# Investigating the transition probability discrepancy along the N=28 neutron shell closure via direct transfer reaction on $^{46}\mathrm{Ar}$

### D. Brugnara et al.

University of Padova - LNL INFN

2020





## The N=28 shell closure



- How does the N=28 shell closure evolve? While <sup>48</sup>Ca is a famous proof of its existence, neutron-rich isotopes below <sup>42</sup>Si no longer demonstrate its presence.
- Transition probabilities measured by (intermediate and near-Coulomb-barrier) coulex measurements hint at a reduced B(E2) compared to the possibility of an onset of collectivity in <sup>46</sup>Ar
- Shell model calculations, which account correctly for the breakdown of the shell gap in S and Si are at odds with B(E2) measurements
- Is there more to be understood of the proton contribution to the wave-function?

# A neutron deponent observable: $D_n$

### D<sub>n</sub> [Z. Meisel et al. PRL 114, 022501 (2015)]



Mass measurements confirm the N=28 shell closure in <sup>46</sup>Ar and its breakdown in the S isotopes (by observing a peaked value of D<sub>n</sub> at N=28 with a sudden drop for more neutron-rich <sup>46</sup>Ar isotopes)

 Experimental data and theory well in agreement ( => SPDF-U describes well the valence-core neutron interaction)

# A neutron deponent observable: $D_n$

### D<sub>n</sub> [Z. Meisel et al. PRL 114, 022501 (2015)]



- Mass measurements confirm the N=28 shell closure in <sup>46</sup>Ar and its breakdown in the S isotopes (by observing a peaked value of D<sub>n</sub> at N=28 with a sudden drop for more neutron-rich <sup>46</sup>Ar isotopes)
- Experimental data and theory well in agreement ( => SPDF-U describes well the valence-core neutron interaction)

# Quadrupole transition probabilities

### B(E2) [A. Gade et al., PRC 68, 014302; S. Calinescu et al., PRC 93, 044333]



- Divergence in trend and in value between shell model calculations and coulomb excitation measurements
- Other isotopes are well reproduced
- A systematic calculation of the contribution on the matrix elements from protons and neutrons hints that a discrepancy that lies on the former would explain the observed divergence in B(E2)

# The physics case

# The aim of the experiment is to probe the proton wavefunction by a transfer direct reaction in ${}^{46}Ar$

- We are looking into a well studied isotope Spin assignments are well established
- The crucial aspect is to investigate the relative s and d transfer
- Spectroscopic factors are known for <sup>48</sup>Ca p removal: what for <sup>46</sup>Ar p addition?



# The physics case

# The aim of the experiment is to probe the proton wavefunction by a transfer direct reaction in $^{\rm 46}{\rm Ar}$

- We are looking into a well studied isotope Spin assignments are well established
- The crucial aspect is to investigate the relative s and d transfer
- Spectroscopic factors are known for <sup>48</sup>Ca p removal: what for <sup>46</sup>Ar p addition?

# (Simplistic) single particle picture



Investigating the transition probability discrepancy along the N=28 neutron shell closure via direct transfer reaction on  $^{46}$  Ar

└─ The Physics Case

L The Observables

## What we expect

▶  ${}^{46}$ Ar( ${}^{3}$ He, d) ${}^{47}$ K @ 10MeV/u. (primary beam  ${}^{48}$ Ca)  $I_{beam} = 3 \times 10^{4}$  Hz



Investigating the transition probability discrepancy along the N=28 neutron shell closure via direct transfer reaction on  $^{46}$  Ar

The Physics Case

└─ The Experimental Setup

# **Detection Arrays**

## ► AGATA+MUGAST+VAMOS+Cryo <sup>3</sup>He Target+Must2+Cats2



- MUGAST: Energy, Position, Particle discrimination
- VAMOS: Reaction fragment identification, Beta
- AGATA: Gamma energy, Position
- CATS2: Time of flight
- Cryogenic target: cooled to 7K, density of  $\approx 1.5 \times 10^{-3}$  g/cm<sup>2</sup>, equiv. to  $\approx 5.0 \times 10^{-3}$  g/cm<sup>3</sup>

Investigating the transition probability discrepancy along the N=28 neutron shell closure via direct transfer reaction on  $^{46}$ Ar

— The Physics Case

A peek at the data

# Double coincidences: VAMOS + AGATA

## $\gamma$ spectrum on AGATA with ${\rm ^{47}K}$ identified in VAMOS



- Many protons are identified in MUGAST (one order of magnitude more respect to the deuterons) ⇒
  Deuteron breackup reaction channel, three body reaction
- ▶ The level scheme of <sup>47</sup>K is observed

э

A peek at the data

# Double coincidences: MUGAST+VAMOS (Kinematics)



- Ice deposition and target thickness affect the energy resolution (an thus the kinematic plot).
- As expected, the transfer to the <sup>1</sup>/<sub>2</sub> g.s. and the 360 keV <sup>3</sup>/<sub>2</sub> level cannot be disentangled based on Excitation energy.

The PDF de-convolution on the angular distribution will assess the ratio between L=0 and L=2 transfer.

## Eff. corrected $d\sigma/d\theta$ .



A peek at the data

# Triple coincidences: AGATA+VAMOS+MUGAST

## Kinematic line in coincidence with 360 KeV $\gamma$ -ray



- ▶ 360 keV γ detected by AGATA+<sup>47</sup>K identified in VAMOS+ deuteron measured in MUGAST ⇒ clear kinematic line
- Narrow  $\gamma$  gate removes background

## $\gamma\text{-rays}$ in coincidence with d and $^{47}\mathrm{K}$



- Despite the majority of counts in the Excitation energy peak are centered at 0 MeV, the only counts in coincidence with AGATA are at 2 MeV
- No gammas seen for Ex< 1MeV</p>
- Suppressed direct transfer to  $\frac{3}{2}^+$  ?

### Perspectives

# Concusions and future perspectives

- The reaction is correctly populated and identified thanks to the granularity of the setup
- Final refinements on the data analysis are needed
- Results will be compared and discussed in the framework of theoretical models, namely the shell model

### Perspectives

## The Collaboration

- G. de Angelis, E. T. Gregor, A. Gottardo, A. Illana, G. Jaworski, D.R. Napoli, M. Siciliano, J.J. Valiente-Dobon, F. Galtarossa, I. Zanon INFN, Laboratori Nazionali di Legnaro, Legnaro, Italy.
- F. Flavigny, M. Assie, D. Beaumel, Y. Blumenfeld, D. Verney, F. Ibrahim, M. Babo, C. Delafosse, F. Hammache, L. Lalanne, I. Matea, N. De Sereville IPN. Orsav. France.
- D. Bazzacco, D.Brugnara, A. Goasduff, S.M. Lenzi, S. Lunardi, R. Menegazzo, D. Mengoni, F. Recchia, D. Testov, Dipartimento di Fisica e Astronomia and INFN, Sezione di Padova, Padova, Italy.
- I. Lombardo INFN, Sezione di Catania, Italy
- A. Matta LPC Caen, Caen, France
- G. Benzoni, A. Bracco, S. Bottoni, F. Camera, F.C.L. Crespi, S. Leoni, B. Million, O. Wieland INFN, Milano, Milano, Italy.
- E. Clement, G. de France, A. Lemasson GANIL, Caen, France.
- M. Zielinska CEA Saclay, IRFU/SPhN, Gif-sur-Yvette, France.
- K. Wimmer University of Tokyo, Tokyo, Japan
- A. Shrivastava, K. Mahata BARC, Mumbai, India

Summary

# Thank you for your attention

### Overview

The Physics Case The Observables The Experimental Setup A peek at the data Perspectives Summary





イロト イヨト イヨト イヨト 二日



