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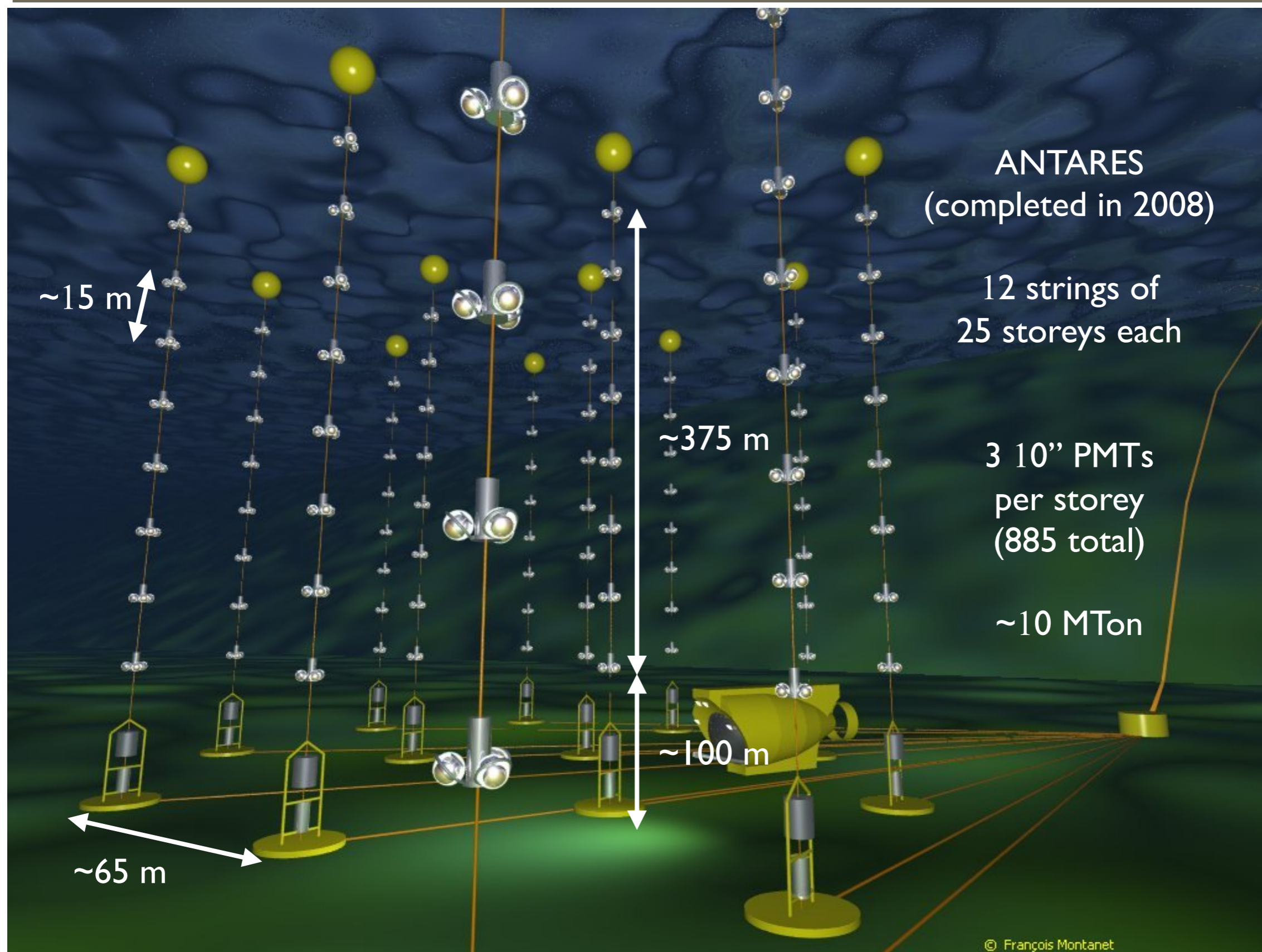
CONSEJO SUPERIOR DE INVESTIGACIONES CIENTÍFICAS

# Search for neutrino non-standard interactions with 10 years of ANTARES data and perspectives for KM3NeT/ORCA

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# ANTARES



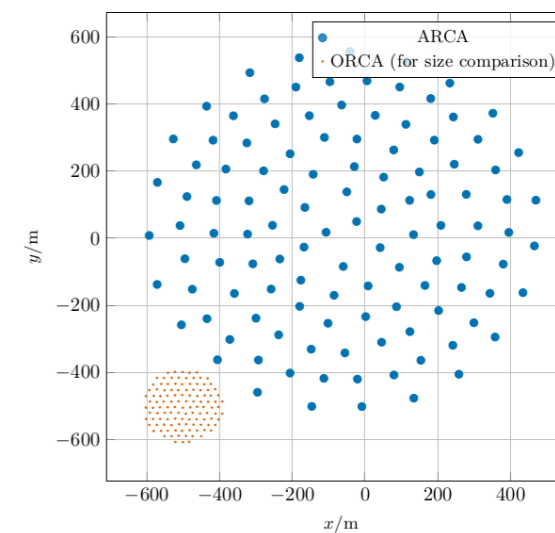
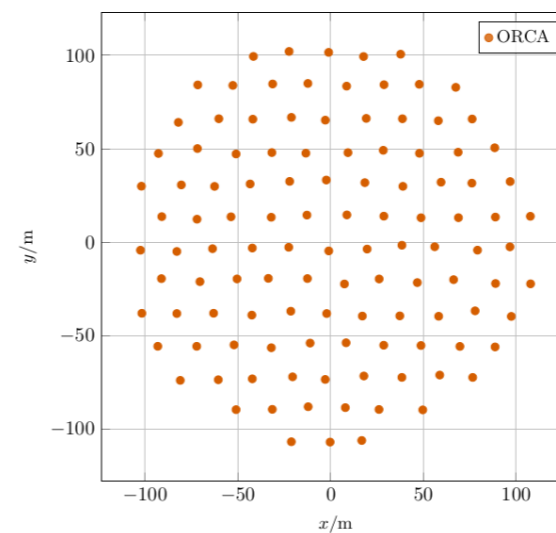
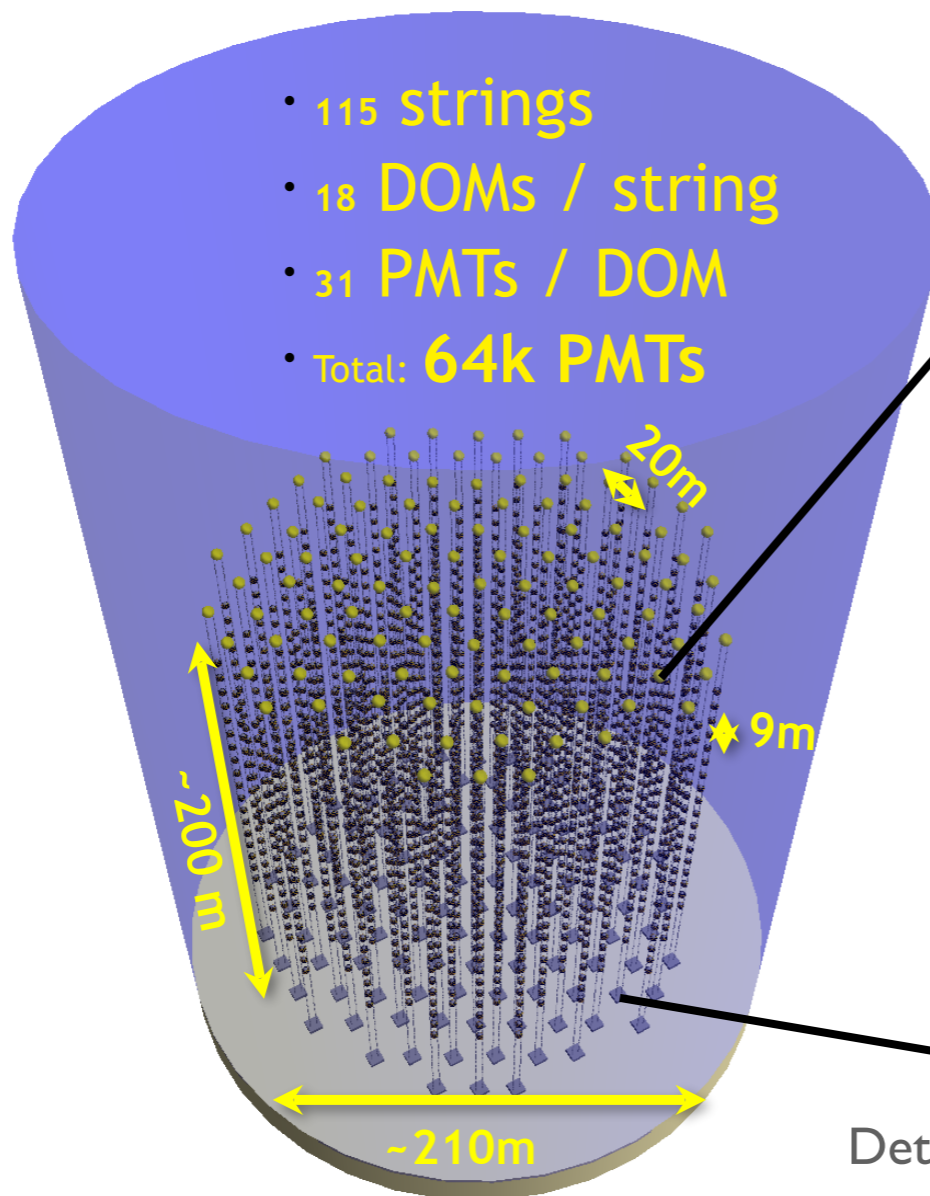
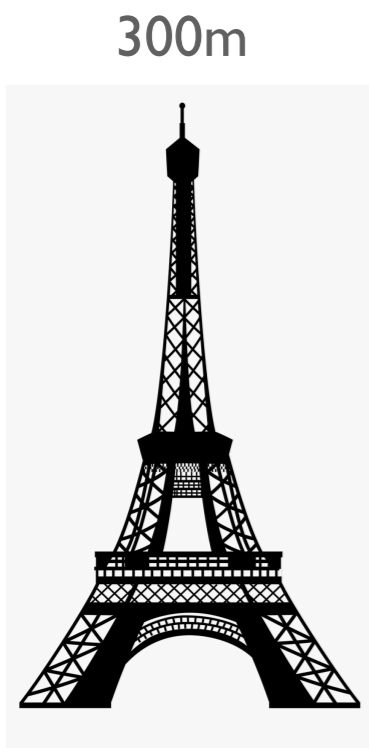
# KM3NeT/ORCA & ARCA

ORCA: Low energy physics  $\sim$  few GeV  
Neutrino Oscillations, Supernovae ....

ARCA: High energy physics  $\sim$  GeV - PeV  
High energy astronomy, DM searches ...

ORCA instrumented Volume  $\sim$  5 Mton

- 115 strings
- 18 DOMs / string
- 31 PMTs / DOM
- Total: 64k PMTs



ORCA

ARCA



Detection Unit (DU)  
aka "string/line"

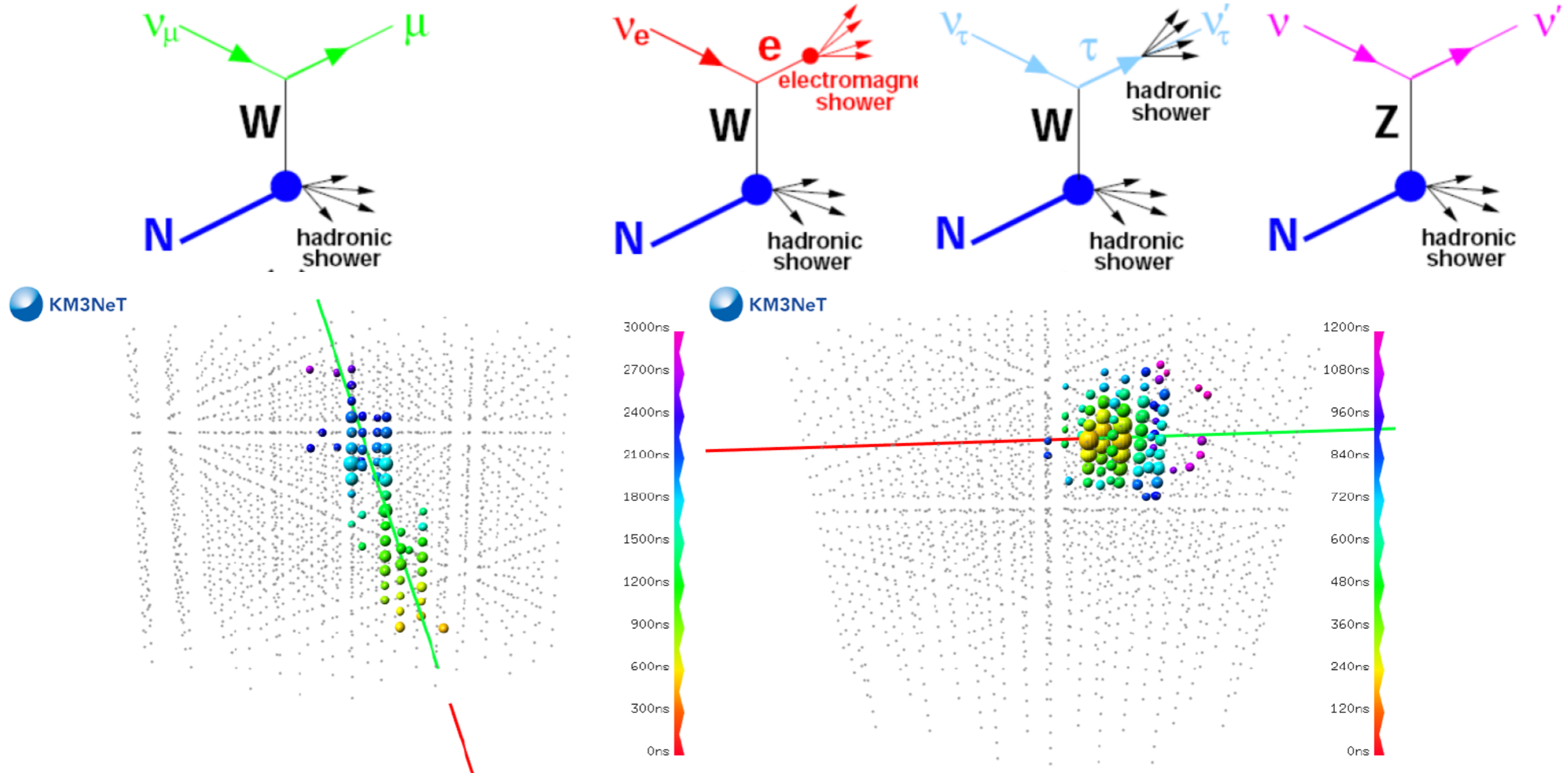


KM3NeT DOM  
housing 31 3" PMTs

# Event Topologies

Tracks:  $\nu_{\mu}^{CC}$ ,  $\nu_{\tau}^{CC}$  (muonic decay)

Showers:  $\nu_e^{CC}$ ,  $\nu_{\tau}^{CC}$ ,  $\nu^{NC}$



Position, time and charge pattern within the detector used to reconstruct particle direction and energy.

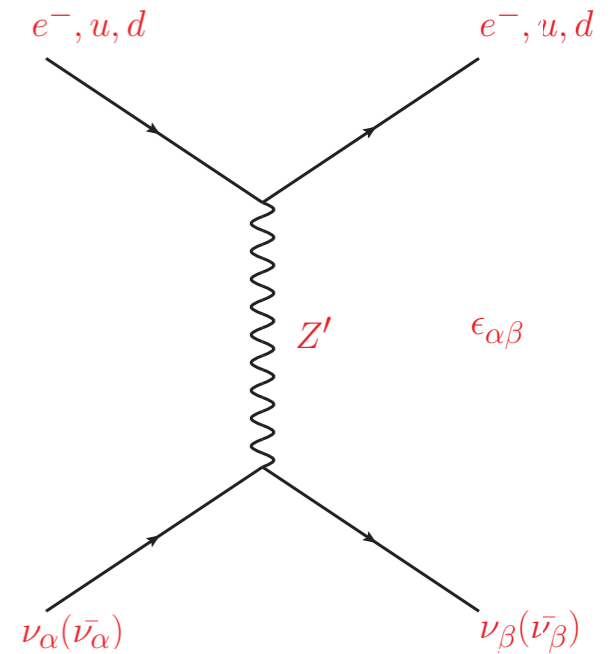
The Cherenkov signature of the outgoing lepton is used for Particle IDentification (PID).

# Non-Standard Interactions (NSIs)

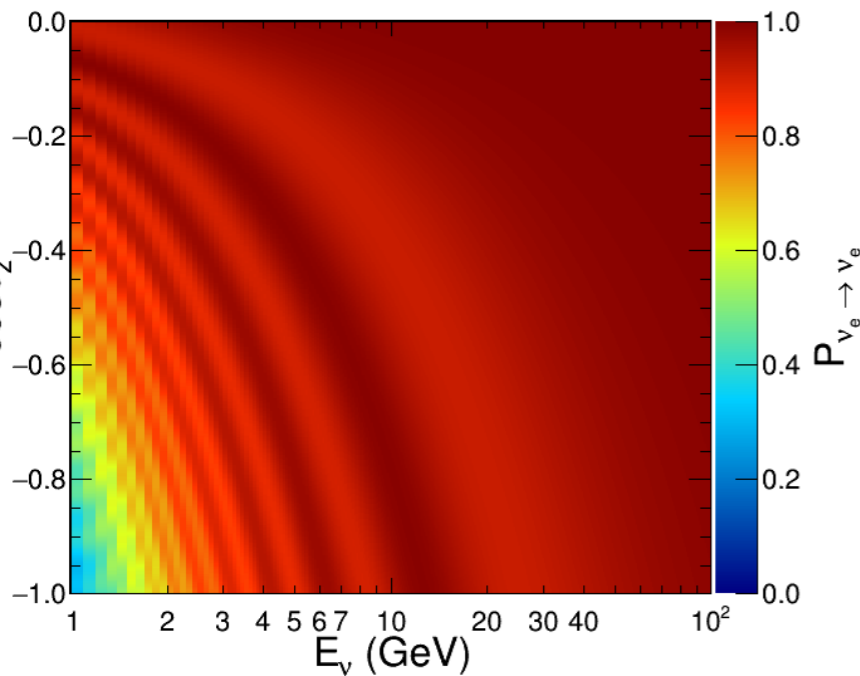
- The neutrino propagation Hamiltonian in presence NSIs:

$$H = \frac{1}{2E} U_{PMNS} \begin{bmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{bmatrix} U_{PMNS}^\dagger + \begin{bmatrix} A_{CC} & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix} + \sqrt{2} G_F N_d(x) \begin{bmatrix} \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{bmatrix}$$

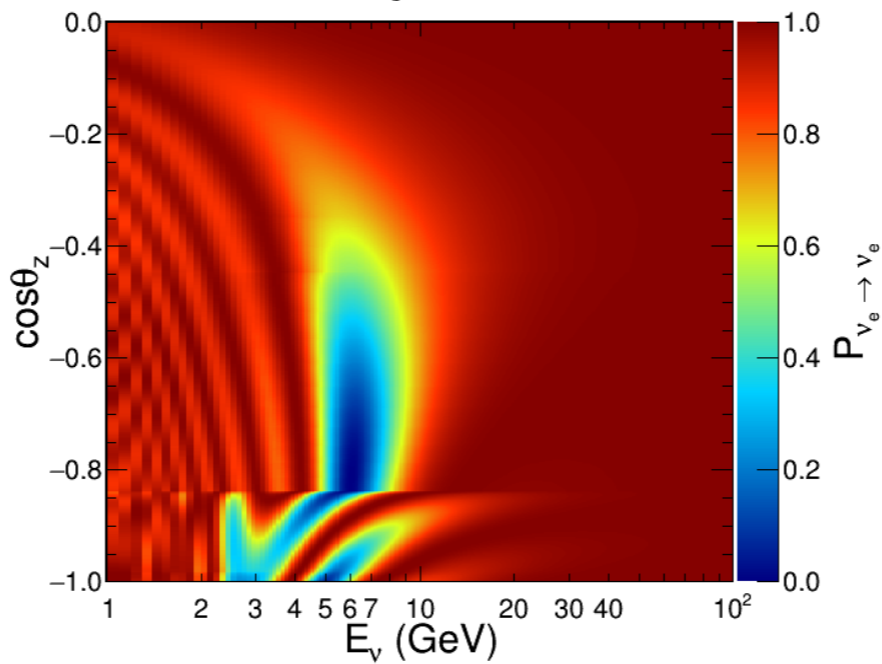
Vacuum                      SM MSW                      NSI



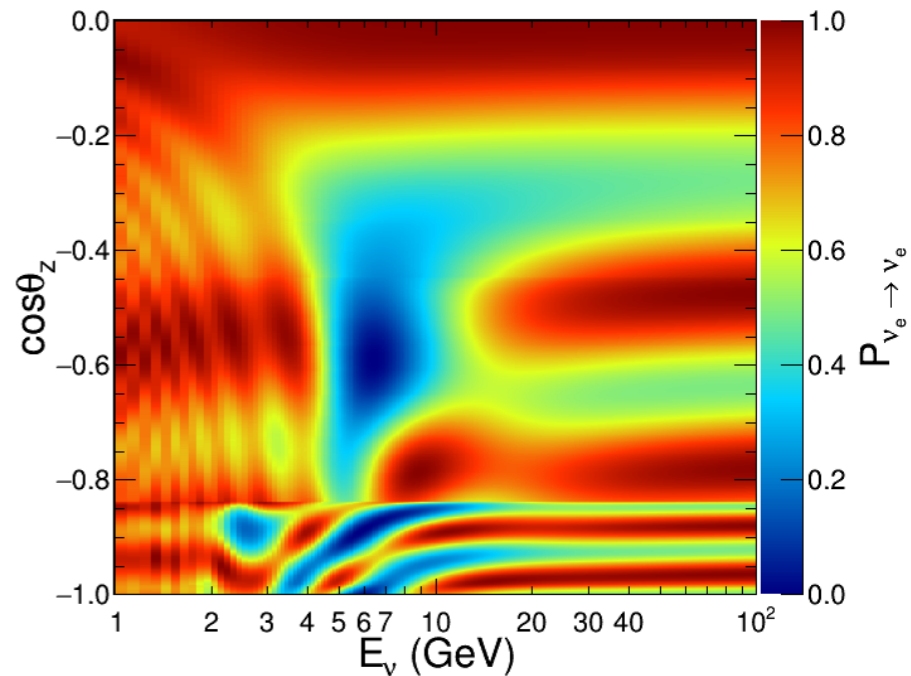
Normal Ordering, Oscillations in vacuum



Normal Ordering, Oscillations in matter



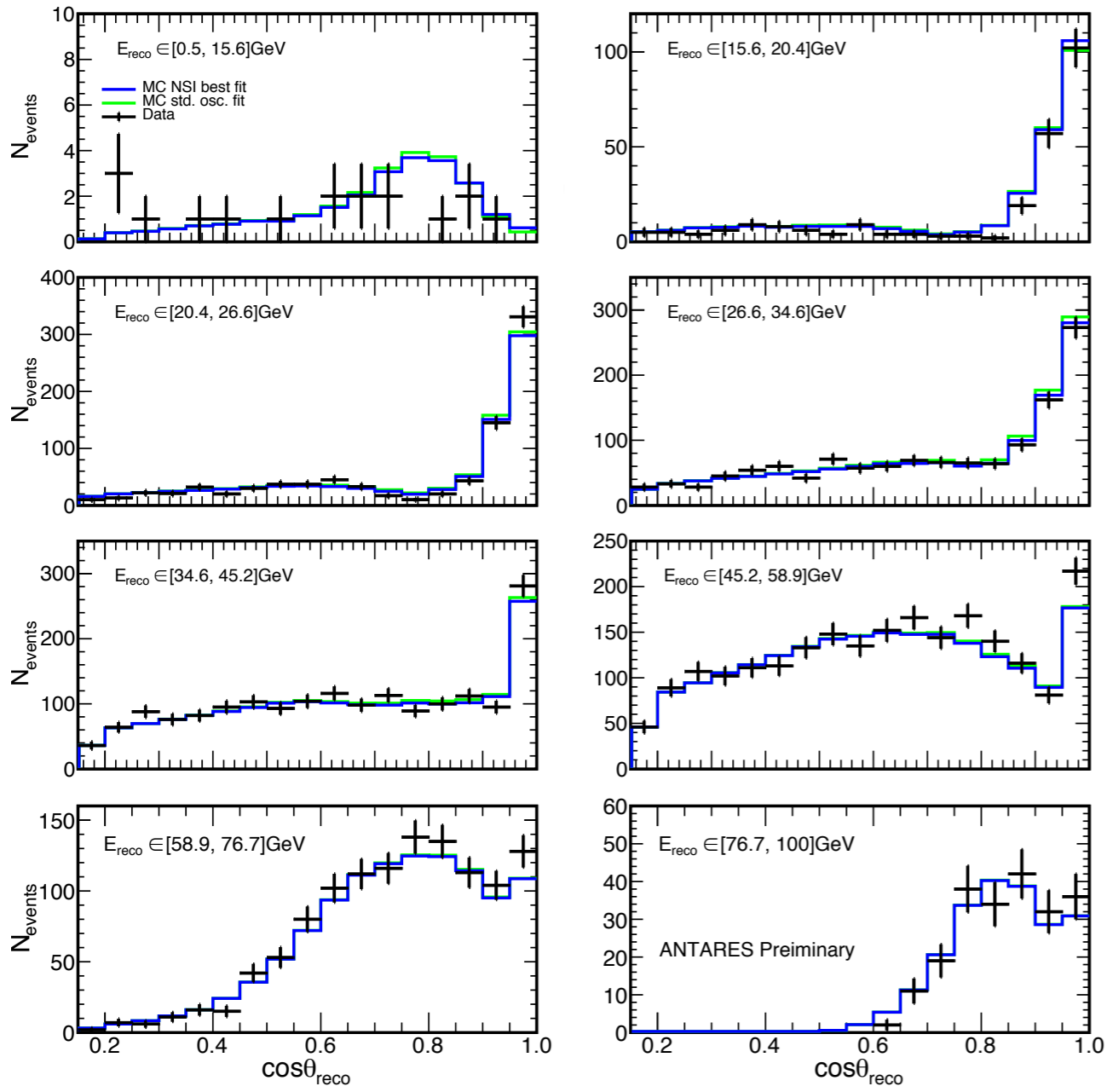
Normal Ordering, Oscillations with NSI



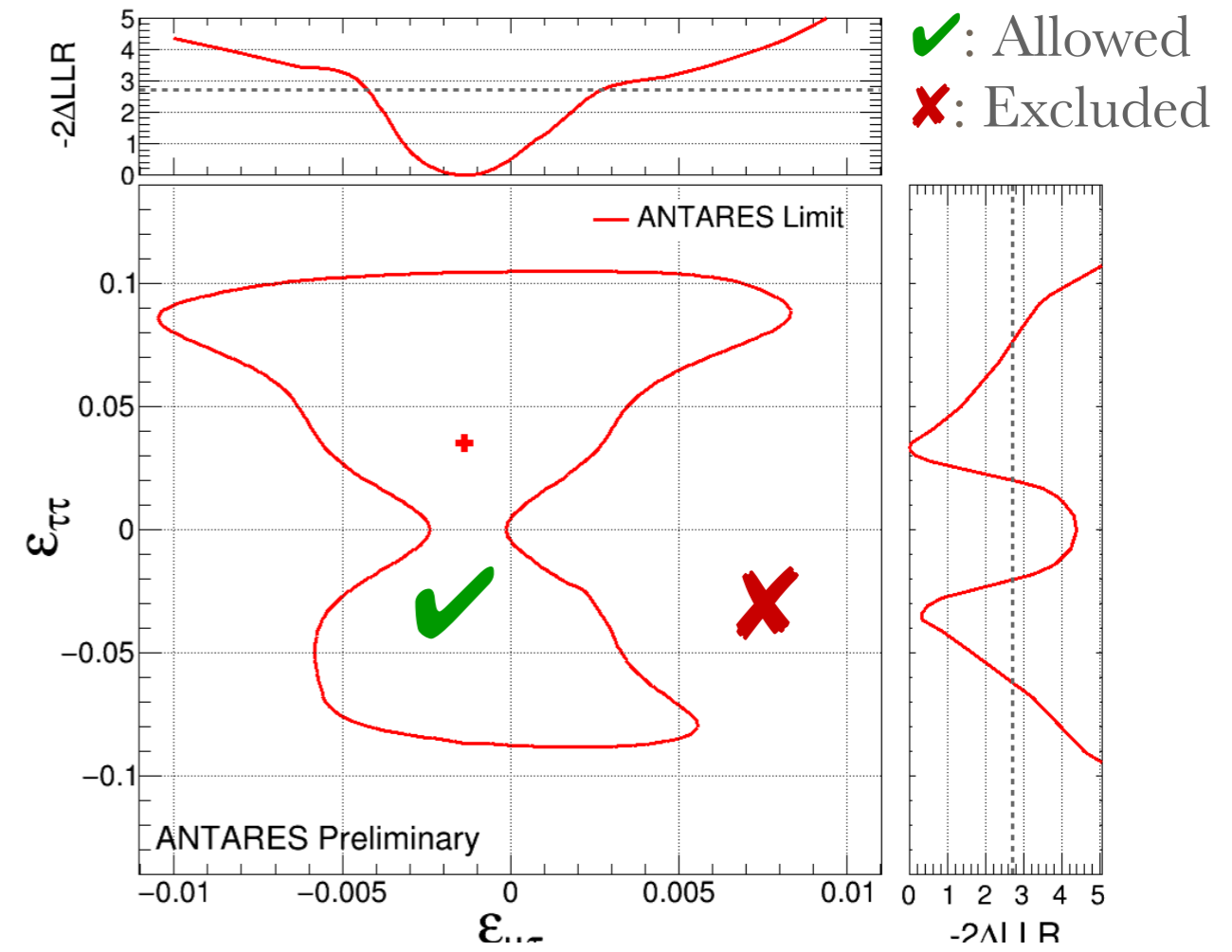
- Signature at detector: a statistical excess/deficit of  $(\nu + \bar{\nu})$  events from anomalous flavour transitions expected with standard oscillations predictions.

# ANTARES Limits

- ANTARES dataset corresponding to 2007-2016 (both years included) has been used.



best-fit points  
 $\epsilon_{\mu\tau} : -1.3e-3$   
 $\epsilon_{\tau\tau} : 3.2e-2$   
 $\chi^2 / \text{d.of.} : 152.36 / 136.$



$$-4.2 \times 10^{-3} < \epsilon_{\mu\tau} < 2.7 \times 10^{-3} \quad (\text{at } 90\% \text{ C.L.}),$$

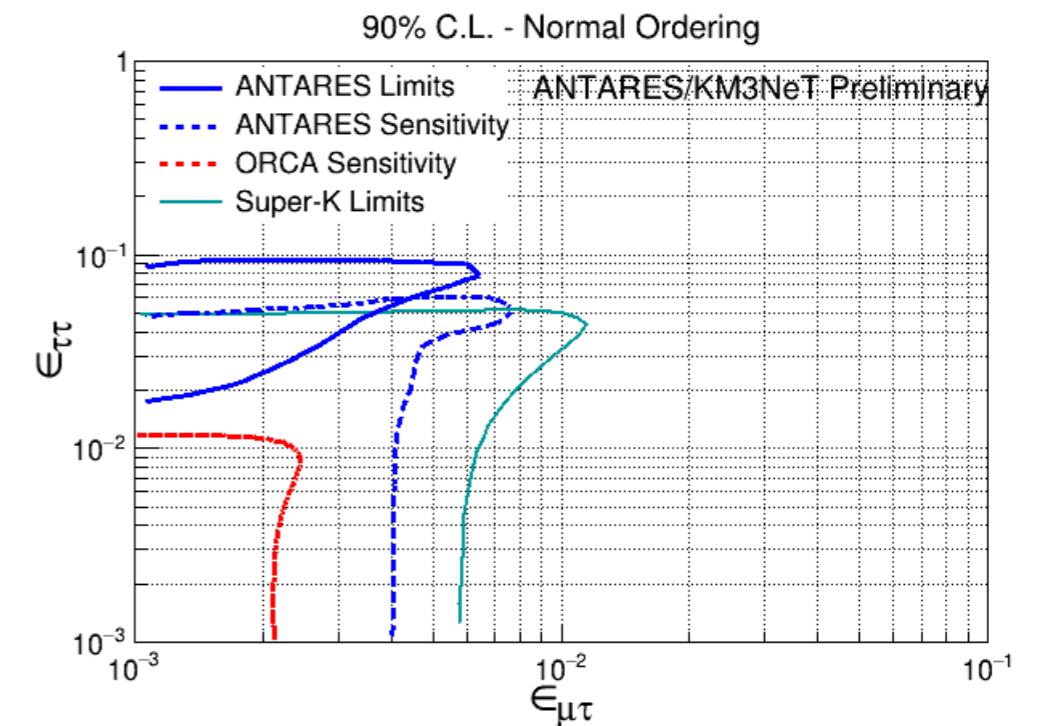
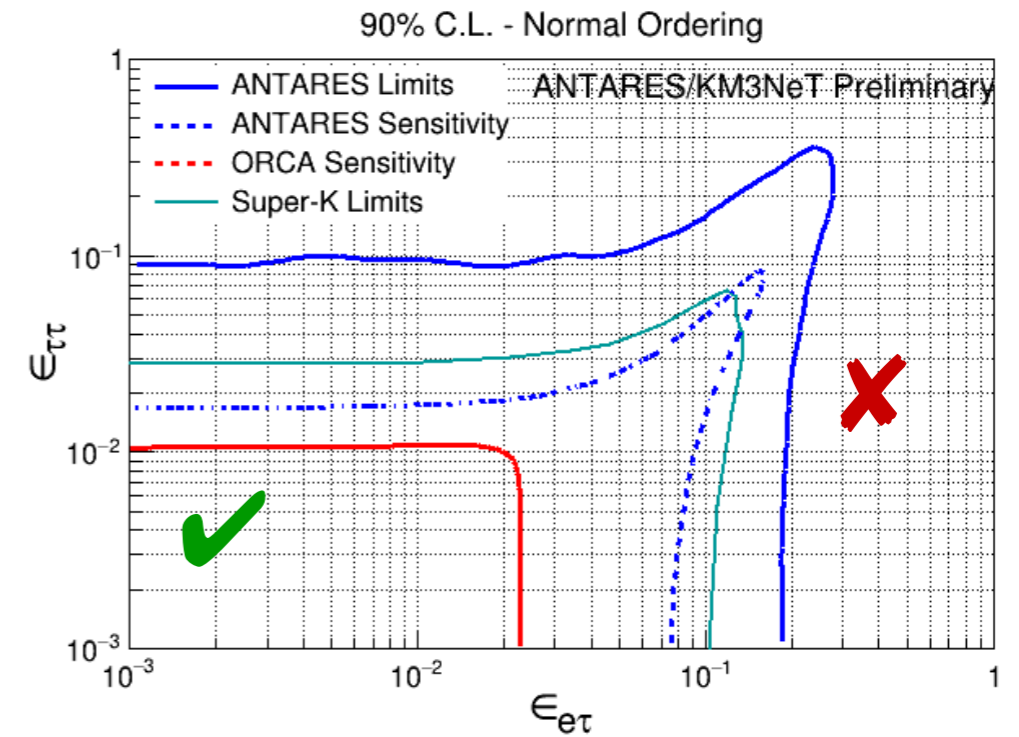
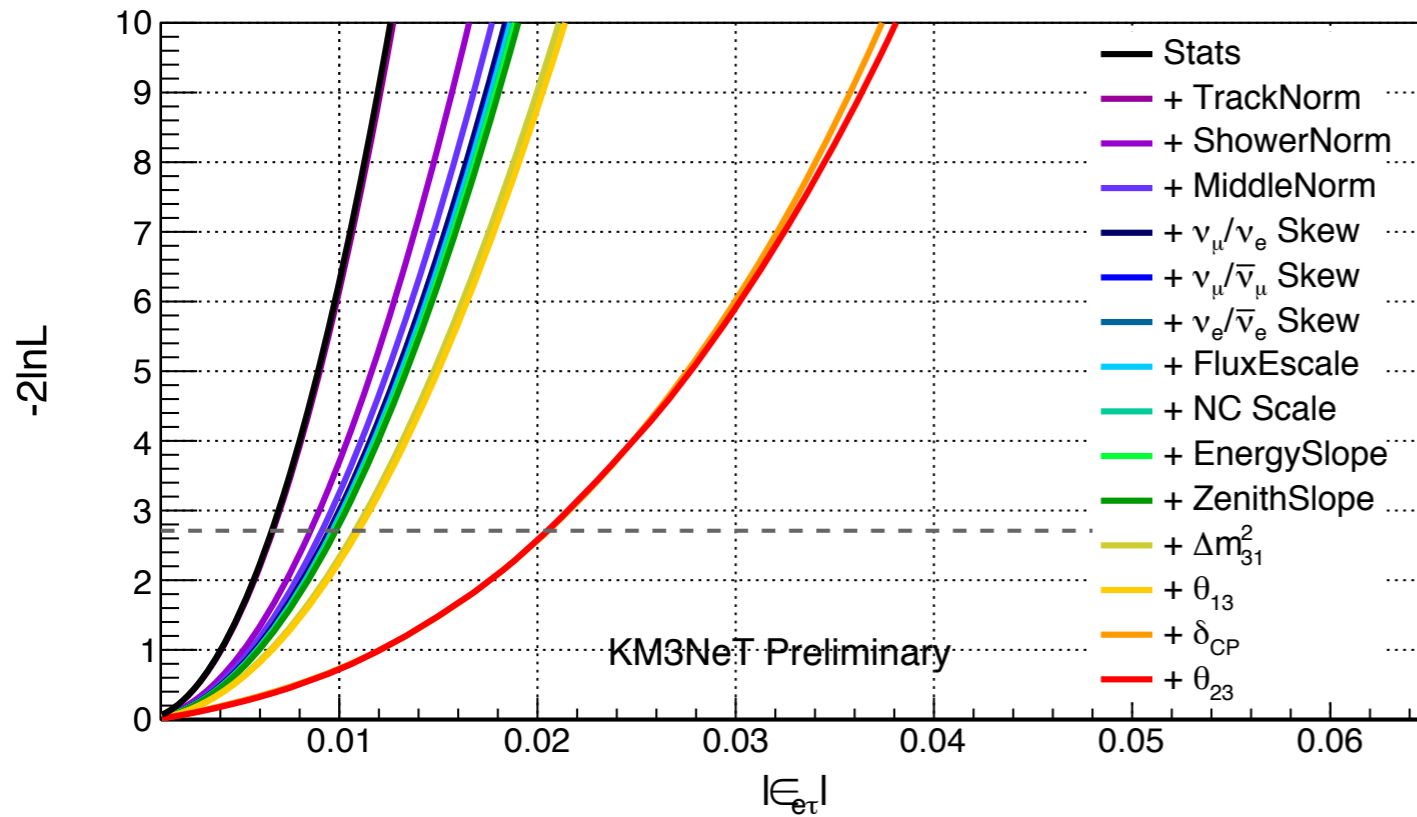
$$-6.1 \times 10^{-2} < \epsilon_{\tau\tau} < -2.1 \times 10^{-2} \quad \text{and} \quad 2.1 \times 10^{-2} < \epsilon_{\tau\tau} < 7.3 \times 10^{-2} \quad (\text{at } 90\% \text{ C.L.}).$$

<https://nusoft.fnal.gov/nova/nu2020postersession/pdf/posterPDF-178.pdf>

**World's best limits!**

# ORCA Sensitivities

ORCA115 - 20m - 3 yr - Normal Ordering



NSI Couplings	Assumed True NMO	Bounds (90% C.L.)
$\epsilon_{e\mu}$	NO	$(-1.7 \times 10^{-2}, 1.7 \times 10^{-2})$
	IO	$(-2.0 \times 10^{-2}, 2.0 \times 10^{-2})$
$\epsilon_{e\tau}$	NO	$(-1.8 \times 10^{-2}, 2.1 \times 10^{-2})$
	IO	$(-3.1 \times 10^{-2}, 2.7 \times 10^{-2})$
$\epsilon_{\mu\tau}$	NO	$(-1.7 \times 10^{-3}, 1.7 \times 10^{-3})$
	IO	$(-1.7 \times 10^{-3}, 1.7 \times 10^{-3})$
$\epsilon_{\tau\tau}$	NO	$(-0.8 \times 10^{-2}, 1.1 \times 10^{-2})$
	IO	$(-1.1 \times 10^{-2}, 0.8 \times 10^{-2})$

3 year projected sensitivities towards NSI parameters at ORCA.

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# Outlook

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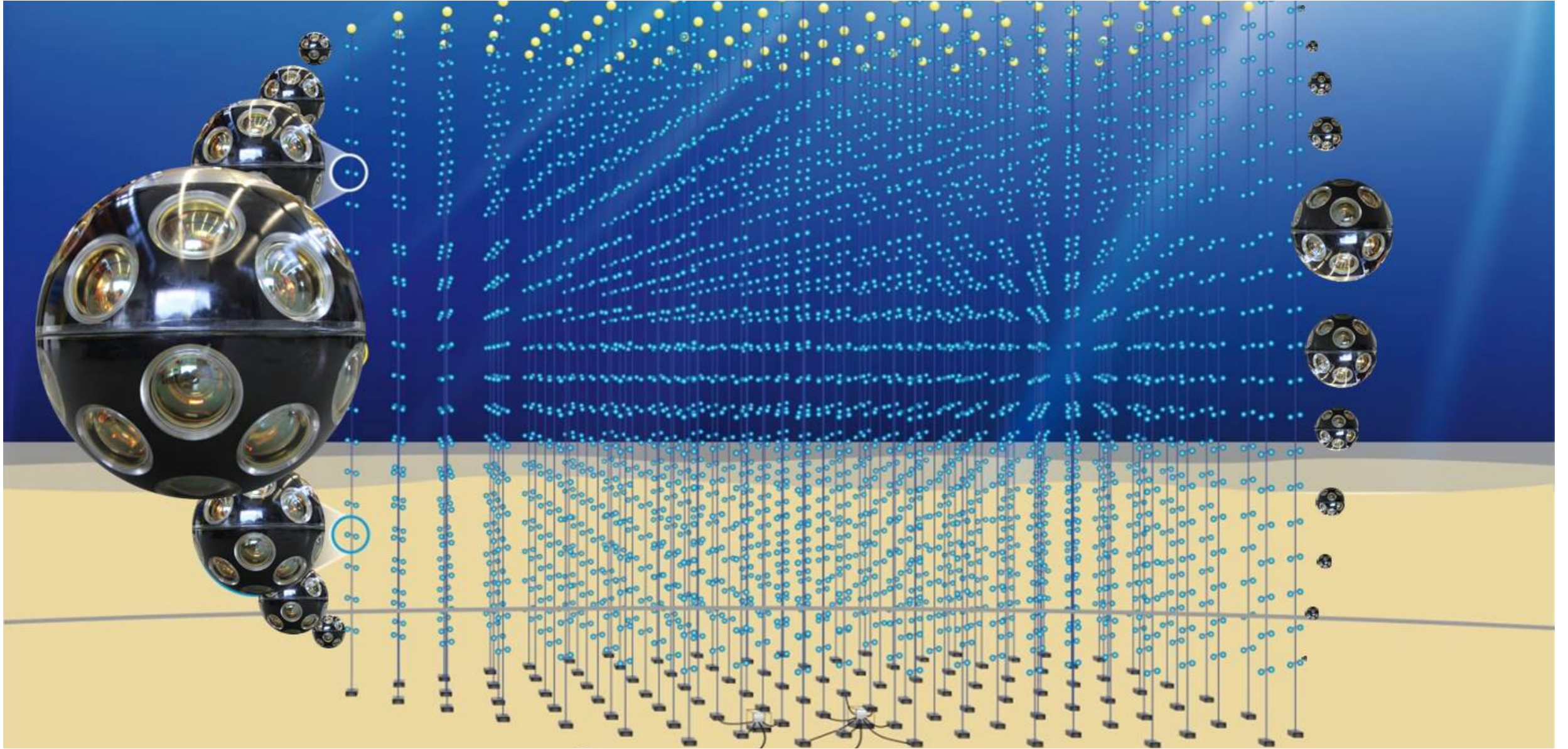
ANTARES has produced significant results in constraining many new physics scenarios by studying neutrinos at the abyss of the Mediterranean Sea.

KM3NeT will make further improvements.

6 ORCA (and 1 ARCA) strings are operational. (Stay tuned for first results!)

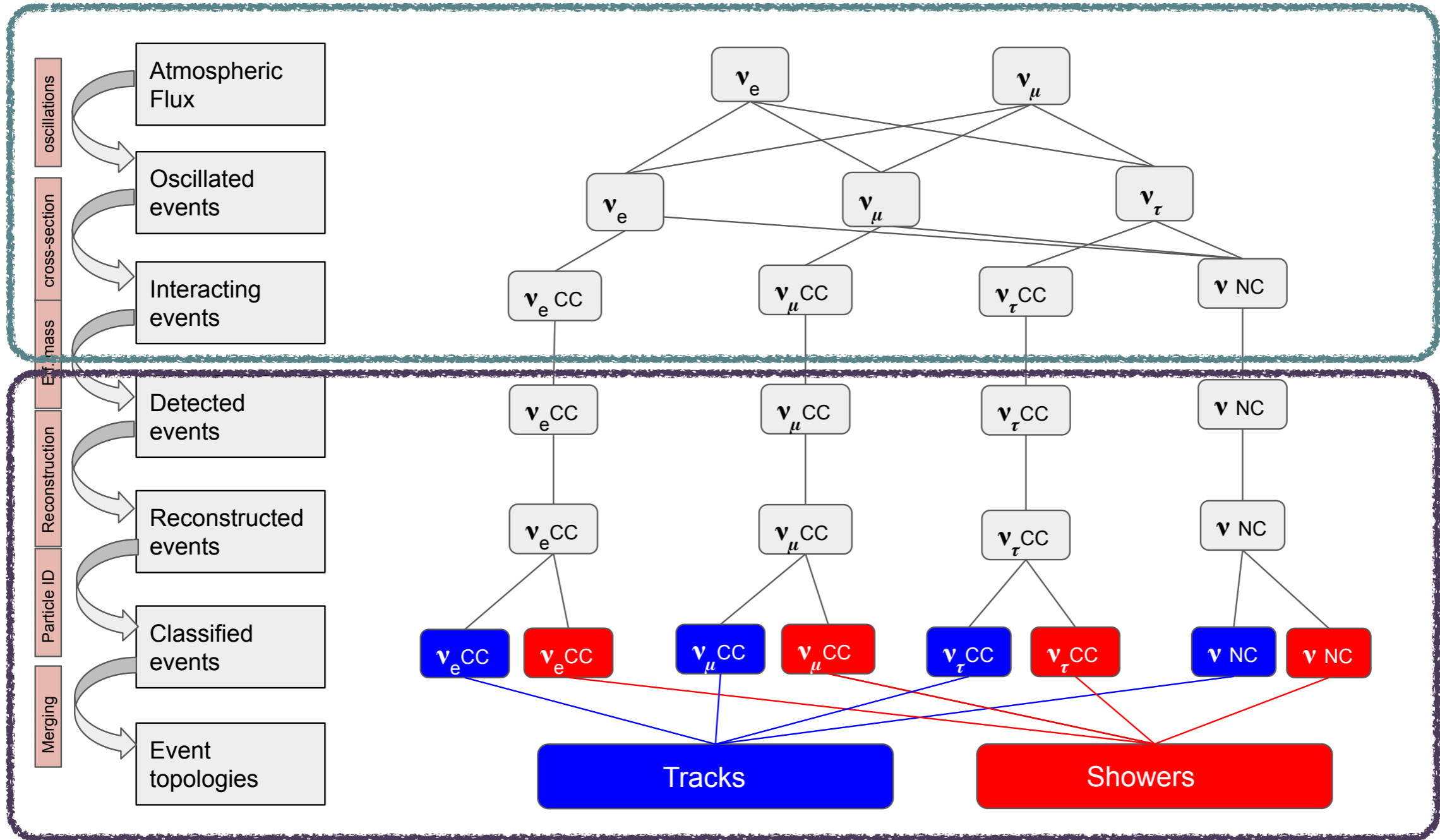






Backup

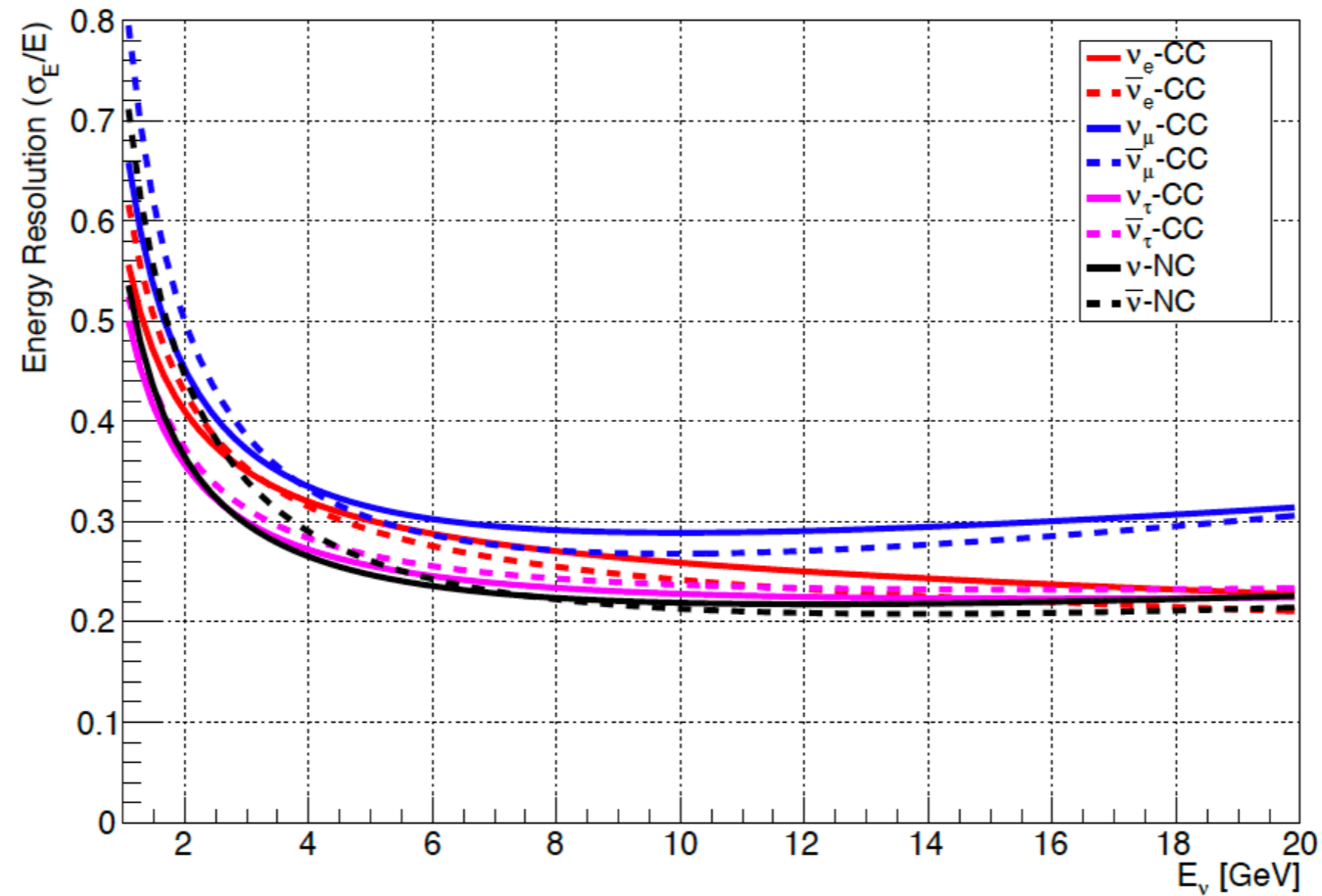
# Computation of event numbers



Model specific  
Detector specific

# Energy resolution

- Energy resolution for different interaction channels.

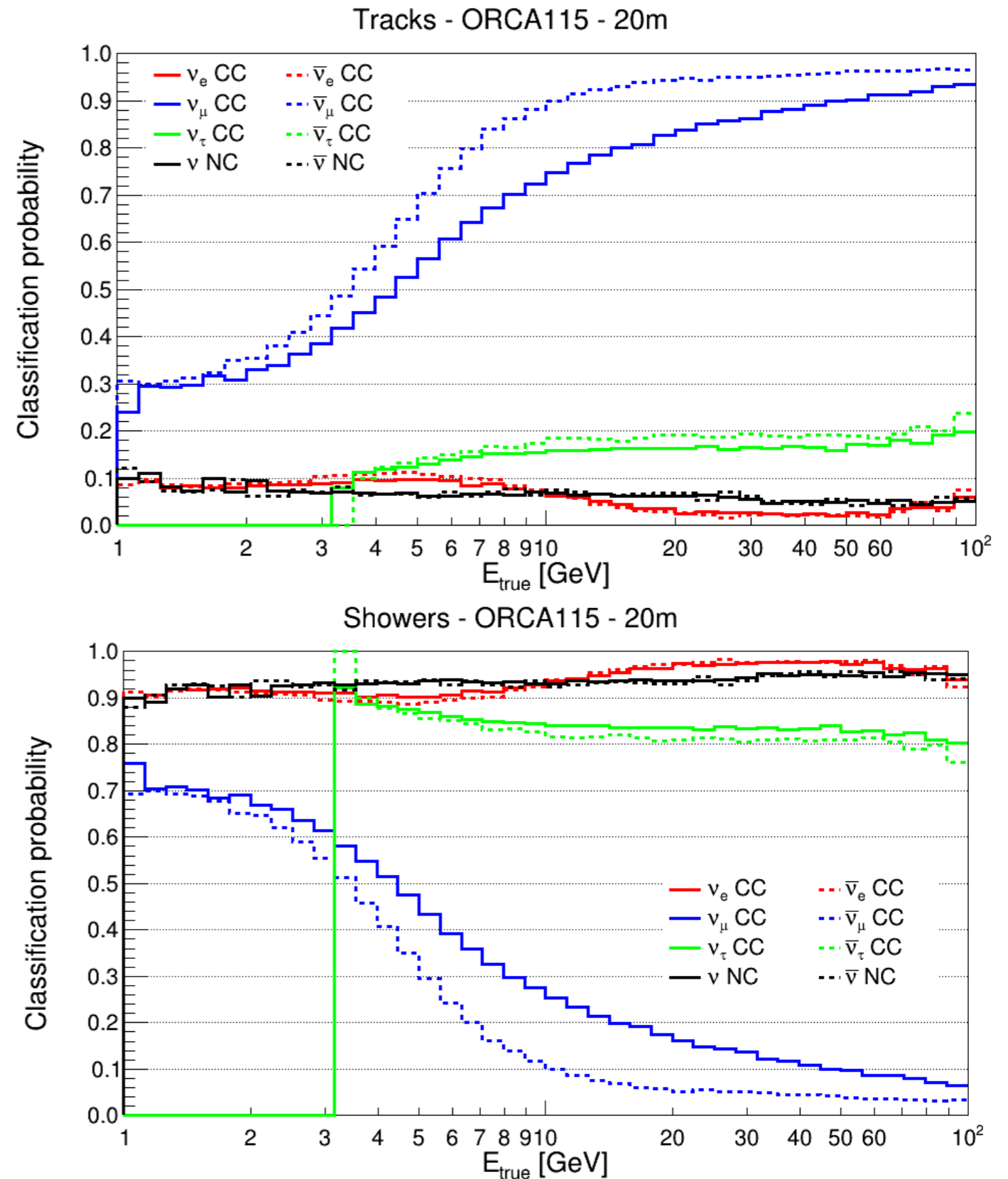


# Particle IDentification (PID)



Event classes based on track probability estimated with Random Decision Forest (RDF) of reconstructed heuristics.

- 1 = Perfect Track
- 0 = Perfect Shower



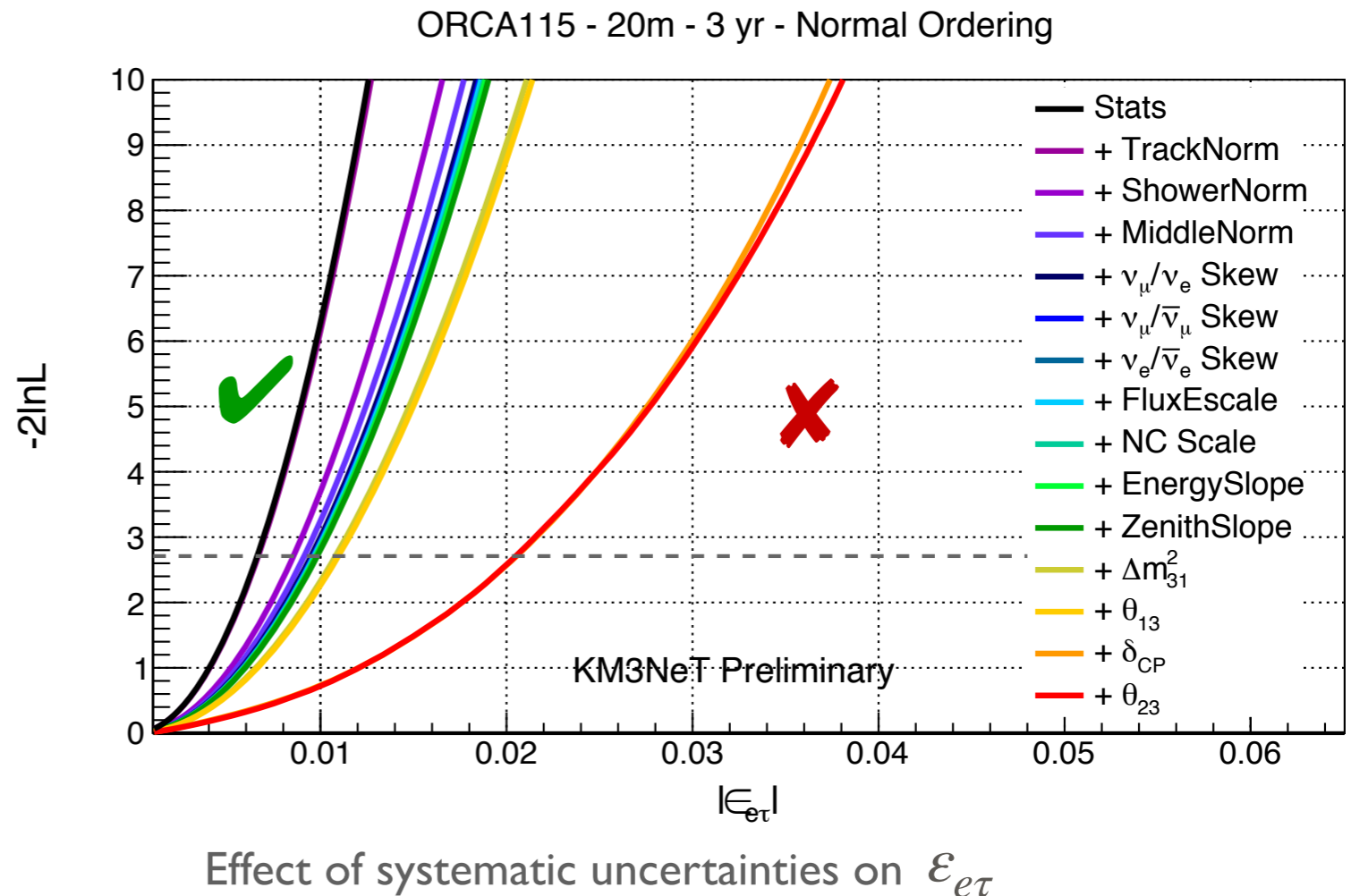
# Systematics:

$$-2\log\mathcal{L}_{NSI} = 2 \sum_{i \in \{bins\}} \left[ N_i^{NSI}(\bar{o}, \bar{s}) - N_i^{SM}(\bar{o}, \bar{s}) + N_i^{SM} \cdot \log \frac{N_i^{SM}(\bar{o}, \bar{s})}{N_i^{NSI}(\bar{o}, \bar{s})} \right] + \sum_{k \in \{syst\}} \frac{(s_k - \hat{s}_k)^2}{\sigma_k^2}$$

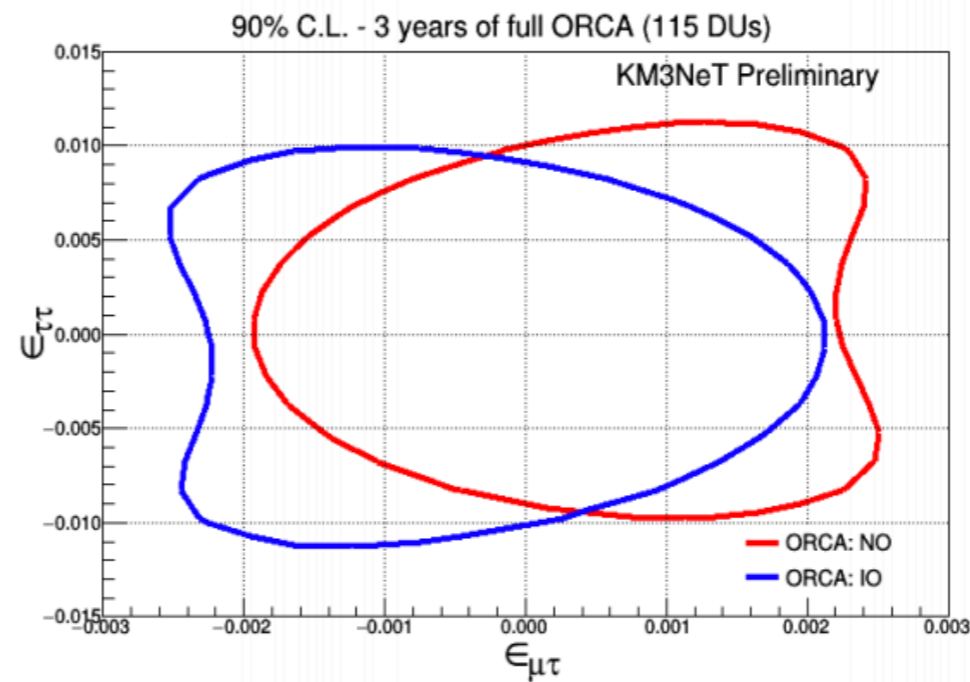
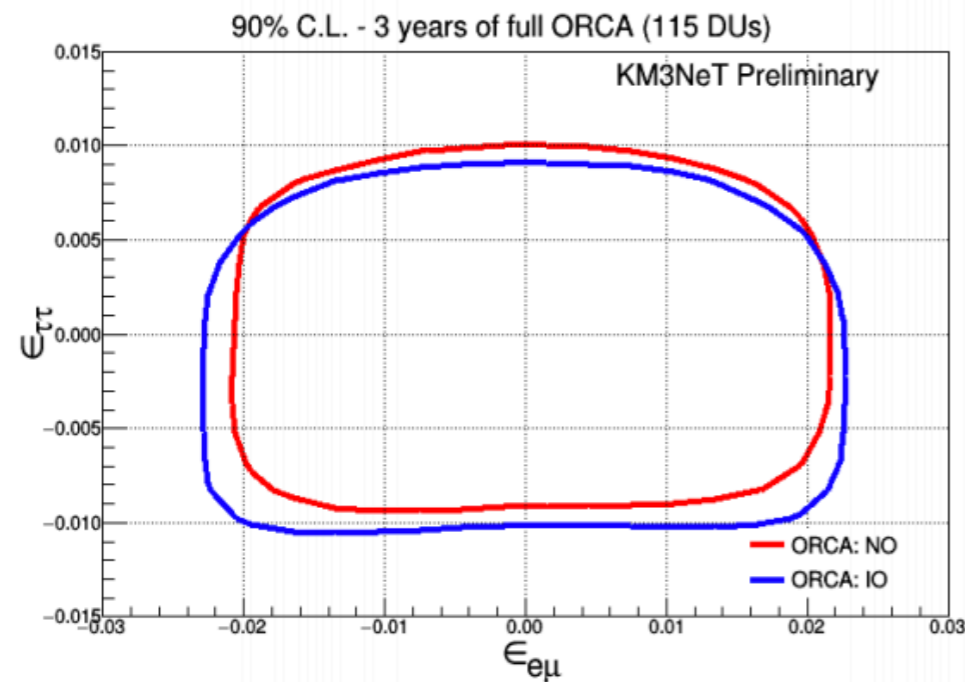
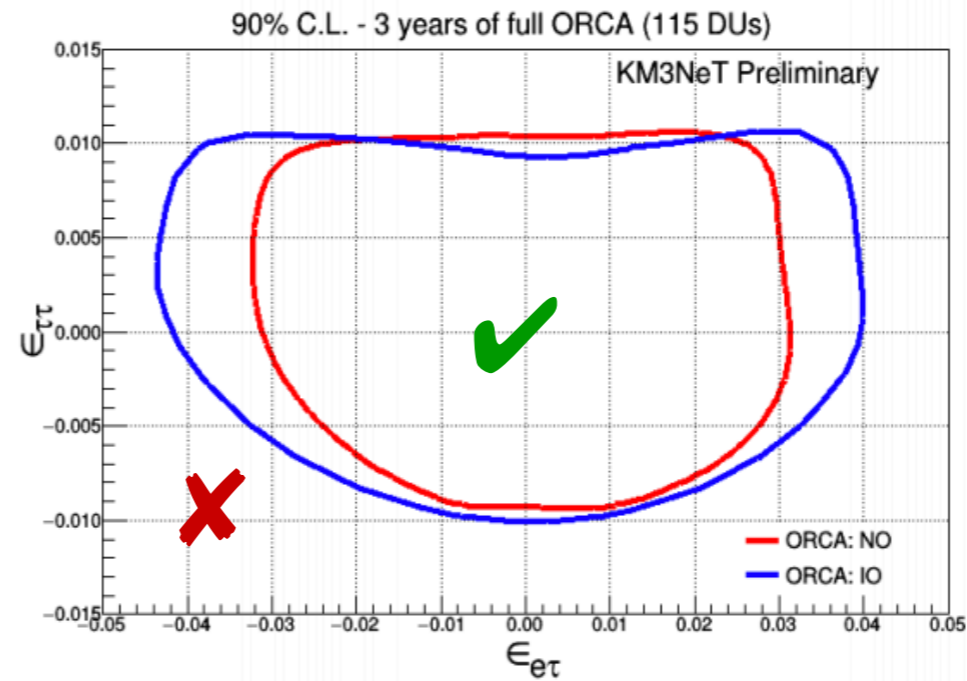
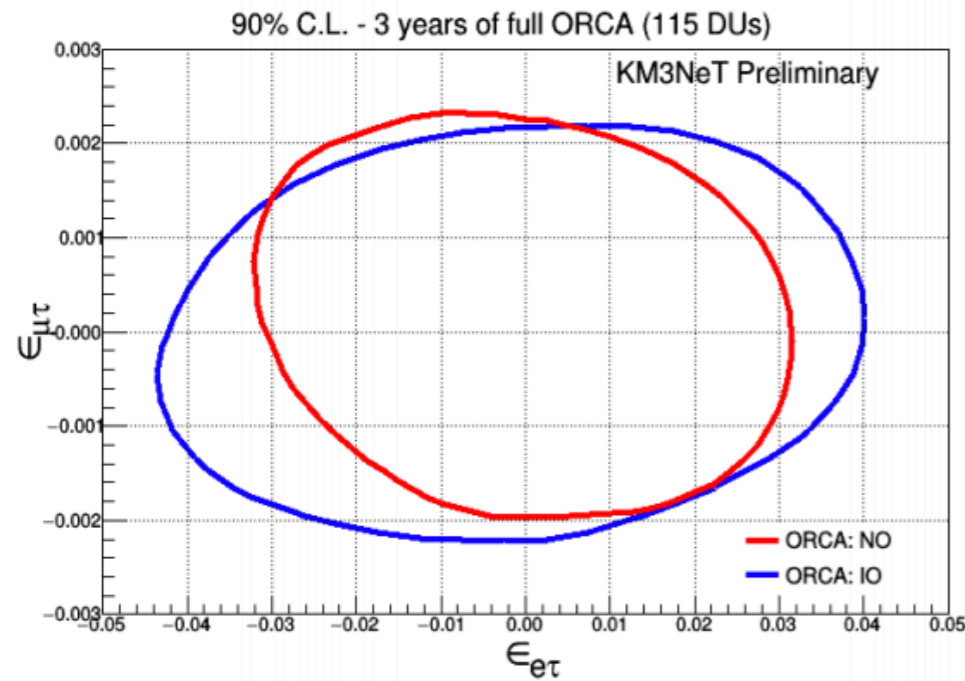
parameters	treatment	true values	priors
Track norm.	fitted	1	free
Shower norm.	fitted	1	free
Middle norm.	fitted	1	free
$\nu_\mu/\nu_e$ skew	fitted	0	5%
$\nu_\mu/\bar{\nu}_\mu$ skew	fitted	0	5%
$\nu_e/\bar{\nu}_e$ skew	fitted	0	5%
Flux Escale	fitted	1	10%
NC scale	fitted	1	5%
Energy slope	fitted	0	5%
Zenith slope	fitted	0	2%
$\delta_{CP}(\circ)$	fitted	221	free
$\Delta m_{31}^2/10^{-3}eV^2$	fitted	2.528	free
$\theta_{13}(\circ)$	fitted	8.60	0.13
$\theta_{23}(\circ)$	fitted	48.6	free

List of systematics

✓: Allowed  
✗: Excluded



# Correlated NSI Sensitivities



90% CL contours in different NSI phase spaces.

The region inside the closed curves are the region allowed by ORCA after 3 years of data taking.

<https://pos.sissa.it/358/931/>