



XIX International Workshop on Neutrino Telescopes

Neutrino Masses and Mixing

- Mixing, squared mass differences and CPV
- $0\nu\beta\beta$ searches
- Direct (kinematic) measurement of m_β

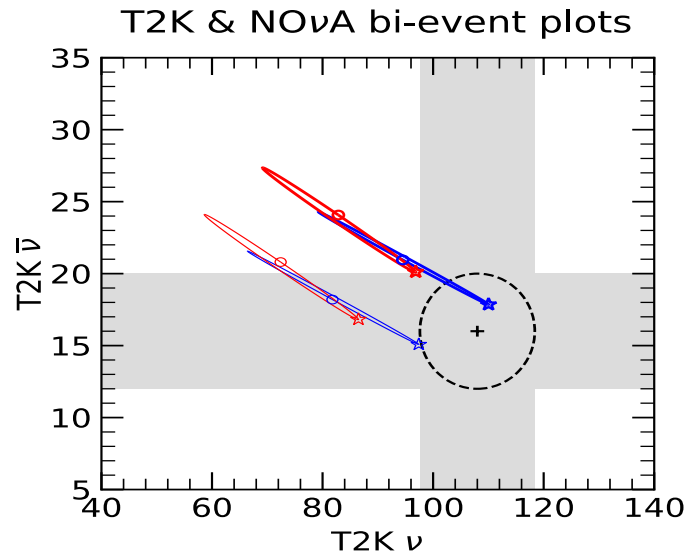
Stefano Ragazzi
Neutel 2021

Mixing, squared mass differences and CPV

A superb update by Antonio Marrone at this Workshop

- Let me highlight ***tensions*** between results of T2K and NO ν A and their consequences: weaker hints on δ **and on NO**

Integrated info on ν and $\bar{\nu}$, stat errors only
(but analysis uses spectral data)



$$S_{23}^2 = \begin{matrix} 0.57 \\ 0.45 \end{matrix} \quad \begin{matrix} \overline{\text{NO}} \\ \overline{\text{IO}} \end{matrix} \quad \delta = \begin{matrix} \pi \\ 3\pi/2 \end{matrix} \quad \begin{matrix} \circ \\ \star \end{matrix}$$

T2K alone prefers:

NO

$\delta \sim 3\pi/2$

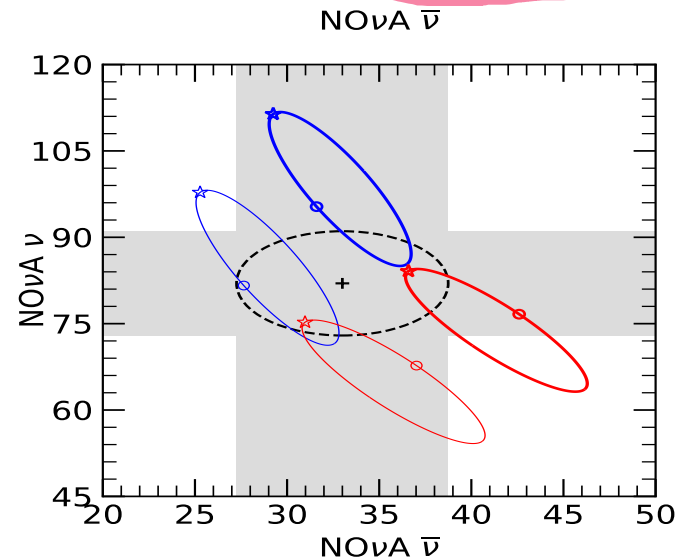
2nd octant of θ_{23}

NOvA alone prefers:

NO

CP conservation

(\sim octant degenerate)



Integrated info on ν and $\bar{\nu}$, stat errors only
(but analysis uses spectral data)

T2K & NOvA bi-event plots

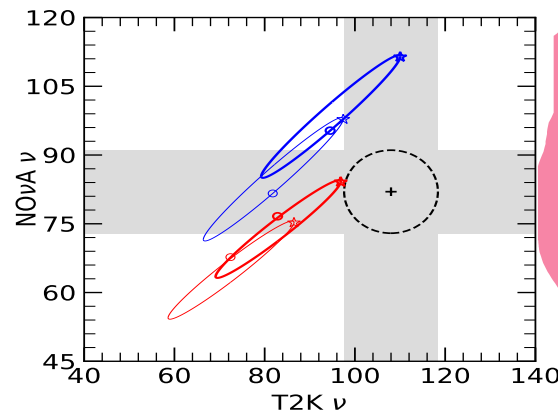
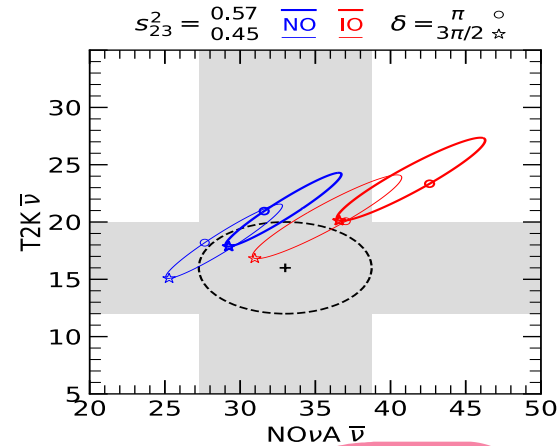
T2K + NOvA ($\bar{\nu}$)

prefer:

IO

$\delta \sim 3\pi/2$

1st octant of θ_{23}



T2K + NOvA (ν)

prefer:

IO

$\delta \sim 3\pi/2$

2nd octant of θ_{23}

→ T2K and NOvA alone:
NO preferred

→ T2K and NOvA combined:
IO preferred

→ In IO:
CP violation preferred

Mixing and CPV

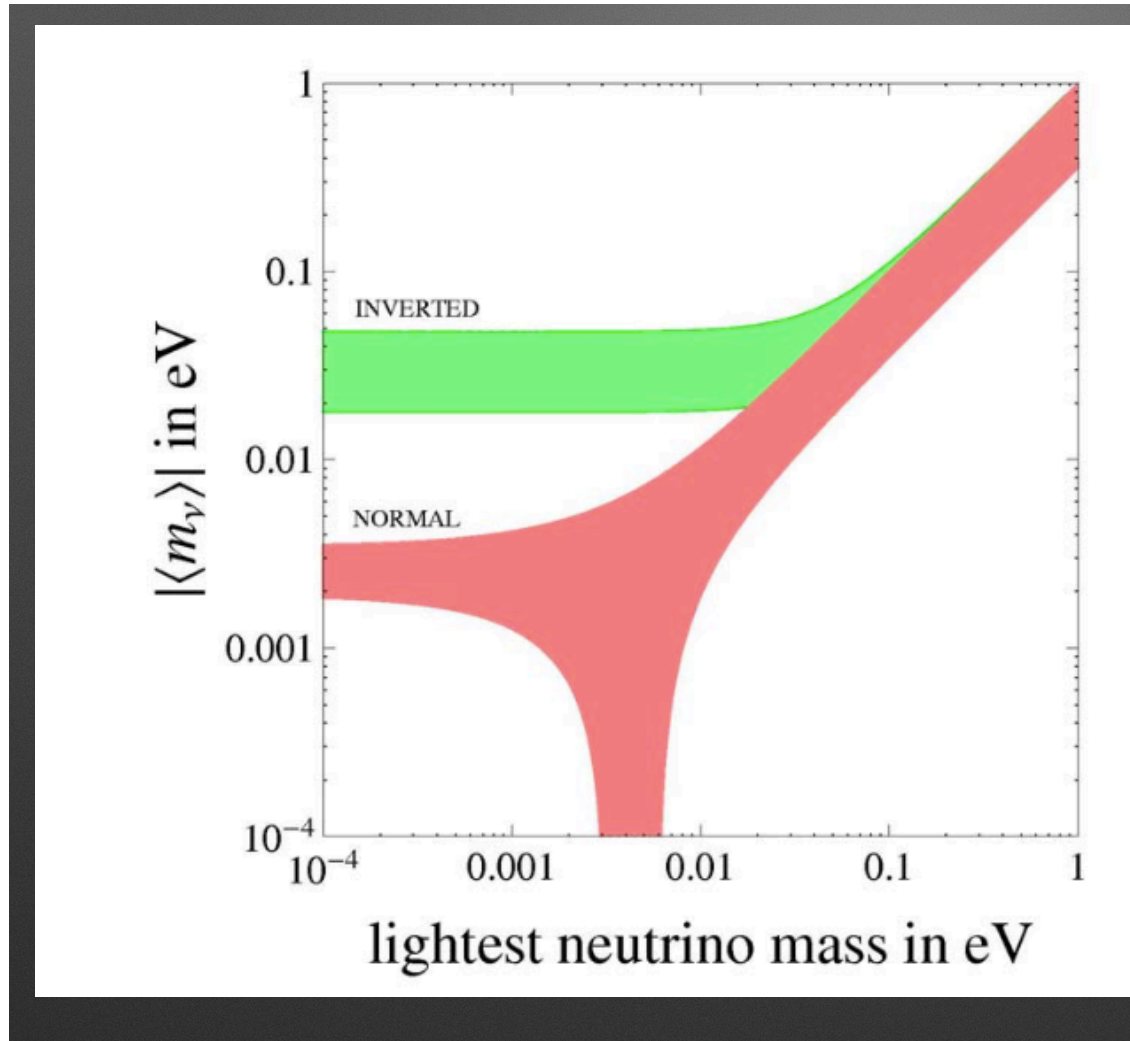
what next

- Two *privileged* players in the future
 - T2K + beam upgrade + HK
 - DUNE
- Wide, long-lasting physics program for both of them
- *Privileged* = fully supported by reference laboratories

$0\nu\beta\beta$ searches

The evolution of the universe of $0\nu\beta\beta$ searches

- Please look at the full presentation by Nando Ferroni in the opening session: I'm going to show 3 slides out of 57



realistic predictions
say:

20 meV for IH
2 meV for NH

now you can design
your new generation
experiment !

The lifetime sensitivity to chase

- For **next to next** generation experiments
- $\tau_{1/2} \sim 10^{29} \text{y}$
- $\sim 100 \text{ t}$ of isotope
- Forget about isotopic enrichment (this technology is not very popular... thus do not expect that it will be cheaper)

The menu

The problem at 10^{29} is rather the signal than the background !!!!

Probing Majorana neutrinos in the regime of the normal mass hierarchy

Steven D. Biller

Department of Physics, University of Oxford, Oxford OX1 3RH, UK

Isotope	Q (MeV)	percent natural abund.	element cost [5] (\$/kg)	$G^{0\nu}$ ($10^{-14}/\text{yr}$) [6]	$M^{0\nu}$ (avg) [7]	$T_{1/2}^{0\nu}$ for 2.5meV (10^{29}yrs)	tons of isotope for 1 ev/yr	equivalent natural tons	annual world production [5] (tons/yr)	natural elem. cost (\$M)	enriched at \$20/g (\$M)	$0\nu/2\nu$ rate [2][8] (10^{-8})
^{48}Ca	4.27	0.19	0.16	6.06	1.6	2.70	31.1	16380	2.4×10^8	2.6	622	0.016
^{76}Ge	2.04	7.8	1650	0.57	4.8	3.18	58.2	746	118	1221	1164	0.55
^{82}Se	3.00	9.2	174	2.48	4.0	1.05	20.8	225	2000	39	416	0.092
^{96}Zr	3.35	2.8	36	5.02	3.0	0.93	21.4	763	1.4×10^6	27	427	0.025
^{100}Mo	3.04	9.6	35	3.89	4.6	0.51	12.2	127	2.5×10^5	4.4	244	0.014
^{110}Pd	2.00	11.8	23000	1.18	6.0	0.98	26.0	221	207	5078	521	0.16
^{116}Cd	2.81	7.6	2.8	4.08	3.6	0.79	22.1	290	2.2×10^4	0.81	441	0.035
^{124}Sn	2.29	5.6	30	2.21	3.7	1.38	41.2	736	2.5×10^5	22	825	0.072
^{130}Te	2.53	34.5	360	3.47	4.0	0.75	23.6	68	~ 150	24	471	0.92
^{136}Xe	2.46	8.9	1000	3.56	2.9	1.40	45.7	513	50	513	914	1.51
^{150}Nd	3.37	5.6	42	15.4	2.7	0.37	13.4	240	$\sim 10^4$	11	269	0.024

The only choice that does not call for an impossible cost for the enrichment points to natTe

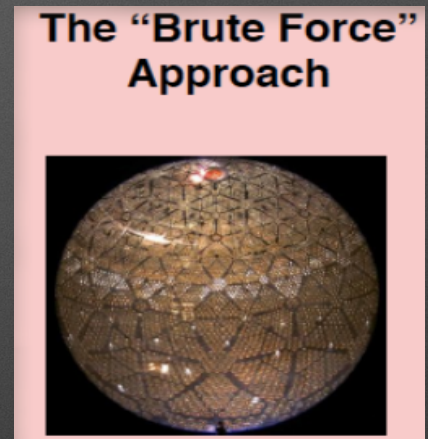
- N.B. forget about the second in the shortlist: ^{100}Mo has a horrible 2ν rate

The (poisonous) Tellurium soup

Te might strikes back

Dissolve a huge quantity of natural Te (few hundred tons)
at the highest concentration allowed by the transmission
of the light in a scintillator

(Juno -20000 tons)
(SuperK -50000tons)



Two backgrounds are serious: $2\nu\beta\beta$ and 8B from the Sun

The neutrinos from the Sun might be tagged if some
directionality could be implemented (Cherenkov !)

No life in the far future for the “peak squeezing” approach?

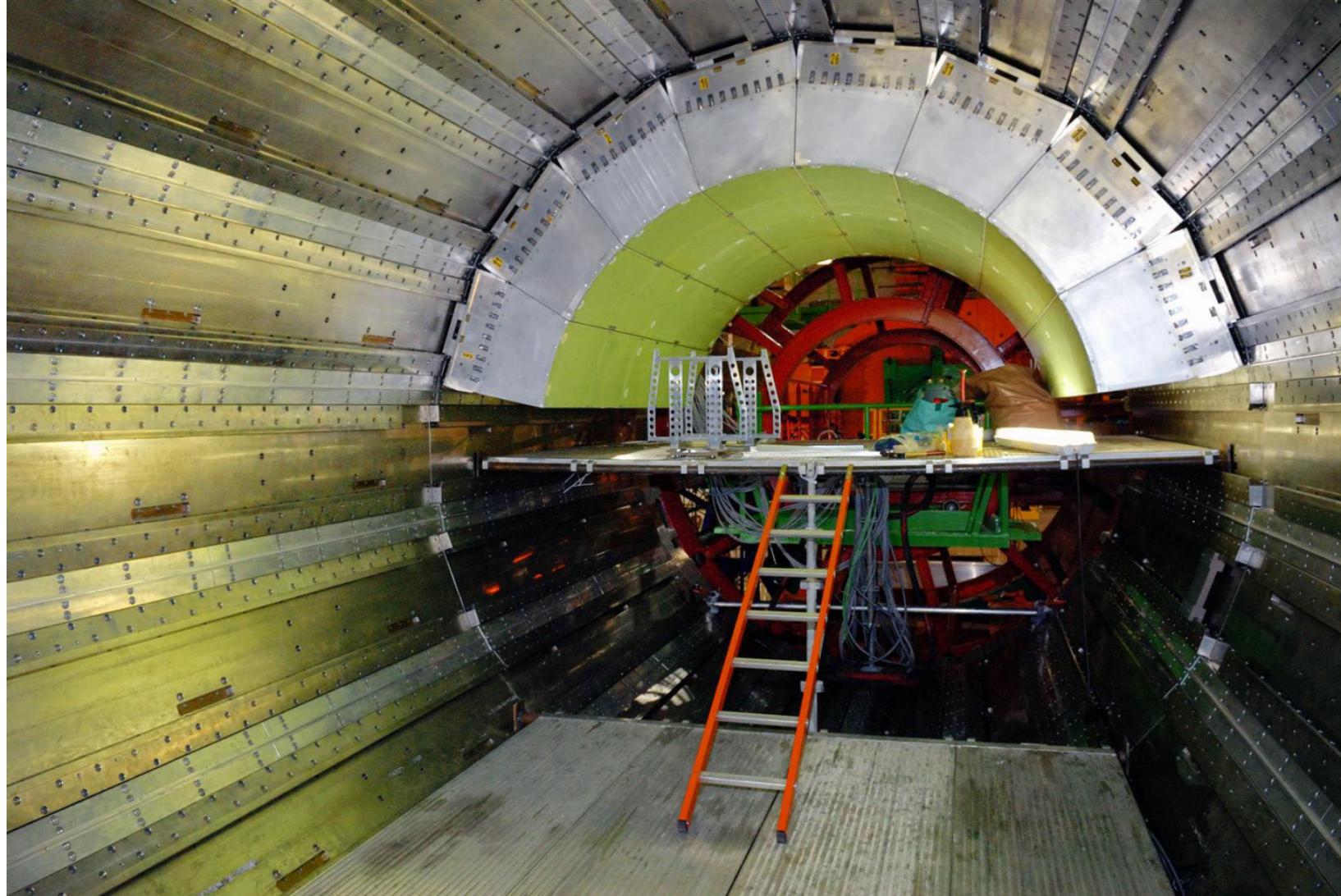
- Keep natural Te \Rightarrow 100 x CUORE

1. \sim 100k Crystals calorimeter
2. At 10 mK
3. 1/100 bkg

- **Impossible or challenging?**

100k Crystal calorimeter

- It is run at 291 K, quite higher than 10 mK



100 x CUORE Cryostat

- CUORE cryostat has been a very challenging project
 - btw: congratulations to CUORE for the impressive duty cycle attained, 2 years ago I could not believe that the cryostat would be useful for CUPID
- Commercial mK cryogenics is ramping up
- What the impact of QC based on superconducting qbits is going to be on cryogenics?
- *Keep calm and carry on*

Background reduction

Nuclear Instruments and Methods in Physics Research 224 (1984) 83–88
North-Holland, Amsterdam

83

- At the beginning it had to be an extreme peak-squeezing approach
- What went wrong?
- Where is energy escaping or trapped?
- If we seek extreme performance we should better understand our detectors
- Keep in mind the long path of bkg reduction of Ge based experiments

LOW-TEMPERATURE CALORIMETRY FOR RARE DECAYS

E. FIORINI

Dipartimento di Fisica dell'Università and INFN, Milano, Italy

T.O. NIINIKOSKI

CERN, Geneva, Switzerland

Received 27 December 1983

The recent developments in underground low-counting experiments give limits to rare decays which are hard to improve since scaling the size and the resolution of the combined source–detector is difficult with the existing techniques. We explore here the possibility of low-temperature calorimetry to improve the limits on processes such as neutrinoless double-beta decay and electron decay.

N.B. Θ_D for Te reported by Fiorini-Niinikoski is $\sim 2x$ lower than Θ_D of TeO_2 crystals

...and things may be different

- Just add one sterile neutrino
- The end of the game may be far away even for IO

Sebastian Böser¹, Christian Buck², Carlo Giunti³, Julien Lesgourgues⁴, Livia Ludhova^{5,6}, Susanne Mertens^{7,8}, Anne Schukraft⁹, Michael Wurm^{1,*}

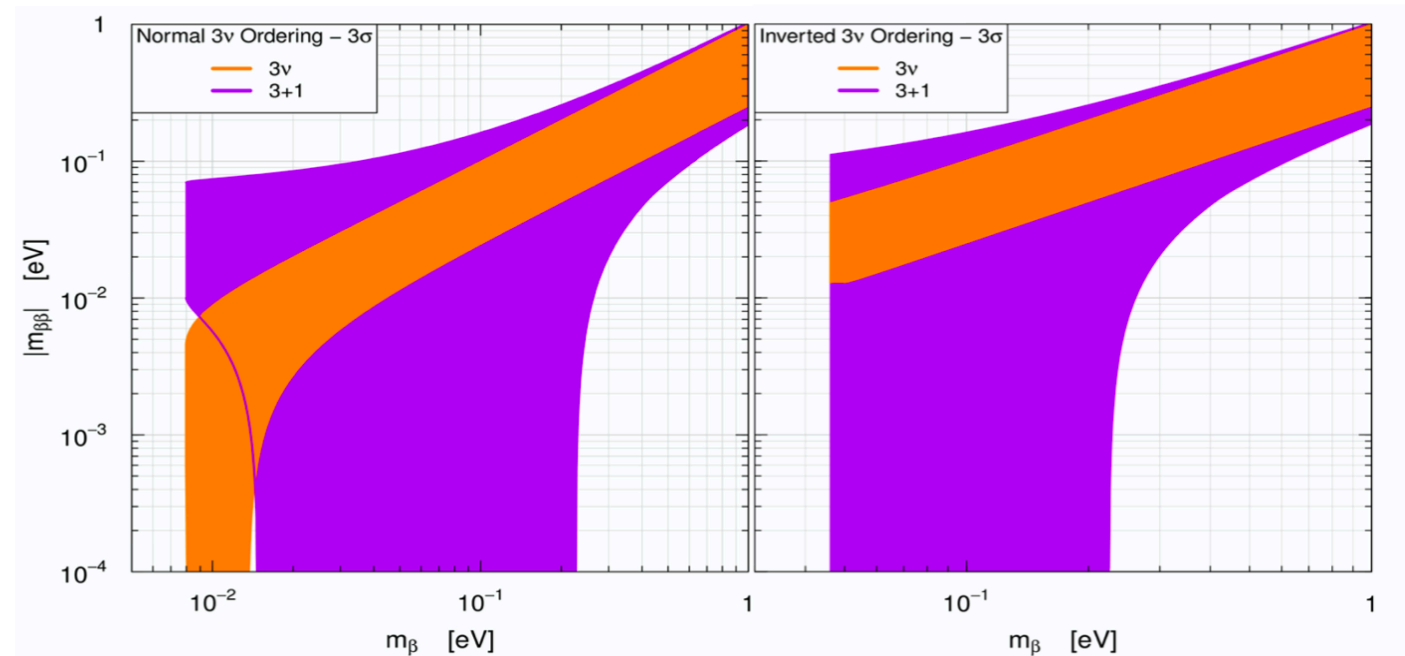


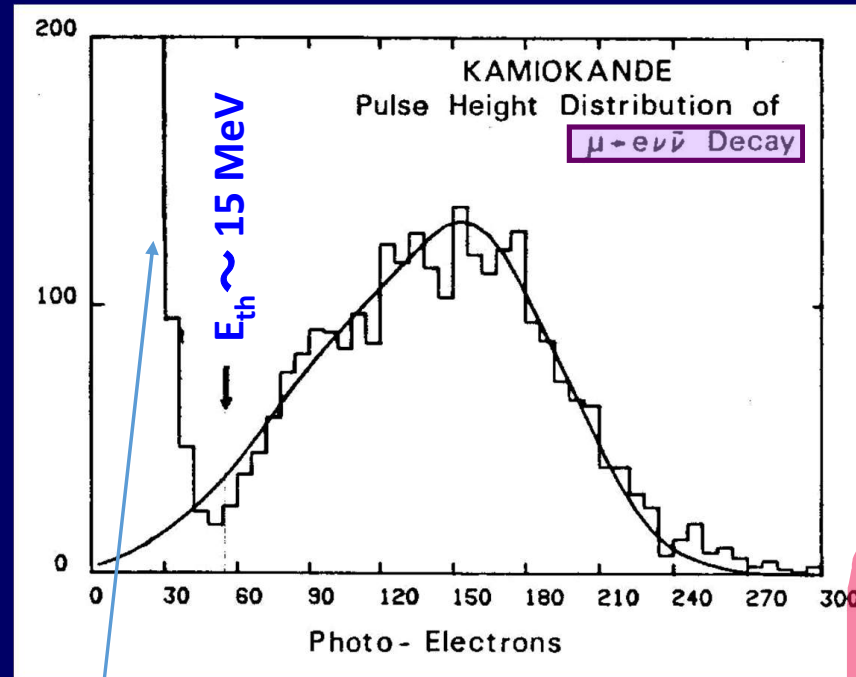
Figure 7: Predicted ranges for the Majorana neutrino mass $|m_{\beta\beta}|$ as a function of the effective electron neutrino mass m_{β} : While left and right panels distinguish normal and inverted neutrino mass ordering, the different-color bands illustrate the different allowed ranges in case of 3-flavor and (3+1) scenarios [69].

A dangerous killer

- From Atsuto Suzuki retrospective on Dr. Masatoshi Koshihara

What invited Kamiokande-II ?

Only **several months** later after data-taking



Kamiokande-II :

huge background mountain stands in front of us

Koshihara :

Why not lower E_{th} down to 10 MeV to detect ^8B solar neutrinos



Proposal : not Improvisation, but his Deep Consideration

Koshihara :

Even before, the start of the experiment, I had been thinking that the Kamiokande should produce significant scientific results, even if proton decays were not observed.

On the origin of the Kamiokande experiment and neutrino astrophysics

T. Kajita^{1,a}, M. Koshihara², and A. Suzuki³

Eur. Phys. J. H
DOI: 10.1140/epjh/e2012-30007-y

THE EUROPEAN
PHYSICAL JOURNAL H

Direct (kinematic) measurement of m_β

A single BIG player today: KATRIN



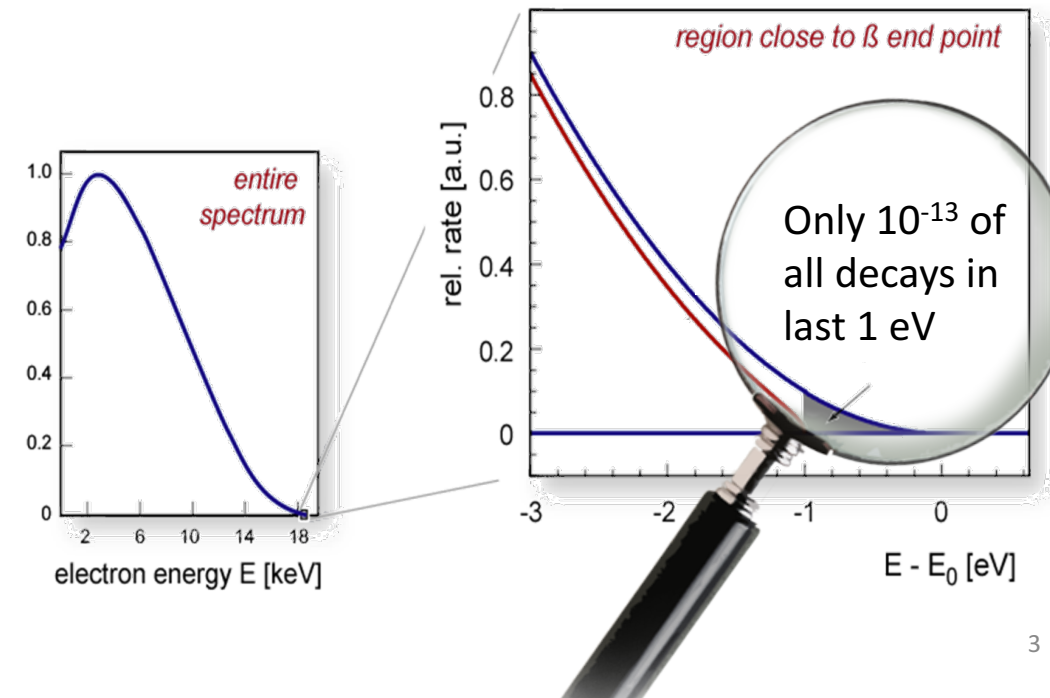
Karlsruhe
Tritium
Neutrino
Experiment

Katrin

The challenge

- Slide by Susanne Mertens @Neutrino 2020

- Ultra-strong β -source: 10^{11} decays/s
- Low background level < 0.1 cps
- Excellent energy resolution ~ 1 eV
- Precise understanding of spectrum



Susanne Mertens (MPP, TUM)

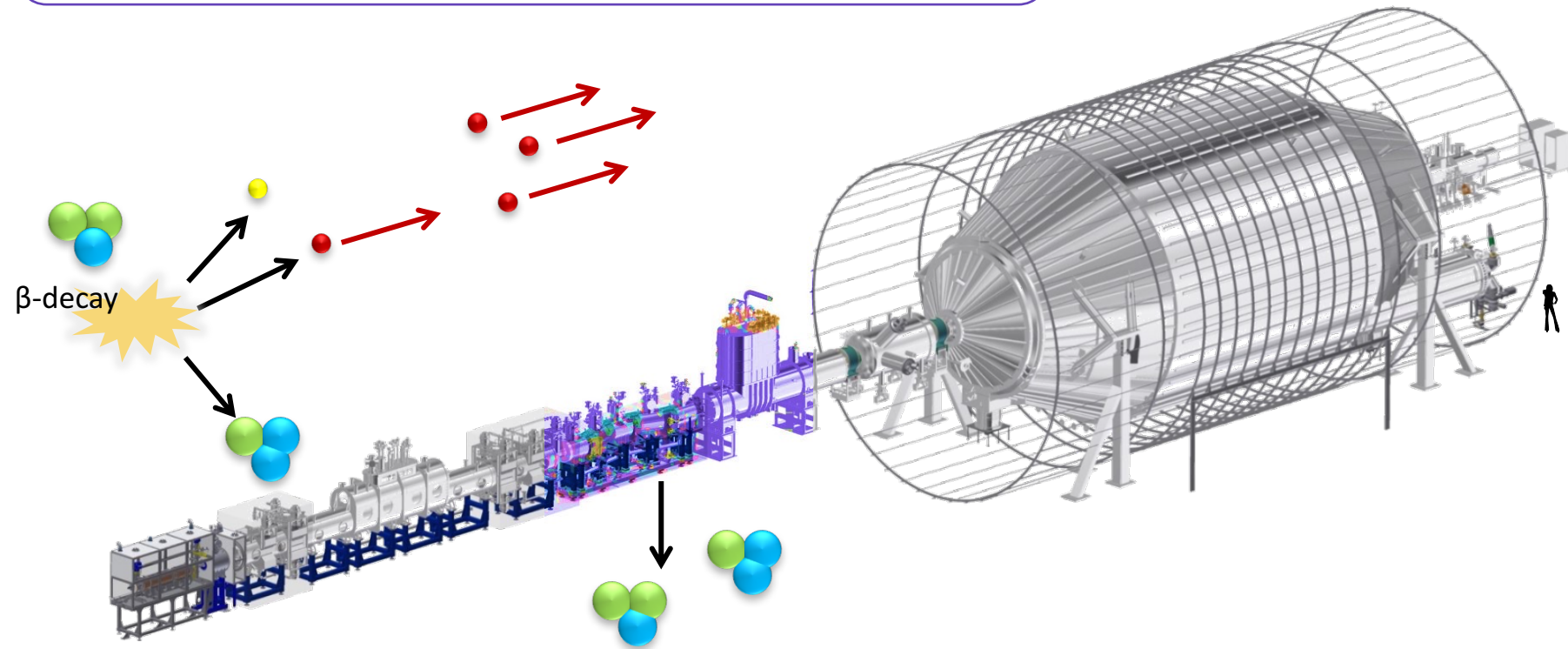
Katrin

- Slide by Susanne Mertens @Neutrino 2020

KATRIN Working Principle

Transport section

- magnetic guidance of electrons (@ 4 T)
- tritium flow reduction by $> 10^{14}$ + tritium ion removal



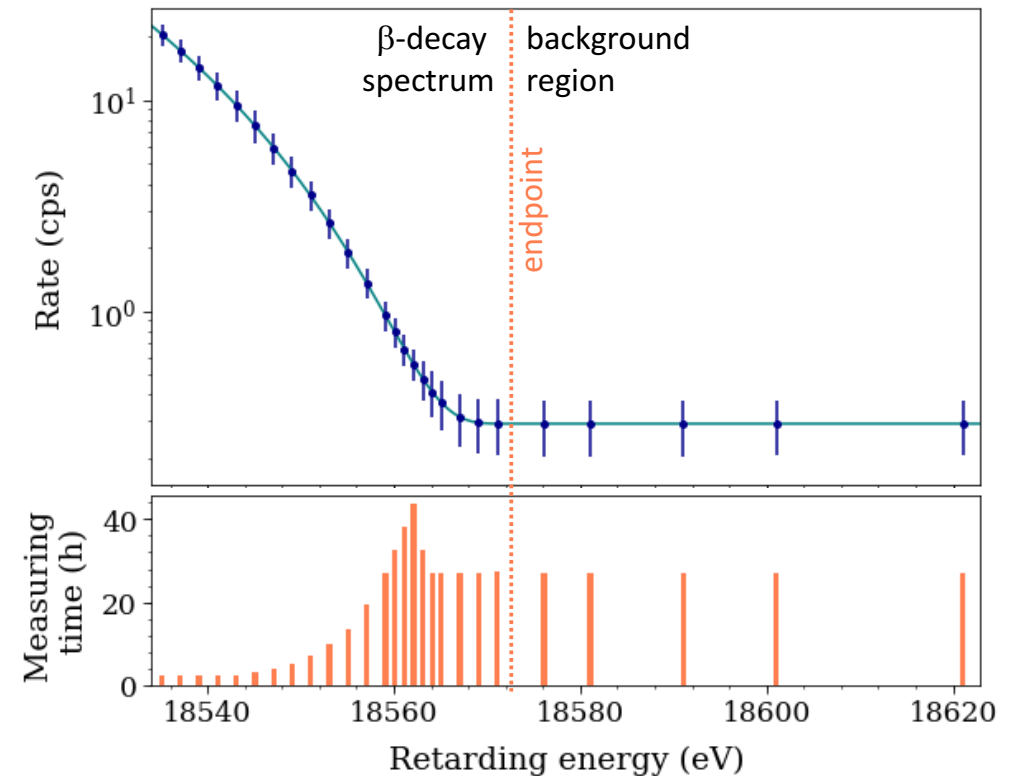
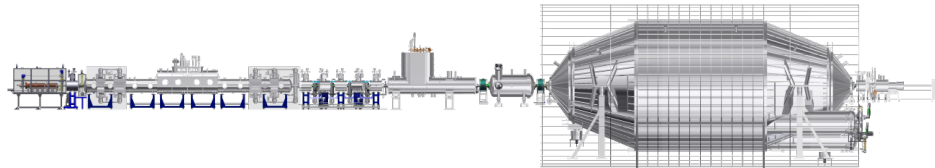
Susanne Mertens

Katrin

- Slide by
Susanne
Mertens
@Neutrino
2020

Measurement strategy

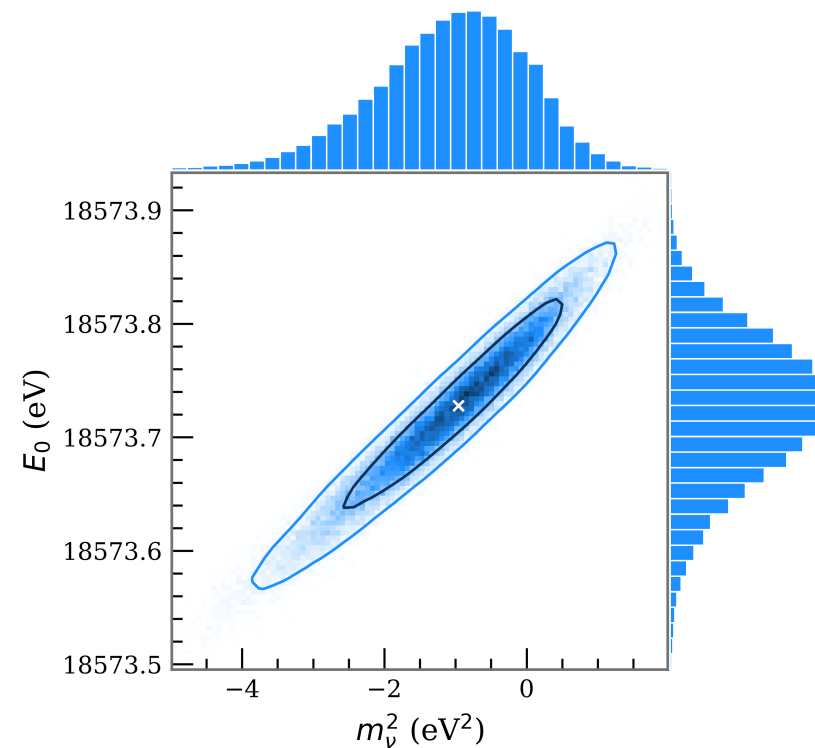
- # HV set points: **27**
- interval: **$E_0 - 40 \text{ eV}$, $E_0 + 50 \text{ eV}$**
- scanning time: **2 hours**
- # scans: **274**
- HV stability: **20 mV (ppm-level)**



Katrin

- Slide by
Susanne
Mertens
@Neutrino
2020

Final fit result



Best fit results:

$$m_\nu^2 = (-1.0^{+0.9}_{-1.1}) \text{ eV}^2$$

→ compatible with zero

→ probability of 16%, if true $m_\nu = 0 \text{ eV}$

$$E_0 = 18573.7 \pm 0.1 \text{ eV}$$

→ Q-value : $18575.2 \pm 0.5 \text{ eV}$

→ good agreement with literature ($Q = 18575.72 \pm 0.07 \text{ eV}$)

E. Myers et al. Phys. Rev. Lett. 114, 013003 (2015)

Katrin – what next

- Search for signatures of m_4 mixing: distortions in beta spectrum
 1. eV scale sterile neutrino -> same data as for standard m_β
 2. keV scale neutrino mass: needs segmented final detector -> TRISTAN
- Further upgrades? Katrin deserves to be fully exploited

m_β from electron capture

Electron capture calorimetric experiments



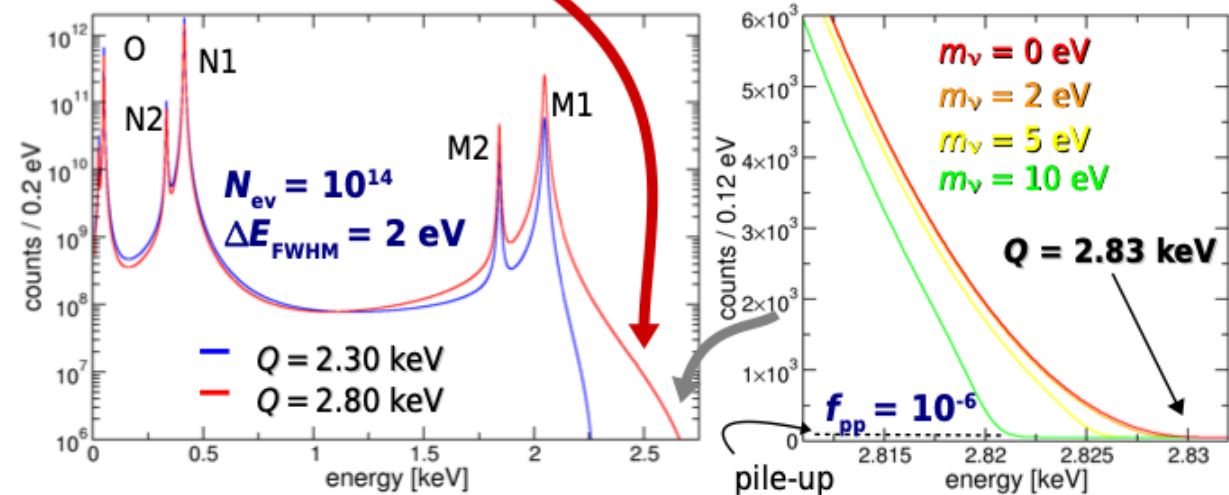
electron capture from shell $\geq M1$

A. De Rújula and M. Lusignoli, Phys. Lett. B 118 (1982) 429

- ECHO and HOLMES
- **Milestone:** demonstrate 1 eV capability
- No general consensus on the costs for a sub-eV detector
 - It depends on number of channels, which depends on maximum tolerable rate per channel
- Be optimistic on major costs reduction from commercial technologies

- **calorimetric measurement of Dy atomic de-excitations** (mostly non-radiative)
- **$Q = 2.83$ keV** (determined with Penning trap in 2015)
 - end-point rate and **ν mass sensitivity** depend on **$Q - E_{M1}$**
- **$\tau_{1/2} \approx 4570$ years** $\rightarrow 2 \times 10^{11}$ ^{163}Ho nuclei \leftrightarrow 1Bq

$$N(E_c) = \frac{G_\beta^2}{4\pi^2} (Q - E_c) \sqrt{(Q - E_c)^2 - m_\nu^2} \times \sum_i n_i C_i \beta_i^2 B_i \frac{\Gamma_i}{2\pi} \frac{1}{(E_c - E_i)^2 + \Gamma_i^2/4}$$



A. Nucciotti, Direct Neutrino Mass Measurements, Università di Roma "La Sapienza", July 8th, 2020

Slide by A. Nucciotti

m_β novel concepts

- Project 8
 - Novel approach to e^- spectrometry: measure frequency, not E
 - In progress
- PTOLEMY
 - Very ambitious: initially proposed as a *Telescope for Relic Neutrinos* – m_β as a by-product
 - Novel concepts for
 - Source: ^3H embedded in a graphene layer, $\sim \times 10^3$ Katrin source intensity
 - Spectrometer: differential high-res spectrometry
 - Detector: low-energy high-resolution detector
 - Every item has to be developed and demonstrated
 - Please refrain from calling it a Neutrino Telescope at least until LIX Neutel
 - Remarkable anyway

Final remarks

- Mixing, phases, squared mass difference experiments are in a *privileged* position
 - Supported by reference laboratories
 - Making measurements
 - Wide physics menu
- Majorana $m_{\beta\beta}$ searches
 - Healthy competition and variety of approaches - some downselection will occur
 - You pay for $(1/m_{\beta\beta})^2$ at least
 - “Zero background” approach has to go ahead if we don’t want to pay for $(1/m_{\beta\beta})^4$ or hit a systematics wall
 - Peak squeezing technique: it is mandatory to take advantage of the time of CUPID construction and run for a better understanding of detectors
 - Beware of single item physics menu! It can be more poisonous than a tellurium soup

Final remarks (continued)

- m_β measurement
 - Measure $(m_\beta)^2$ pay for $1/(m_\beta)^2$
- The shores of neutrino mass spectroscopy are at 50 meV or nearer: a tiny effect for NO a major shoulder for IO
- A single big operating experiment: Katrin
 - It was a tremendous effort: congratulations to Katrin
 - Fully exploit Katrin + Upgrades (Tristan ...)
- The shores of neutrino mass spectroscopy are at 50 meV or nearer: a tiny effect for NO a major shoulder for IO

Final remarks - m_β measurement (continued)

- EC (Holmium) experiments
 - Costs would be unsustainable **today** for a sub-eV experiment, but they will tremendously benefit of development of commercial technologies
 - There still is a long way to go along the path followed by Mac-E Spectrometers: important milestone at 1 eV
- New spectrometers
 - Project-8 - measure frequency not energy
 - Ptolemy - very ambitious final goal: detect relic neutrinos (m_β as a by-product): novel concept for source, spectrometer, detector
 - It needs a supporting laboratory (at least)

Thanks

- Many thanks to NeuTel organizers for inviting me
- Many thanks to O. Cremonesi, A. Nucciotti, S. Pirro, F. Terranova for enlightening discussion