The international journal of science / 16 April 2020

THENRY OF AN INDICATION OF MATTER

symmetry violation in neutrinos

Coronavirus The models driving the global response to the pandemic Hot source Remnants of primordial nitrogen in Earth's mantle

Origin of a species Revised age for Broken Hill skull adds twist to human evolution



T2K & beyond : δ_{cp} measurements in T2K and prospects in neutrino physics in Japan.



Maria Gabriella Catanesi INFN Bari LNF 21 Maggio 2020





- Neutrino Oscillation in a "Nutshell"
- The T2K Experiment
- Data Analysis
- T2K result on "δ_{cp}"
- Medium term prospects (2022-2026)
 T2K-II
- Long term (>2026)
 - HYPER-K (T2HK)

Neutrino Oscillation

If mass and weak eigenstates are different:

- * Neutrino is produced in a weak eigenstate
- It travels a distance L as a mass eigenstate
- It will be detected in a (possibly different) weak eigenstate

 \mathcal{V}



Bruno Pontecorvo 1969

$$\rightarrow V_{\mu}, V_e \text{ or } V_{\tau}$$

$$\begin{pmatrix} v_{\mu} \\ v_{x} \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} v_{1} \\ v_{2} \end{pmatrix} P(v_{\mu} \rightarrow v_{x}) = \sin^{2}(2\theta)\sin^{2}\left(\frac{1.27\Delta m^{2}L}{E_{\nu}}\right)$$

 V_1, V_2, V_3

Evidence for neutrino oscillations (Super-Kamiokande @Neutrino '98)



Summary Evidence for Vu oscillations レルッル 90%. C.L. $\Delta M^2 \left(eV^2 \right)$ Kam up M SK UP M Cam. contain 10-210-2 Super-Kamiokande concluded that the observed zenith angle dependent deficit (and the other supporting data) gave evidence for neutrino oscillations.

Y. Fukuda et al., PRL 81 (1998) 1562

 $\Delta m^2 \sim 3 \times 10^{-3} \, eV^2$

L/E ≈ 1000 (KM/GeV)





Pontecorvo-Maki-Nakagawa-Sakata (PMNS) Matrix < 2011

$$\begin{pmatrix} v_{e} \\ v_{\mu} \\ v_{\tau} \end{pmatrix} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \times \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \times \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{pmatrix} v_{1} \\ v_{2} \\ v_{3} \end{pmatrix}$$

$$c_{ij} = \cos \theta_{ij}$$

$$atmosferic v$$

$$accelerator v$$

$$SBL reactor v$$

$$accelerator v$$

$$Interference$$

$$sin^{2}\theta_{23} \sim 1/2$$

$$\Delta m^{2} \sim 3 \times 10^{-3} eV^{2}$$

$$sin^{2}\theta_{13} \sim 0$$

$$\delta m^{2} \sim 8 \times 10^{-5} eV^{2}$$

Source of neutrino:

- Solar & Atmospheric
- Reactors
- Accelerators («Long Baseline»)



Everything changes: 2011/2012









v Experiments: LBL vs Reactors

• In LBL APP
$$\nu_{\mu} \rightarrow \nu_{e}$$

 $P_{\mu e} \simeq s_{23}^{2} \sin^{2} 2\theta_{13} \left(\frac{\Delta_{31}}{B_{\mp}}\right)^{2} \sin^{2} \left(\frac{B_{\mp}L}{2}\right)$
 $+\tilde{J} \frac{\Delta_{12}}{V_{E}} \frac{\Delta_{31}}{B_{\mp}} \sin\left(\frac{V_{E}L}{2}\right) \sin\left(\frac{B_{\mp}L}{2}\right) \cos\left(\frac{\Delta_{31}L}{2} \pm \delta_{CP}\right)$
 $B_{\pm} = \Delta_{31} \pm v_{E} \ \bar{J} = c_{13} \sin^{2} 2\theta_{13} \sin^{2} 2\theta_{23} \sin^{2} 2\theta_{12}$
So $\sin^{2} 2\theta_{APP} = 2 \sin^{2} \theta_{23} \sin^{2} 2\theta_{13}$

• In Reactor
$$P_{ee} \simeq \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta_{31} L}{2}\right)$$

So $\sin^2 2\theta_{\text{REAC}} = \sin^2 2\theta_{13}$

Concha Gonzalez-Garcia

T2K 2013



v_e appearance by T2K



28 events observed over 4.92 ± 0.55 bkgs $\rightarrow 7.3\sigma$ excess First Confirmation of 'Appearance phenomenon' w/ > 5σ significance.

Breakthrough Prize 2016



T2K & K2K	KamLAND	Daya Bay	SNO	Super-Kamiokande
Δm^2_{32}	Δm^2_{21}	Δm^2_{31}	Δm^2_{21}	Δm^2_{32}
θ_{23} θ_{13}	θ_{12}	θ_{13}	θ_{12}	θ_{23}

.... "For the fundamental contributions to the discovery of neutrino Oscillation"

Everything solved ???

PDG2018

Parameter	best-fit	3σ
$\overline{\Delta m_{21}^2 \ [10^{-5} \text{ eV}^2]}$	7.37	6.93 - 7.96
$\Delta m^2_{31(23)} \ [10^{-3} \ {\rm eV}^2]$	2.56(2.54)	2.45 - 2.69 (2.42 - 2.66)
$\sin^2 \theta_{12}$	0.297	0.250 - 0.354
$\sin^2\theta_{23},\Delta m^2_{31(32)} > 0$	0.425	0.381 - 0.615
$\sin^2 \theta_{23}, \Delta m^2_{32(31)} < 0$	0.589	0.384 - 0.636
$\sin^2 \theta_{13}, \Delta m^2_{31(32)} > 0$	0.0215	0.0190 - 0.0240
$\sin^2 \theta_{13}, \Delta m^2_{32(31)} < 0$	0.0216	0.0190 - 0.0242
δ/π	1.38(1.31)	2σ : (1.0 - 1.9)
		$(2\sigma: (0.92-1.88))$
	Parameter $ \frac{\Delta m_{21}^2 [10^{-5} \text{ eV }^2]}{\Delta m_{31(23)}^2 [10^{-3} \text{ eV }^2]} $ $ \sin^2 \theta_{12} $ $ \sin^2 \theta_{23}, \Delta m_{31(32)}^2 > 0 $ $ \sin^2 \theta_{23}, \Delta m_{32(31)}^2 < 0 $ $ \sin^2 \theta_{13}, \Delta m_{31(32)}^2 > 0 $ $ \sin^2 \theta_{13}, \Delta m_{32(31)}^2 < 0 $ $ \delta/\pi $	$\begin{array}{ c c c c c } \hline Parameter & best-fit \\ \hline \Delta m_{21}^2 \ [10^{-5} \ eV \ ^2] & 7.37 \\ \Delta m_{31(23)}^2 \ [10^{-3} \ eV \ ^2] & 2.56 \ (2.54) \\ \sin^2 \theta_{12} & 0.297 \\ \sin^2 \theta_{23}, \ \Delta m_{31(32)}^2 > 0 & 0.425 \\ \sin^2 \theta_{23}, \ \Delta m_{32(31)}^2 < 0 & 0.589 \\ \sin^2 \theta_{13}, \ \Delta m_{31(32)}^2 > 0 & 0.0215 \\ \sin^2 \theta_{13}, \ \Delta m_{32(31)}^2 < 0 & 0.0216 \\ \delta/\pi & 1.38 \ (1.31) \\ \hline \end{array}$

Most of the parameters measured with <10% precision

 θ_{23} is known with 15% precision

What is the value of δ_{CP}??
What is the mass hierarchy?

PNMS vs CKM



- What is the absolute mass scale?
- Why so small??

CKM





PNMS



Mostly diagonal

Unitarity enforced by construction

Is the PMNS parameterization correct?

The Tokai to Kamioka (T2K) Experiment

Super-K Detector





- * The T2K experiment searches for neutrino oscillations in a high purity ν_{μ} beam
- * A near detector located 280 m downstream of the target measures the un-oscillated neutrino spectrum
- * The neutrinos travel 295 km to the Super-Kamiokande water Cherenkov detector
- v_e ($\bar{v_e}$) appearance
- ν_{μ} ($\bar{\nu_{\mu}}$) disappearance
- δ_{cp}

• X-sections +exotics

J-PARC Accelerator





T2K Collaboration



<u>T2</u>





- 30 GeV proton beam generated by J-PARC Main Ring (MR) directed to the graphite target
- Secondary pions collected and focused by the magnetic horns
 - ν beam: $\pi^+ \rightarrow \mu^+ + \nu_{\mu}$ (Forward horn current)
 - $\overline{\nu}$ beam: $\pi^- \rightarrow \mu^- + \overline{\nu}_{\mu}$ (Reverse horn current)
- Uses off-axis method to make the spectrum peak at 600 MeV
 Expected oscillation maximum at L=295 km
 - near the flux peak. The uncertainty on the ratio of the flux predictions at the far and near detectors Beam Dump Near Detectors Decav Primary 3 Horns protons Beamline in energy. Pha of the neutrino beam car varied by changing the off-axis angle as illustrated in lower panel of Fig. 1. In the case utrino flux and en angle is set at 2.5° so that the neutrin important component of analyses in accelerator neutrino a pealement at about 0.6 GeV, near th experiments [1-4]. However, it is difficult to simulate oscillation maximum (Fig. 1). This maximizes the el of the neutrino oscillations at 295 km as well as redu the flux precisely de itons certainties in the underly-ing physical processes, particularly hadron production Muon Monitor in proton-nucleus in united Sis. To reduce flux-related background events. Since the energy spec um cha depending on the off-axis uncertainties, neutrino schlation experiments are some-times conducted by comparing measurements between a rection has to be precisely



J-PARC neutrino beamline overview





T2K Near Detector pit houses both the off-axis (ND280) and on-axis (INGRID) detectors





On-Axis near Detector



±5m

On-axis Detector (INGRID)







Off-axis near detector:ND280





Event number : 27404 | Run number : 8115 | Spill : 51004 | Time : Mon 2012-01-23 06:04:28 JST [Trigger: Beam Spill



- <u>UA1 magnet (0.2 T)</u>
- Fine Grained Detector (FGD) (target)
- <u>Time Projection Chambers (TPC)</u> Momentum Res. < 10% PID (< 10% dE/dx Res.)

•ND280 @ 2.5 degree off-axis

- Normalization of Neutrino Flux
- Measurement of neutrino cross sections.





T2K-ND280: INFN Contributions (Ba, LNL, Na, Pd, Rm1)



"Inclusive CC anti-v Differential Cross-Section on Carbon "





First large TPC with MPGD





INFN in T2K

- TPC design, assembling, calibration, maintenance and operation
- Initial idea and calculations for a magnetized detector
- ✓ Leading role in anti-v Analysis @ ND280
- Several Management & Coordination Roles

TPC assembling

TPC design with advanced detectors (MPGD)

ND280 off-axis event gallery Tzk



vA cross-sections



Large discrepancy in the models predictios



Cross sections poorly known at low energy





A couple of examples : many papers published in the last years



* The Far detector: Super-Kamiokande



Probability to mis-id μ as electron is ~1%.

- 50 kton Water Cherenkov detector 1 km underground
- ^{*} Typically 61% v_e signal eff.
- 95% π_0 rejection
- ⁶ 32 kton inner volume (22.5 kton fiducial volume)
- ^{*} 2 m outer volume to identify entering particles

All triggers in +/- 0.5 ms of neutrino arrival time are recorded

Data Set





Beam model

Beam model is obtained from a full GEANT simulation of the particle transport reweighed by the NA61 (Shive) results



* ND280 data for Oscillation Analysis

- 14 total ND280 data samples used by oscillation analysis fit
- v-mode (FHC)
 - sort by π^{\pm} multiplicity
 - 2 FGDs (C,O)
 - ➡6 samples
- $\overline{\nu}$ -mode (RHC)
 - sort by muon charge
 - sort by number of tracks
 - 2 FGDs (C,O)









Neutrino mode





e-like



Anti-neutrino mode

ents/100 MeV Ras5.9 Data (1635-10¹⁰ POT) 7,0006 V, CC QE 20 v.+7. (1' nm-OE ¥,+₹,CC MC w T2K-DB bestfit 5 Number 1000 2000 000 Reconstructed v energy (MeV) ents/125 MeV - Rand-9 Data (16.35-30²⁰ POD) 0sc. W, CC Osc.v, CC ¥,/7, CC Beamy /T, CC - NO 5 MC w/T2K+DB besili 5 Number 500 1000 Reconstructed v energy (MeV)

No CC1 π sample in antineutrino mode because π^- produced in $\bar{\nu}$ interaction are mostly absorbed before decay.



5 SK samples

- 1. muon neutrinos
- 2. muon antineutrinos.
- 3. electron neutrinos.
- electron neutrino + pion (Michel electron)
- 5. electron antineutrino.





*Oscillation fits

 $v_{\mu} \rightarrow v_{e}$ and $\bar{v}_{\mu} \rightarrow \bar{v}_{e}$ combined analysis within the 3v oscillation paradigm (PMNS).

Other oscillation parameters from 2018 PDG values.

$$-2 \ln \lambda(\overline{\delta_{CP}}; \boldsymbol{a}) = 2 \sum_{i=1}^{N} \left[n_i^{\text{obs}} \ln \left(\frac{n_i^{\text{obs}}}{n_i^{\text{exp}}} \right) + n_i^{\text{exp}} - n_i^{\text{obs}} \right] \\ + (\boldsymbol{a} - \boldsymbol{a}_0)^T \mathbf{C}^{-1} (\boldsymbol{a} - \boldsymbol{a}_0)$$

Binned likelihood comparing data to MC predictions.

Bins of reconstructed energy from lepton kinematics assuming CCQE two body interactions.

 v_e sample also bins in θ_e

Bayesian Markov Chain MonteCarlo and 2 frequentist approach.

Frequentists confidence intervals (grid search) agree with the Bayesian factors and credible intervals.



<u>T2K highlights</u>





Word best measurements in disappearance mode



 Θ_{13} measurements (appearance mode) by T2K

Ϋμ & **Ϋ**e

160 140



ν_µdisappearance (best word result)



0.25

0.2

No ν_e appearance disfavoured to 2.4 σ

*CP violation phase



c	1e0de v-mode	1e0de $\bar{\nu}$ -mode	1e1de v-mode
$v_{\mu} \rightarrow v_{e}$	59.0	3.0	5.4
$\bar{\nu}_{\mu} ightarrow \bar{\nu}_{e}$	0.4	7.5	0.0
Background	13.8	6.4	1.5
Total predicted	73.2	16.9	6.9
Systematic uncertainty	8.8%	7.1%	18.4%
Data	75	15	15

v_e/v_e Systematic Uncertainty

Type of Uncertainty	$\nu_e/\bar{\nu}_e$ Candidate Relative Uncertainty (9		
Super-K Detector Model	1.5		
Pion Final State Interaction and Rescattering Model	1.6		
Neutrino Production and Interaction Model Constrained by ND280 Data	2.7		
Electron Neutrino and Antineutrino Interaction Model	3.0		
Nucleon Removal Energy in Interaction Model	3.7		
Modeling of Neutral Current Interactions with Single γ Production	1.5		
Modeling of Other Neutral Current Interactions	0.2		
Total Systematic Uncertainty	6.0		



δ_{CP} measurement

Fit uses the value of θ_{13} from reactor experiments

Data also prefers Normal Hierarchy with a posterior probability of 89%



*CP violation phase



CP violation phase

T2K result excludes most of the δ_{CP} >0 values @ 99.7% CL



Technically not a *discovery* (it is not 5 sigma), but the first step in the long path towards the measurement of leptonic CP violation



*CPV: what's next ?



Before HYPER-K/DUNE Era : 2022-2026
*J-PARC neutrino beam upgrade

- Continuous upgrade of neutrino beam up to 2030
- Present beam power ~470 kW
- New MR power supply for 750kW by 2021
- Repetition rate increase to 0.86 Hz for 1.3MW by 2026

Continuous beam upgrade @ J-PARC



J-PARC upgrade for Hyper-K is top priority in KEK Project Implementation Plan (KEK-PIP)

Strong commitment for future neutrino program

*Sources of systematic errors

- * Different Acceptance Near/Far detector
- * Cross sections poorly known at low energy (in particular ν_e, anti- ν_µ and ν_e/ν_µ ratio)
 * Different target material Near/Far (CH/H₂0)
 * Models









- \blacktriangleright Designed to improve systematics (from 6% => 3-4%)
- > **2019-2021** Production, integration at CERN. System test (cosmics).
- > **2021-22** Shipment to Japan, installation, commissioning.

*NP280 upgrade







ND280 upgrade goals

- quasi-3D imaging.
 - Improved target tracking.
 - Improved proton detection threshold.
 - neutron detection capabilities
- Improved high angle acceptance:
 - High Angle TPC's.
- x 2 in statistics for equal p.o.t.
- Time of Flight for background reduction.

INFN Contribution to T2K-II & beyond



INFN Role in T2K-II

- ✓ Coordination of the TPC Project
- ✓ Field Cages construction (2018-2021) (0.5 MEur investment)
- ND280 Installation, maintenance & operation (2022 & beyond)

 We are interested in future ND280 upgrades for HK



TPC prototype



MM



Test-Beam event

*SK-X with Gd

- Loading Gd to Super-K (SK-Gd) to significantly enhance neutron detection capabilities.
- Aiming for the first detection of Supernova Relic Neutrinos (SRNs)
- Also aiming for many new measurements with T2K beam:
 - Neutron multiplicity and kinematics measurements from neutrino interactions
 - Improved oscillation measurements with neutrinoantineutrino separation and further background rejection
 - Non-standard oscillation searches with additional Neutral-Current samples with neutrons



SK-V with GD



Plan to start 0.01% Gd run in

(Adjusting schedule with T2K)

early 2020.

0.1%Gd run ~90% n cap. eff.

INFN is involved in SK-V

*Looking forward

* Hyper-Kamiokande project Covering both accelerator and non-accelerator physics

	4mφ ble- sitivity τ Ο Ο Ο Ο Ο Ο Ο Ο Ο Ο Ο Ο Ο	entioka 295km entioka entioka entioka entioka entioka entioka	
	Super-K	Hyper-K (Ist wnk)	
Site (depth)	Mozumi (1000 m)	Tochibora (650 m)	
Number of ID PMTs	11,129	40,000	
Photo-coverage	40%	40% (×2 sensitivity)	Design Report arXiv:
Mass / Fiducial Mass	50 kton / 22.5 kton	260 kton / 188 kton	1805.04163

- 1. Hyper-K detector with **8.4 times larger fiducial mass** (190 kiloton) than Super-K and using **double-sensitivity PMTs**
- 2. J-PARC neutrino beam to be upgraded from 0.5 to **1.3 Mega Watt**
- 3. New and upgraded **Near Detectors** to control systematic errors

* Hyper-K: a multi pourpose Experiment

* Neutrino oscillation physics

- * CP violation
- $^{*}\Theta_{23}$ octant determination
- * Mass hierarchy with beam and atmospheric ν's

* Nucleon decay discovery potential

* Possible discovery with ~×10 better sensitivity than Super-K

* Neutrino astrophysics

- * Precision measurements of solar v
- ^{*} High statistics measurements of SN burst ν
- Detection and study of relic SN neutrinos
- *Unexpected....



Extend highly successful program of Super-K

* Physics with J-PARC neutrino beam

- Reliable sensitivity estimate based on T2K Experience directly applicable
- \succ 1,000-2,000 events for both of v_e and \bar{v}_e
- appearance
- cf. current T2K: 90 and 15 events Firm discovery and measurement of CP violation in neutrino oscillation
- Precision disappearance measurements



1.3MW×10years(10⁸s)

Reconstructed Energy

Appearance v mode Number of events/50 MeV 250 Total Signal $v_{\mu} \rightarrow v_{e}$ 200 Signal $\overline{v}_{..} \rightarrow \overline{v}_{..}$ BG Reconstructed Energy E^{re} Appearance \overline{v} mode Number of events/50 MeV Total 200 Signal $v_{\mu} \rightarrow v_{\mu}$ Signal V BG BG 100

*Non-beam physics

- * Very broad science targets!
- * Nucleon decay searches
- * Atmospheric neutrinos
- * Solar neutrinos

....

* Supernova neutrinos
 * Indirect dark matter search







* Newly developed 50cm PMT



Higher pressure tolerance (> 80m)

Charge [photoelectron]21

* HYPER-K DETECTOR: HYBRID PMT CONFIGURATION

Possible hybrid configuration with 20k 20"PMTs and 5k mPMTs



Complement 20" PMTs:

 Max number of mPMTs limited by the production sites.
 Better directionality, granularity and timing resolution , less magnetic field sensitivity, pressure tolerance

* Multi PMT Modules

Derived from KM3NET mPMTs. ~5,000 mPMTs are now included in the HK baseline design.
 Proposed and designed by the Italian groups



- ✓ 3" PMTs are Hamamatsu R14374 (or ETEL D794KFL or HZC XP82B20)
- ✓ Improved timing, spatial accuracy, dynamic range, lower dark rate with +HV



19 x 3" PMTs + reflectors

* Near Detectors for Hyper-K



- Upgrade of existing (T2K) detectors + new IWCD
- Started discussion of possible INGRID/ND280 upgrade for HK
- Keeping ND280 infrastructure for HK era is already a big challenge...

* IMPACT ON SENSITIVITY



- To achieve 5σ for ~60% of δ_{cp} need **4% error on relative cross section**
- Other sources may introduce error on the metry
 - Expect signal cross section error to dominate budget
 - Keep other sources to the 1% level is possible

*Project timeline



Hyper-K was approved in 2019 Construction from 2020 \rightarrow Operation from 2027



Hyper-K vs Pune

Experiment	Status	E _v (GeV)	L (Km)	E/L (eV ²)	ν beam	v type
DUNE 40KT Liquid Argon	Future (2026)	5	1300	3.8x10 ⁻³	Fermilab newbeam	ν_{μ} /anti- ν_{μ}
HYPERK 190KT WC	Future (2026-27)	0.6	295	2x10 ⁻³	KEK J-PARC (improved)	ν_{μ} /anti- ν_{μ}



Matter effects are large (big sensitivity on MO), all the neutrino interaction modes contribute (quasi-elastic, resonances, deep-inelastic). The Liquid Argon far detectors can precisely reconstruct the event energy for all the topologies



Matter effects are small (optimized for CP), neutrino interaction are mostly quasi-elastic. The water Cherenkov far detector can precisely reconstruct quasi-elastics and is very massive.

Great complementarity also in many astrophysics measurements





**HyperK 10 y, Dune 7y full conf.



Thanks !





	Predicted rates				Observed
Sample	$δ_{CP} = -π/2$	$\delta_{CP} = 0$	δCP = π/2	δCP = π	Events
CCQE 1-Ring e-like v	74.46	62.26	50.59	62.78	75
CCQE 1-Ring mu-like v	272.34	271.97	272.30	272.74	243
CC1pi 1-Ring e-like v	7.02	6.10	4.94	5.87	15
CCQE 1-Ring e-like anti-v	17.15	19.57	21.75	19.33	15
CCQE 1-Ring mu-like anti-v	139.47	139.12	139.47	139.82	140

* Nova vs T2K











Combined analysis

Sensitivity T2K+Nova



Sensitivity T2K+SK



T2K formed working groups with Nova and SK to provide combined oscillation analyses

*SK-GD: Impact for oscillation analysis

- Improvements for neutrino oscillation study:
 - CP violation search with neutrino-antineutrino separation with neutrons.
 - v_µ disappearance measurement with reduced NC or CC-nonQE events
- Would provide important cross- check for neutrino interaction systematics and background assumptions.







Reconstructed neutrino energy (MeV)



*CPV: what's next ?



High Power J-PARC Secondary Beamline

J-PARC secondary beamline infrastructure (shielding, decay volume, hadron absorber) were all designed for 3–4 MW

Component	Limiting Factor	Current	Upgraded
		Acceptable Value	Acceptable Value
Target	Thermal Shock	$3.3 imes10^{14}$ ppp	$3.3 imes10^{14}$ ppp
	Cooling Capacity	0.75 MW	>1.5 MW
Horn	Conductor Cooling	2 MW	2 MW
	Stripline Cooling	0.54 MW	>1.25 MW
	Hydrogen Production	1 MW	>1 MW
	Operation	2.48 s & 250 kA	1 s & 320 kA
He Vessel	Thermal Stress	4 MW	4 MW
	Cooling Capacity	0.75 MW	>1.5 MW
Decay	Thermal Stress	4 MW	4 MW
Volume	Cooling Capacity	0.75 MW	>1.5 MW
Beam	Thermal Stress	3 MW	3 MW
Dump	Cooling Capacity	0.75 MW	>1.5 MW
Radiation	Radioactive Air Disposal	1 MW	>1 MW
	Radioactive Water	0.5 MW	0.75 \rightarrow 1.3 or 2 MW

J-PARC Secondary Beamline Upgrades

However, need upgrades to improve cooling capacity, radiation containment, and irradiated cooling water disposal for $1+\ MW$

Component	Limiting Factor	Current	Upgraded
		Acceptable Value	Acceptable Value
Target	Thermal Shock	$3.3 imes10^{14}$ ppp	$3.3 imes10^{14}$ ppp
	Cooling Capacity	0.75 MW	>1.5 MW
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Decay	Thermal Stress	4 MW	4 MW
Volume	Cooling Capacity	0.75 MW	>1.5 MW
Beam	Thermal Stress	3 MW	3 MW
Dump	Cooling Capacity	0.75 MW	>1.5 MW
Radiation	Radioactive Air Disposal	1 MW	>1 MW
	Radioactive Water	0.5 MW	$0.75 \rightarrow 1.3 \text{ or } 2 \text{ MW}$

* hyper-kamiokande sensitivity

arXiv:1805.04163 [physics.ins-det]

Using neutrino beam



X-section measurements in T2K



Nuclear Effects



Diagrams by Patrick Stowell

Neutrino beams for precision physics: the ENUBET Project

- The next generation of short baseline experiments for cross-section measurement and for precision v physics at short baseline (e.g. sterile neutrinos and NSI) should rely on:
- a high precision, direct measurement of the fluxes
 a beam covering the region of interest from
 - sub- to multi-GeV a narrow band beam where the energy is
 known a priori from the beam width

the ENUBET facility fulfills simultaneously all these requirements



Enhanced NeUtrino BEams from kaon Tagging

ERC-CoG-2015, G.A. 681647 (2016-21)

A. Longhin, INFN

CERN-EoI: 41 physicists, 10 institutions:

CERN, IN2P3 (Bordeaux), INR, INFN (Bari, Bologna, Insubria, Milano-Bicocca, Napoli, Padova, Roma-I + **NUTECH** funding from the Italian Min. of Research (MIUR)



Reference parameters: **100 m baseline, 500 t detector** (e.g. ICARUS@FNAL or Protodune-SP/DP@CERN)

Intermediate Water Cherenkov Detector (IWCD)

- Movable water Cherenkov detector to sample different off-axis angle beams
 - Measure E_ν-(p,θ) relation independent of neutrino interaction models
 - Measure v_e and neutrons
- Smaller detector → utilize finer granularity with mPMT modules





Near Detectors (High Pressure TPC)

2) Add new detectors in the 280m pit: High pressure TPC to study low momentum final state particles and in particular resolve vertex HPTPC detector design to reduce cross-sections systematics





Significant discrepancies on proton multiplicity and momentum distributions
 Need low momentum thresholds to reduce cross-sections systematics
 Important difference lie below threshold for liquid detectors



T2K has pioneered (atm. pressure) gas TPCs for accelerator neutrinos
Need a path to high pressures for sufficient statistics

- •Generally applicable to next generation LBL
- experiments

Gas versus liquid argon

liquid Ar


CC events assuming a 8m ³ detector & full FV.		
2x2x2 m ³ 20°C	5 bars	10 bars
He	6.65 kg	13.3 kg
	520 evt/10 ²¹ pot	1040 evt/10 ²¹ pot
Ne	32.5 kg	67.1 kg
	2543 evt/10 ²¹ pot	5086 evt/10 ²¹ pot
Ar	66.5 kg	133 kg
	5203 evt/10 ²¹ pot	10406 evt/10 ²¹ pot
CF ₄	146.3 kg	293 kg
	11450 evt/10 ²¹ pot	22893 evt/10 ²¹ pot
F.Sanchez, Jennifer meeting 22 th September 2016, London		