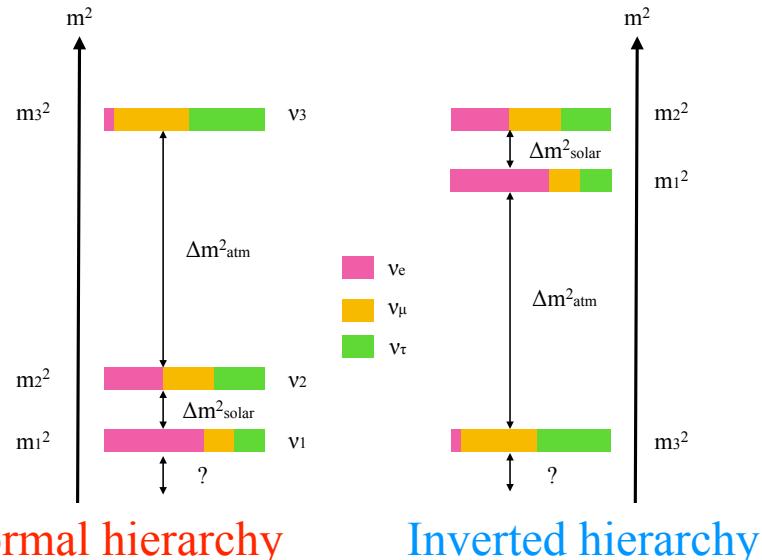
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{\text{PMNS}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



Δm^2 , θ SK, SNO, KamLAND, etc δCP T2K, NOvA

Neutrino mass hierarchy

$$\Delta m^2_{\text{atm}} \sim 2.5 \times 10^{-3} \text{ eV}^2, \Delta m^2_{\text{solar}} \sim 7.4 \times 10^{-5} \text{ eV}^2$$



Dirac or Majorana?

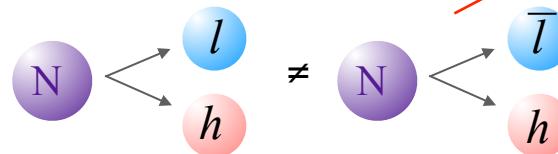
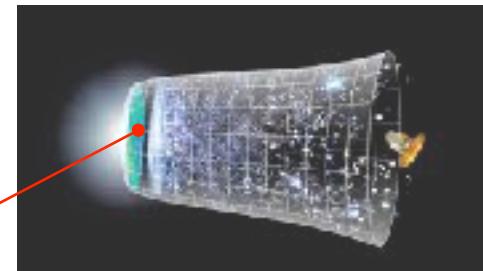
If Majorana neutrinos exist

Seesaw mechanism

Extremely small neutrino mass

Leptogenesis

Matter dominant universe



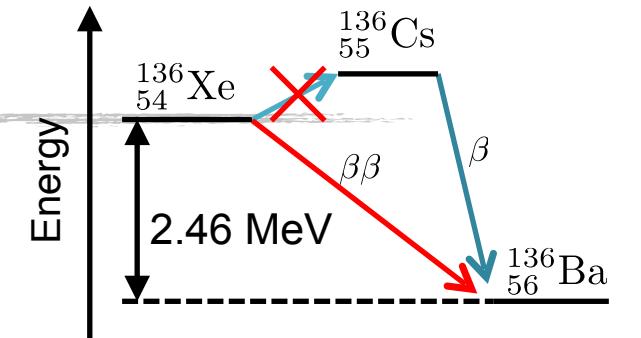
$$\Delta L \neq 0$$

How to find Majorana nature of neutrinos
-> Neutrinoless Double-Beta Decay($0\nu\beta\beta$)

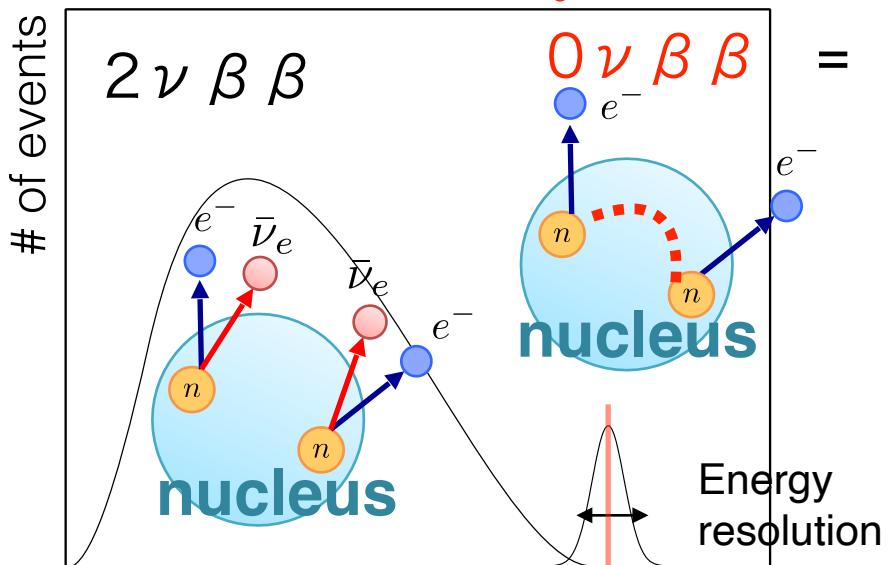
Neutrinoless Double-Beta Decay

Double-Beta Decay

Half-life $O(10^{18-24})$ years (^{48}Ca , ^{76}Ge , ^{130}Te , ^{136}Xe etc.)



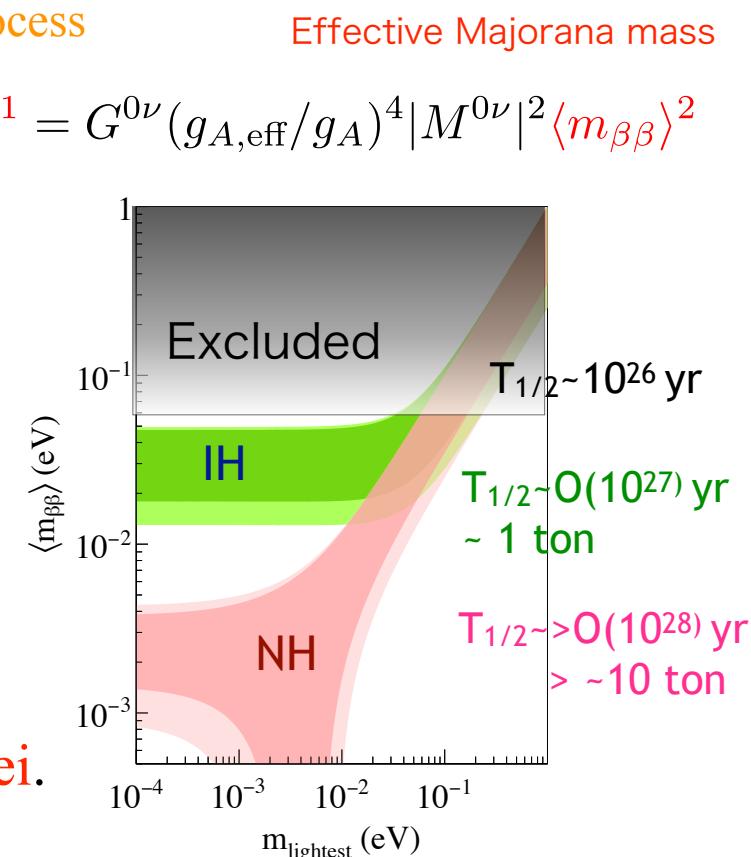
If Majorana neutrinos



$$\langle m_{\beta\beta} \rangle = \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|$$

Total kinetic energy of two electrons

What is important is to construct an **ultra-low background** detector and to monitor **a large number of $0\nu\beta\beta$ decaying nuclei**.





KamLAND

Kamioka Liquid Scintillator
Anti-Neutrino Detector

Located underground near Kamioka in Japan ~1,000 m (2,700 m w.e.)

Inner detector (ID)

- Liquid scintillator (LS) ~1 kton

Decane ~ 80%

Pseudo-cumene ~20%

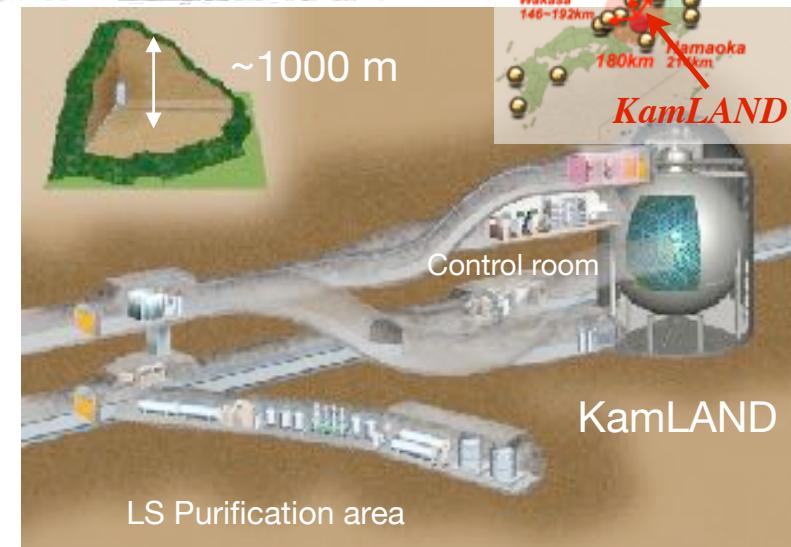
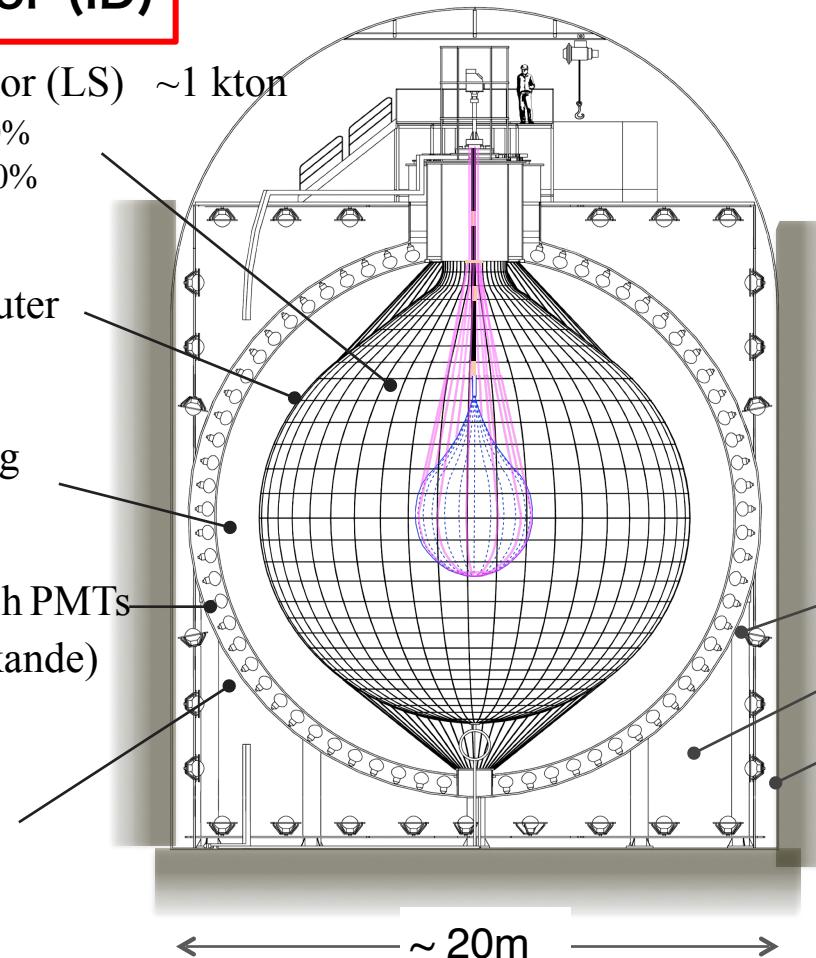
PPO 1.36 g/l

- 13 m-diameter outer balloon

- Non-scintillating buffer oil

- 1325 B&L17-inch PMTs
554 SK(Kamiokande)
20-inch PMTs

- 18 m-diameter stainless tank



Water Cherenkov Outer Detector(OD)

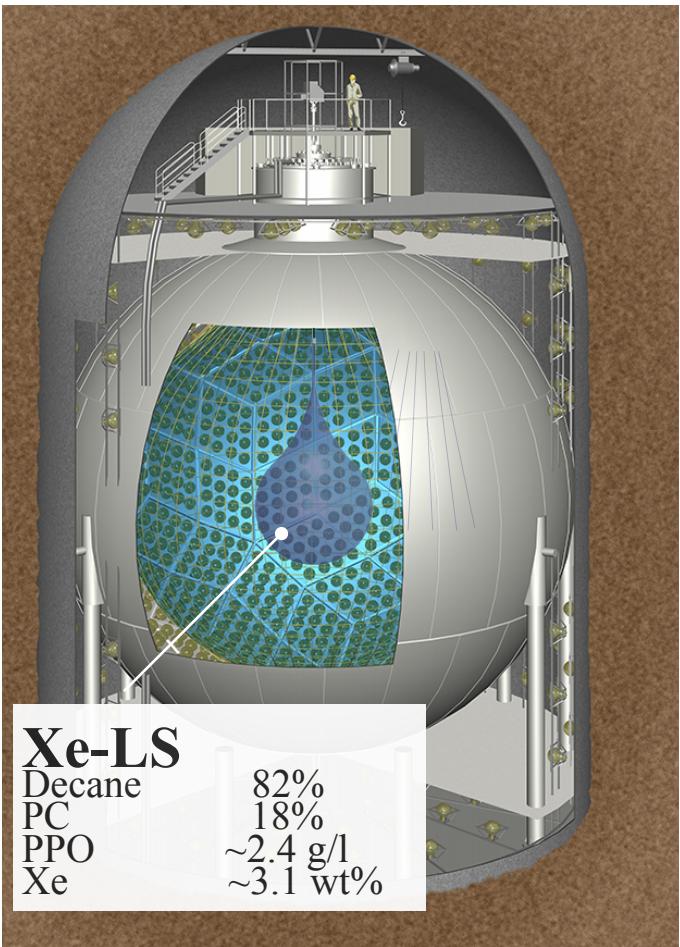
- 140 SK type 20-inch PMTs

Refurbished in 2016

- ~ 3.2 kton pure water

- Tyvek sheets for high reflectivity

Photo coverage ~34%
~ 400 p.e. / MeV
 $\Delta E \sim 7\%/\sqrt{\text{MeV}}$ @2020



Inner-balloon (IB)

Zen 400 ~3.0 m-diameter V=16.7 m³

Zen 800 ~3.8 m-diameter V=30.5 m³

×2 larger

KamLAND + Xe

- **KamLAND Ultra-low BG detector**

- Low radioactivity $^{238}\text{U} \sim 5.0 \times 10^{-18}$ g/g, $^{232}\text{Th} \sim 1.3 \times 10^{-17}$ g/g
- Cosmic-ray muon ~ 0.2 Hz
- Active shielding

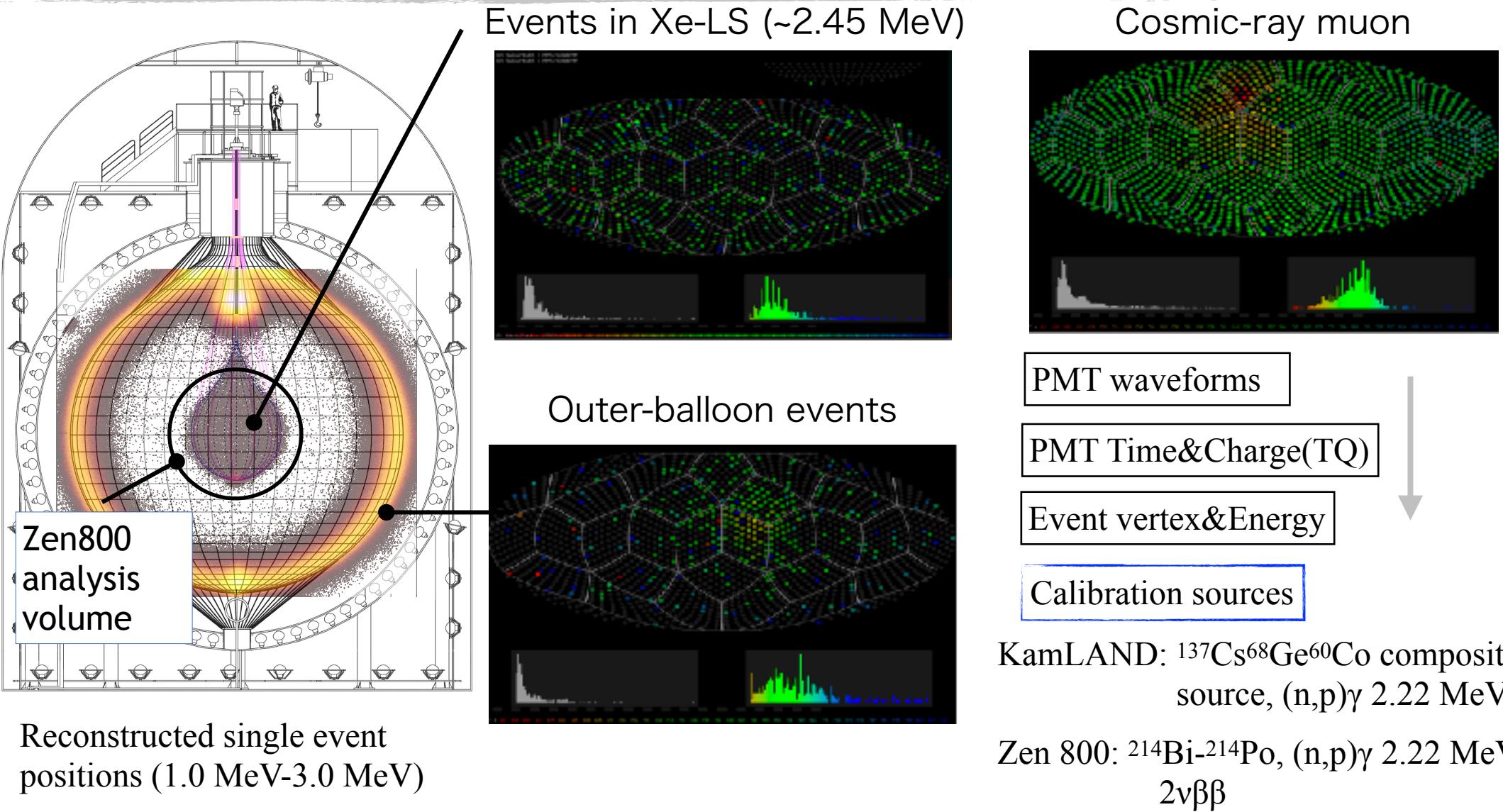
- **Xenon**

- Chemically stable noble gas
- Solubility: Xe-LS ~ 3 wt%
- Large mass&enrichment 745 ± 3 kg of ^{136}Xe ~ 91 wt% enriched Xe
- High Q-value: 2.458 MeV (Low BG in KamLAND)
- Slow ^{136}Xe $2\nu\beta\beta$ decay \rightarrow less $2\nu\beta\beta$ tail due to energy resolution

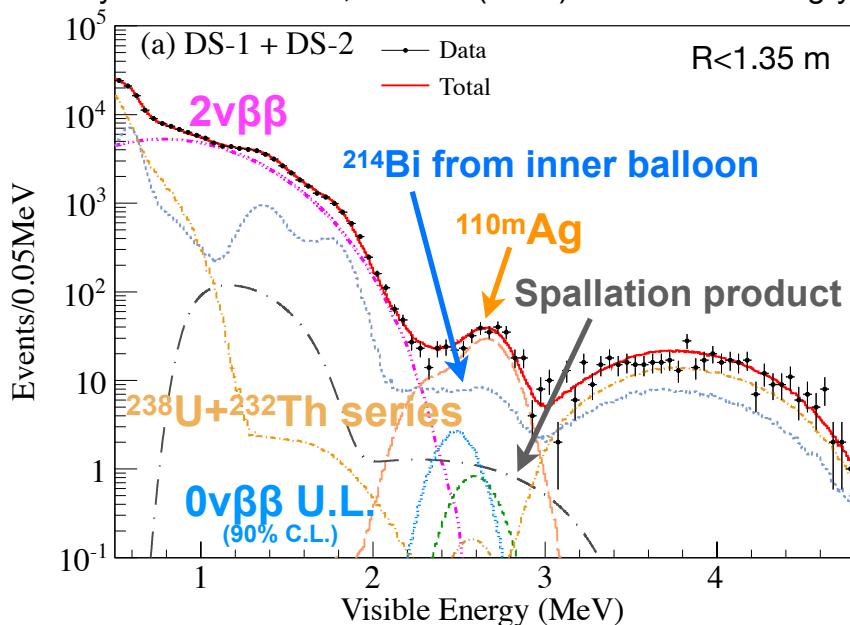
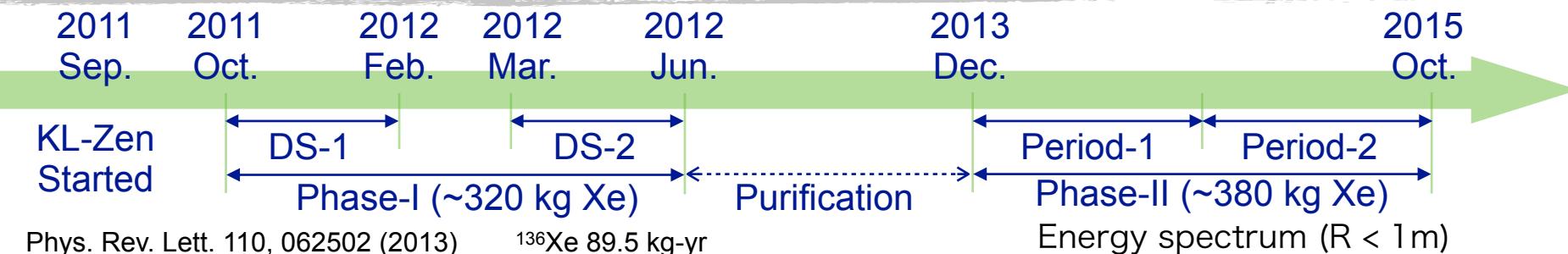
- **KamLAND-Zen**

- Scalability
- ^{136}Xe On-off

Event reconstruction



History of KamLAND-Zen 400

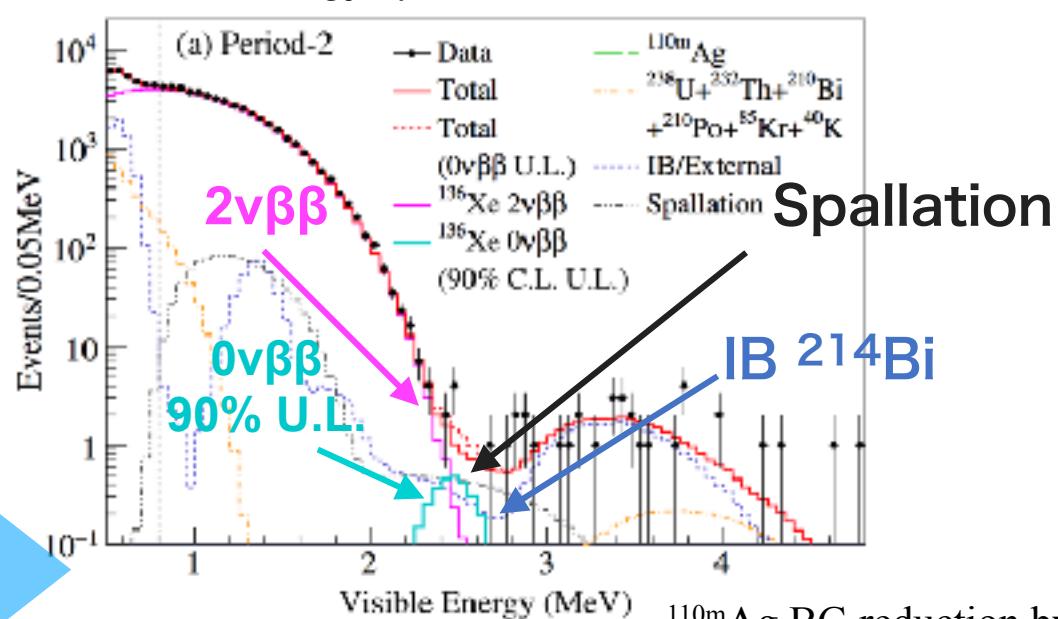
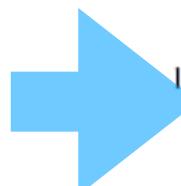


Lower limit for $0\nu\beta\beta$ in phase I

$$T^{1/2} > 1.9 \times 10^{25} \text{ yr (90% C.L.)}$$

Unexpected ^{110m}Ag BG

LS and Xe gas purification



$0\nu\beta\beta$ result (only Phase II)

$$T^{1/2} > 9.2 \times 10^{25} \text{ yr}$$

$0\nu\beta\beta$ result (Phase I+II)

$$T^{1/2} > 1.07 \times 10^{26} \text{ yr}$$

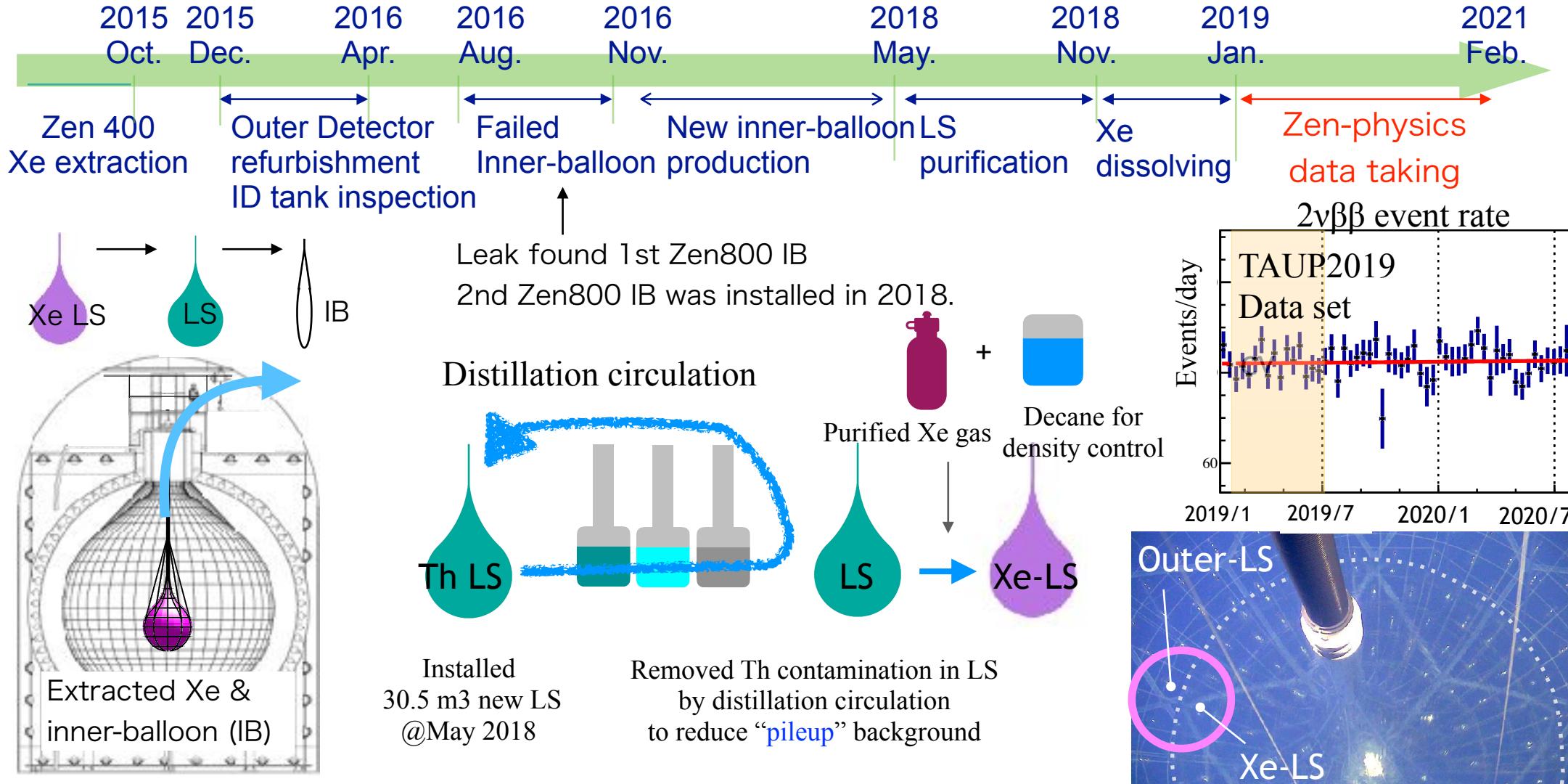
Phys- RevLett.117.082503 (August 2016) pp.1-6.

^{110m}Ag BG reduction by LS and Xe gas purifications.

Very stringent limit

KamLAND-Zen 800 construction

Target:
Doubling Xe amount
Reducing ^{214}Bi from IB



Ultra-low BG Inner-balloon(IB)

Fabricated in class-1
cleanroom in Sendai

25 μm -thick nylon film



3.8 m diameter inner-balloon

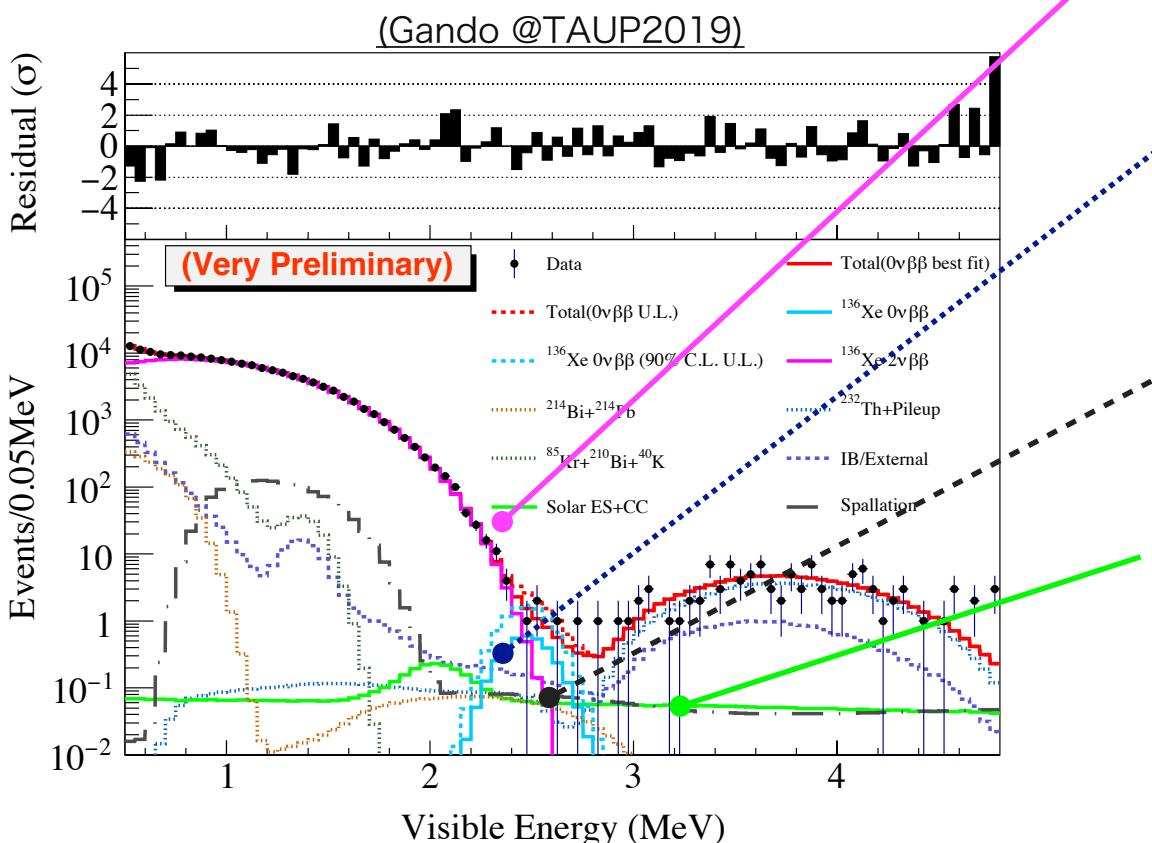
Dust from workers is the main source of contamination.



Backgrounds ($R < 157$ cm)

Zen 800 ROI: $2.35 \text{ MeV} < E < 2.70 \text{ MeV}$,
 0v $\beta\beta$ efficiency $\sim 90\%$

Source: $2\nu\beta\beta$
 Radioactive impurities: Xe-LS, IB and outer-LS
 Cosmogenic: muon-spallation Solar neutrino ES



(Gando @TAUP2019)

$2\nu\beta\beta$ tail due to energy resolution
 Most dominant & inevitable BG
 $\Delta E \sim 4.5\% @ 2.45 \text{ MeV} \sim 5.1 \text{ events/ROI}$

^{214}Bi from IB film

Vertex resolution $\sim 15 \text{ cm}/\sqrt{E(\text{MeV})}$
 Limit sensitive region to 0v $\beta\beta$ decay.
 $\sim 0.9 \text{ events/ROI}$

^{12}C spallation products

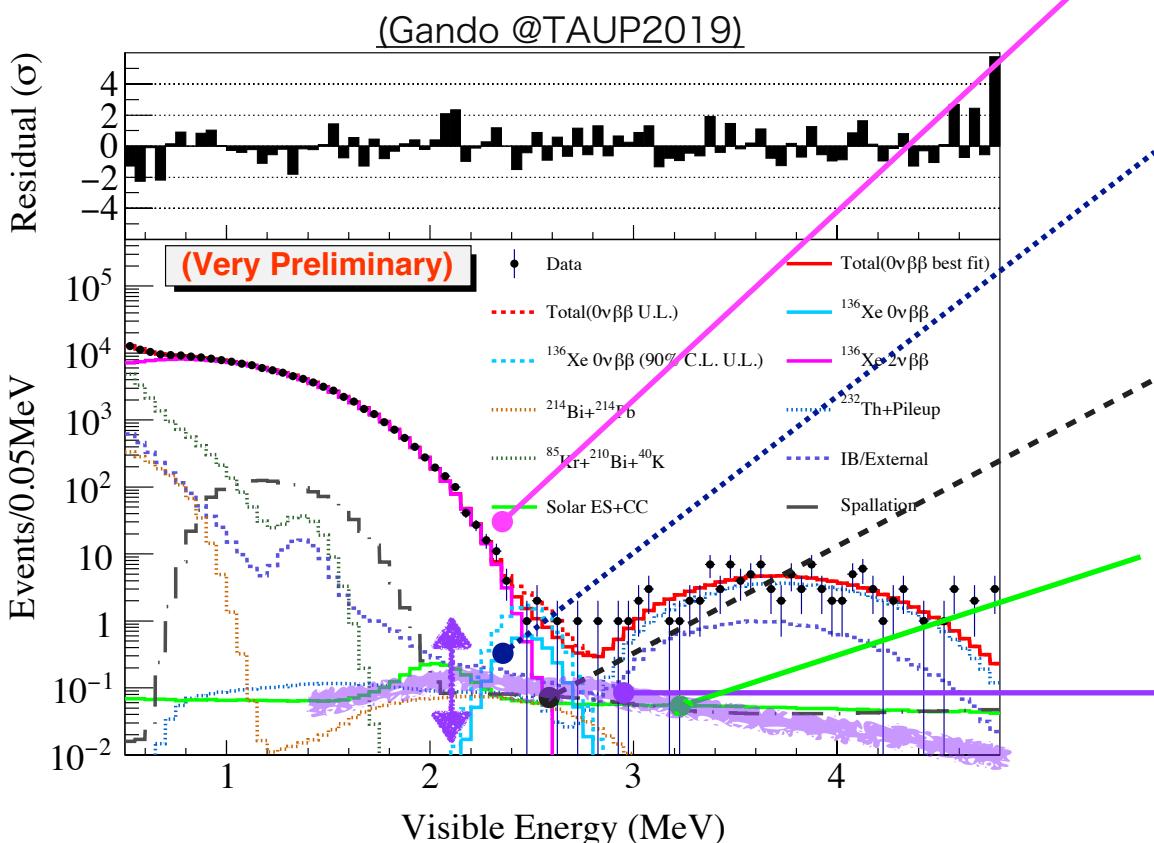
Significantly reduced by new analysis methods
 $\sim 0.2 \text{ events/ROI}$

Solar neutrino electron scattering + C.C. (^{136}Cs)
 Inevitable BG
 $\sim 0.4 \text{ events/ROI}$

Backgrounds ($R < 157$ cm)

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Solar neutrino electron scattering + C.C. (^{136}Cs)

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^{136}Xe spallation products

Nuisance BG in Zen800

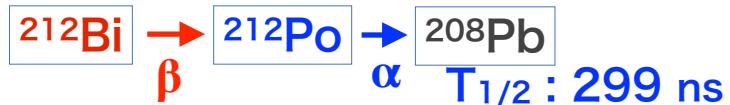
Hard to reject because of their long-lives

Radioactive impurity (Xe-LS)

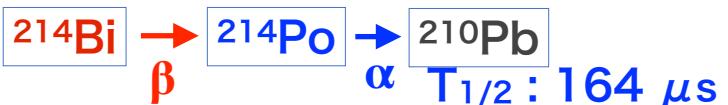
BGs ^{238}U series ... ^{214}Bi
 ^{232}Th series ... Pileup($^{212}\text{Bi} + ^{212}\text{Po}$)

U and Th contamination in Xe-LS measured using delayed coincidence ($1.\beta + 2.\alpha$)

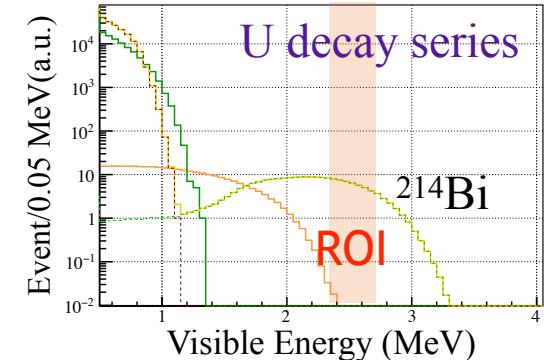
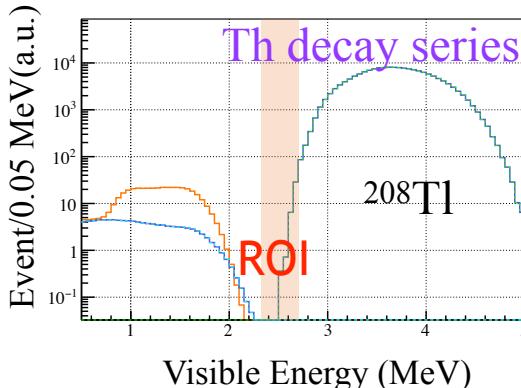
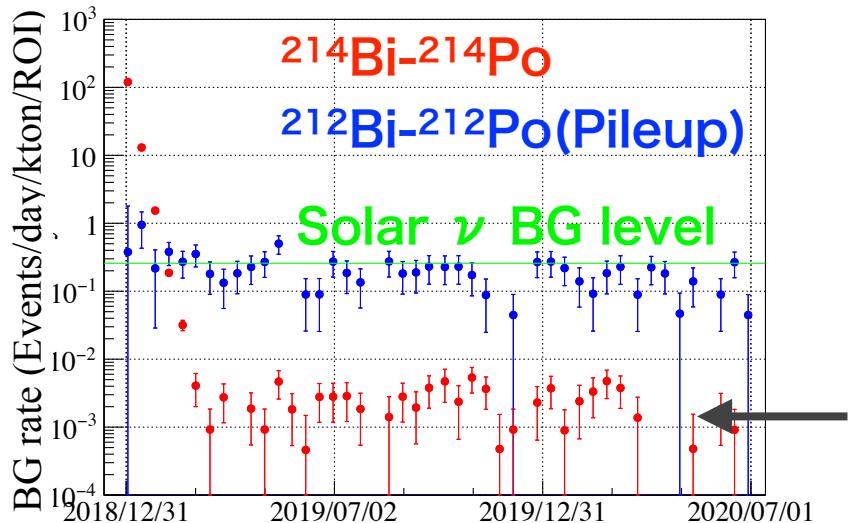
^{232}Th daughters ~70% eff.



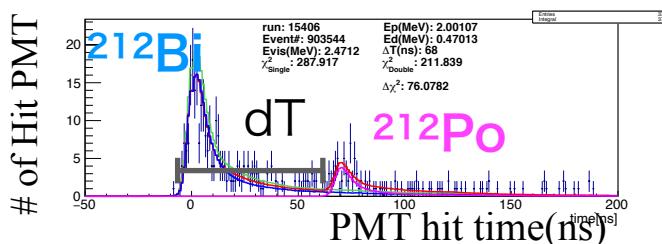
^{238}U daughters > 99.9% eff.



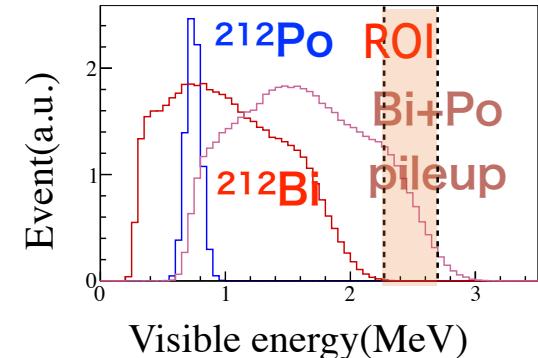
Trends of DC events



Pileup BG from Th daughters



Pileup if $dT \sim 10$ nsec (~2%)



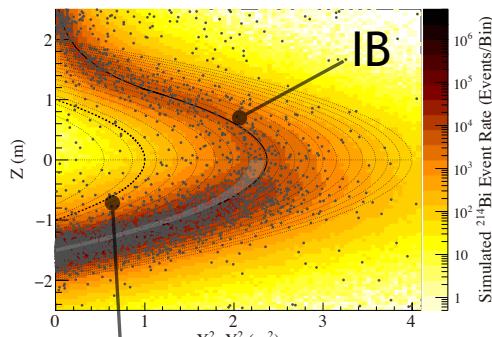
BG from RI in Xe-LS is negligibly small!!

Radioactive impurity in Xe-LS
 $^{238}\text{U} \sim \text{O}(10^{-17}) \text{ g/g Xe-LS}$
 $^{232}\text{Th} \sim 3 \times 10^{-16} \text{ g/g Xe-LS}$

Radioactive impurity (IB film)

^{214}Bi distribution from IB

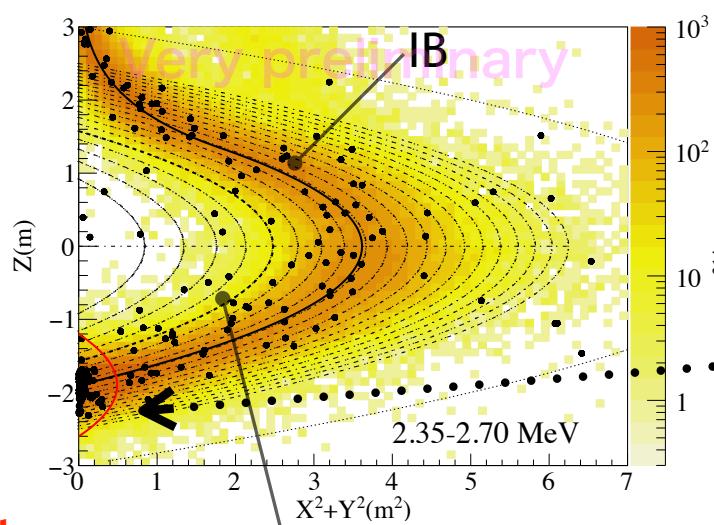
Zen 400 Phase-II



Sensitive region
 $r < 1.06 \text{ m } z < 0$
 $r < 1.26 \text{ m } z > 0$

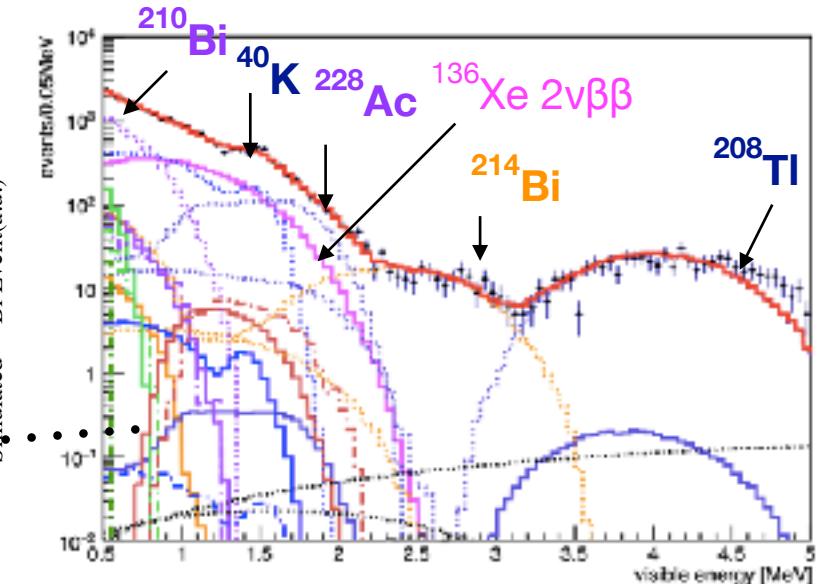
$\times >3$ scale-up

Zen 800 (Gando @TAUP2019)



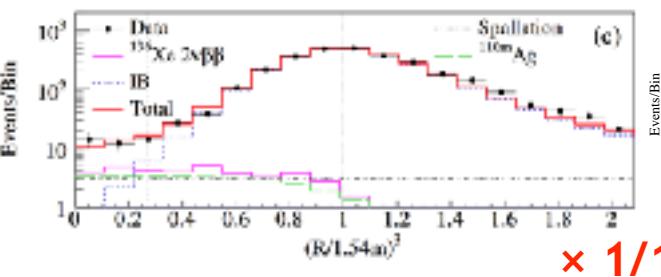
Sensitive region $r < 1.57 \text{ m}$

Bottom Hot Spot



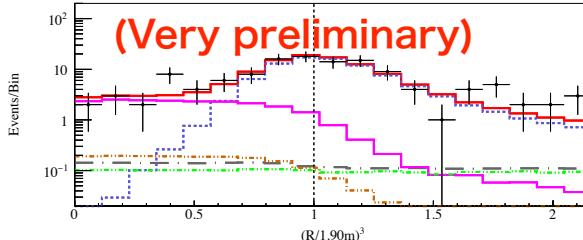
Sinking dusts?
 Will be removed from analysis volume.

0ν region ($2.30 < E < 2.70 \text{ MeV}$)



$\times 1/10 \text{ BG}$

0ν region ($2.35 < E < 2.70 \text{ MeV}$)

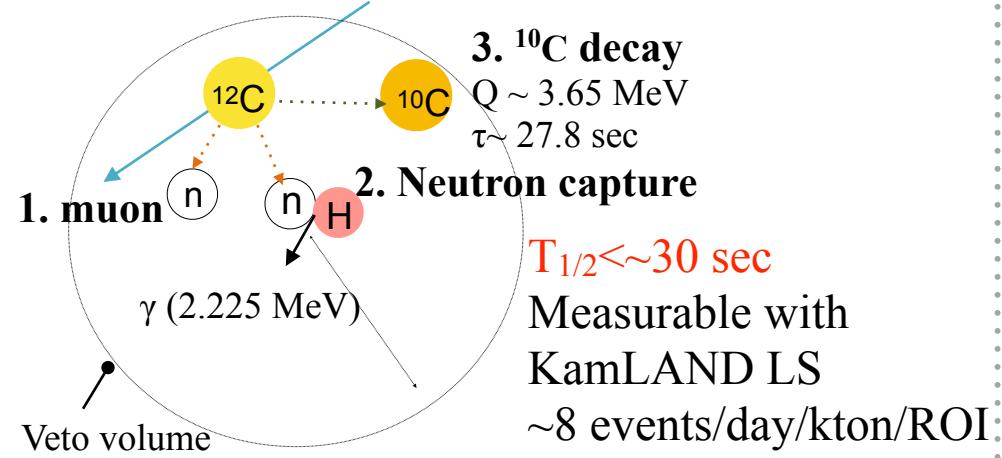


- $2\nu \beta\beta$
- - Balloon
- - $\text{Solar } \nu$
- - Spallation
- - U,Th in

Radioactive impurity in IB
 $^{238}\text{U} \sim 3 \times 10^{-12} \text{ g/g IB}$
 $^{232}\text{Th} \sim 4 \times 10^{-11} \text{ g/g IB}$

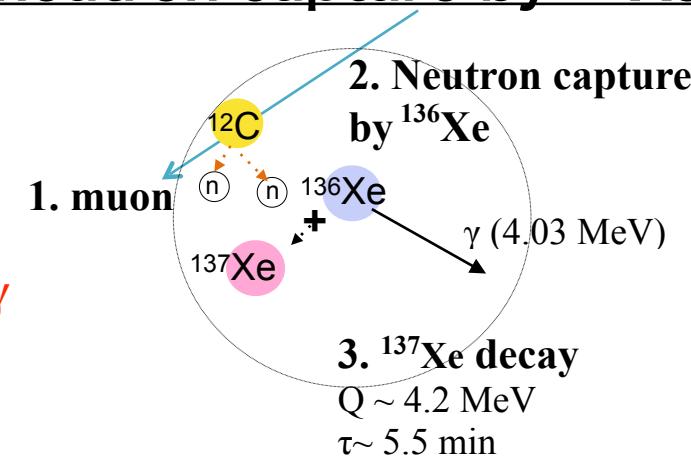
Backgrounds from cosmic-ray muons

^{12}C spallation



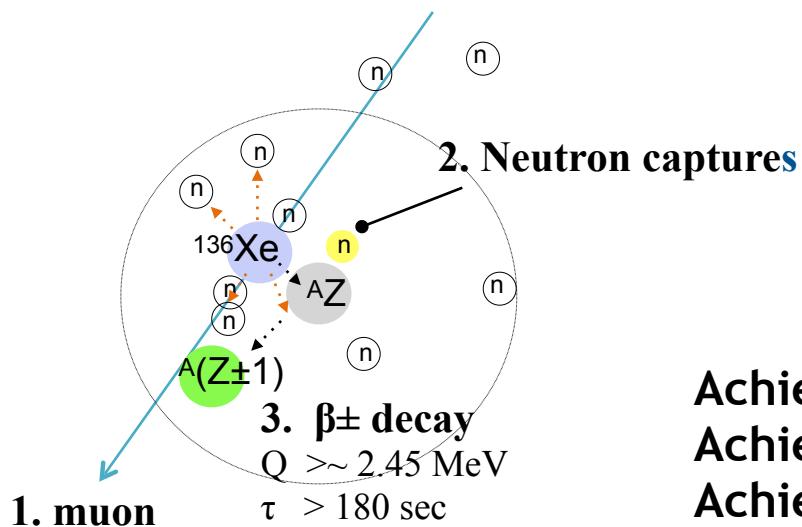
^{137}Xe decay from neutron capture by ^{136}Xe

$\sim 0.5 \pm 0.2$ events/day/ton/ROI
Long-life $\tau \sim 5.5$ min
Tagged by **high energy γ**
 ~ 4 MeV



^{136}Xe spallation

More than 30 main nuclei
Long-life $T_{1/2} \sim 0$ (hours)
High neutron multiplicity



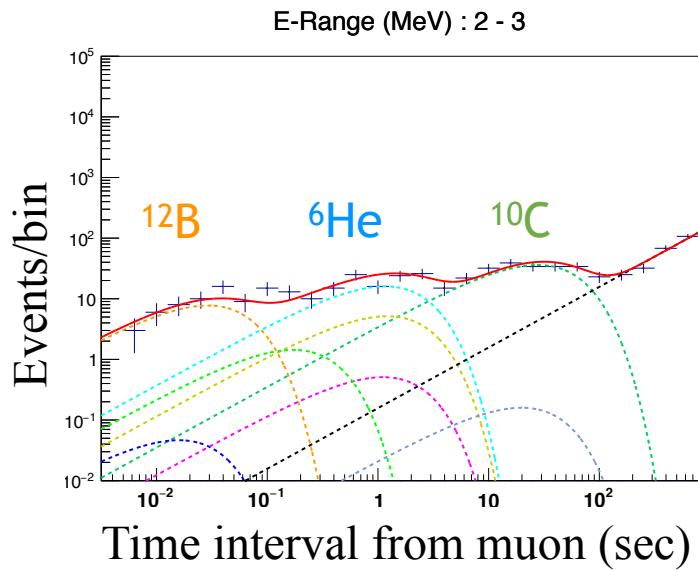
Background rejection methods

1. After-muon full volume veto
2. Muon-neutron tagging
3. Muon-shower tagging (New)

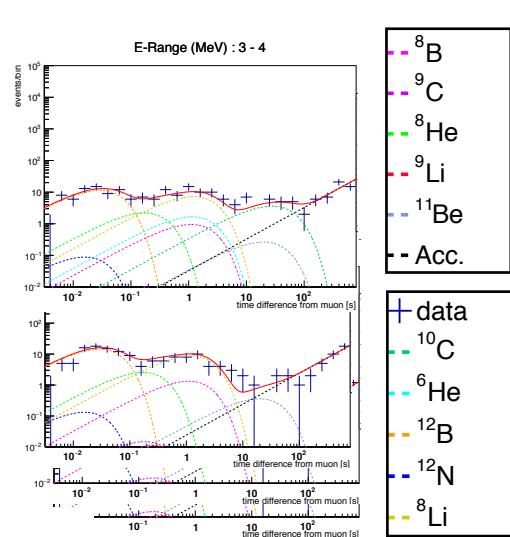
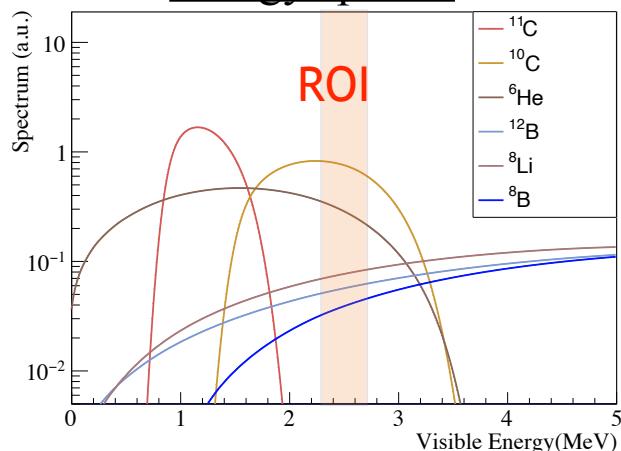
Achieved **more than 90%** ^{10}C rejection
Achieved **~80%** ^{137}Xe rejection
Achieved **~40%** ^{136}Xe spallation rejection
with **~10%** $0\nu\beta\beta$ inefficiency

^{12}C -spallation

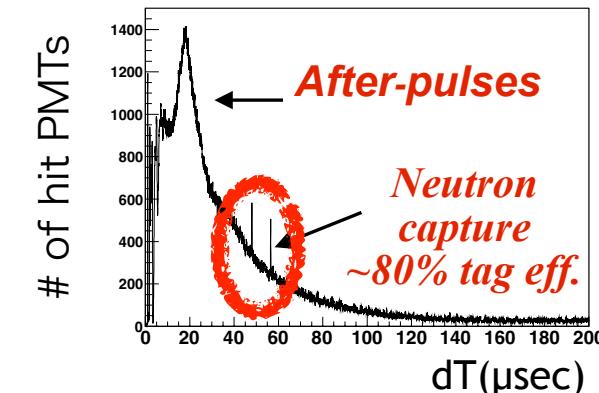
Simultaneous fit with energy and dT spectra



Energy spectra



PMT hit timing after muon

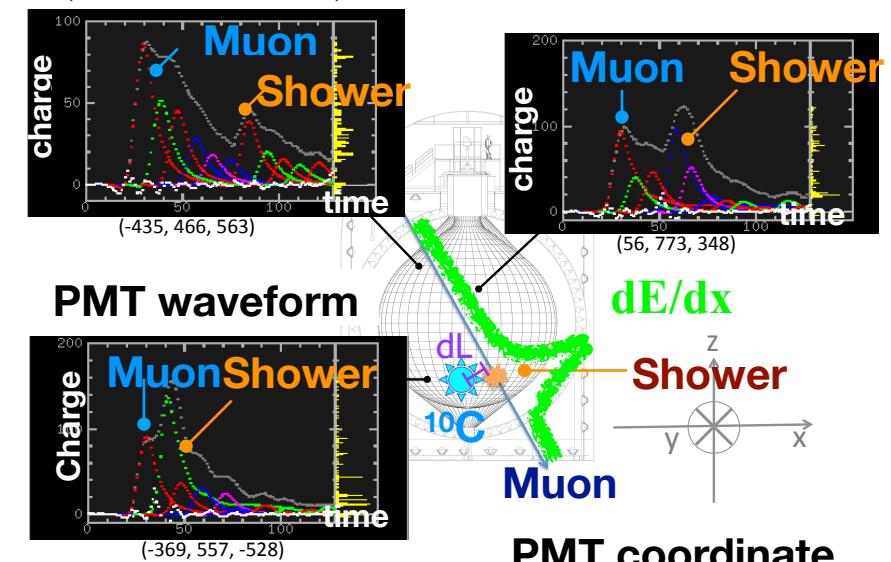


| | Production rate [/day/kton] |
|-----------------|-----------------------------|
| ^6He | 15 ± 2 |
| ^8Li | 12 ± 3 |
| ^{10}C | 20 ± 1 |
| ^{12}B | 51 ± 2 |

After muon 150 ms veto
+ muon-neutron-spallation
+ muon-shower-spallation

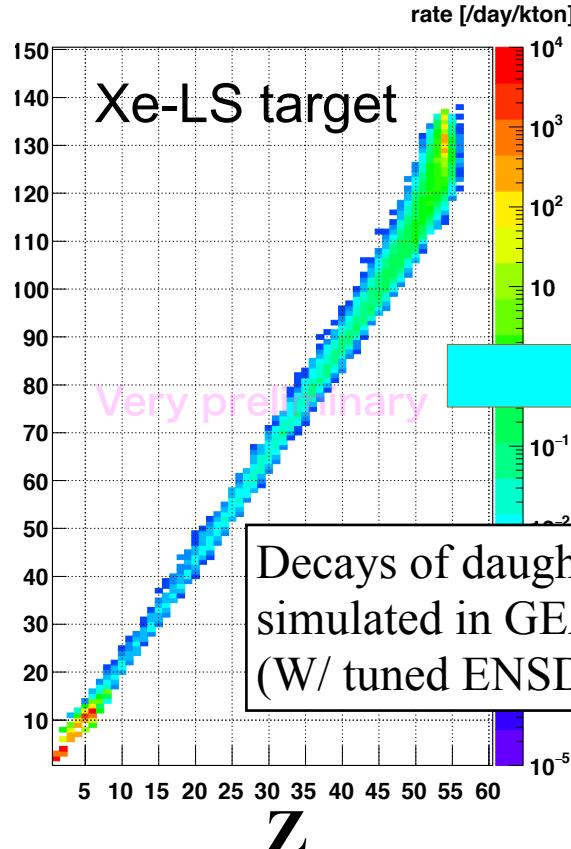
Remaining
 $\sim <0.1$ /day/kton/ROI
 $> 90\%$ rejection
Less than solar ν BGs
(~ 0.3 /day/kton/ROI)

Spallation veto with shower correlation (dE/dx and dL)

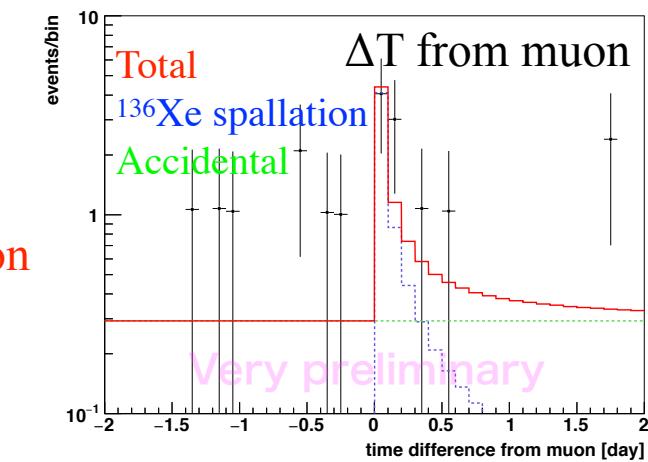
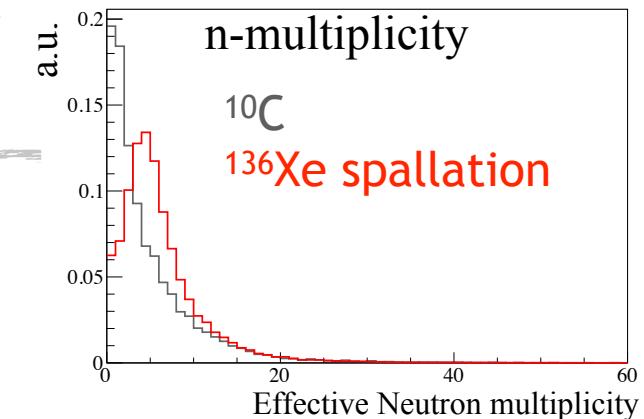
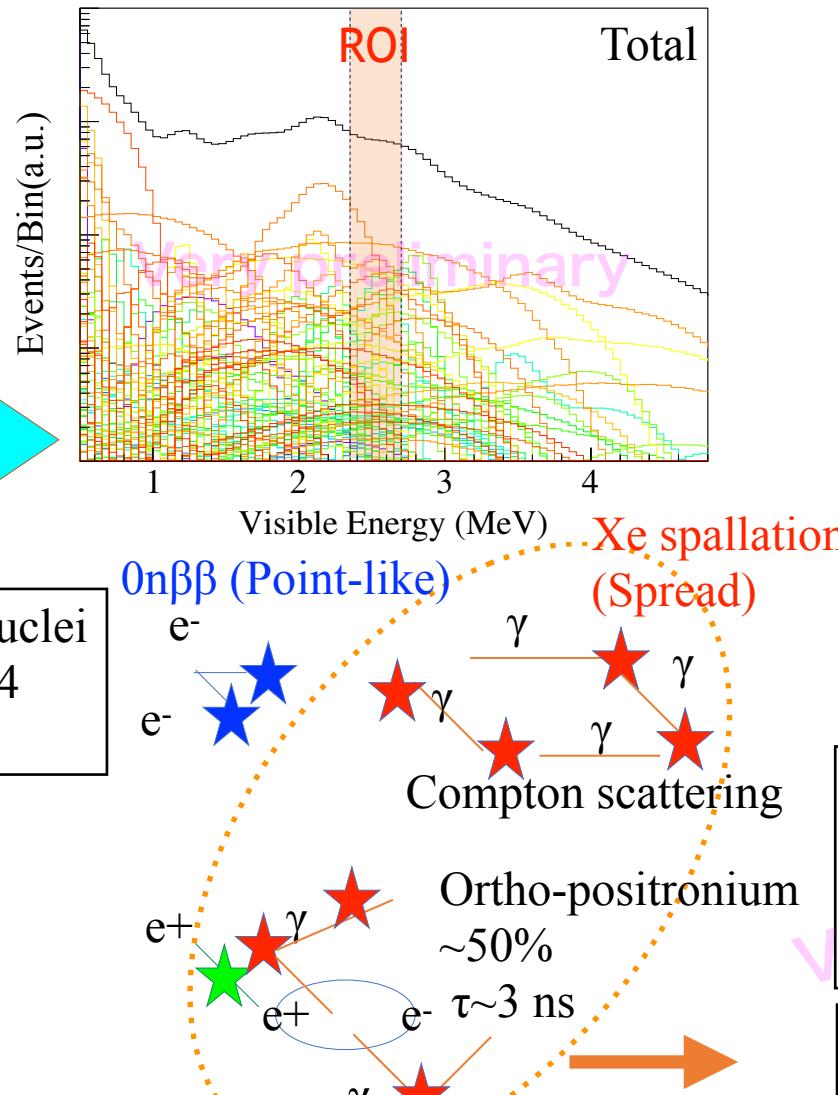


^{136}Xe -spallation

^{136}Xe spallation products



Production rate was estimated by FLUKA.



Neutron tag with n-multiplicity, dR, dT
~40% rejection with ~10% 0v $\beta\beta$ ineff.

PID by neural network
Under development

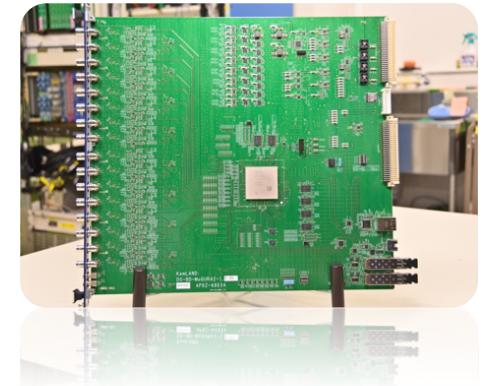
Prospects for Improvements and future plans

Started Jan. 2019



KamLAND-Zen ~ 800 kg phase
with new larger and cleaner inner-balloon

- PMT AMP for better energy resolution
- New electronics to collect all neutron captures
- Source calibration



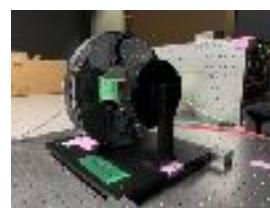
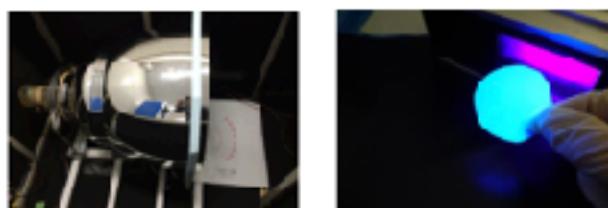
~5 years observation

Future plan

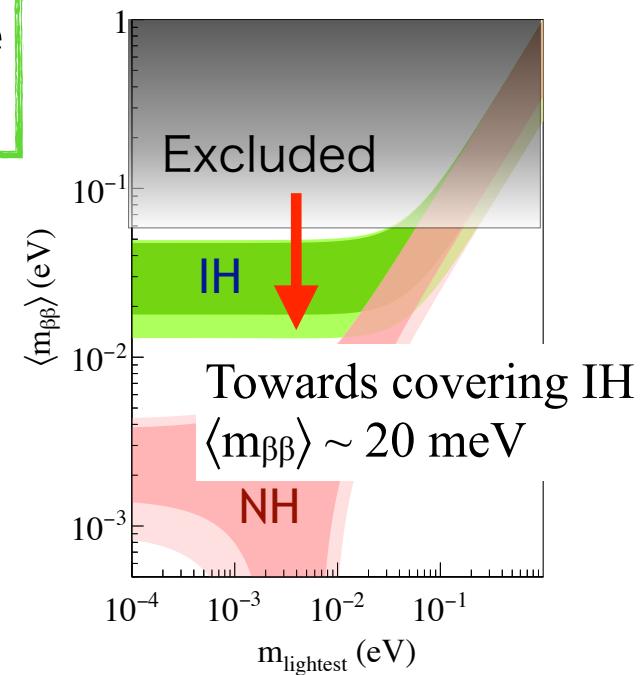
KamLAND2-Zen >1000 kg phase
with remodeled detector

- HQE PMT
- LAB LS
- Mirror for PMT
- Scintillating inner balloon
- Imaging camera

} > $\times 5$ brighter
Energy resolution
 $\sim 2\%$ @ 2.45 MeV
> $\times 2$ improvement



R&D
Ongoing



Summary

- KamLAND-Zen 800 is stably running with the goal of searching for $0\nu\beta\beta$ within the IH region.
- We successfully characterized and reduced backgrounds.
 - Constructed an ultra-low background IB
 - Strongly reduced ^{12}C -spallation by analysis
 - Successfully characterized ^{136}Xe -spallation
- The latest results will be published very soon.