



cherenkov  
telescope  
array



# lstchain developments

Rubén López-Coto  
INFN Padova, 30/06/20  
INFN Analysis meeting



Funded by H2020 Marie Skłodowska  
Curie FELLINI - Grant 754496



- cta-lstchain (<https://github.com/cta-observatory/cta-lstchain>):
  - ctapipe-based pipeline that includes all the tools for LST data processing from R1 up to DL2.
- This presentation: Updates in the new software release

Pre-release

v0.1.0

5f4710a

Verified

Compare ▾

## First release of cta-lstchain



rlopezcoto released this on 18 Oct 2019 · [979 commits](#) to master since this release

First tagged release of `cta-lstchain`, necessary to keep track of a version with which data taken in La Palma were analyzed.

# cta-lstchain recap

- **R1**  $\Rightarrow$  **DL1** (from camera raw waveforms to calibrated & parametrized images)

- Low level (DRS4) waveform corrections.
- Pulse integration & calibration (pixel-wise charge in p.e., time in ns)
- Image cleaning & parametrization  $\Rightarrow$  DL1
- Identification and muon analysis

Notes:

- currently R1 is stored permanently (all pixels, all waveforms)
- no official DL0 data format & model yet
- no lossy data reduction scheme established yet

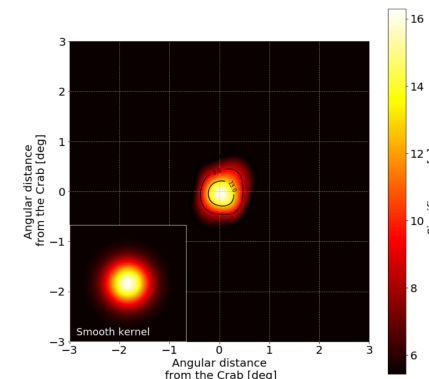
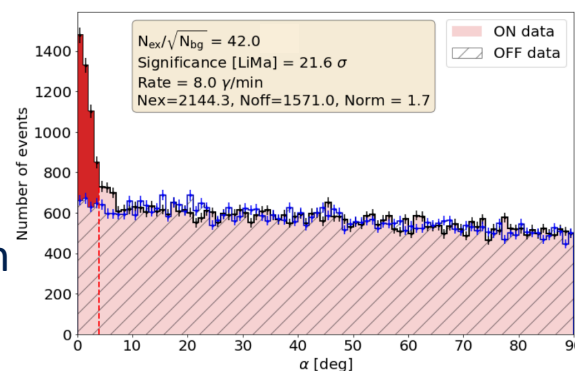
- **DL1**  $\Rightarrow$  **DL2**

Use scikit-learn's Random Forest, trained on MC, for gamma/hadron separation and direction and energy reconstruction

- **DL2**  $\Rightarrow$  **DL3**

apply gammaness and direction cut to produce list of gamma candidates

First Crab LST detection  
December 2019



# lstchain evolution

Oct 2019

v0.1.0

Nov 2019

v0.2.0

First  
release to  
analyze  
real data

v0.2.1

Nov 2019

v0.3.0

v0.3.1

First Crab  
campaign

Jan 2020

v0.4.0

v0.4.1

...

v0.4.5

May 2020

v0.5.0

**Now**

v0.5.1

# Current version

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- Istchain v0.5.0 released a few weeks ago:
  - Many code and repo improvements.
  - Volume reduction.
  - Improved muon analysis.
  - Usage of interleaved information in the calibration.
  - Integrated datacheck for data at different data levels.
  - Source-dependent analysis.
  - Proper timing and pointing information from DL1 level on.

Pre-release

 v0.5.1

 2b06f89

Verified

## v0.5.1

 release-drafter released this 27 days ago ·

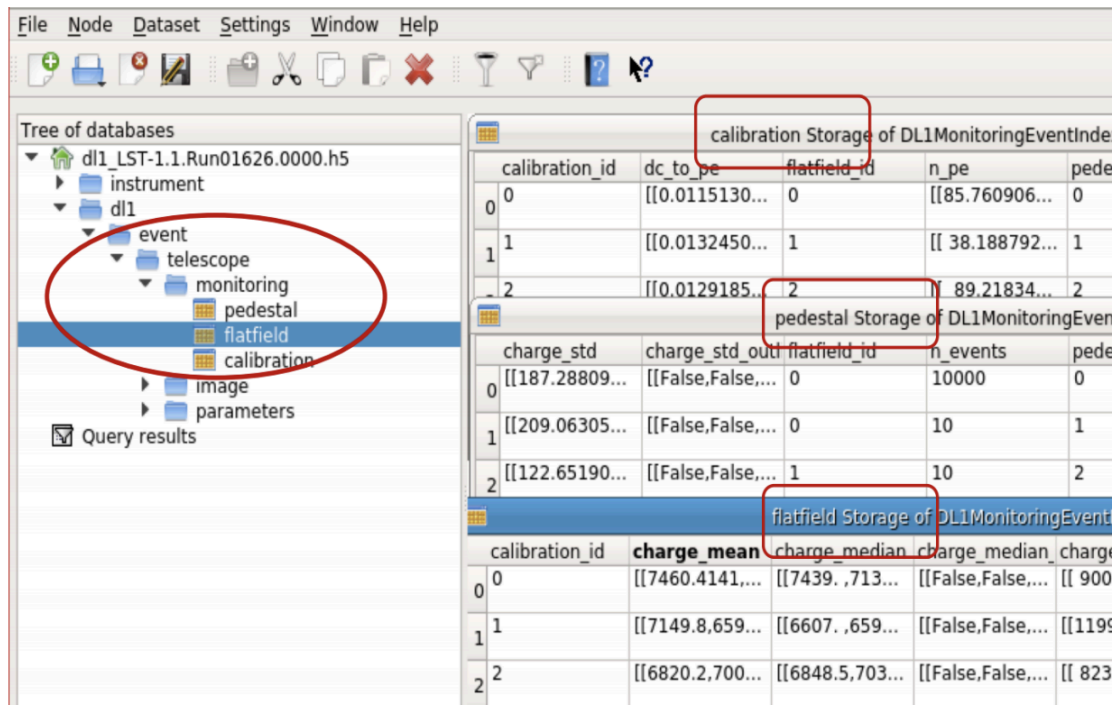
## What's Changed

Currently following  
**v0.5.1** after a few  
bugfixes

# Calibration

- Interleaved information included in the DLx files now (F. Cassol)
  - Unfortunately tagging information of current interleaved data is mostly non reliable.
- HG vs LG agreement pretty good after the implementation of the latest low level correction curves. (Pawel, Yukiho, Seiya, Mitsunari...)

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Tree of databases

- dl1\_LST-1.1.Run01626.0000.h5
  - instrument
  - dl1
    - event
      - telescope
        - monitoring
        - pedestal
        - flatfield
        - calibration
        - image
        - parameters

Query results

calibration Storage of DL1MonitoringEventIndex

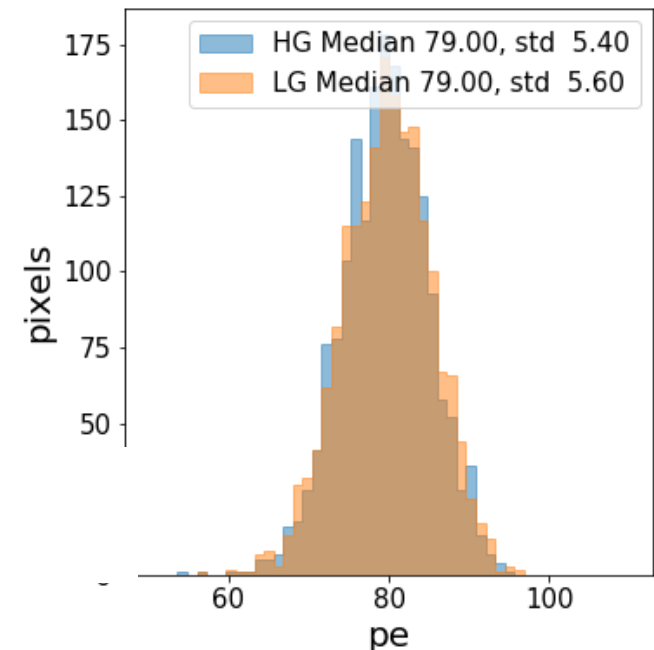
calibration_id	dc_to_pe	flatfield_id	n_pe	pedestal_id
0	[[0.0115130...	0	[[85.760906...	0
1	[[0.0132450...	1	[[ 38.188792...	1
2	[[0.0129185...	2	[[ 89.21834...	2

pedestal Storage of DL1MonitoringEventIndex

charge_std	charge_std_out	flatfield_id	n_events	pedestal_id
[[187.28809...	[[False,False,...	0	10000	0
[[209.06305...	[[False,False,...	0	10	1
[[122.65190...	[[False,False,...	1	10	2

flatfield Storage of DL1MonitoringEventIndex

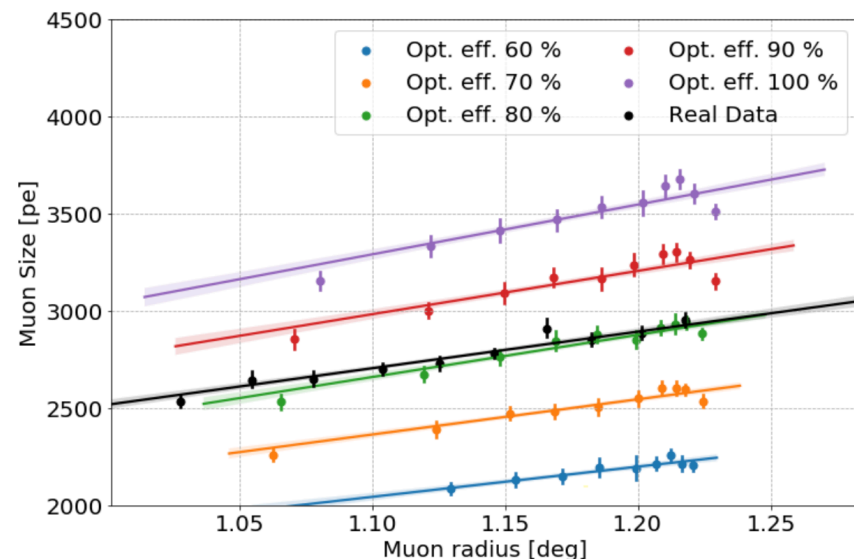
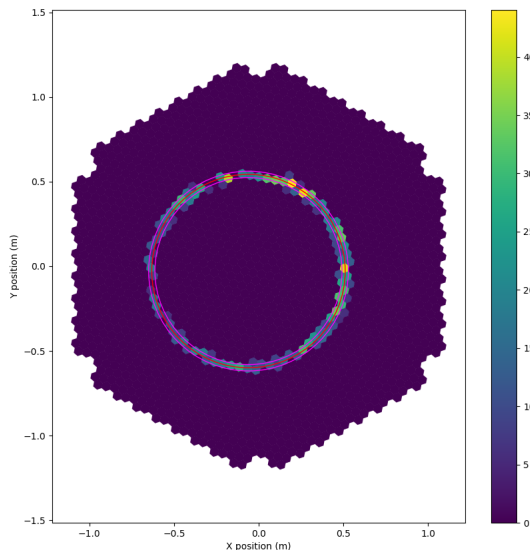
calibration_id	charge_mean	charge_median	charge_median	charge
0	[[7460.4141...	[[7439. ,713...	[[False,False,...	[[ 900.
1	[[7149.8,659...	[[6607. ,659...	[[False,False,...	[[1199
2	[[6820.2,700...	[[6848.5,703...	[[False,False,...	[[ 823.



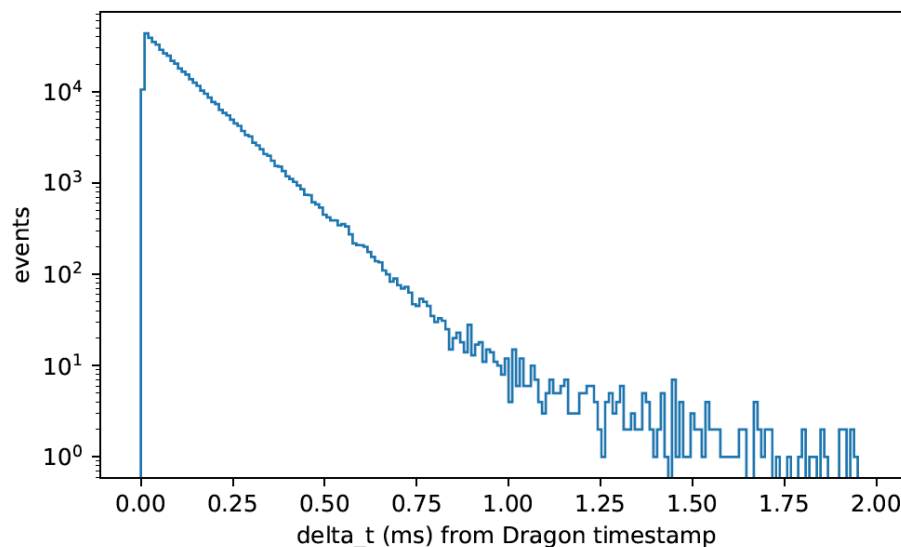
HG/LG comparison

# Muon analysis

- Now performed together with the R1 -> DL1 step (no significant increase in processing time) (A. Moralejo)
  - To avoid signal extraction biases, the GlobalPeakWindowSum is now used.
  - At the moment, an ~80% overall Optical efficiency is reached using the same signal integration for Data/MC, with respect to Prod3b simulations -> currently working on more realistic, measured parameters for Prod5.



- Quantities to be checked are written into an h5 file per subrun analyzed -> Later merged and plotted together per run (A. Moralejo)
  - Many checks currently implemented -> necessary to spot data problems the day after the data taking.
  - Includes rates, timestamps, flatfield, pedestal, cosmic and muon event checks.

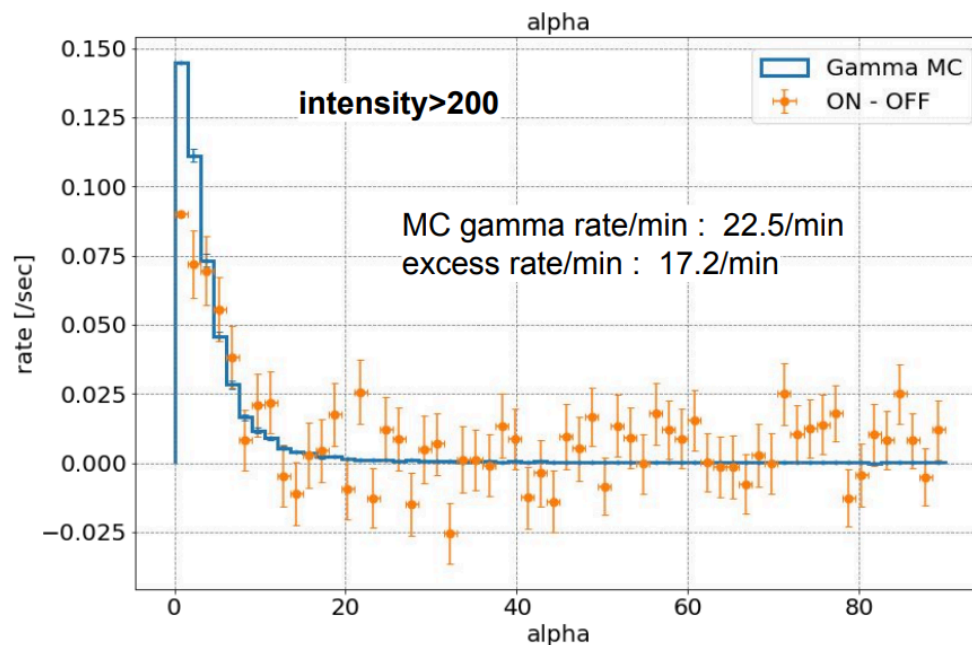
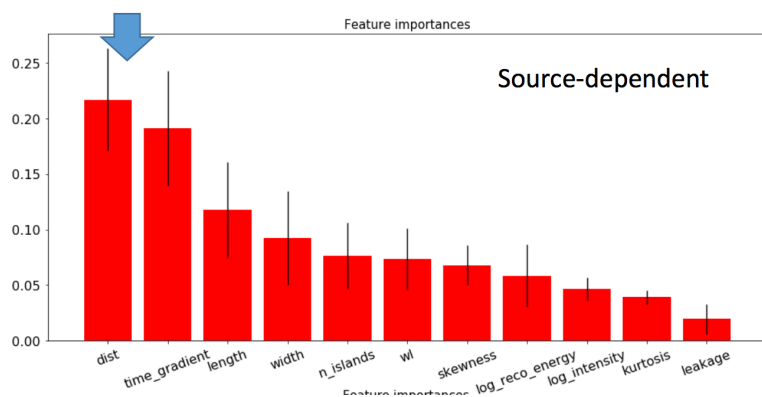




# Source-dependent analysis

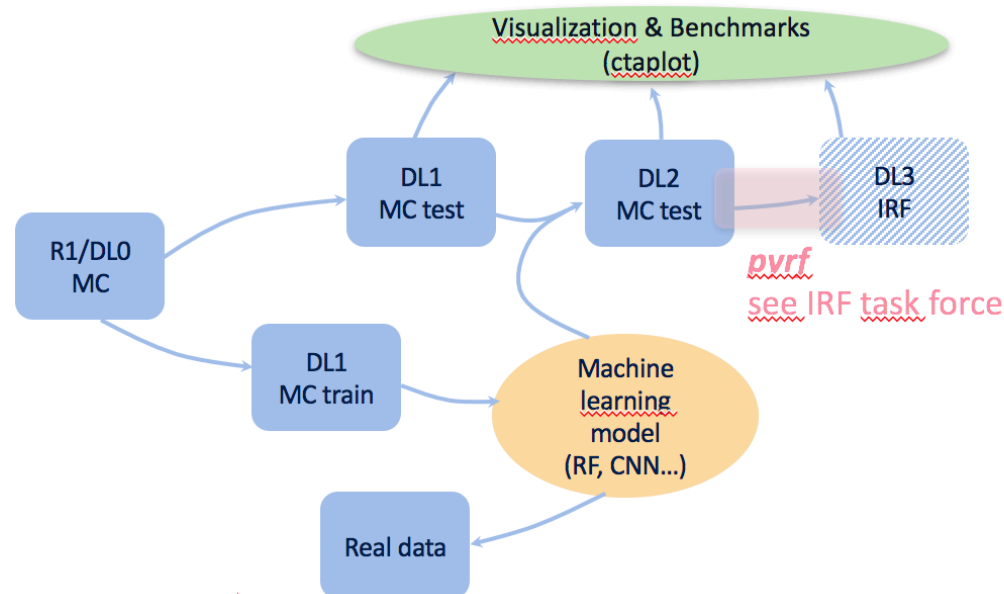


- Single telescope source dependent analysis (S. Nozaki)
  - Usage of the information of the assumed location of the source in the analysis.
  - Source-dependent parameters as “dist” are very useful for image reconstruction



# High level analysis

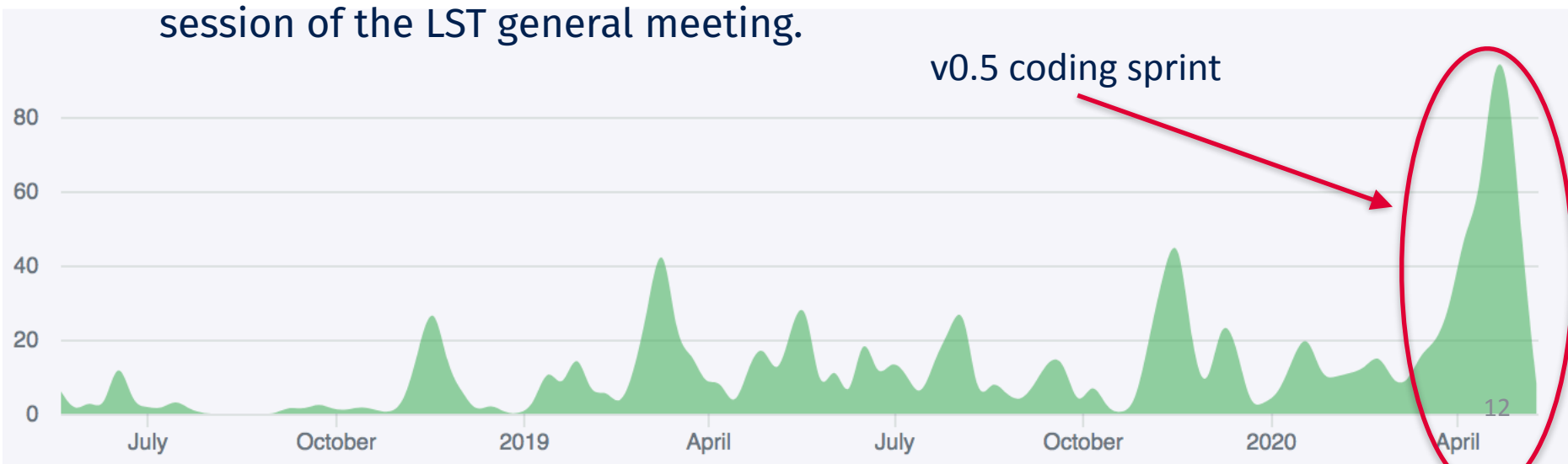
- Automatized training of RF models after software releases (T. Vuillaume, E. Garcia)
  - Preliminary results of training using real data as OFF to improve gamma/hadron separation (A. Baquero).
  - Standard analysis uses models trained on MC (Gammas/Protons).
  - IRF production:
    - Currently custom-made formats.
    - Development of pyIRF within the IRF working group to produce IRFs in FITS format to be analyzed with standard high level software (T. Vuillaume).



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- Towards lstchain v0.6
  - Follow ctapipe v0.8.0 -> Many changes expected (possibly a remote coding sprint needed?)
  - Implement latest low level calibration corrections
  - Optimize cleaning levels using interleaved information.
  - Update calibration constants using interleaved information.
  - Implement the same signal extraction for MC and Data.
  - Start thinking on new, fine tuned MC production.
- News on data analysis using all these new tools shown in the software session of the LST general meeting.



# Coding sprint week 6-10 July



- If you want to contribute, save this week in your agendas
  - mainly experts will be working on adapting the code to ctapipe 0.6. See list: <https://github.com/cta-observatory/cta-lstchain/issues/357>
  - if you are not a lstchain expert you can:
    - start with small implementations.
    - review code.
    - write test units, documentation...



# Contributors

