

LUCA LATRONICO, INFN TORINO

INFN LNF WEBINAR

30 APRIL 2020

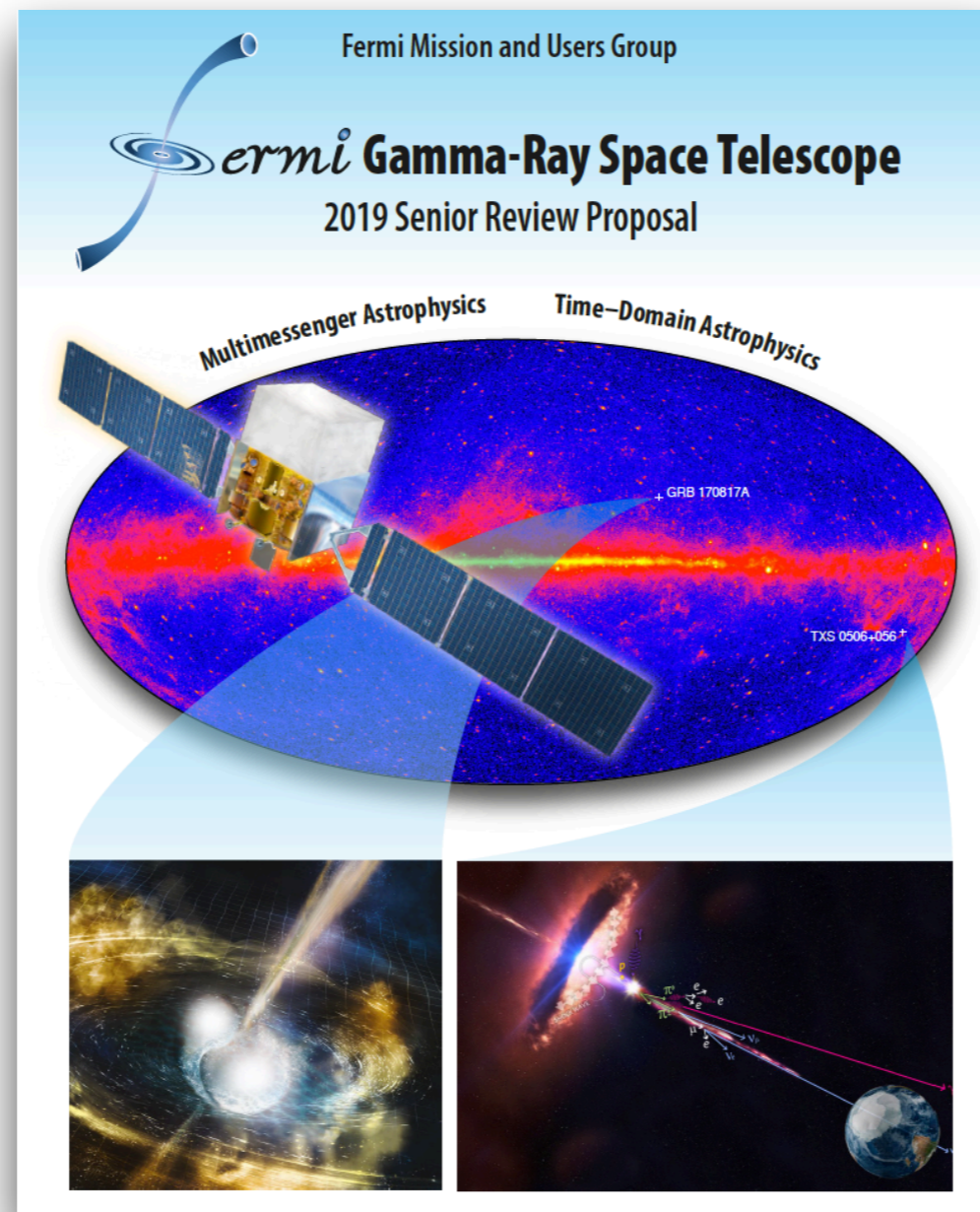


image credit Paz Beniamini

FERMI AND MULTI- MESSENGER OBSERVATIONS

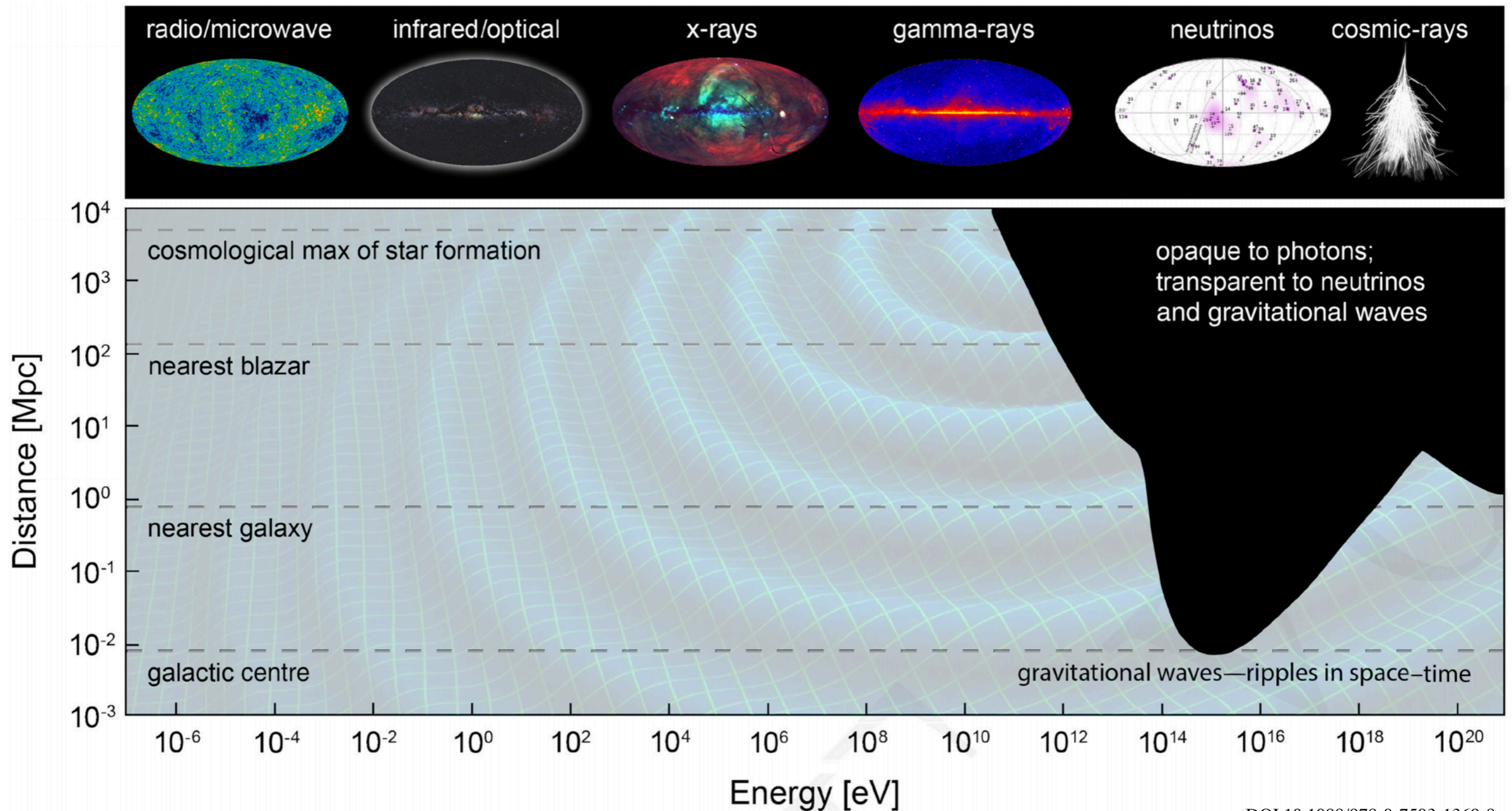
FERMI AND MULTIMESSENGER OBSERVATIONS

1 – BIRTH OF MULTI MESSENGER ASTRONOMY



FERMI AND MULTIMESSENGER OBSERVATIONS

MESSENGERS



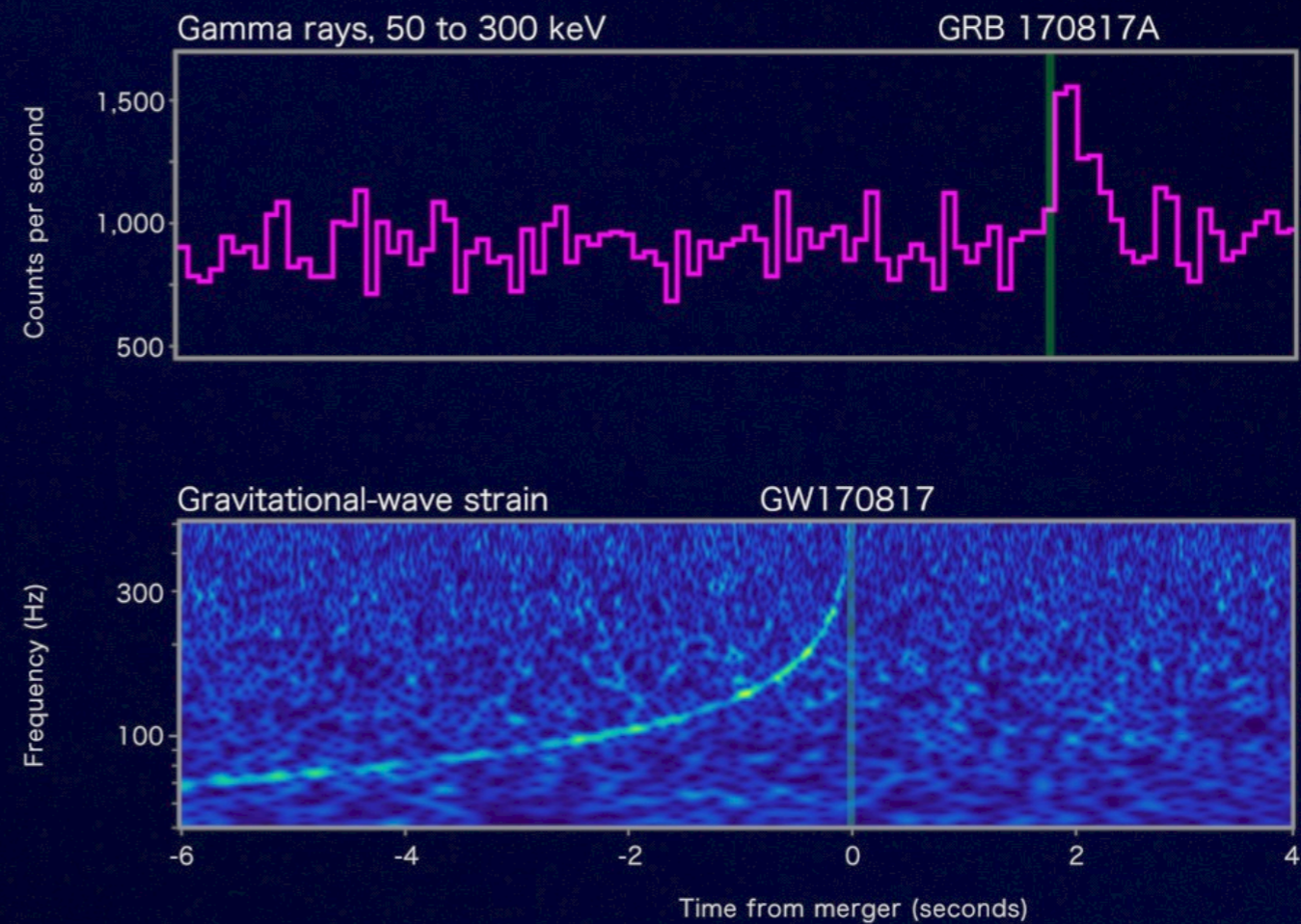
FERMI AND MULTIMESSENGER OBSERVATIONS

17 AUGUST 2017 – WAKE UP!

Fermi

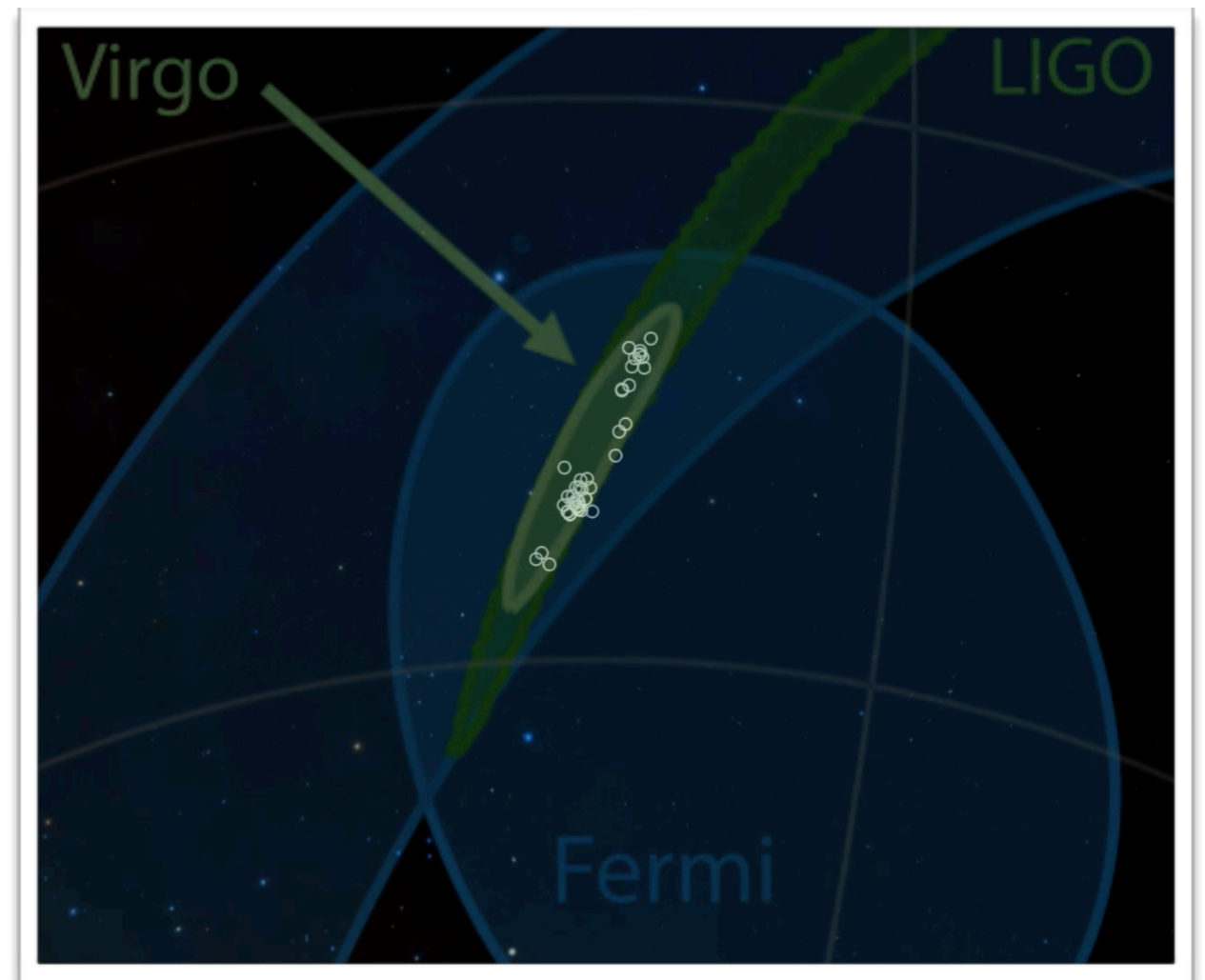
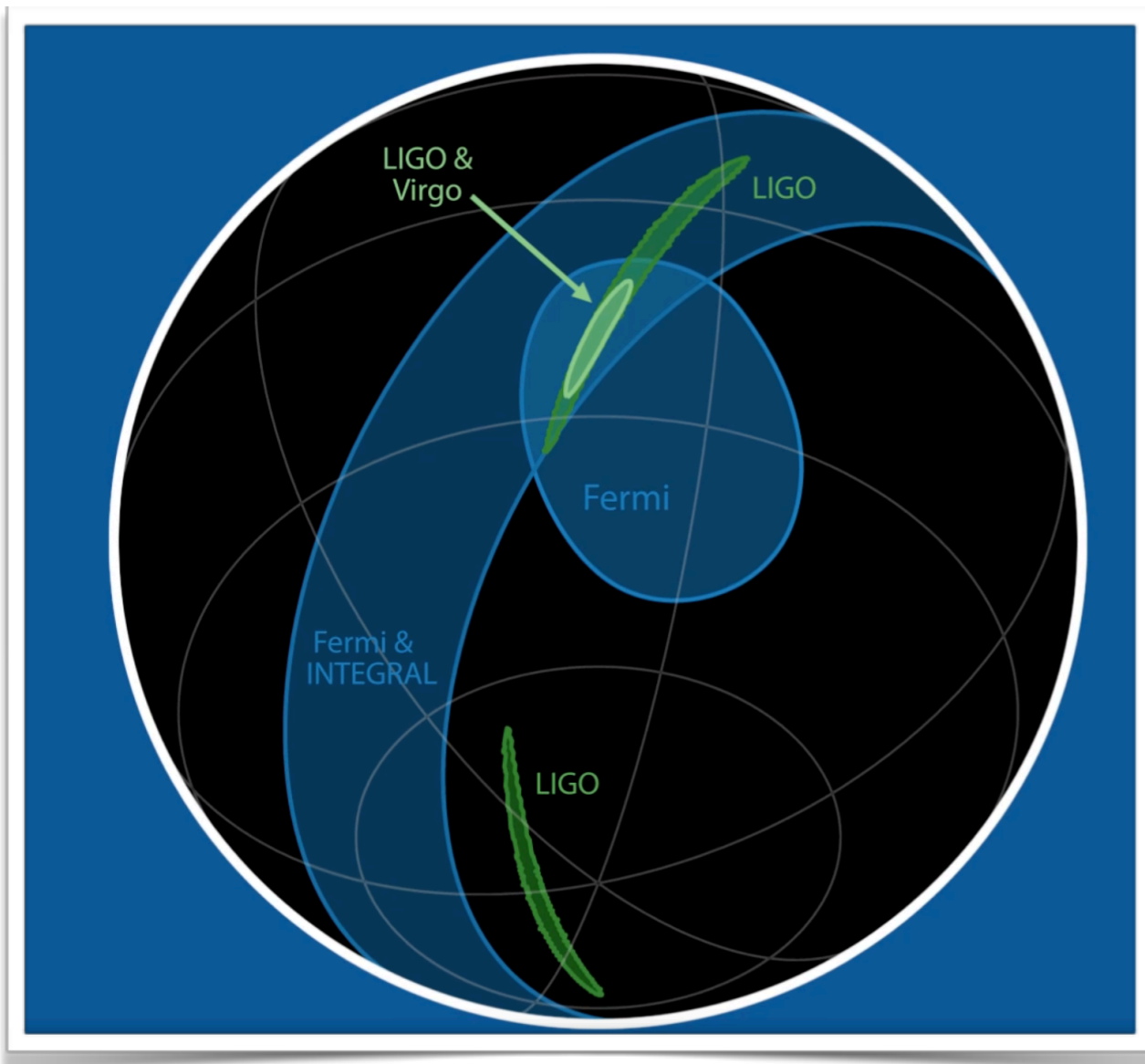
Reported 16 seconds
after detection

LIGO-Virgo

Reported 27 minutes
after detection

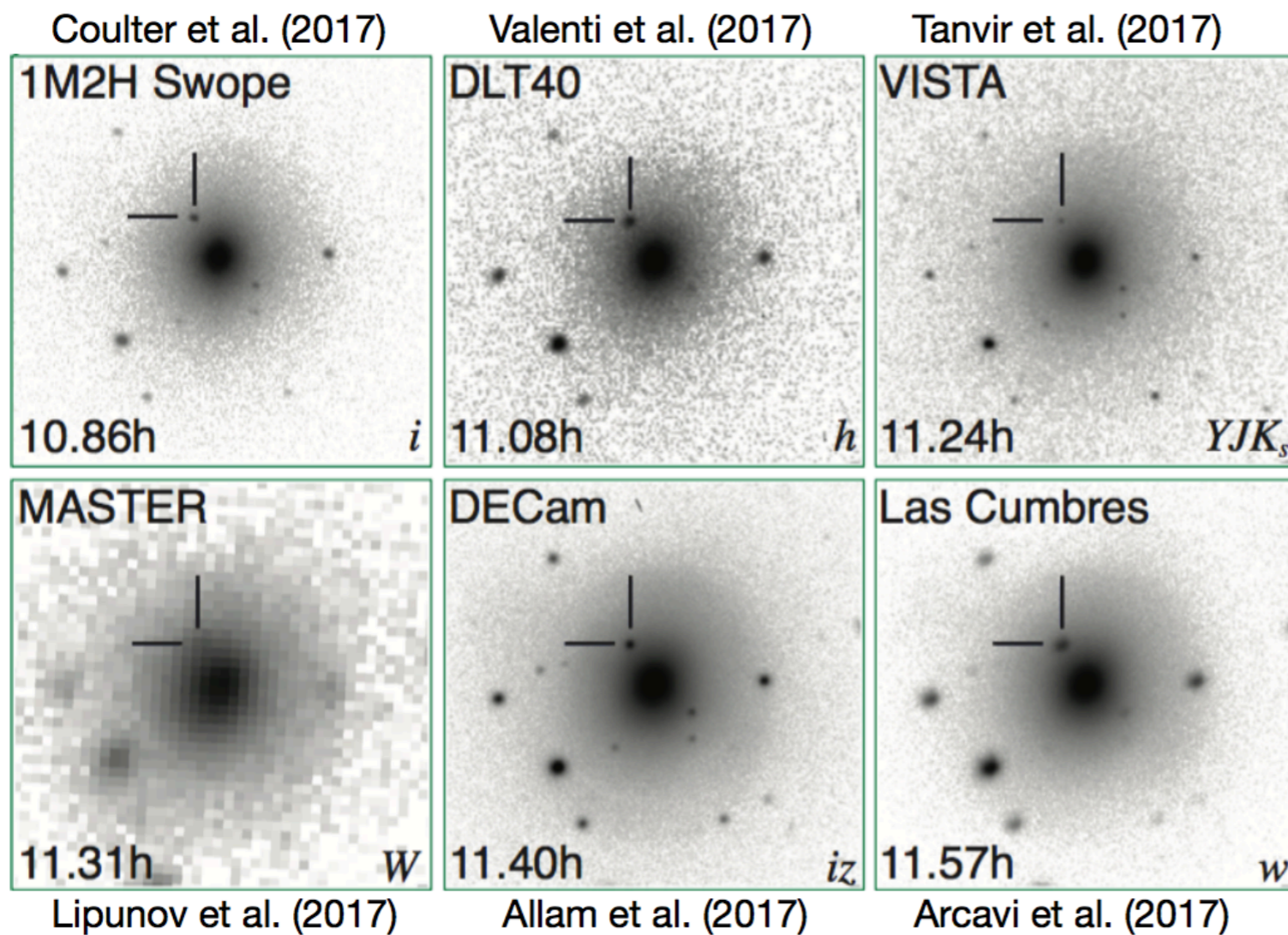
FERMI AND MULTIMESSENGER OBSERVATIONS

LOCALIZATION IN THE SKY



FERMI AND MULTIMESSENGER OBSERVATIONS

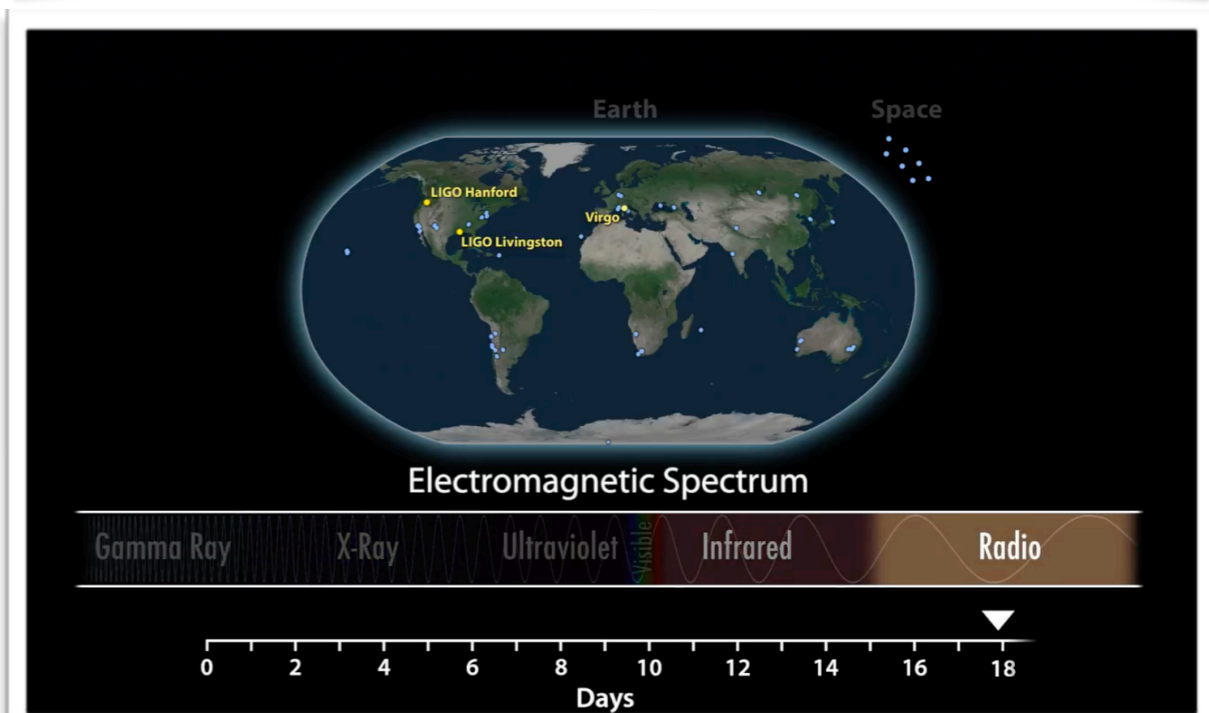
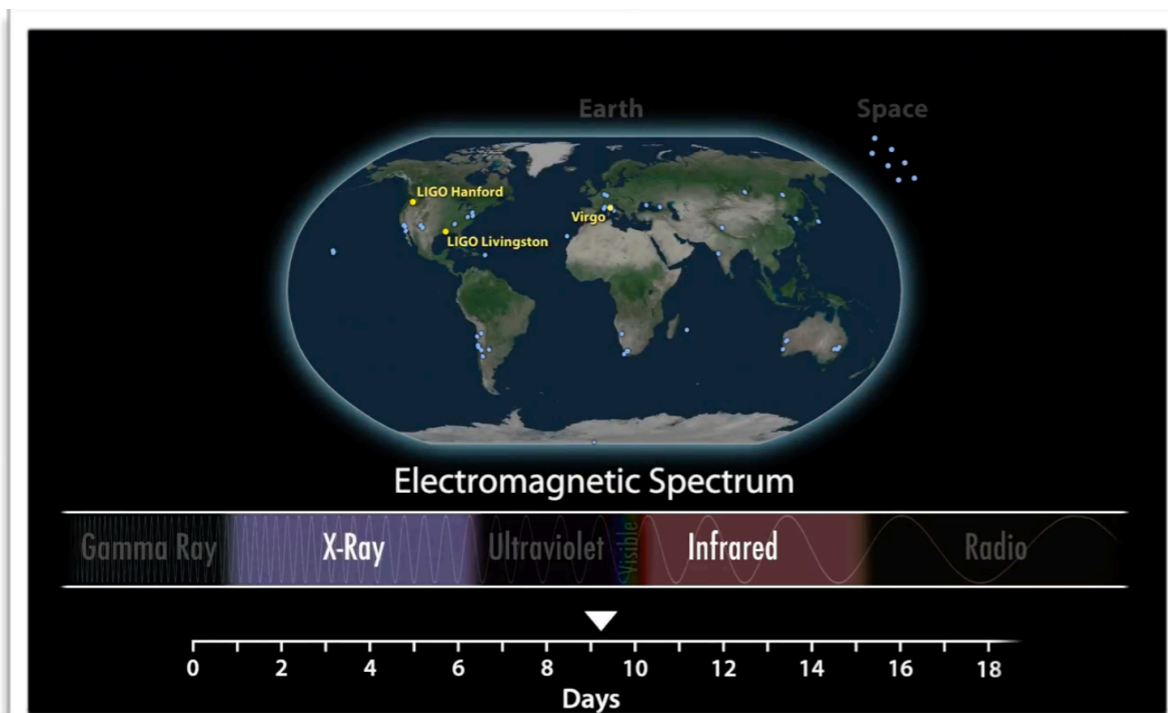
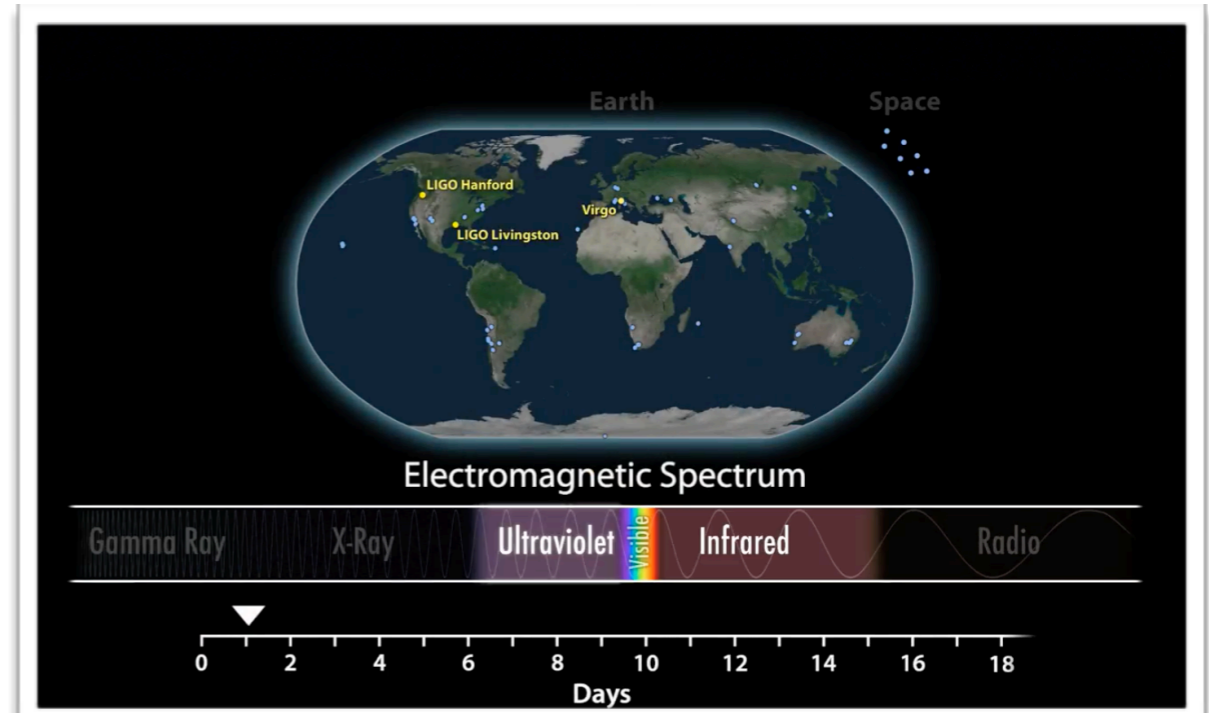
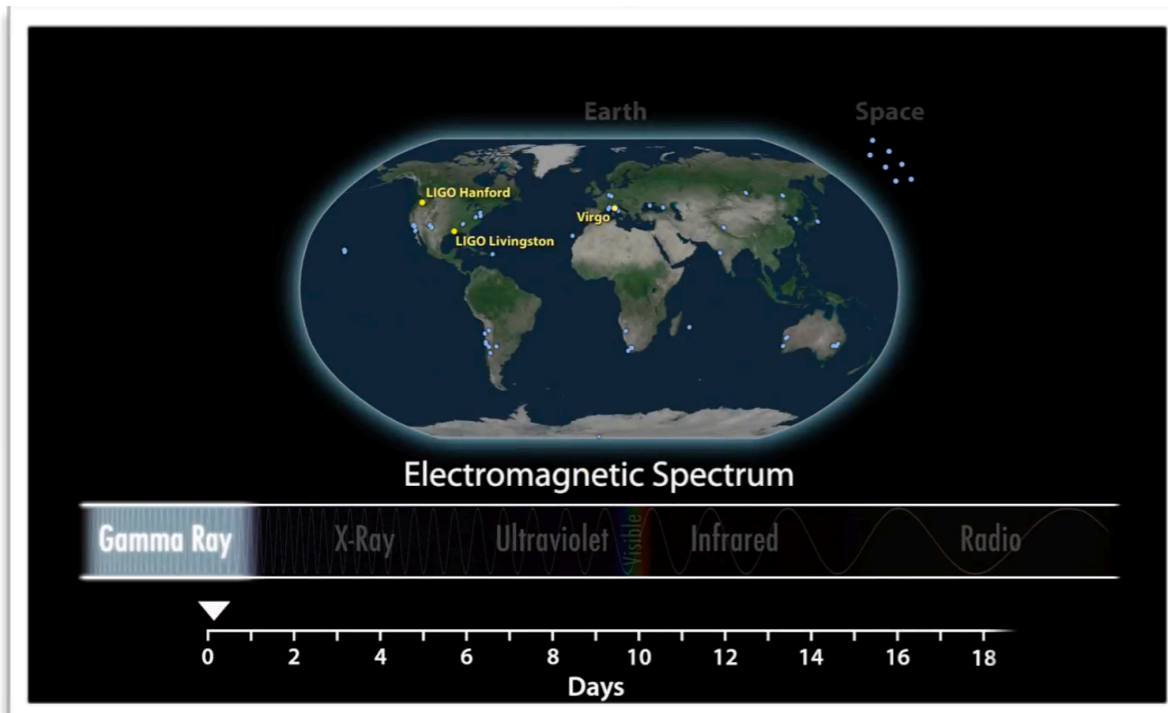
HOST GALAXY IDENTIFICATION VIA OPTICAL OBSERVATIONS



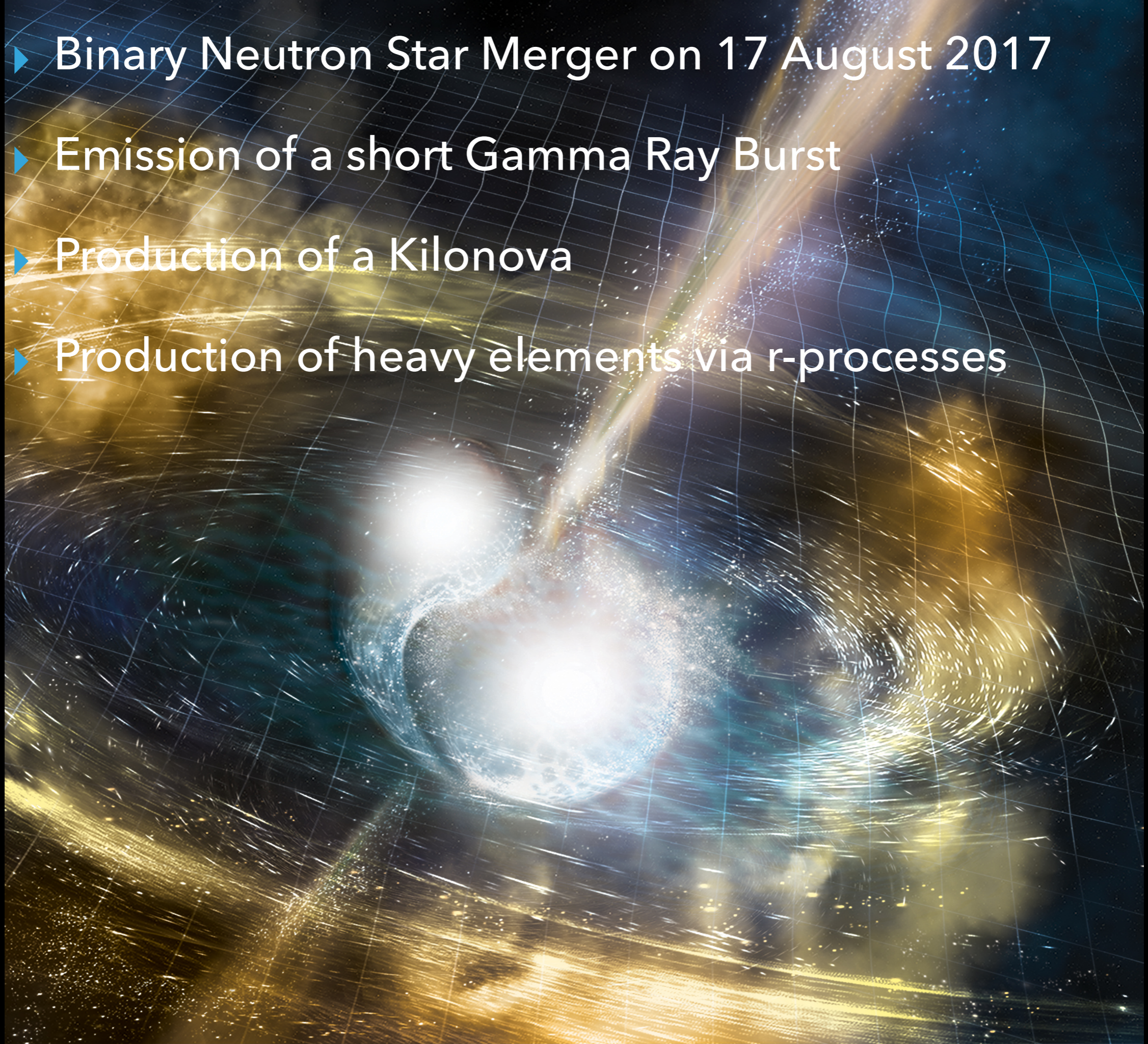
NGC 4993, S0 galaxy @ $D = 41$ Mpc, $z = 0.00968$ (Hjorth et al. 2017)

FERMI AND MULTIMESSENGER OBSERVATIONS

MULTIFREQUENCY OBSERVATIONS

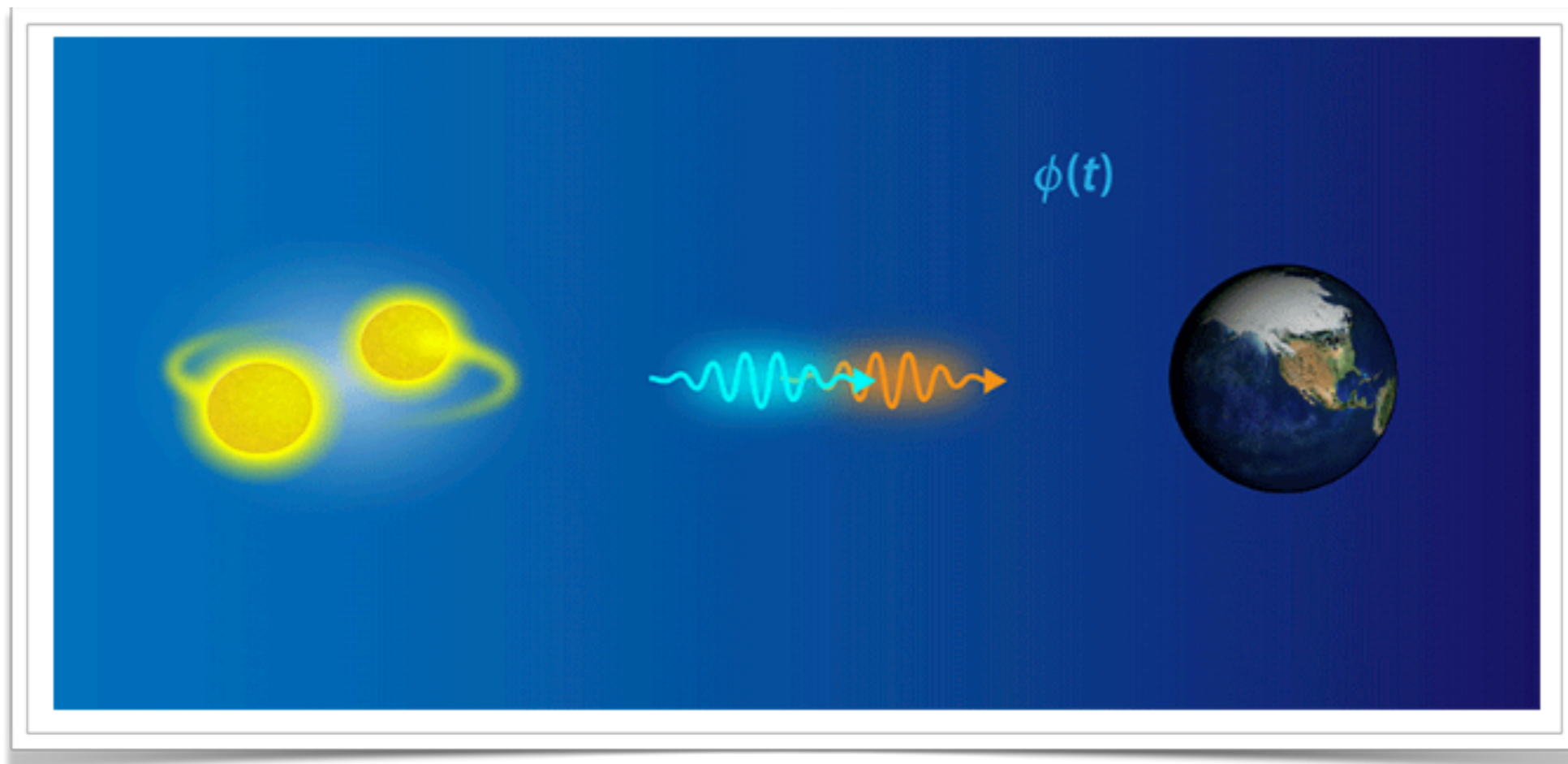


- ▶ Binary Neutron Star Merger on 17 August 2017
- ▶ Emission of a short Gamma Ray Burst
- ▶ Production of a Kilonova
- ▶ Production of heavy elements via r-processes



FERMI AND MULTIMESSENGER OBSERVATIONS

FUNDAMENTAL PHYSICS IMPLICATIONS



- ▶ GW and photons speed = c with $\sim 1/10^{15}$ error (2s/130Mly)
 - ▶ Lorentz Invariance and Equivalence Principle verified (ApJL 848:L13, 2017)
 - ▶ limits to alternative cosmology other than GR + cosmological constant (<https://physics.aps.org/articles/v10/134>)

FERMI AND MULTIMESSENGER OBSERVATIONS

19 SEPTEMBER 2017



FERMI AND MULTIMESSENGER OBSERVATIONS

REMARKS

- ▶ Observatories
 - ▶ complex systems, decades long operations
- ▶ Multiple experimental techniques and operating environments
 - ▶ multi-decades R&D programs
- ▶ Complementary communities
 - ▶ Particle physics, Astrophysics, Cosmology
- ▶ Open questions
 - ▶ what are the sources of Cosmic Rays ? how are they accelerated ?
 - ▶ what is the nature of Dark Matter ?

FERMI AND MULTIMESSENGER OBSERVATIONS

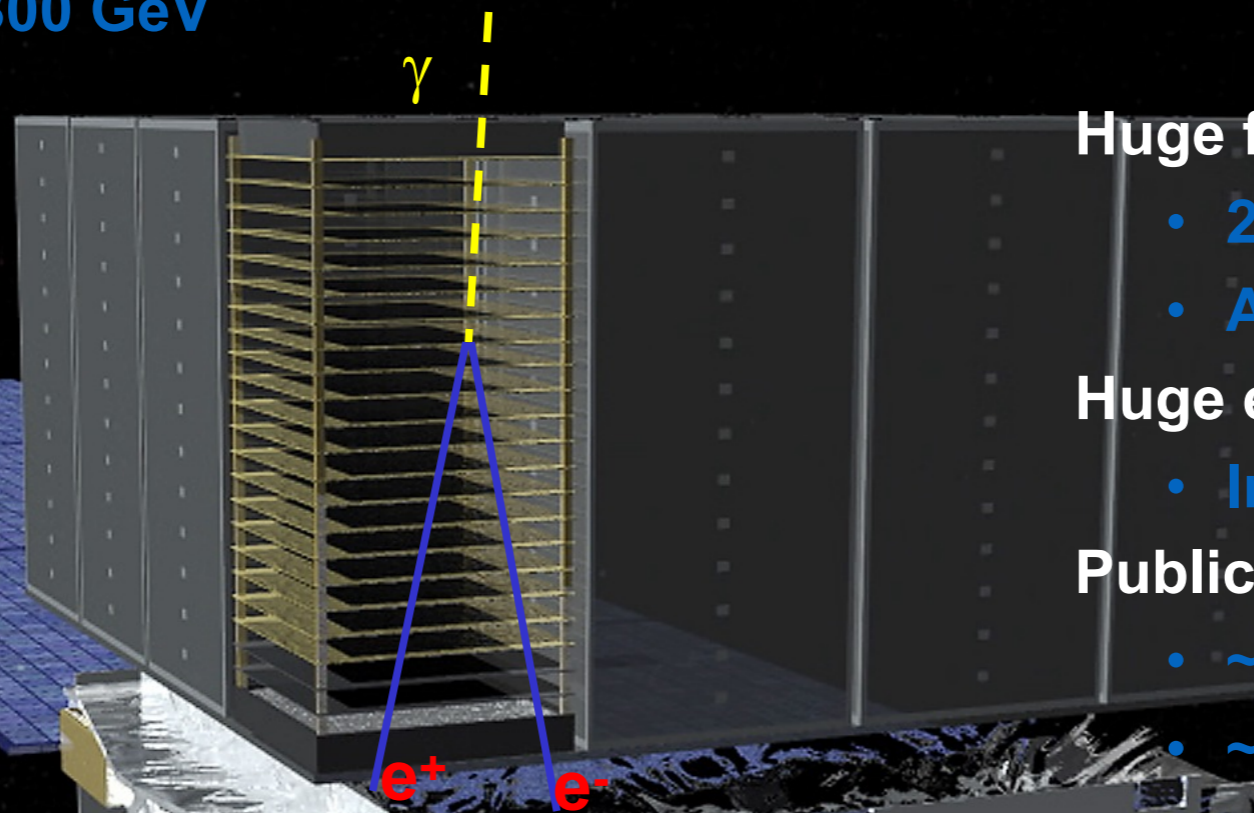
2 – THE FERMI OBSERVATORY



The Fermi Observatory

Large Area Telescope (LAT) - pair conversion telescope

- 20 MeV – > 300 GeV



Huge field of view (2.4sr)

- 20% sky any instant
- All sky for 30' every 3h

Huge energy range

- Including 10-100 GeV

Public data

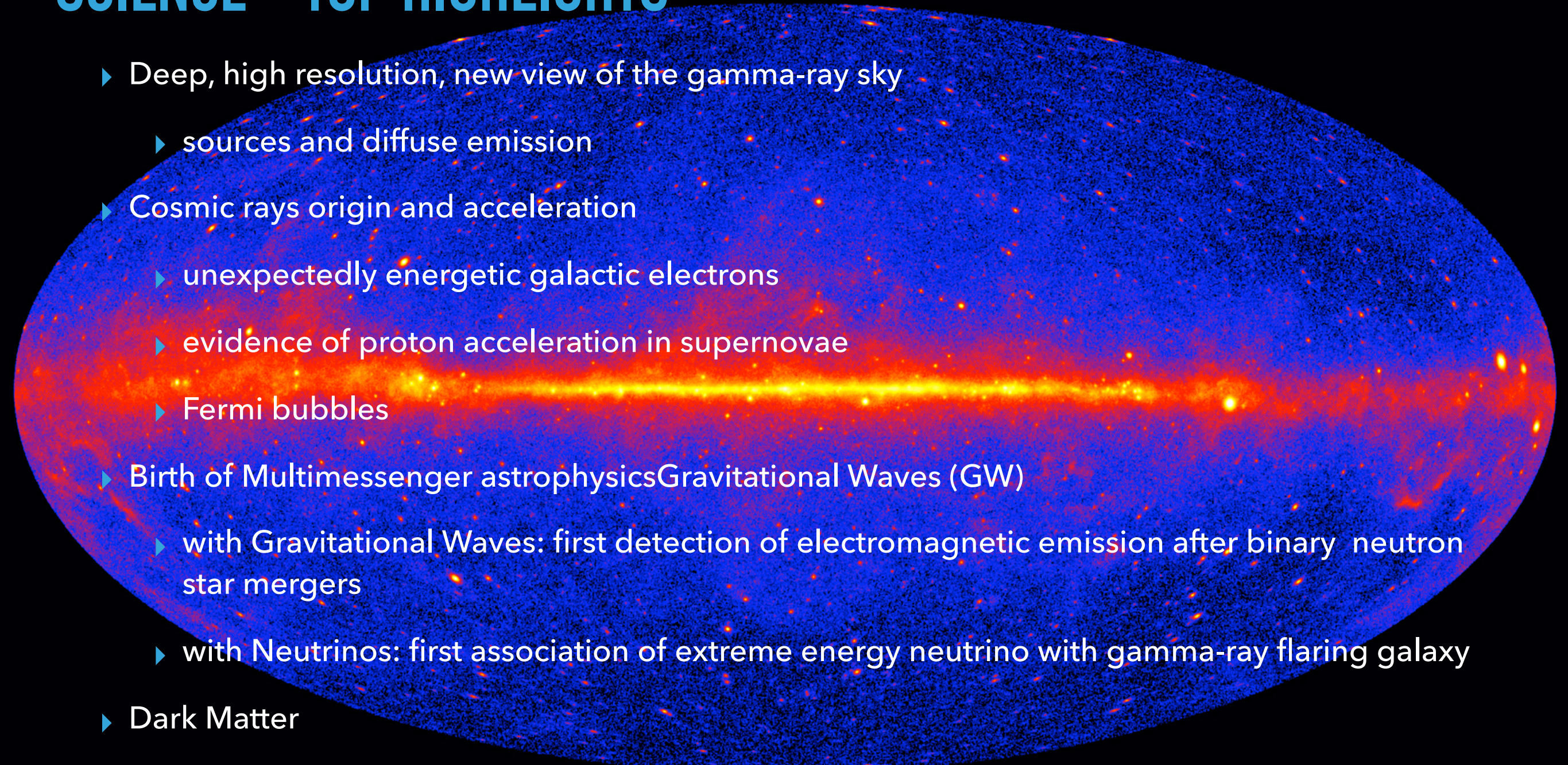
- ~400 collaboration papers
- ~2400 total nb of papers

Gamma Burst Monitor (GBM) - counters

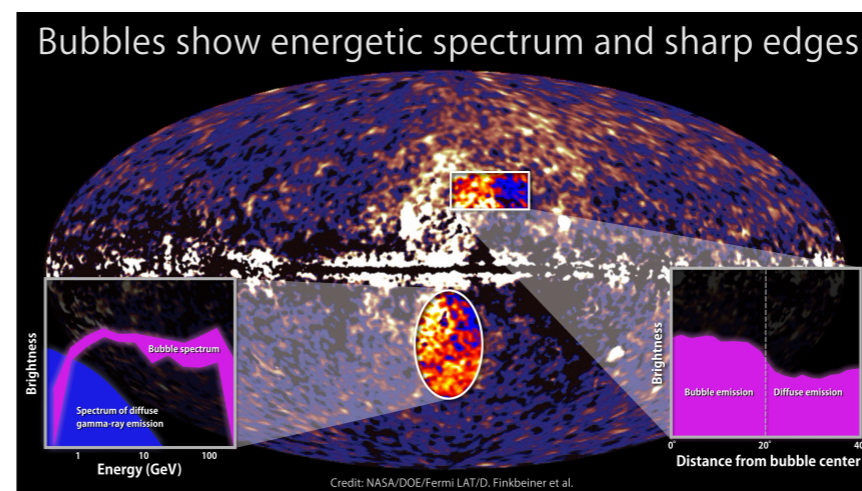
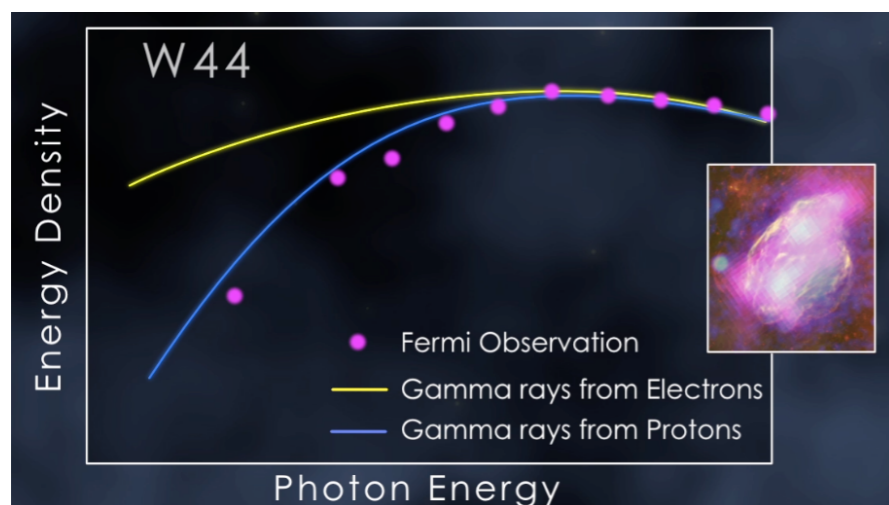
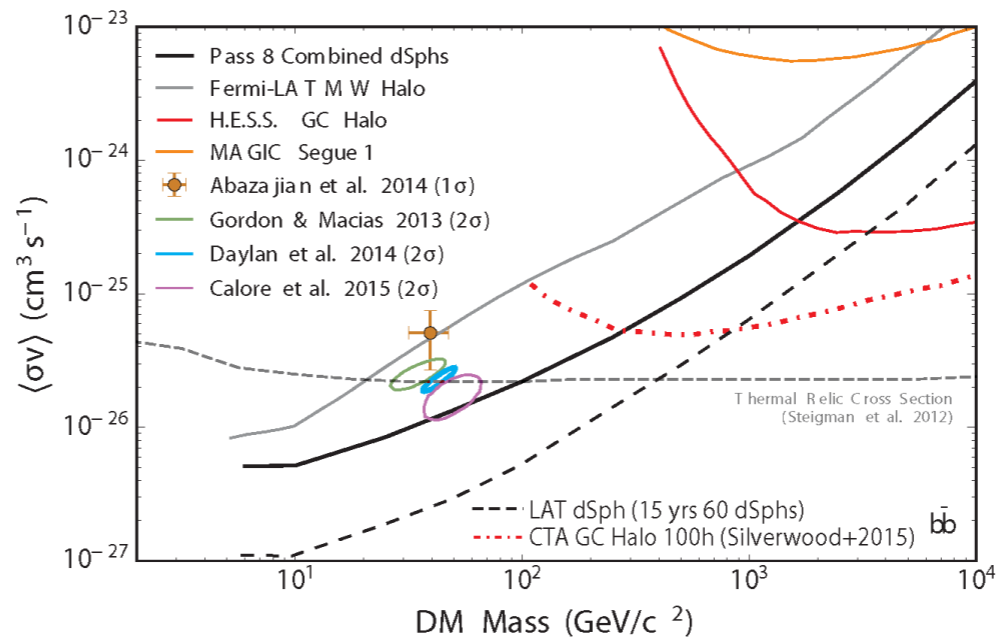
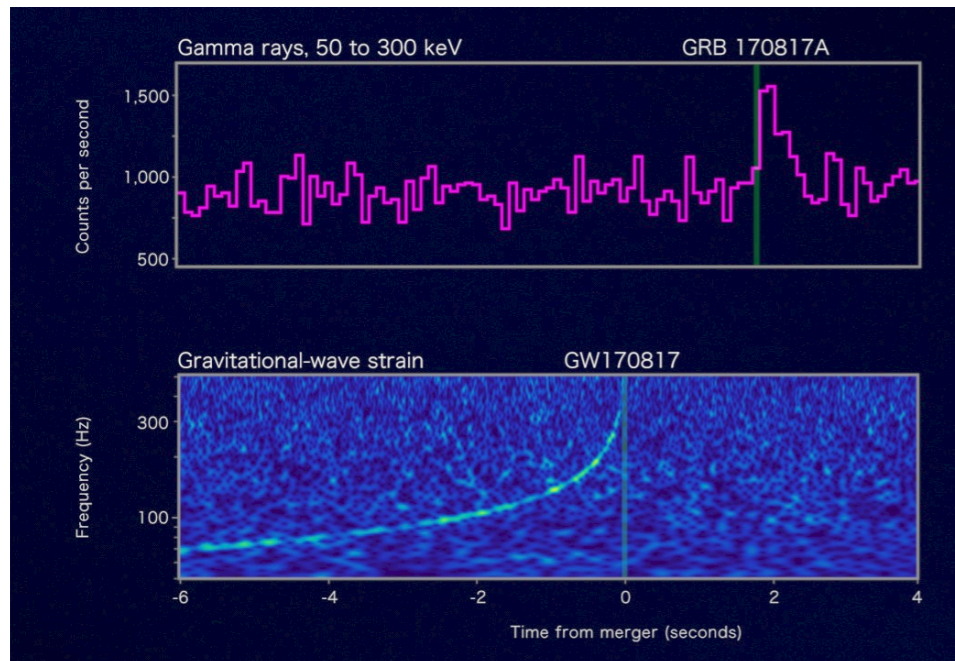
- 8 keV – 40 MeV

launch from KSC 11-6-2008

SCIENCE – TOP HIGHLIGHTS

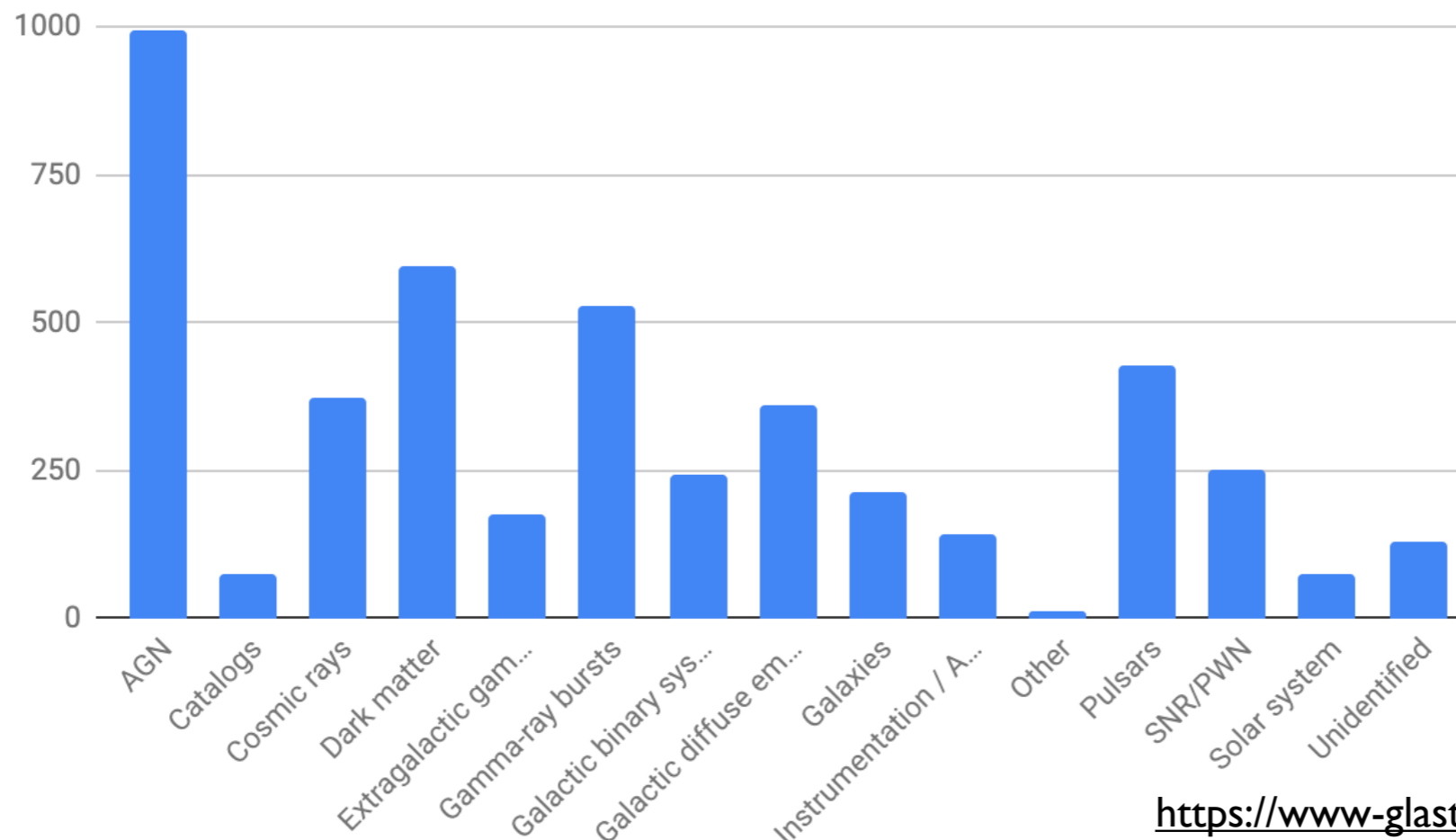
- 
- ▶ Deep, high resolution, new view of the gamma-ray sky
 - ▶ sources and diffuse emission
 - ▶ Cosmic rays origin and acceleration
 - ▶ unexpectedly energetic galactic electrons
 - ▶ evidence of proton acceleration in supernovae
 - ▶ Fermi bubbles
 - ▶ Birth of Multimessenger astrophysics
 - ▶ Gravitational Waves (GW)
 - ▶ with Gravitational Waves: first detection of electromagnetic emission after binary neutron star mergers
 - ▶ with Neutrinos: first association of extreme energy neutrino with gamma-ray flaring galaxy
 - ▶ Dark Matter
 - ▶ most stringent limits on generic particle candidate (WIMP)

SCIENCE HIGHLIGHTS GALLERY



Science highlights

- **Fermi data - 4photons/sec**
 - **3.13B public photons, 1.19B source-class photons**
- **Fermi products - 1 LAT paper/week**
 - **3438 papers, 129667 citations, 563 LAT papers**



Science Highlights - 8 years catalog

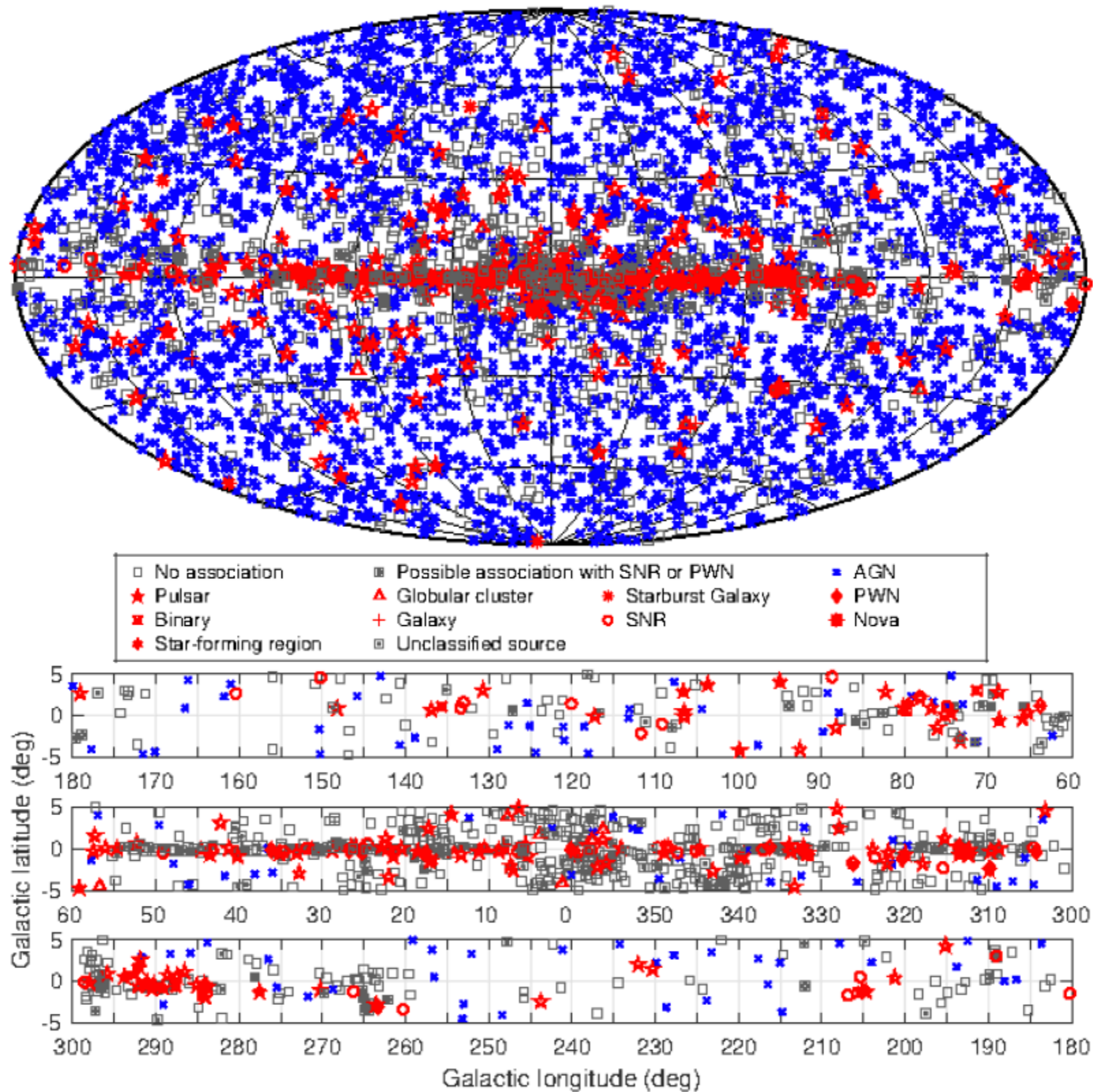


Figure 2. Full sky map (top) and blow-up of the Galactic plane split into three longitude bands (bottom) showing sources by source class (see Table 7, no distinction is made between associations and identifications). All AGN classes are plotted with the same blue symbol for simplicity. Other associations to a well-defined class are plotted in red. Unassociated sources and sources associated to counterparts of unknown nature are plotted in black.

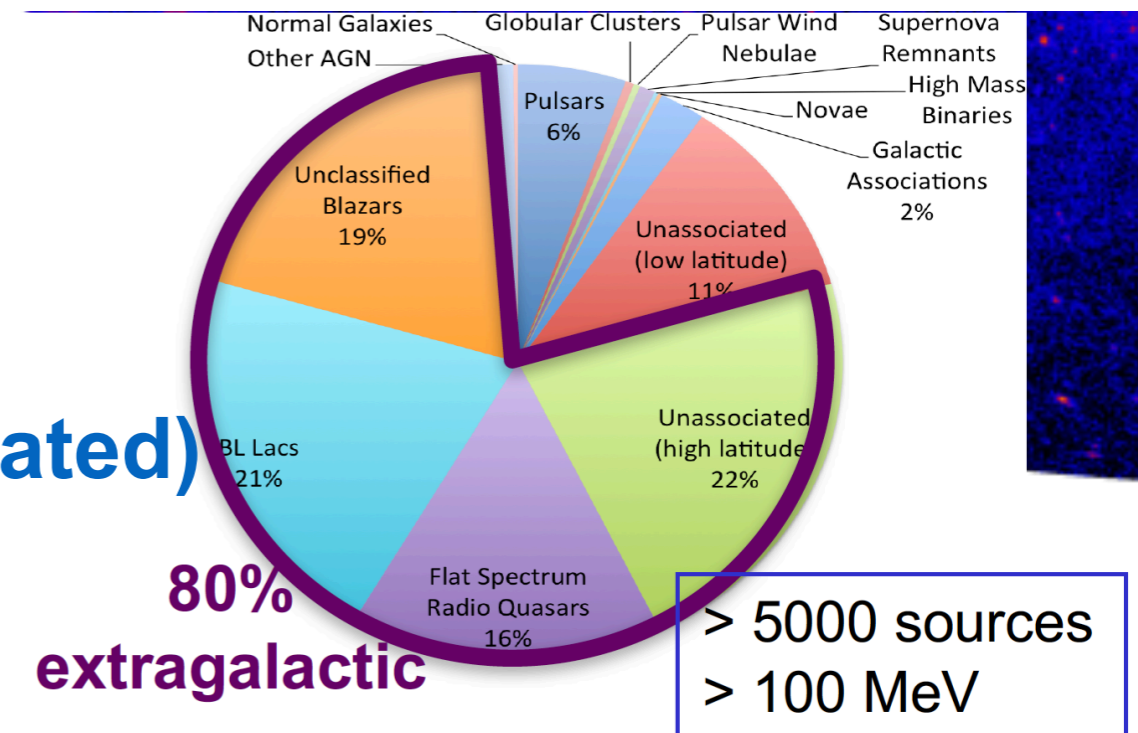
4th LAT Catalog

- **Ultimate effort - beyond 5 years catalog legal obligation**
 - 8 years of data, latest Pass8 IRFs, improved diffuse emission model, updated multi-wavelength catalogs, search for spatial extension, multiple spectral models

- **Known findings**

- ~30% unidentified sources
- >60% (of total) / >80% (of associated) are blazars

- multiple classes

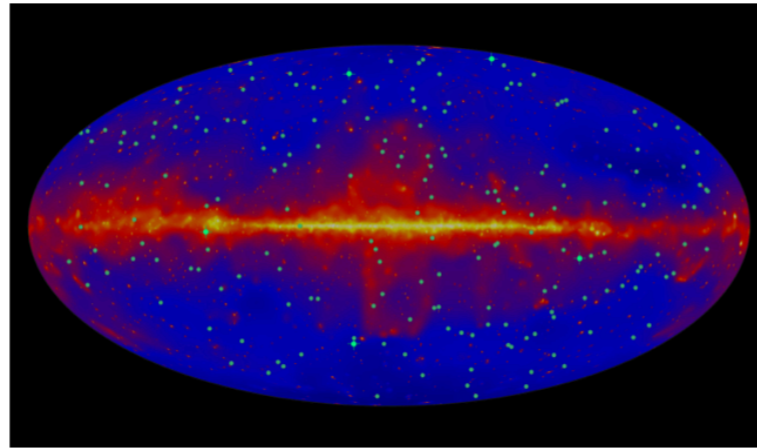


Comunicazione

- Ufficio Comunicazione ▾
- Comunicati stampa ▾
- News ▾
- Newsletter ▾
- Rassegna Stampa ▾
- Mostre e Installazioni ▾
- Eventi ▾
- Approfondimenti ▾
- Infografiche e poster ▾
- Brochure
- Editoria
- Asimmetrie
- Immagini
- Video
- Social network ▾

📅 13 GIUGNO 2019

LAMPI GAMMA: ECCO IL CATALOGO DEI PIÙ ENERGETICI OSSERVATI DA FERMI



186 lampi gamma di alta energia registrati dal telescopio satellitare per raggi gamma Fermi in dieci anni di osservazione: sono questi i protagonisti del [catalogo](#) pubblicato oggi, 13 giugno, su *The Astrophysical Journal*. Questi lampi gamma, anche noti come Gamma Ray Burst (GRB), sono stati rivelati dal Large Area Telescope (LAT), strumento di Fermi progettato e realizzato con un contributo decisivo dell'Italia, grazie all'Agenzia Spaziale Italiana ASI, all'Istituto Nazionale di Fisica Nucleare INFN e l'Istituto Nazionale di Astrofisica INAF.

“Ogni Gamma Ray Burst è in qualche modo unico. È solo quando siamo in grado di studiarne tanti, come abbiamo fatto in questo catalogo, che iniziamo a comprenderne le caratteristiche comuni,” racconta Elisabetta Bissaldi, ricercatrice dell'INFN e del Politecnico di Bari. “Il primo catalogo LAT, pubblicato nel 2013, comprendeva solo 35 GRB. Grazie a un netto miglioramento delle tecniche di analisi dati, abbiamo identificato un numero di GRB cinque volte maggiore in questo nuovo catalogo, imparando così a conoscere meglio i meccanismi fisici all'opera. Ad esempio, abbiamo confermato che l'emissione di raggi gamma ad alta energia dura più a lungo rispetto all'emissione a bassa energia e che la succede,” conclude Bissaldi.

Il catalogo che fornisce nuove indicazioni su origine ed evoluzione dei lampi gamma è il frutto del lavoro di 120 scienziate e scienziati della collaborazione Fermi coordinati da Bissaldi, da Magnus Axelsson dell'Università di Stoccolma e dagli italiani Nicola Omodei e Giacomo Vianello dell'Università di Stanford.

La maggior parte dei lampi gamma nasce quando alcuni tipi di stelle massive esauriscono il proprio combustibile e collassano generando buchi neri. Altri invece hanno origine dalla collisione di due stelle di neutroni, oggetti densissimi residuo di esplosioni stellari. Sia il collasso di una stella sia la collisione di due stelle di neutroni danno, infatti, origine a jet relativistici di particelle che si muovono a una velocità prossima a quella della luce. Quando le particelle all'interno dei jet si scontrano tra di loro o interagiscono con l'ambiente intorno alle stelle, nascono i raggi gamma che sono poi rivelati da Fermi grazie ai suoi strumenti principali: il LAT e il GBM.

A Decade of Gamma-Ray Bursts Observed by Fermi-LAT: The Second GRB Catalog

Ajello, M. (Fermi LAT Collaboration) et al. 2019, *ApJ*, 878, 52 [Hide links](#)

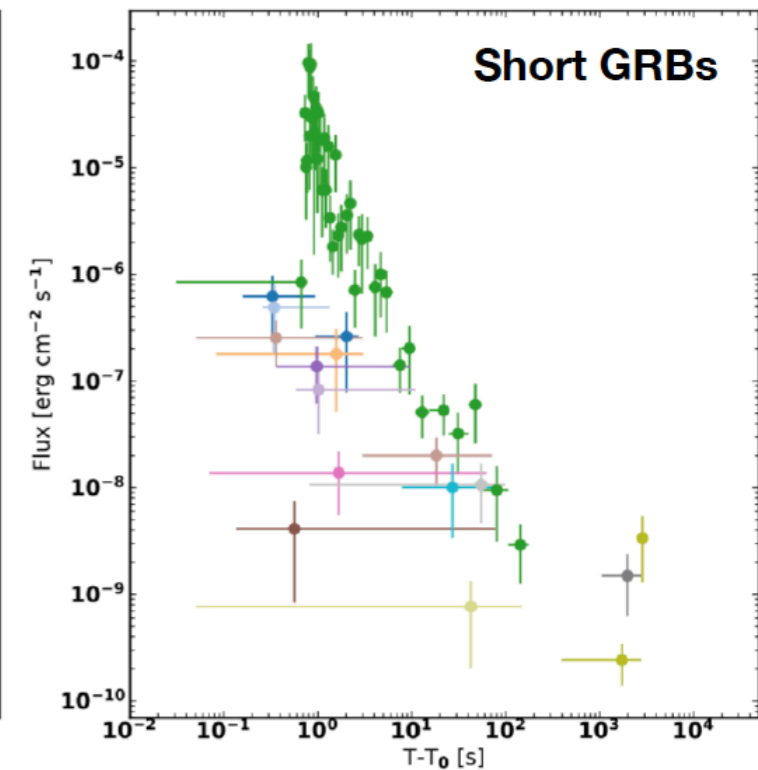
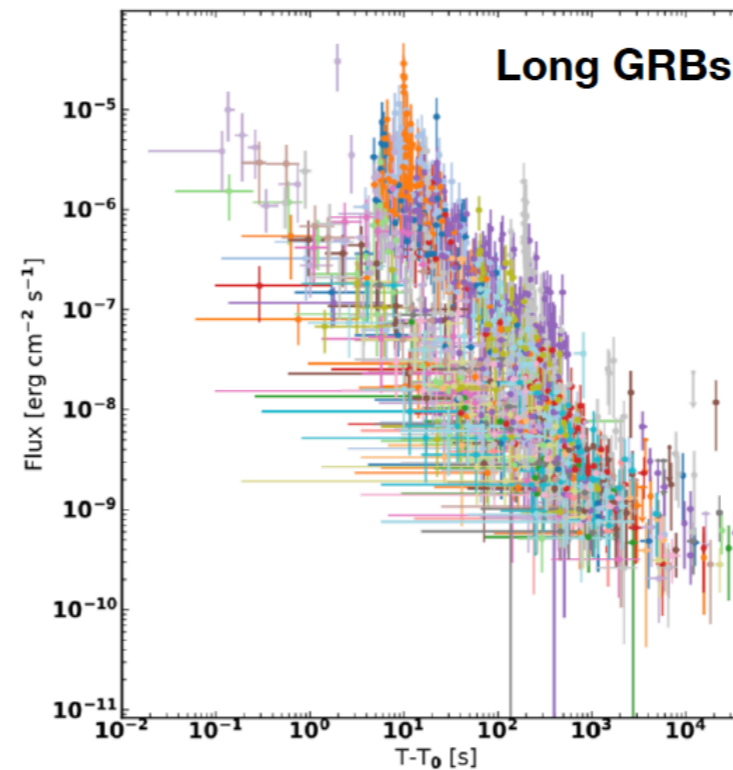
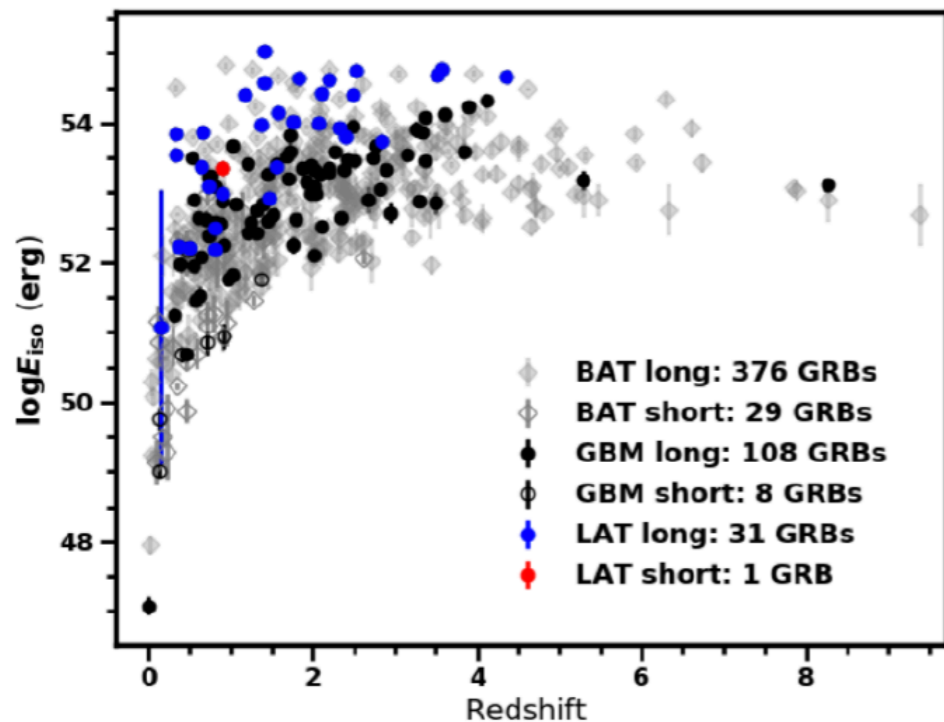
doi: [10.3847/1538-4357/ab1d4e](https://doi.org/10.3847/1538-4357/ab1d4e)

arXiv: [1906.11403](https://arxiv.org/abs/1906.11403) [INSPIRE](#)

Fermi GRBs



- GBM has detected ~2650 GRBs in 11 years
 - ~240 GRBs/yr (40 short/yr)
- LAT has detected ~200 GRBs in 11 years
 - ~18 GRBs/yr (1.7 short/yr)
 - closest and/or most energetic events
- Common features in High Energy GRB emission
 - delayed onset (GeV relative to keV)
 - extra components (thermal, extra power laws)
 - extended emission - GeV afterglows



Ajello et al. 2019, ApJ, 878, 52

Operational Context

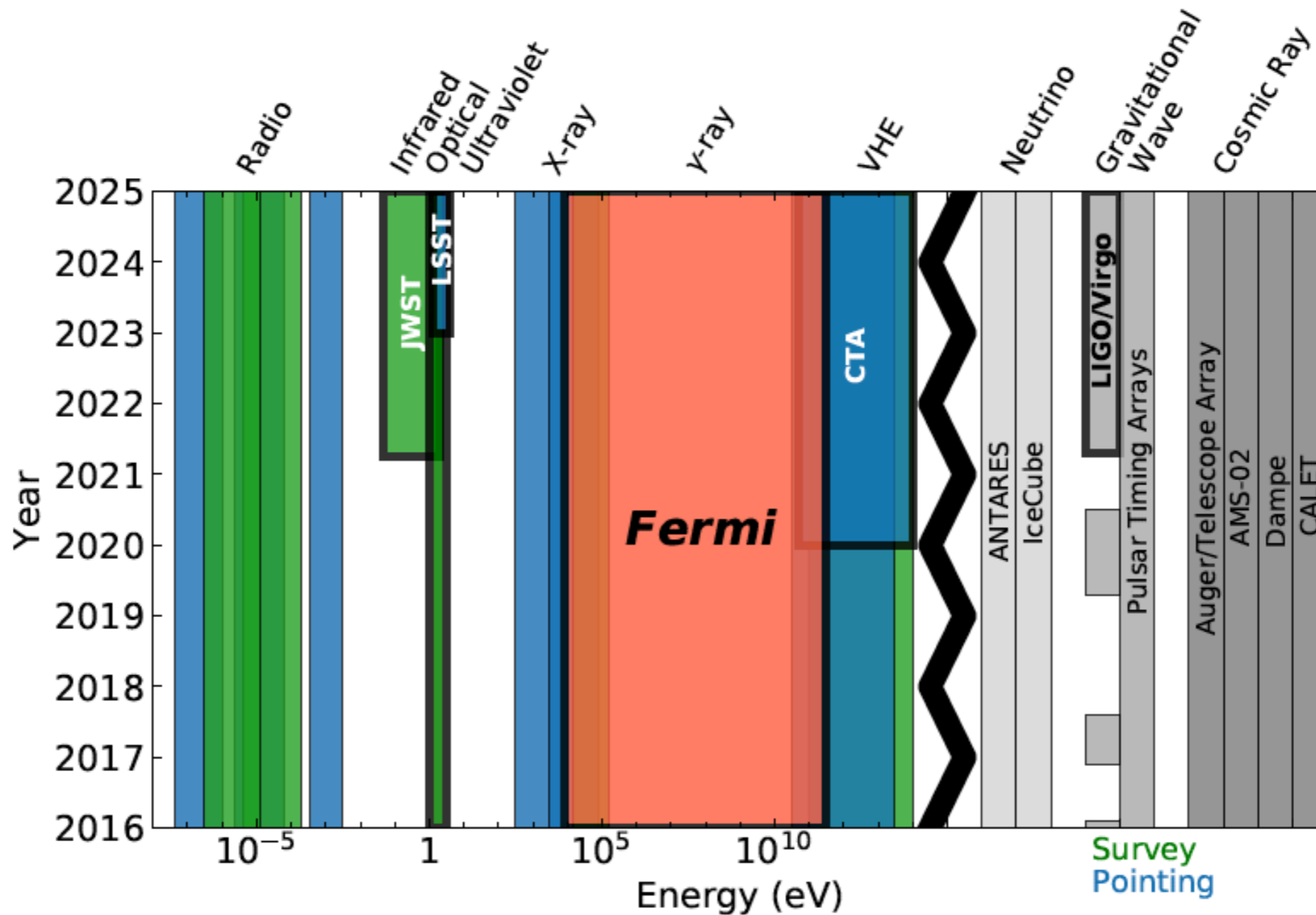


Figure 7: *Fermi* fills a unique part of the electromagnetic spectrum, especially vital in the coming years as new or enhanced facilities (bold outlined) come online presenting new opportunities in cooperation with *Fermi*.

Fermi renewed status of CERN recognized experiment through 2022

NASA Senior Review

- 2019 Proposal well received
- SR panel and NASA HQ confirmed funds through 2022 and invite Fermi at next 2022 SR



The screenshot shows the NASA Science website. The top navigation bar includes 'Science Topics', 'Science News', 'For Researchers', and 'Learn More'. The main header is 'Universe' with a sub-navigation bar for 'Overview', 'What We Study', 'Programs', 'Missions', 'Science Questions', 'Astro Data', and 'Documents'. The main content area features the article '2019 Senior Review of Operating Missions'. The article text states: 'NASA's Science Mission Directorate (SMD) periodically conducts independent, comparative reviews of its operating missions. NASA uses the findings from these reviews to define an implementation strategy and give programmatic direction to the missions and projects concerned for the next five fiscal years. These reviews of operating missions are NASA's highest form of peer review, as the subject is not a single science investigation, or even a single space mission, but rather a portfolio of operating missions. The reviews of operating missions are referred to as Senior Reviews, in recognition of the high level of the peer review. The next Senior Review will be conducted in 2019. The following missions will be included: Chandra X-ray Observatory, Fermi Gamma-ray Space Telescope, Hubble Space Telescope, Neutron Star Interior Composition Explorer, Nuclear Spectroscopic Telescope Array, Neil Gehrels Swift Observatory, Transiting Exoplanet Survey Satellite, and X-ray Multi-Mirror Mission-Newton'. On the right side, there are two sections: 'What We Study' with links to 'Dark Energy', 'Black Holes', 'The Big Bang', 'Galaxies', 'Stars', and 'Exoplanets'; and 'Helpful Information' with links to 'Organizations', '2020 Distinguished Lecturers', '2019 Senior Review of Operating Missions', 'Astrophysics Chart', 'Astronomy', 'Night Sky', and 'Spacecraft'.

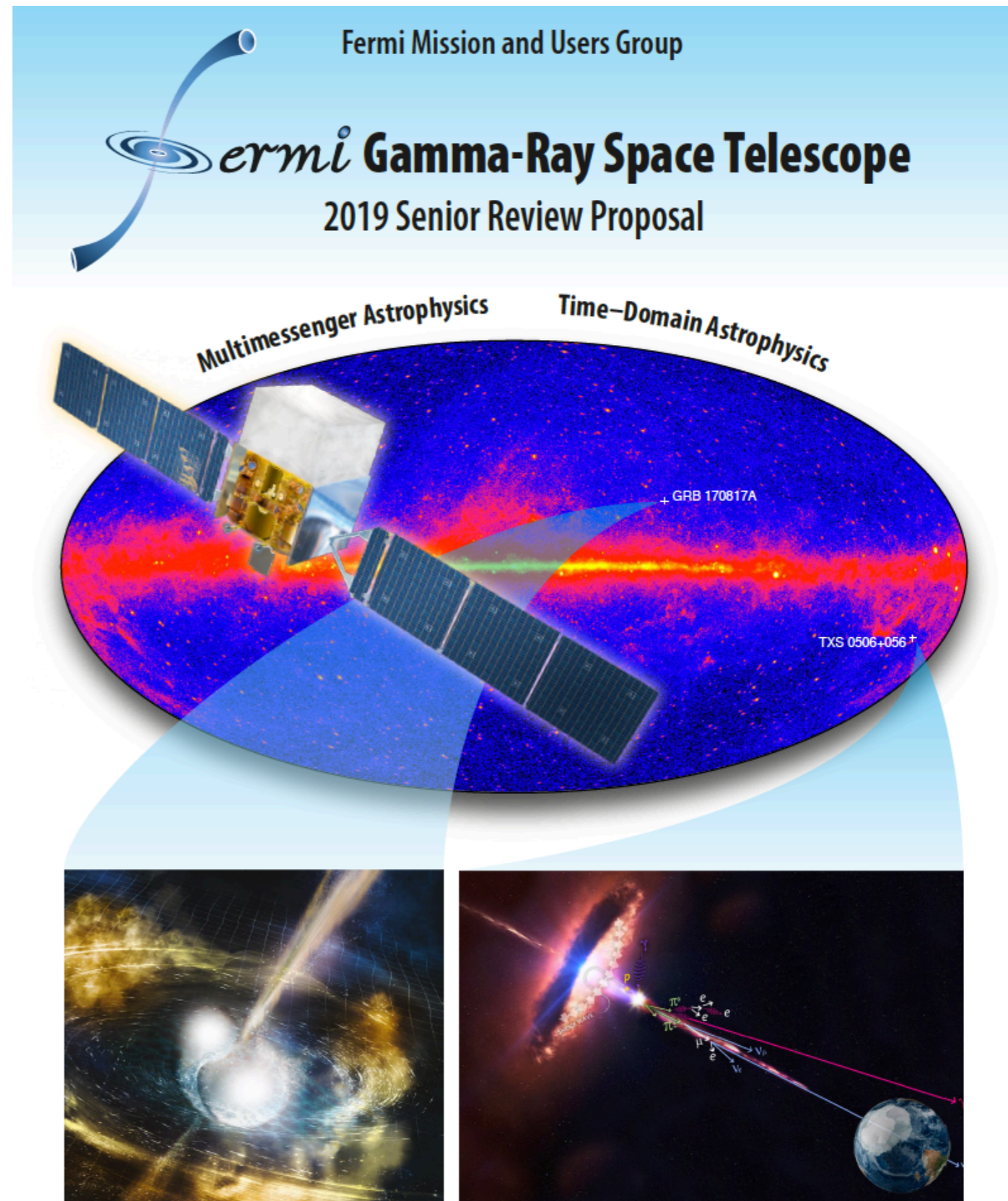
<https://science.nasa.gov/astrophysics/2019-senior-review-operating-missions>

Fermi 2019 SR Proposal

**Focus on multi-messenger
and time-domain**

astrophysics after

- **first observation of BNS
gravitational merge in
gamma-rays**
- **first association of flaring
AGN with high energy
neutrino**

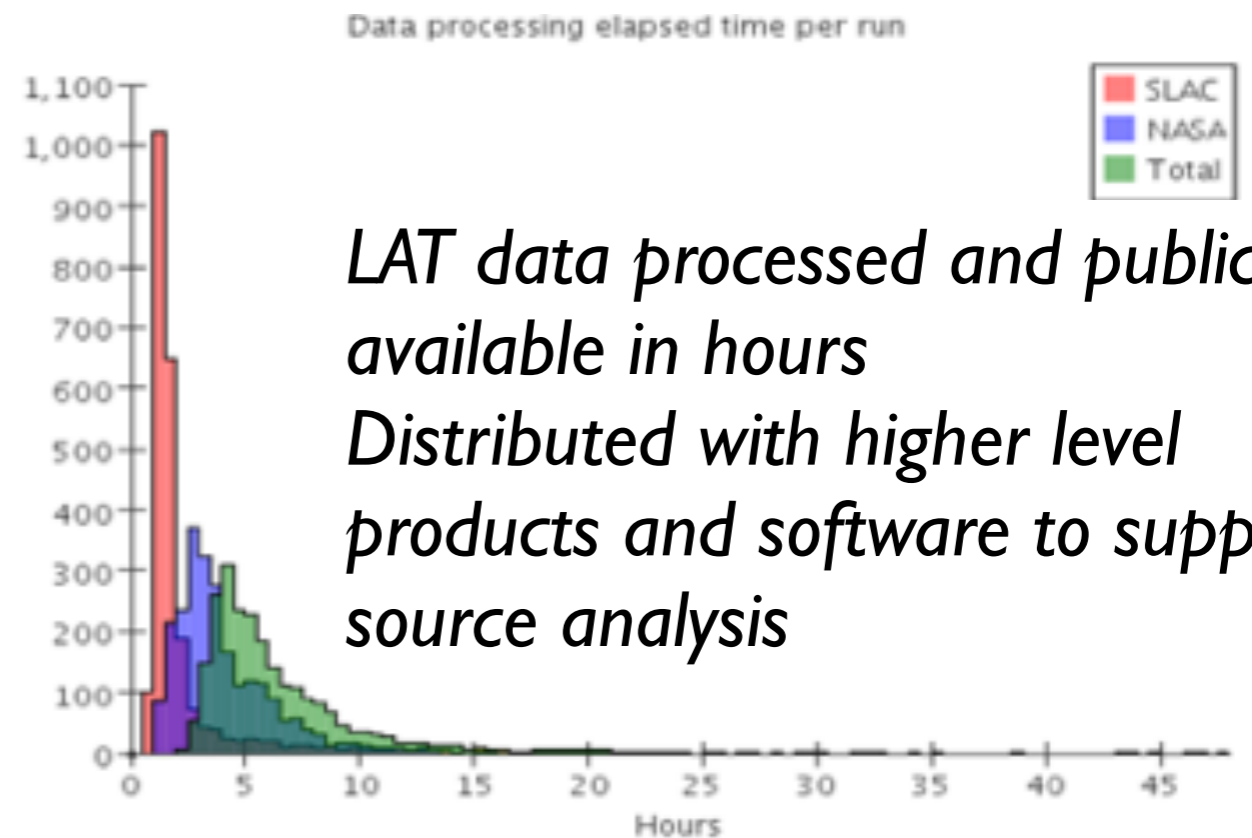


Mission status

- **Observatory running smoothly**
 - **one solar array drive damaged in March 2018, since then oriented at fixed position**
 - **modified rocking profile to recover exposure uniformity**
- **All LAT subsystems working with no degradation**
 - **CAL light output reduced by ~6% since launch, as expected from irradiation**
 - **TKR has only 0.07% strips masked**

Upcoming challenges

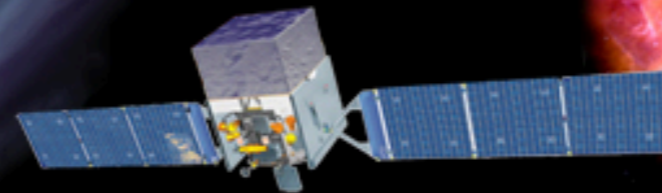
- **Open data and alert systems**
 - **LAT collaboration prepared pipelines and staffed shifts to cover multi-messenger opportunities**
 - **LAT committed to continue supporting smooth telescope operations and serve high quality gamma-ray data products to the community**



*LAT data processed and publicly available in hours
Distributed with higher level products and software to support source analysis*

Fermi

Gamma-ray Space Telescope



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+ [GBM Data](#)

▶ [Data Analysis](#)

▶ [Caveats](#)

▶ [Newsletters](#)

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Currently Available Data Products

The Fermi data released to the scientific community is governed by the [data policy](#). The released instrument data for the GBM, along with LAT source lists, can be accessed through the [Browse interface specific to Fermi](#). LAT photon data can be accessed through the [LAT data server](#).

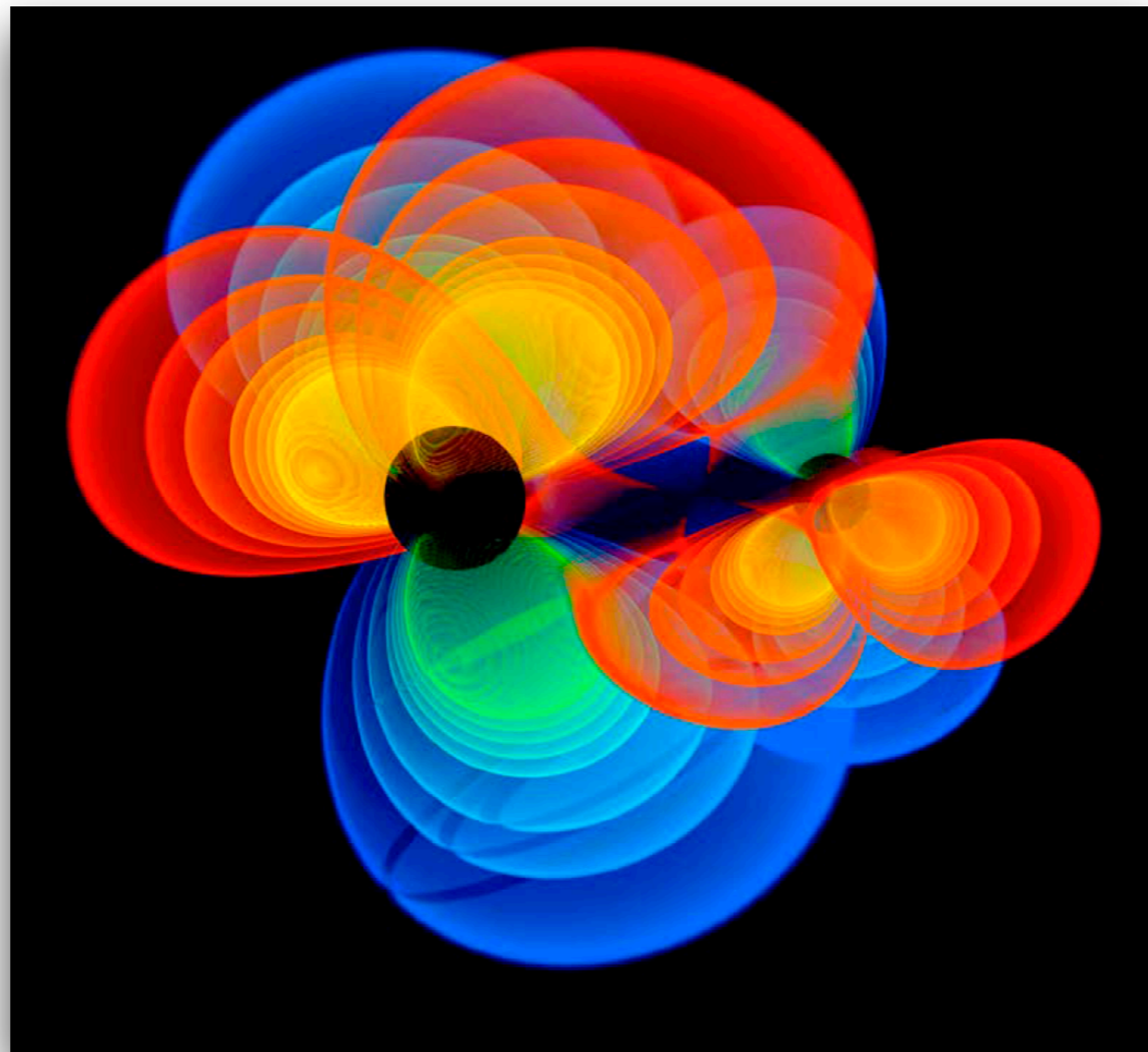
The FITS files can also be downloaded from the Fermi [FTP site](#). The file version number is the 'xx' in the characters before the extension in each filename; you should keep track of the version numbers of files you analyze since the instrument teams may update them.

Note that the LAT and GBM data are accompanied by [caveats](#) about their use.

- LAT Photon and Extended Data
 - [LAT Data Server](#) (updated with P8R3 data 26-Nov-2018)
 - [LAT Low-Energy \(LLE\) Data](#) (Browse table)
 - Products available on the [FTP Site](#) (current processing version of the data).
 - [Weekly Photon Files](#)
 - [Weekly Spacecraft Files](#)
 - [Mission Long Spacecraft File](#)
 - [Weekly 1-second Spacecraft Files](#)
 - [Filtered Weekly Photon Files with Diffuse Response Columns](#)
 - Previous processing versions available on the FTP site
 - [Pass 8 \(P8R2\) Weekly Files](#)
 - [Pass 7 \(V6d\) Weekly files](#)
 - [Pass 7 \(V6\) Weekly files](#)
 - [Pass 6 \(V11\) Weekly files](#)
 - [Pass 6 \(V3\) Weekly files](#)
 - [ASDC data server](#) (external)
- LAT catalogs and associated products (high-level products only)
 - LAT Point Source Catalog
 - [LAT 8-year Point Source Catalog \(4FGL\)](#)
 - [Preliminary LAT 8-year Source List \(FL8Y\)](#)
 - [LAT 4-year Point Source Catalog \(3FGL\)](#)
 - [LAT 2-year Point Source Catalog \(2FGL\)](#)
 - [LAT 1-year Point Source Catalog \(1FGL\)](#)
 - [LAT 3-month Bright Source List \(0FGL\)](#)

FERMI AND MULTIMESSENGER OBSERVATIONS

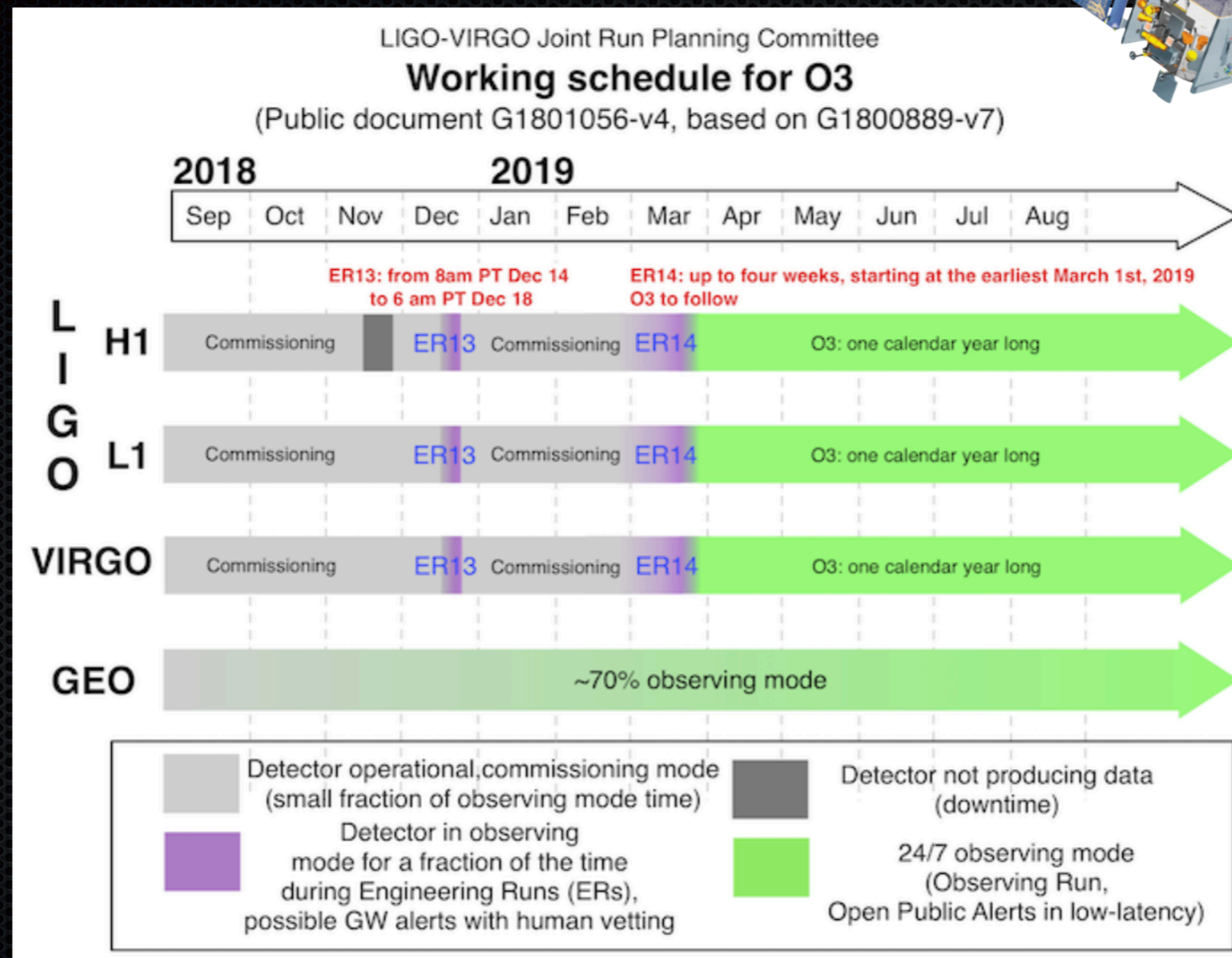
3 – MULTIMESSENGER ASTRONOMY – STATUS AND PROSPECTS



FERMI AND MULTIMESSENGER OBSERVATIONS

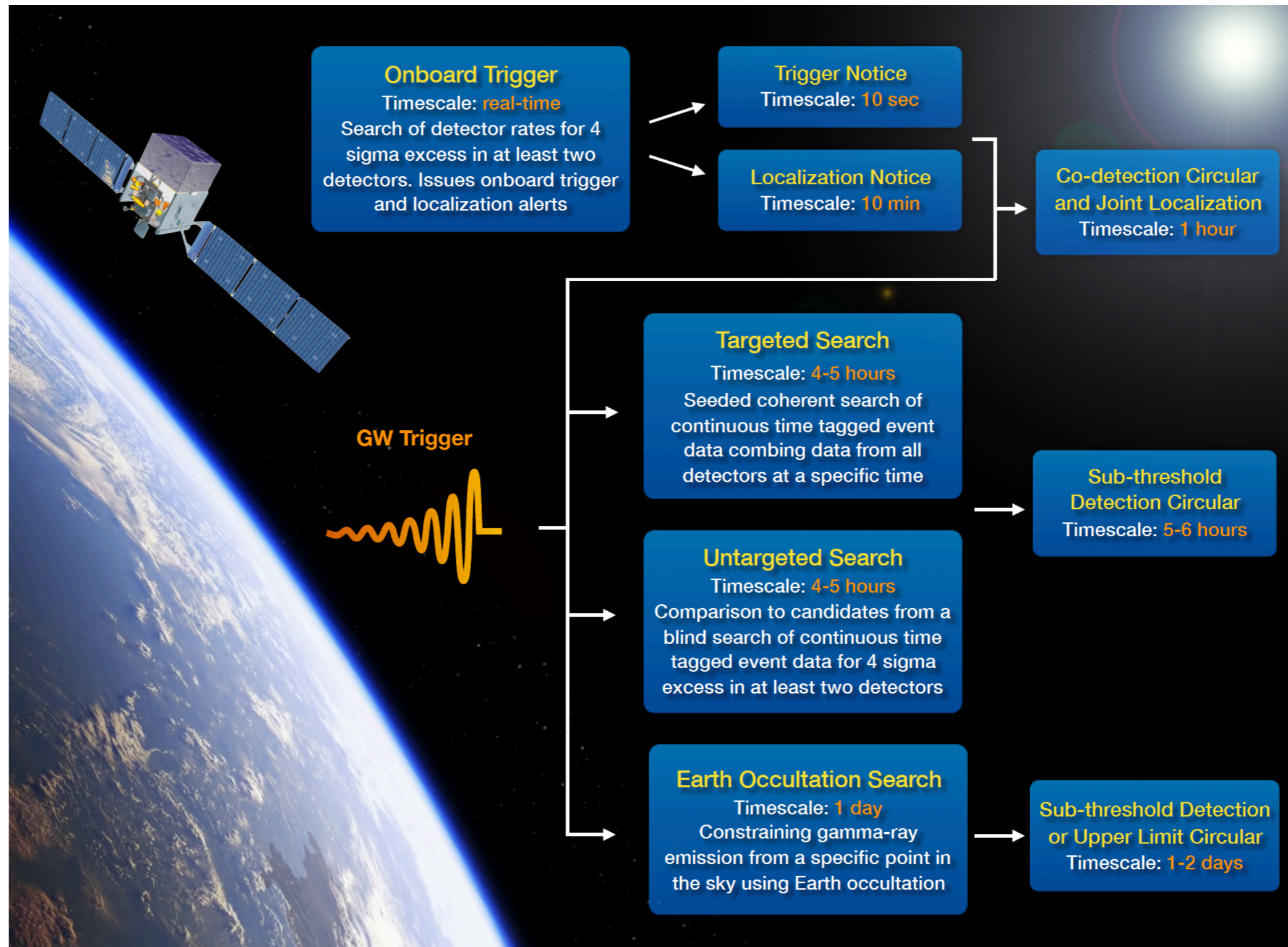
LVC OPERATIONS

LIGO & VIRGO O3



FERMI AND MULTIMESSENGER OBSERVATIONS

FERMI GBM OPERATIONS



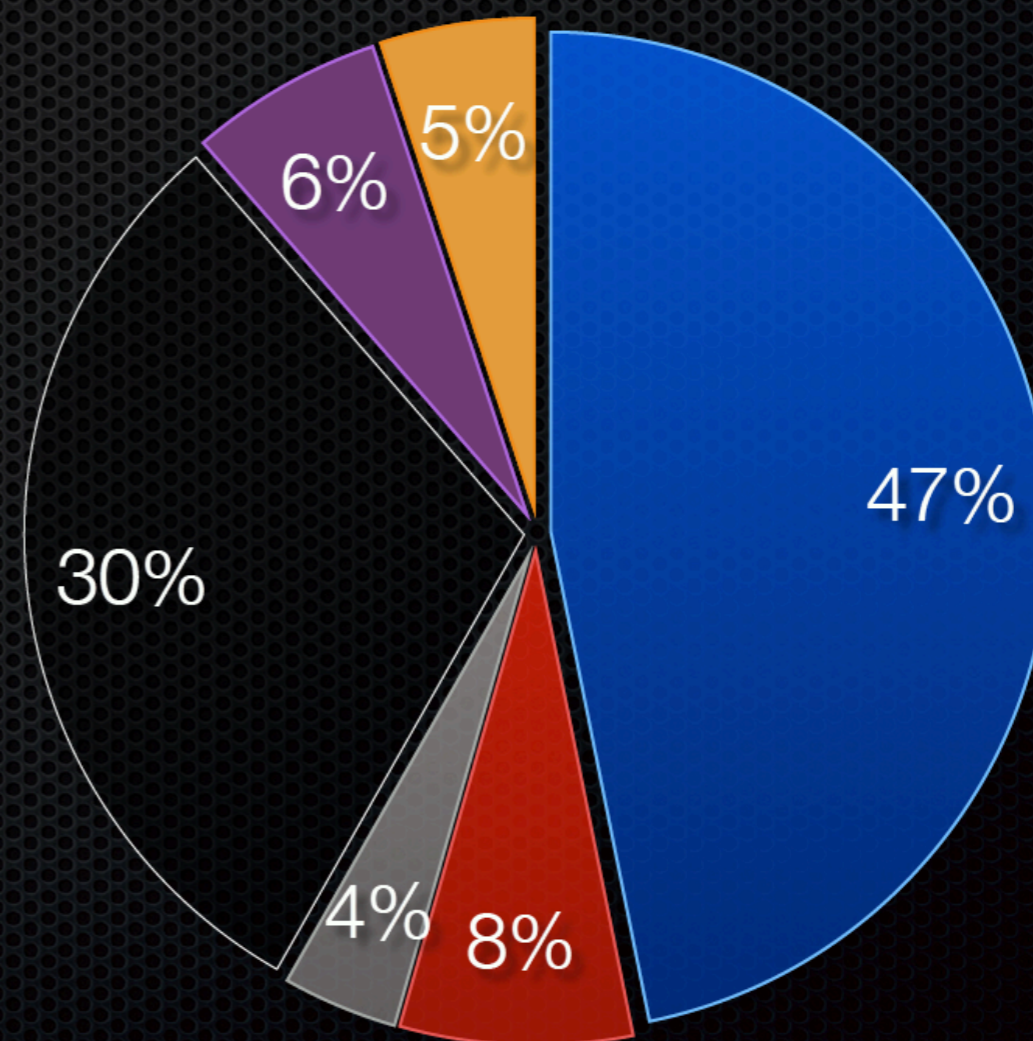
FERMI AND MULTIMESSENGER OBSERVATIONS

GW OBSERVATIONS

- LVC issued 80 public detection alerts via GCN since April 5th, 2018

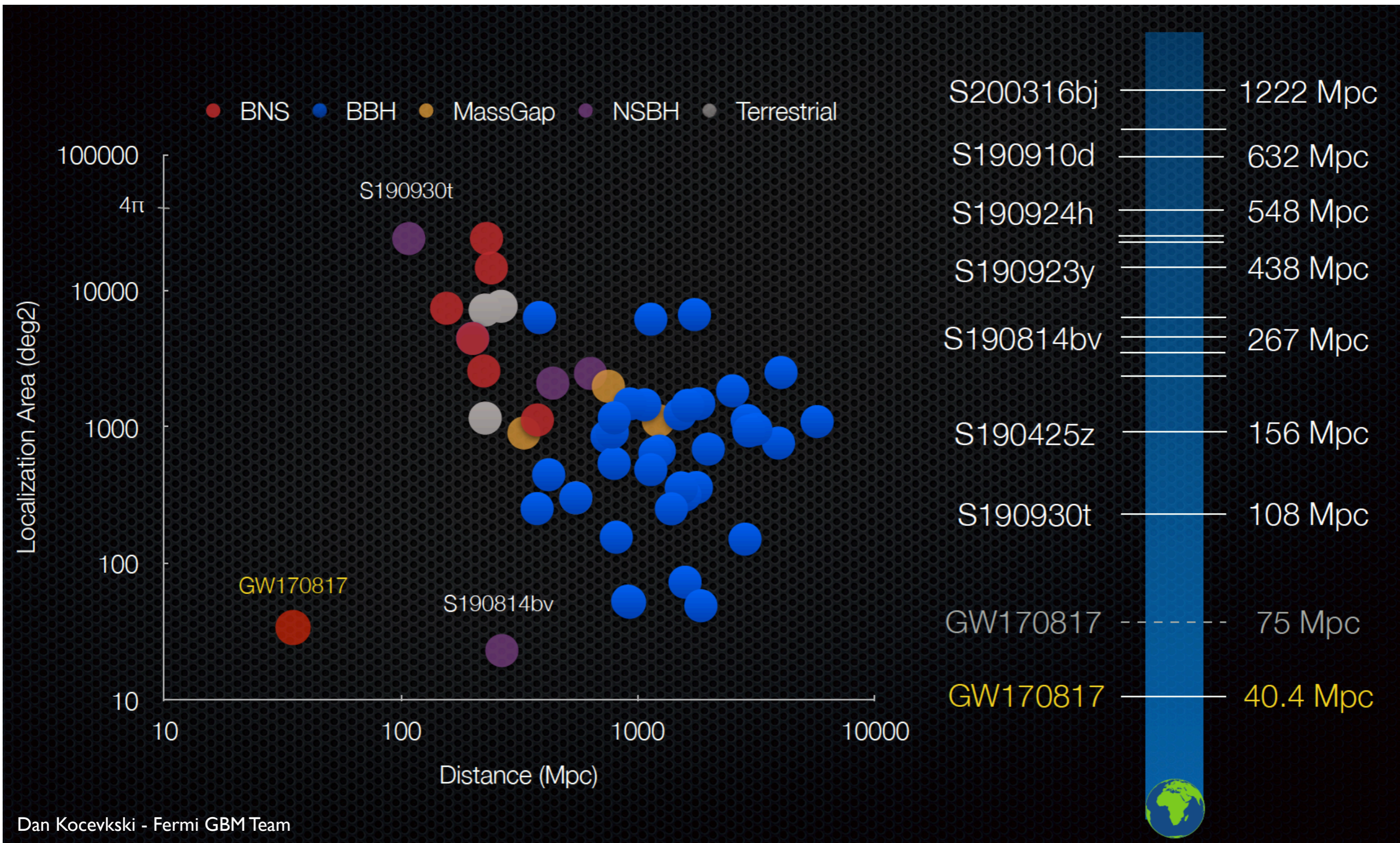
- 37 BBH
- 6 BNS
- 5 NSBH
- 4 MassGap
- 3 Terrestrial?
- 24 Retractions

- Fermi was in SAA for 9 of 55 un-retracted triggers, or roughly 16% of the time



FERMI AND MULTIMESSENGER OBSERVATIONS

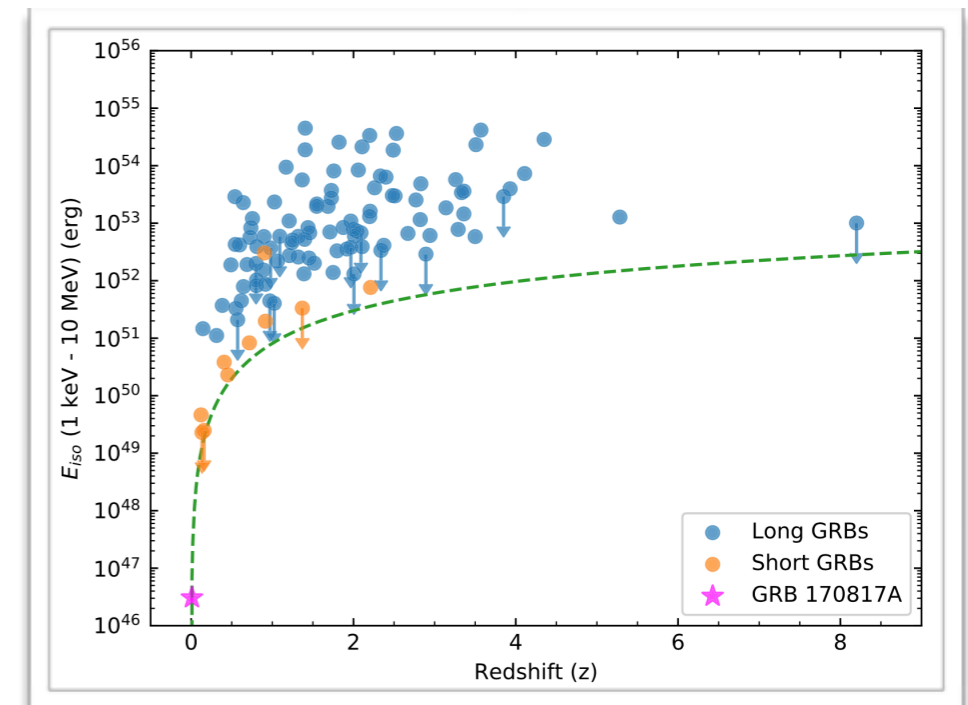
GW OBSERVATIONS



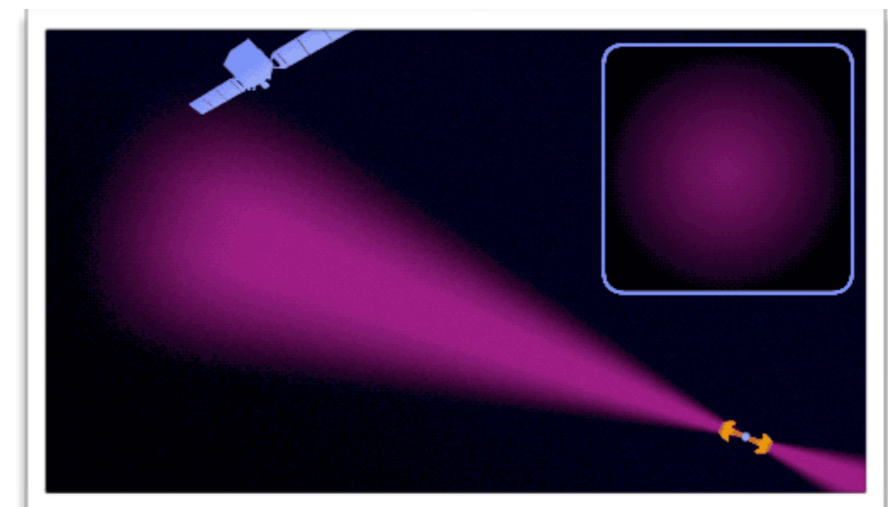
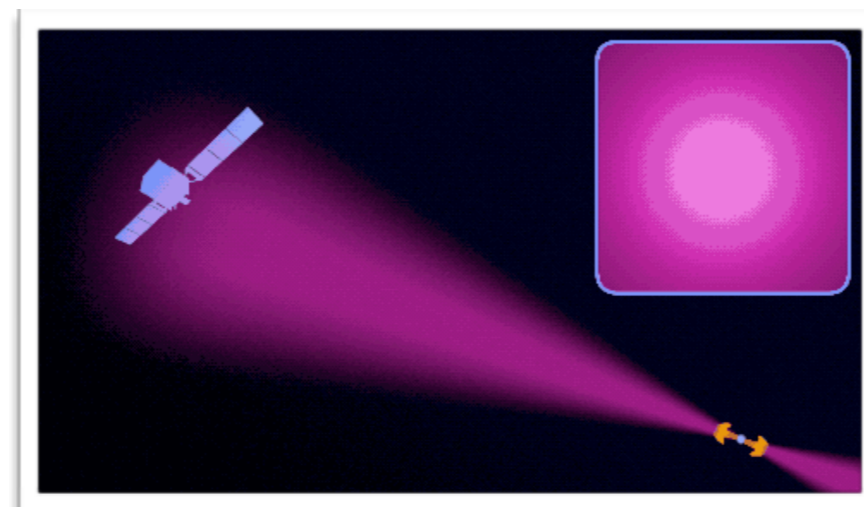
FERMI AND MULTIMESSENGER OBSERVATIONS

GRB170817 – A SPECIAL EVENT?

- ▶ $\sim 10^4$ dimmer than ordinary GRBs
- ▶ Circumstantial evidence of a jet seen off axis
- ▶ long term multi-wavelength observations and jet modeling to assess this picture



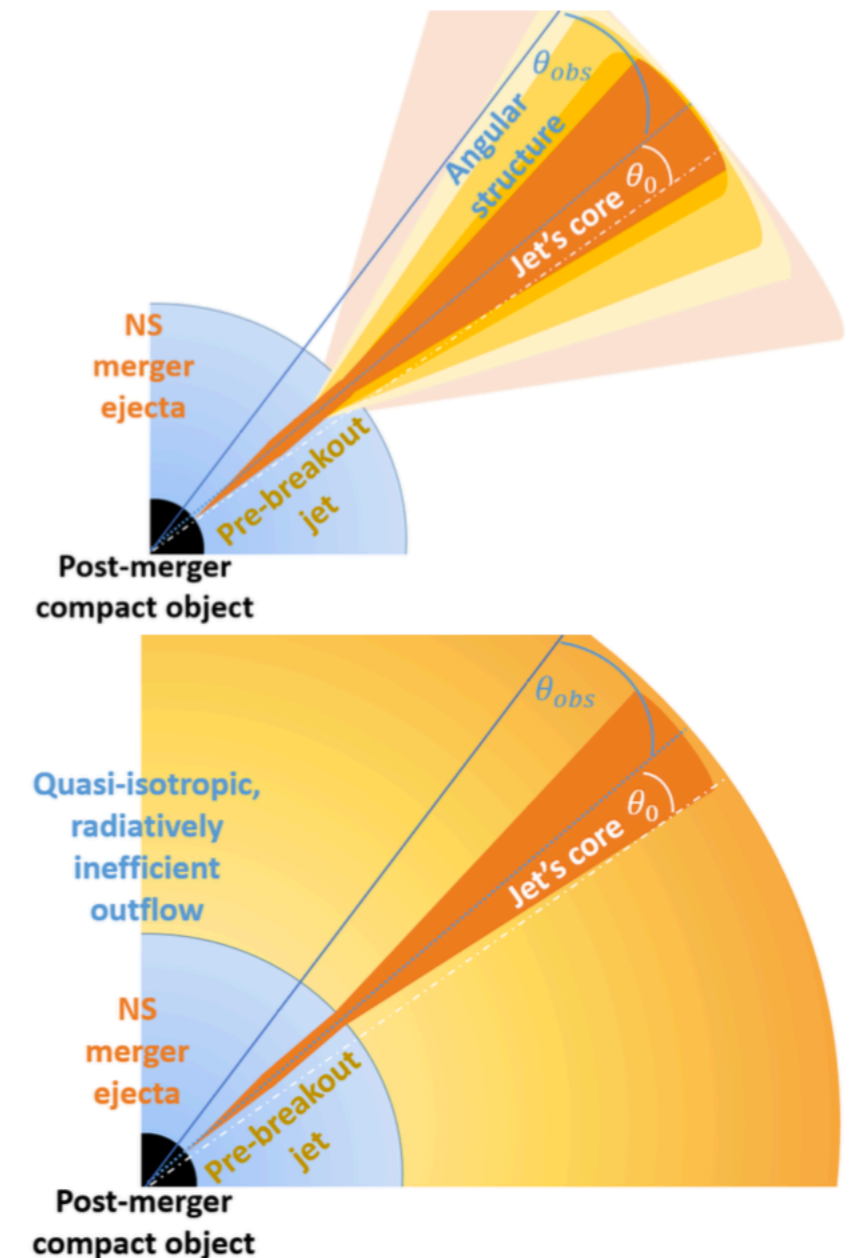
Abbott et al, 2017, arXiv:1710.05834



FERMI AND MULTIMESSENGER OBSERVATIONS

GRB170817 EMISSION MODELS

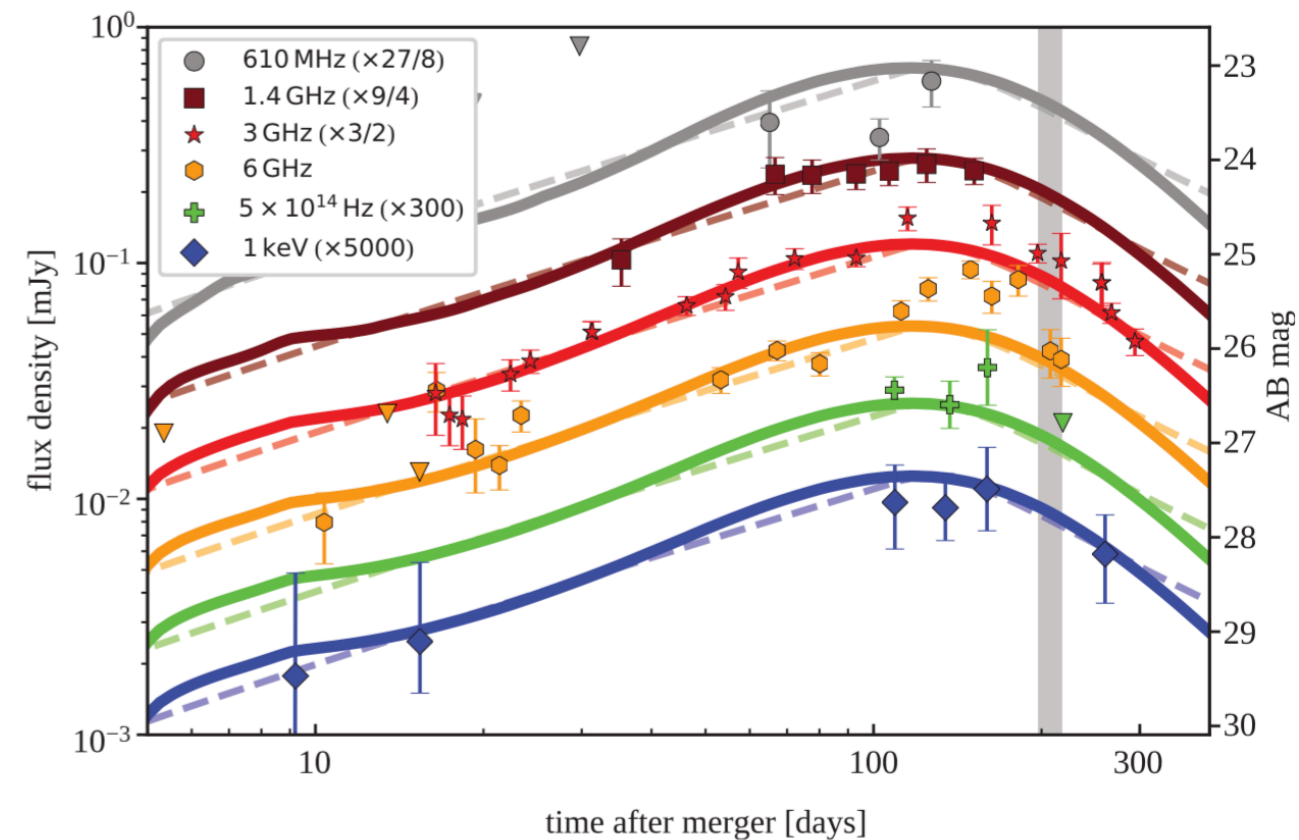
- ▶ Top: **structured jet**
 - ▶ jet with core luminosity (power-law / gaussian) breaks the ejecta
 - ▶ observer sees dimmer / less energetic emission
- ▶ Bottom: **choked cocoon**
 - ▶ jet with radiatively inefficient outflow



FERMI AND MULTIMESSENGER OBSERVATIONS

GRB170817 – MULTIFREQUENCY OBSERVATIONS

- ▶ Long-term observations of the multifrequency emission confirm non thermal emission
- ▶ Flux information compatible with a structured jet (solid lines) as well as with a choked cocoon (dashed)

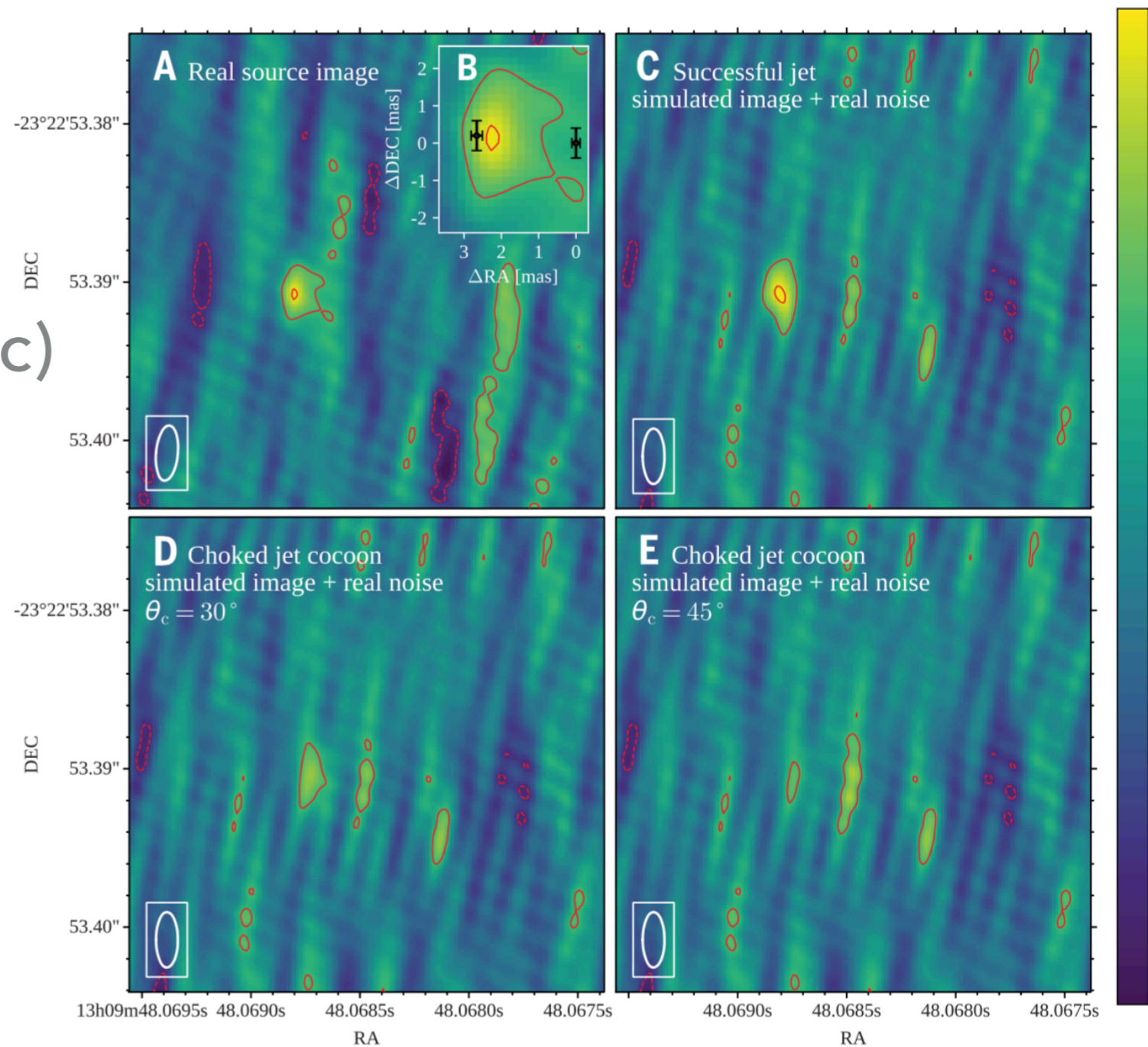


Ghirlanda et al, 2019, Science 363, 968-971

FERMI AND MULTIMESSENGER OBSERVATIONS

GRB170817 – RADIO VLBI OBSERVATIONS – IMAGING

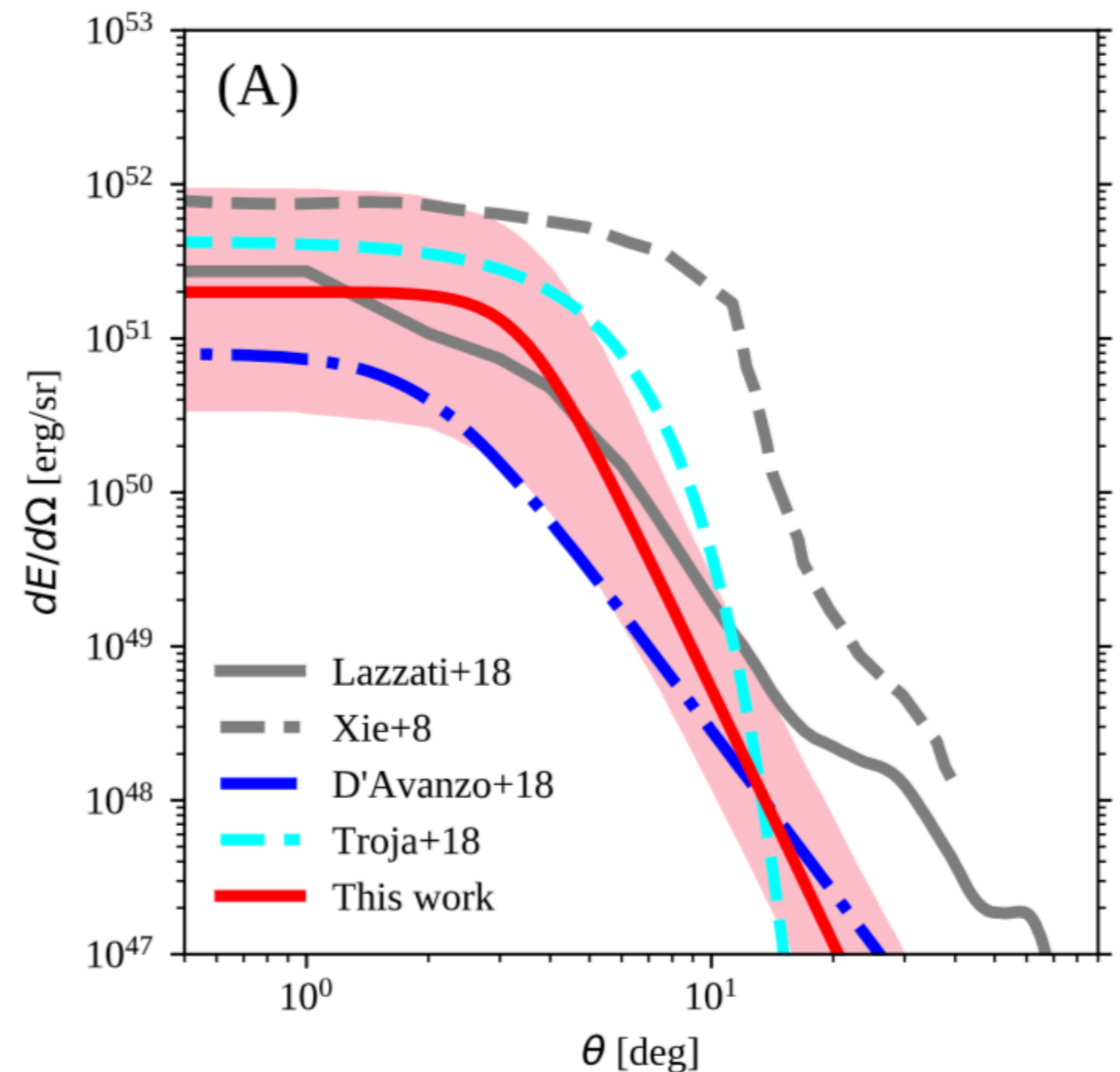
- ▶ Structured jet preferred to choked cocoon
- ▶ **compact** ($< \sim 2$ marcsec)
- ▶ peaked brightness



FERMI AND MULTIMESSENGER OBSERVATIONS

GRB170817 – RADIO VLBI OBSERVATIONS – JET MODEL

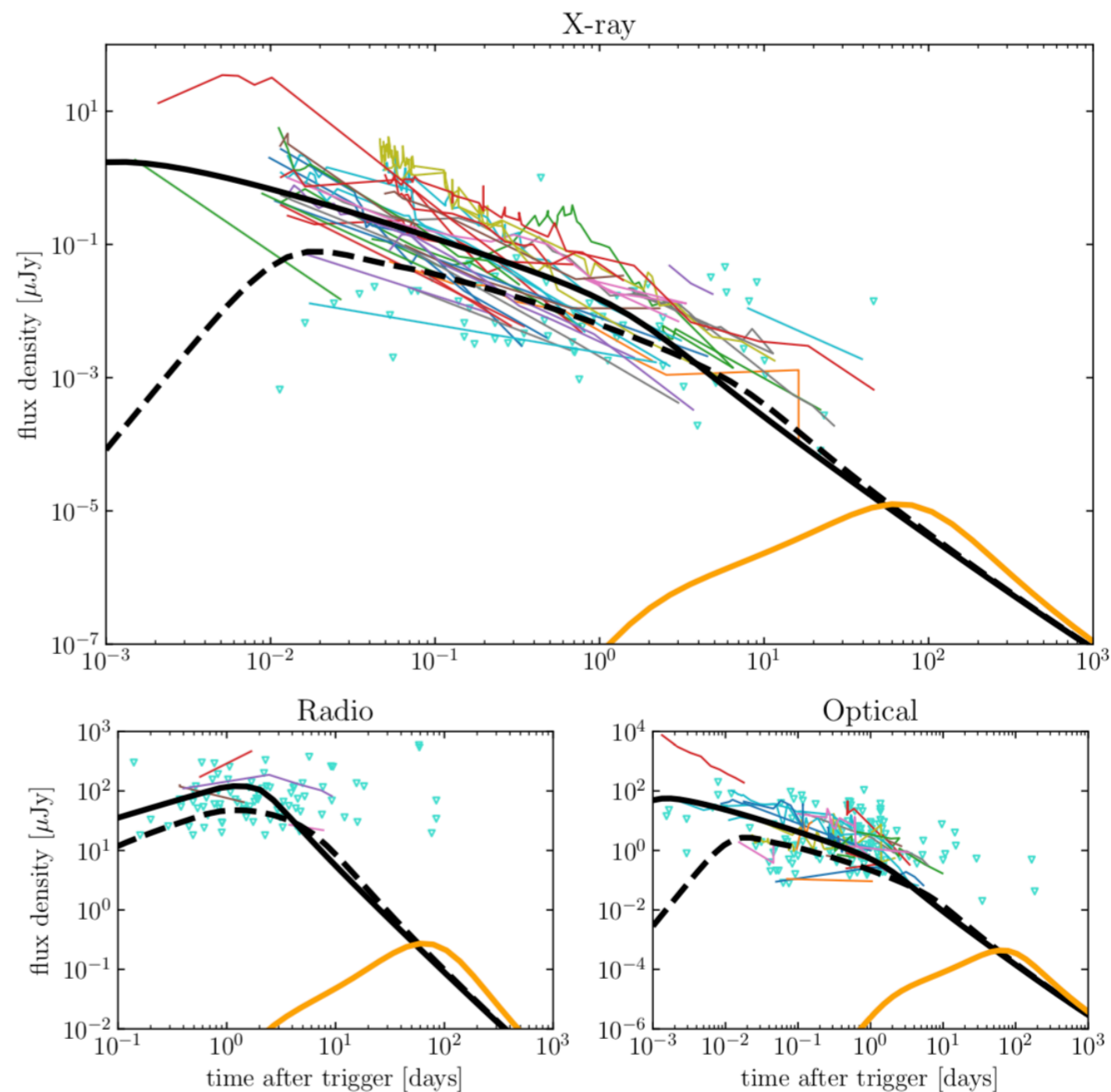
- ▶ Structured jet preferred to choked cocoon
- ▶ compact ($< \sim 2$ marcsec)
- ▶ **peaked brightness**



FERMI AND MULTIMESSENGER OBSERVATIONS

A UNIVERSAL SHORT-GRB JET STRUCTURE ?

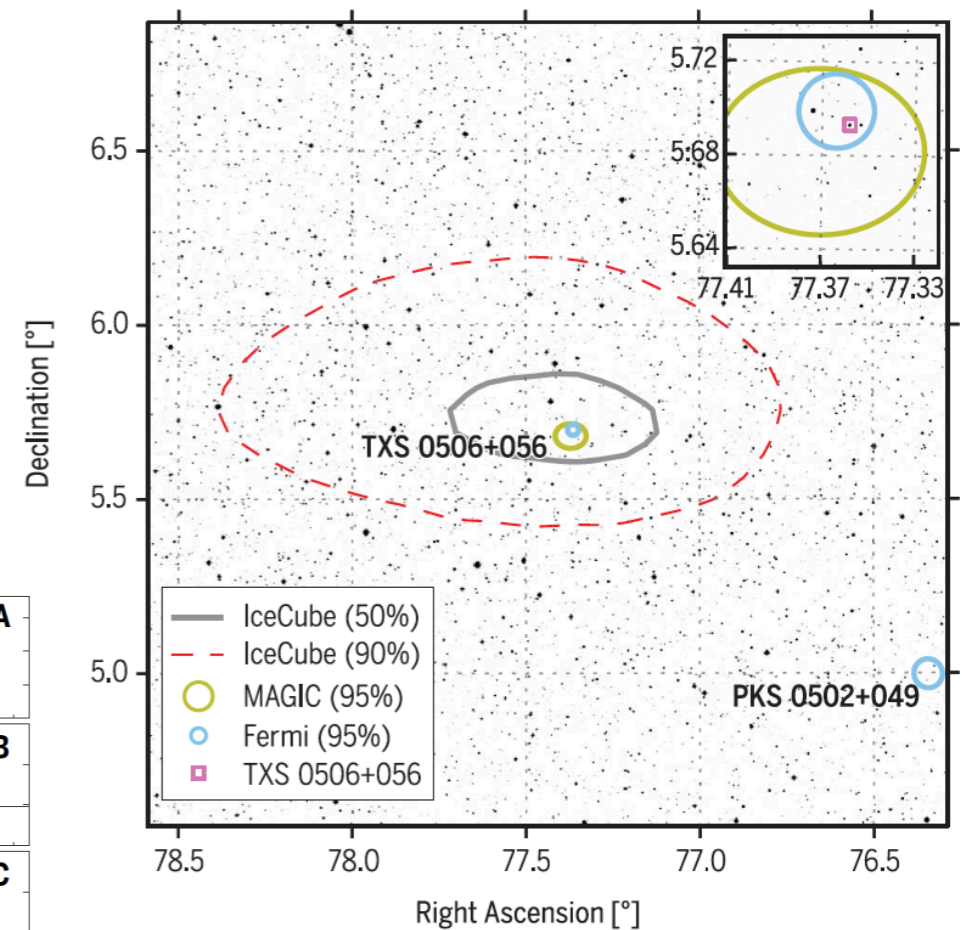
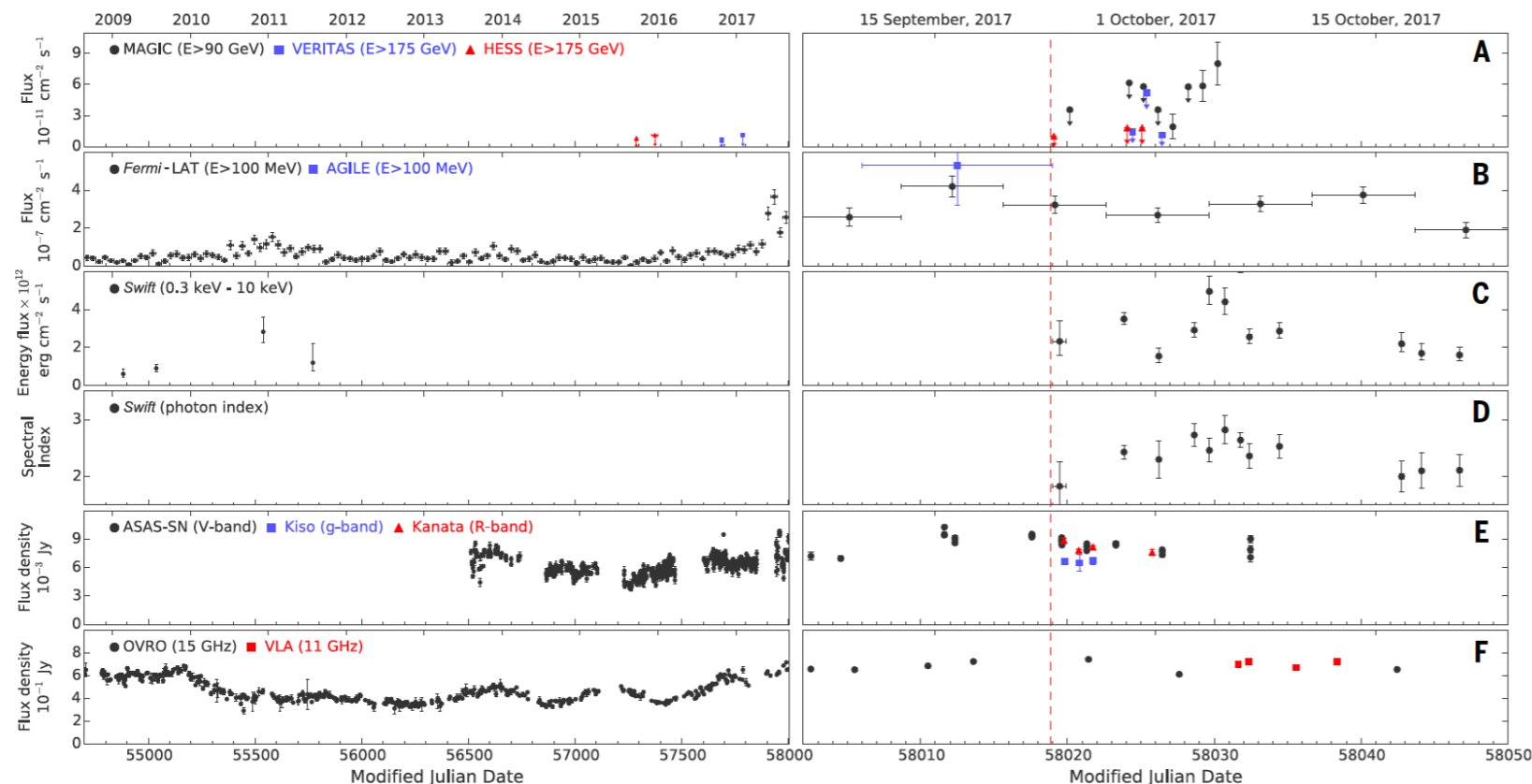
- ▶ sGRBs afterglows: archival data for ordinary GRBs with a GRB170817-like structured jet seen on-axis (black lines) or off-axis (orange line)
- ▶ diversity of sGRB afterglows attributed to external properties



FERMI AND MULTIMESSENGER OBSERVATIONS

NEUTRINO SOURCES ?

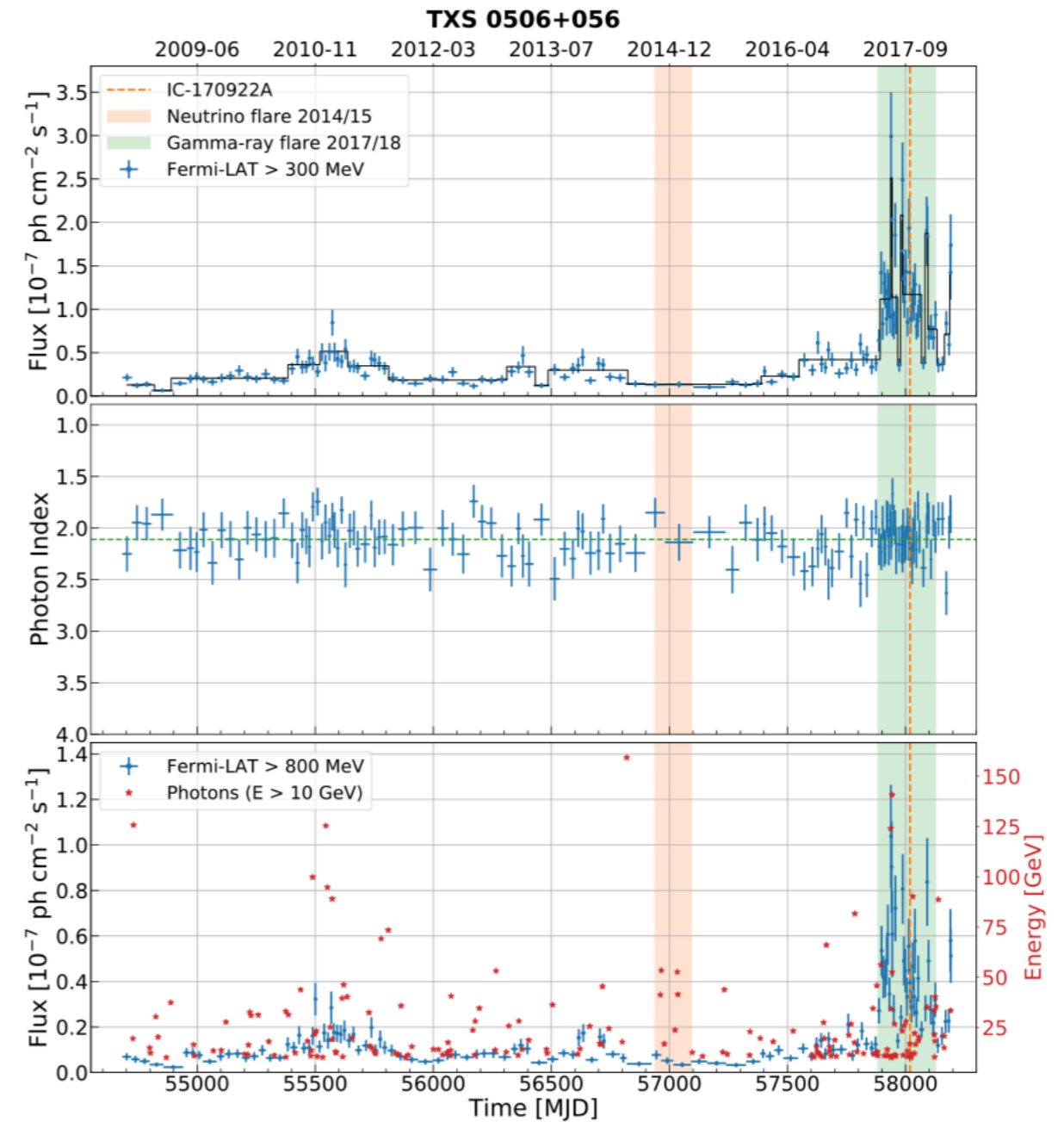
- ▶ TXS 0506+056 and IceCube 170922A association inferred from positional coincidence and concurrent high energy flare



FERMI AND MULTIMESSENGER OBSERVATIONS

NEUTRINO SOURCES ?

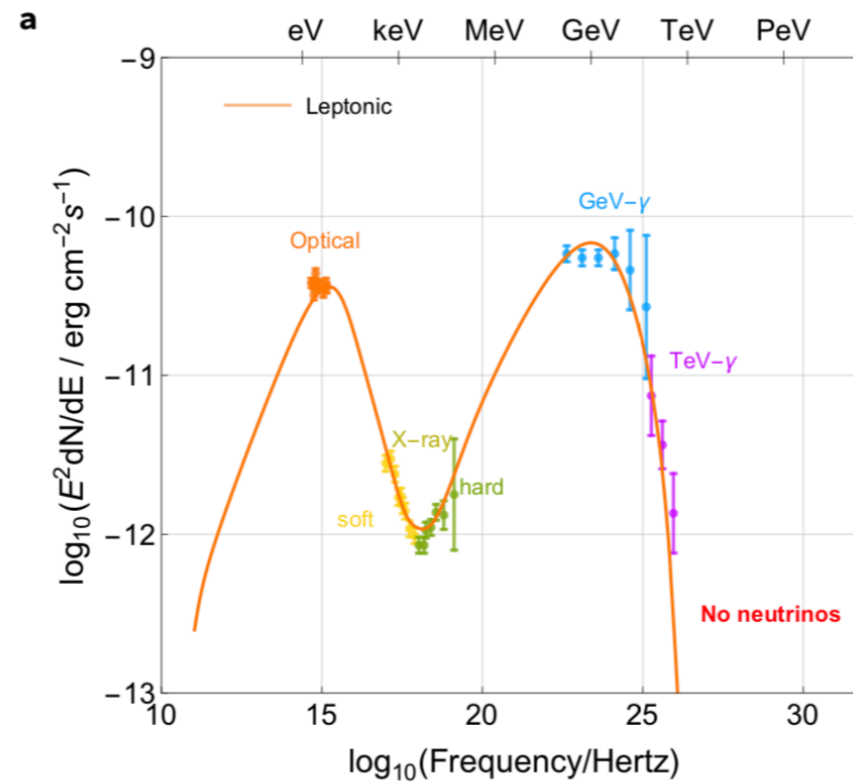
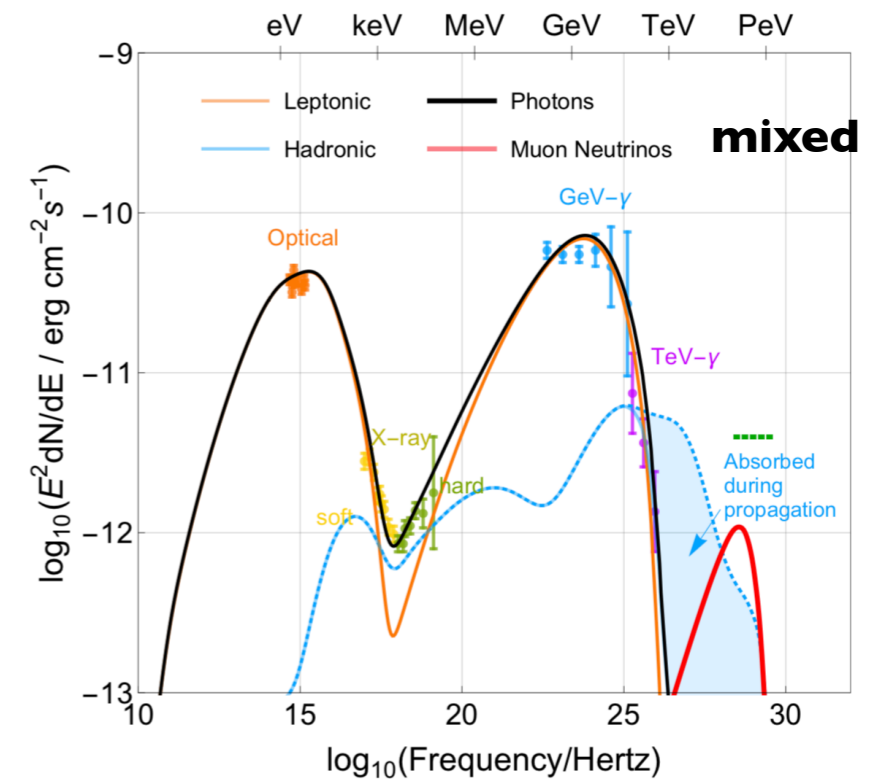
- ▶ Absence of gamma-ray emission with archival neutrino flare 2014/2015



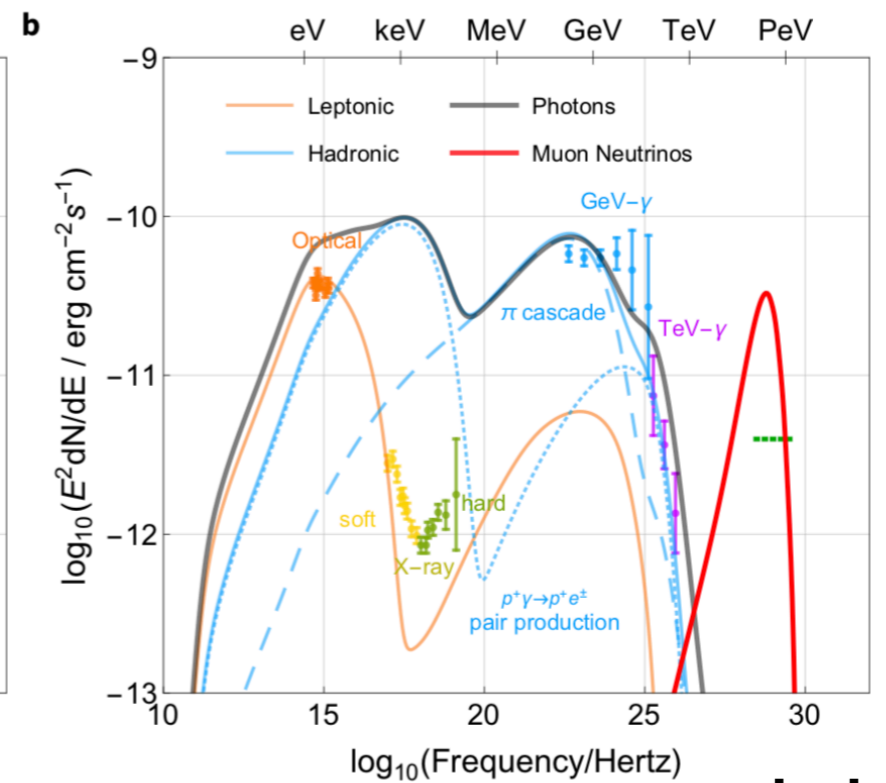
FERMI AND MULTIMESSENGER OBSERVATIONS

NEUTRINO SOURCES ?

- ▶ Difficulties in reproducing photon and neutrino measured fluxes with simple models



leptonic



hadronic

FERMI AND MULTIMESSENGER OBSERVATIONS

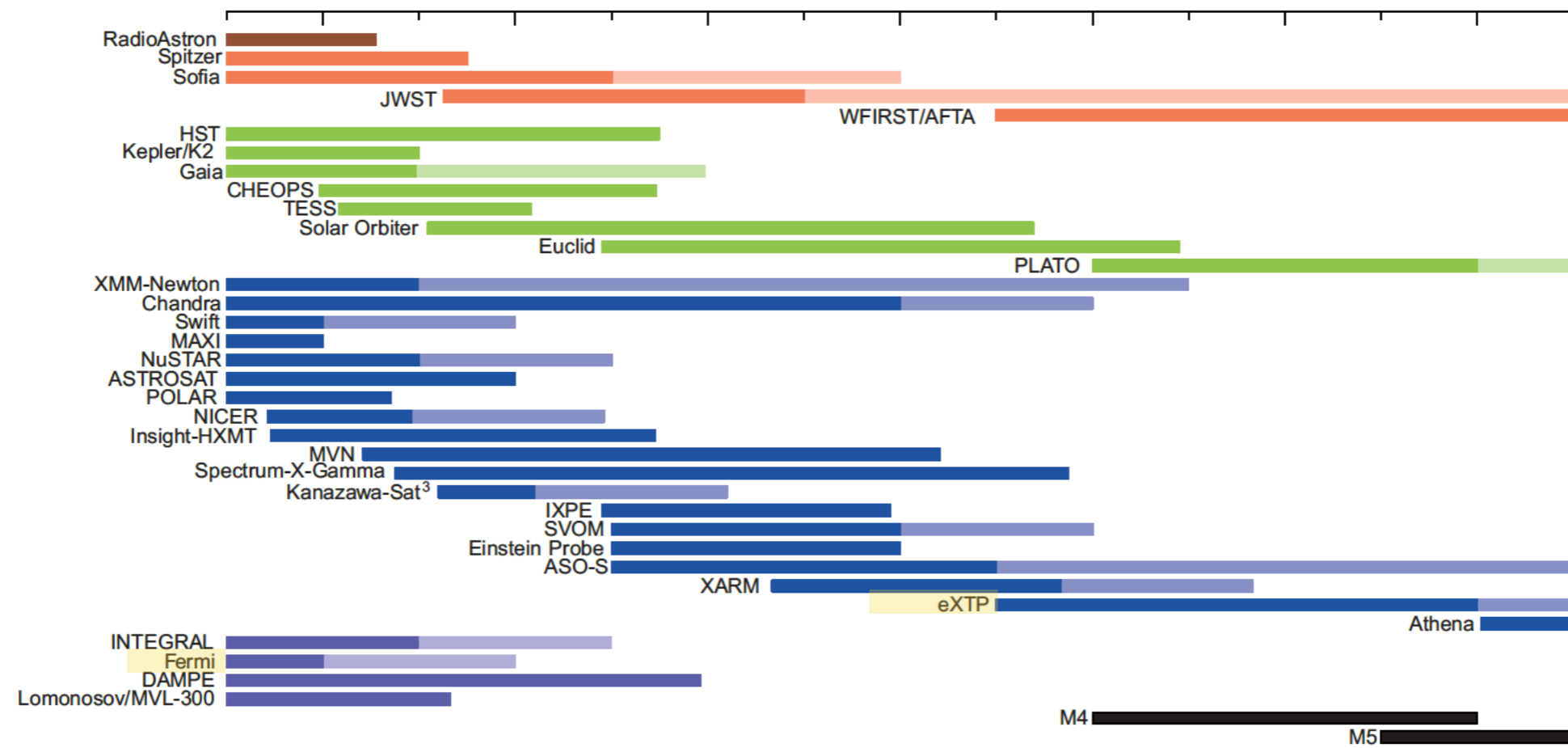
4 - PROSPECTS



FERMI AND MULTIMESSENGER OBSERVATIONS

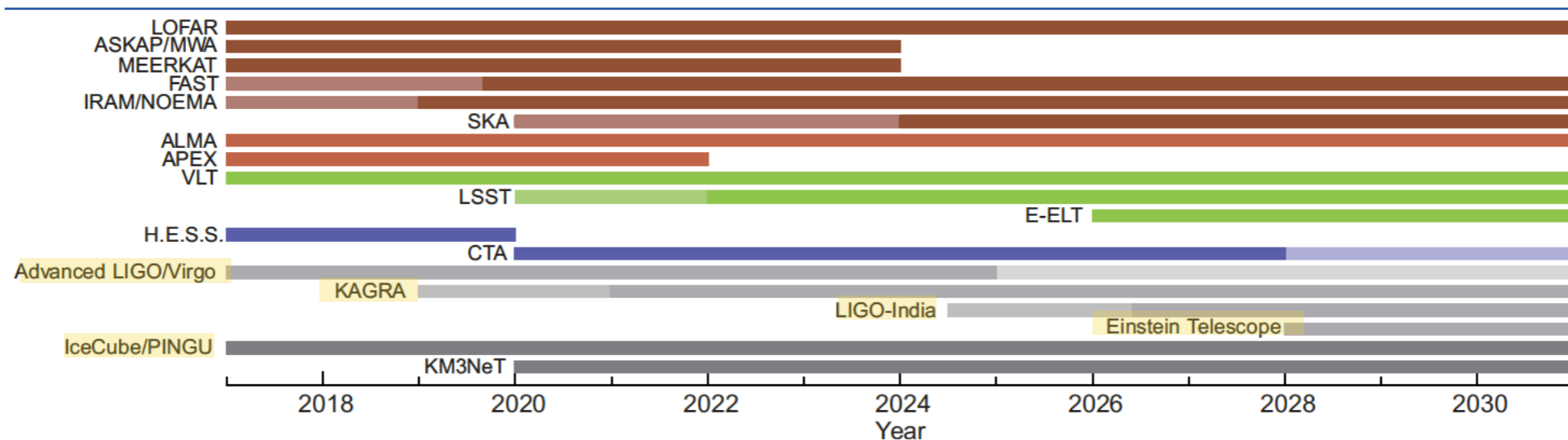
OBSERVATORIES

Space



Dark - funded
Light - lifetime

Ground

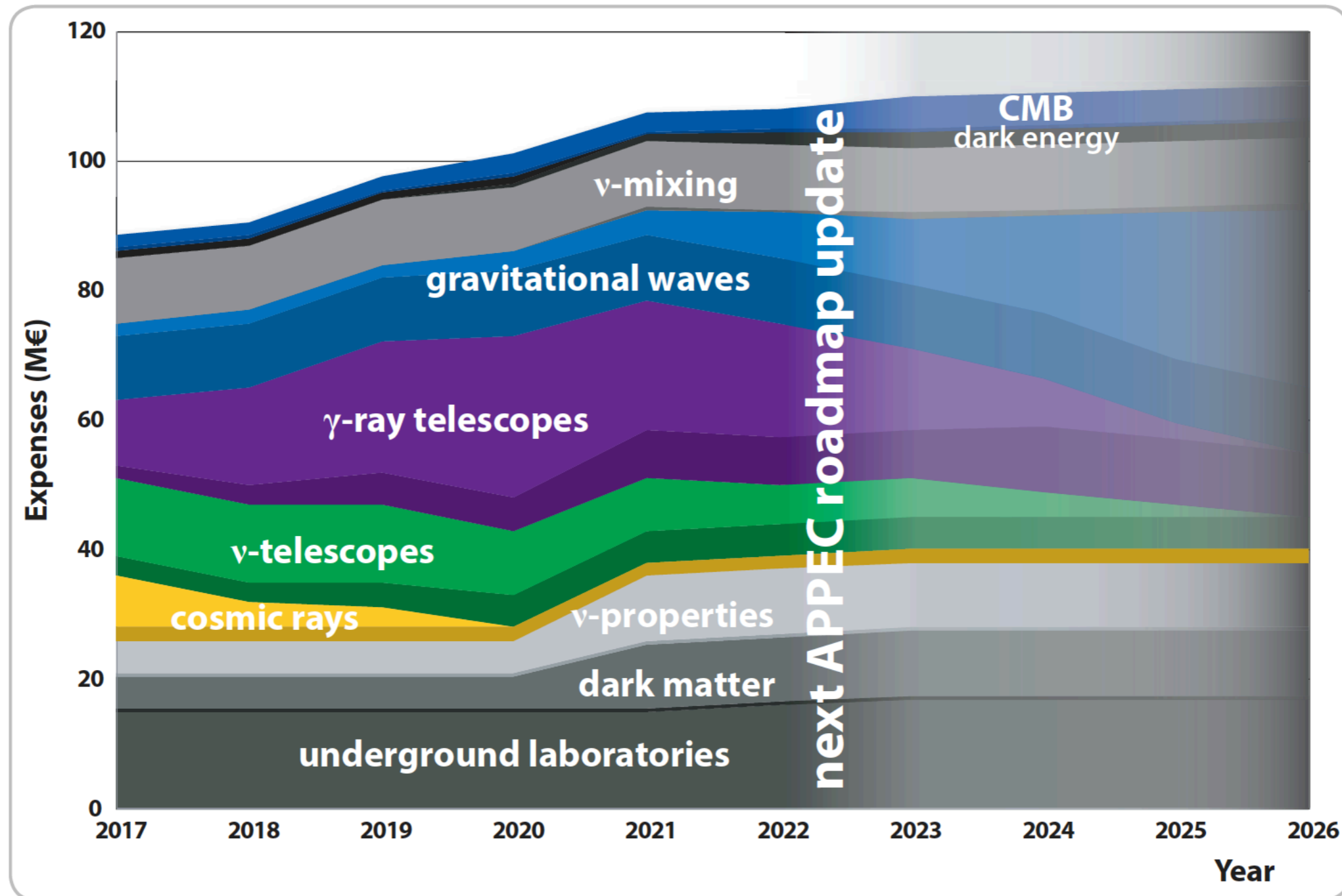


Red - radio
Green - OIR
Blue - X
Purple - gamma
Grey - GW nu

Survey - large FOV

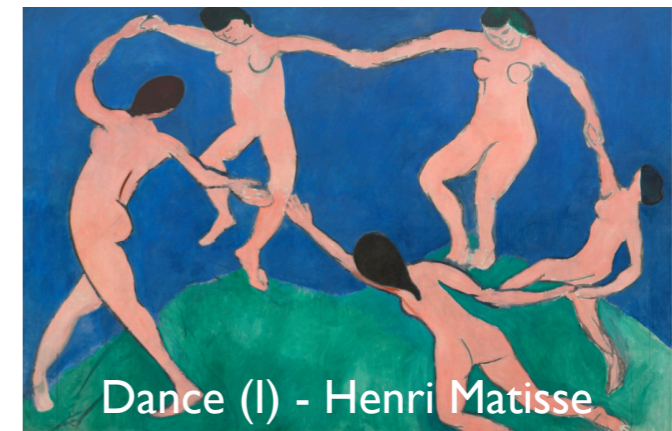
FERMI AND MULTIMESSENGER OBSERVATIONS

FUNDS



FERMI AND MULTIMESSENGER OBSERVATIONS

CONCLUDING REMARKS



- ▶ Observational multi-messenger astronomy starts in 2017
 - ▶ still only two concurrent observations of major events
- ▶ Progress comes from a new interdisciplinary community
 - ▶ data from many observatories must continue to flow
 - ▶ requires dedicated efforts and investments
 - ▶ complementary scientific backgrounds and cultures are key
 - ▶ to complete the broad picture from major events and non concurrent multi-messenger data