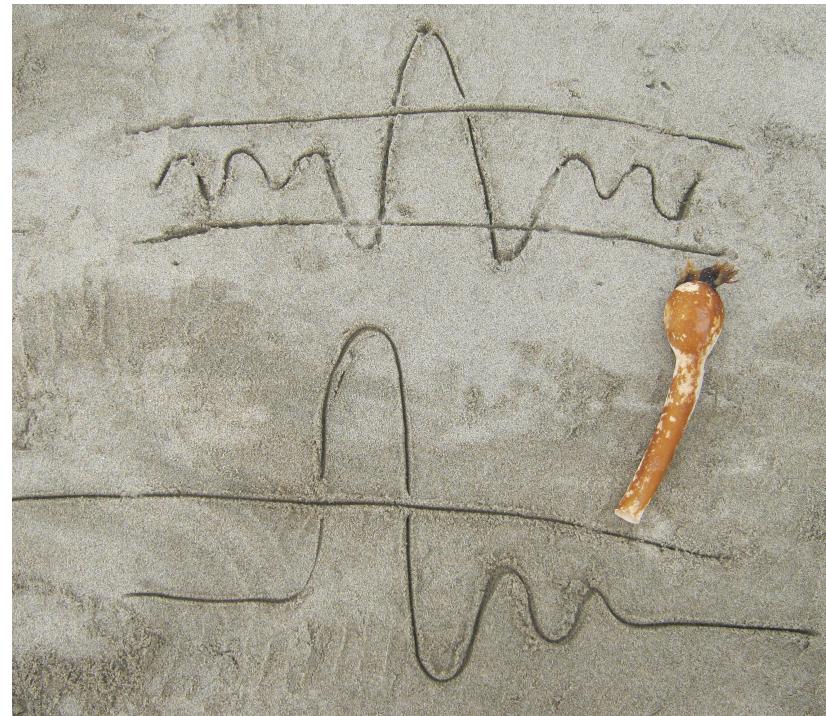


# First results from fTOF prototype test at SLAC CRT

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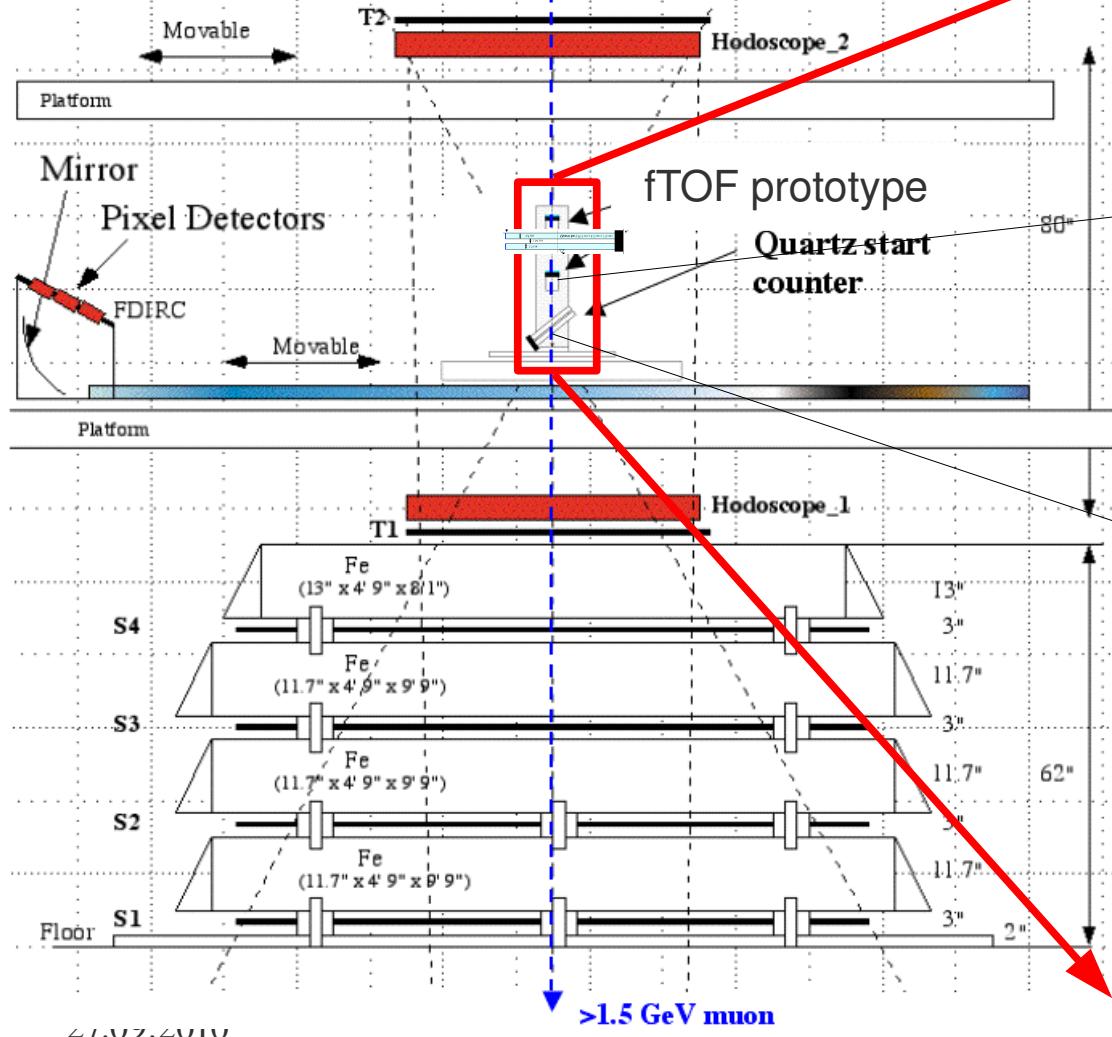


# Outline

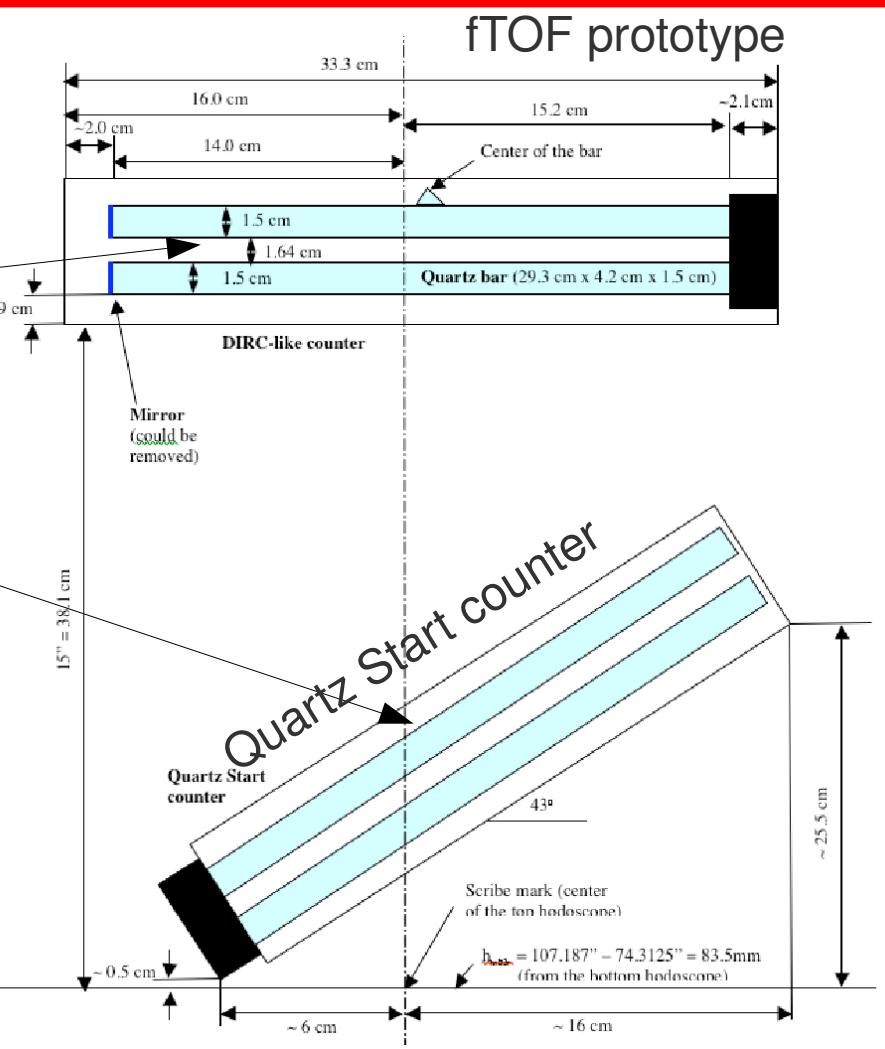
- Experimental Setup
  - DAQ
  - Time synchronization monitoring
  - Run with laser
  - Calibration of the system without laser
  - Run with cosmic muon  
results **without data from CRT**
-

# SLAC Cosmic ray telescope

Side view



## Experimental Setup(1)



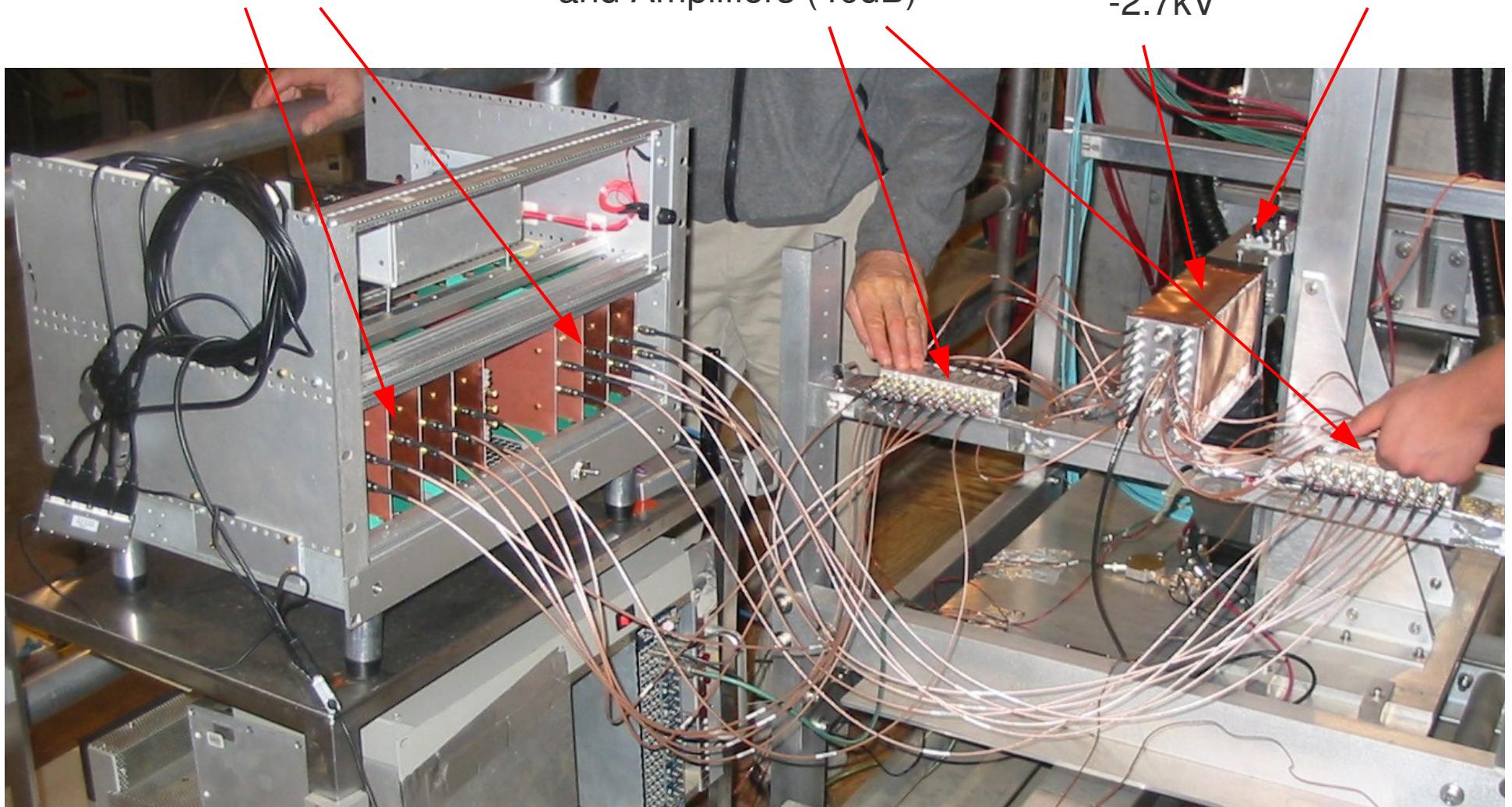
# Experimental Setup(1)

8 USBWC = 16 Channels

Filters (600MHz bandwidth )  
and Amplifiers (40dB)

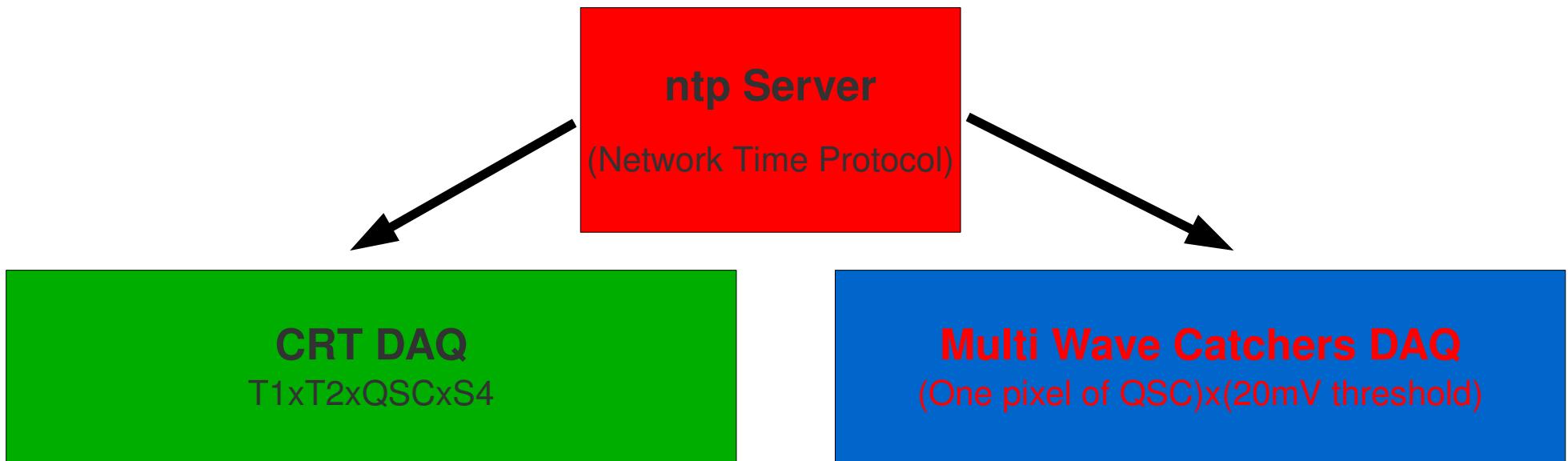
MCP-PMT  
-2.7kV

Quartz Bars



# DAQ

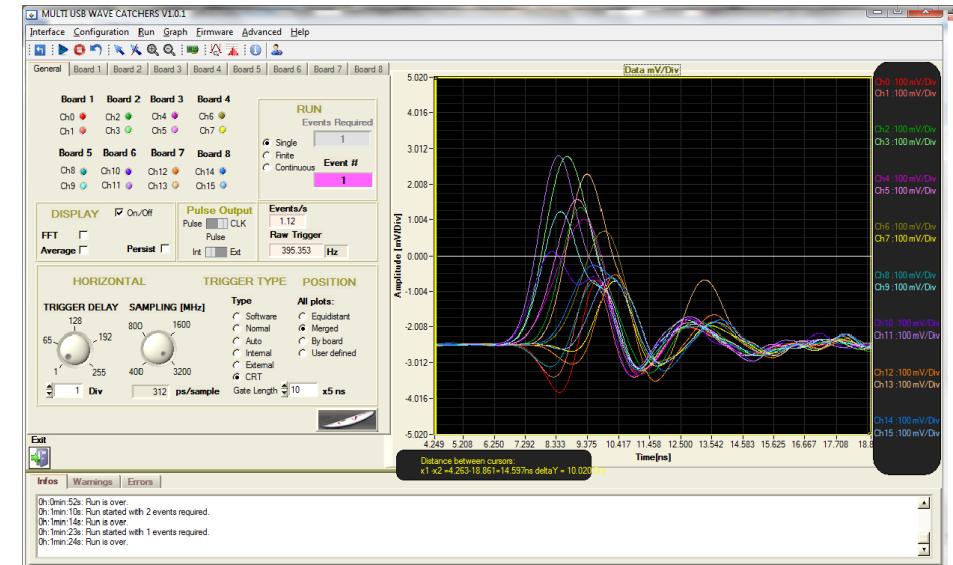
CRT and Multi Wave Catchers DAQ are running independently. It is possible to merge them in time. For this we need precise time mark.



Data rate in CRT 0.15 event/s

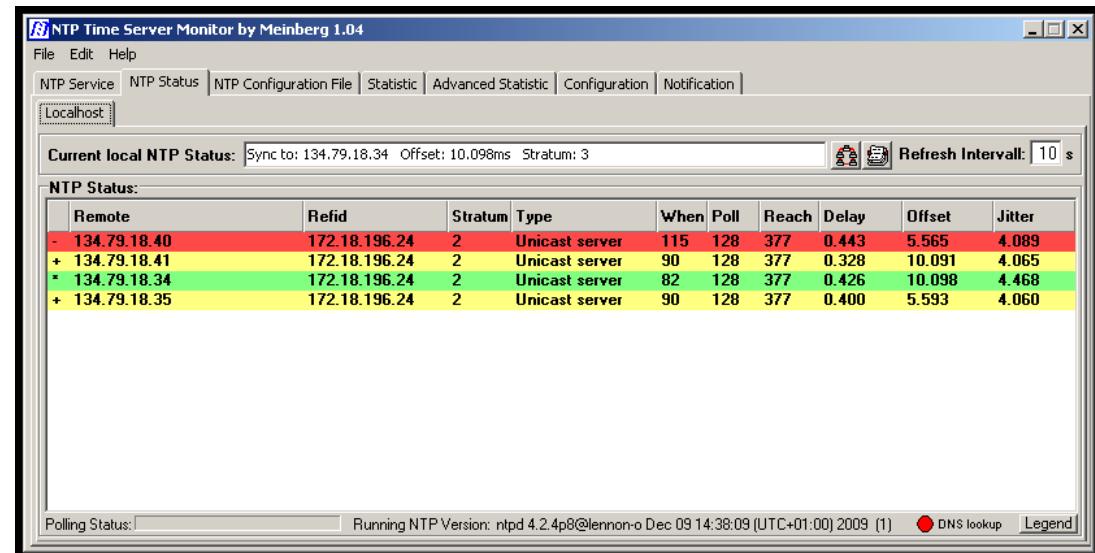
Data rate in USBWC 0.1event/s

dst file

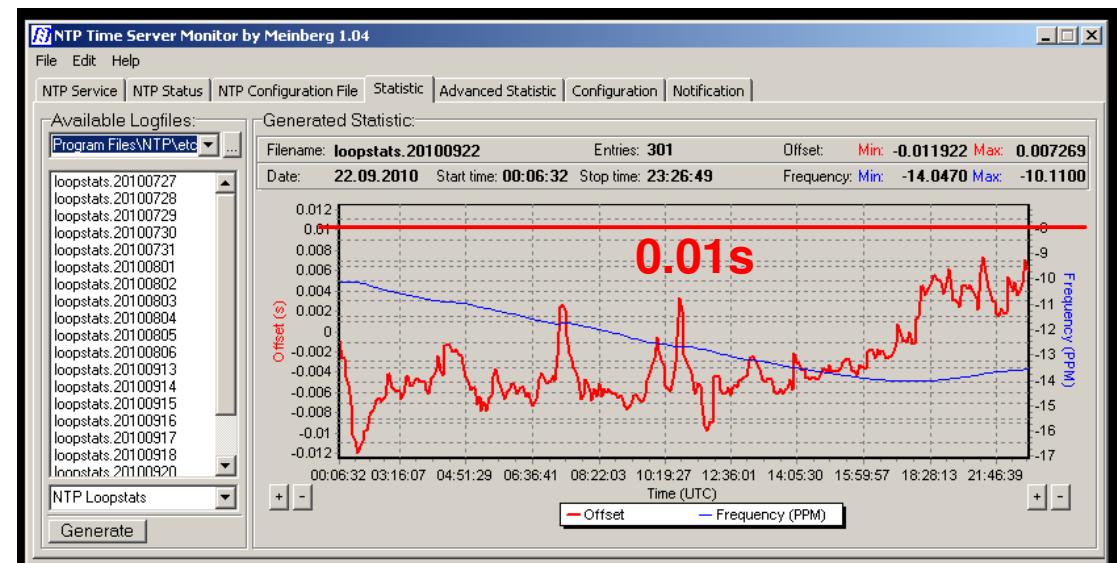


# Time synchronization monitoring

- Time can be updating from 4 different ntp servers.



- Log information

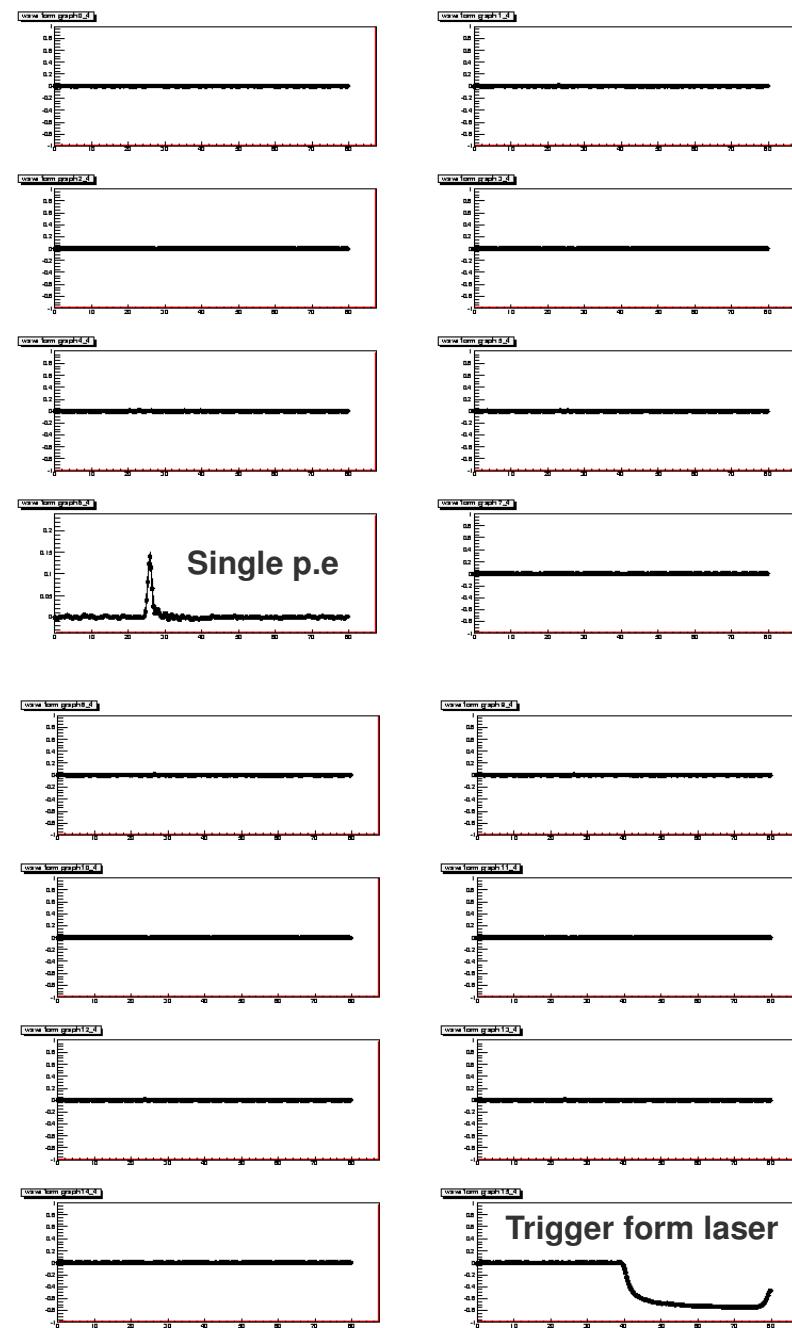
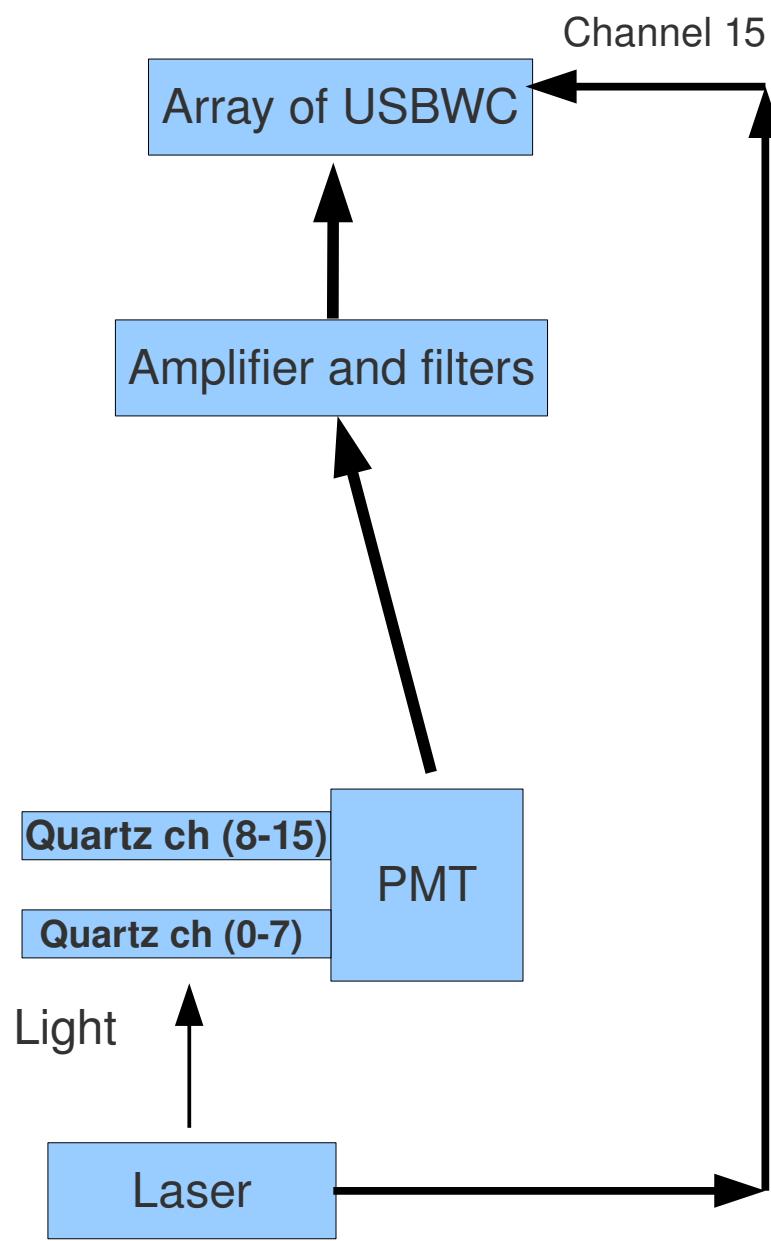


As I was told CRT DAQ have same precision

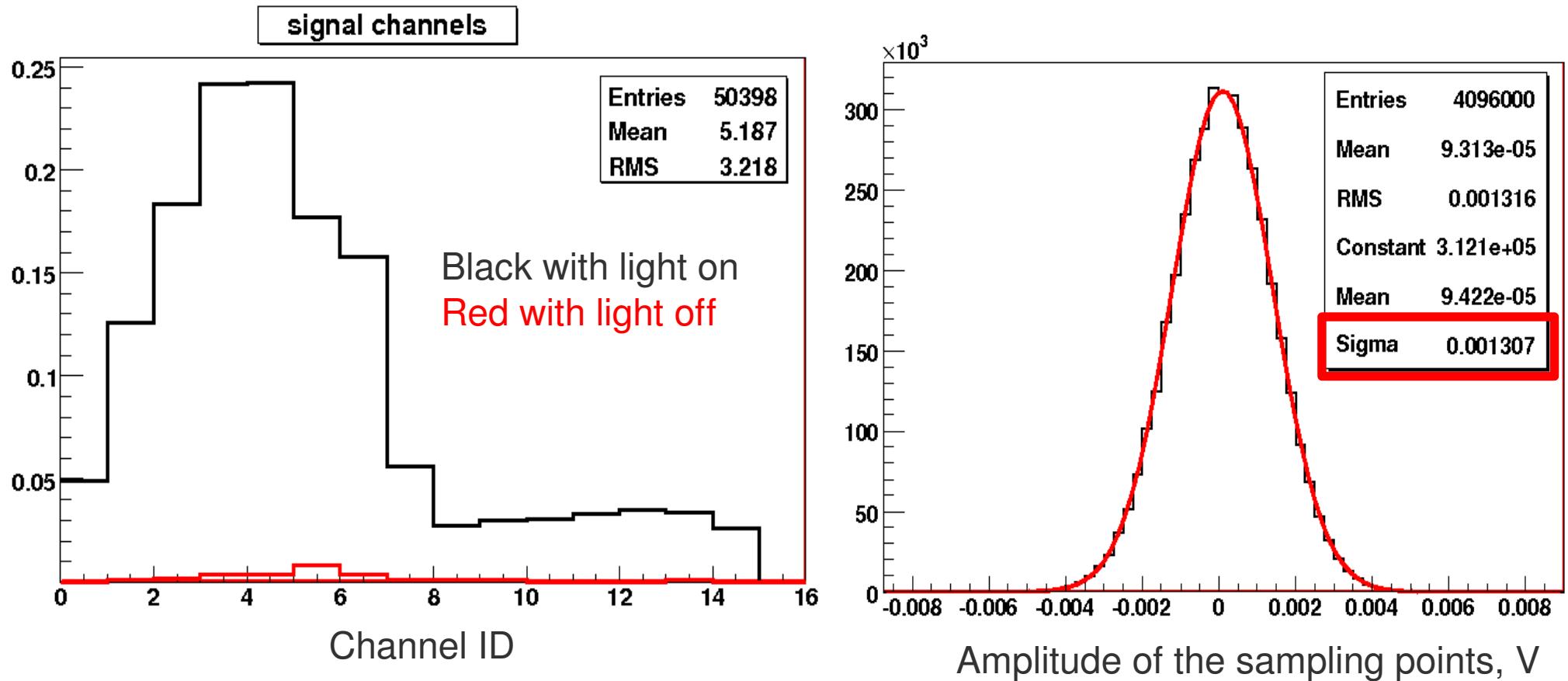
Monitoring of the system time updating is very important since we need to be synchronized with CRT DAQ

# **First results from fTOF prototype test at SLAC CRT**

# Run with laser(0)



# Run with laser(1)

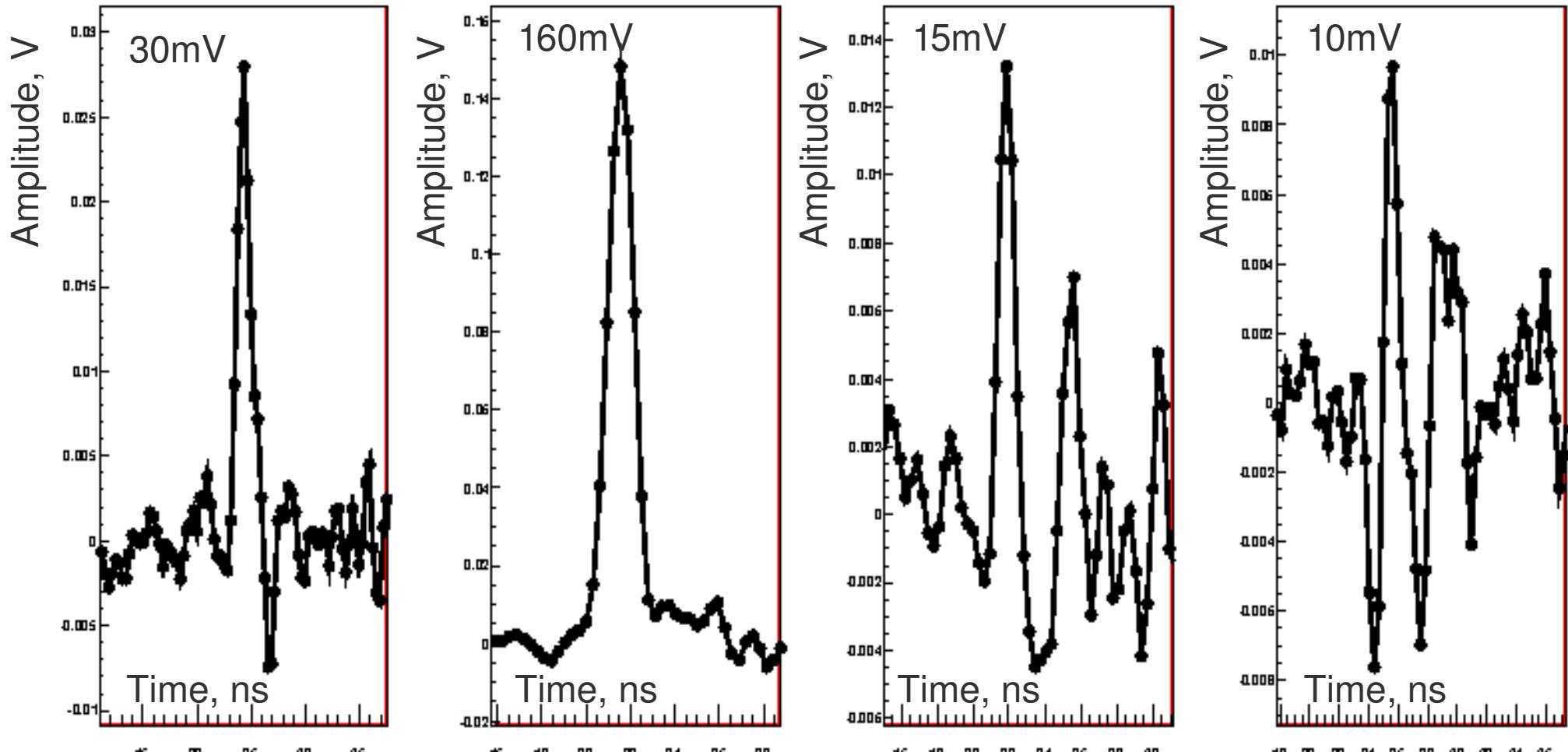


Histograms normalized by number of events

We select signals with amplitude > 50mV

RMS of the noise 1.3mV

# Run with laser(2), crosstalk and charge sharing



Charge sharing

● Signal amplitude  
more than 50mV  
(40 noise RMS)

27.09.2010

Single p.e signal

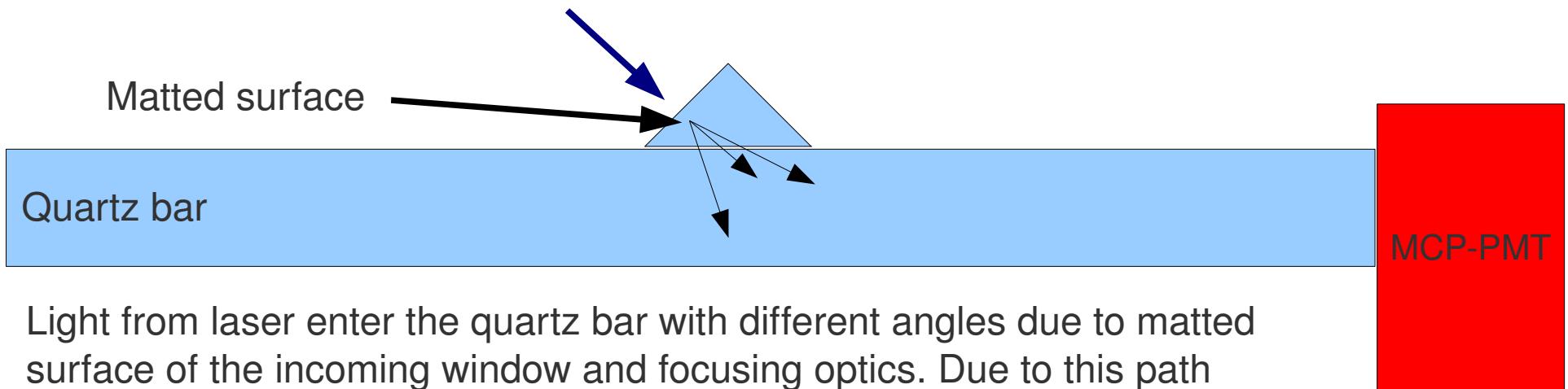
● Charge sharing  
amplitude is changing a  
lot , average (10-20)mV  
(8 noise RMS)

Charge sharing

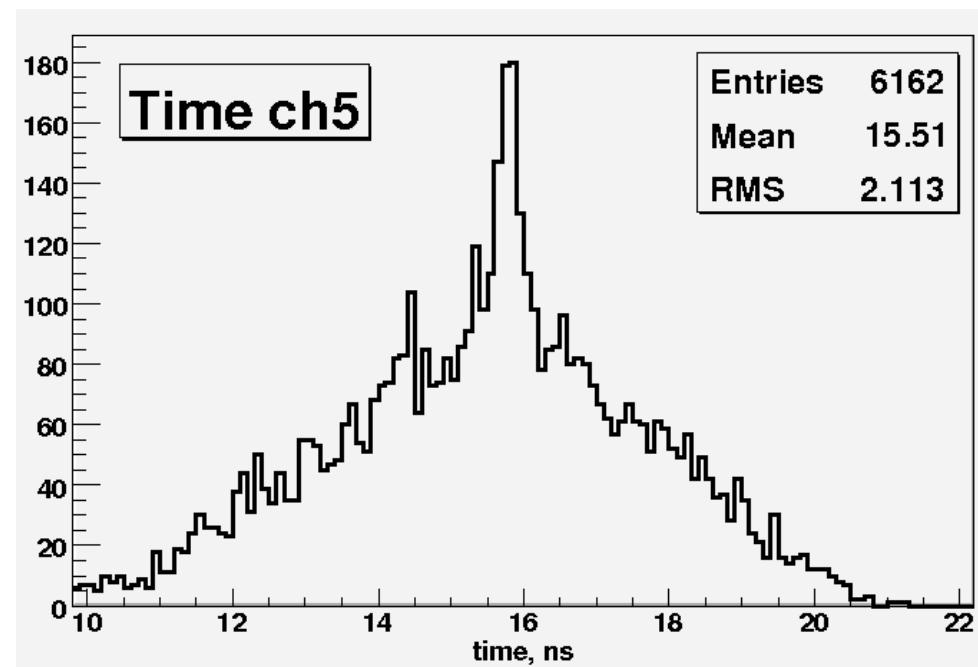
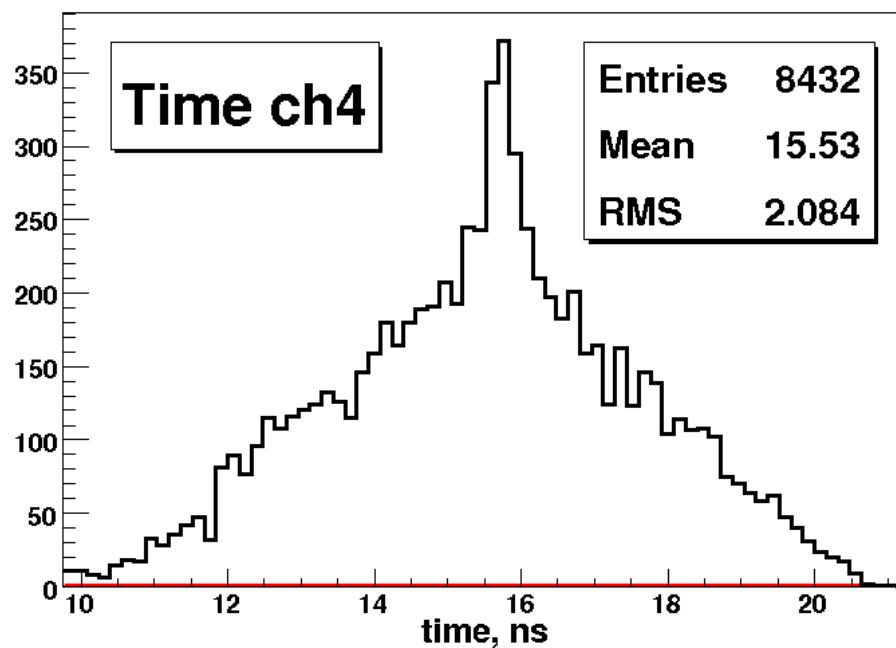
● Crosstalk  
amplitude is  
around 6-8 mv  
(6 noise RMS)

10

## Run with laser(3), time distribution

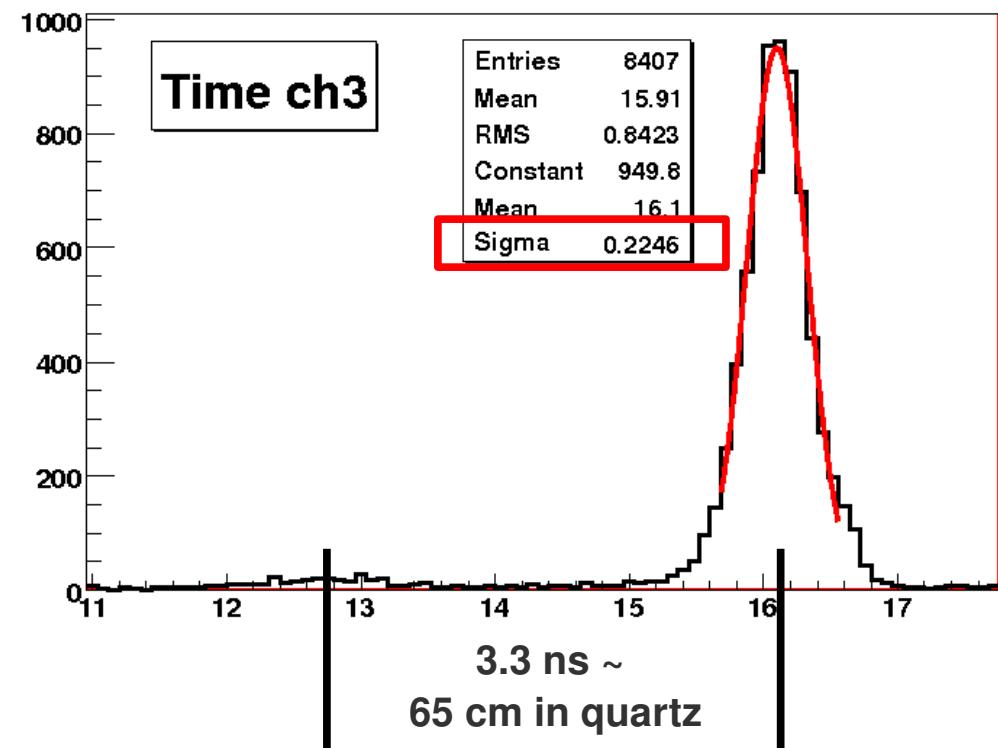
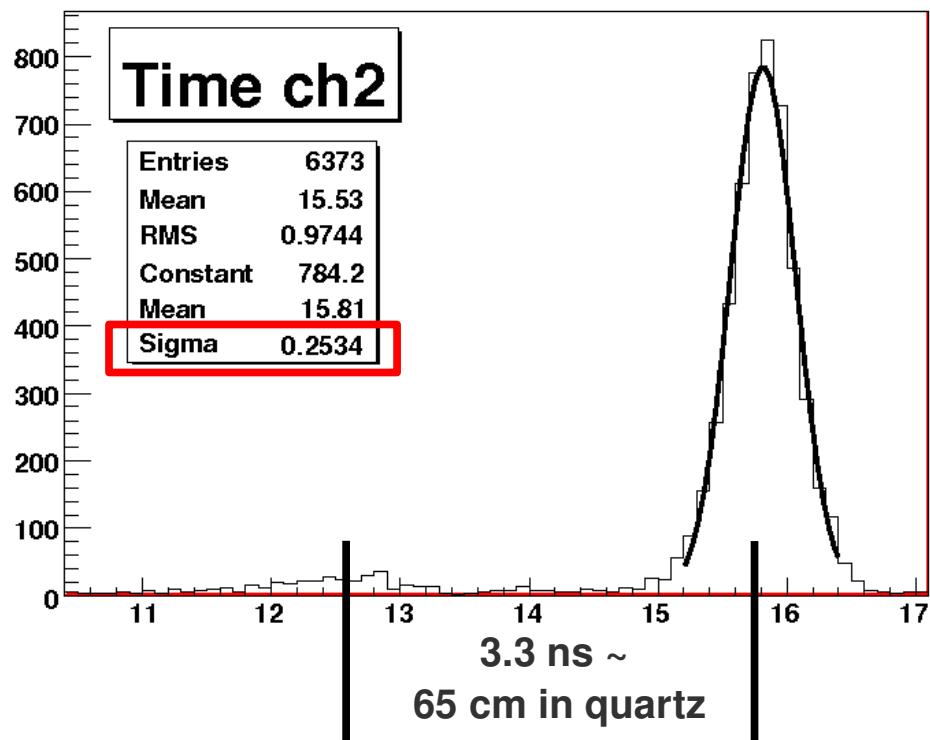


Light from laser enter the quartz bar with different angles due to matted surface of the incoming window and focusing optics. Due to this path propagation of the light touching given channel is very different. So the time difference between trigger and signal from given channel have very wide distribution.



# Run with laser(4), time distribution

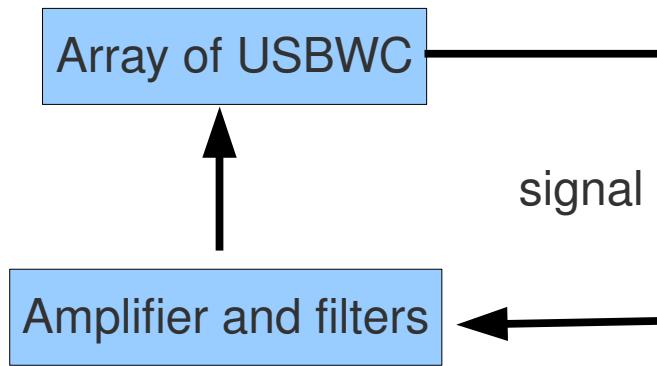
But some channels have not very wide distributions.



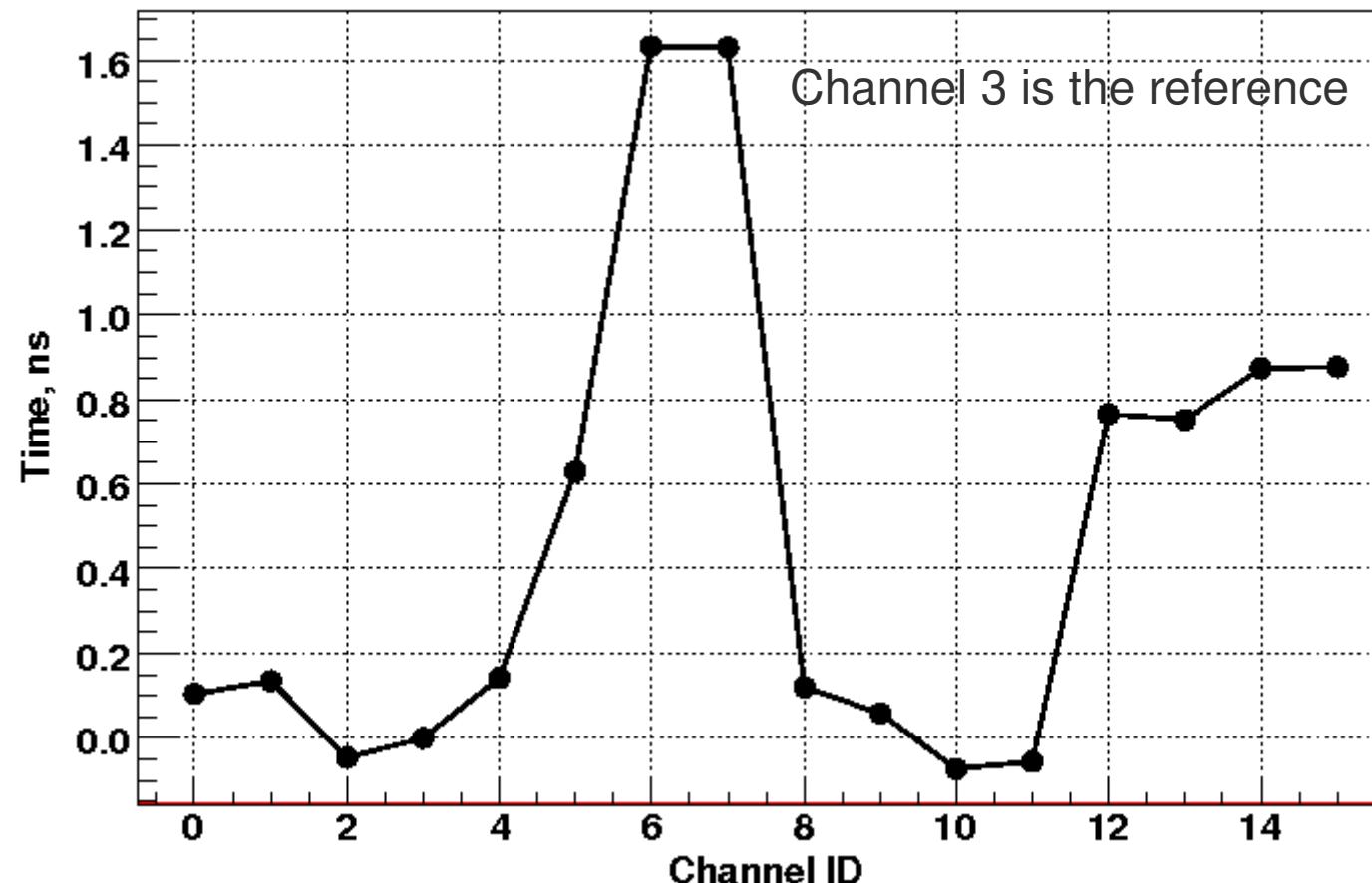
This can be scattered back from the mirror photons

# Calibration of the system without laser

Each channel has his own very stable bias. If we want to know which signal was first we need to have a map of this constant biases.

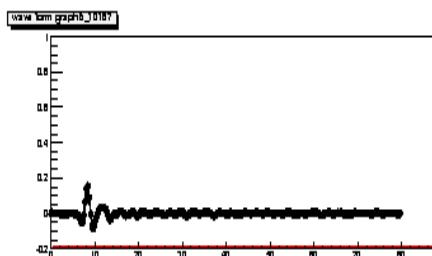
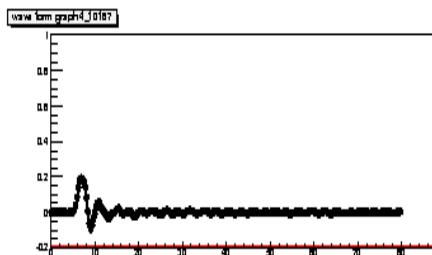
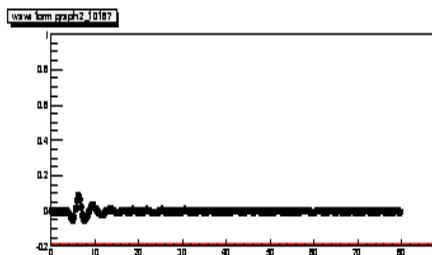
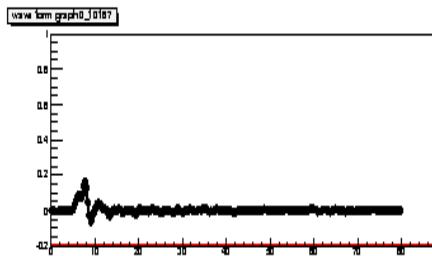


With this technique all system started from amplifiers and filters can be calibrated

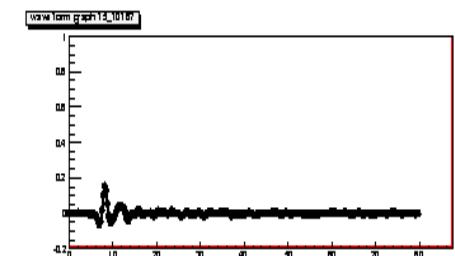
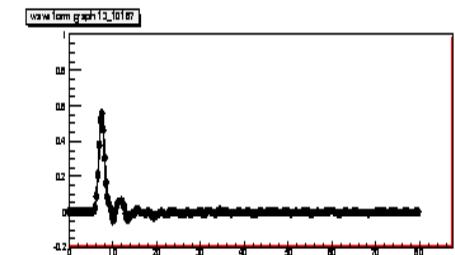
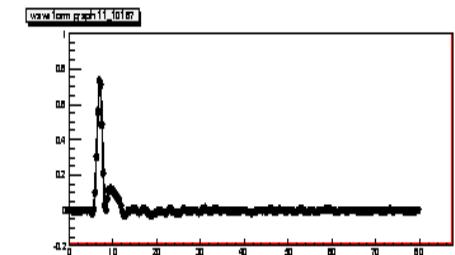
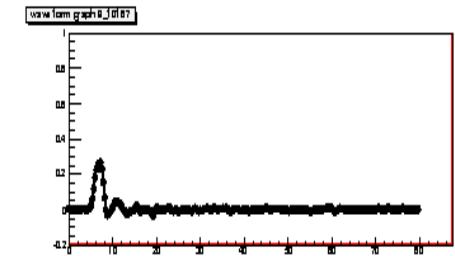
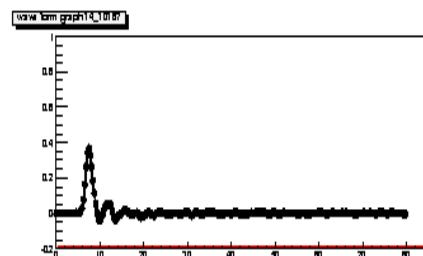
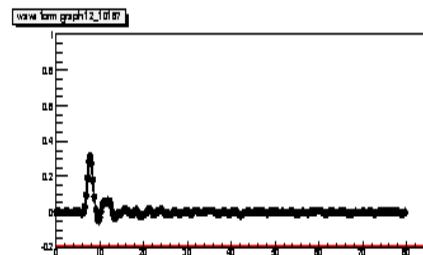
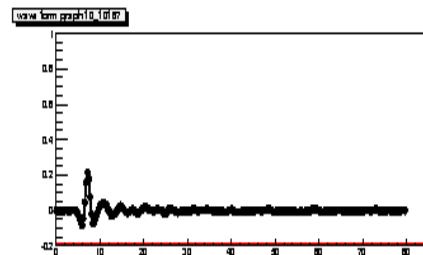
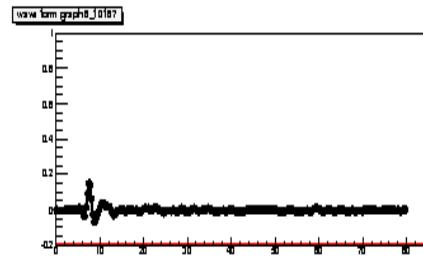


# Run with cosmic muons(0)

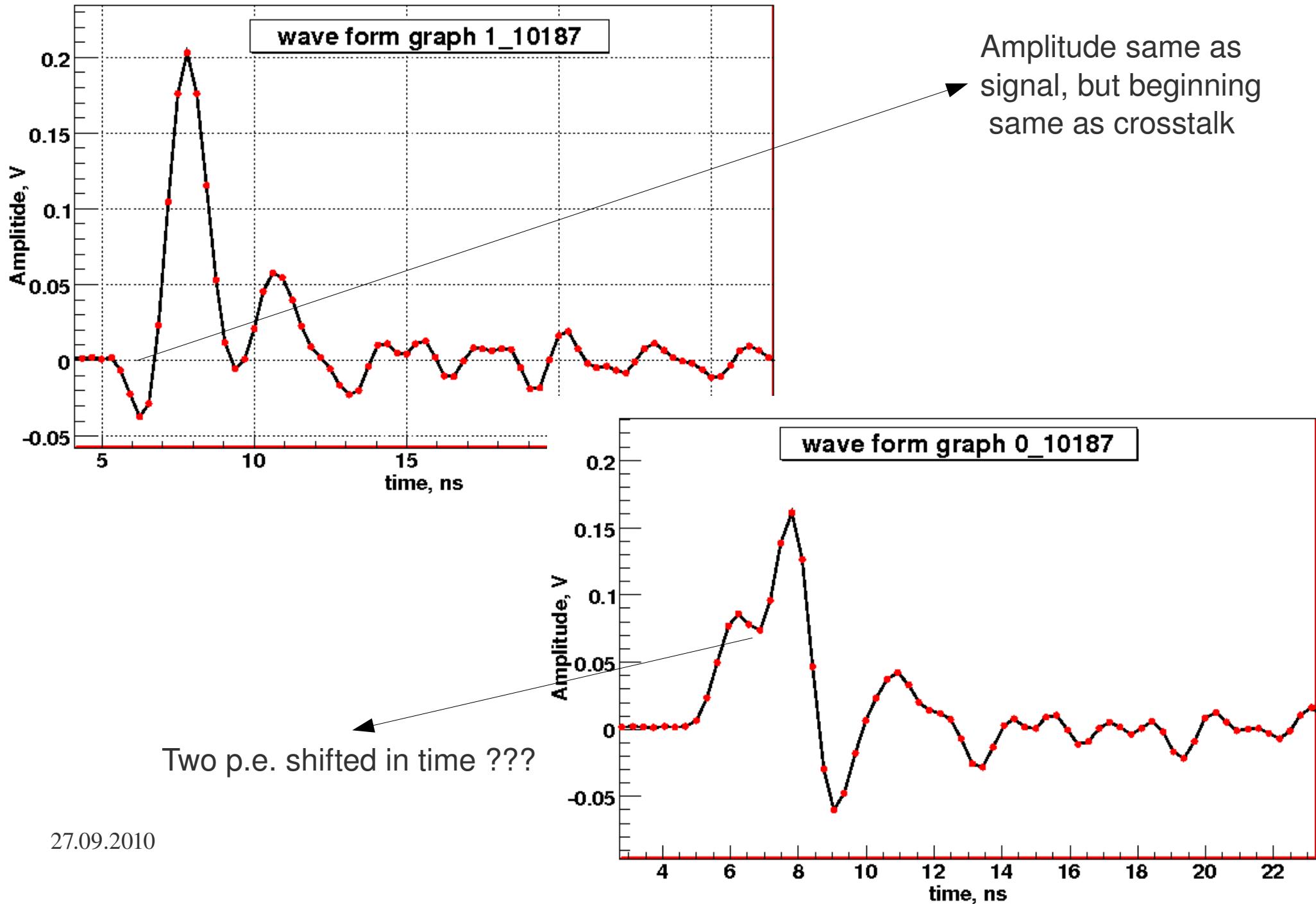
0 to 7



8 to 15

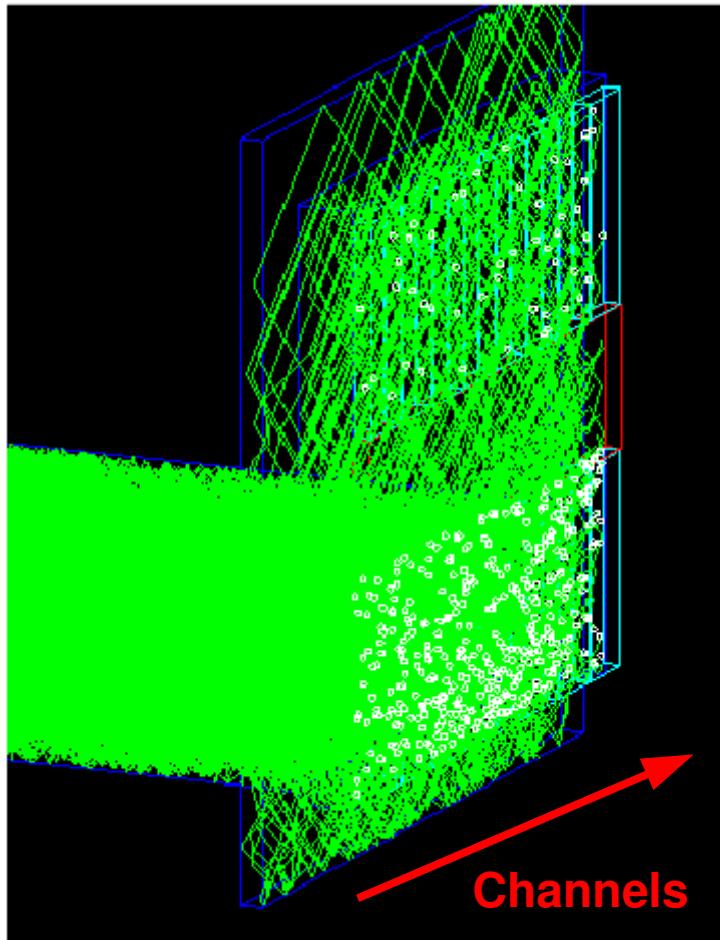


# Run with cosmic muons(1)

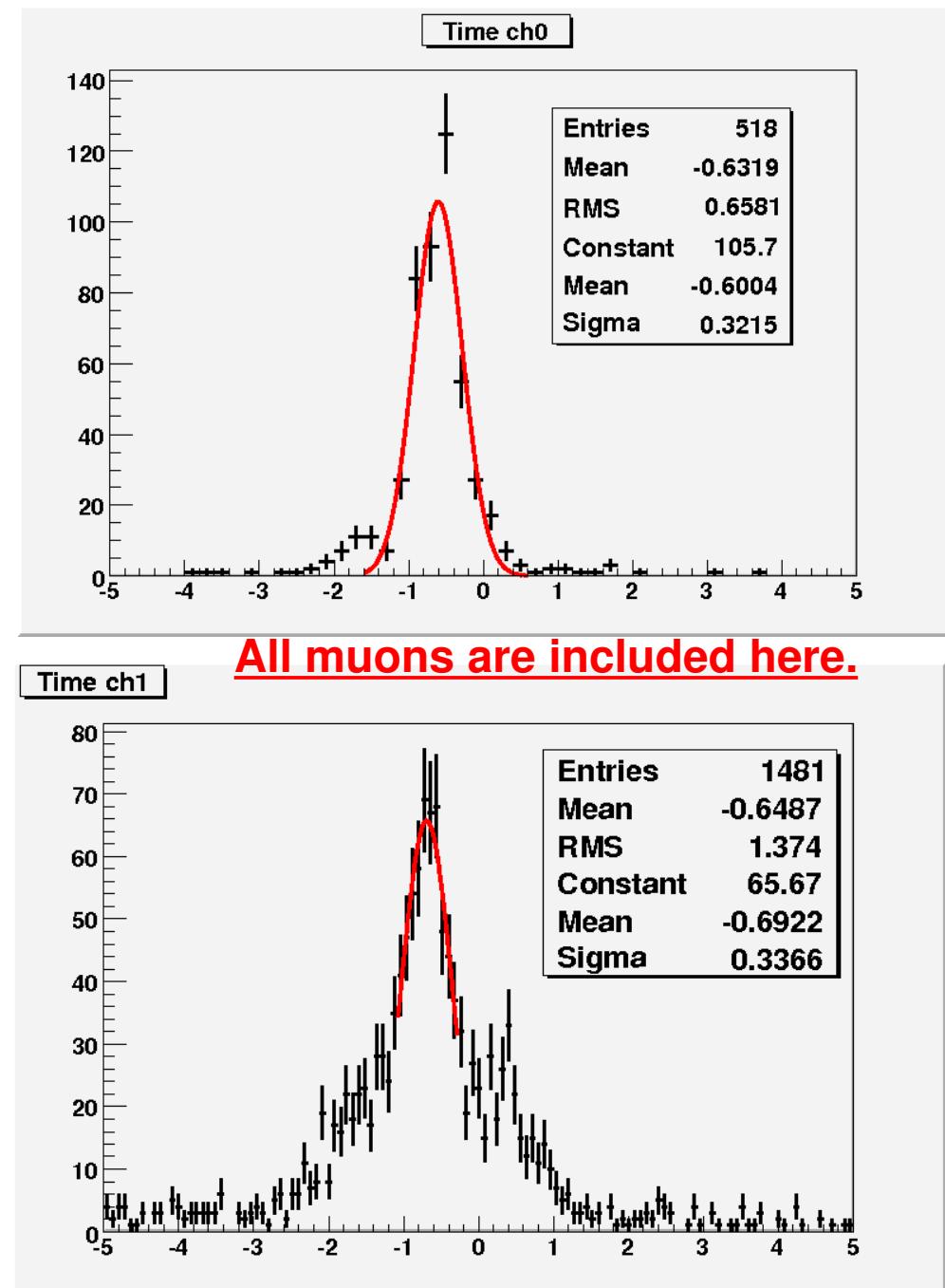


# Run with cosmic muon(2)

We substruct  $\text{ch}_0 - \text{ch}_8; \text{ch}_1 - \text{ch}_9 \dots$



**NOTE:**  
Information from  
CRT does not used



# Conclusion

- 1) fTOF prototype was installed at SLAC CRT. It taking date at present moment.
- 2) Electronics connected, checked and calibrated.
- 3) First analysis of data from laser run has been started, rough information about signal, noise, crosstalk, charge sharing were obtained.
- 4) Analysis of the cosmic muon data has been started

# **BACKUP**