Direct Neutrino Mass Measurements Coherent Neutrino Scattering

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Neutrino Telescopes 2021

## A bit of context behind the two topics...

We would like to invite you to give one of these Challenge talks. The argument(s) would be "New technologies for measuring neutrino masses" together with "Neutrino coherent scattering".

-- Mauro Mezzetto

Yes, I am happy to partake in the Challenge talks! One question of clarification: are you asking me to give a talk on **both** of these topics [...]? -- Me

Yes – Mauro

Ok <gulp> — Me

Luckily, there are some similarities...

#### ... for one, these challenges are long standing

(1974)



Mass (1934)

#### Both measurements access key neutrino properties...

#### v Mass Scale

$$m_{\beta}^2 \equiv \sum_{i=1}^3 |U_{ei}|^2 m_i^2$$

Incoherent sum of neutrino mass eigenstates.

Relates directly to the energy-momentum dispersion of neutrinos.



Probes the weak coupling of neutrinos to nuclei.

Coherent interaction with entire nucleus.

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#### and in principle, the measurements are simple.



Measure the kinetic energy spectrum of the decay electron (for beta decay), or de-excitation energy (for electron capture).



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Measure the kinetic energy of the outgoing recoil.

Cross-section proportional to (roughly) the square of the number of neutrons.





and also connect to other branches of physics.



Neutrino mass connects back to cosmological observations.

3  $m_i$ 

PLANCK & Sky Map

Limits on the neutrino mass scale from cosmology are affected by model extensions (such as additional parameters) or new physics (neutrino self-interaction or dark matter interactions).



and also connect to other

#### branches of physics.

The coherent process likewise provides allows insight into new physics at small momentum exchange.





Also gives insight into supernova physics, nuclear structure, solar physics and much more.





#### A World-wide effort is already underway for direct

measurements.





Electron transfers all of its energy to the absorbing medium.

#### Calorimetric (Cryogenic Bolometers)

Electromagnetic filtering of electrons of selected energy.

#### Electromagnetic Collimation (MAC-E Filter)





Use photon spontaneous emission from electron in magnetic field.

### Frequency-Based

(Cyclotron Radiation Emission Spectroscopy)

### ECHo & HOLMES



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#### Project 8



Use photon spontaneous emission from electron in magnetic field.

### **Frequency-Based**

(Cyclotron Radiation Emission Spectroscopy)

#### A World-wide effort is also underway for coherent

neutrino scattering, too.





#### Ionization Detectors (Germanium & Si-CCD)

Large mass, keV scale

#### CONUS, CONNIE, TEXONO





MINER, Ricochet, NuCLEUS



### **Cryogenic Bolometers**

Smaller detectors, sub-keV scale

particle discrimination via ionization, photon tagging We have made great strides in last few years...





2 million events, 780 hours of data.

Excellent goodness-of-fit: p-value=0.56.

 $m^2(\nu_e) = (-1.0^{+0.9}_{-1.1}) \text{ eV}^2 (90\% \text{ C.L.})$ 

 $m_{\beta} \leq 1.1 \text{ eV} (90\% \text{ C.L.})$ 

We have made great strides in last few years...



Process finally measured at the spallation neutron source by the COHERENT collaboration.

Now measured for Cs and Ar targets.



#### but both programs have ambitious goals.

#### <u>v Mass</u>

Resolve whether the neutrino mass scale is quasi-degenerate in the next few years.

And eventually push down to the inverted ordering scale.



#### v Coherent Scattering

A first detection from reactor neutrinos in the next few years.

#### Build O(10 kg) detectors over the next decade.



Does the technology being pursued get us there?

So, both these measurements have a

common wish list...

Sufficient statistics

**Background Reduction** 

Superior Energy Resolution

## Sufficient Statistics



Only a tiny fraction of the spectrum yields information about the neutrino mass.

Large activities are required in order to probe the spectrum at the few eV level.

## Sufficient Statistics

In the case of coherent scattering, coherence helps a lot.

That said, the crosssections are still small.

Two work-arounds: operate at higher energies (COHERENT), or with large fluxes (nuclear reactors).







## Overcoming Backgrounds

#### **CENNS** Detectors **Direct Measurements** KATRIN data with 1 $\sigma$ error bars $\times$ 50 Count rate (cps) Background Model Fit result Analysis Threshold (60 eV) Number of counts [evts/keV] Data Excluded WIMP model: 0.7 GeV/ $c^3$ , 9.8 × 10<sup>-25</sup> cm<sup>2</sup> Excluded WIMP model: 2.0 GeV/ $c^2$ , 4.5 × 10<sup>-37</sup> cm<sup>2</sup> Excluded WIMP model: 10.0 GeV/ $c^2$ , 1.1 × 10<sup>-37</sup> cm<sup>2</sup> $10^{4}$ 10<sup>0</sup> ' Standard Spectra $10^{3}$ 10 20 30 40 -30 -20 -10 0 $10^2$ 0.03n Energy [keV] Rydberg atoms Background **Radon decays** "wall"

Reduction of backgrounds is crucial for experiments of both sorts, though the causes for these backgrounds depends on the specific technology deployed.

## Overcoming Backgrounds

#### **Direct Measurements CENNS** Detectors Nuclear recoil equivalent energy [keVnr] 0.15 0.2 0.25 0.3 0.35 Magnetic field Magnetic **Active Veto Electron/recoil** optimization shielding from Passive event Radon removal surfaces shielding discrimination

Different mitigation strategies are used which also depend on the technology that is deployed.





For direct measurements, energy resolution is needed to extract the spectral shape near the endpoint. Order few eV resolution now achieved by major efforts.

## Superior energy resolution



For CENNS measurements, energy resolution is needed to move the energy threshold for recoils as low as possible. Sufficient statistics

**Background Reduction** 

Superior Energy Resolution

In certain cases, experiments in both categories implement very

similar technology to achieve their goals.



e.g.: Magnetic µ-calorimters, Transition Edge Sensors, and NTDs

Sufficient statistics

**Background Reduction** 

Superior Energy Resolution

In certain cases, experiments in both categories implement very

similar technology to achieve their goals.



e.g. Multiplexing of signals to increase target mass

Sufficient statistics

**Background Reduction** 

Superior Energy Resolution

In certain cases, experiments in both categories implement very

similar technology to achieve their goals.



e.g. Operating at the quantum noise limit for microwave detection



Much of these advances are due to great advances in superconducting technology.

High quality production of JJ junctions, superconducting leads.

Driven by strong recent development for superconducting circuits for quantum computing.

However, unique challenges still remain for using these new technologies for neutrino physics:



Do these technology scale?

Can they operate in harsh environments?

What backgrounds lurk in this new space?



# But half the fun is the challenge.

"We choose to go to the Moon. We choose to go to the Moon in this decade and do the other things, not because they are easy, but because they are hard."

Thank you for your attention.

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**Backup Slides** 

(In case people ask questions)

## Modern Calorimetric Experiments



Micro calorimeters which are sensitive to changes in temperature (energy deposition). Contain the full decay energy.



High Magnetic Field (Bs) Low Field B<sub>A</sub>

High Magnetic Field (Bs)

#### Magnetic Adiabatic Collimation with Electrostatic Filtering

(only electrons with enough energy can overcome potential barrier)





## Cyclotron Radiation Emission Spectroscopy (CRES)





Frequency Approach  ${}^{3}\mathrm{H} \rightarrow {}^{3}\mathrm{He}^{+} + e^{-} + \bar{\nu}_{e}$ 



## KATRIN Outlook

KATRIN continues to collect data (Runs 2 & 3 already obtained). A 1000 day data set is expected to reach design goal.

Comprehensive campaign to reduce and mitigate backgrounds, including radon and Rydberg events.

Better measurement/control of plasma instabilities in source.





**Increased** statistics



**Background mitigation** 



Assessment of plasma effects

## Beyond the Degeneracy Scale

Any experiment with a molecular tritium (T<sub>2</sub>) source will have a systematic penalty associated with uncertainty in the width of rotational and vibrational states of the daughter <sup>3</sup>HeT<sup>+</sup> populated in the decay.

In order to push to the next target mass scale (IO), one will need to switch to an atomic tritium source.



## Transition to an Atomic Source (Phase III)

The endpoint for  $T_2$  is higher than for atomic tritium. Thus, any atomic tritium experiment must be extremely pure ( $T_2$  / T< 10<sup>-5</sup>).

At low densities, recombination occurs mainly on surfaces. Thus a magnetic trap is necessary to prevent recombination. Can utilize magnetic moment of atomic tritium.

 $\Delta E = -\vec{\mu} \cdot \vec{B}$ 

Ioffe traps and Halbach arrays can have large fields near surfaces, with a large uniform region in the center suitable for CRES.





<u>Solution</u>: A large volume magnetic trap for T atoms

## Toward an Inverted Ordering Experiment



Ultimate atomic tritium experiment combines R&D from Phase III into large RF array tritium trap.

- Conversion from molecular to atomic.
- Demonstrate transport, cooling and trapping of tritium.
- Detection via CRES antenna array

Current Project 8 R&D effort (Phase III) focused on developing these technologies.

