

Firenze, 23 Aprile 2010

# LHCf: un esperimento per la fisica dei raggi cosmici ad LHC

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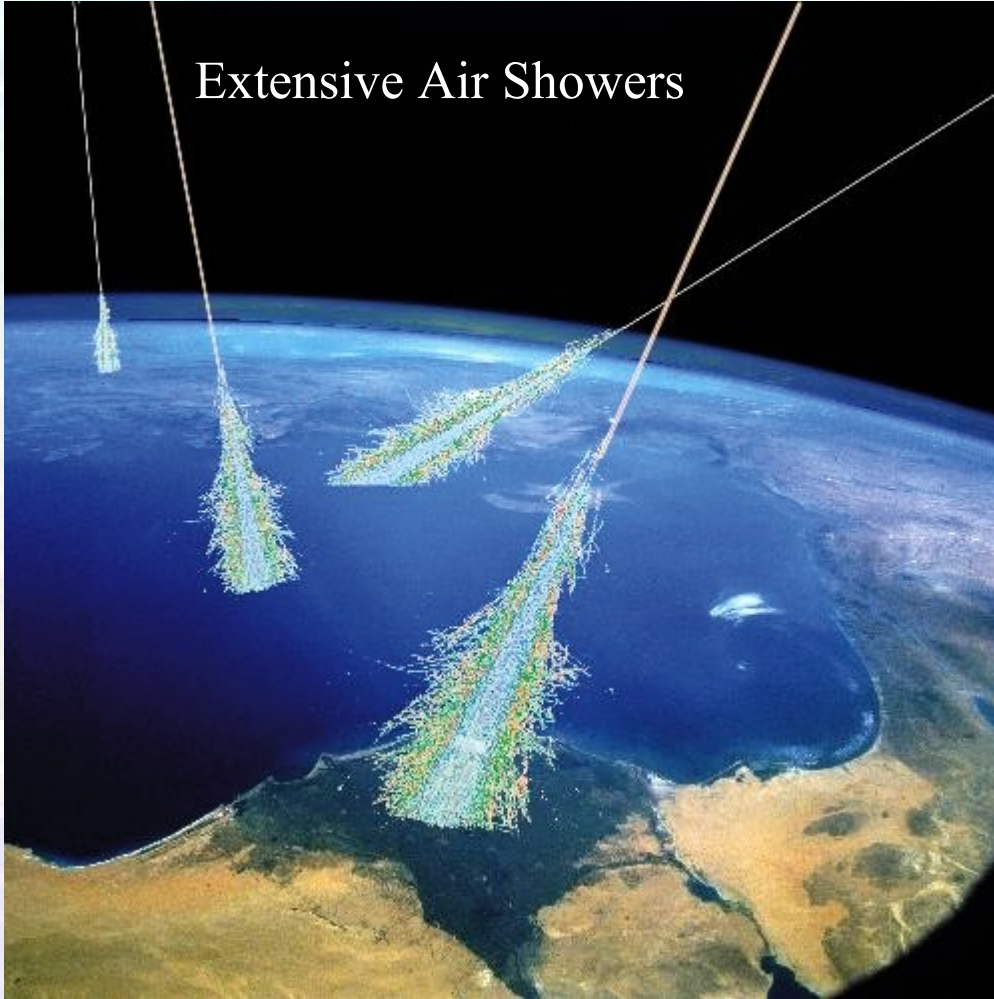


# Why LHCf?

## Physics Motivations

# Ultra High Energy Cosmic Rays

Extensive Air Showers



Experimental observations: at  $E > 100$  TeV only EAS  
(shower of secondary particles)

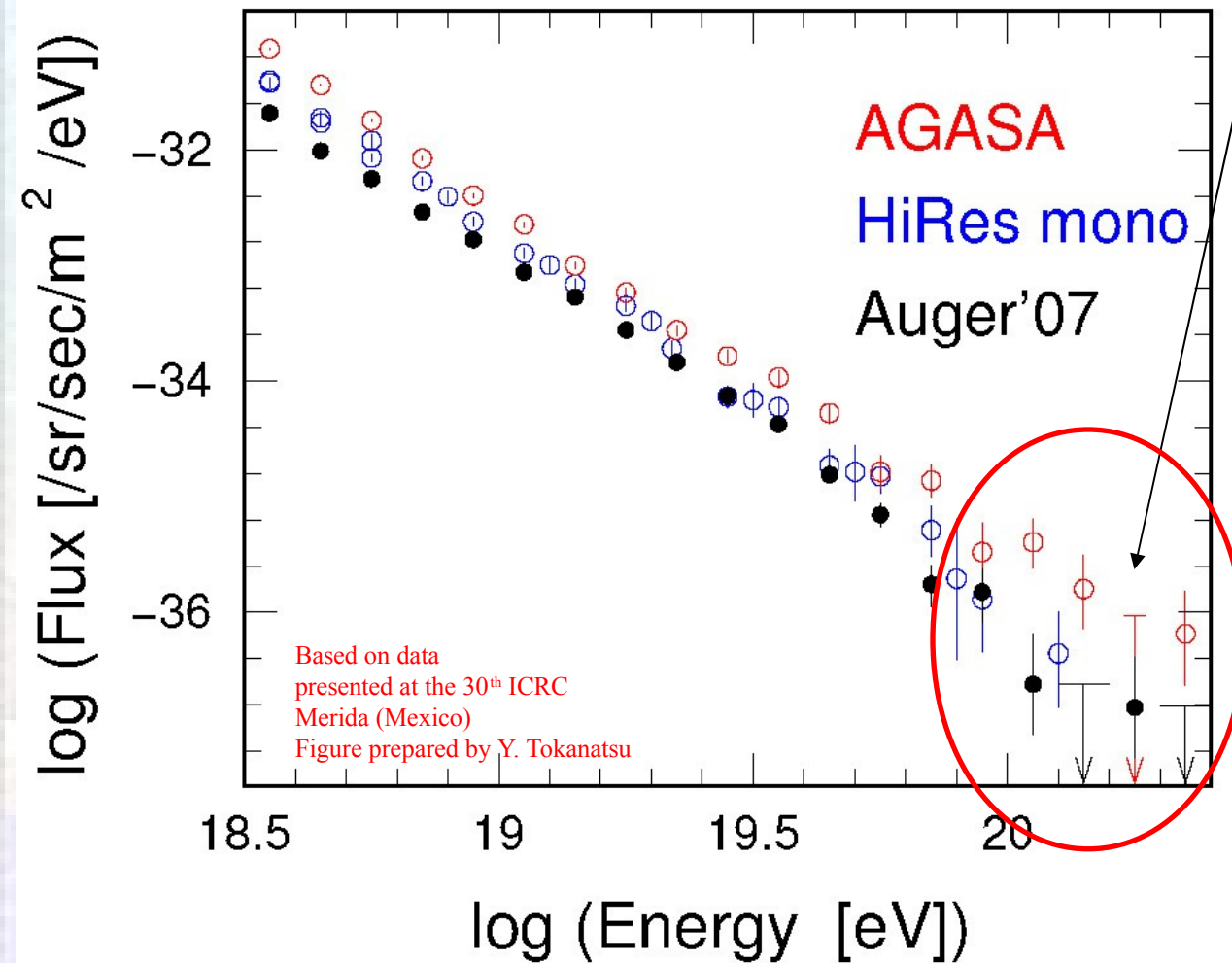
- lateral distribution
- longitudinal distribution
- particle type
- arrival direction

**Air shower development  
(particle interaction in the  
atmosphere)**

Astrophysical parameters:  
(primary particles)

- spectrum
- composition
- source distribution
- origin and propagation

# The Cosmic Ray Spectra



GZK cutoff:  $10^{20}$  eV

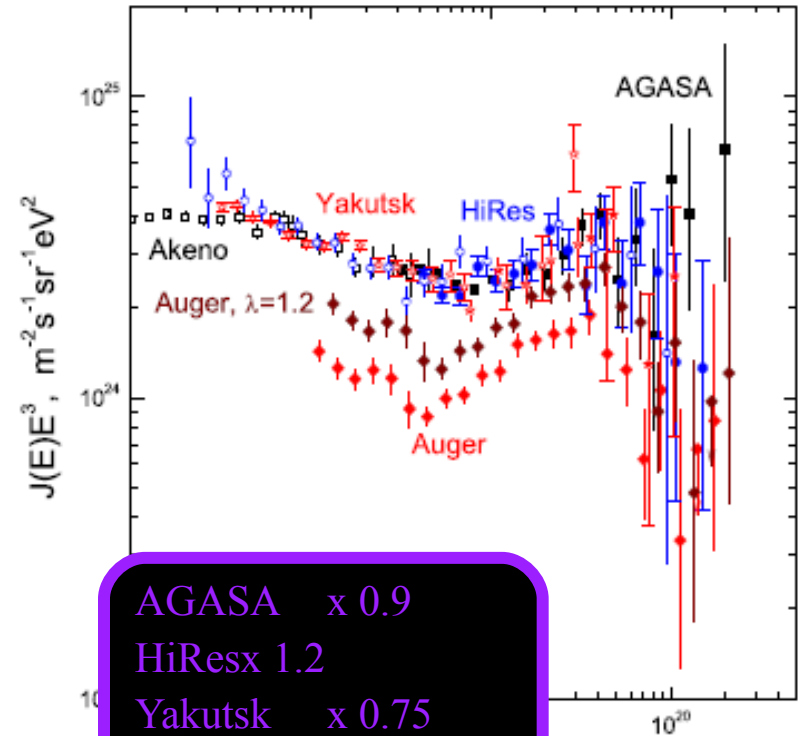
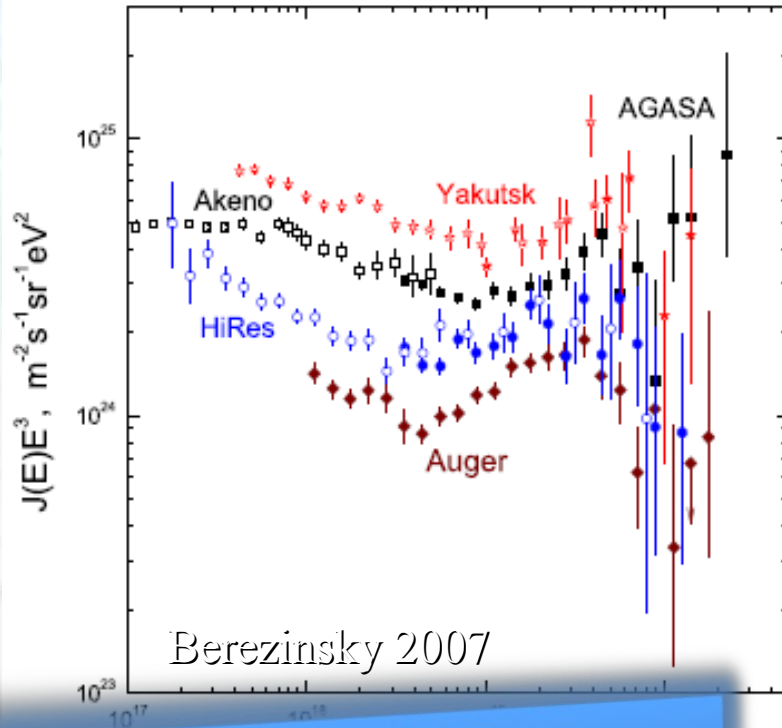
GZK cutoff would limit energy to  $10^{20}$  eV for protons, due to Cosmic Microwave Background

$p\gamma(2.7K) \rightarrow \Delta \rightarrow N\pi$

super GZK  
events?!?

Different results  
between different  
experiments

# The Cosmic Ray Spectra



AGASA x 0.9  
 HiResx 1.2  
 Yakutsk x 0.75  
 Auger x 1.2

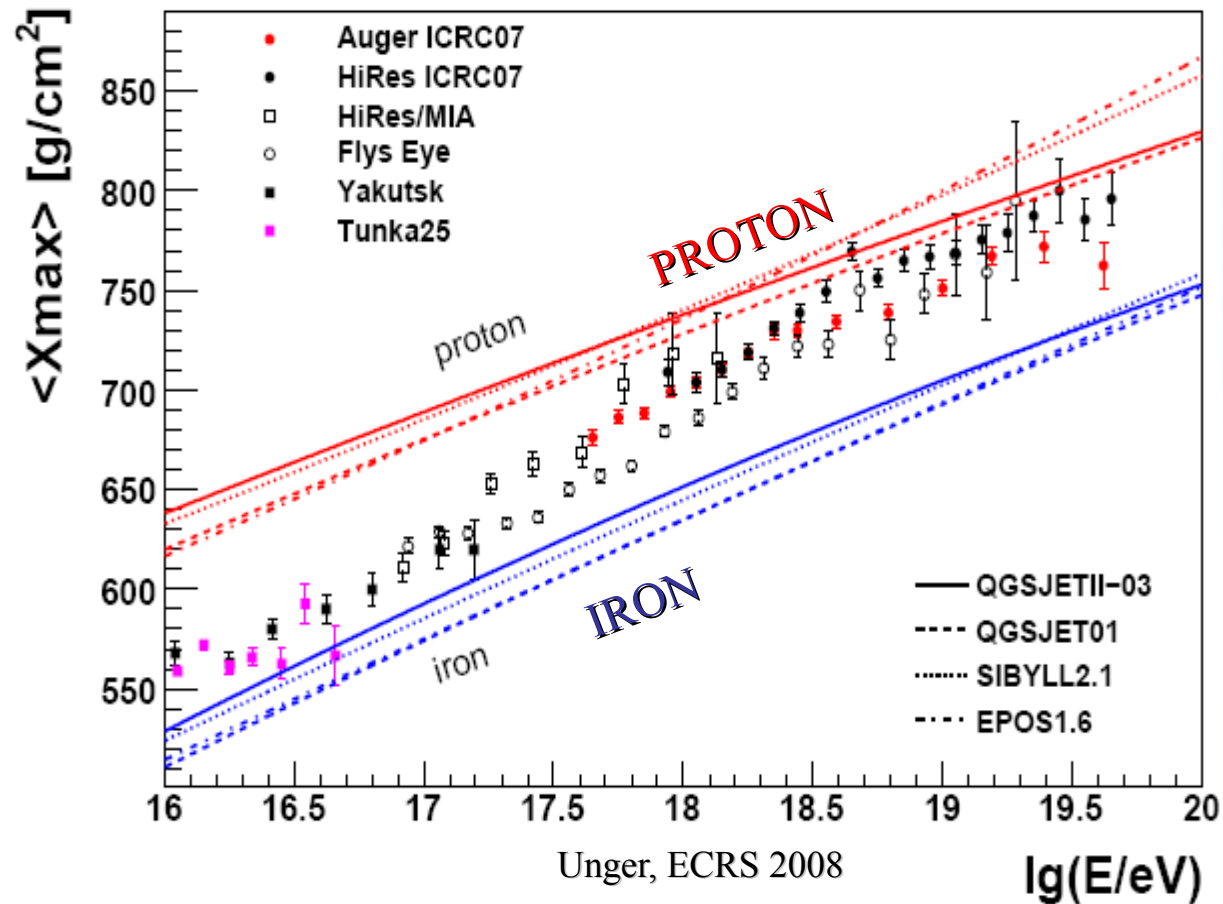
**AGASA Systematics**  
 Total  
 $\pm 18\%$   
 Hadron interaction  
 (QGSJET, SIBYLL)  $\sim 10\%$   
 (Takeda et al., 2003)

Difference in the energy scale between different experiments???

# HECR composition

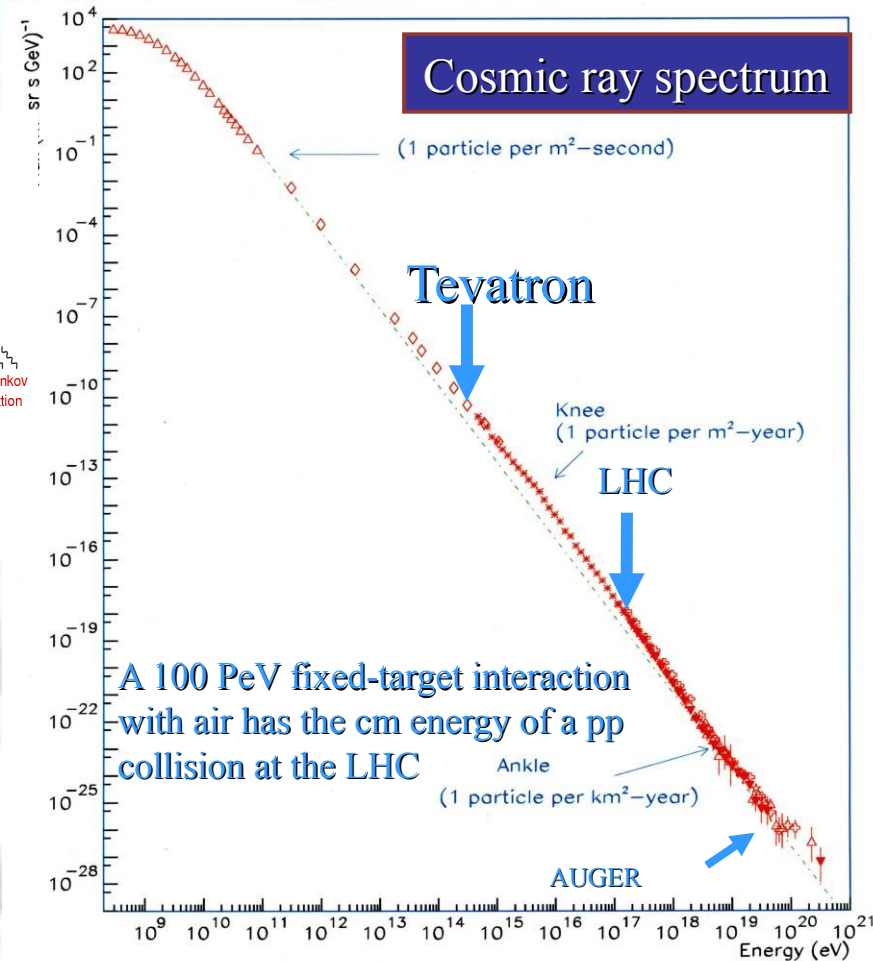
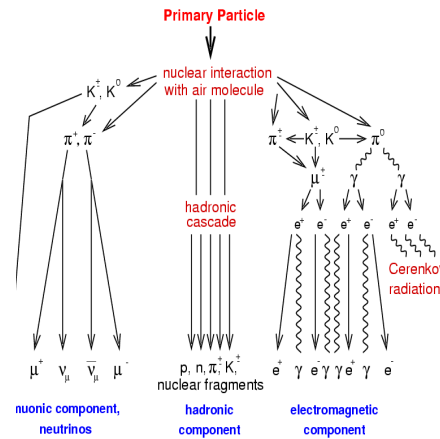
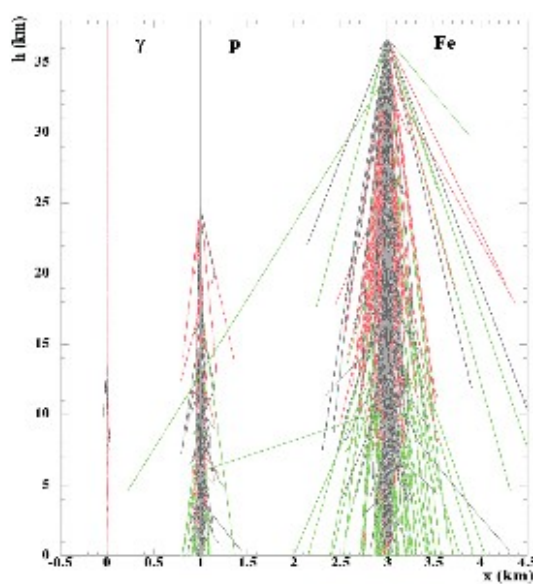
The depth of the maximum of the shower  $X_{\max}$  in the atmosphere depends on energy and type of the primary particle

Different hadronic interaction models give different answers about the composition of HECR





# Development of atmospheric showers



Determination of E and mass of cosmic rays depends on description of primary UHE interaction

Hadronic MC's need tuning with data

The dominant contribution to the energy flux is in the very forward region ( $\theta \approx 0$ )

In this forward region the highest energy available measurements of  $\pi\pi^0$  cross section done by UA7 ( $E \approx 10^{14}$  eV,  $y \approx 3.5$ )



**LHCf: use LHC**

$\sqrt{s} = 14 \text{ TeV} \Rightarrow E_{\text{lab}} = 10^{17} \text{ eV}$

**to calibrate MCs**

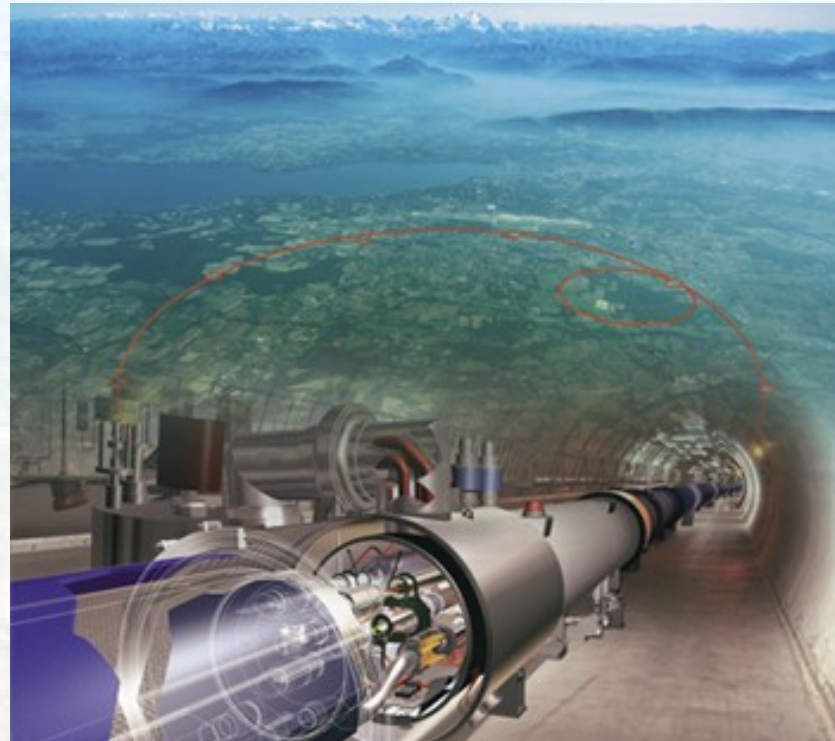


# What is LHCf? The Detectors



# LHCf: an Astroparticle Experiment at LHC

- LHCf is the smallest of the six LHC experiments and is a fully dedicated collider experiment for HECR Physics
- LHCf will use the highest energy particle accelerator to provide useful data to calibrate the hadronic interaction models used in Monte Carlo simulations of atmospheric showers



7TeV+ 7 TeV proton collisions at LHC

Two independent electromagnetic calorimeters equipped with position sensitive layers, on both sides of IP1 will measure energy and position of  $\gamma$  from  $\pi^0$  decays and neutrons produced in pp interaction at LHC

# Experimental set-up

Detector I

Tungsten

Scintillator

Scintillating fibers

INTERACTION POINT

IP1 (ATLAS)

Detector II

Tungsten

Scintillator

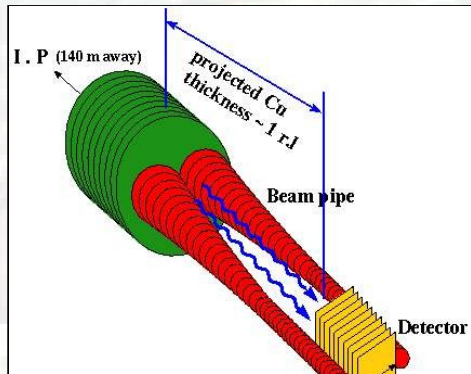
Silicon  $\mu$ strips

140 m

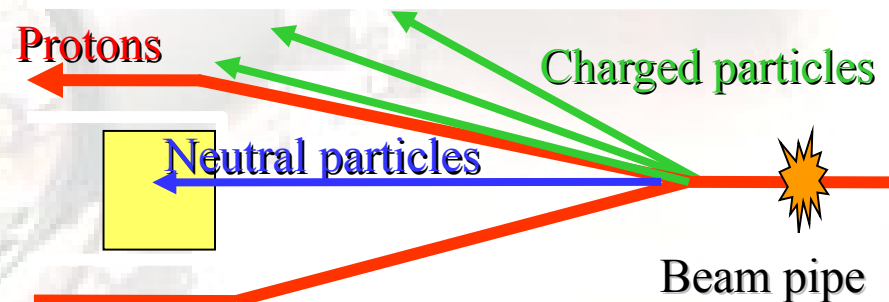
140 m

Beam line

Detectors installed in the TAN region, 140 m away from the Interaction Point 1



- \* Here the beam pipe splits in 2 separate tubes.
- \* Charged particle are swept away by magnets
- \* We cover up to  $y \rightarrow \infty$



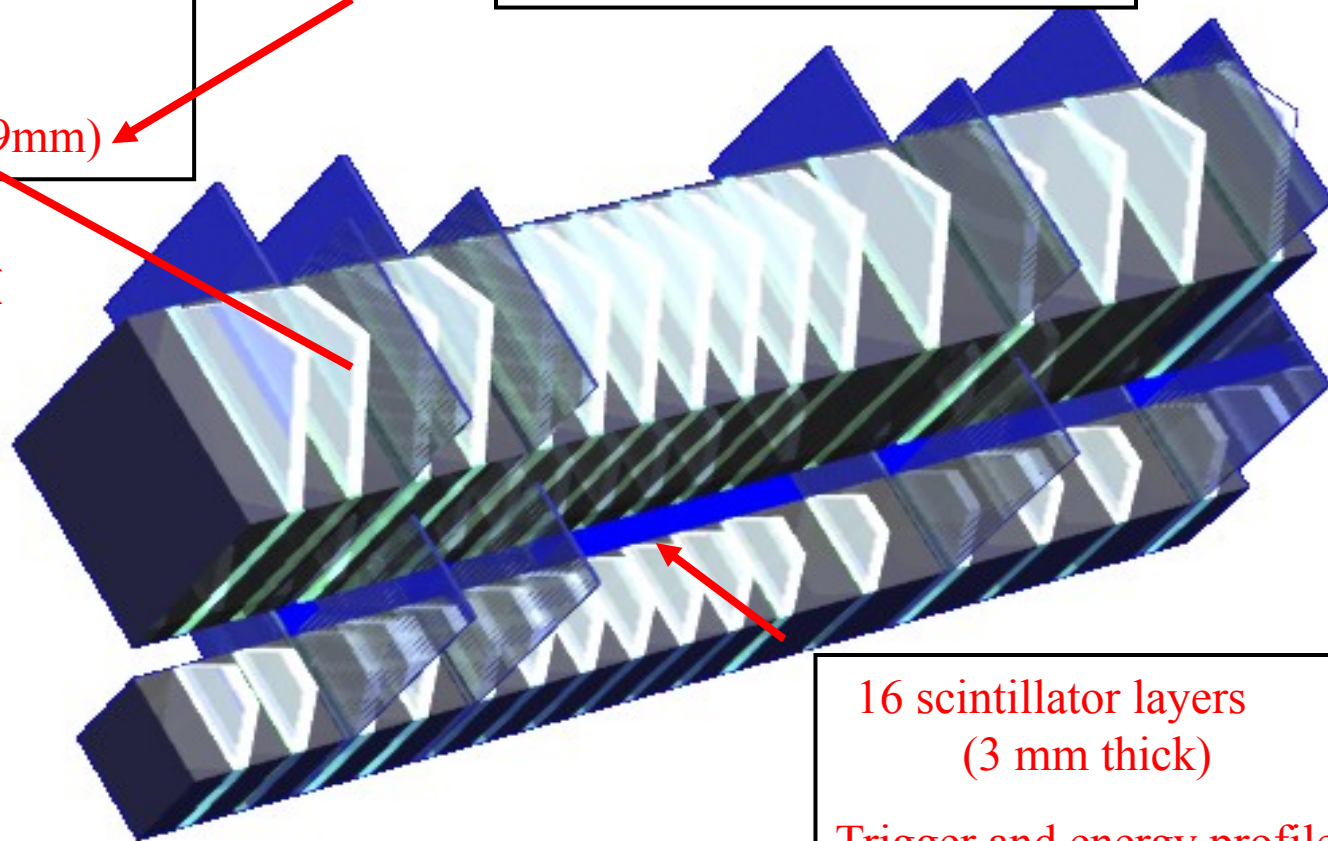
Absorber

22 tungsten layers  
7mm – 14 mm thick

(W:  $X_0 = 3.5\text{mm}$ ,  $R_M = 9\text{mm}$ )

4 pairs of scintillating fiber  
layers for tracking purpose  
(6, 10, 32, 38  $X_0$ .)

Arm1



16 scintillator layers  
(3 mm thick)

Trigger and energy profile  
measurements

2 towers 24 cm long stacked vertically with a 5 mm gap

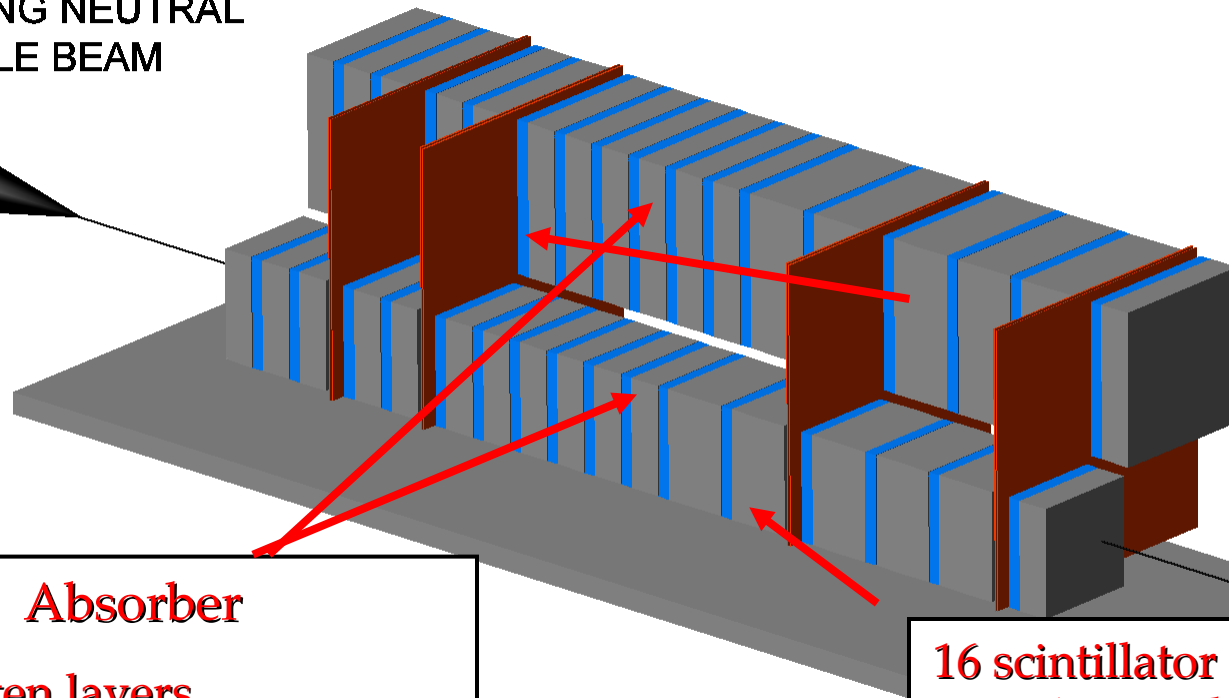
Lower: 2 cm x 2 cm area

Upper: 4 cm x 4 cm area

# Arm2

INCOMING NEUTRAL  
PARTICLE BEAM

4 pairs of silicon microstrip layers  
(6, 10, 30, 42  $X_0$ ) for tracking purpose  
(X and Y directions)



**Absorber**

22 tungsten layers  
7mm – 14 mm thick

(W:  $X_0 = 3.5\text{mm}$ ,  $R_M = 9\text{mm}$ )

16 scintillator layers  
(3 mm thick)

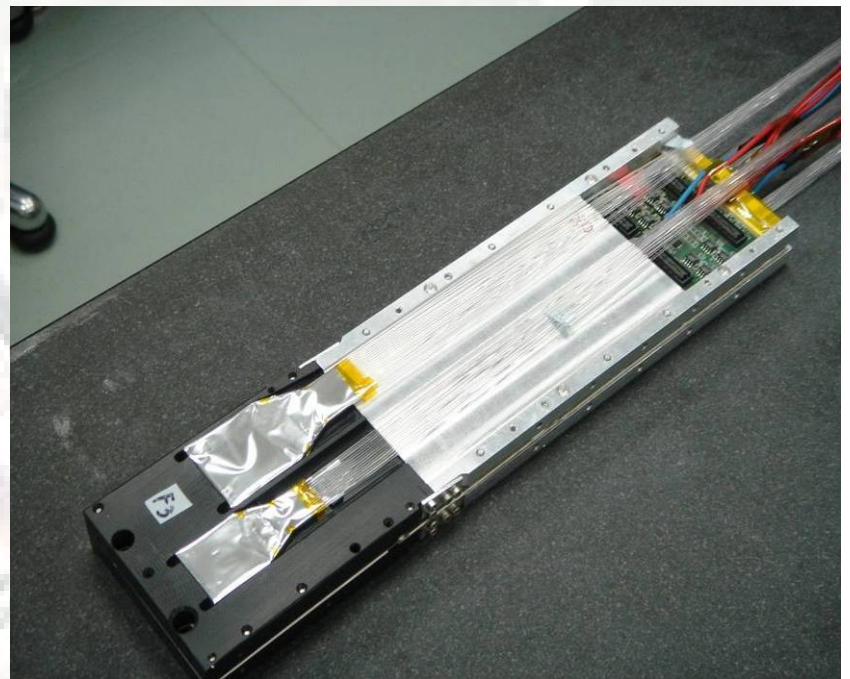
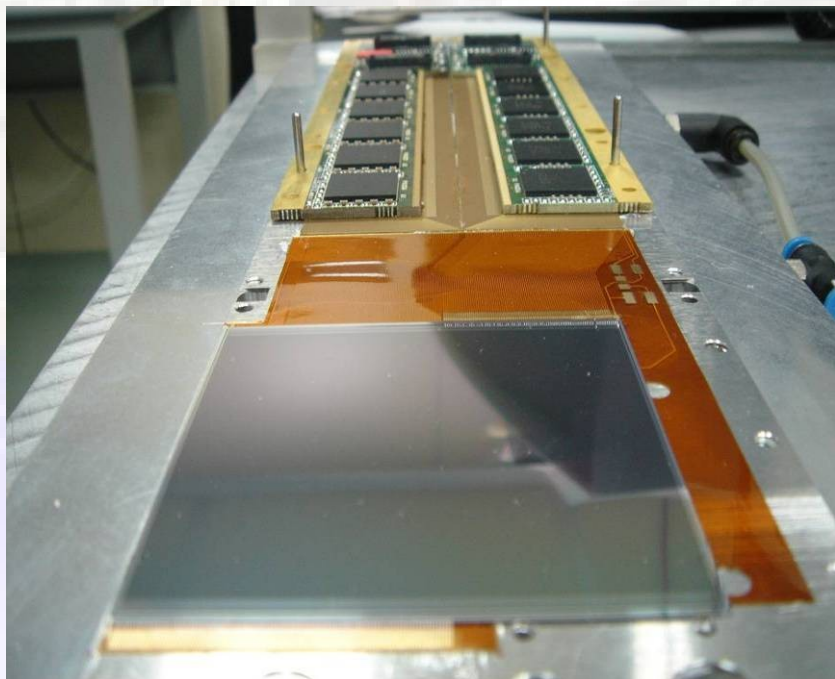
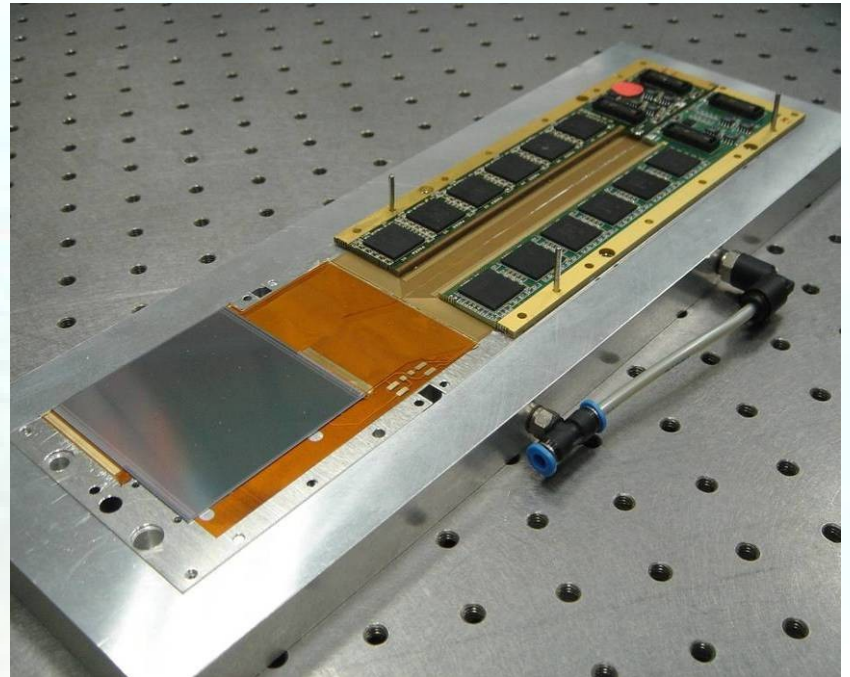
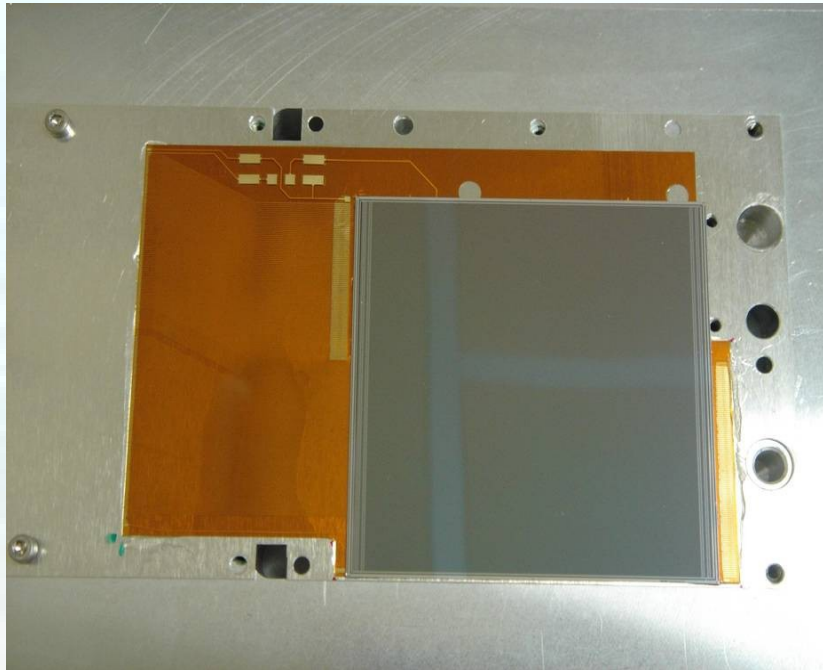
Trigger and energy  
profile measurements

2 towers 24 cm long stacked on their edges and offset from one another

Lower: 2.5 cm x 2.5 cm

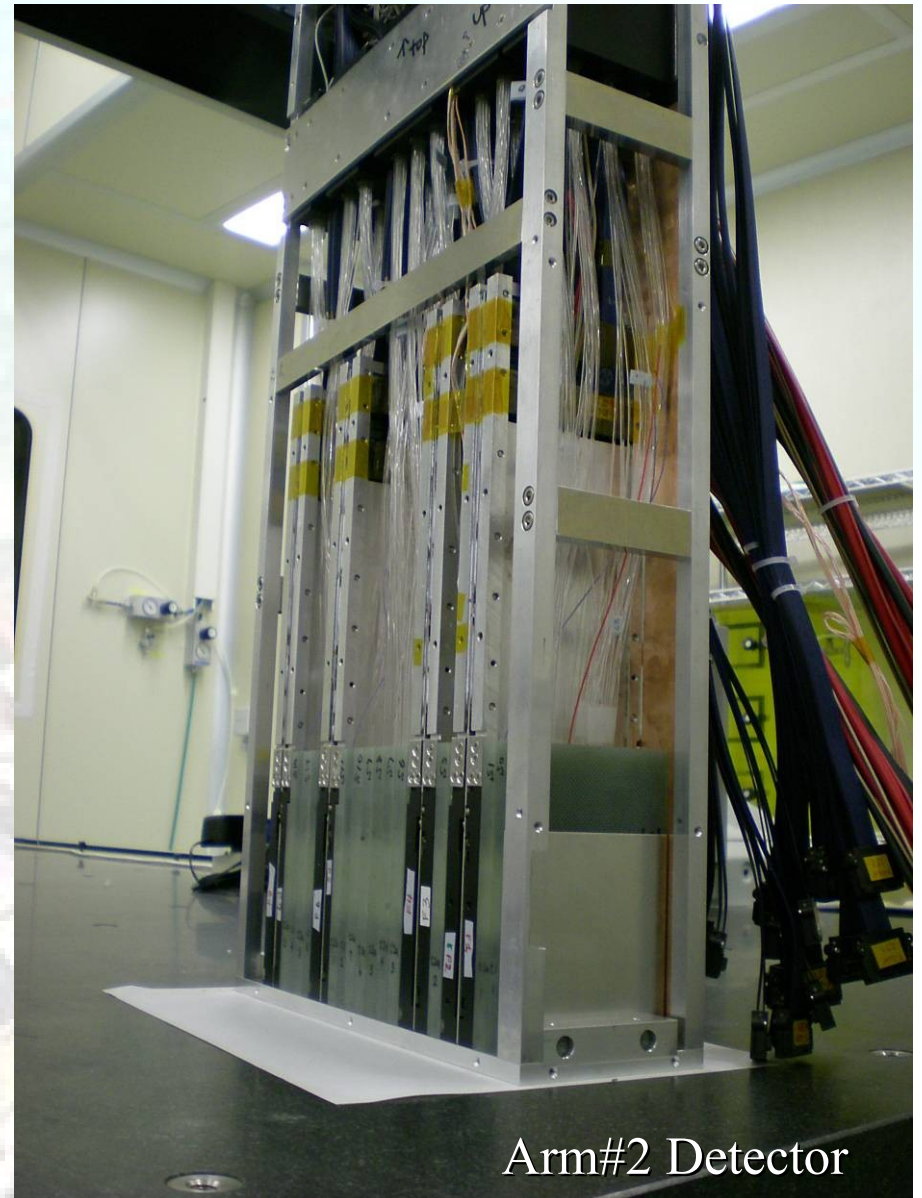
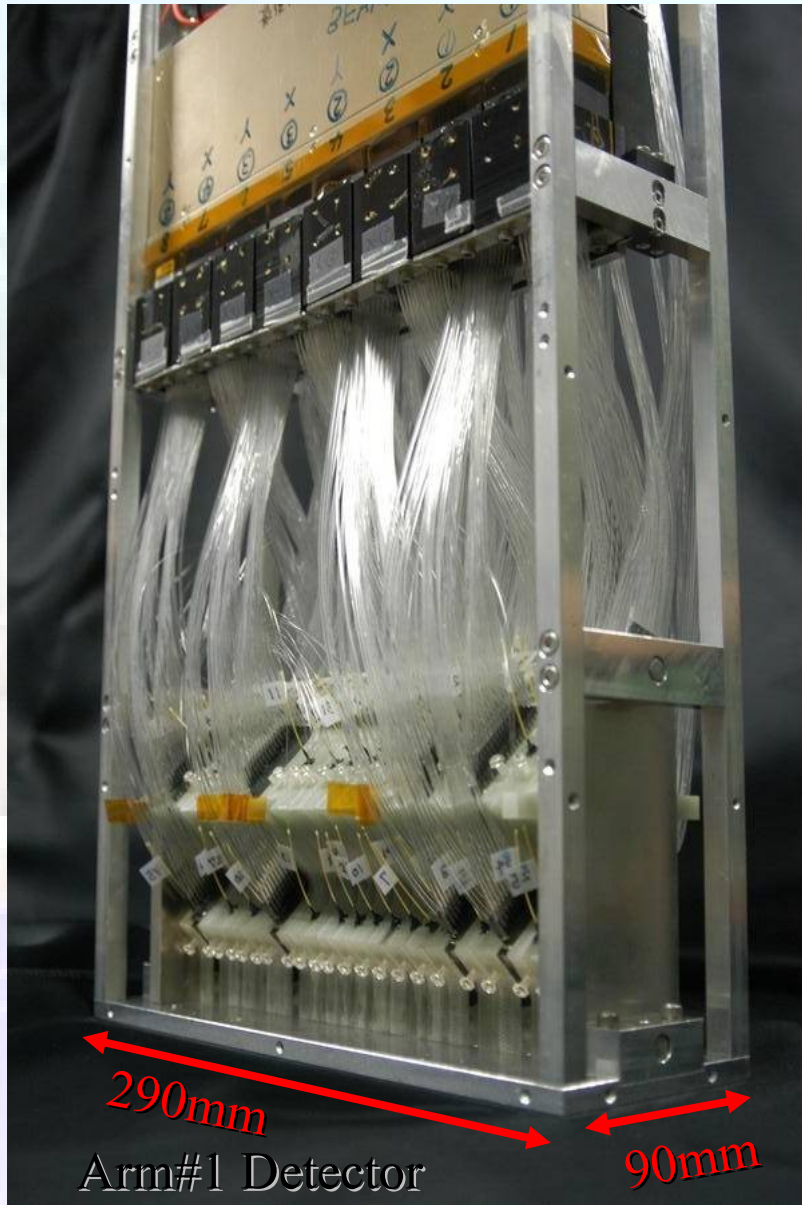
Upper: 3.2 cm x 3.2 cm







# Double ARM Detectors





# LHCf detectors in the LHC tunnel



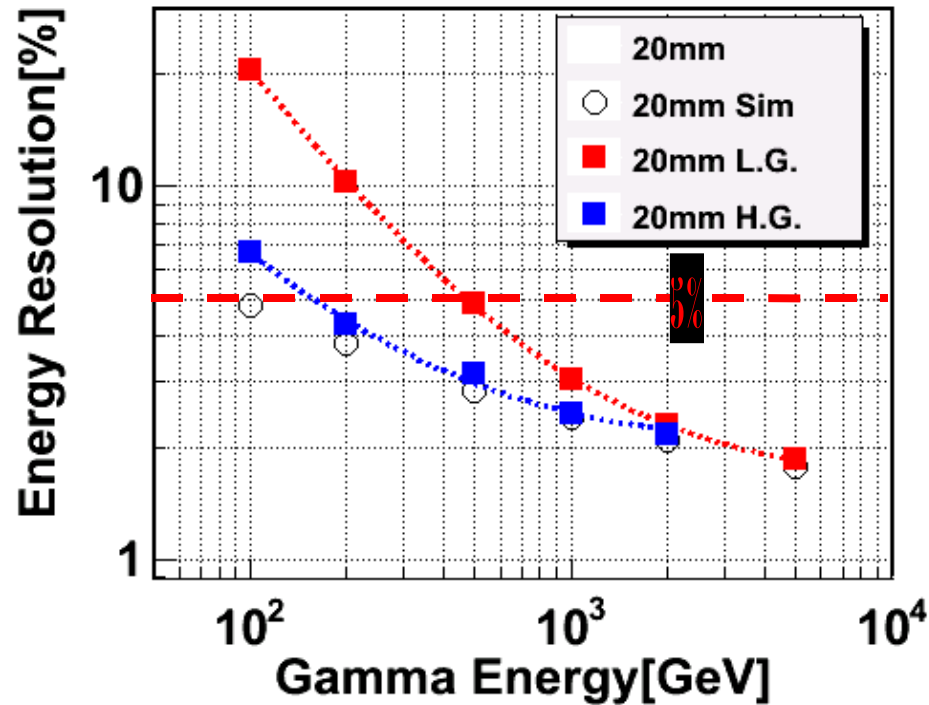


# **How can LHCf calibrate MCs?**

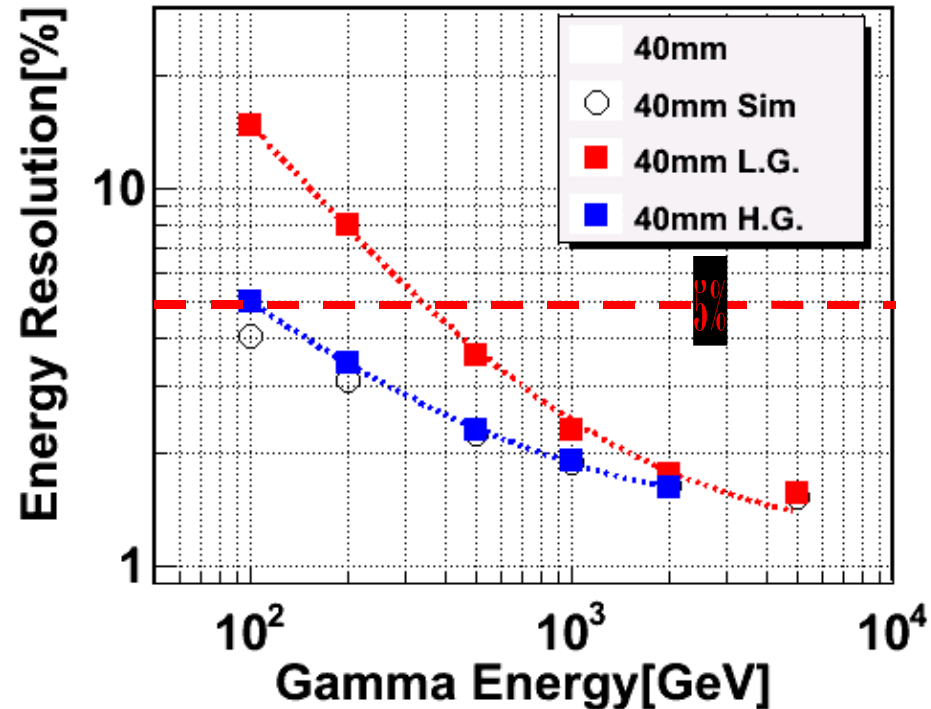
## **Physics Performance**

# Energy resolution for $\gamma$

For 20mm



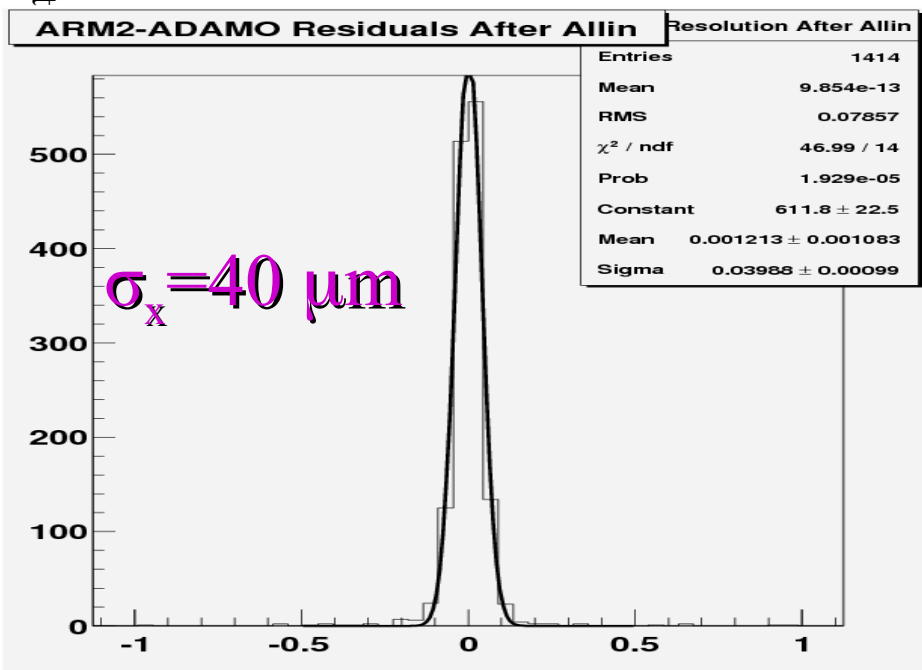
For 40mm



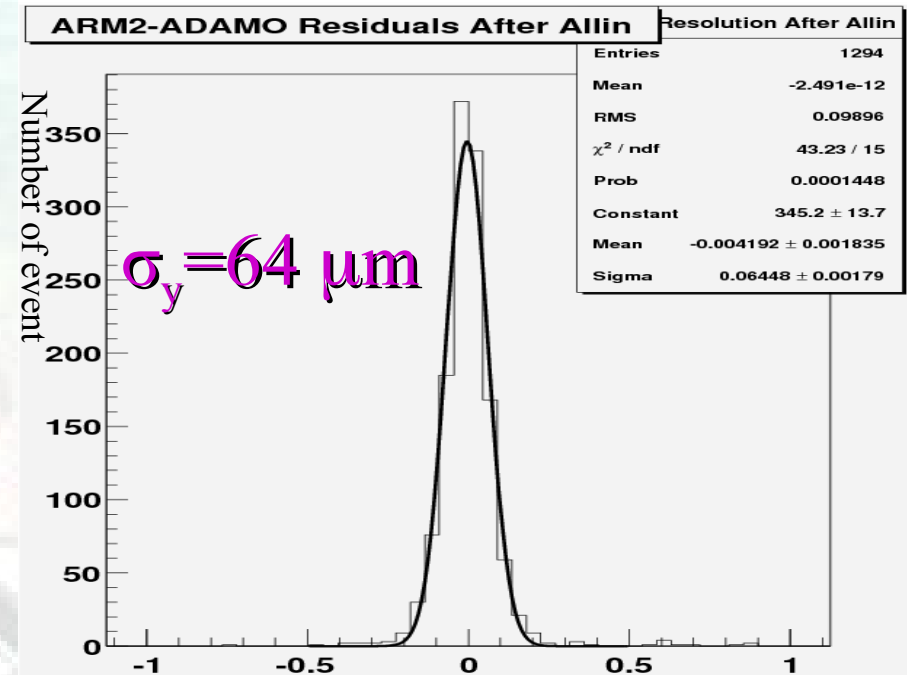
# ARM2 Position Resolution

200 GeV electrons

Number of event



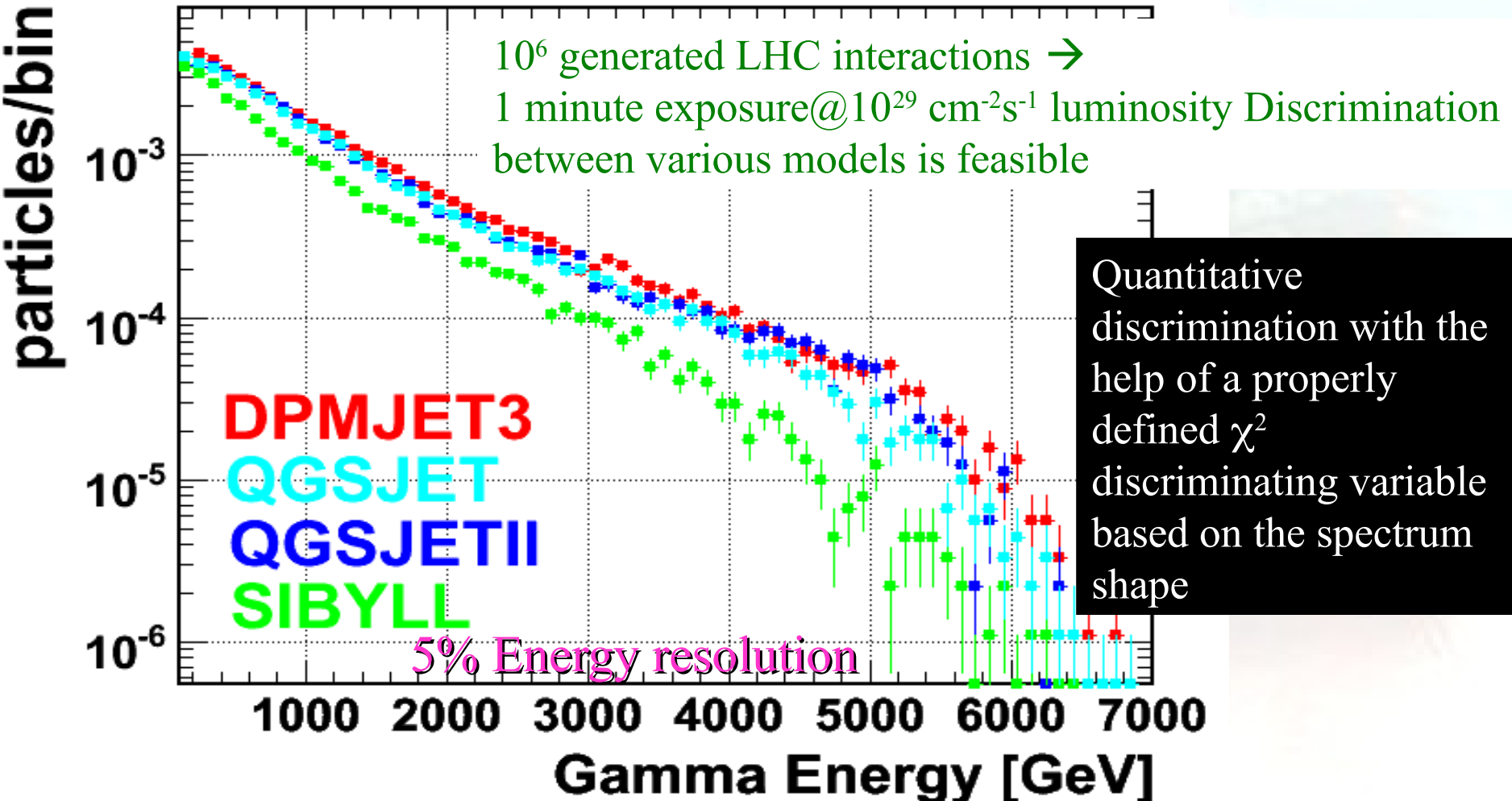
x-pos[mm]



y-pos[mm]

# LHCf: Monte Carlo discrimination

## Gamma Energy Spectrum of 20mm square at Beam Center

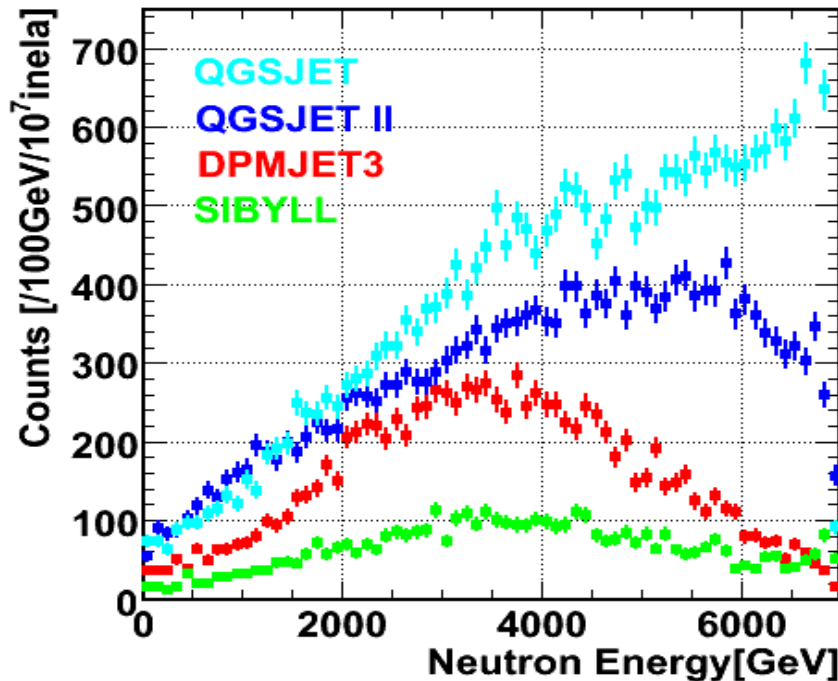




# LHCf: model dependence of neutron energy distribution

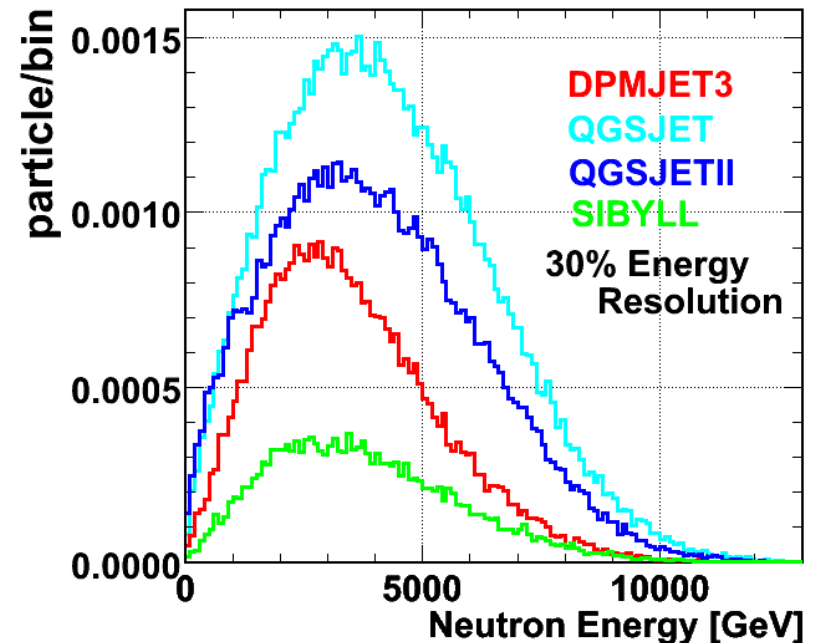
Original n energy

Neutron Energy Distributions



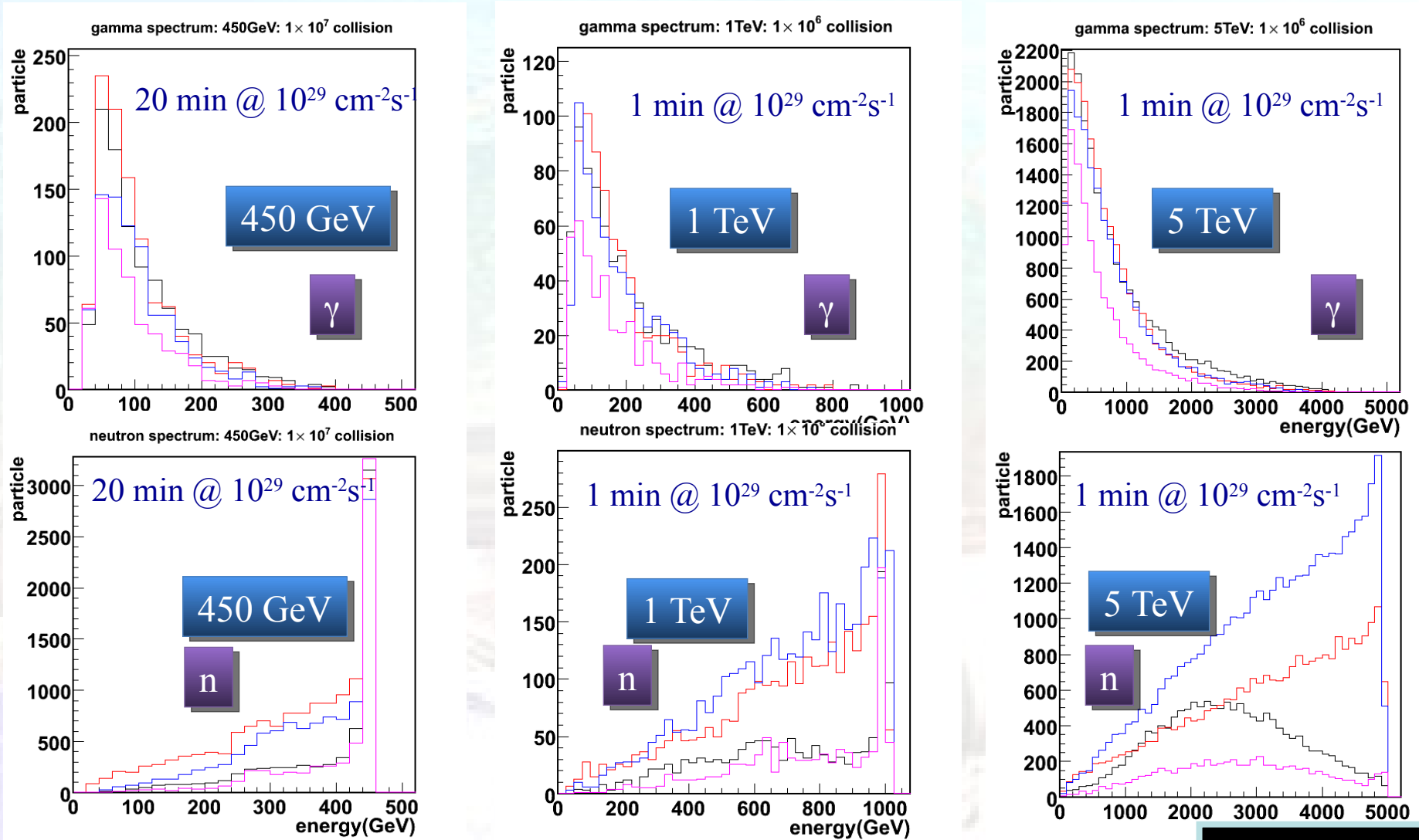
30% energy resolution

Neutron Energy Spectrum  
of 20mm Calorimeter at beam center





# $\gamma$ and n spectra at different energies



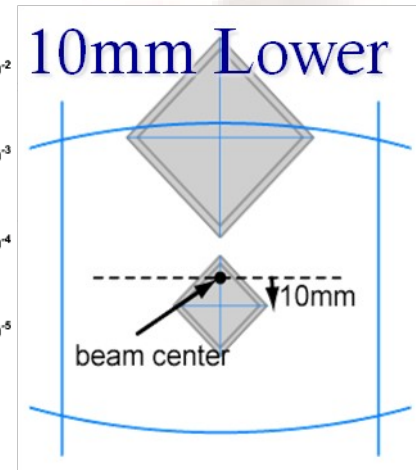
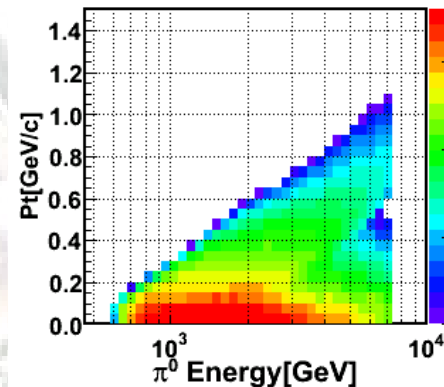
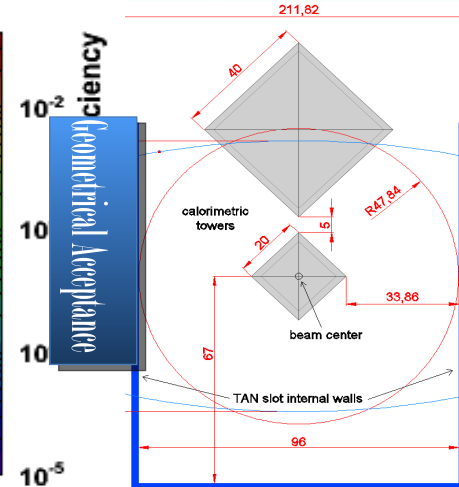
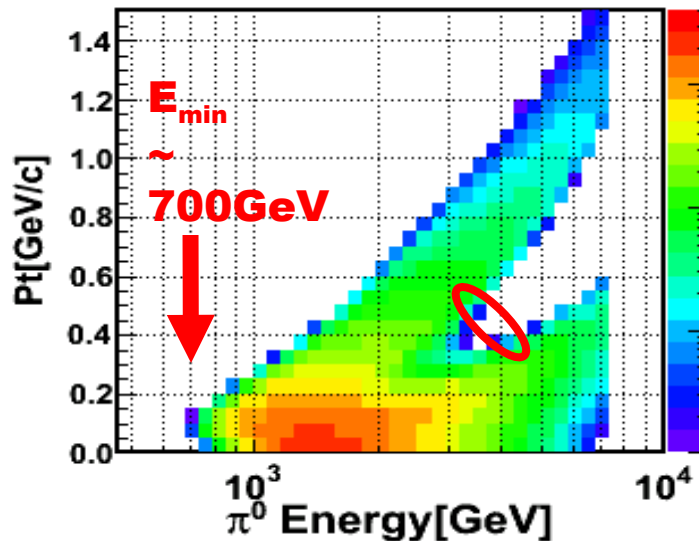
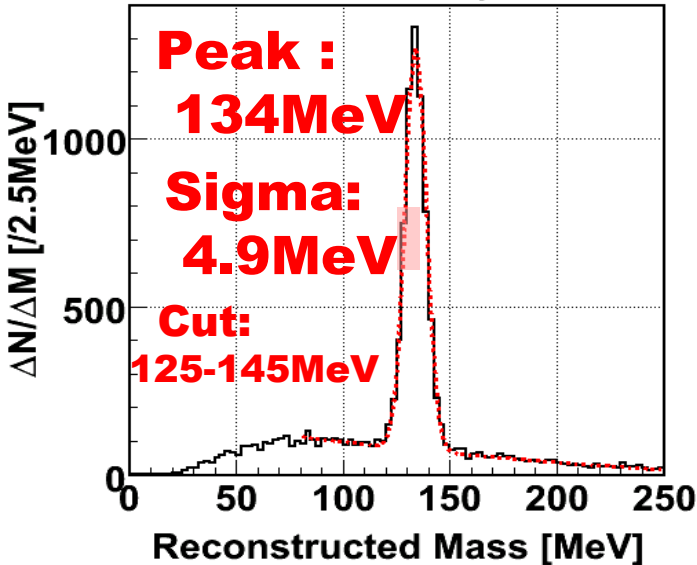
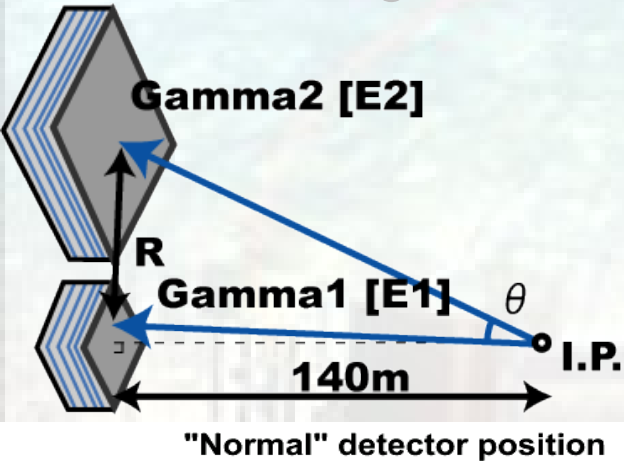
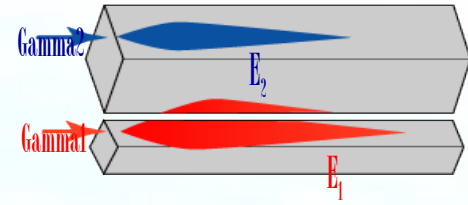
No detector resolution taken into account

DPMJET3  
QGSJET2  
QGSJET1  
SIBYLL

# $\pi^0$ spectra

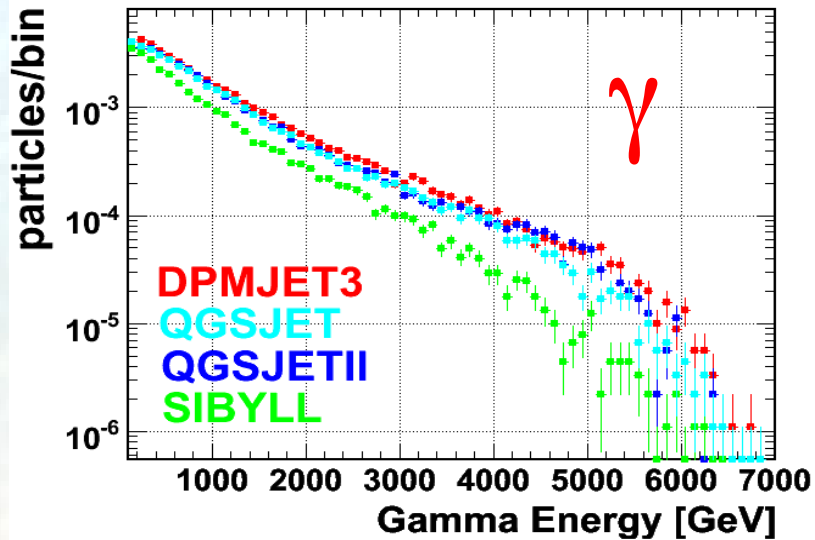
$\pi^0$  produced at collision can be extracted by using gamma pair events

Powerful tool to calibrate the energy scale and also to eliminate beam-gas BG



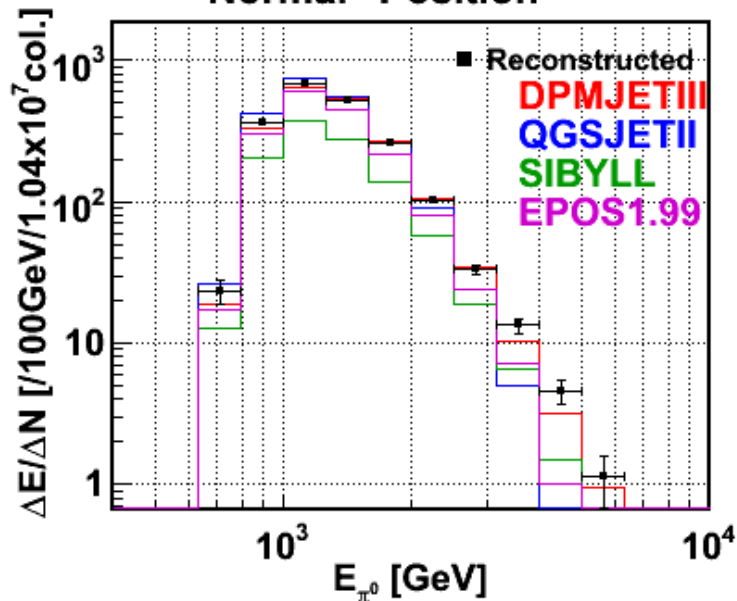
# LHCf : Monte Carlo discrimination

Gamma Energy Spectrum  
of 20mm square at Beam Center



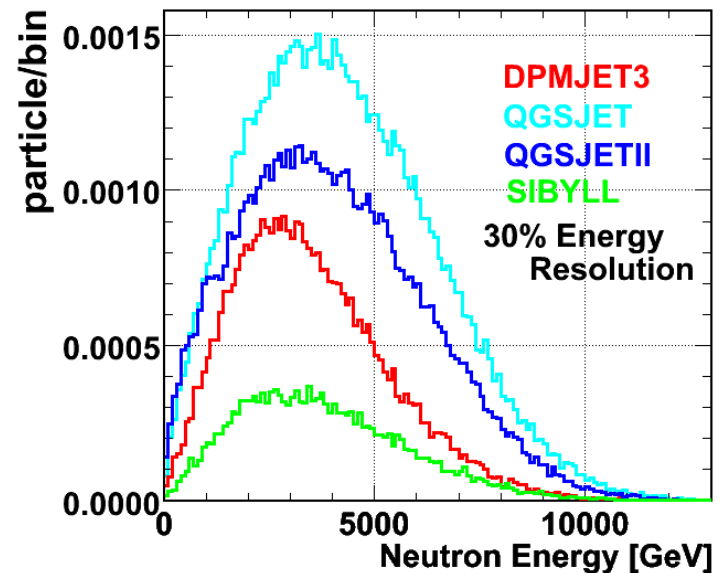
$10^6/10^6$  generated LHC interactions  
at  $\sqrt{s} = 7$  TeV  $\rightarrow$   
1 minute exposure  $\approx 10^{29} \text{ cm}^{-2} \text{ s}^{-1}$  luminosity


"Normal" Position



$\pi^0$

Neutron Energy Spectrum  
of 20mm Calorimeter at beam center





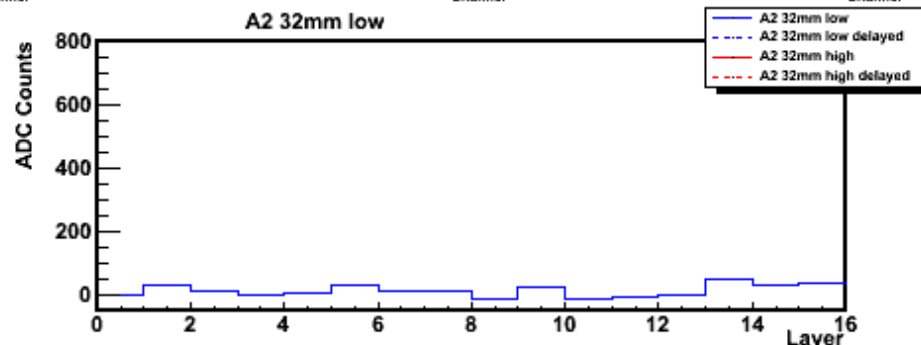
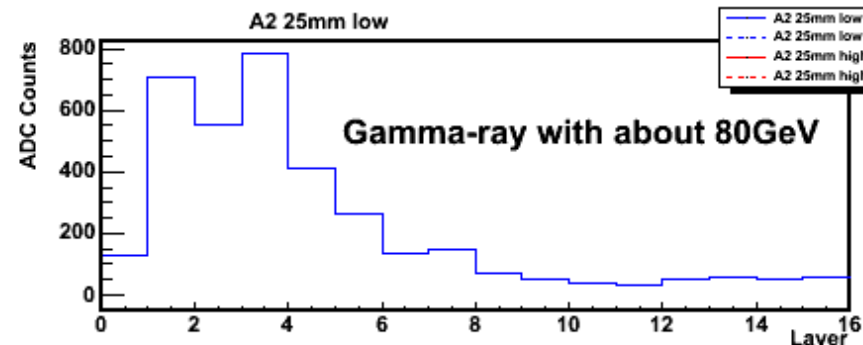
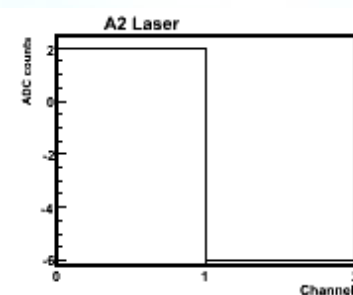
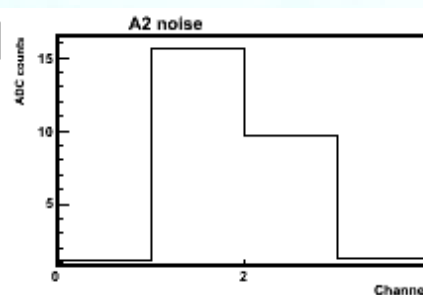
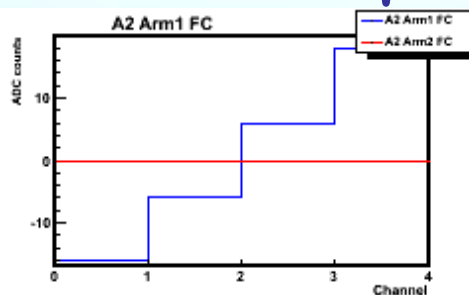
**When LHCf?  
NOW!!!!!!!!!!!!!!**

# 2009 LHC Operation

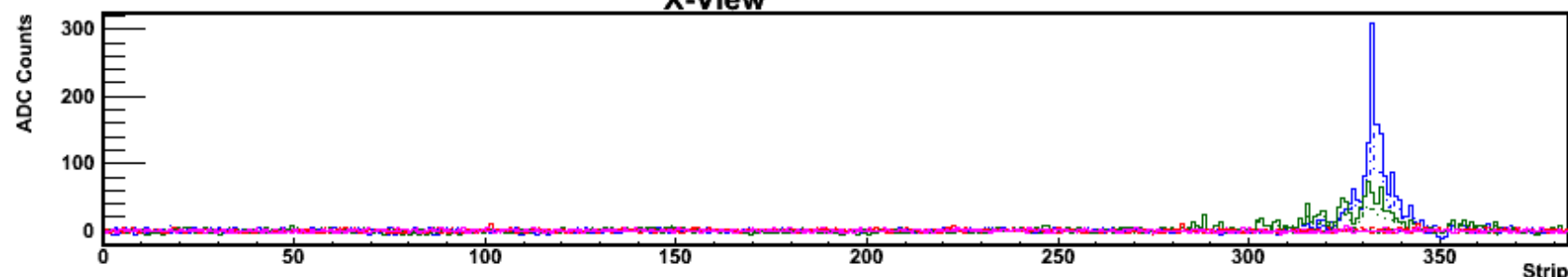
- From End of October 2009 LHC restarted operation
- $450 \text{ GeV} + 450 \text{ GeV} \rightarrow 1.2 \text{ TeV} + 1.2 \text{ TeV}$
- Exceptional effort and success from LHC!!!
- Few weeks of 'smooth' running allowed LHCf to collect some statistics at  $450+450 \text{ GeV}$  in stable beam conditions (Moving from garage to running position) 😊 😊 😊 😊
- No stable beam at  $1.2+1.2 \text{ TeV} \rightarrow$  No data at this energy for this year ☹

# Arm2 $\gamma$ event

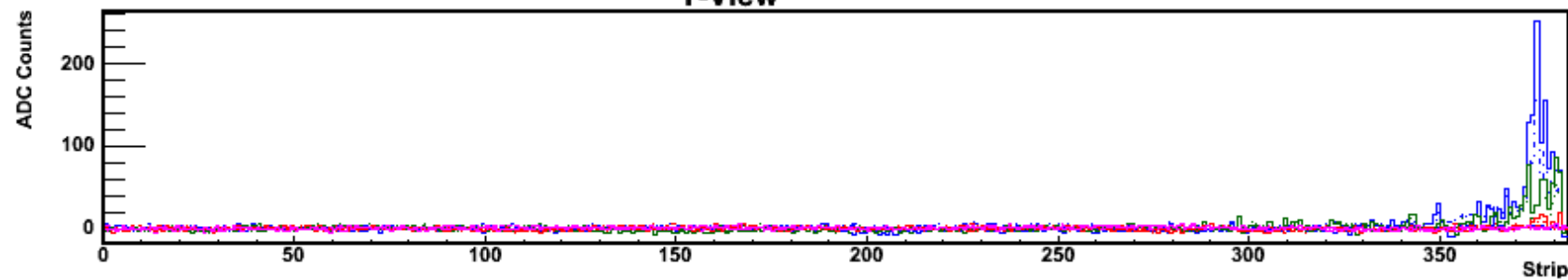
RUN: 2342  
 NUMBER: 506  
 GNUMBER: 1154  
 TIME: 1260085179  
 FLAG0: 00009557  
 FLAG1: 000009ff  
 FLAG2: 00a02371



X-View



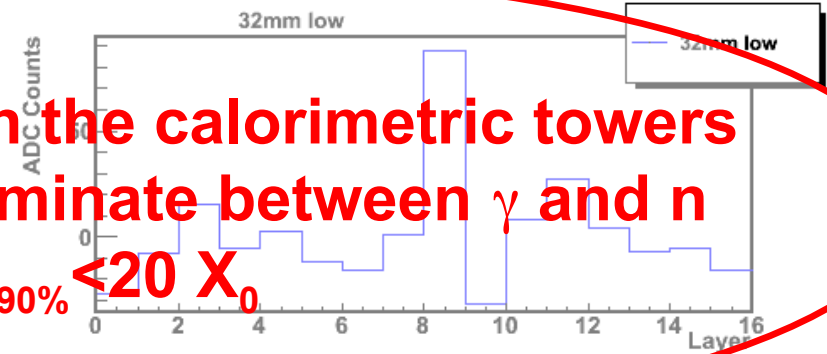
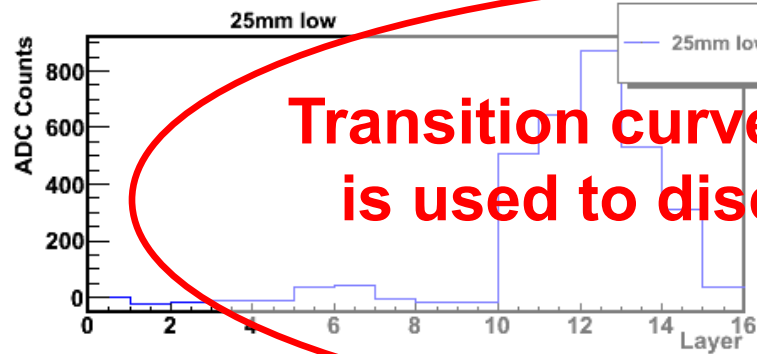
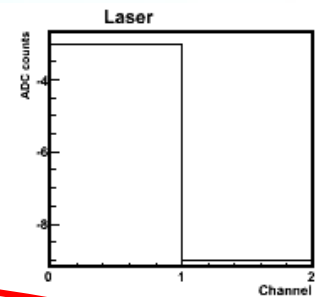
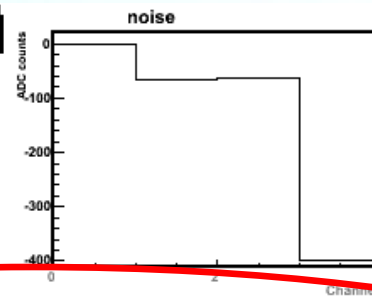
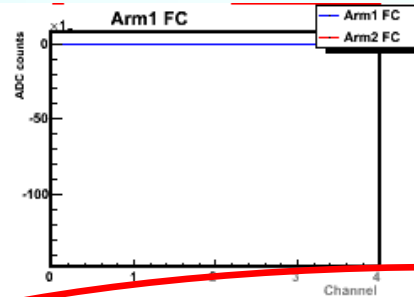
Y-View





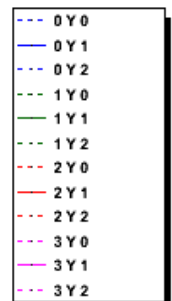
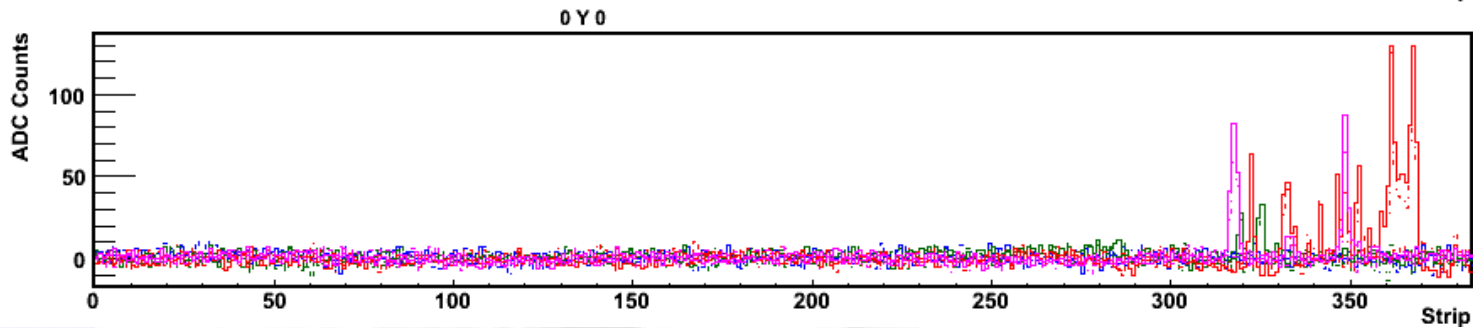
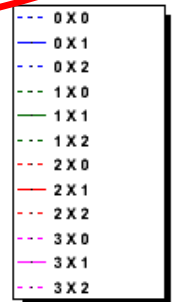
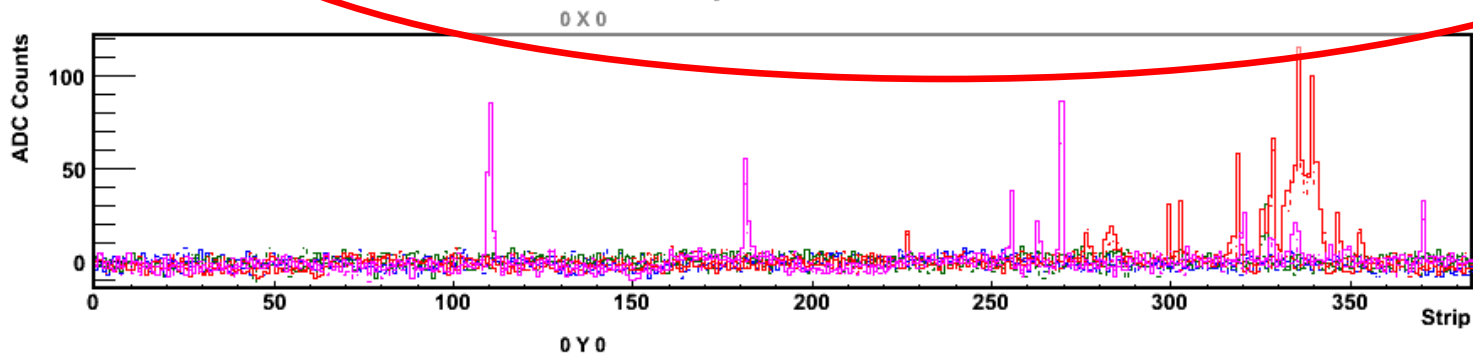
# Arm2 neutron event

RUN: 2274  
NUMBER: 1076  
GNUMBER: 2568  
TIME: 1259889430  
FLAG0: 00009517  
FLAG1: 0100fc60  
FLAG2: 00f02375

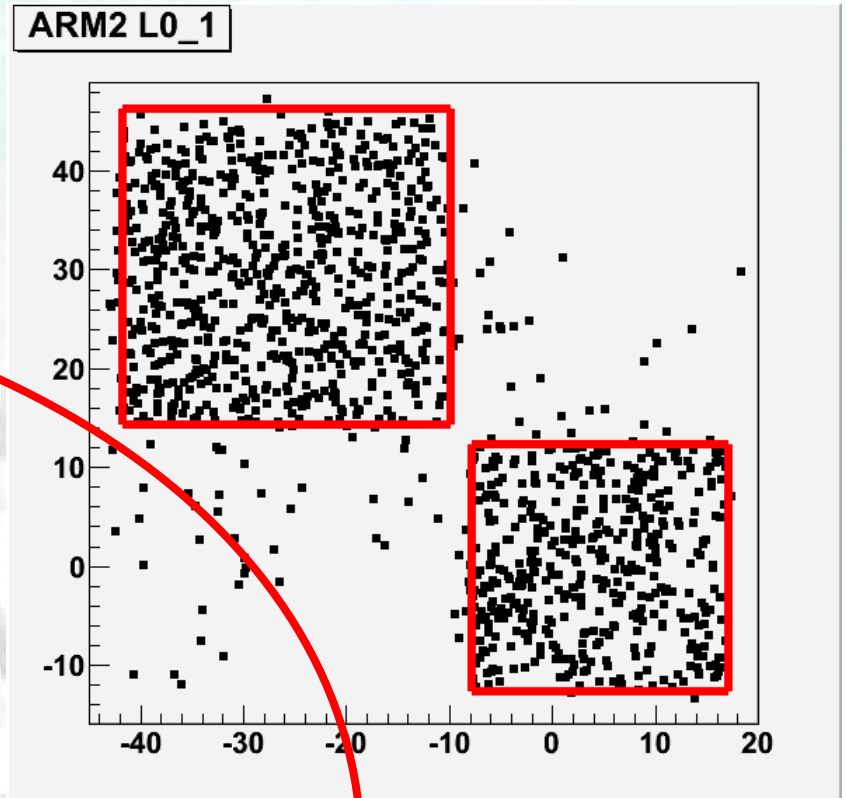
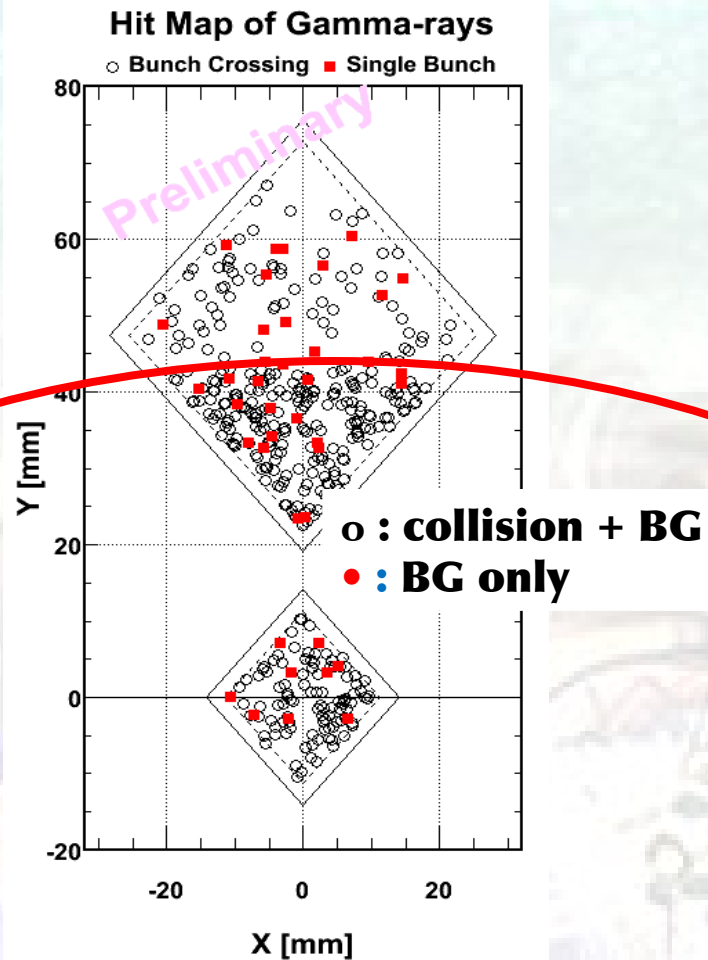


Transition curve in the calorimetric towers  
is used to discriminate between  $\gamma$  and  $n$

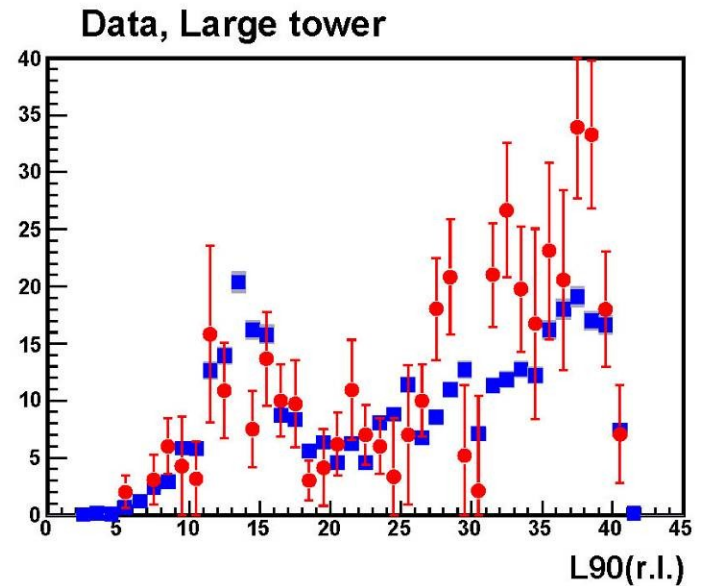
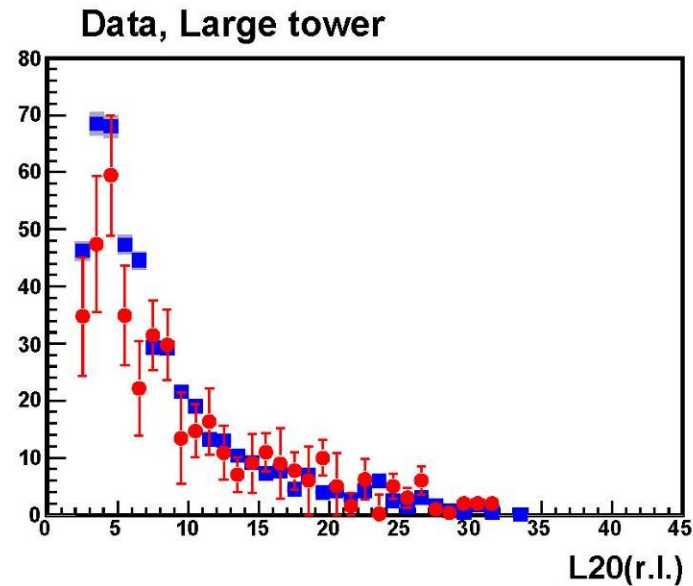
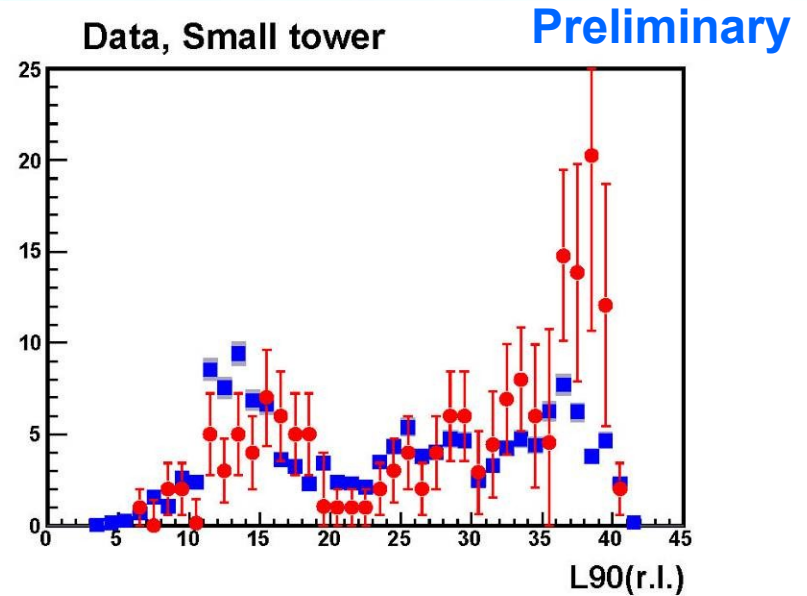
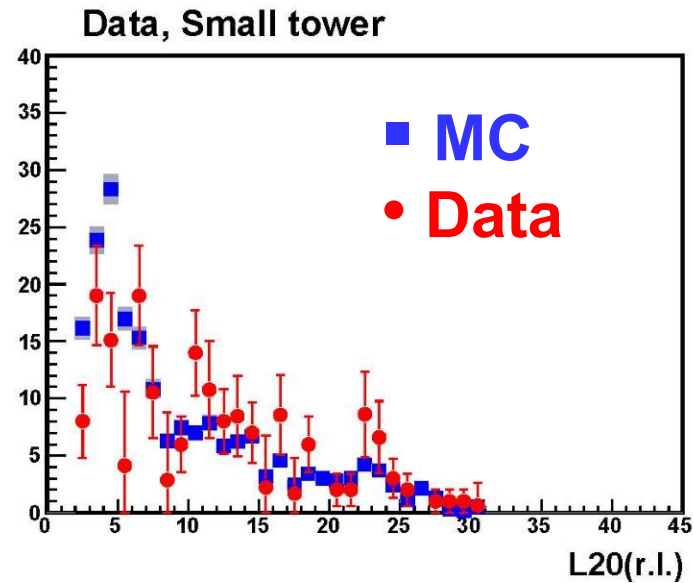
$$L_{90\%} < 20 X_0$$



# Hit map on ARM1 and ARM2

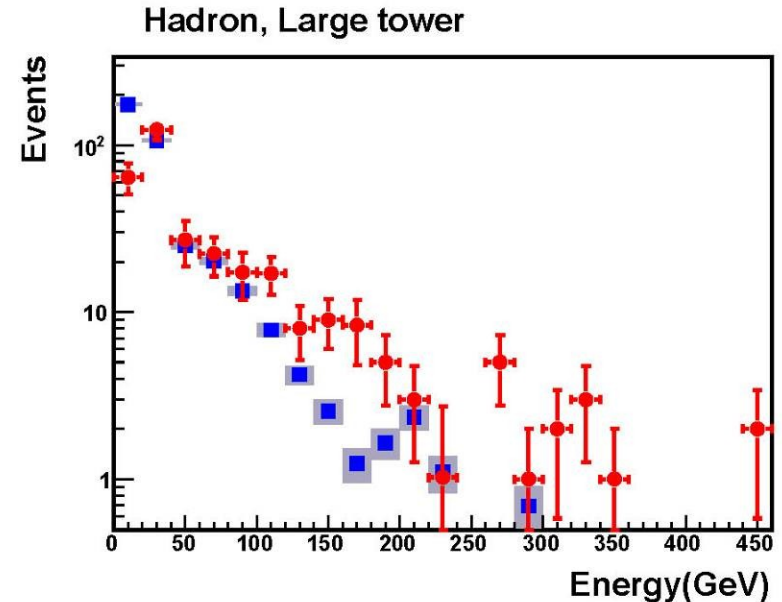
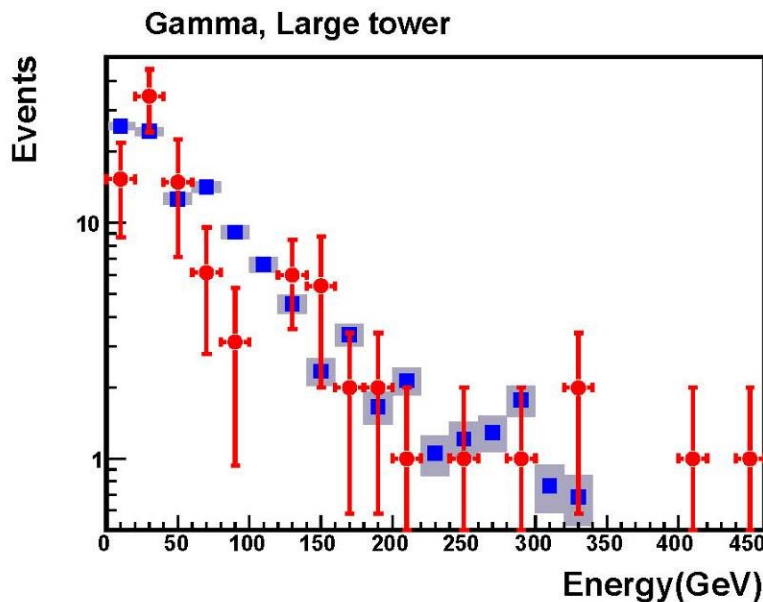
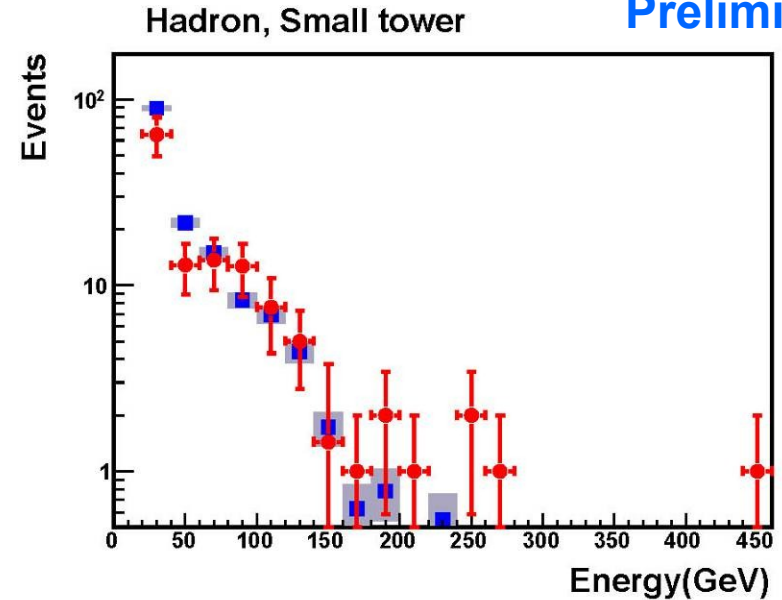
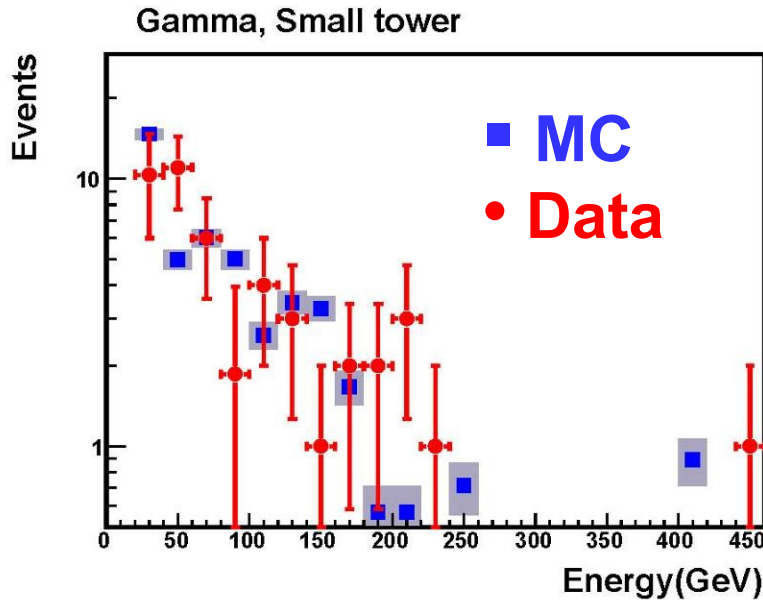


# Arm1 Results: L20 and L90



# Arm1 Results: $\gamma$ and n spectra on the 2 towers

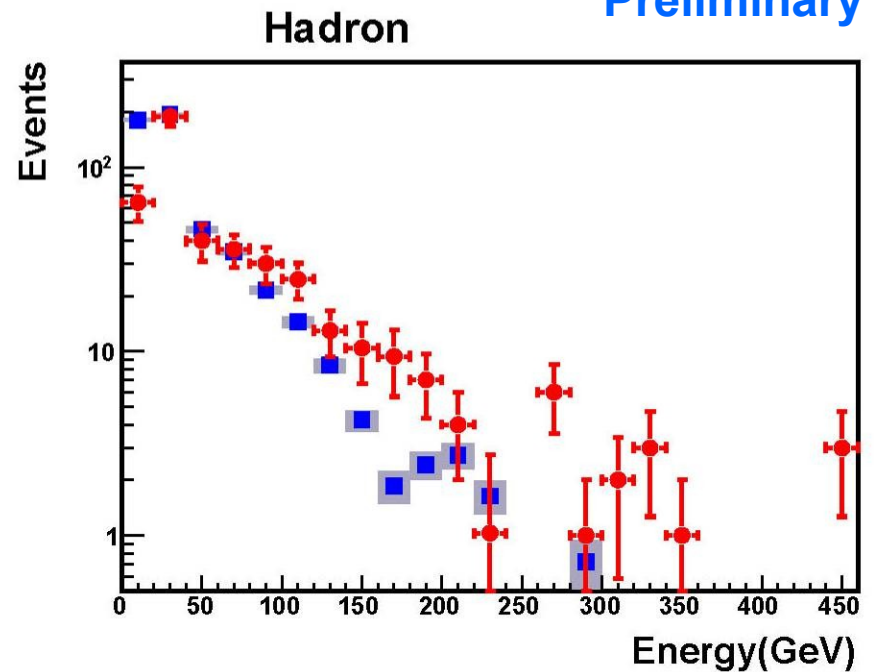
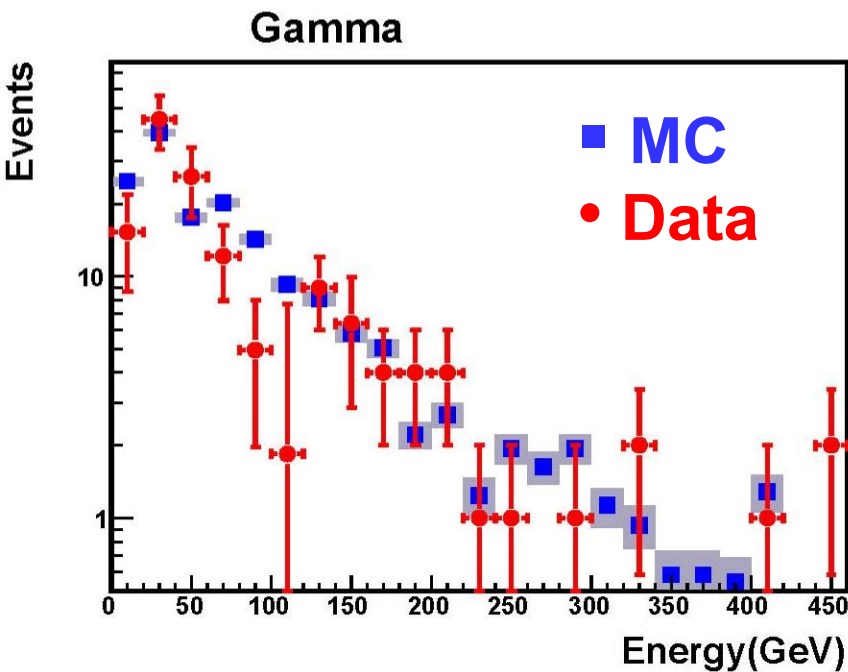
Preliminary



After the subtraction of the Not Colliding Bunches properly normalized

# Arm1 Results: combined $\gamma$ and n spectra

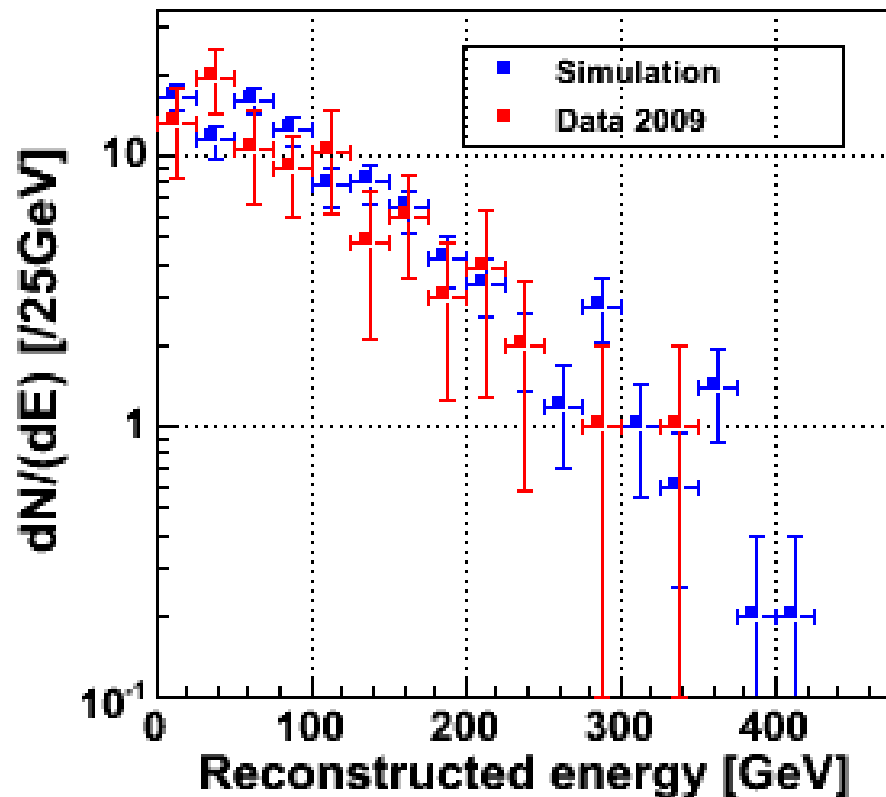
Preliminary



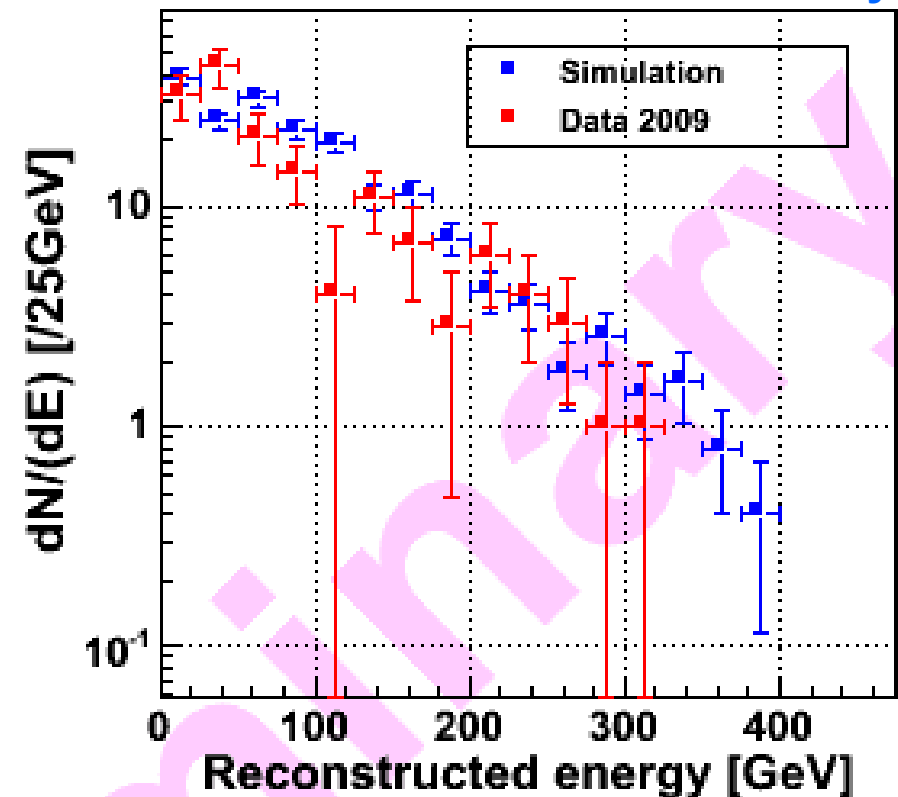


# Arm2 Results: $\gamma$ spectra on the 2 towers

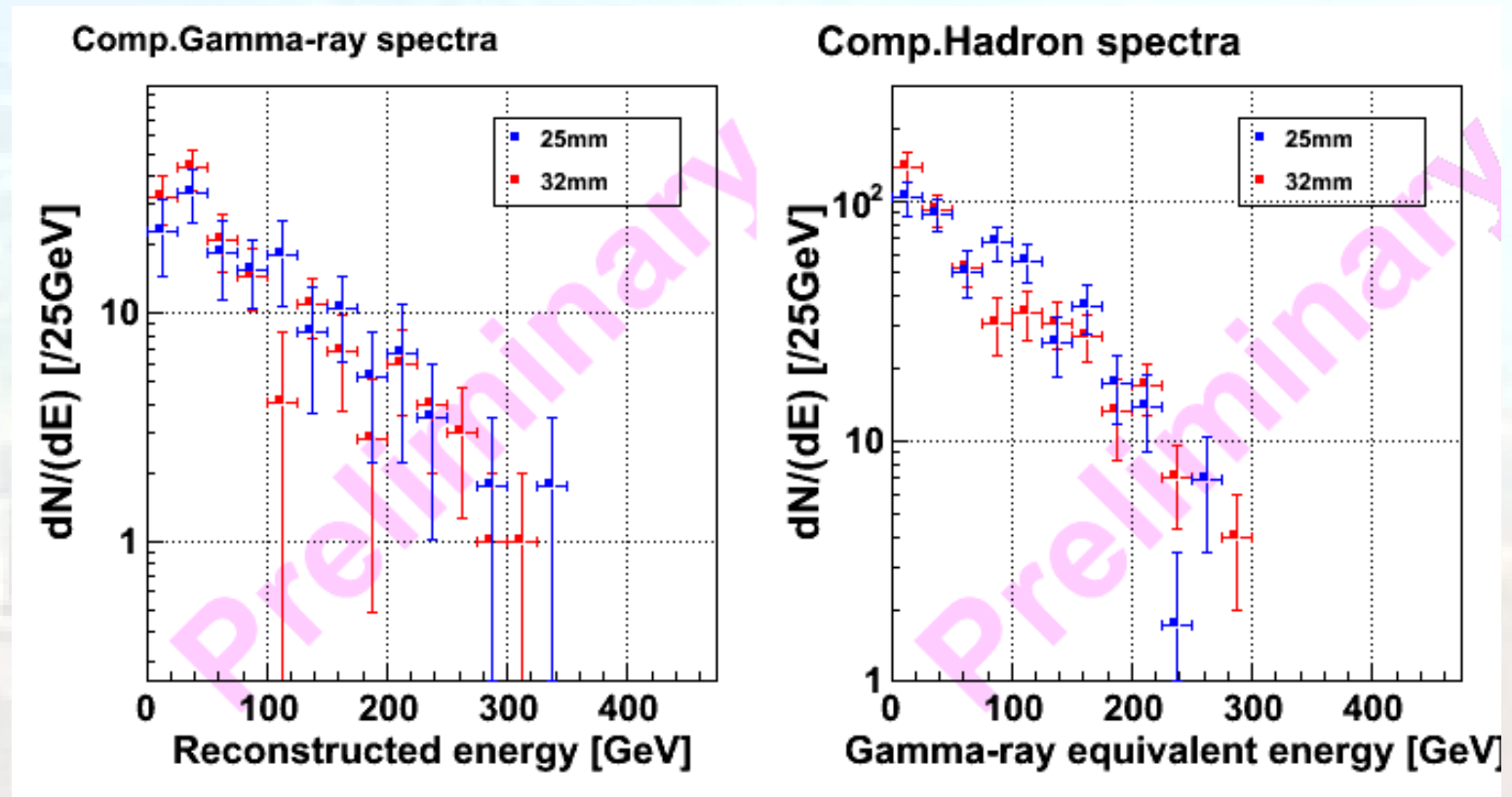
Gamma-ray @ 25mm



Gamma-ray @ 32mm Preliminary



# Arm2 Results: Comparison of 25mm and 32mm



Chi2 : 13.3 (DOF=10)

Chi2 : 7.7 (DOF=10)

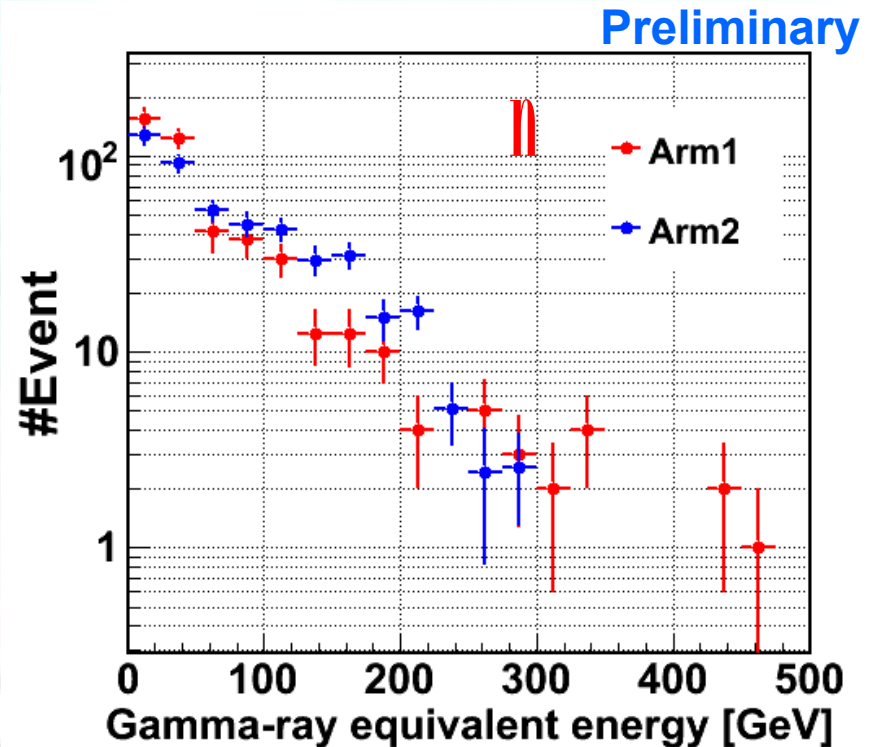
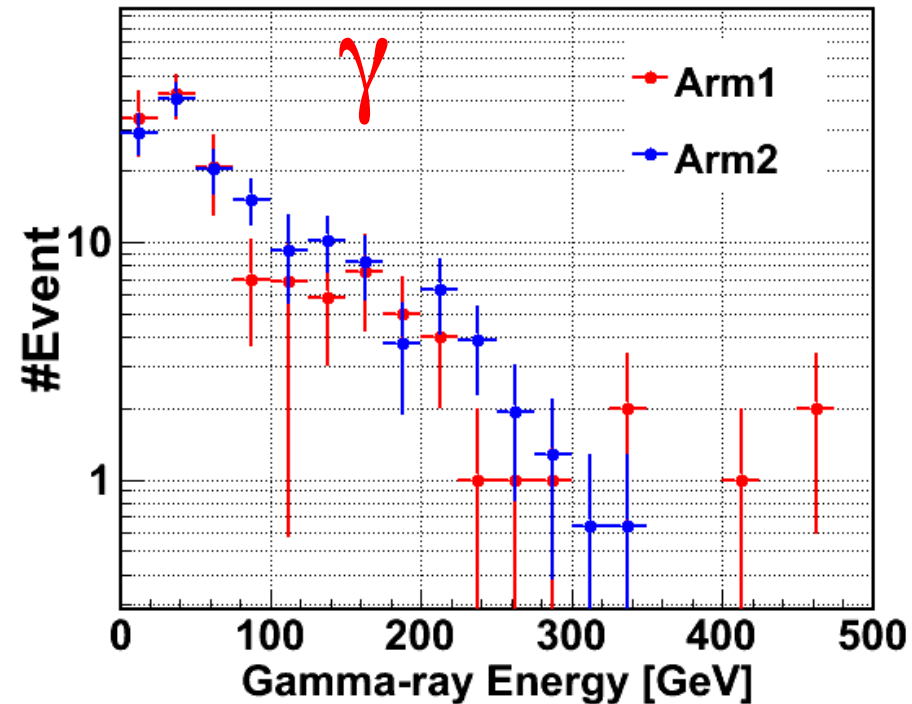
The spectra of 32 mm are normalized by the relative acceptance (factor 1.77)

No significant difference between 25mm and 32mm spectra.

It is consistent with the expectation by simulation:

**Flat distributions at 450 GeV**

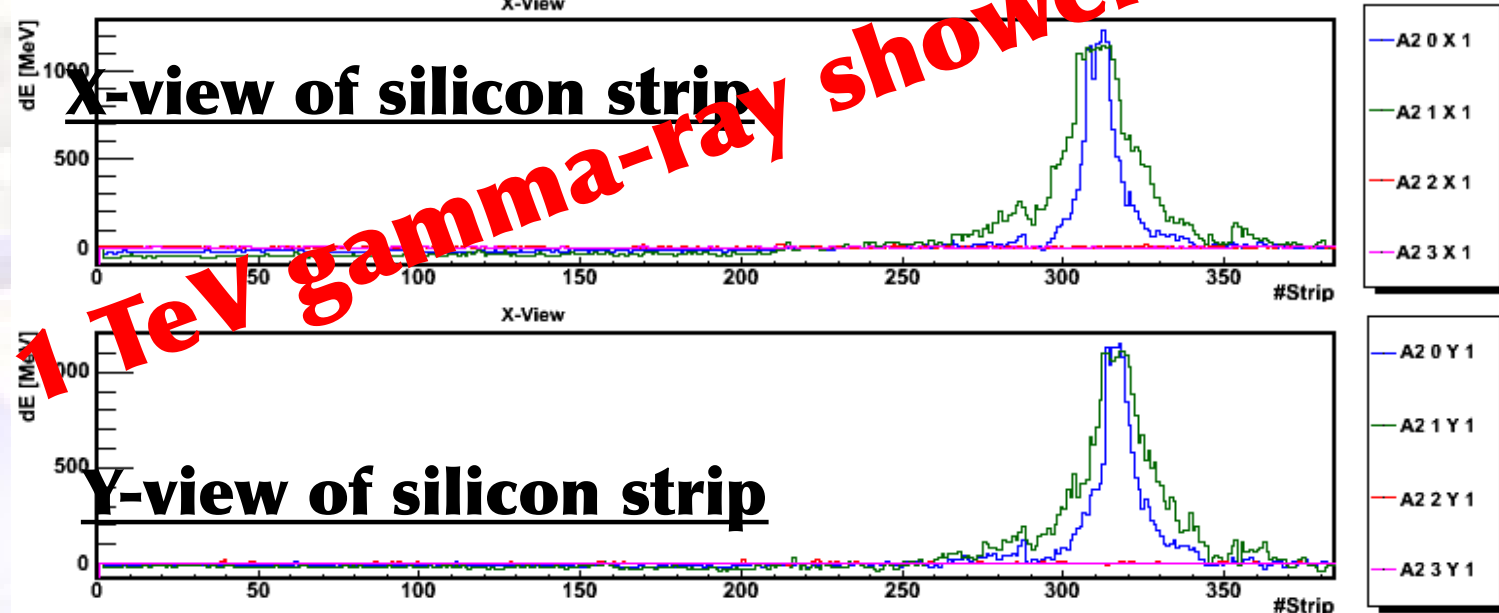
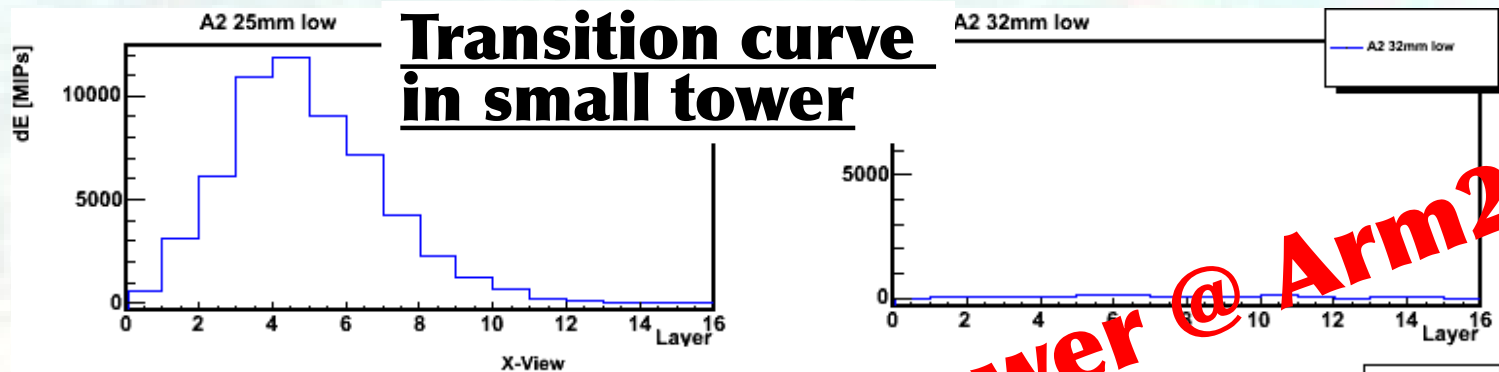
# Arm1 & Arm2 comparison



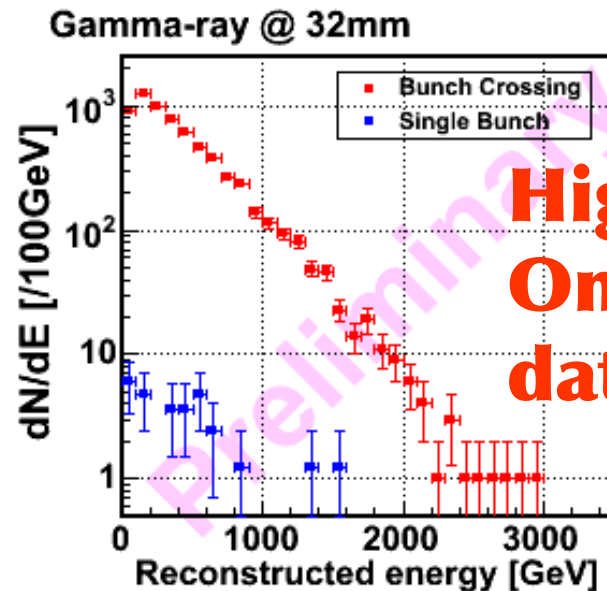
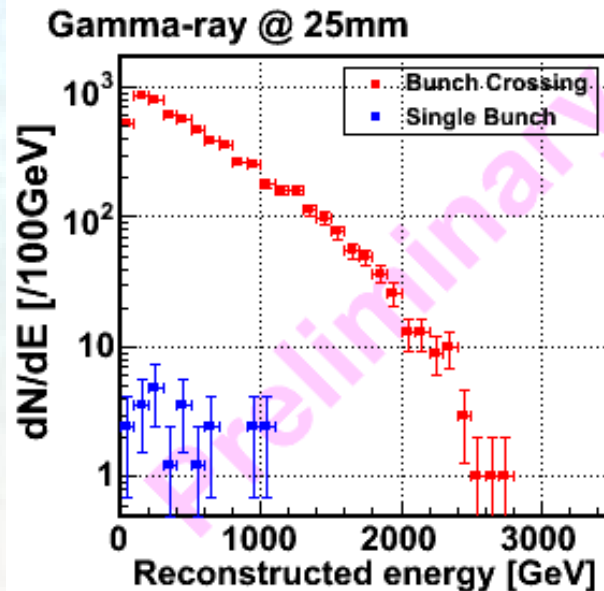
**Arm1 and Arm2  $\gamma$  spectra are normalized to the ratio of the fiducial volumes surfaces**

# = Operations in 2010 =

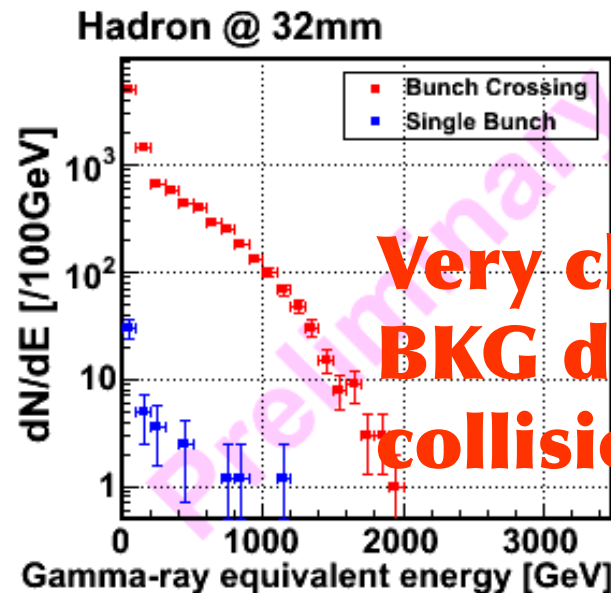
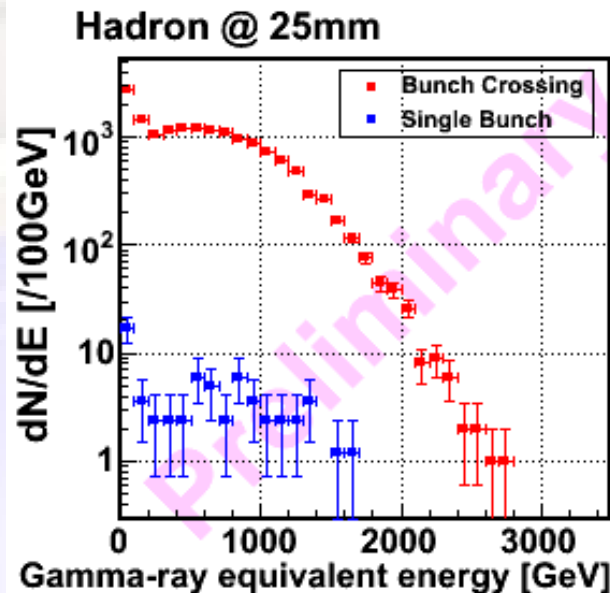
Data taking at 7TeV collisions is ongoing !!!  
Already  $7 \times 10^6$  shower events in Arm1 and Arm2  
have been collected (30<sup>th</sup> Mar. - 19<sup>th</sup> Apr. 2010)



# Very preliminary results



**High statistics !!**  
**Only 1.5% of total data are used**



**Very clean data!!**  
**BKG due to beam-gas collisions is ~ 1%**





**Which is the future of LHCf?**

# Present running and Plans for the future

- Now we are taking data at 3.5+3.5 TeV
- When luminosity will become too high ( $>10^{31} \text{ cm}^{-2}\text{s}^{-1}$ ,  $2 \text{ pb}^{-1}$ ) we will go out from the TAN (Radiation damage of the plastic scintillator is significant, LHCf has been designed to run at low luminosity/high energy!)

# Improve the radiation resistance of LHCf

## Basic idea:

- Replace the plastic scintillator with more RadHard scintillators
  - GSO
- Rearrange the order of silicon sensors to improve the silicon energy measurement
  - Cross check for scintillator measurement
- Go back in the TAN for 7+7 TeV run when it will be done
  - Removal when Luminosity will be too high
- This work will be done in the Florence INFN clean room in Summer 2010

## Chi

- **Dipendenti Università**
  - Oscar Adriani
  - Raffaello D'Alessandro
- **Dipendenti INFN**
  - Paolo Papini
- **Dipendenti CNR**
  - Guido Castellini
- **Assegnisti/Borsisti**
  - Lorenzo Bonechi
  - Massimo Bongi
  - Menjo Hiroaki
  - Sergio Ricciarini

## Dove

- Laboratori nell'edificio di Fisica Sperimentale
- Camera Pulita nell'edificio di Fisica Sperimentale

## Quando

- Attività prevista per i prossimi 3-4 anni
  - Modifiche hardware
  - Presa dati
  - Analisi dati

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