

# CYGN O P R O J E C T

**E. Baracchini,<sup>a,b</sup> L. Benussi,<sup>c</sup> S. Bianco,<sup>c</sup> C. Capoccia,<sup>c</sup> M. Caponero,<sup>c,d</sup> G. Cavoto,<sup>e,f</sup> A. Cortez,<sup>a,b</sup> I. A. Costa,<sup>g</sup> E. Di Marco,<sup>e</sup> G. D'Imperio,<sup>e</sup> G. Dho,<sup>a,b</sup> F. Iacoangeli,<sup>e</sup> G. Maccarrone,<sup>c</sup> M. Marafini,<sup>e,h</sup> G. Mazzitelli,<sup>c</sup> A. Messina,<sup>e,f</sup> R. A. Nobrega,<sup>g</sup> A. Orlandi,<sup>c</sup> E. Paoletti,<sup>c</sup> L. Passamonti,<sup>c</sup> F. Petrucci,<sup>i,j</sup> D. Piccolo,<sup>c</sup> D. Pierluigi,<sup>c</sup> D. Pinci,<sup>e,1</sup> F. Renga,<sup>e</sup> F. Rosatelli,<sup>c</sup> A. Russo,<sup>c</sup> G. Saviano,<sup>c,k</sup> and S. Tomassini<sup>c</sup>**

# CYGNO PROJECT

The aim of CYGNO project is the development and realisation of a GEM-based Optically Readout Time Projection Chamber for the study of rare events with energy releases in the range 1-100 keV.

Expected performance is:

- High detection efficiency down to 1 keV;
- Directionality at 10 keV;
- Background rejection below 10 keV;

Main ideas of the technology are:

- He/CF<sub>4</sub> based gas target (atmospheric pressure);
- GEM amplification stage;
- Combined optical readout (CMOS for granularity + PMT for timing);

# PHASES OF PROJECT

## PHASE 0: R&D

## PHASE 1: ~1M<sup>3</sup> DEMONSTRATOR

2018

2019

2020

2021/22

2023 ...

@ ROMA1/LNF

@ LNF

@ LNF/LNGS

@ LNF/LNGS

@ LNGS

### ORANGE

### LEMON

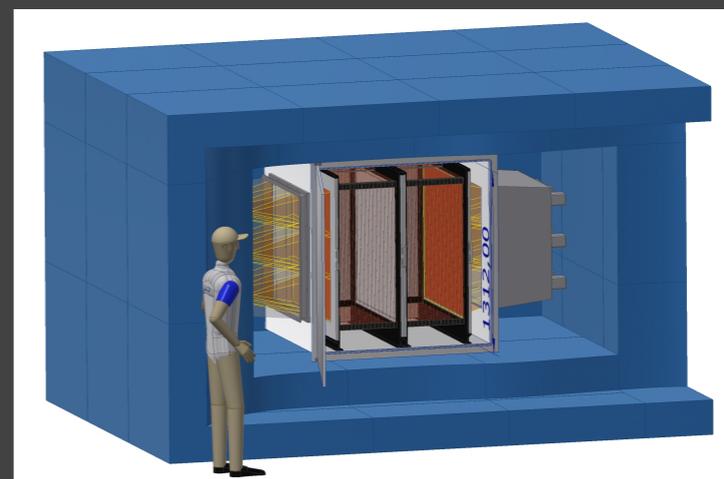
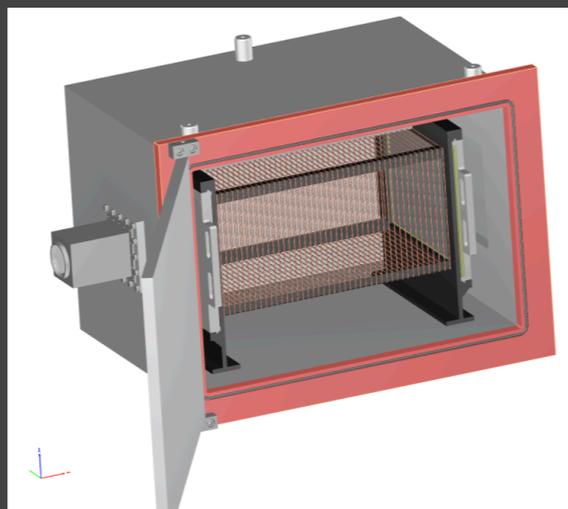
### LIME

Construction & test

Installation & commissioning

### CYGNUS

CYGNUS 30-100 m<sup>3</sup>



- OPT readout
- 1 cm drift
- 0 resolution

- OPT readout
- 3D printing
- 20 cm drift
- PID
- tracking
- drift resolutions

- 50 cm drift
- materials test
- underground tests
- shielding

- background
- materials test
- gas purification
- shielding
- stability
- scalability
- reliability

**He:CF<sub>4</sub>:SF<sub>6</sub>**  
**0.8-1 atm**

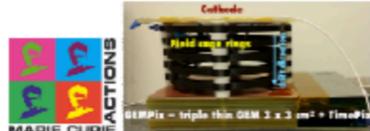
4 Timepix chips  
3 x 3 cm<sup>2</sup> area

1 sCMOS  
10 x 10 cm<sup>2</sup> area

**NITEC**

**MANGO**

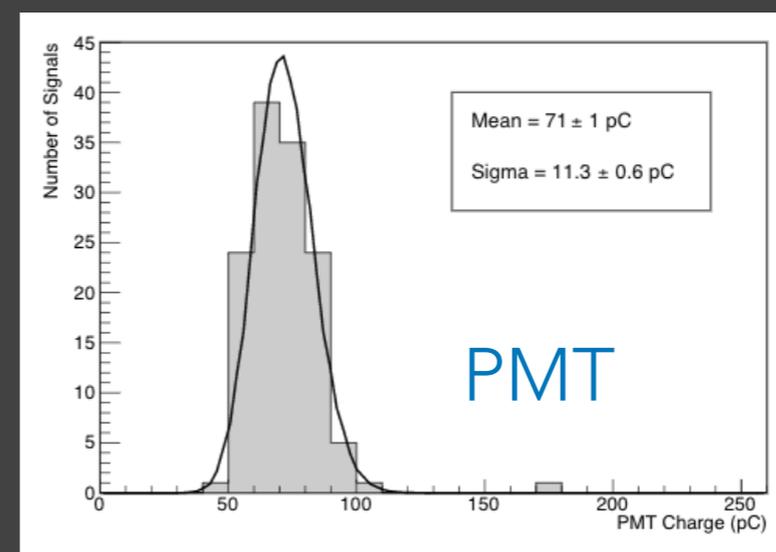
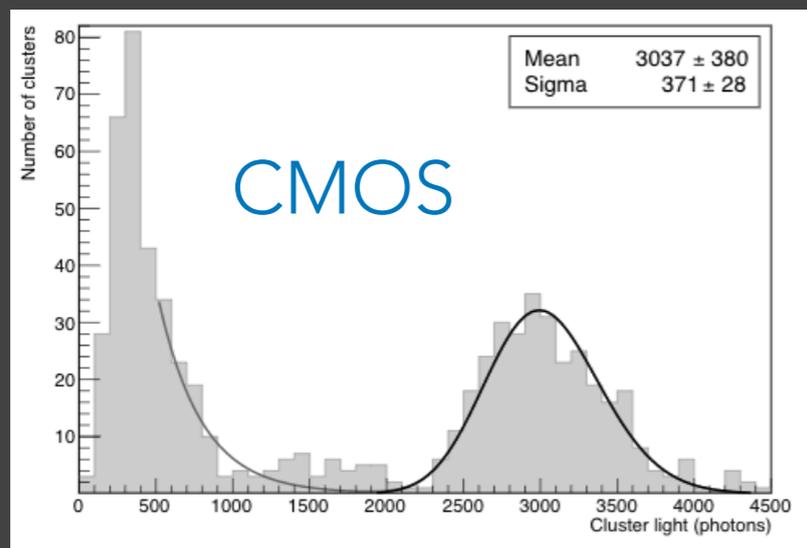
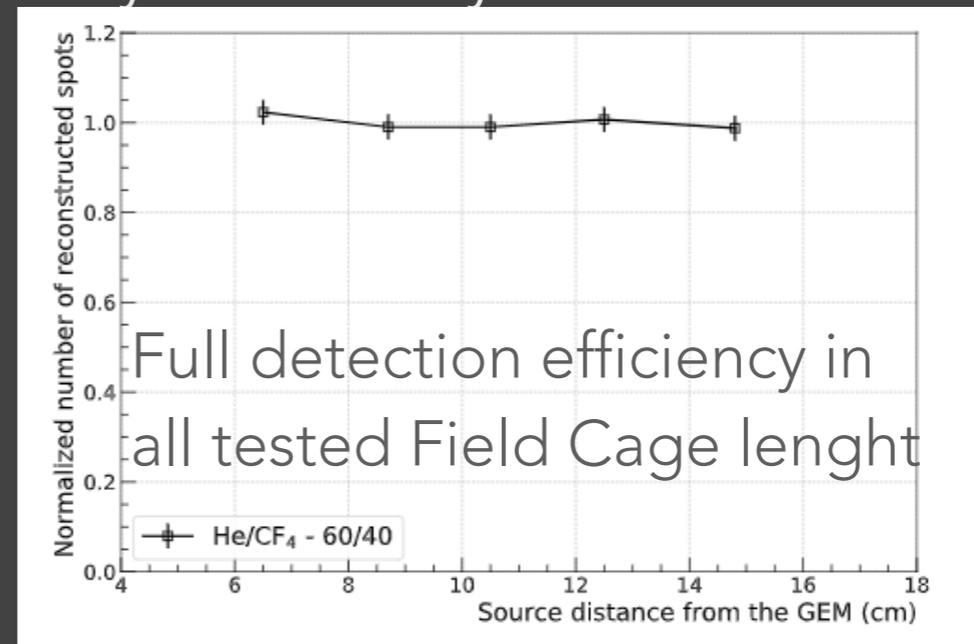
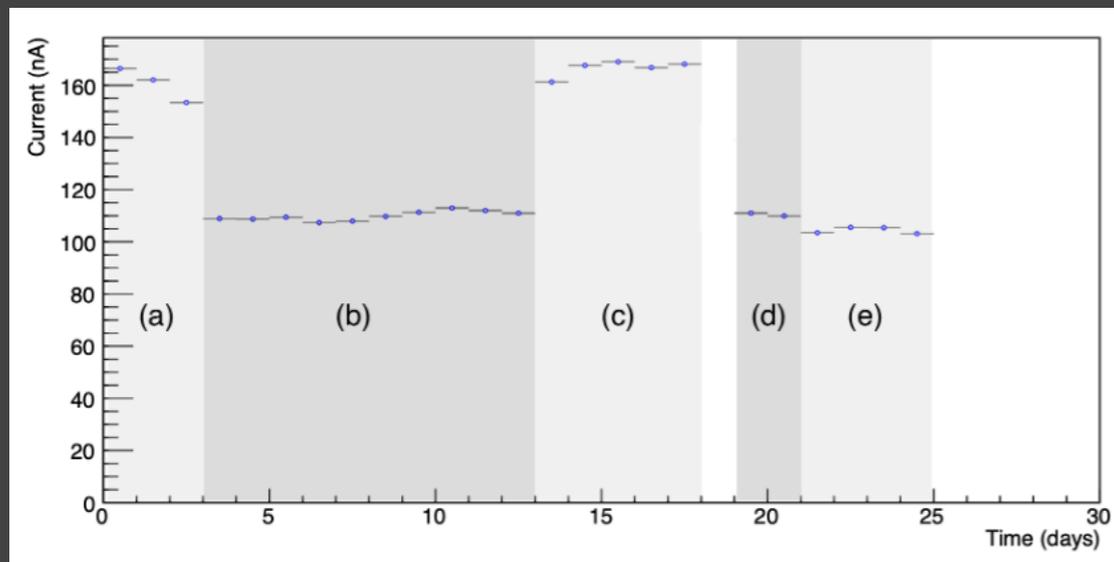
Negative ion drift



WHERE DO WE ARE

# PERFORMANCE WITH $^{55}\text{Fe}$

5.9 keV photons from  $^{55}\text{Fe}$  source were used to test detection efficiency and light yield. Detector run continuously for 25 days



500 photons collected per keV  
Energy resolution of 15%

Similar resolution with PMT

# BACKGROUND REJECTION

To evaluate the background (electron recoils) rejection capability and the signal (nuclear recoils) efficiency, we tested LEMON (a 7 litre prototype) with  $^{55}\text{Fe}$  and AmBe sources.

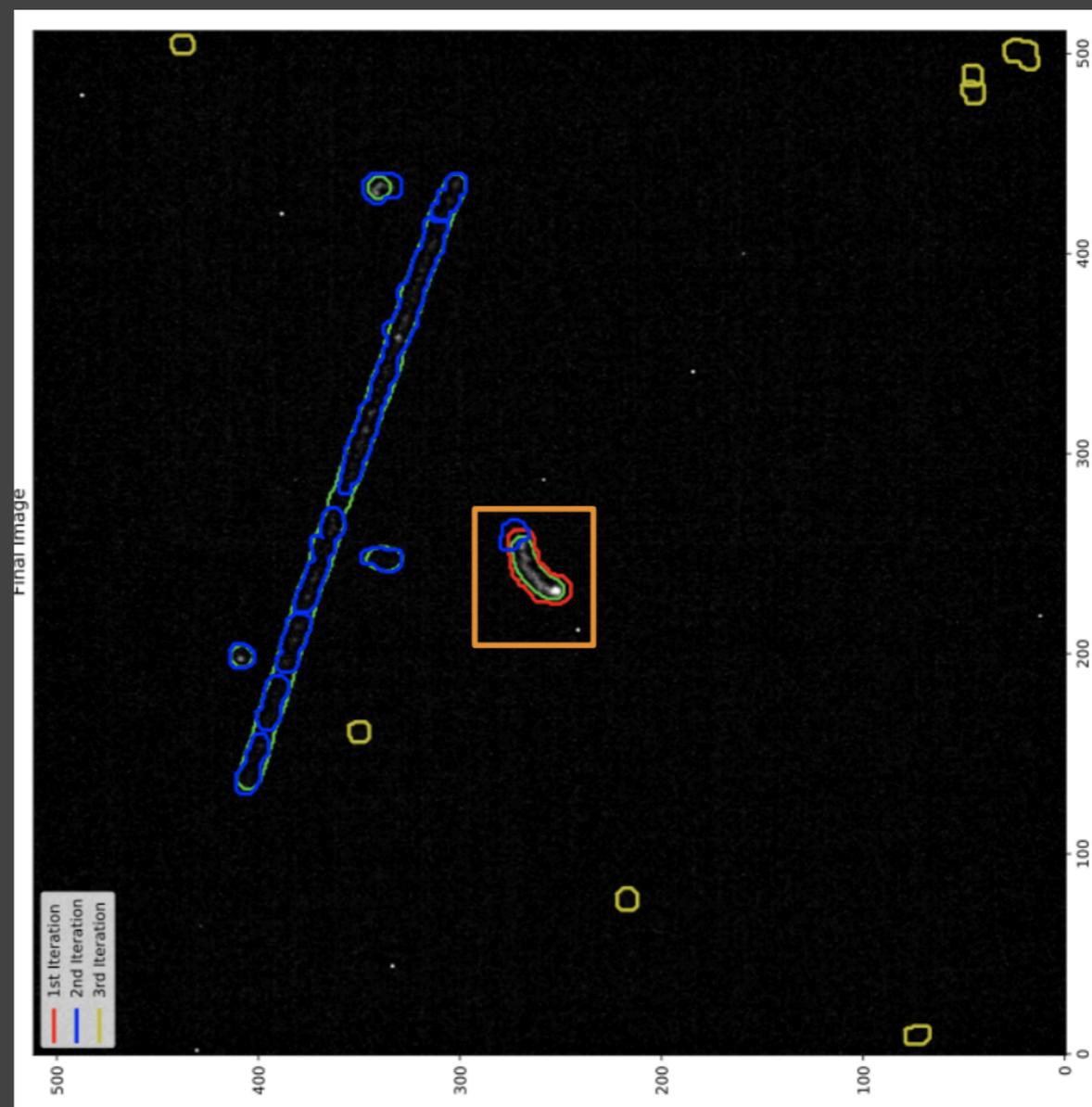
AmBe produces:

- 59 keV photons;
- 4 MeV photons;
- 1-10 MeV neutrons;

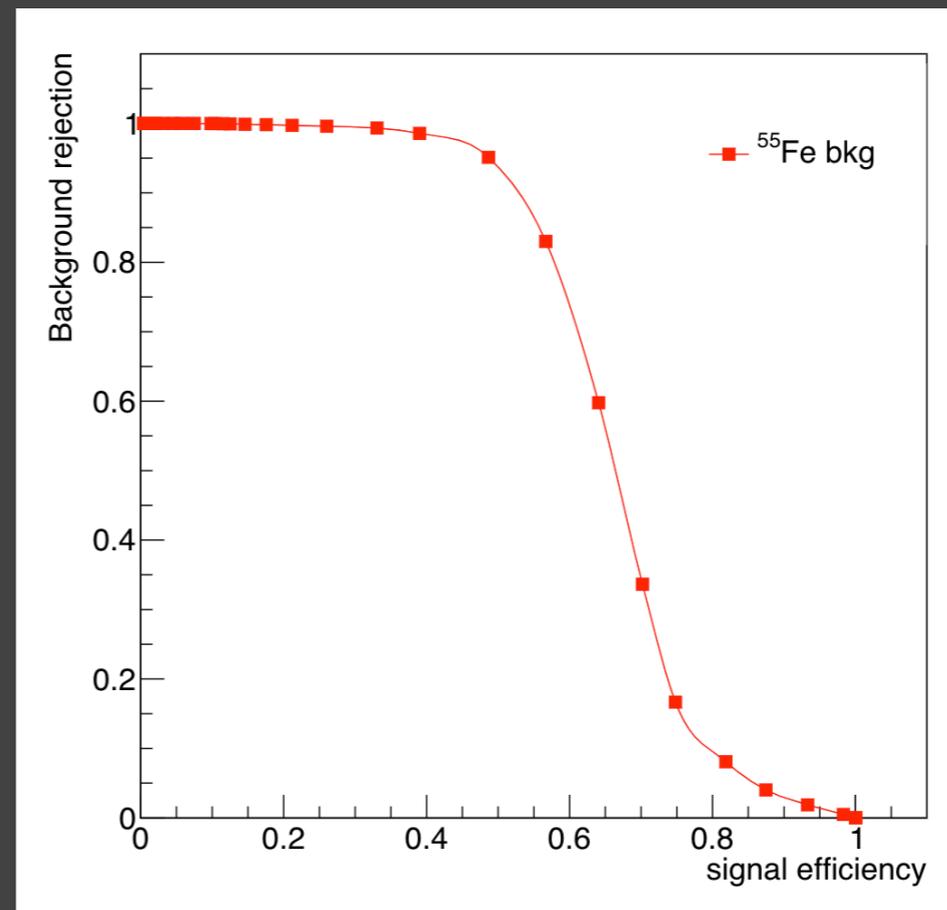
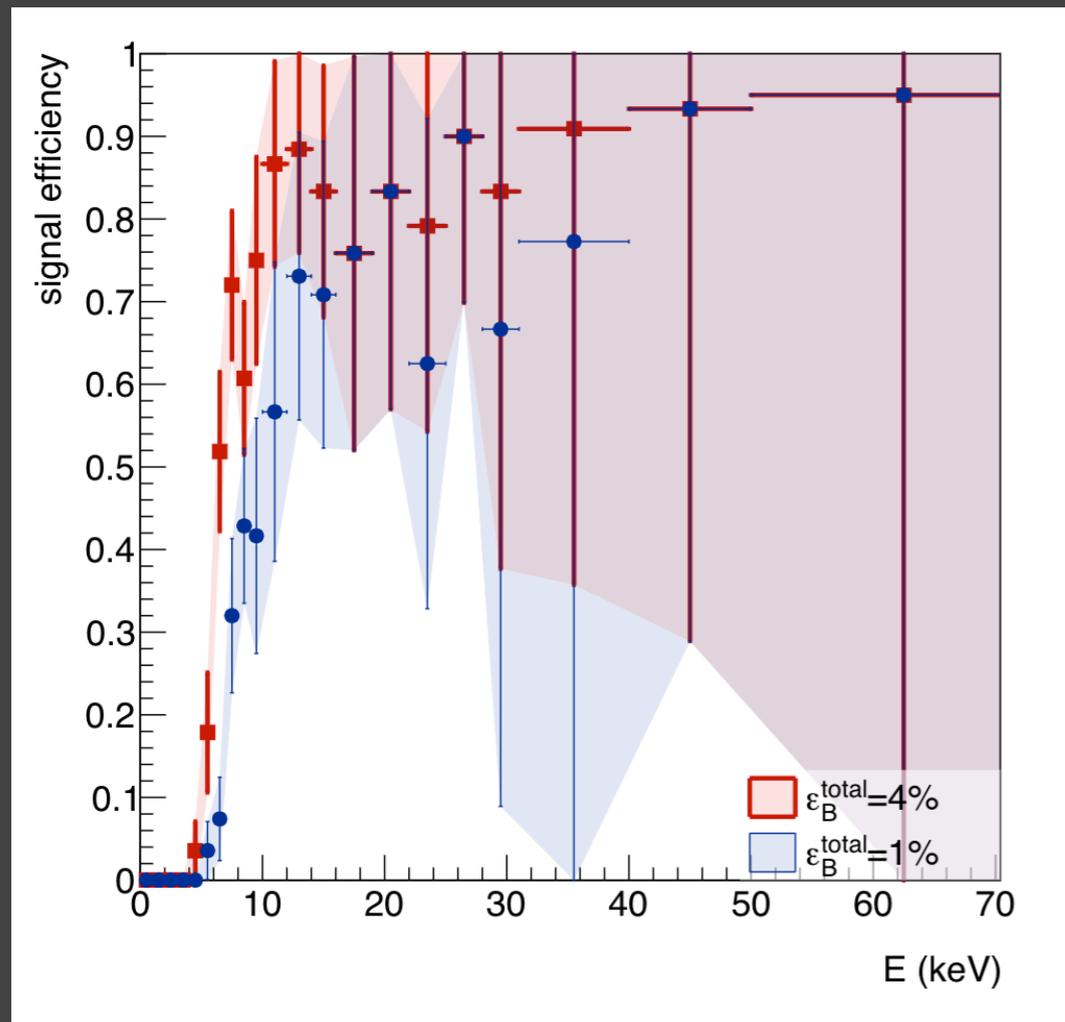
A 5 cm Pb shield was used

Unfortunately in all cases cosmics and natural radioactivity produce an unknown background that piles-up to signals.

Need to go underground.



# SIGNAL EFFICIENCY



working point	Signal efficiency			Background efficiency		
	$\epsilon_S^{presel}$	$\epsilon_S^\delta$	$\epsilon_S^{total}$	$\epsilon_B^{presel}$	$\epsilon_B^\delta$	$\epsilon_B^{total}$
WP <sub>50</sub>	0.98	0.51	0.50	0.70	0.050	0.035
WP <sub>40</sub>	0.98	0.41	0.40	0.70	0.012	0.008

A sizeable efficiency in the range 5-10 keV was measured while more than 95% (99%) <sup>55</sup>Fe photons were rejected

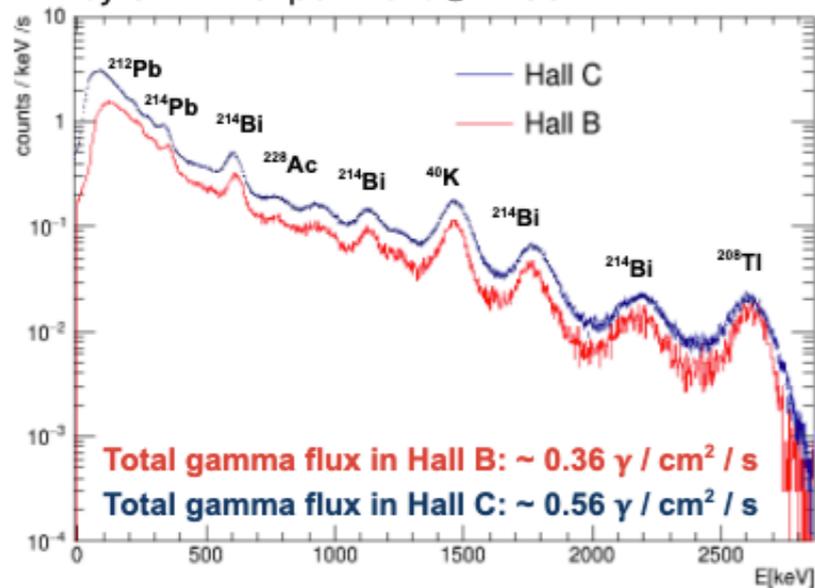
# BACKGROUND SIMULATION

Full Detector simulation in GEANT4

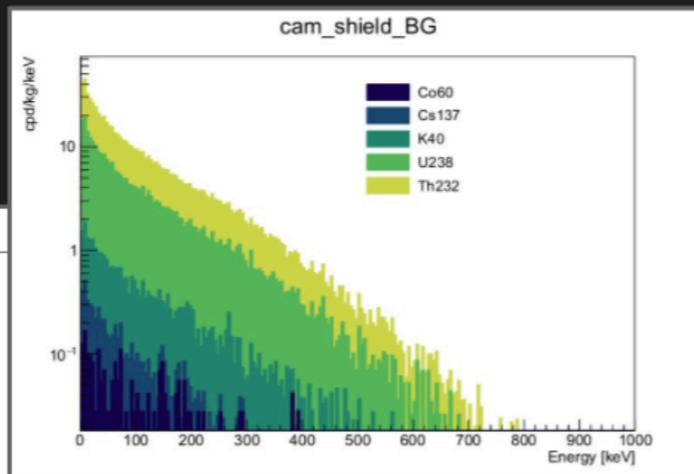
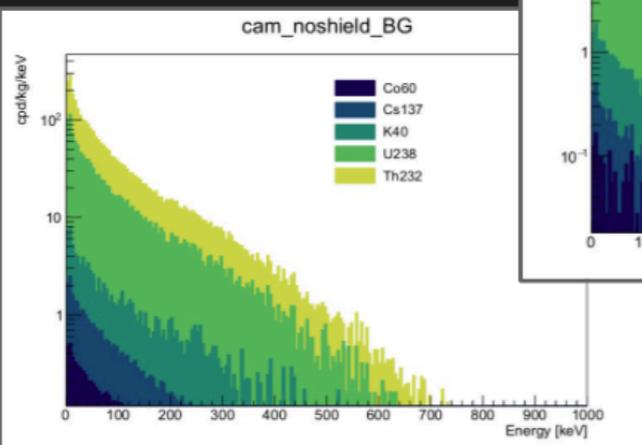
Gamma and neutron background due to internal radioactivity simulated

Measured radioactivity of GEM  
and different Cameras

Input → gamma flux measured  
by SABRE experiment @LNGS



Even with shield, the rate is far  
beyond the threshold we want to  
reach ( $\sim 10^4$  cpy from ER)



in [0-20] keV

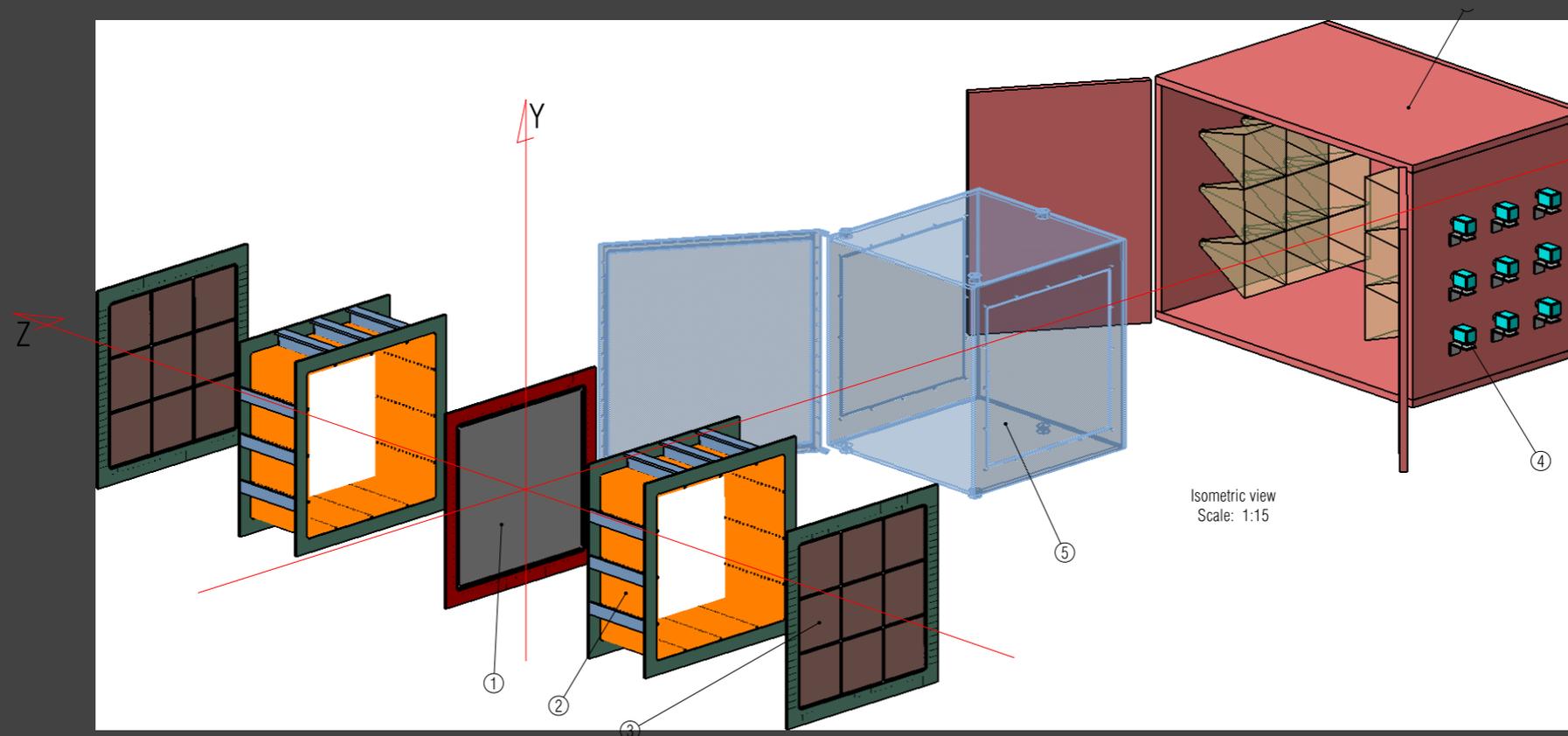
- No shield:  $2.05 \times 10^7$  cpy
- Shield:  $6.60 \times 10^6$  cpy

Camera	Sensitivity (eV/count)	Resolution (%)	Noise (eV)	<sup>228</sup> Ra (Bq)	<sup>228</sup> Th (Bq)	<sup>226</sup> Ra (Bq)	<sup>234</sup> Pa (Bq)	<sup>40</sup> K (Bq)	Total activity
Hamamatsu ORCA FLASH 4.0	2.96	15.2	4.6	2.1	2.1	1.9	7.0	1.9	15.0
ORCA FLASH sensor	2.6	15.2	8	1.0	1.0	1.1	1.1	4.3	8.5
Photometrics Prime BSI Mode 1	3.3	19.0	9.7	-	-	-	-	-	tbm
Photometrics Prime BSI Mode 2	1.12	16.4	4.5	-	-	-	-	-	tbm
Photometrics BSI Express Mode 2	0.84	13.4	3.0	1.3	1.8	1.0	6.0	3.6	13.7
Hamamatsu Fusion Closer (LEMON)	0.65	17.5	1.58	-	-	-	-	-	tbm
Hamamatsu Fusion Farther (LIME)	0.85	16.4	2.06	-	-	-	-	-	tbm
Thorlab Quantalux	tbm	tbm	tbm	0.3	0.6	0.2	3.0	1.2	5.3

Included in GEANT4, resulted  
dominant with respect to external  
shielded

# CYGNO BASELINE DESIGN

1 m<sup>3</sup> of He/CF<sub>4</sub> 60/40 (1.6 kg) at atmospheric pressure subdivided in two 50 cm long parts by the cathode with a drift field of about 1 kV/cm

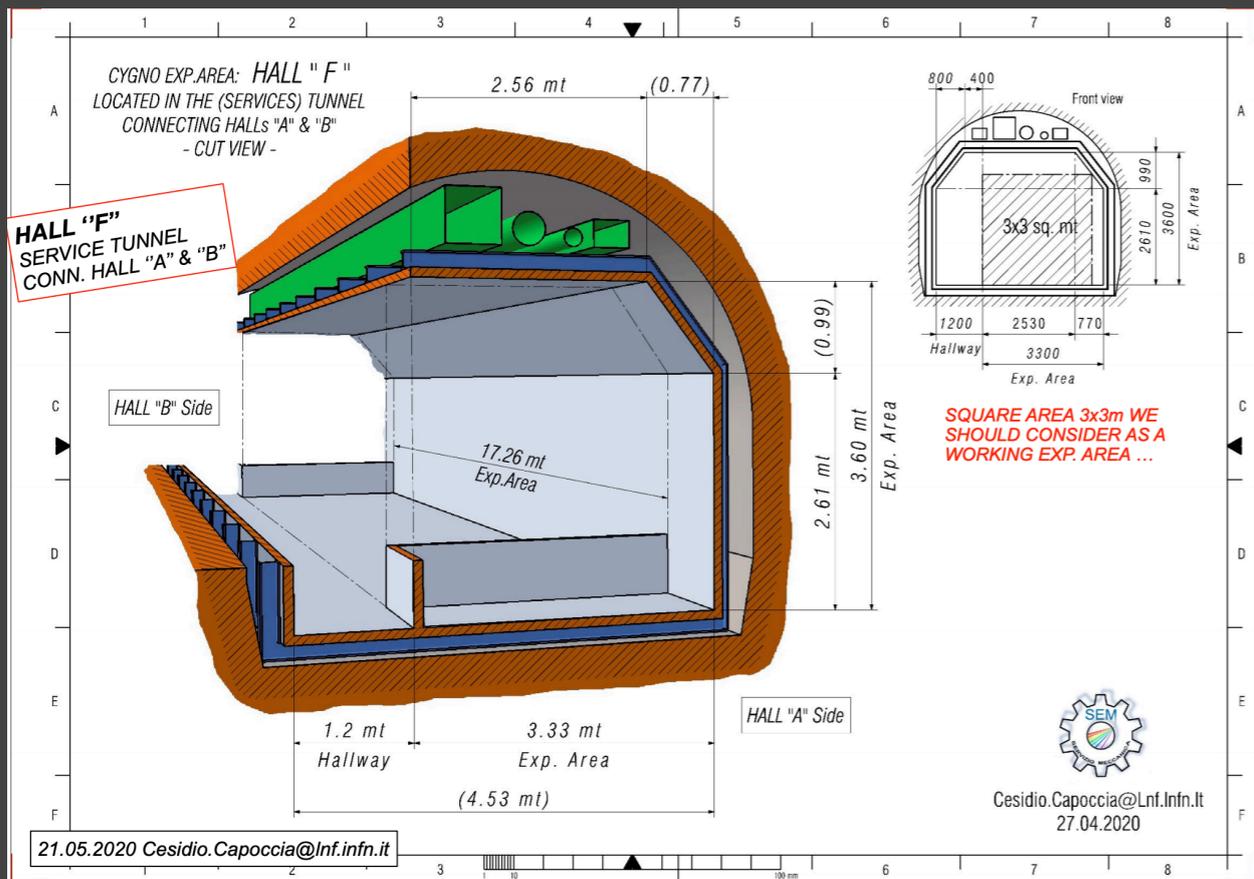


Each side equipped by a 3x3 matrix of LIME-like:

- sCMOS sensor 65 cm away;
- Fast light detector (PMT or SiPM).

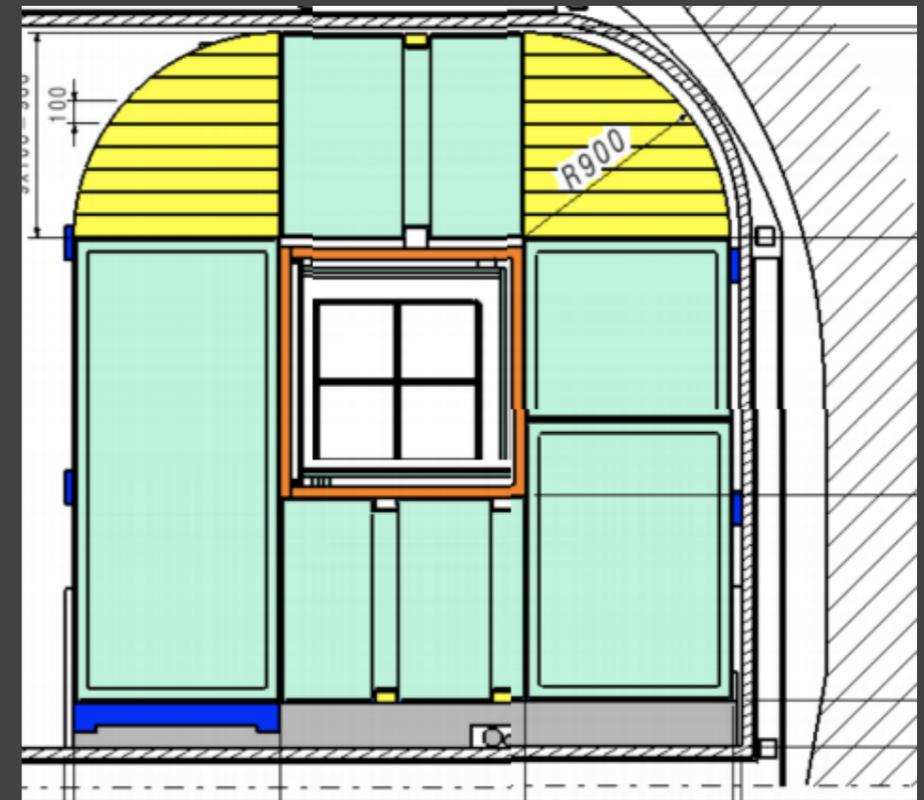
Shielding of 200 cm water and 5 cm copper giving a bkg cpy [1-20] keV of few 10<sup>2</sup>;

# HALL-F: CYGNO IN THE BOTTLE



A possible room for CYGNO is in Hall-F (tunnel connecting A and B halls);

Available cross section is 3x3 m<sup>2</sup> (l=17 m)

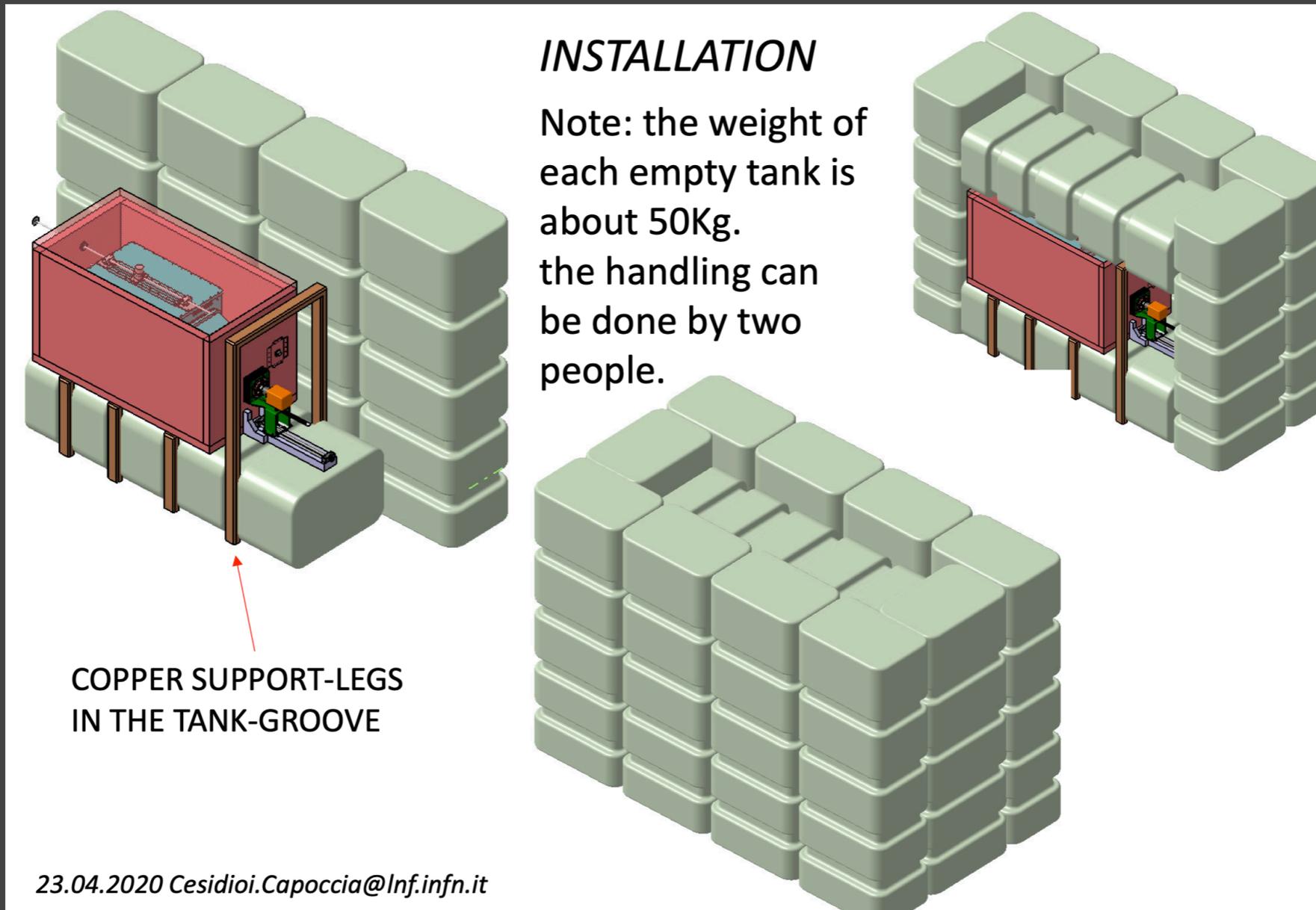


It will be possible to house 0.4 m<sup>3</sup> - CYGNO with a 110/10 (water/Cu) shielding scheme;

Background rate of **2x10<sup>4</sup>** cpy, will be lower than internal bkg (currently **5x10<sup>5</sup>** cpy) and will allow to study it;

With a rejection of **10<sup>2</sup>-10<sup>3</sup>** these numbers will result in an avoidable bkg of few **tens** of events/year **external** and few **hundreds** of events/year **internal**;

# SHIELD



## PRODUCTION COSTS

	MOLD-TOOLING	DRAW-DOCUM	ROT. MOULDING	N.TANK	TOTAL
VERZELLESI SRL	Refused to carry out an evaluation due to the low quantity				
TELCOM SPA	€ 42,500.00	€ 3,000.00	€ 260.00	12	€ 48,620.00
MOULDING SERV. SRL	€ 15,900.00	€ 1,200.00	€ 235.00	12	€ 19,920.00
Diff (%)	267%	250%	111%		244%

We are investigating the possibility of building custom tanks

# BACKGROUND- OPEN POINTS

$^{14}\text{C}$  contamination in gas can give important bkg. Need to know the origin of C in  $\text{CF}_4$ .

Copper and plexiglass can be produced with low-radioactivity content. Need to get in contact with producers, develop special procedures, measure quality of materials.

Need to study a low radioactive GEM production;

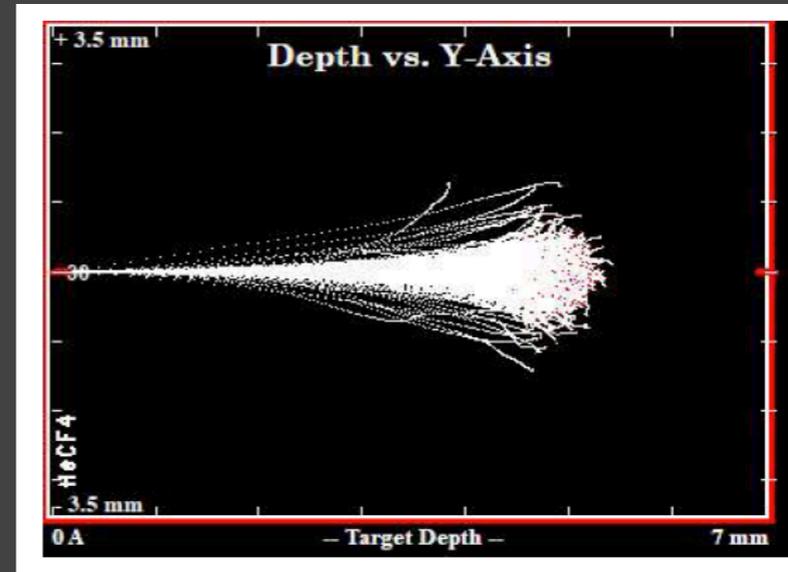
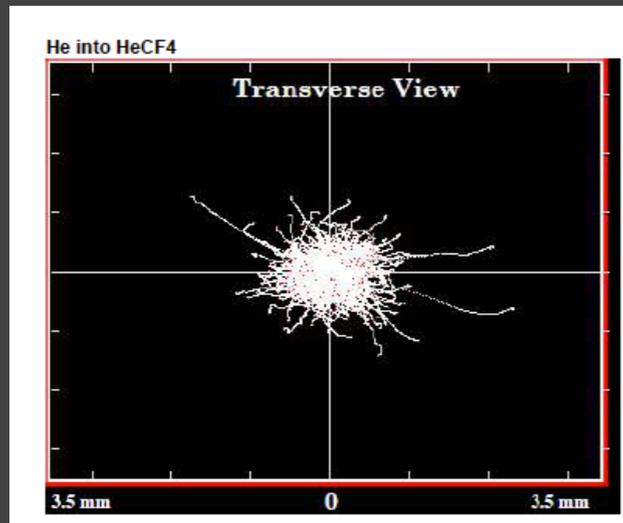
Need to find low radioactive CMOS sensors;

Need to purify and monitor gas mixture.

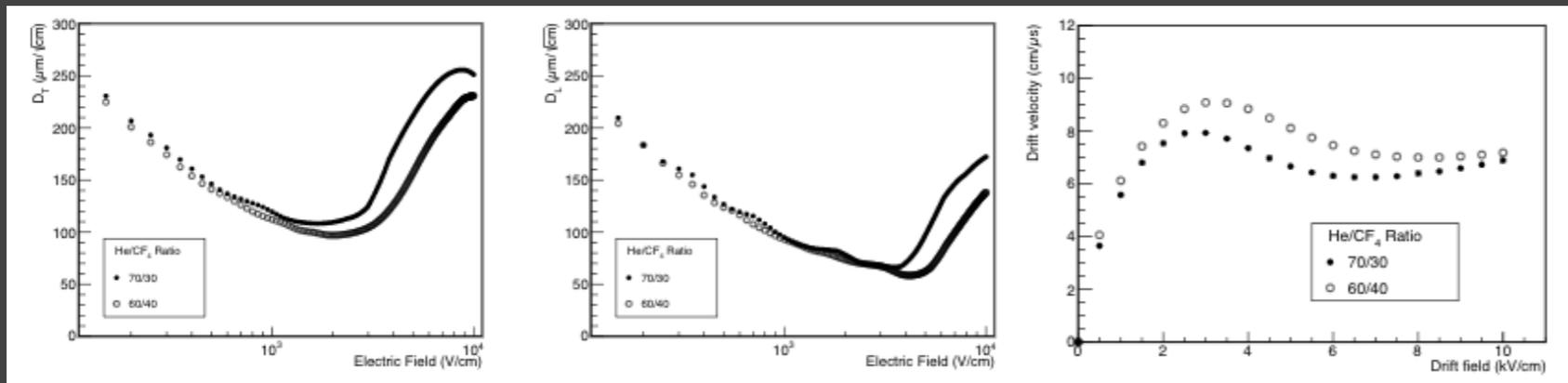
Some study already started about radioactive source to be used. Any advice?

# SIGNAL SIMULATION

Interactions of recoils in gas simulated: GEANT for electrons and SRIM for nuclei;



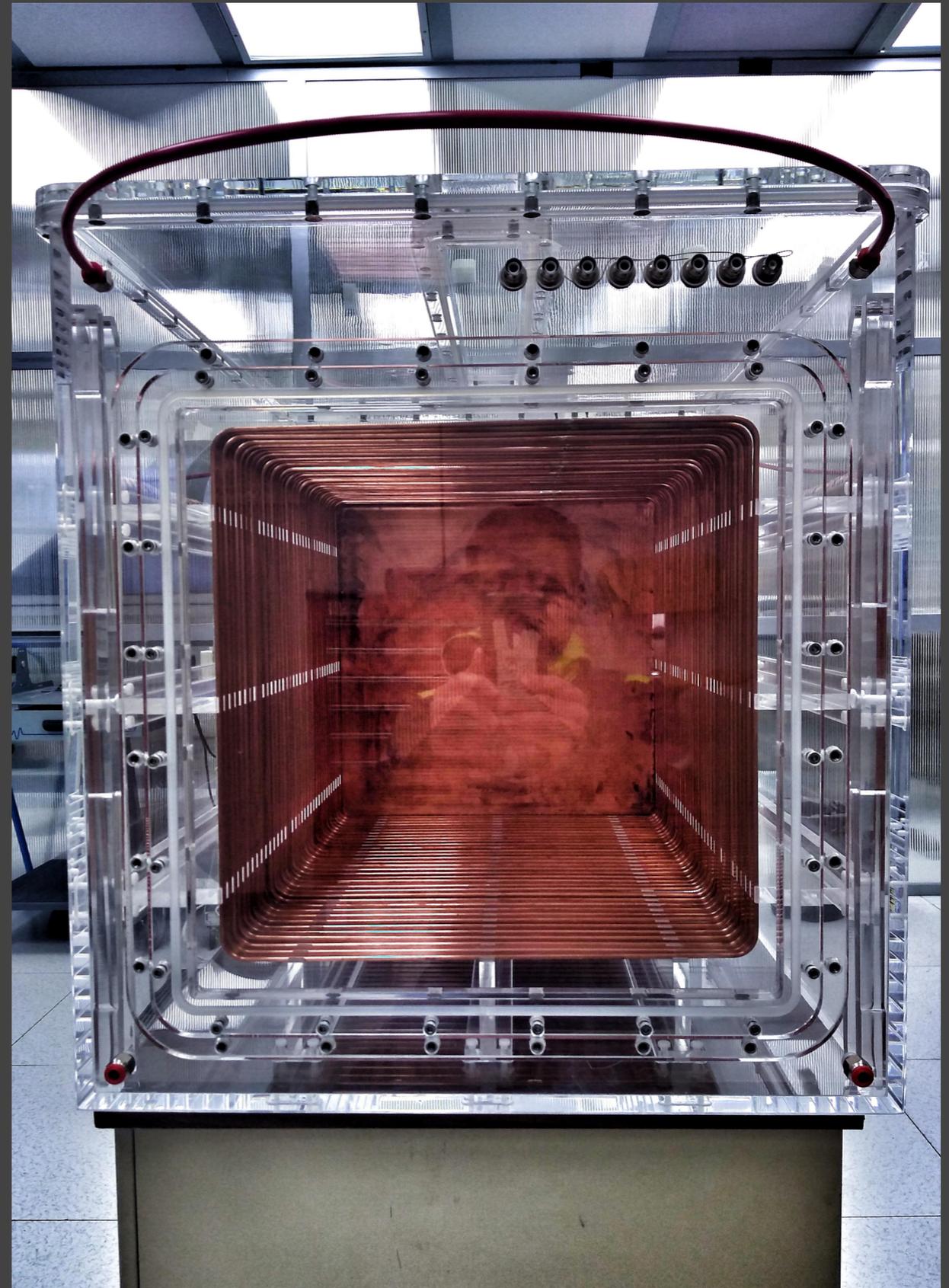
Gas parameters (ionization, diffusion, drift velocity, attachment) simulated with GARFIELD;



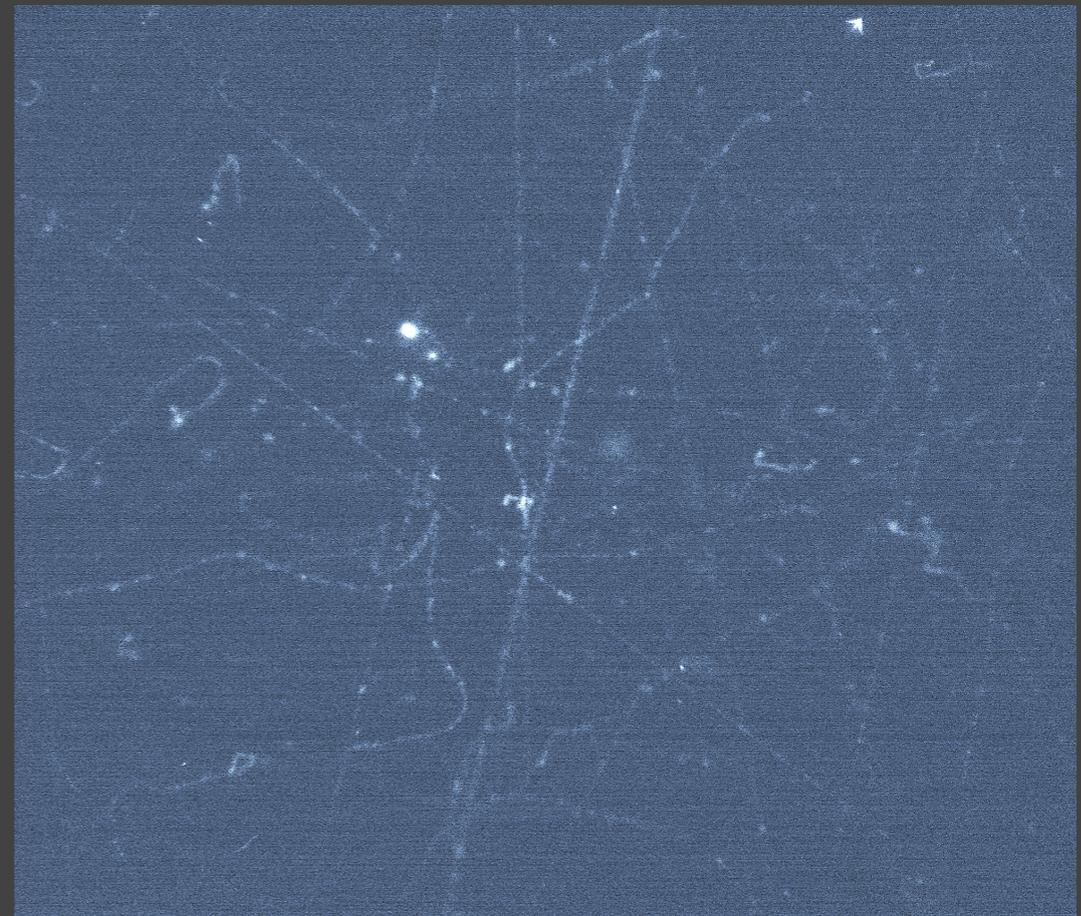
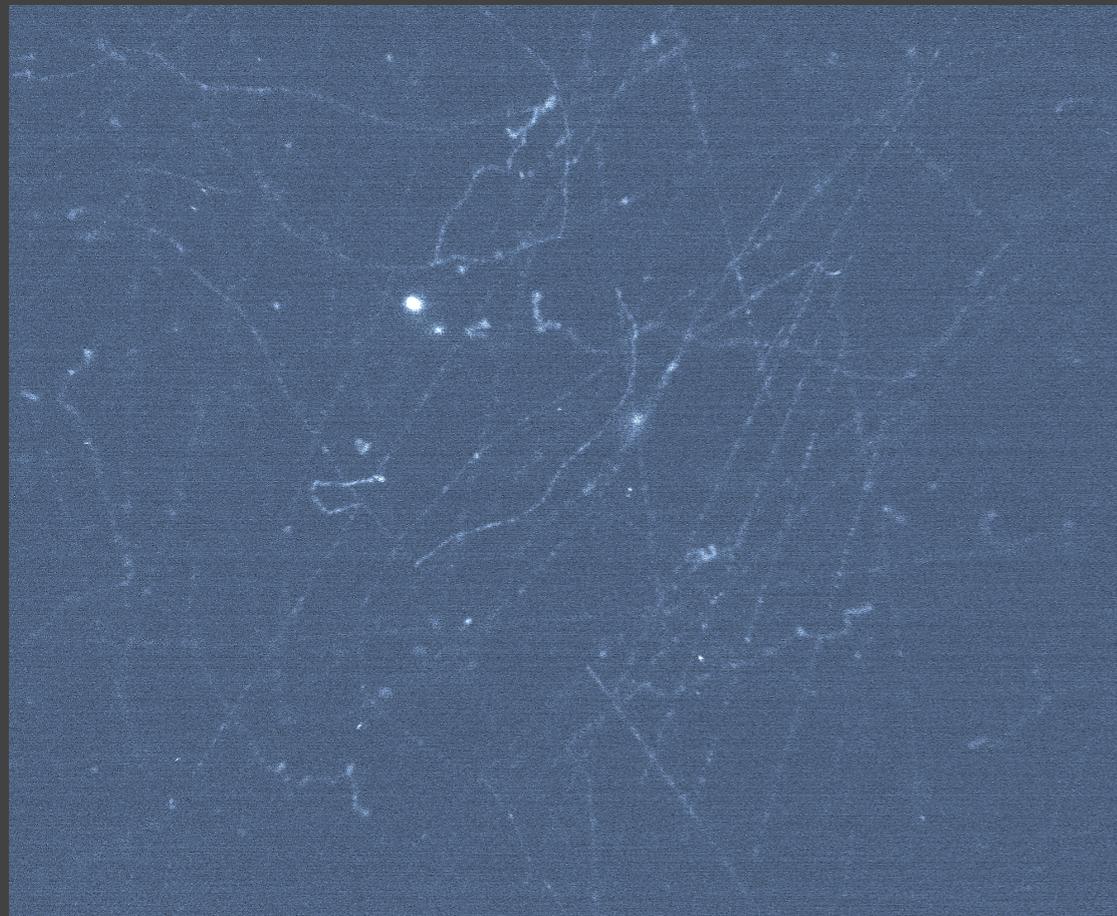
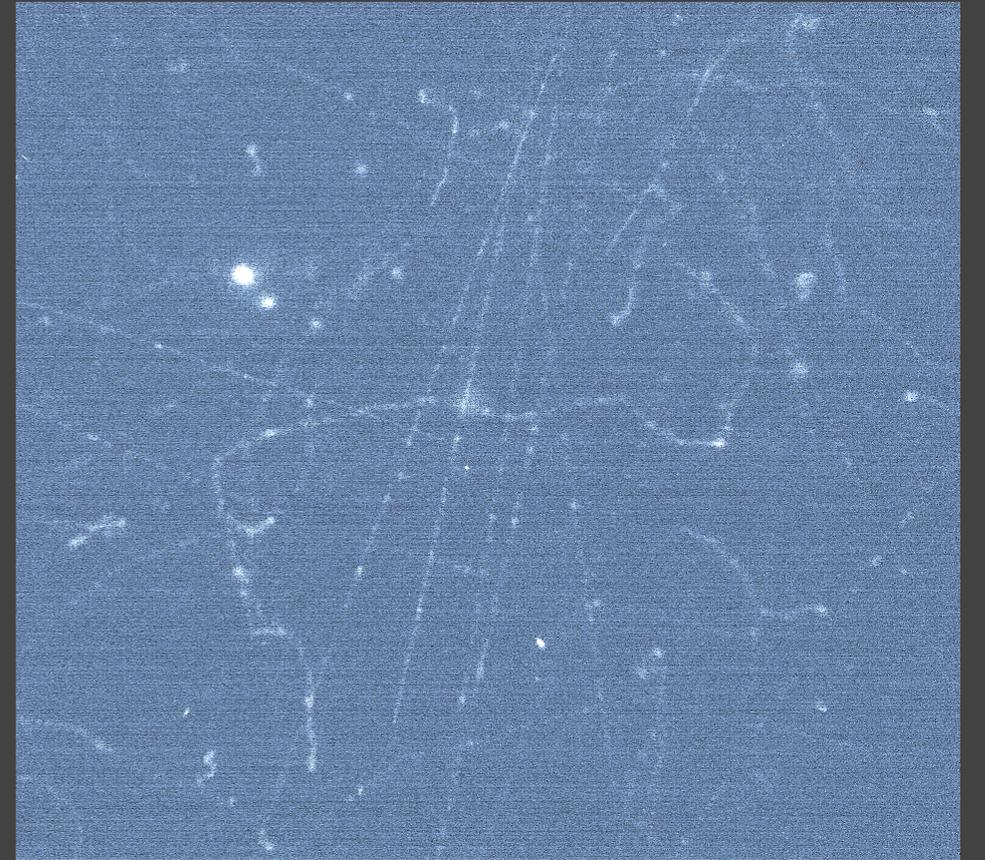
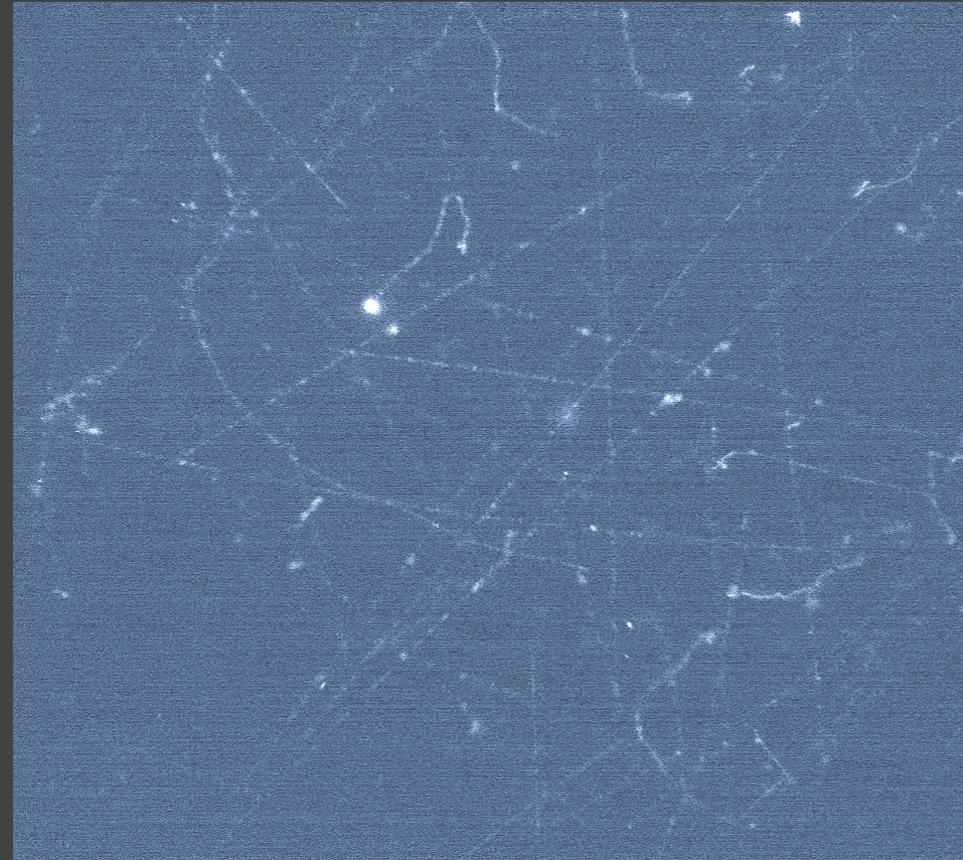
Response and noise behavior of PMT and CMOS simulation on-going

# LIME

50 litre sensitive  
volume



# LIME

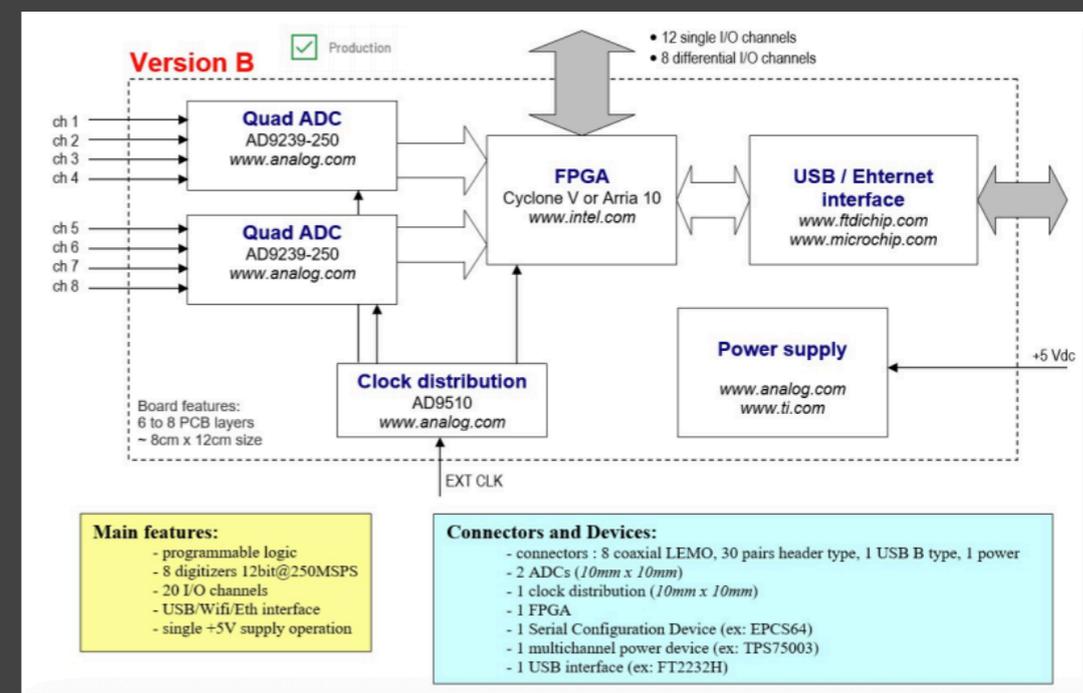
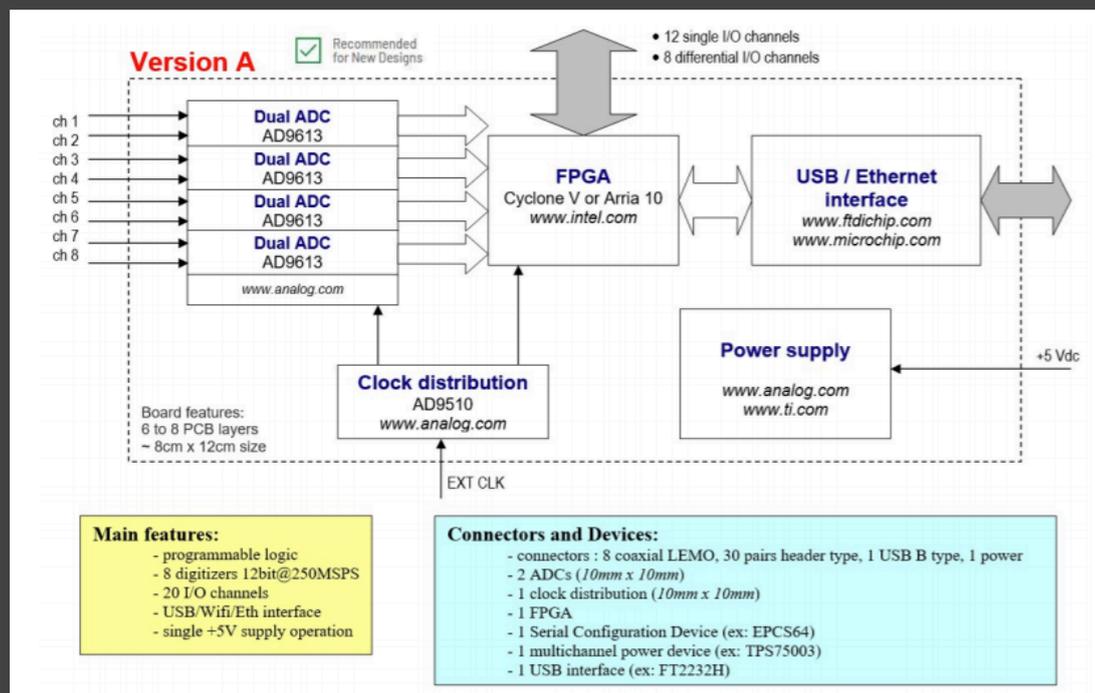


# DAQ

Data acquired so far with a DAQ prototype able to synchronise CMOS and PMT;

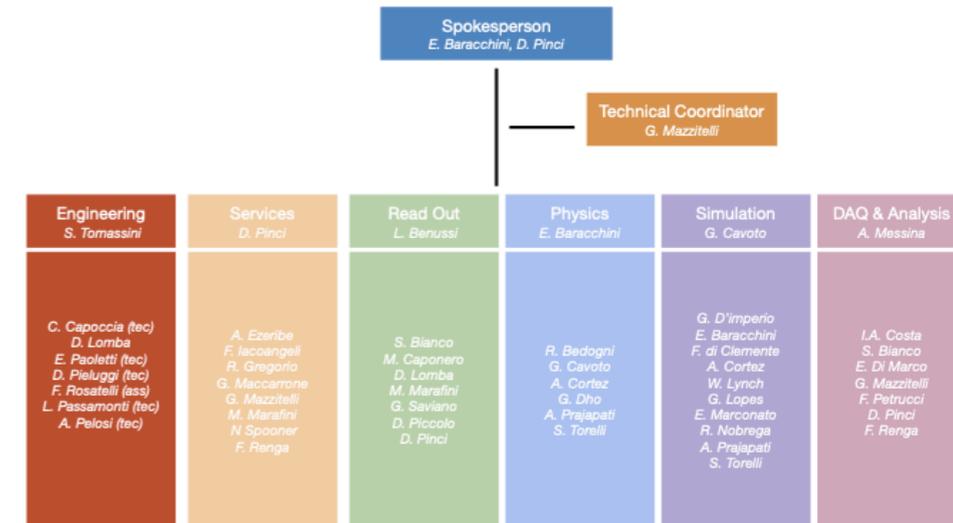
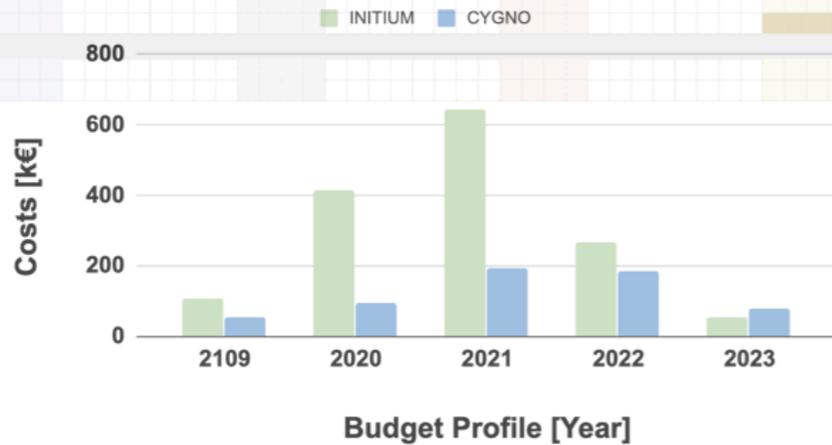
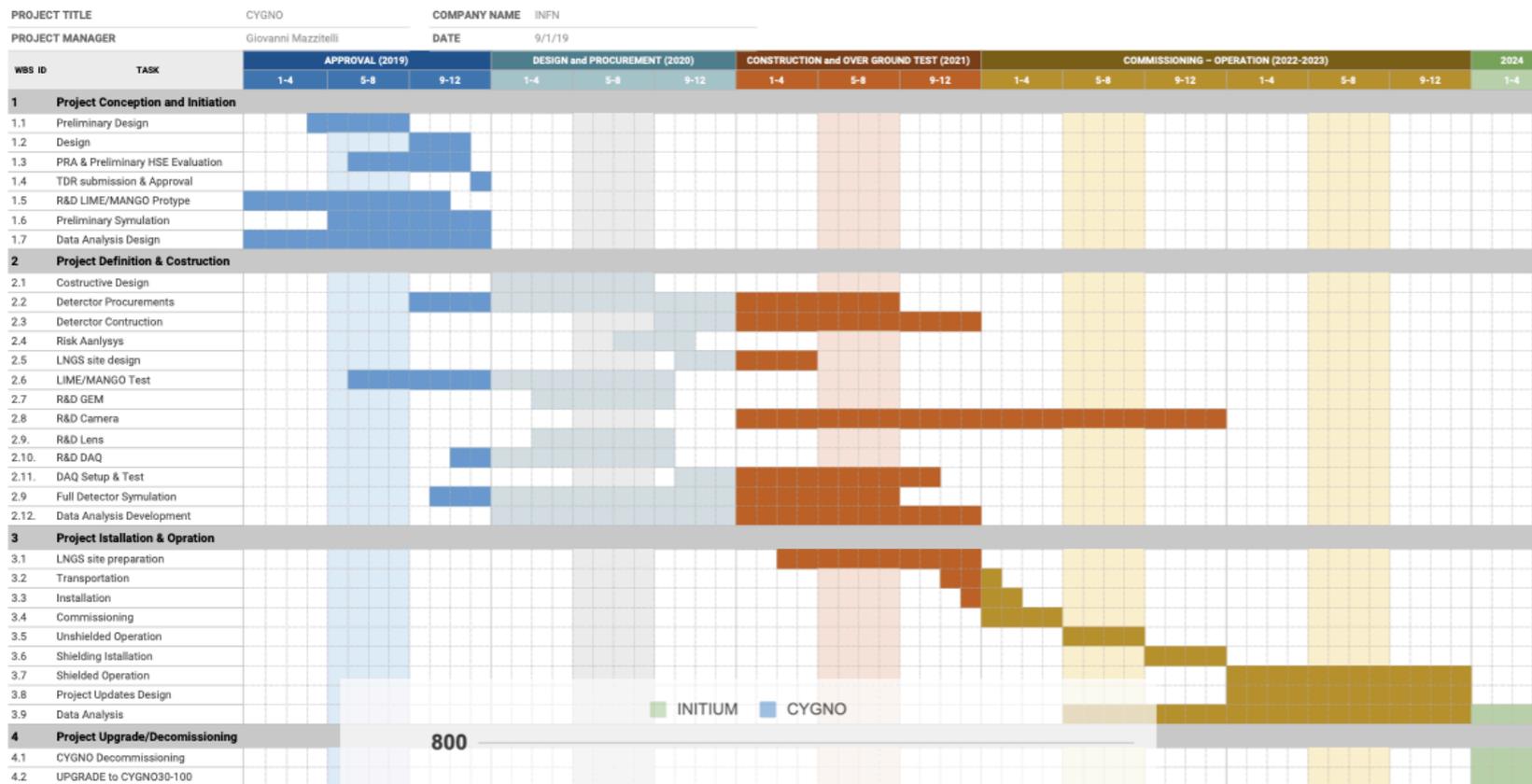
Need to design and assemble a reliable DAQ system, running underground for several CMOS and PMT;

Brazilian group is working on this.



# CYGN0 Schedule

## GANTT



Analisi	Emanuele DM
Gas	Francesco R
HV	Davide Pinci
GEM	Luigi B
CMOS	Davide Pinci
PMT - SiPM	Francesco I
Ottica (lenti, specchi)	Davide Pinci
Progettazione Mecc	Sandro T
Integrazione	Giovanni M
DAQ	Andrea M
Fisica	Elisabetta B
Simulazione	Giulia D'I
Rapp. Internazionali	Elisabetta B
Calibrazione	Gianluca C
Eco gas	Davide Piccolo

G. Mazzitelli-LNF-INFN, UFJF, Oct. 2019

