

M.D.I.

Summary

E.P.
for

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Simone, and the computing team

General Considerations

- **13** talks on background on the meeting agenda
 - 8 speakers + several coworkers
 - Critical mass reached by many subsystems
- **1** Fast Sim to Bruno event converter (Dana)
 - Bruno will be shortly able to simulate generic BBar events
- **5** talks on the IR side of the MDI
 - 1 speaker
- **4** talks on the MDI
 - 1 speaker

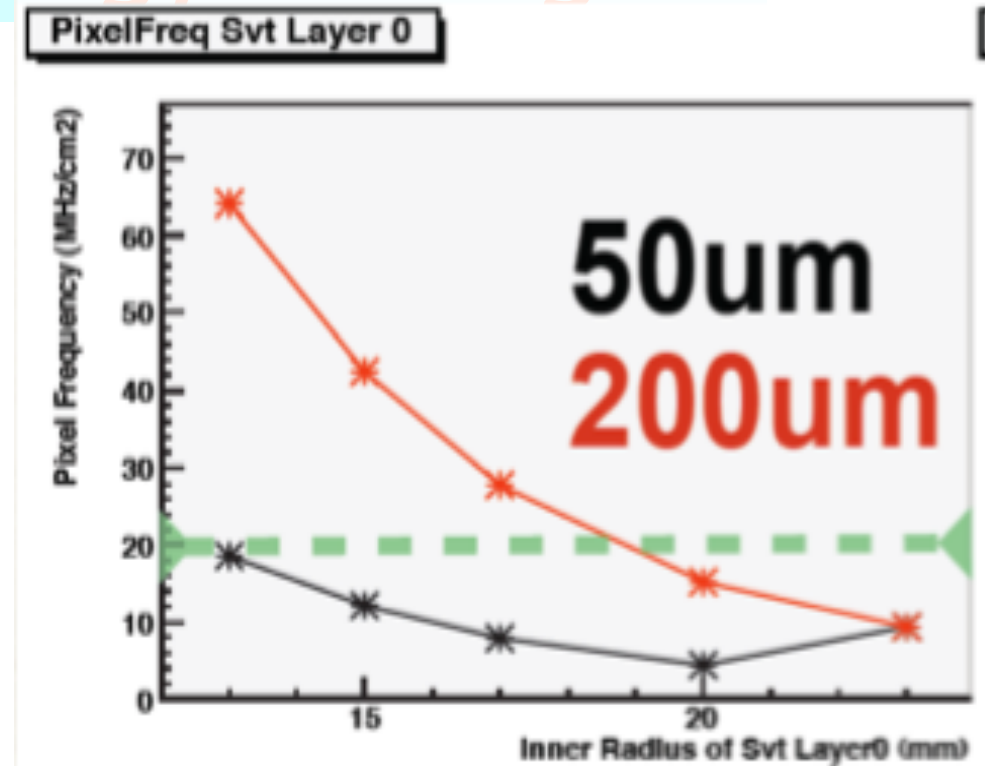
General Considerations II

- **Up to now the MDI Detector Side focused his attention on background simulation, lot of progresses**
 - Simulation Software
 - Computing Infrastructure
 - First steps toward Candidate Reconstruction
- **Time to aggress the MDI from the IR Detector Side**
 - Negotiation of the Machine, Detector vital spaces
 - Current assumption
 - Machine: $r < 24$ cm and $|\theta| < 0.300$ radian
 - Detector: $r > 25$ cm or $|\theta| > 0.300$ radian
 - Many missing elements: SVT read out, services, radiation monitors, luminosity monitors etc...

Layer0 radius & technology vs bkg.

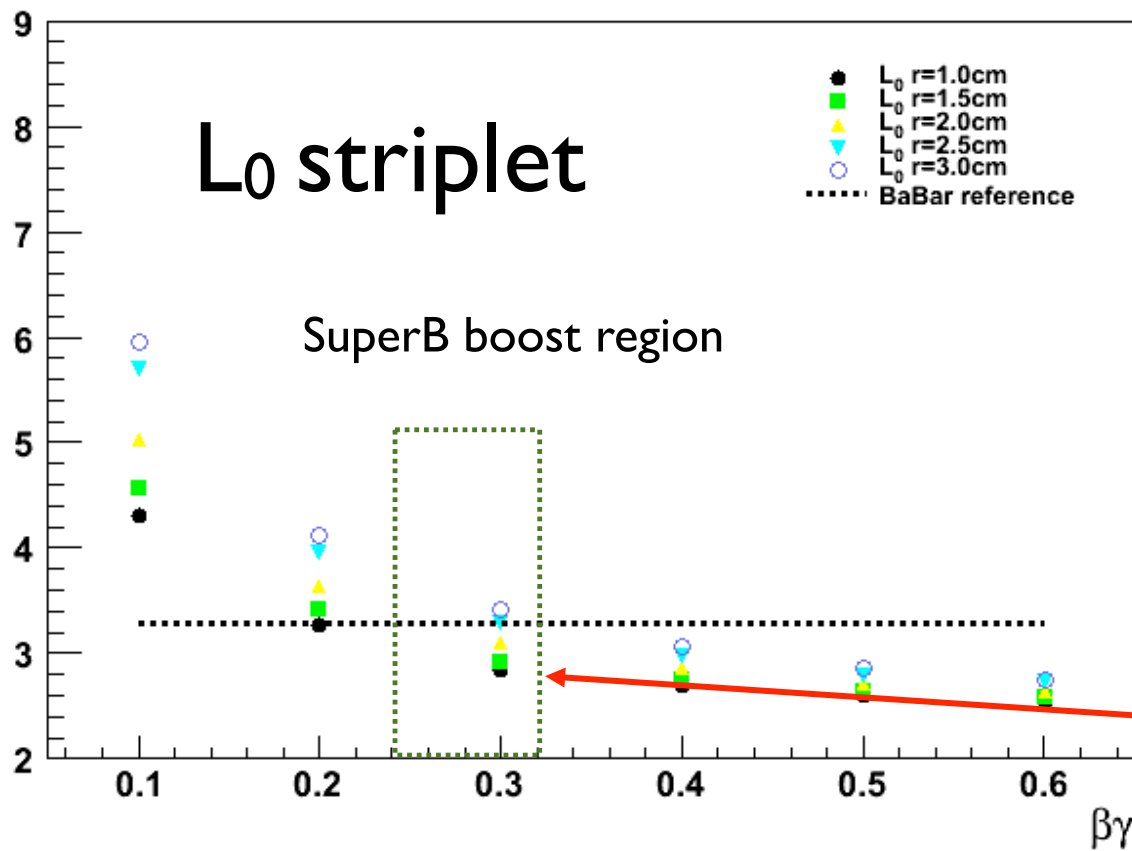
Update on background:

- Hit rate vs Layer0 radius from pairs production depends strongly on sensor thickness:
 - on thick sensor larger cluster width for low momentum tracks with large crossing angle
- Large difference for thin pixels (50 um) and triplets (200 um)
- Hybrid pixel with 200 um sensor will be like triplets, unless thinner sensor can be used



Sustainable background hit rate (radius) depends on technology: triplets vs pixel area and readout chip.

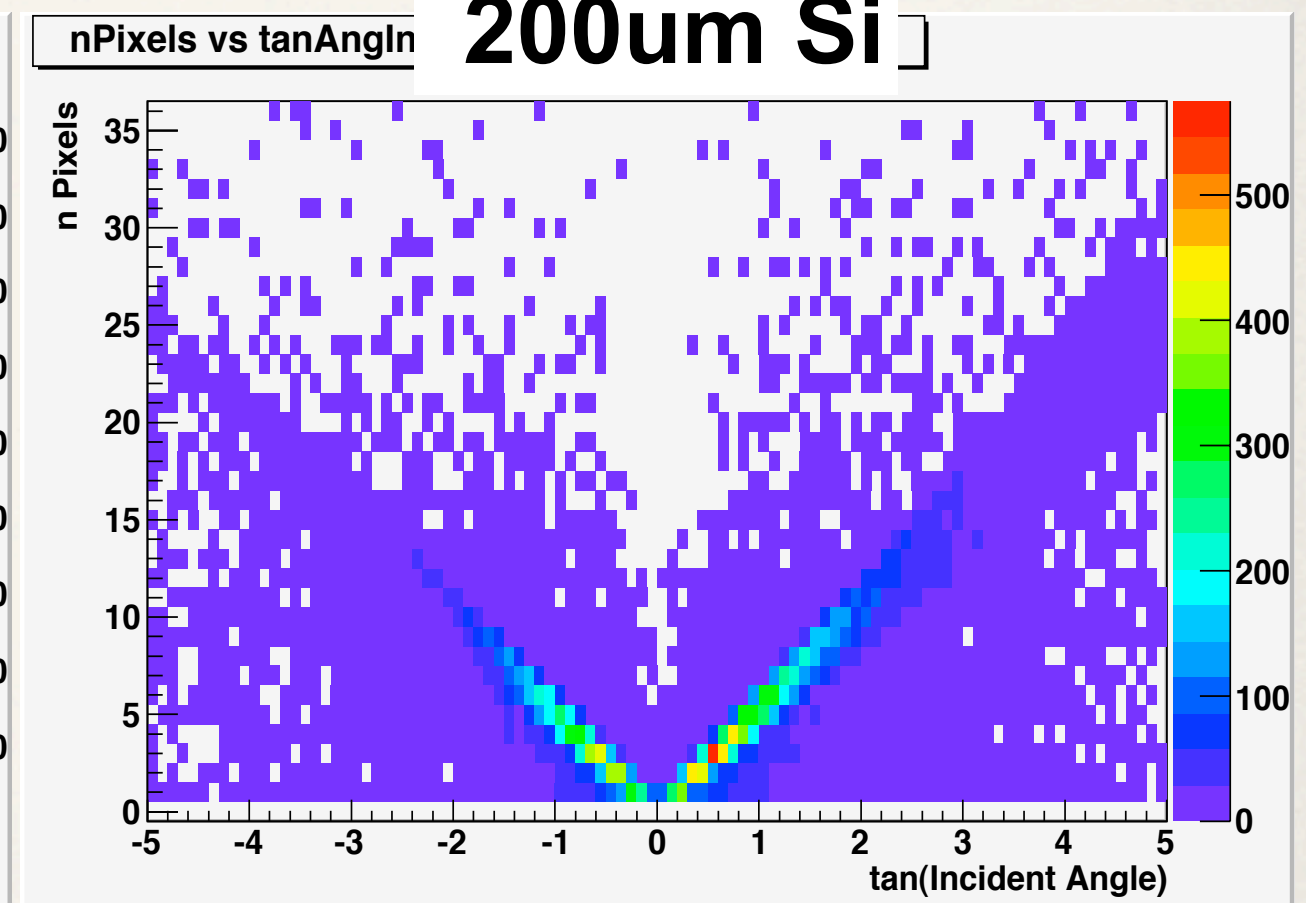
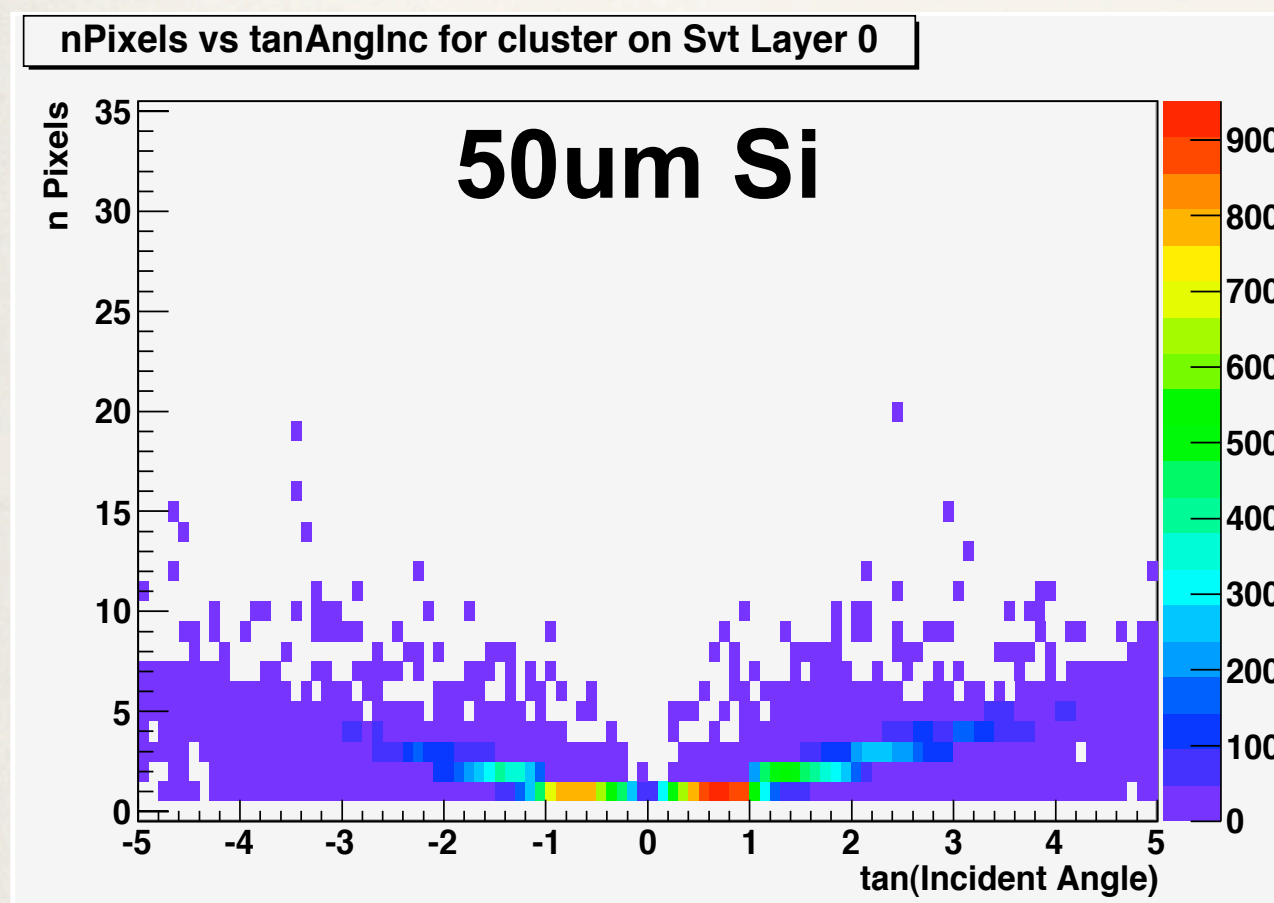
- Development of thin pixel chip readout architecture continue: data push and triggered with target 100MHz/cm² (safety x5 included) with timestamp 100 ns. → $R \sim 1.3\text{cm}$
 - Still to demonstrate: scaling to large matrix, rad hardness for MAPS,
- Assumed 100MHz/cm² hard limit for triplets (~10% occupancy in 100 ns, area $\sim 10^{-2}\text{cm}^{-2}$) → $R \sim 2\text{cm}$
 - performance similar to BaBar and thin pixel at lower radius. No margin left!



Pixel multiplicity

Will be compared
with test beam real data

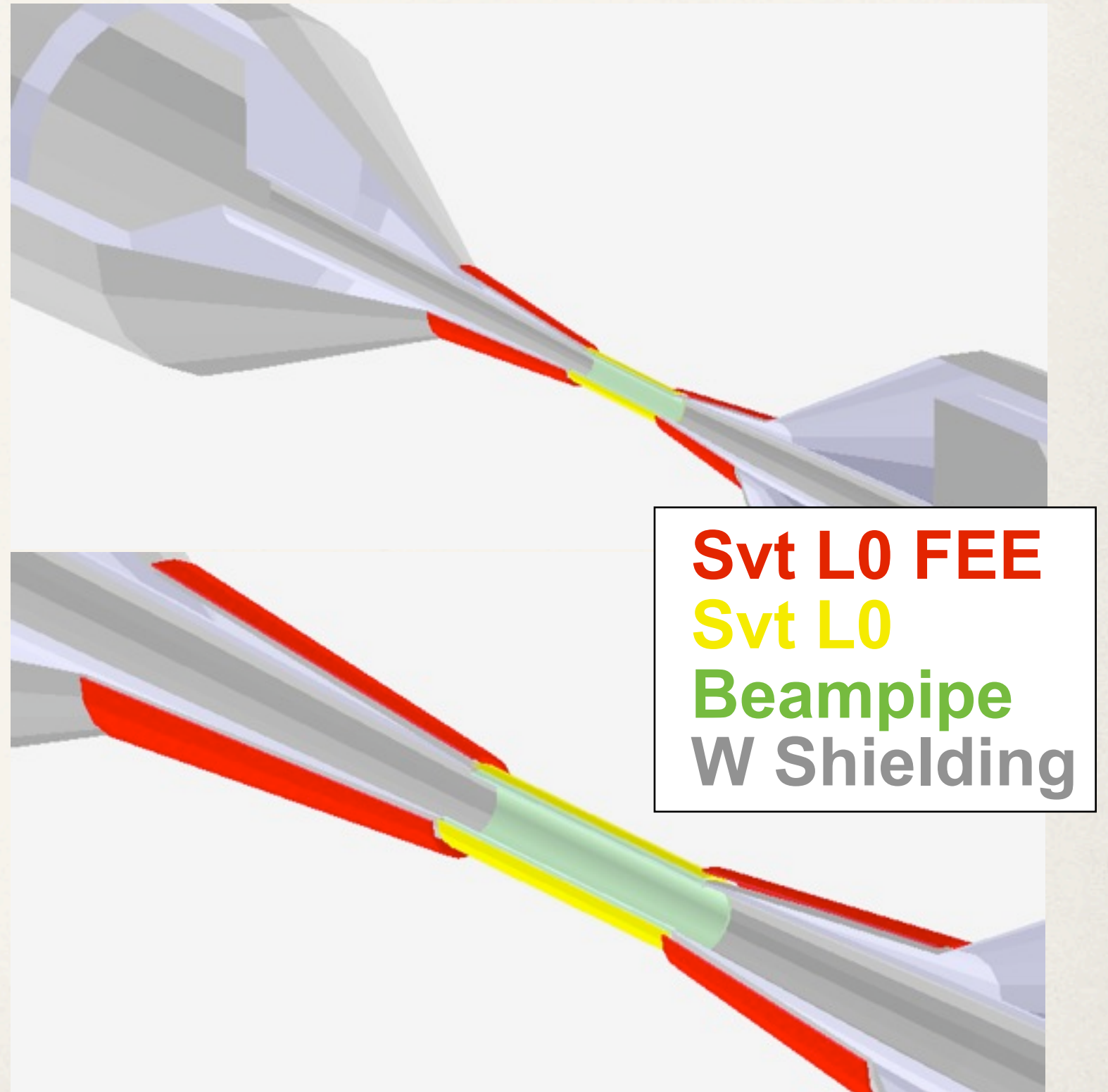
- Multiplicity proportional to tangent of incident angle on Si
- Simulation with different thickness of L0 silicon:
 - 50um, monolithic pixels device
 - 200 um, striplets device



- $\text{Tan} < 0$ when particle is going from outside to inside L0

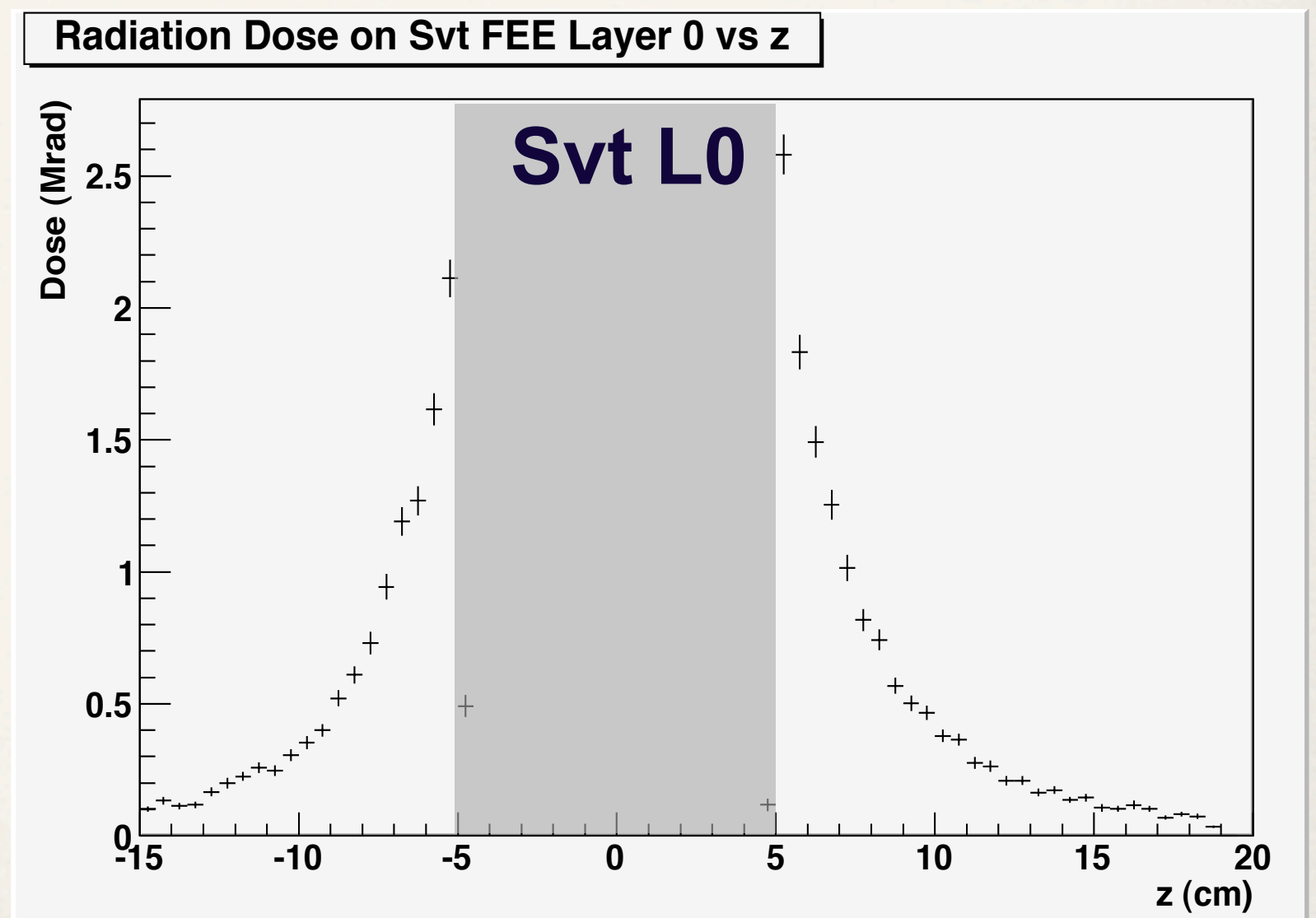
L0 electronics position

- Additional 2 volumes
 - Cones around IR tungsten shielding close to L0
 - 1mm of Si at 2mm from shields (radiation probes)
- Sensitive volumes: additional BrnRootHits list dumped by RooEvt object



Radiation on L0 electronics

- Relevant information: Integrated Dose (1 nominal year)
- First test with pairs bkg (40k evts)
- Average dose:
460 krad
- Much higher close to L0 edges: >2 Mrad
- Same technique can be used for other areas or using also more realistic materials

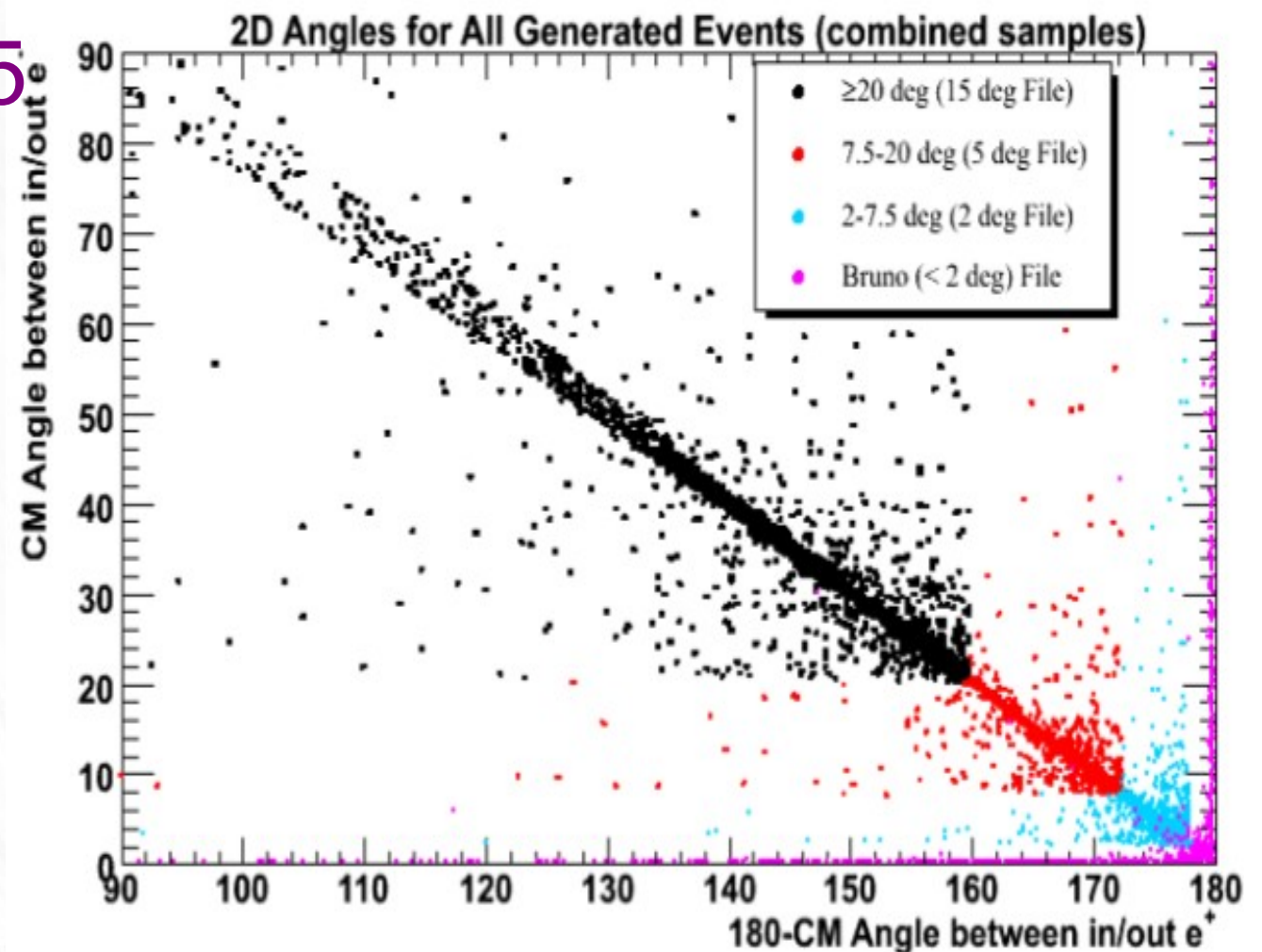


Drift Chamber

- **Bhwide** generator in **FastSim** (No Bhwide with Bruno)
- Transfer events to FullSim by converting **StdHepAsciiDump** output to **guinea generator** input
- Create tuples with e+e- at **2-178, 5-175, 15-165 degrees** (CM frame)

Degrees	Cross-Section (nb)
2-178	7171.77
5-175	876.348
15-165	81.6761

- Combine with tuples created using **Bruno's Bbbrem** generator.
weight = 4.644 ns^{-1}

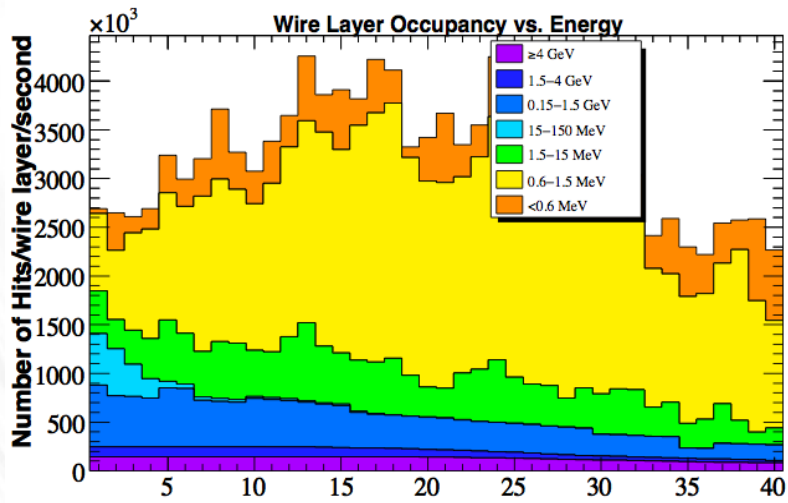


June 2, 2010

Dana Lindemann - McGill

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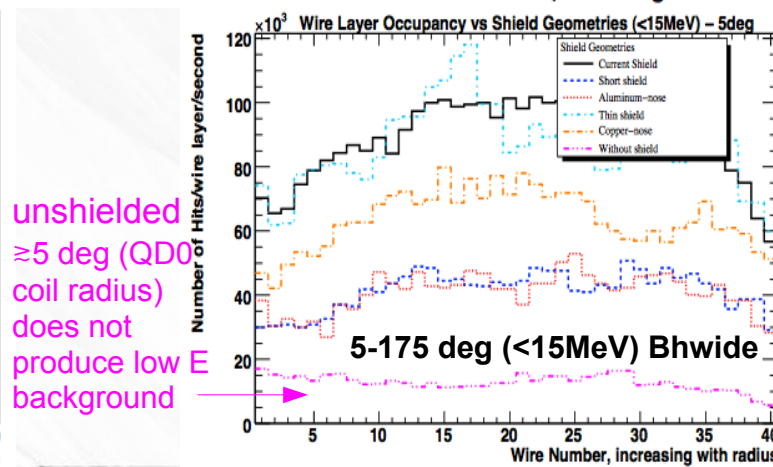
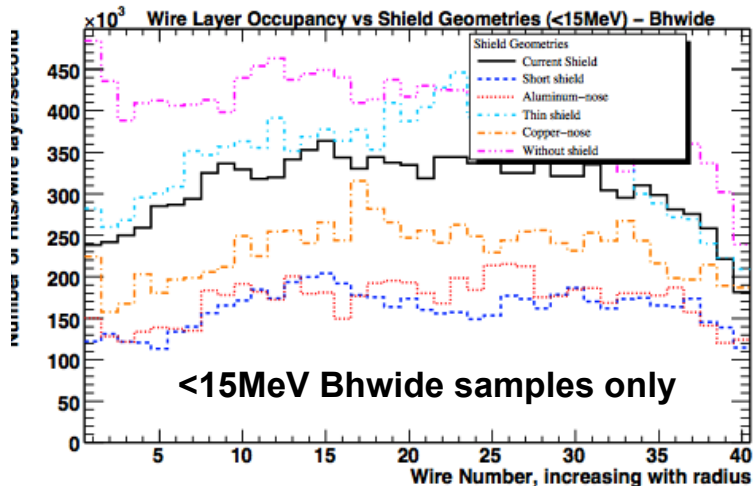
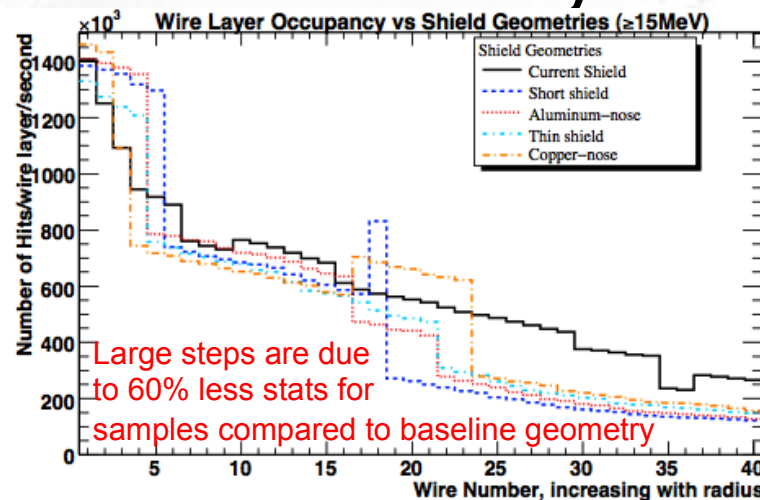
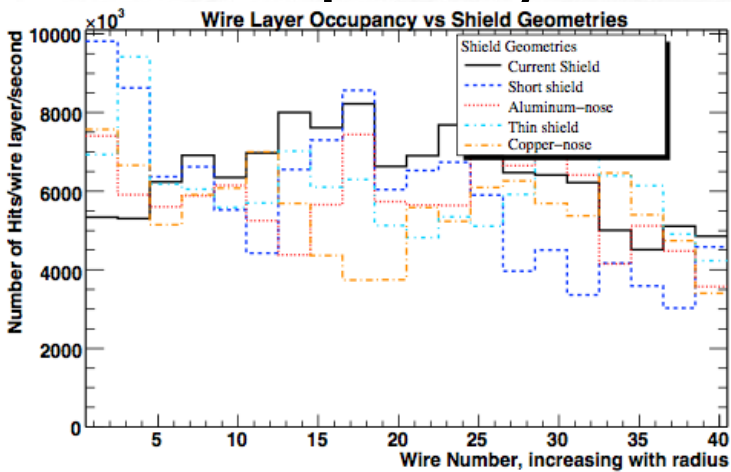
Track reco.!! Shield optimization



Tracking Algorithm

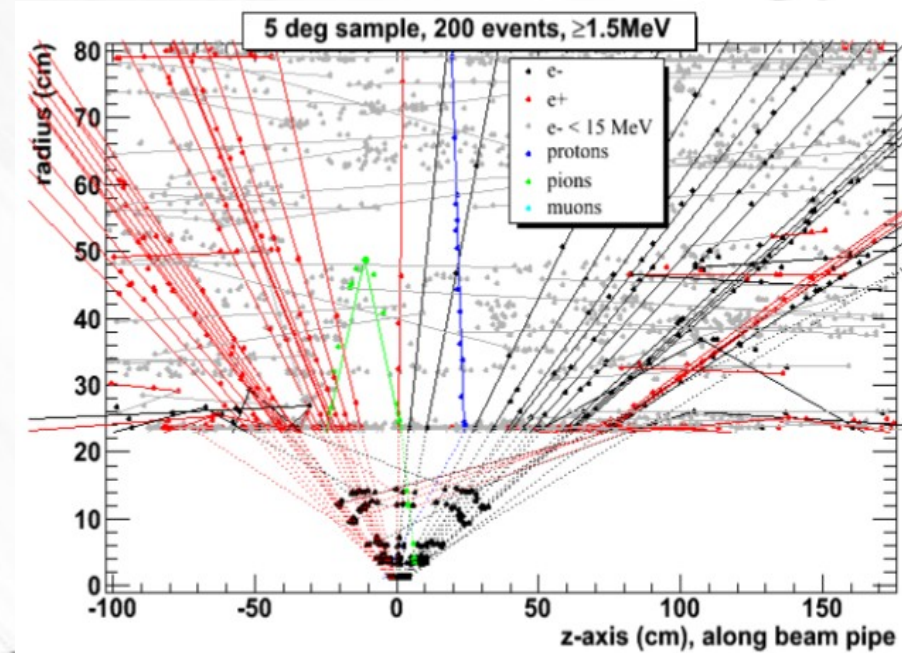
- Bruno only provides deposited energy (hit) information within a chamber that's void of wires. All wires are assumed to be axial and uniformly spaced.
- Using the TrackID of time-ordered hits, I define a track and extrapolate the number of wires the track would cross.

Occupancy vs. Shield Geometry



ne where track enters/exits DCH:

CH are straight
 line
 x radius,
 tion
 CH
 straight
 hit.
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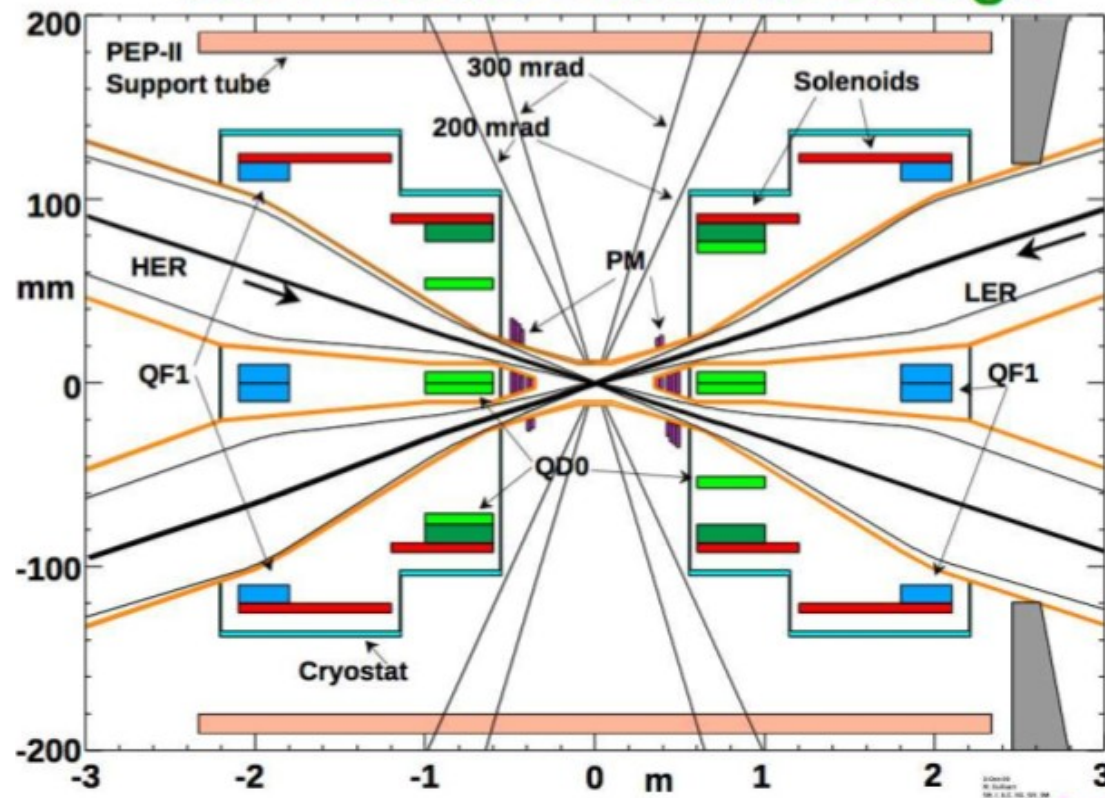


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Hot Spots along the beam line

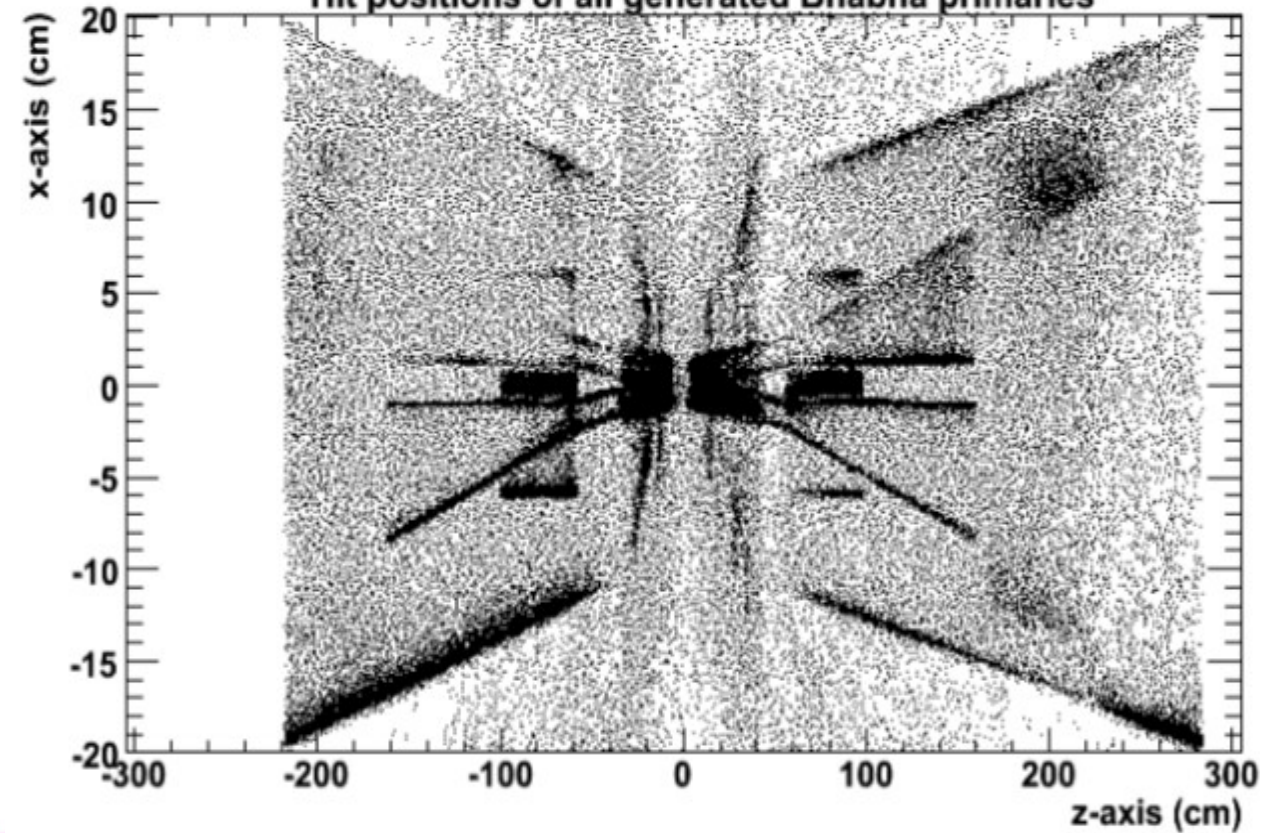
All Daughter vertices of All Bhabhas

The Present Baseline Design

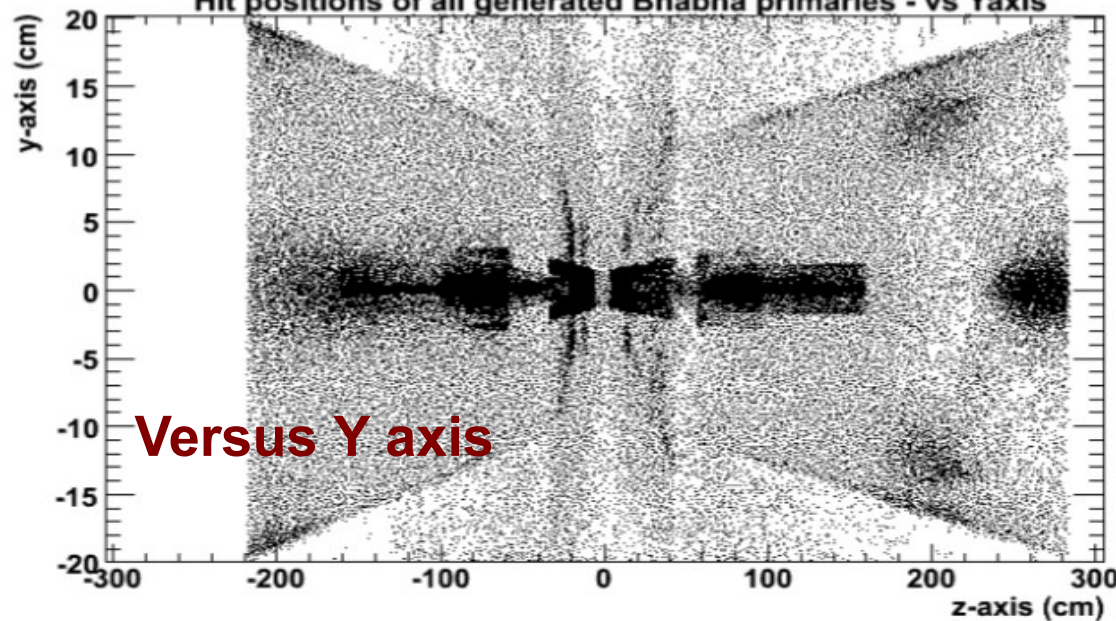


SuperB Workshop
March 16-19, 2010

Hit positions of all generated Bhabha primaries

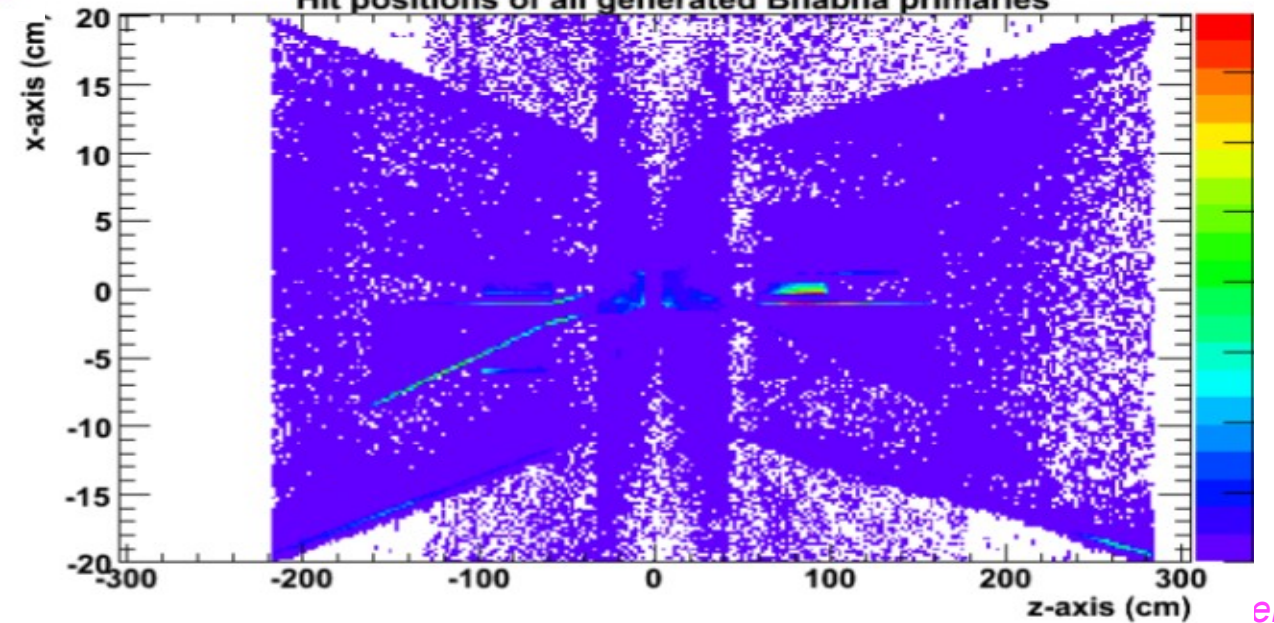


Hit positions of all generated Bhabha primaries - vs Yaxis



Versus Y axis

Hit positions of all generated Bhabha primaries



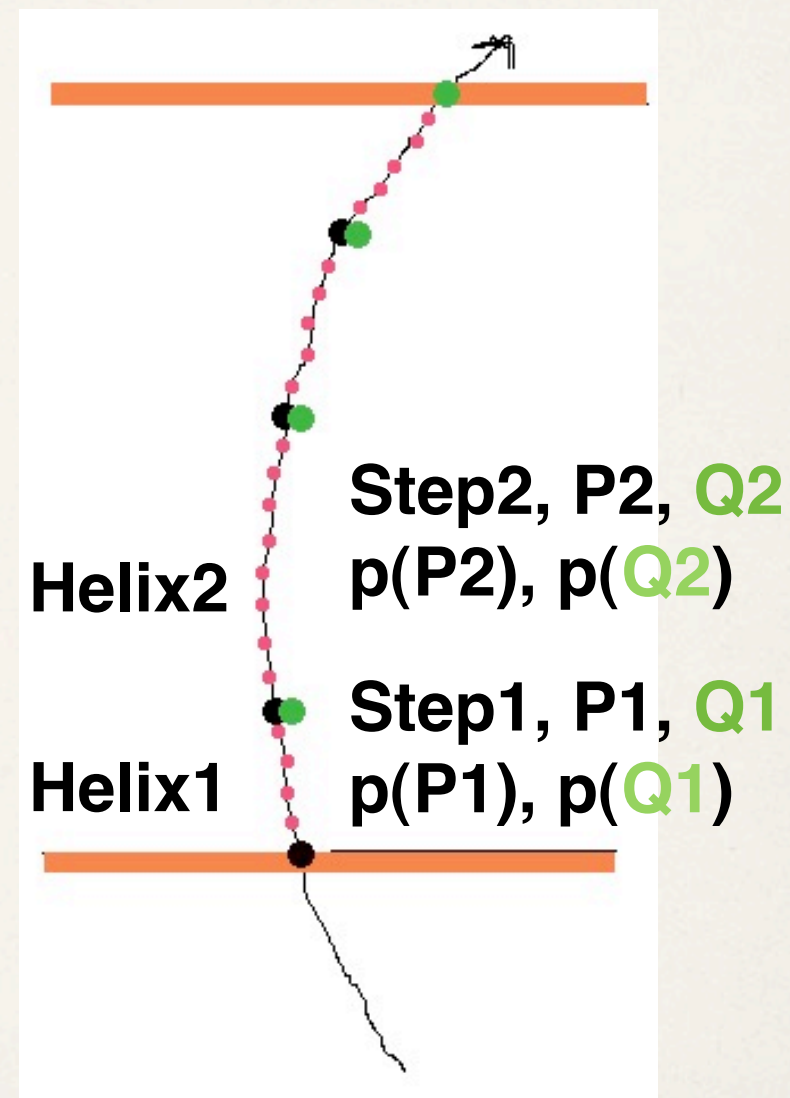
erB XIII
1-4, 2010

Track Reconstruction II

	Old method	New Method
Occ (no step limit)	2.9%	4.7%
Occ (max step 5cm)	2.9%	3.3%
Occ (max step 1mm)	1.35%	1.36%

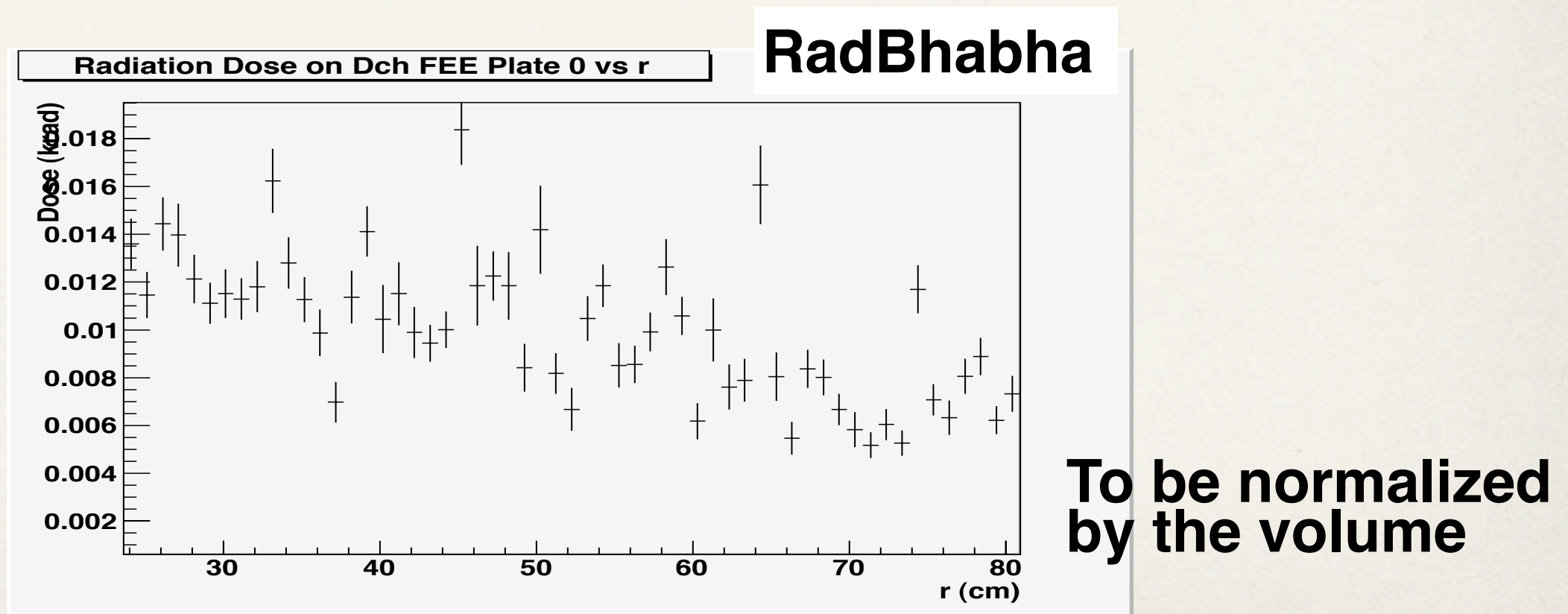
Tracking in a B field

- **Start** and **end point** are not enough for Dch hits, trajectories are helix
- Using the momentum direction and particle charge the helix parameters can be computed (standalone macro after the simulation)
- Then the helix can be **sampled** at a smaller sub-step (3 mm) and we got all the cells crossed by the particle in the step (sub-step energy is assigned to each cell)
- Steps that are shorter than 3 mm or with radius less than 6 mm are approximated with straight lines and sampled as well
- Last point of helix not always exactly match with step end point (multiple scattering), additional sampling of straight line again that connect them



Radiation dose on electronics

- 3 Aluminum plates behind backward endcap by Giuseppe
- Change to make them sensitive, additional list of hits, DCHFEEHits
- Radiation Dose in kRad, 1 nominal year
 - RadBhabha, P0 **0.57** krad, P1 **0.60** krad, P2 **0.69** krad
 - 2 photons, < 50 rad for all the plates
- Any number from Babar for check the consistency?



DIRC

Scaling from Belle-I

Blair's idea:

- **Perhaps, we could scale the total quartz rate in DIRC-like TOF counter from the known Belle endcap Aerogel PMT's window rate by using (a) luminosity ratio, and (b) glass/quartz volume ratios.**

Does it make sense to scale it from Belle-I ? We do not know for sure. But they have single pe response Aerogel detectors around the IP, which can be used for scaling. Most of the background is caused by soft a few MeV gammas, which deposit a few photoelectrons in the in windows, quartz, Aerogel, etc. One needs a single photoelectron detector to see it. Background in EMC, drift chamber or Si will not help us.

- 2 scenarii investigated:
 - Background scales like luminosity
 - Background scales like luminosity / 4 [Belle's current assumption]
 - Next slide shows results for this second hypothesis
 - Jerry is in contact with Japanese colleagues to find out where their assumption is coming from

DIRC

Scaling from Belle-I (cont'd)

- **Contribution from bars in active volume only, and using the scaling from the rate in Belle Forward Aerogel PMT at large radius and using 1/4 of the Lumi increase, one gets a rate per single H-8500 PMT of ~2.2 MHz/PMT, or ~69 kHz/double-pixel.**
- **We neglect the FBLOCK contribution because we can shield it well. We cannot shield the bar box inside the Magnet.**
- **We also neglect bar section located inside the steel, as it is well shielded. This is about ~1.2 meter length.**

⇒ Results agree within a factor 2 with Riccardo's calculations!

- Integrated charge:

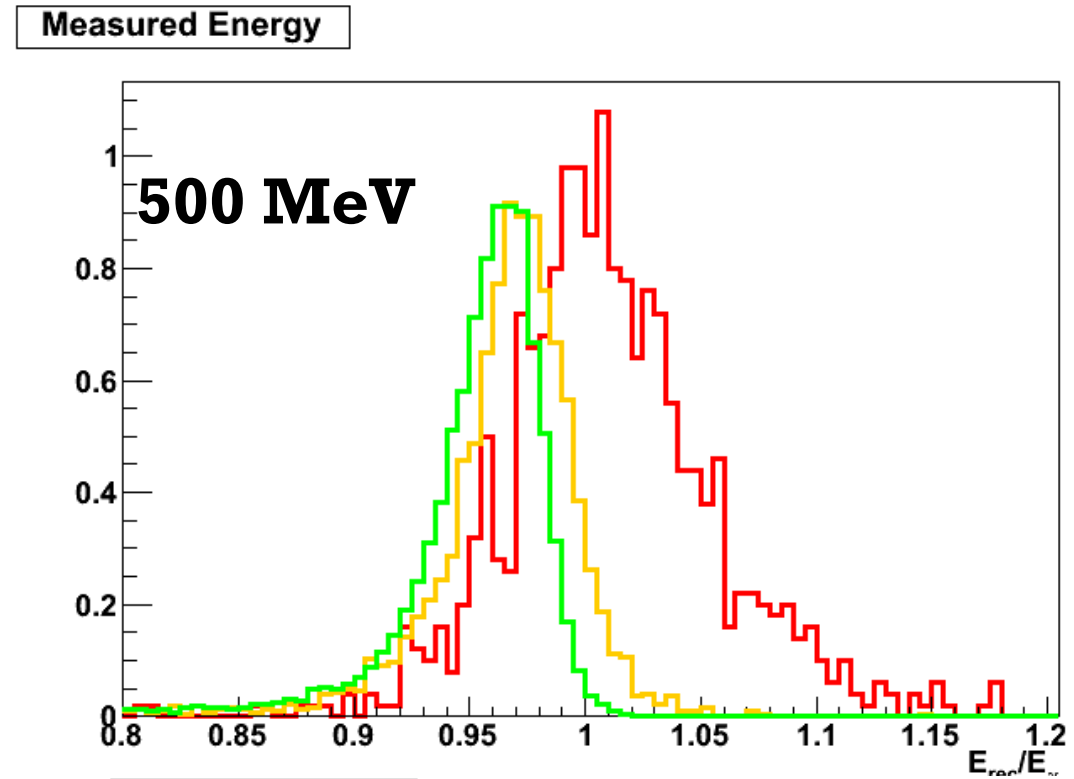
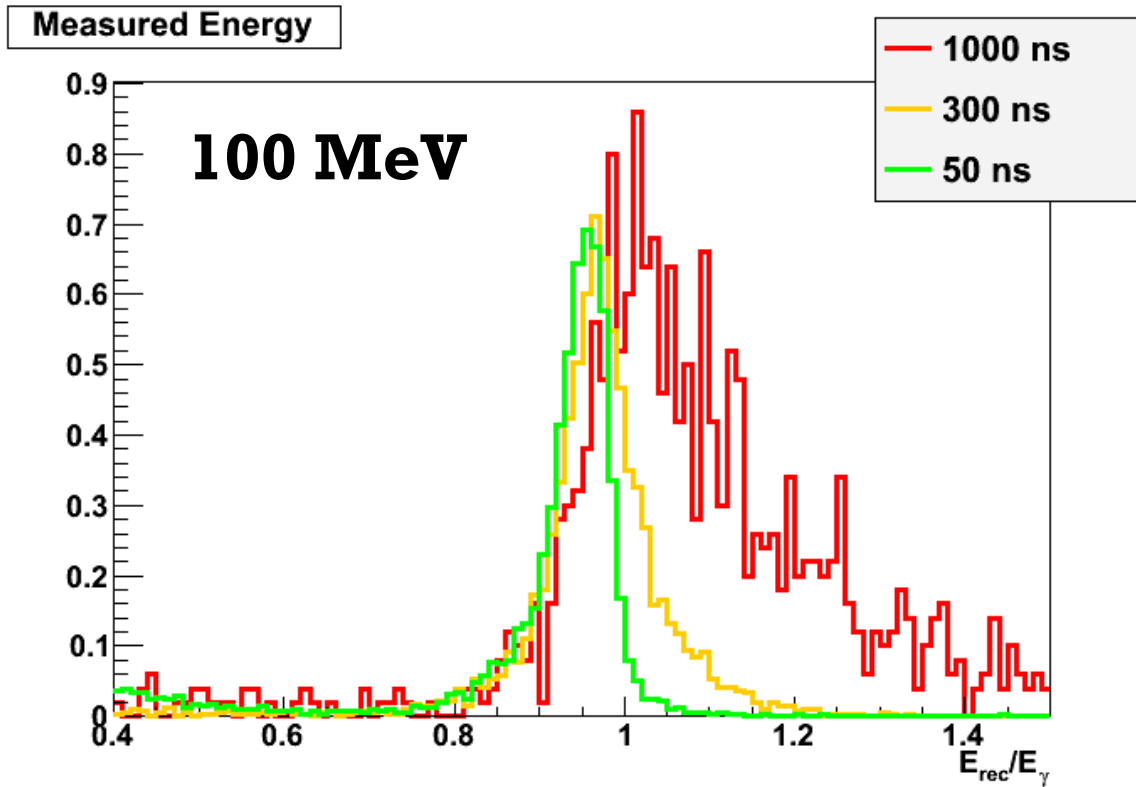
$$Q_{\text{DIRC PMT}} = 10 \text{ (years)} \times 2.2 \times 10^6 \text{ (Hz)} \times 2 \times 10^7 \text{ (sec/year)} \times 5 \times 10^5 \text{ (gain)} \times 1.6 \times 10^{-19} / (5^2) \sim 14.7 \text{ C/cm}^2 / 10 \text{ years !!!}$$

(MaPMT active size)

→ High number! Long-term consequences (ageing, etc.) will have to be studied

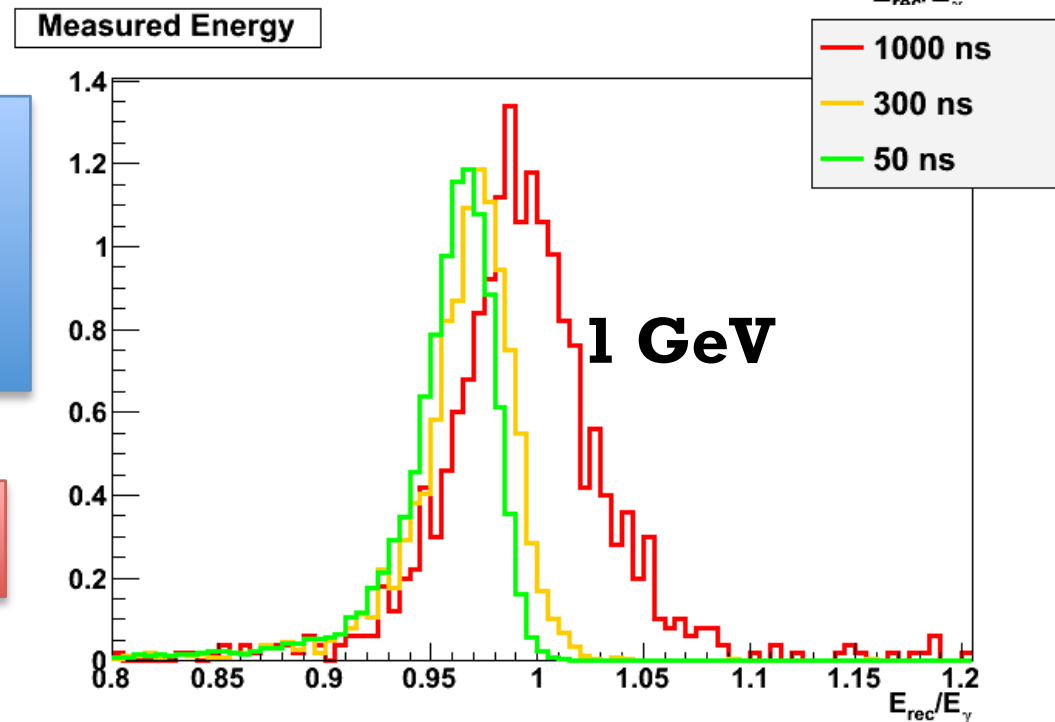
EMC

Time Window Width



For large windows resolution and distribution shape entirely dominated by background

From Now On Time Window = 300 ns



EMC II

Conclusions

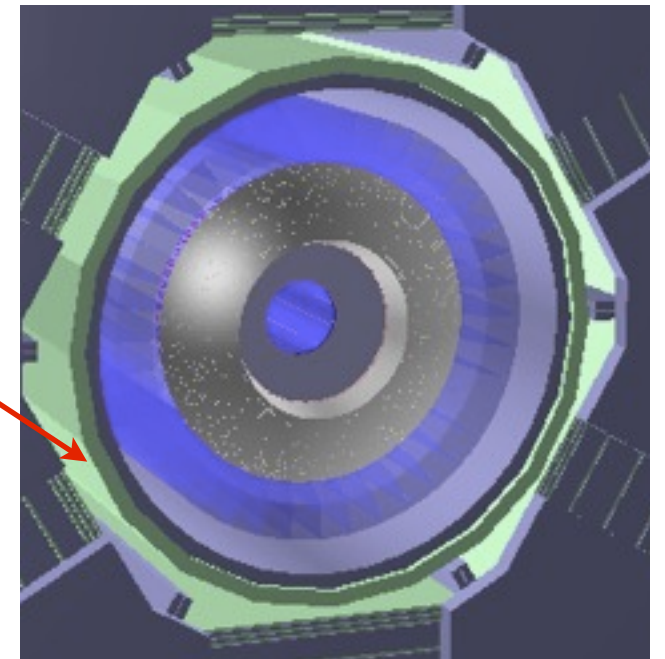
- Background has not negligible effect on reconstructed energy distribution
- Clustering optimization has an impact on the resolution and on the number of clusters
 - Emax cut affects heavily the number of clusters
- Time window width has a dominant effect at all energies
- BaBar Time Window is 240 ns
 - Already a tradeoff between light yield and background reduction
 - Something similar may be needed again
 - Readout chain simulation has to be more realistic

EMC III

- **Present Bruno Digi structure for EMC does not provide enough information to model the time response of the detector**
- **Plan to write the relevant information for the next round of simulation**
- **General comment**
 - **From now on we do not want to reenvent the wheel. Please read BaBar internal notes, ask BaBar expert the way things get done in the good old times.**

SHIELD SETUP

A polyethylene shield has been inserted between solenoid magnet and barrel.
We want study the impact of this shield on barrel rate.



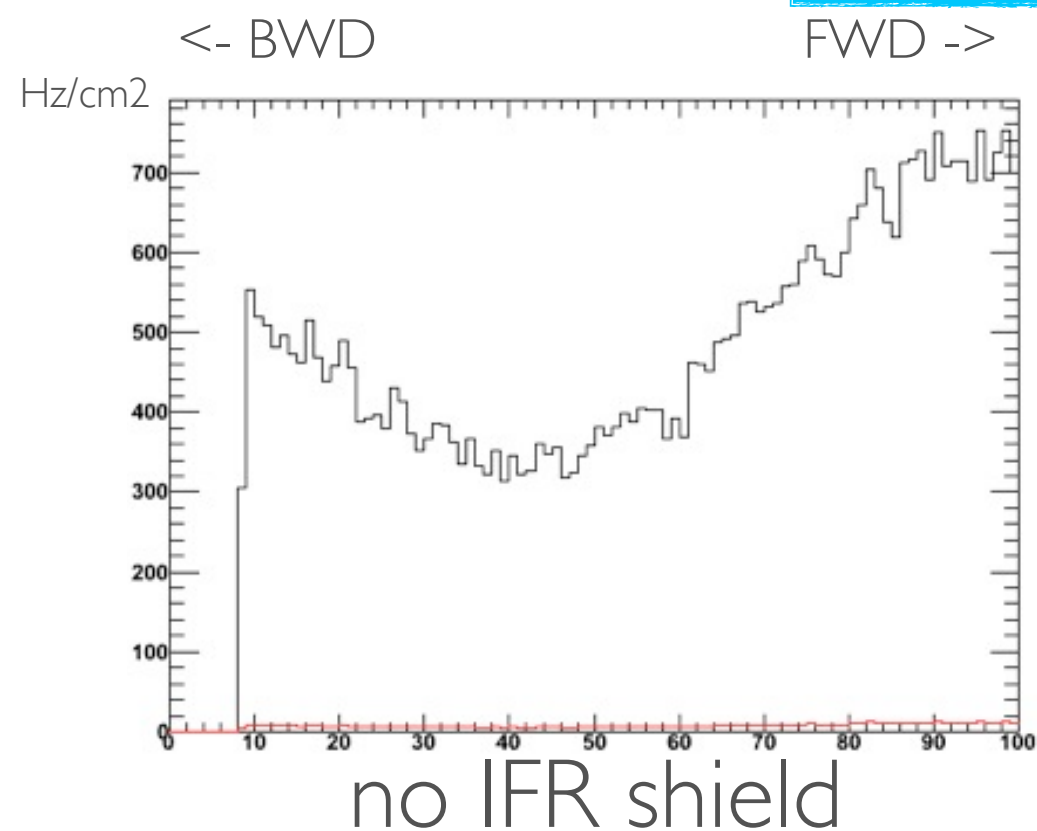
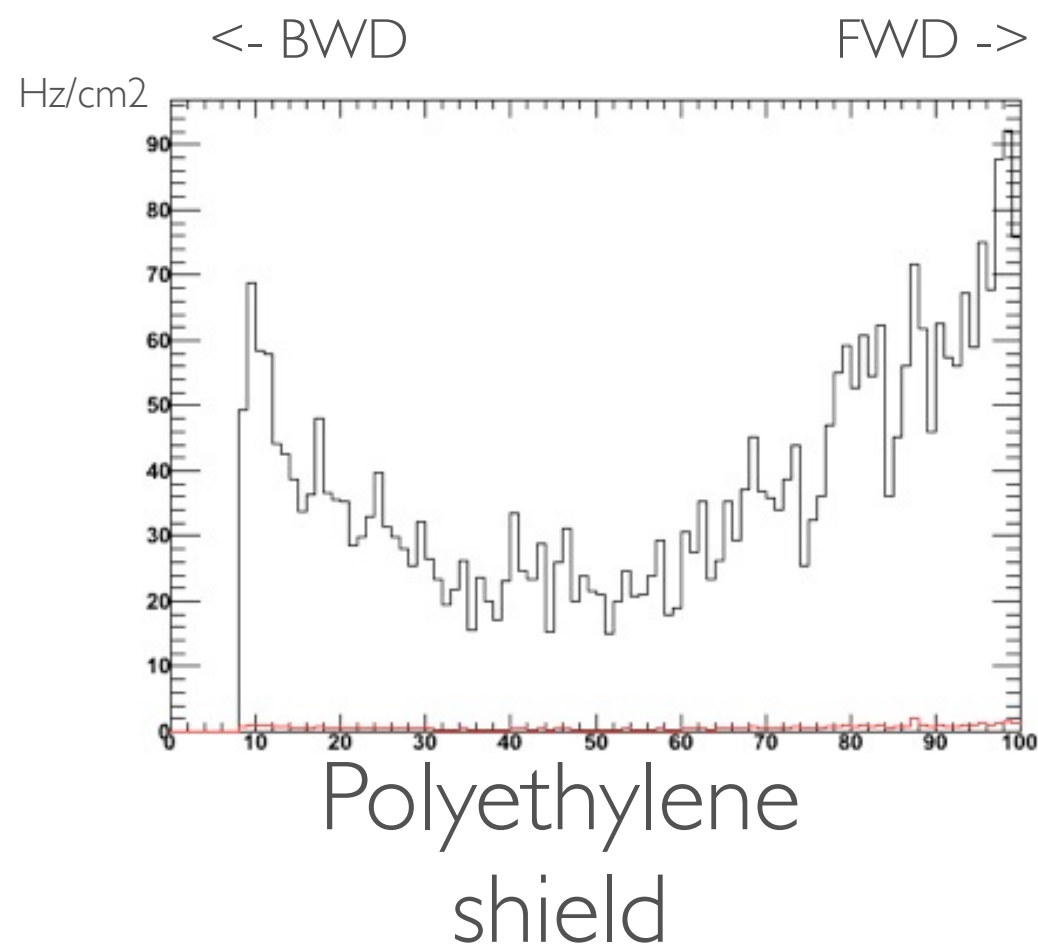
We compare the previous production in february, with a 100k events RadBhabha produced with polyethylene shield

~Factor 10 reduction as expected

IFR SHIELD VS NO SHIELD

Barrel Rate vs Z-Strips

Wolfram-Shielded
QGSP_BERT



Single Beam simulation

- **Manuela Boscolo simulated and cured the Touschek, Coulomb and brehmstrahlung single beam backgrounds**
- **Single beam backgrounds still negligible with respect to lumi terms (need to be verified with Bruno)**
- **Beam tails from her simulation in fairly good agreement with Mike Synchrotron radiation assumptions**
- **Please take your time to read her beautiful results.**

Conclusions

- **Things are proceeding smoothly on the background side of the MDI**
- **We are still learning**
- **We have to improve the bkg. rates optimizing the Shield**

- **Bring young peoples and their enthusiasm into this superb project**
-