

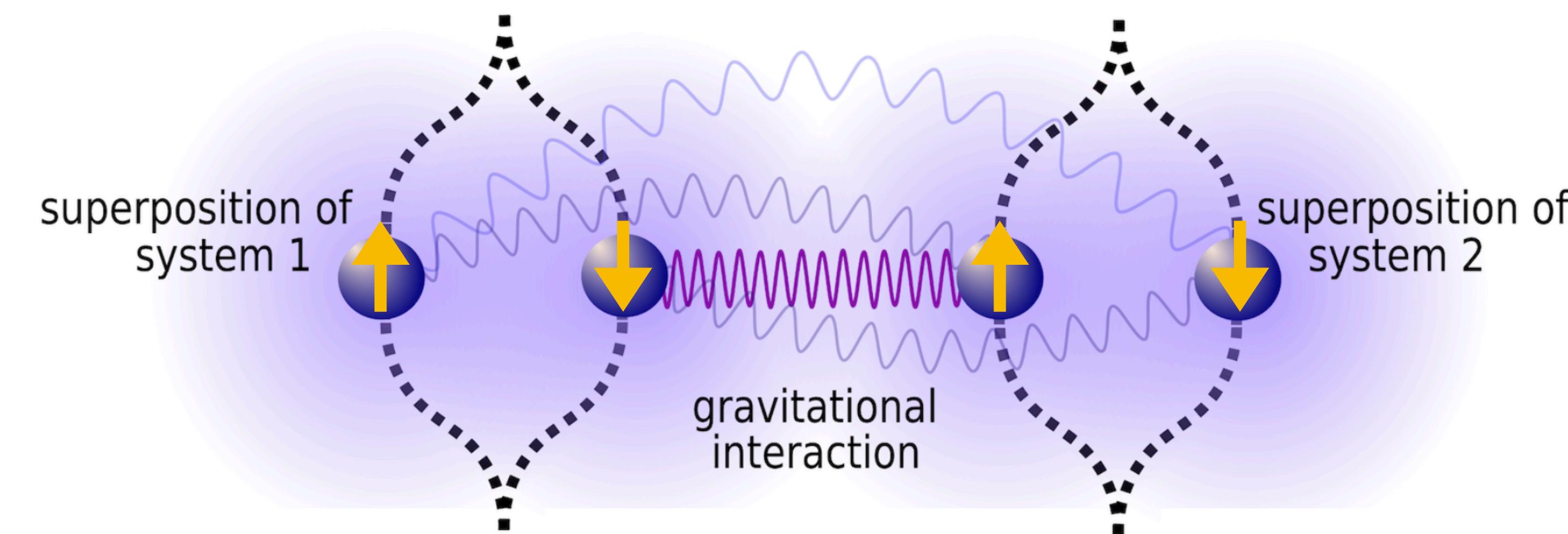
Quantum Gravity witness via Entanglement of Masses:



rijksuniversiteit
groningen

QGEM protocol

Anupam Mazumdar



Quantum
Superpositions of
Geometries

Bose + Mazumdar + Morley + Ulbricht + Toros + Paternostro + Geraci + Barker + Kim + Milburn, Phy. Rev. Lett. [ArXiv: 1707.06050]

Marshman +Mazumdar+Bose, [ArXiv: 1907.01568] Phys. Rev. A.

Van de Kamp+Marshman +Bose+Mazumdar [[2006.06931](#) [quant-ph]]

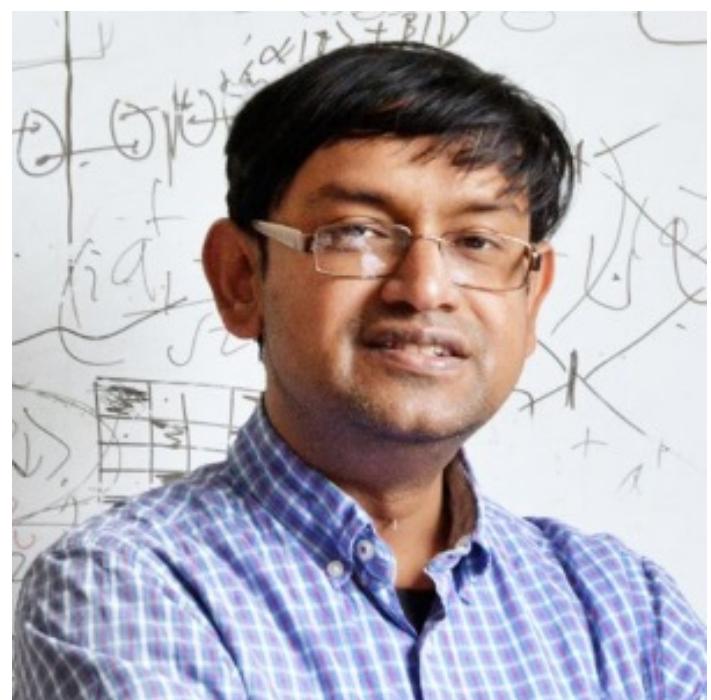
Toros + Van de Kamp + Marshman + Kim + Mazumdar + Bose [[arXiv:2007.15029](#) [gr-qc]]

Toros + Mazumdar + Bose [2008.08609](#) [gr-qc]

See Also: Marletto and Vedral appeared on the same day [1707.06036] (half of the story only)

Collaborators

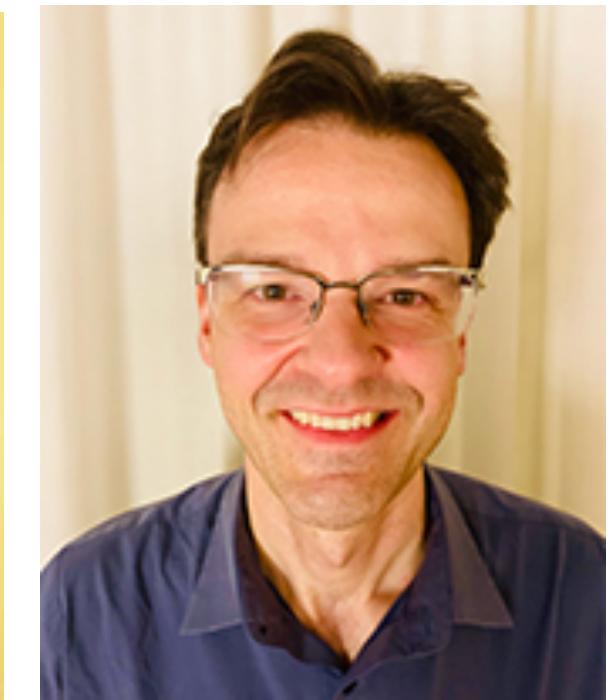
Sougato Bose Ryan Marshman



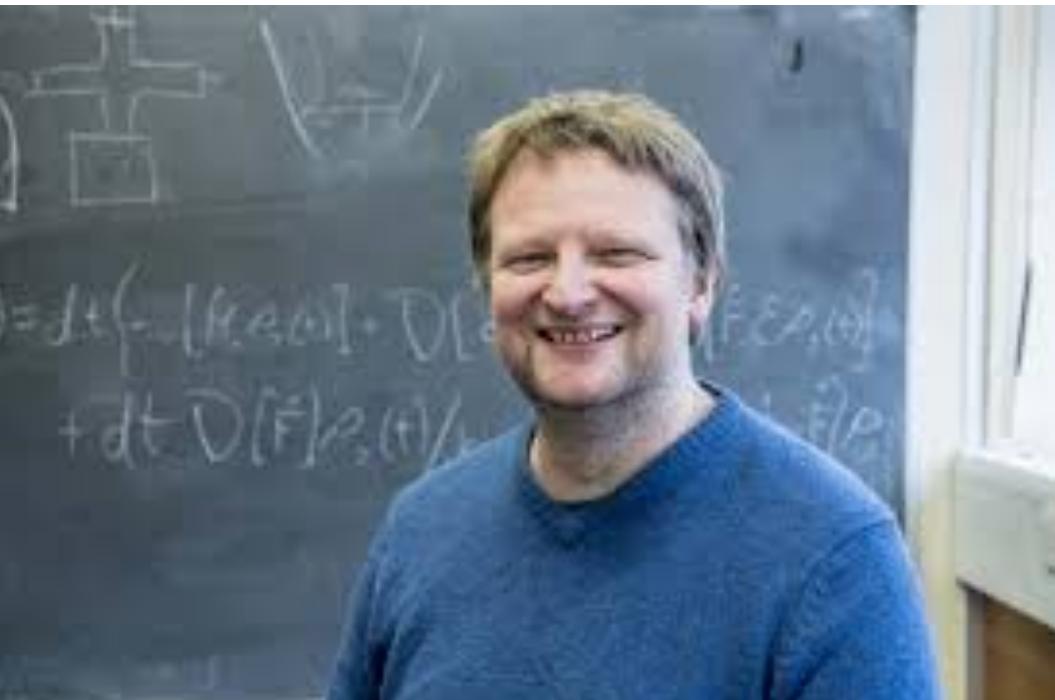
Myungshik Kim



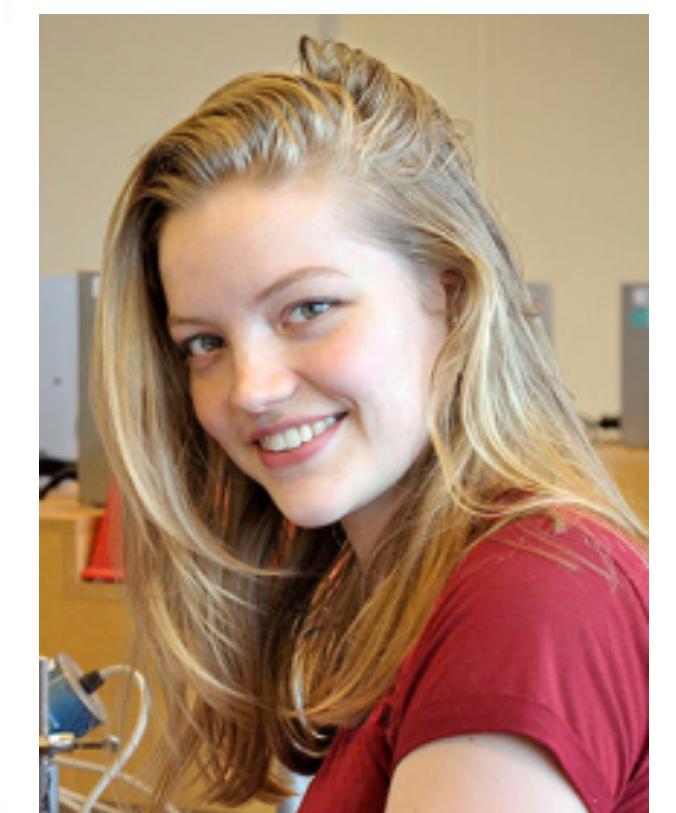
Andrew Geraci



Hendrik Ulbricht



Marko Toros



Mauro Paternostro

Gerard Milburn

Peter Barker

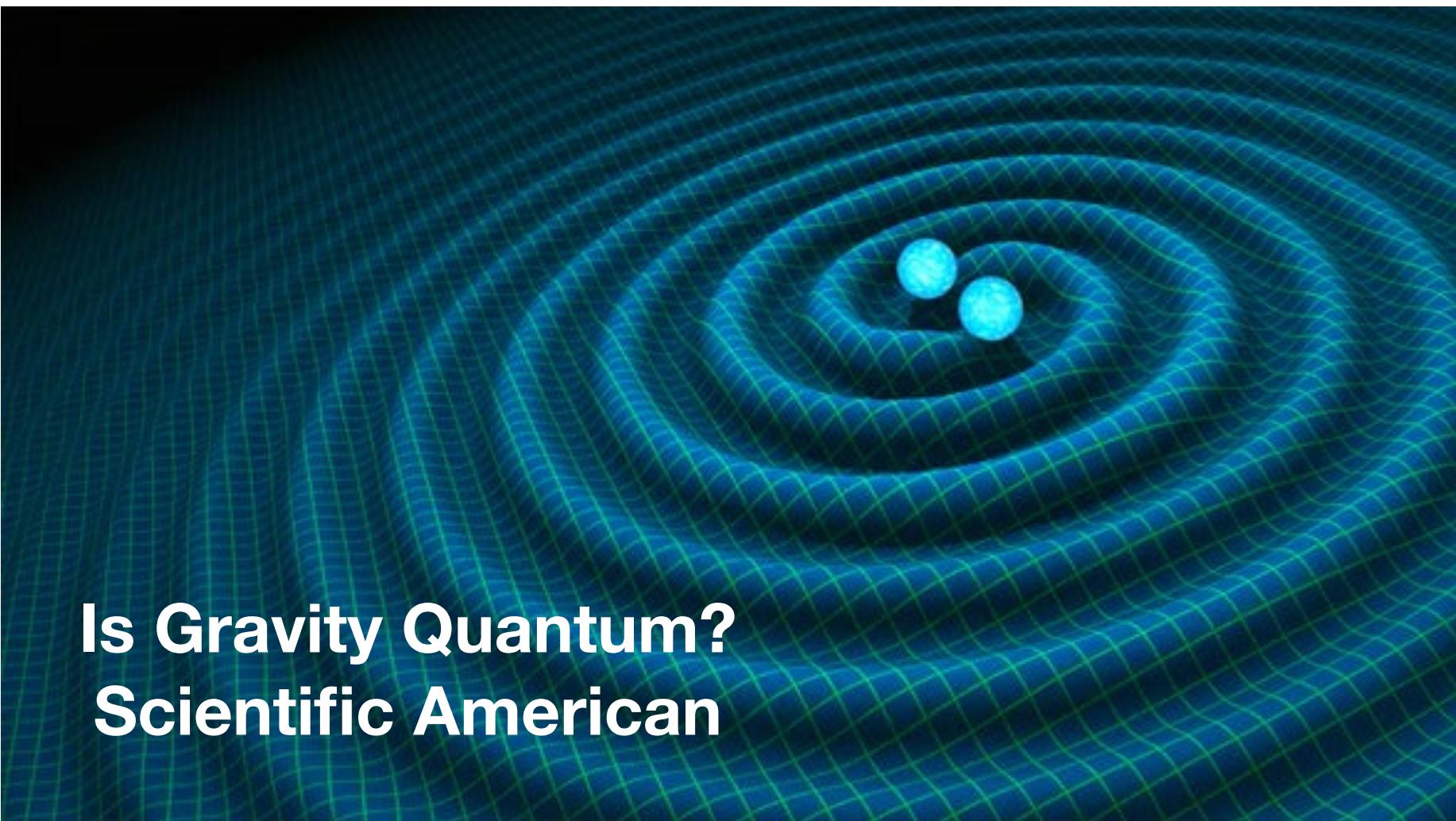
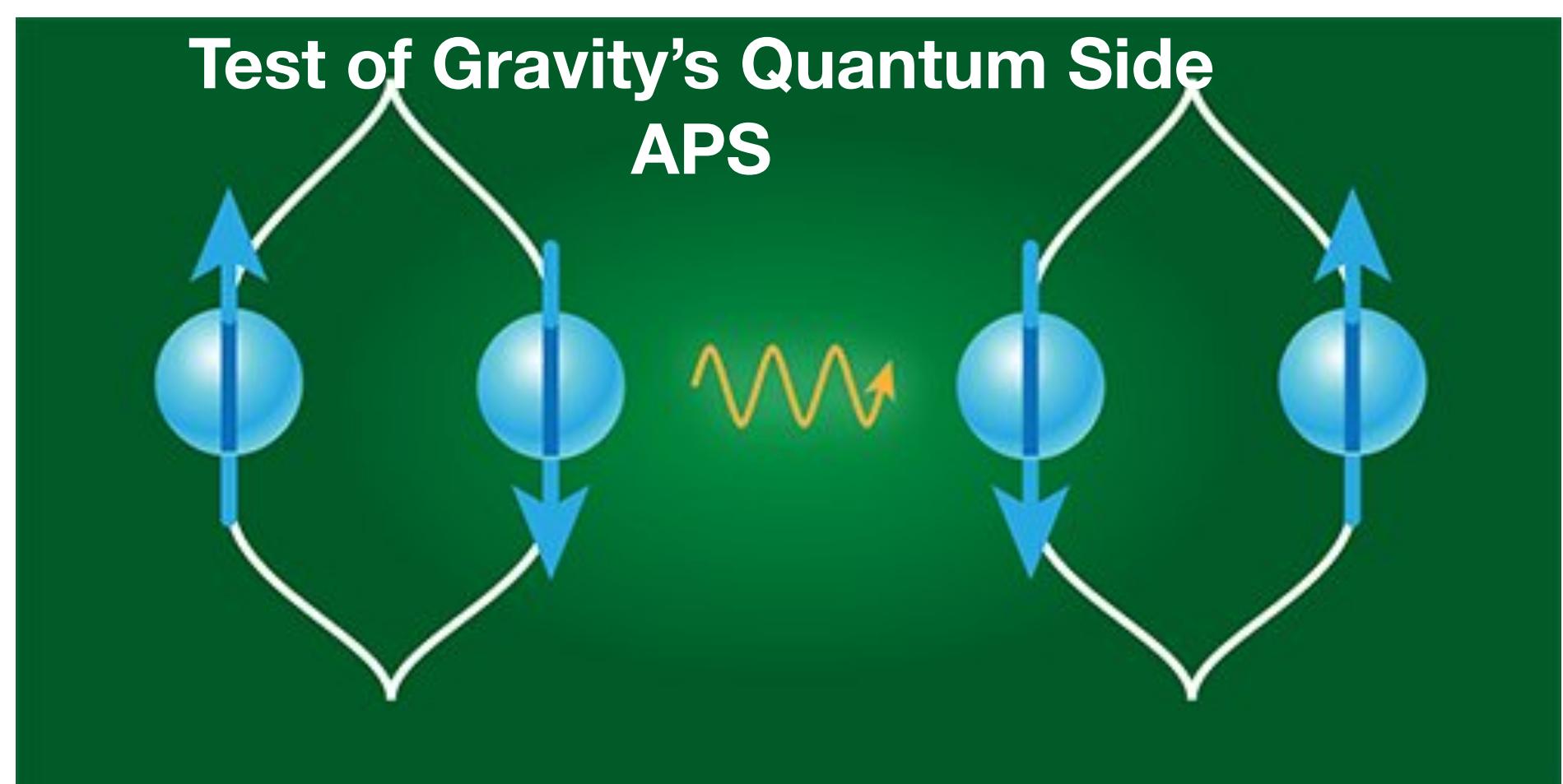
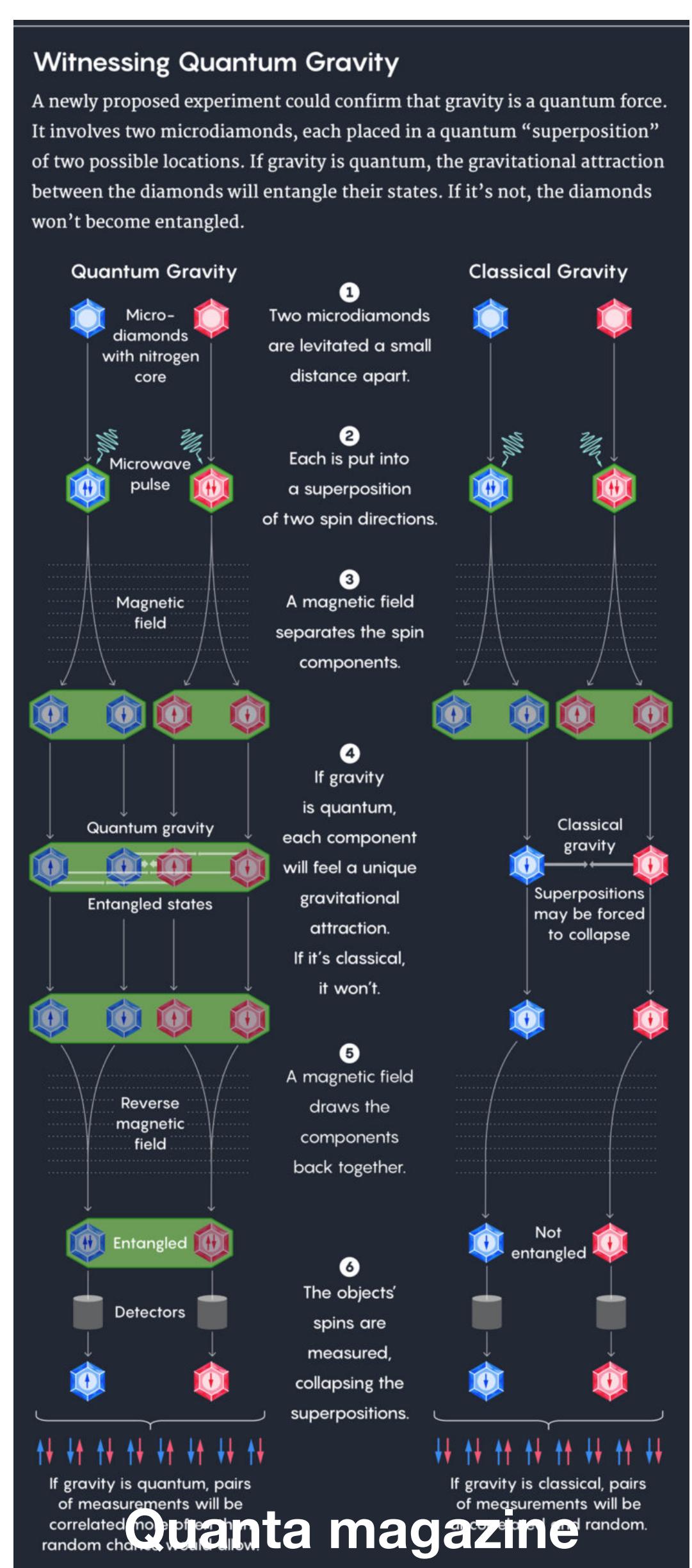
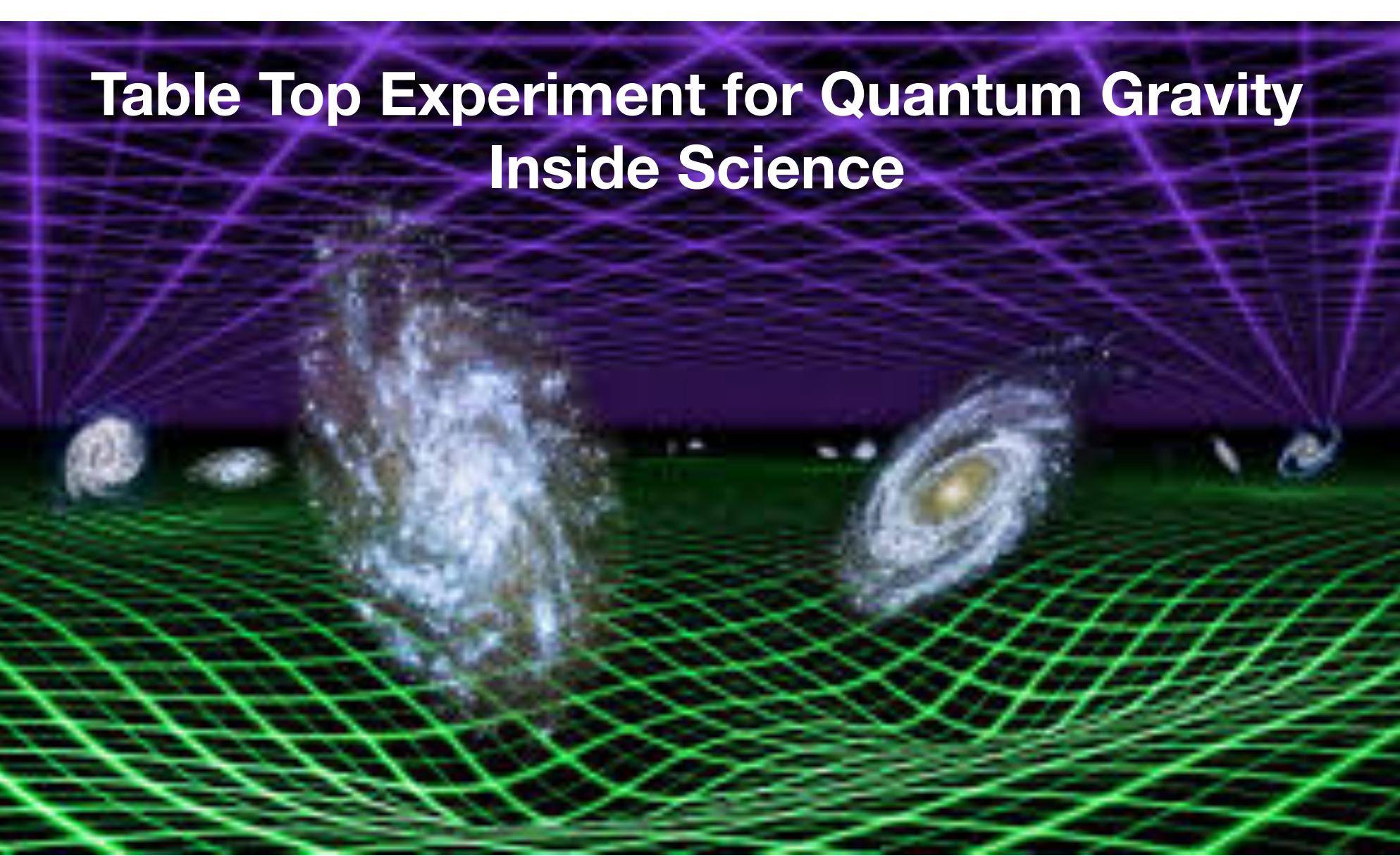
Gavin Morley

Ron Folman

Martine Schut



Anupam Mazumdar



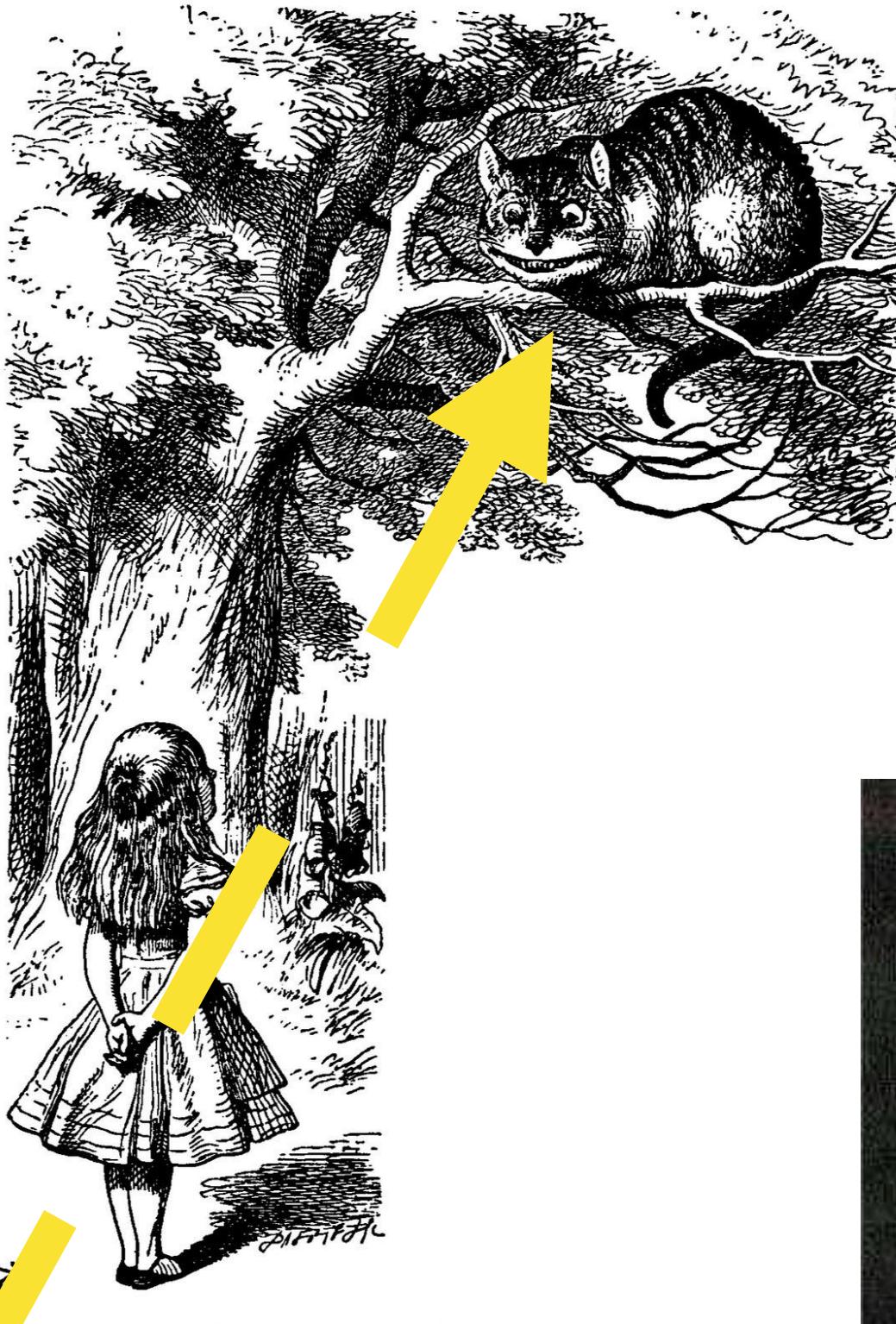
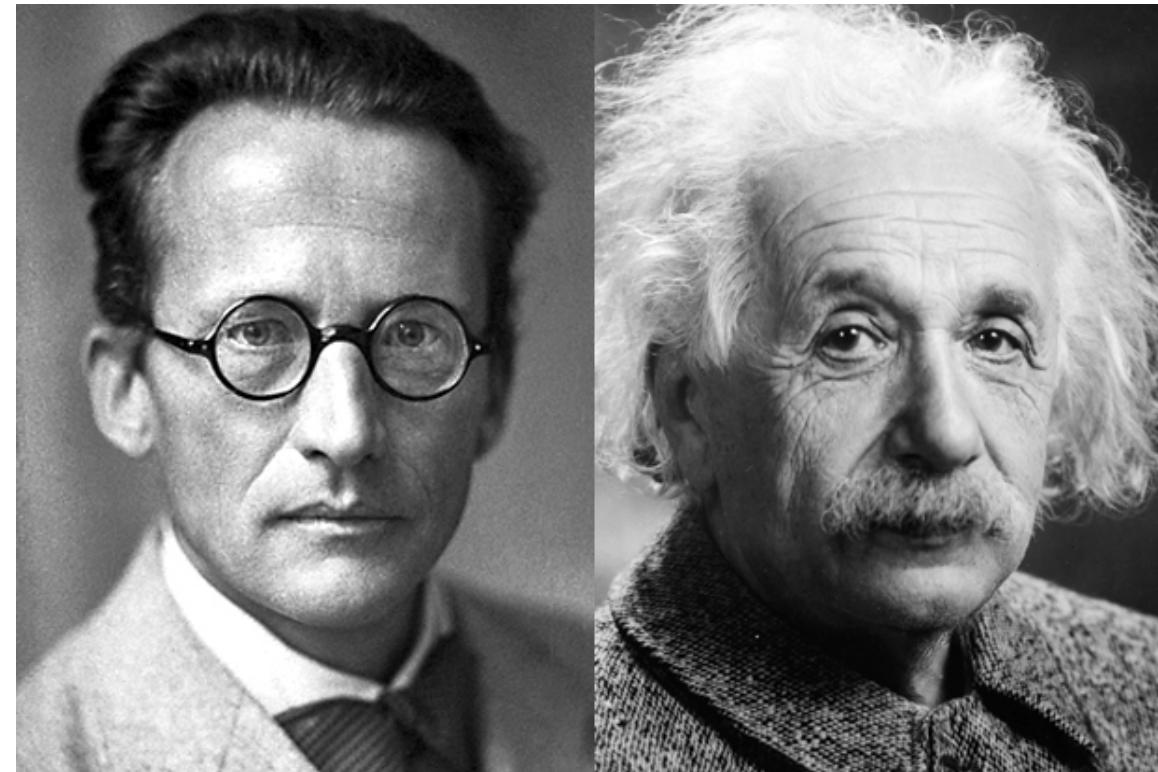
Quantum Gravity Meets Quantum Information & Quantum Technologies
Perfect Avenue for Ground Breaking Contributions

Is Gravity Classical or Quantum?



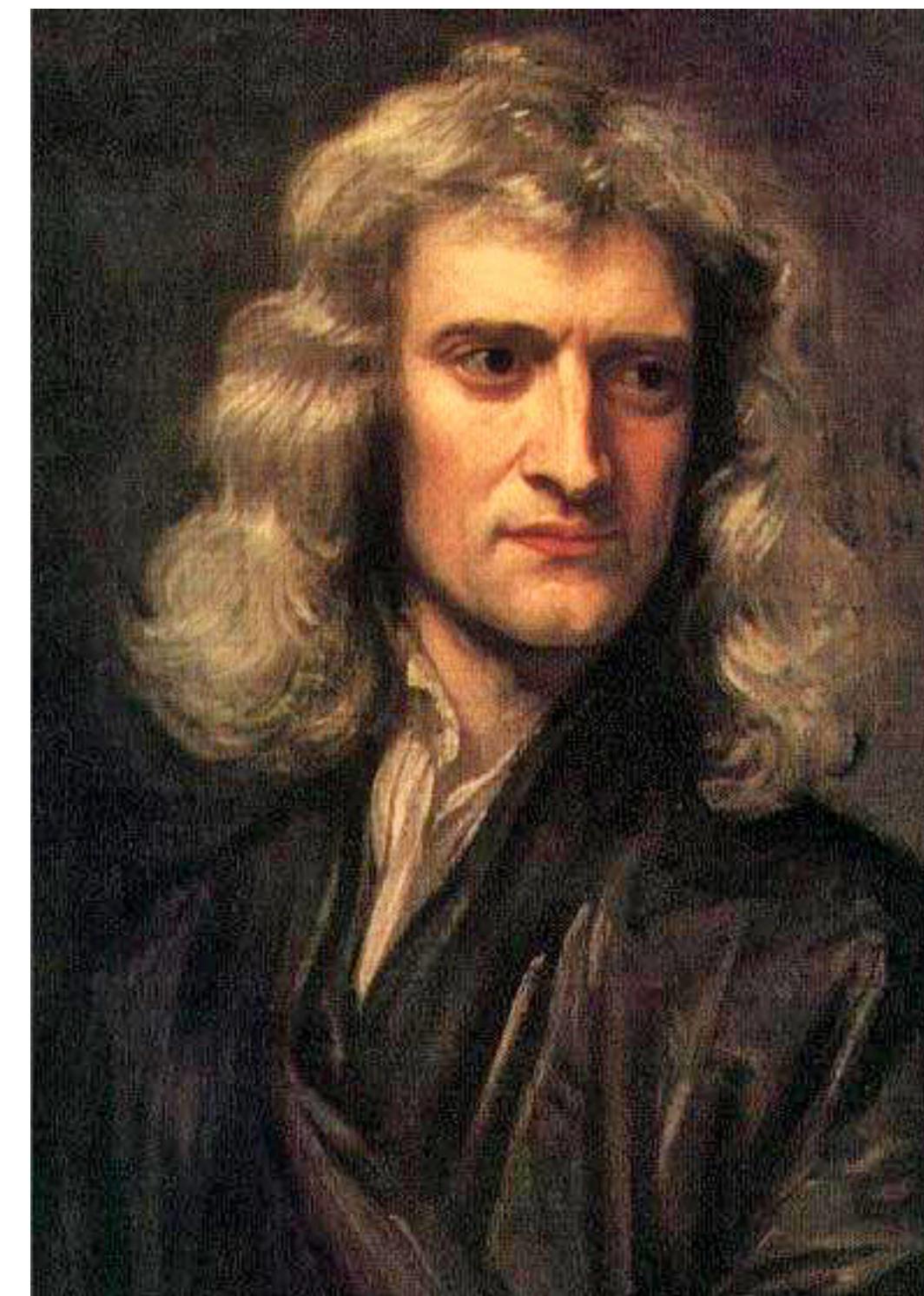
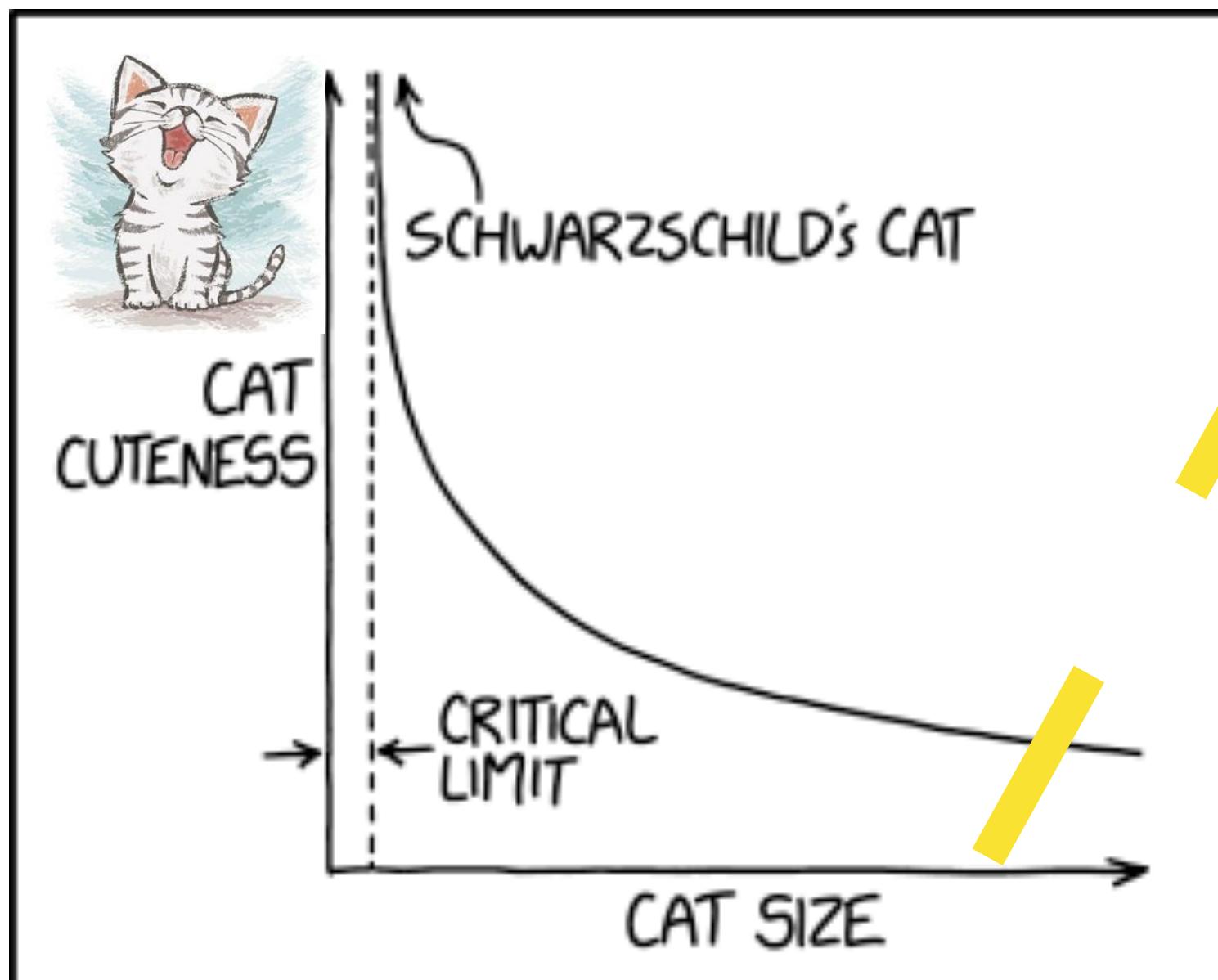
A very simple question, but extremely hard to answer !

Gravity's Cheshire Cat !



Universal Laws
of Gravitation

Microscopic
Origin of Forces



vs.



Today's Science Fiction is Tomorrow's Reality

- Can we put a graviton in a quantum superposition in a lab? **No!**
- Can we study coalescing atoms, and study the loss of gravitons ? **No!**
- Can we witness Quantum Entanglement due to Quantum Nature of a Graviton ?
Yes!

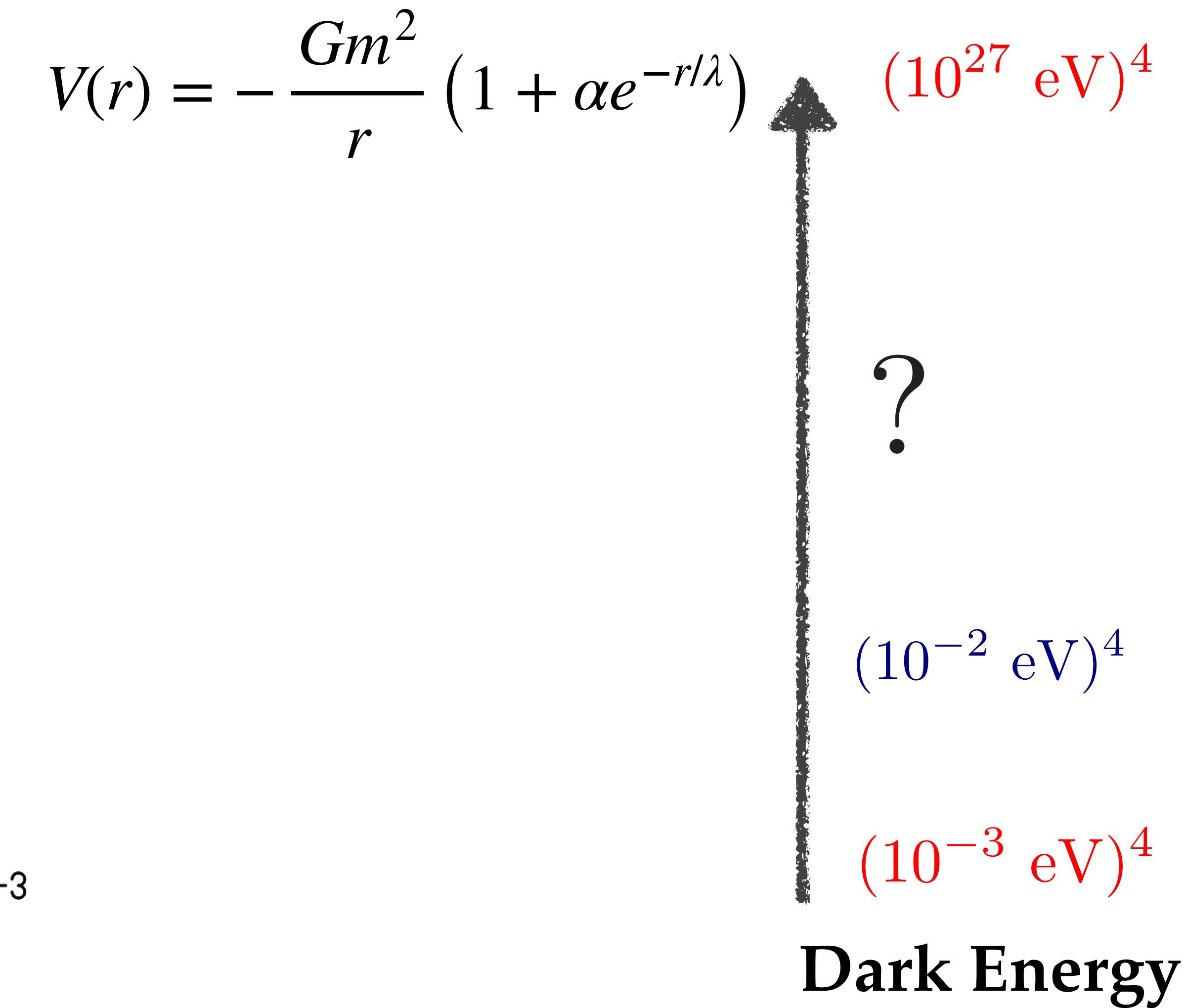
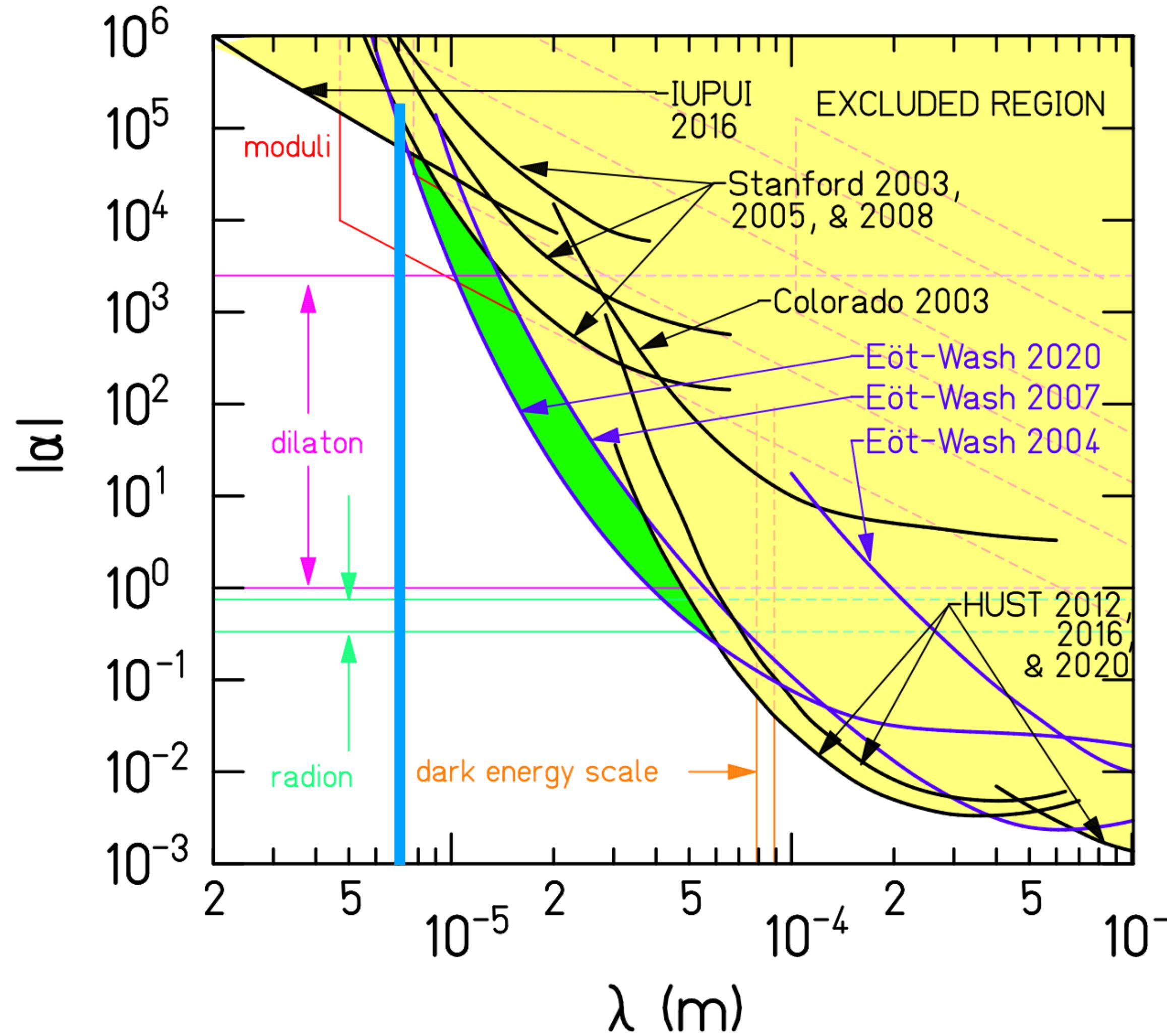
Lively discussions with

Sougato Bose, Jan Ambjorn, Abhay Ashtekar, Markus Aspelmeyers, Markus Arndt, Valeri Frolov, Gary Horowitz, Rob Myers, Don Page, Roger Penrose, Carlo Rovelli, Ashoke Sen, Gabriele Veneziano, Bob Wald, ...

Things should be made as simple as possible, but no simpler! — Albert Einstein

- 
- The background features a black cat walking on a seesaw. The seesaw is set against a vibrant, multi-colored gradient background with glowing energy spheres and particles. The cat is positioned on the right side of the seesaw, with its front paws on the higher end and its back paws on the lower end. The seesaw itself is a light blue color.
- QG-Witness protocol for Quantum Gravity
 - Reading the Witness
 - Associated Challenges & Future Program
 - Quantum Gravity

Constraints on Gravity, Newton's potential



J. G. Lee, E. G. Adelberger, T. S. Cook, S. M. Fleischer, B. R Heckel 2002.11761 [hep-ex]

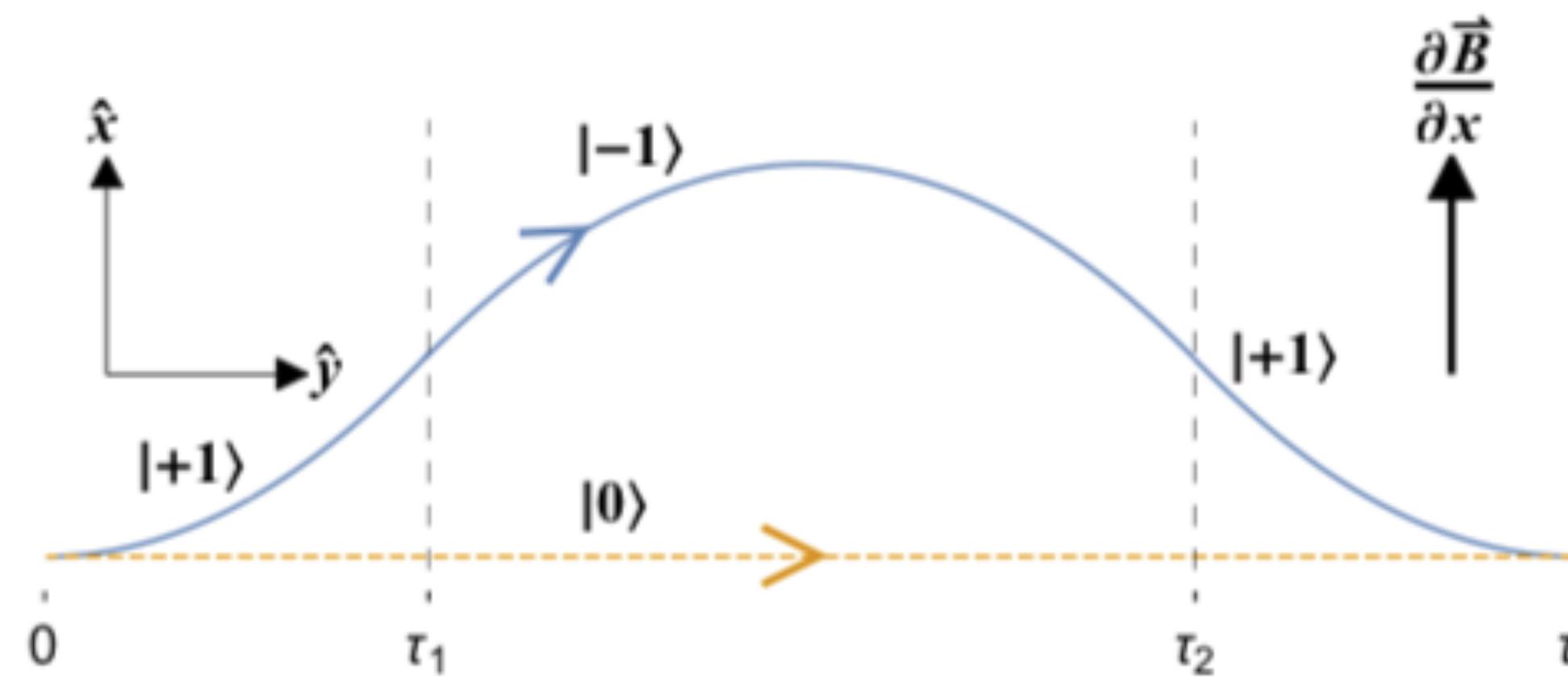
New effects in gravity could be around the corner !

Classical Gravity Induced Quantum Phase

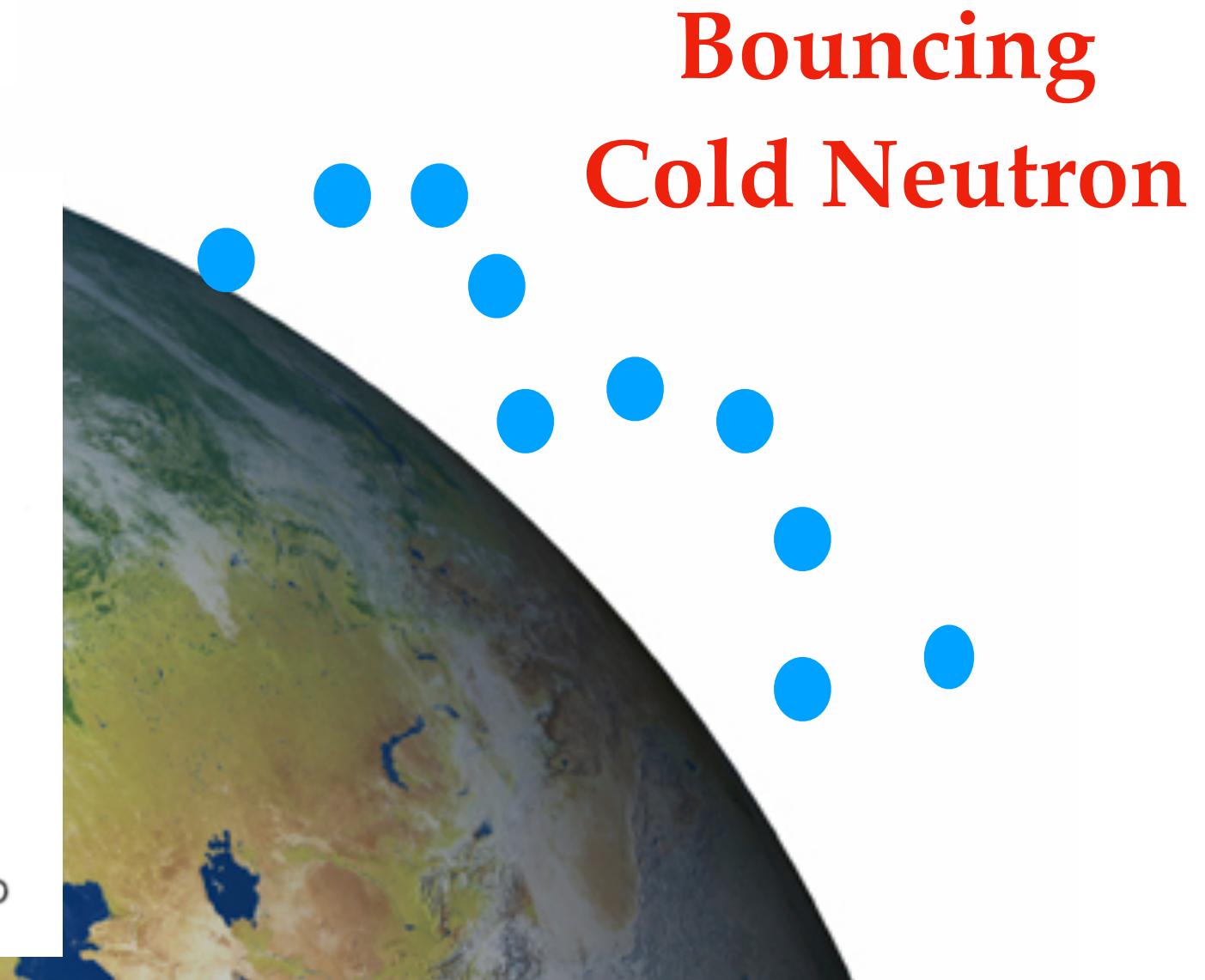
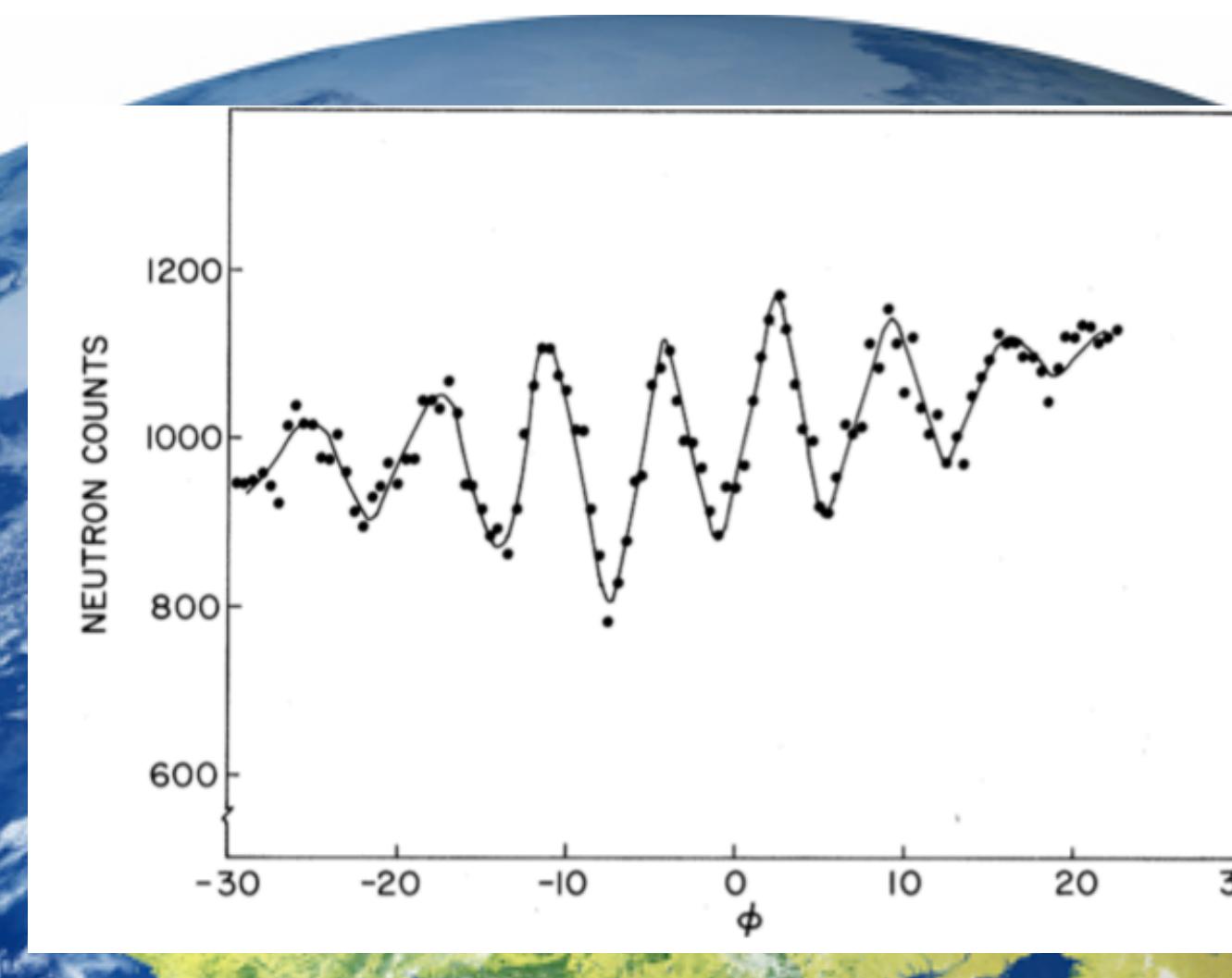
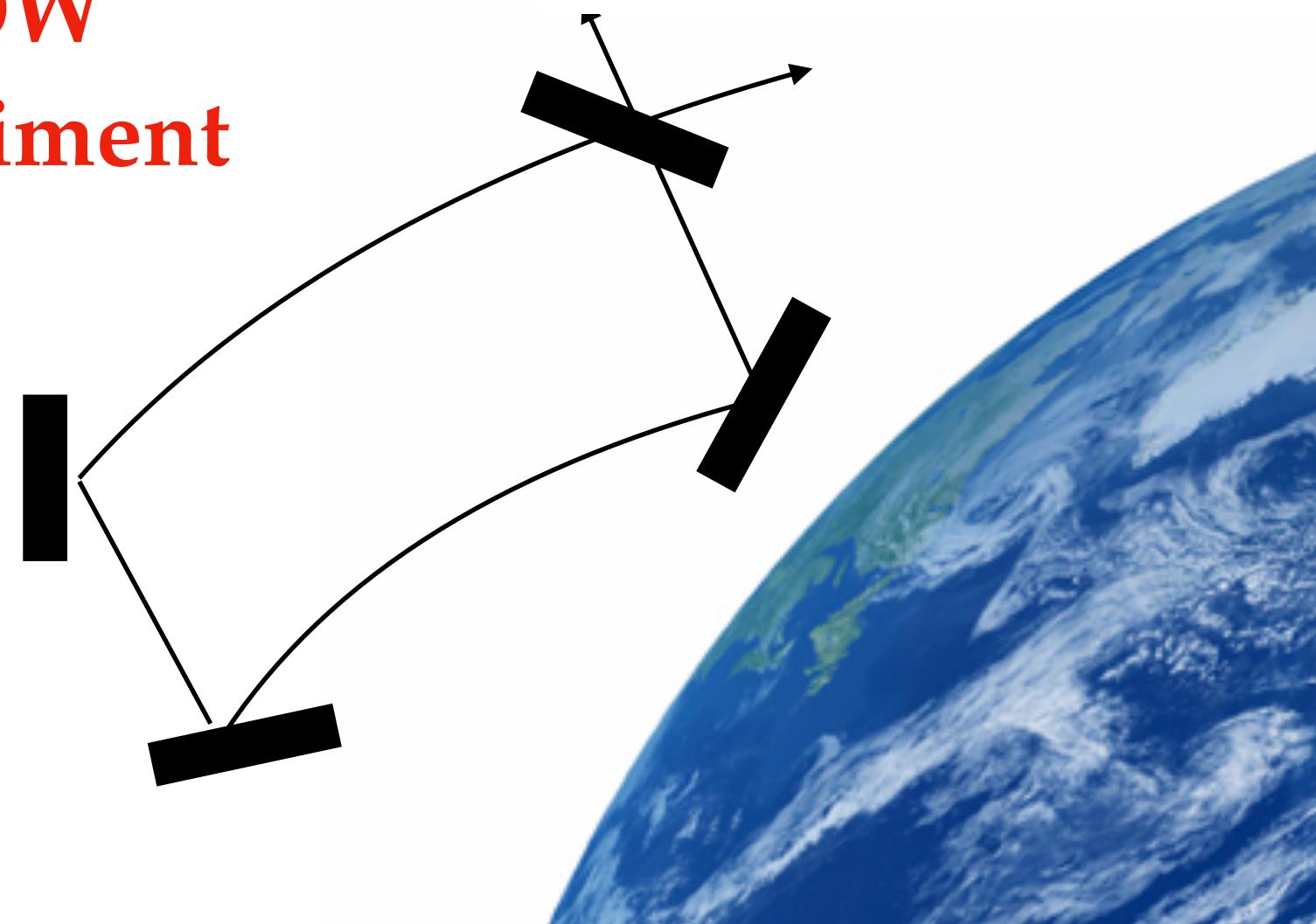
$$\Delta\phi \sim i \frac{S(G, G^2, G^2\hbar \dots)}{\hbar}$$

MIMAC. Mesoscopic Interferometer for Metric and Curvature

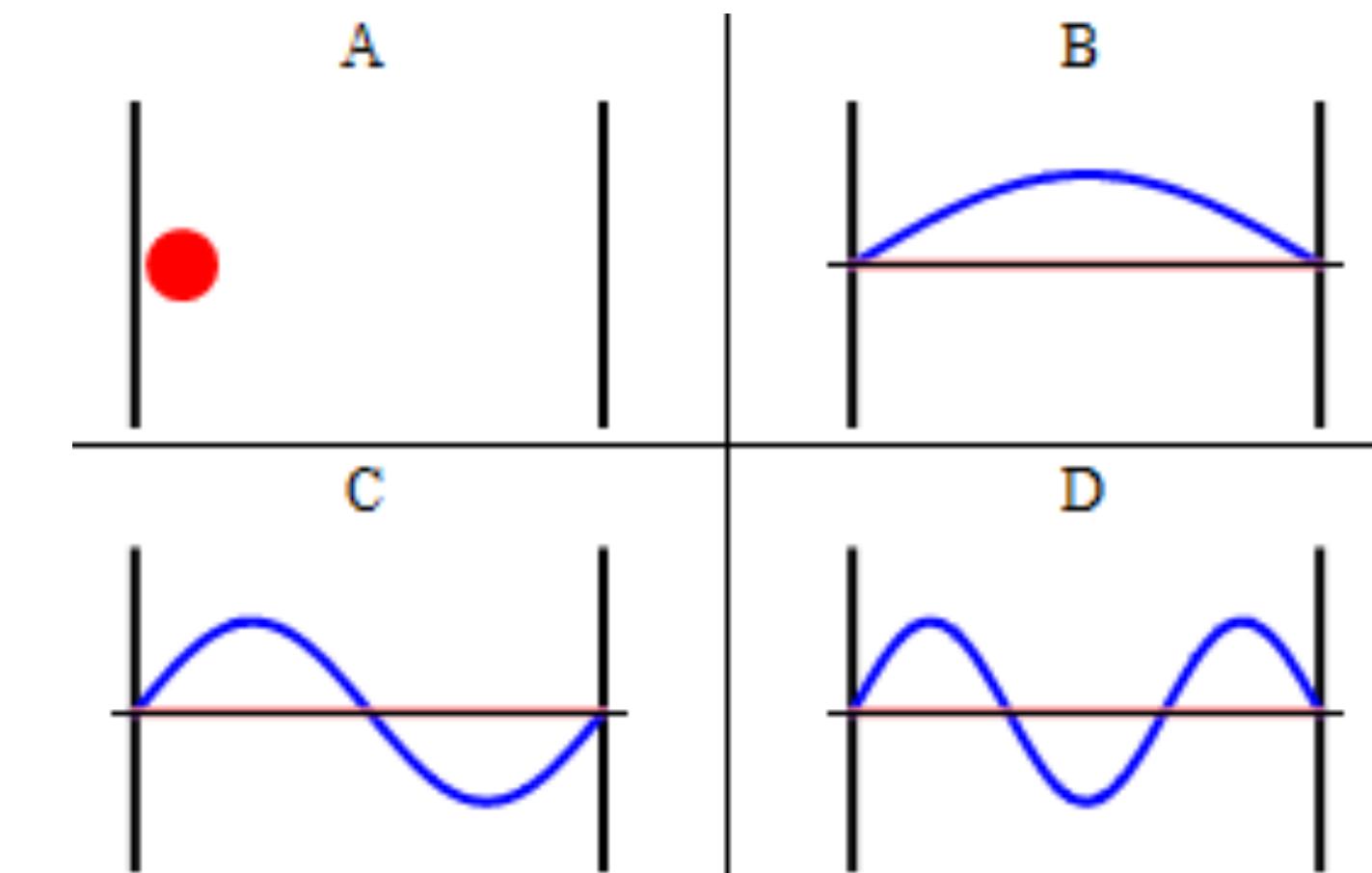
Marshman, Bose,
Morley, Barker,
Hoekstra



COW
Experiment



Bouncing
Cold Neutron



Initial Initiatives against Semi-Classical Gravity

VOLUME 47, NUMBER 14

PHYSICAL REVIEW LETTERS

5 OCTOBER 1981

Indirect Evidence for Quantum Gravity

Don N. Page

Department of Physics, The Pennsylvania State University, University Park, Pennsylvania 16802

and

C. D. Geilker

Department of Physics, William Jewell College, Liberty, Missouri 64068

(Received 9 June 1981)

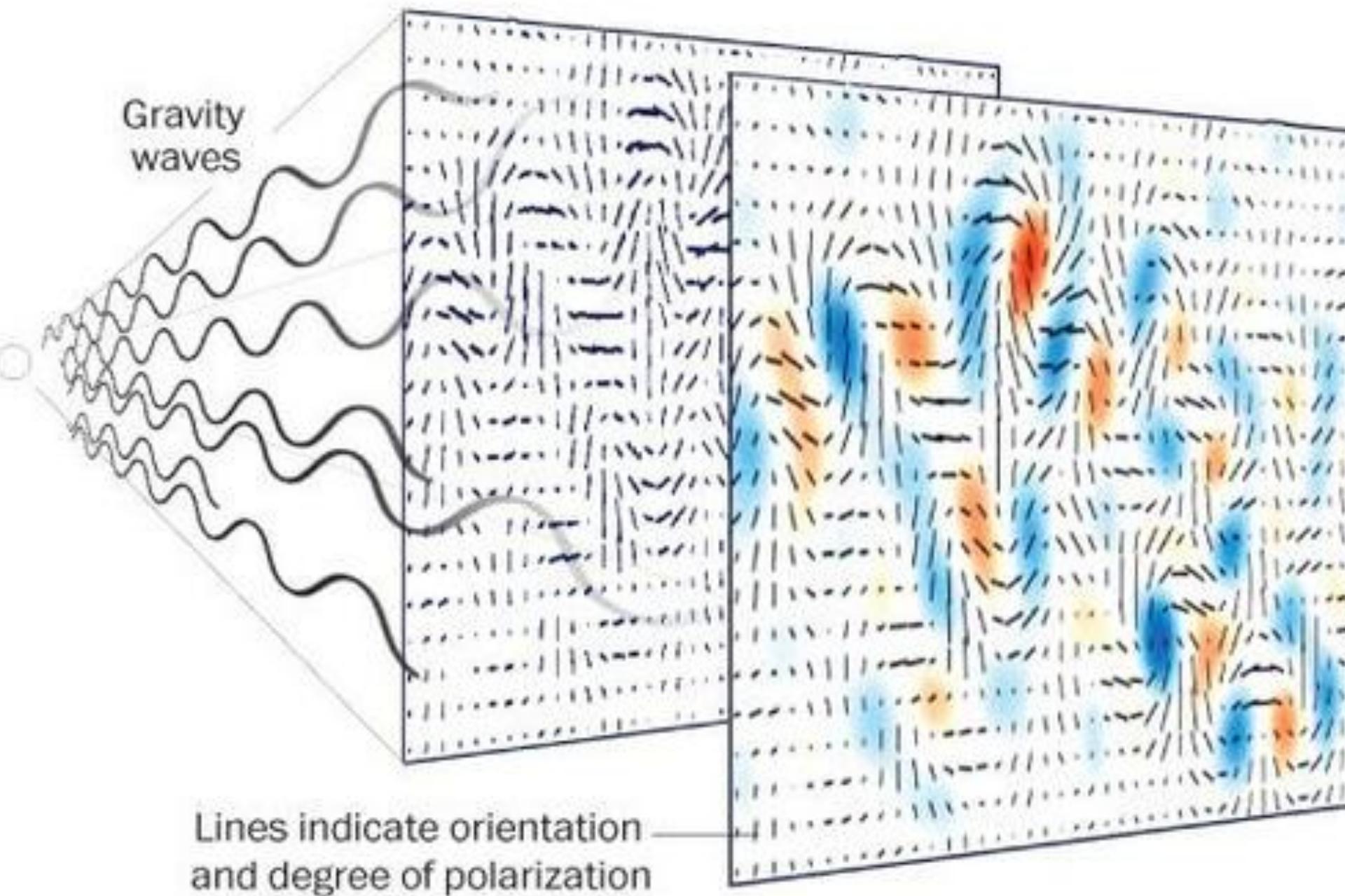
An experiment gave results inconsistent with the simplest alternative to quantum gravity, the semiclassical Einstein equations. This evidence supports (but does not prove) the hypothesis that a consistent theory of gravity coupled to quantized matter should also have the gravitational field quantized.



$$G_{\mu\nu} = \kappa^2 \langle T_{\mu\nu} \rangle$$

Primordial Gravitational Waves: is it a signature of QG?

Caution!



Positive Detection of a B-mode polarization by BICEP (2013)

But the origin was not found to be primordial in nature.

$$\mathcal{P}_g = \frac{64[|\alpha_k|^2 + |\beta_k|^2]\hbar^2 H^2}{M_p^2}$$

Bunch-Davies quantum vacuum:

$$\alpha_k = -\frac{\sqrt{\pi}}{2}, \beta_k = 0$$

Initial Conditions: Classical or Quantum?

Mere presence of \hbar is not sufficient to say that gravity is quantum !

L. M. Krauss and F. Wilczek,
"Using Cosmology to Establish the Quantization of Gravity,"
[arXiv:1309.5343 [hep-th]].

A. Ashoorioon, P. S. Bhupal Dev and AM,
"Implications of purely classical gravity for inflationary tensor modes,"
[arXiv:1211.4678 [hep-th]].

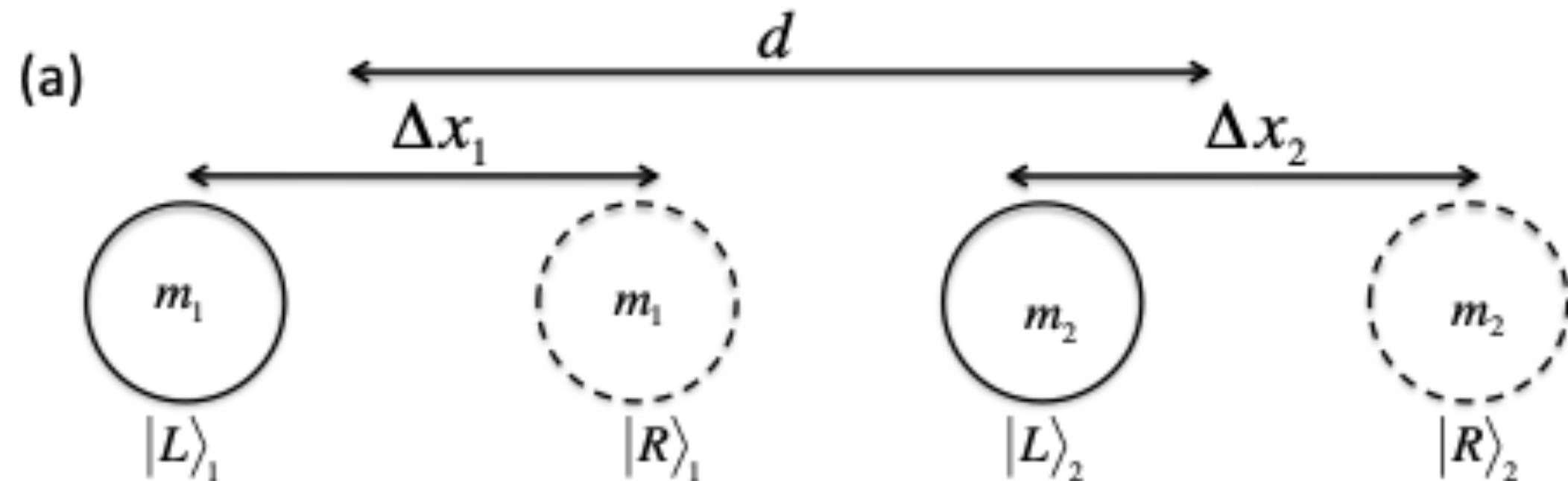
Witnessing Quantum Gravity



Razor Sharp Argument

- If the Witness is Positive then Gravity is Quantum
- If the Witness is Negative: Inconclusive !

Key Ingredients



○ Macroscopic Quantum Superposition of Localised Objects

Curvatures are Localised

10^{-21}Kg [Markus Arndt's Lab, University of Vienna]

Product State



Entangled State



A

B

A

B

$$|\Psi\rangle = |\Psi\rangle_A \otimes |\Psi\rangle_B$$

$$|\Psi\rangle \neq |\Psi\rangle_A \otimes |\Psi\rangle_B$$

Pure

Trace

Mixed

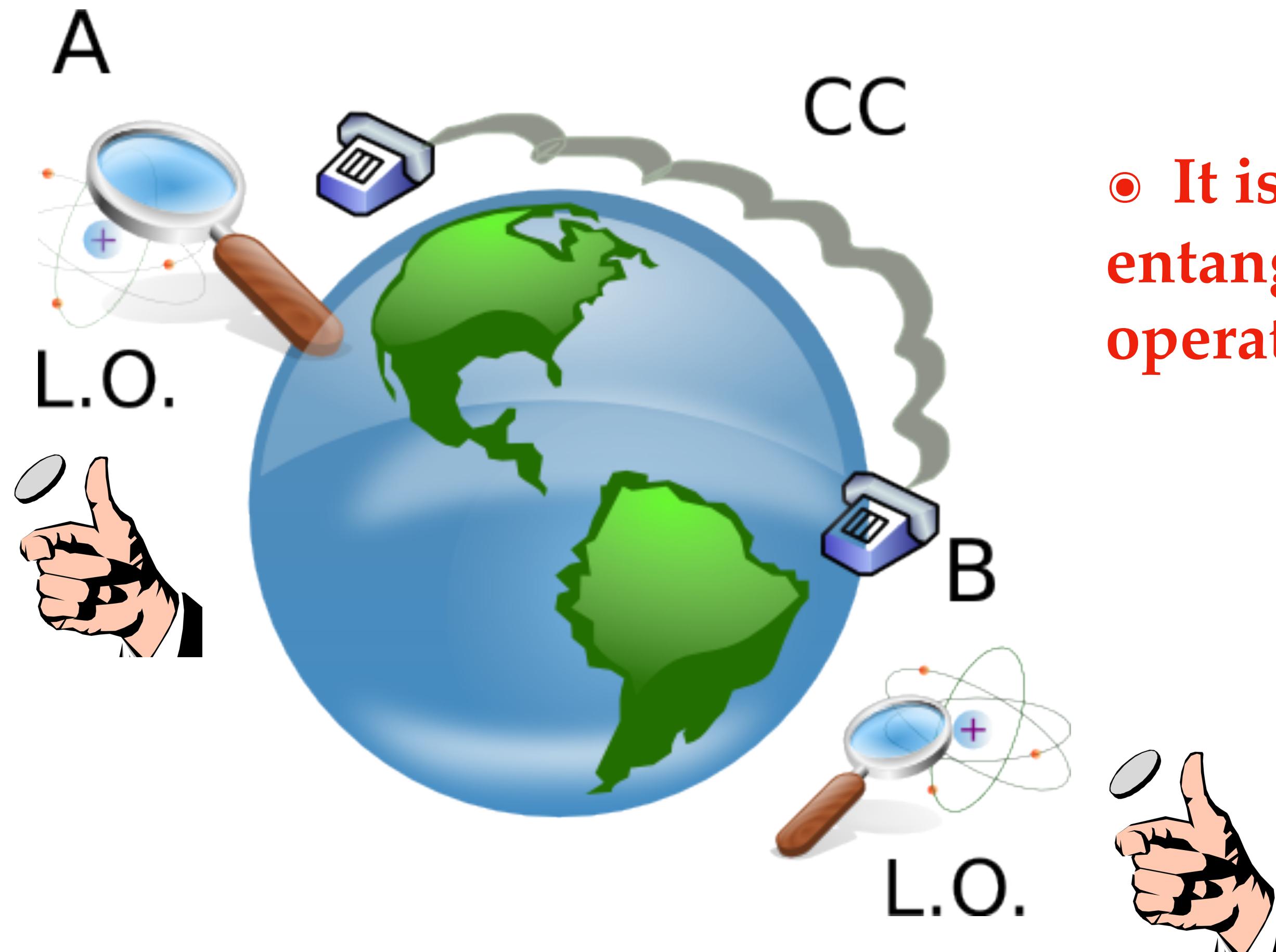
Trace

○ Quantum Entanglement

$$e^+ e^- + e^- e^+$$



Local Operations & Classical Communication (LOCC)



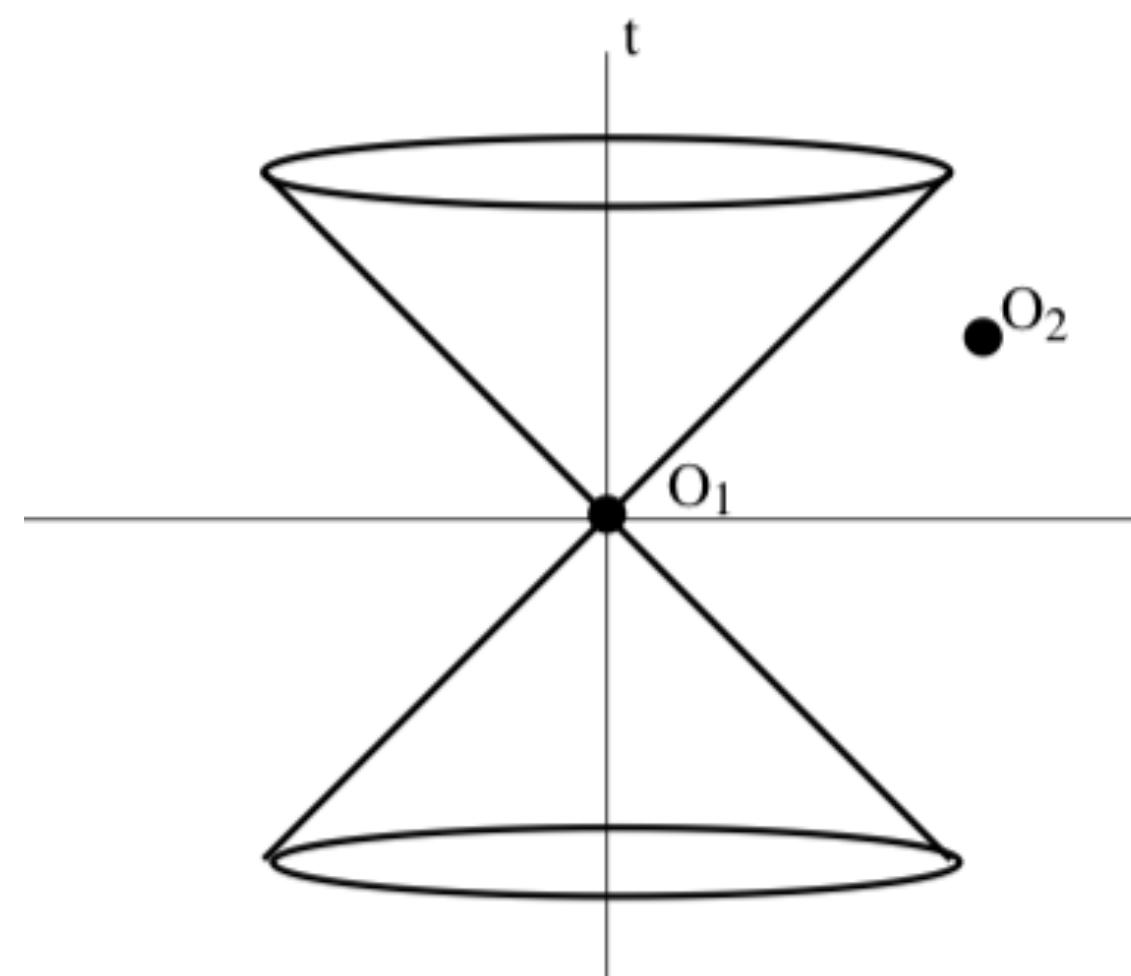
- It is impossible to generate/increase entanglement between A and B by local operations and classical communications

Bennett, et.al, (1996)

LOCC keeps Separable state remains Separable (Cannot Create Entanglement)

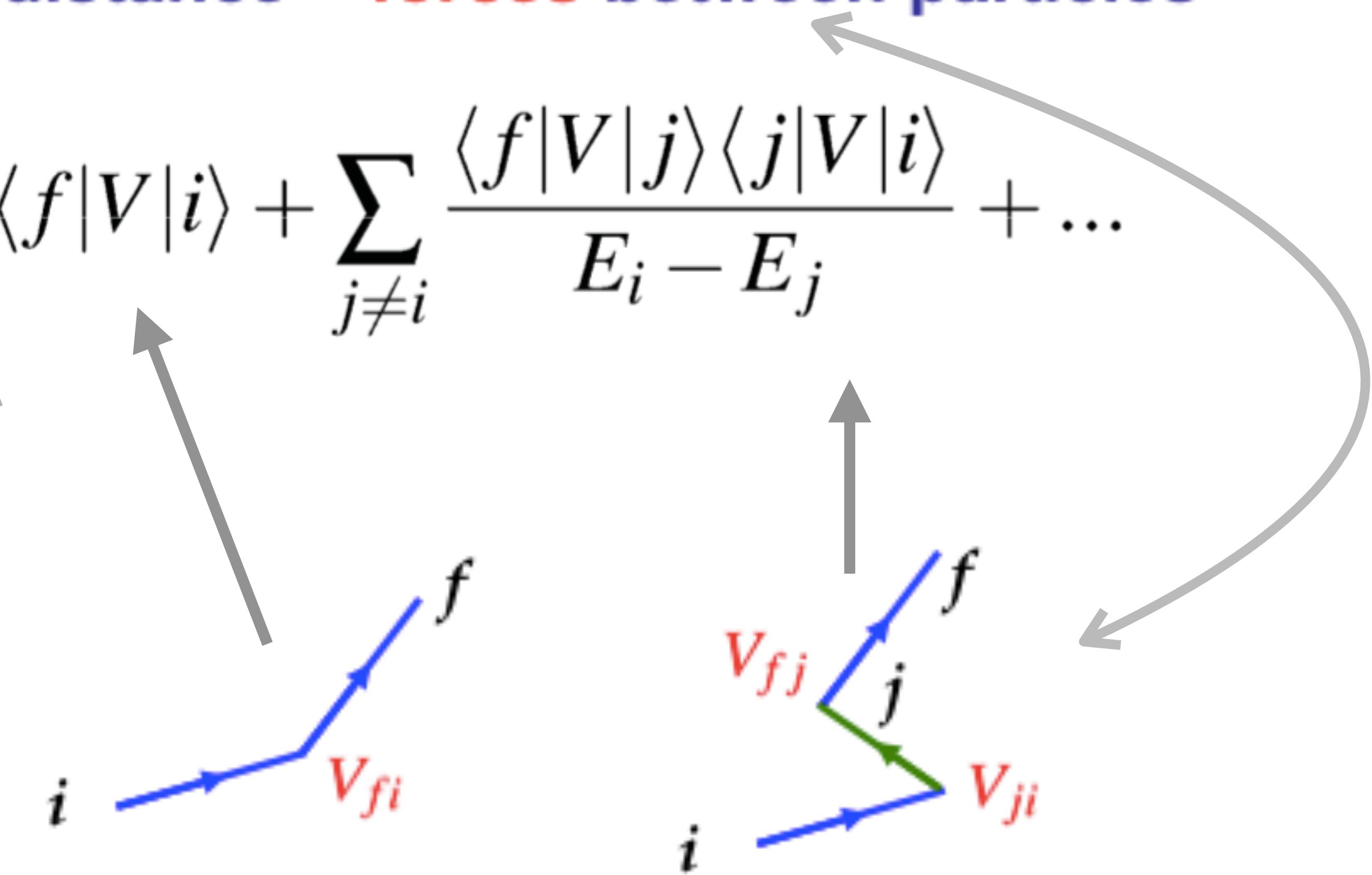
Quantum origin of $V(r)$

- “Classical picture” – particles act as sources for fields which give rise a potential in which other particles scatter – “action at a distance”
- “Quantum Field Theory picture” – forces arise due to the exchange of virtual particles. No action at a distance + forces between particles now due to particles



$$T_{fi} = \langle f | V | i \rangle + \sum_{j \neq i} \frac{\langle f | V | j \rangle \langle j | V | i \rangle}{E_i - E_j} + \dots$$

Violates Special Relativity

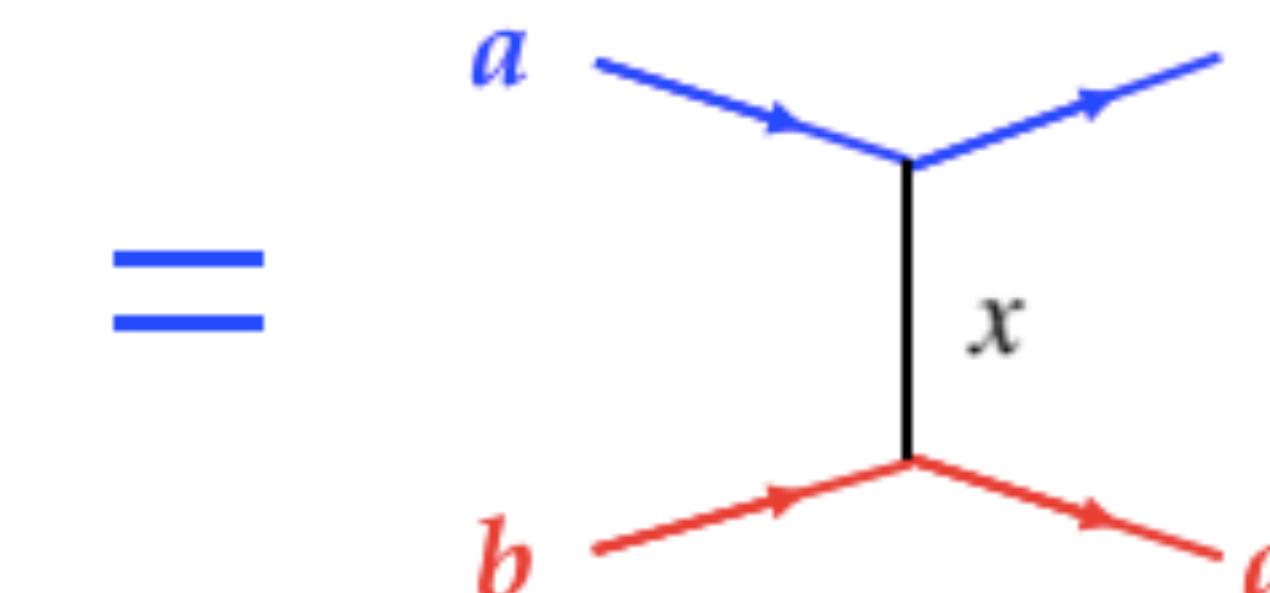
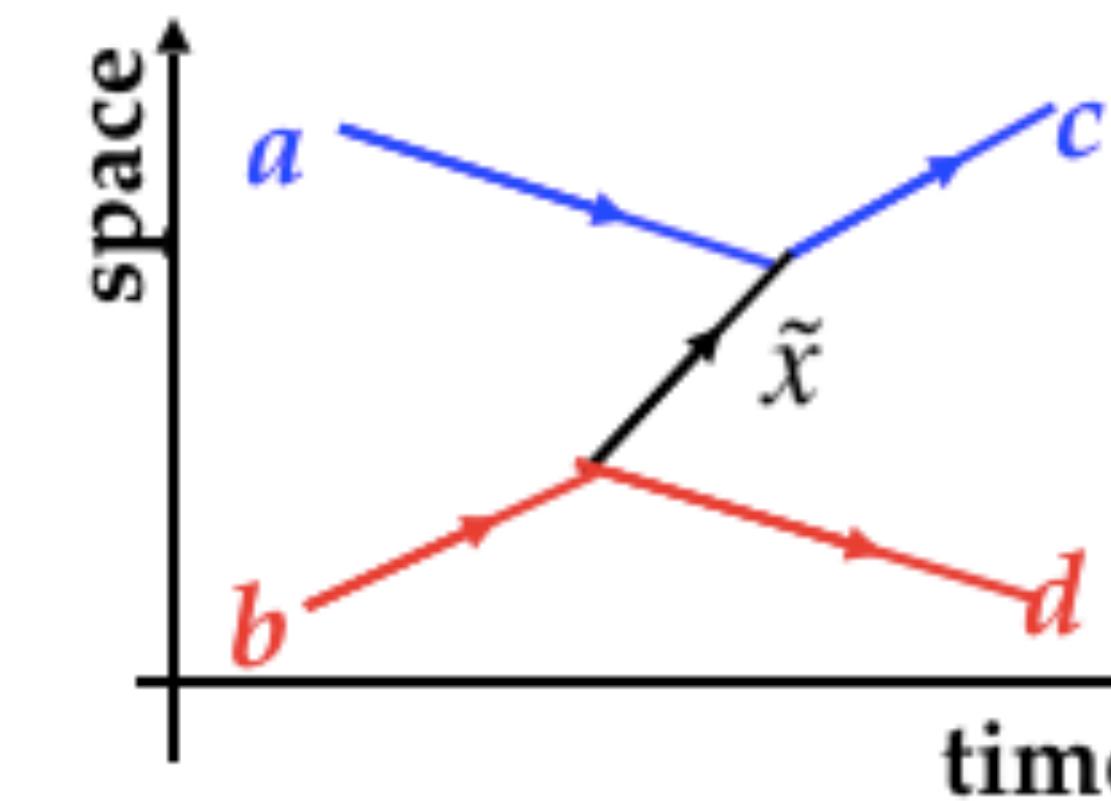
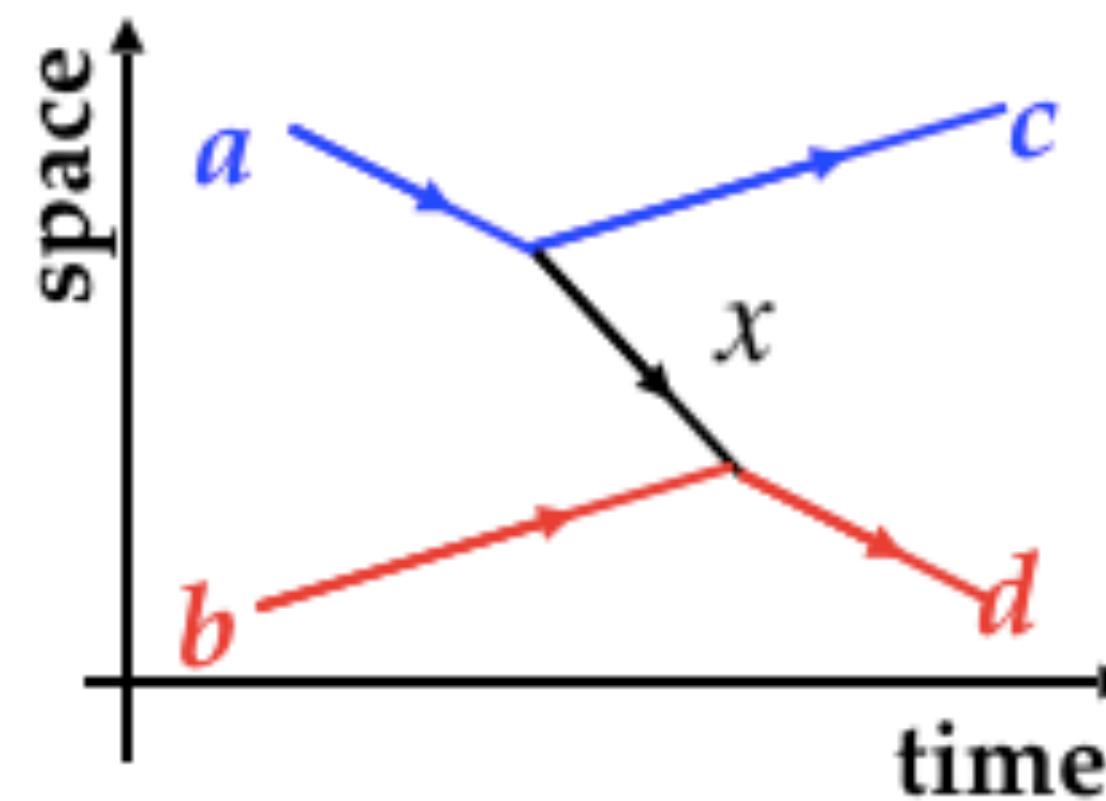


There are papers by Bei-lok Hu and Anastopoulos, who believe that tree level scattering is classical & not quantum!

These papers do not have the concept of an off-shell graviton mediator!

Quantum Mechanics to Quantum Field Theory

- The sum over all possible time-orderings is represented by a **FEYNMAN diagram**



$$M_{fi} = \frac{g_a g_b}{q^2 - m_x^2}$$

- Momentum conserved at vertices
- Energy **not** conserved at vertices
- Exchanged particle "**on mass shell**"

$$E_x^2 - |\vec{p}_x|^2 = m_x^2$$

- Momentum **AND** energy conserved at interaction vertices
- Exchanged particle "**off mass shell**"

$$E_x^2 - |\vec{p}_x|^2 = q^2 \neq m_x^2$$

VIRTUAL PARTICLE

$$\Delta x \Delta p_x \geq \hbar$$

- On-shell (Follows Classical Equations of Motion): $E^2 = (pc)^2 + (mc^2)^2$
- Off-shell does not follow Classical Equations of Motion

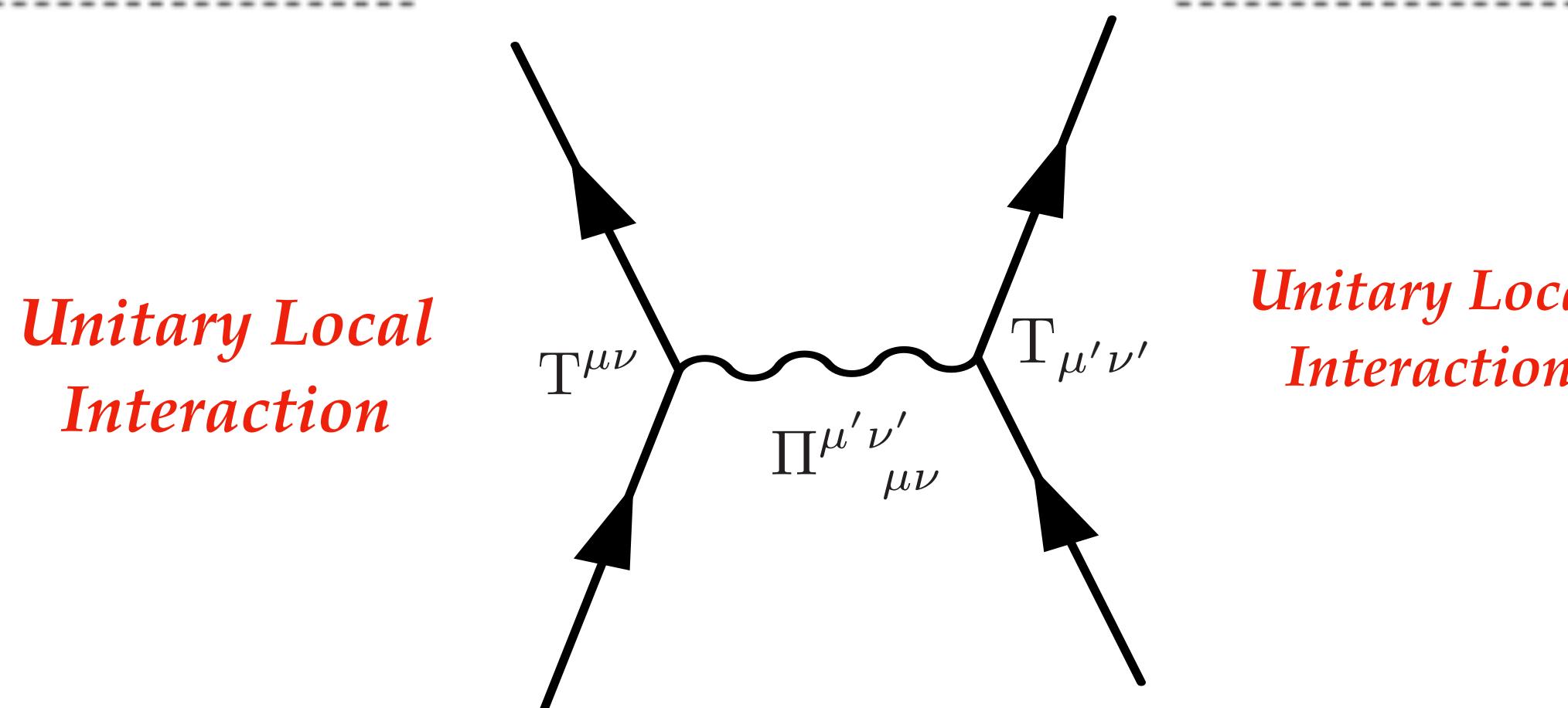
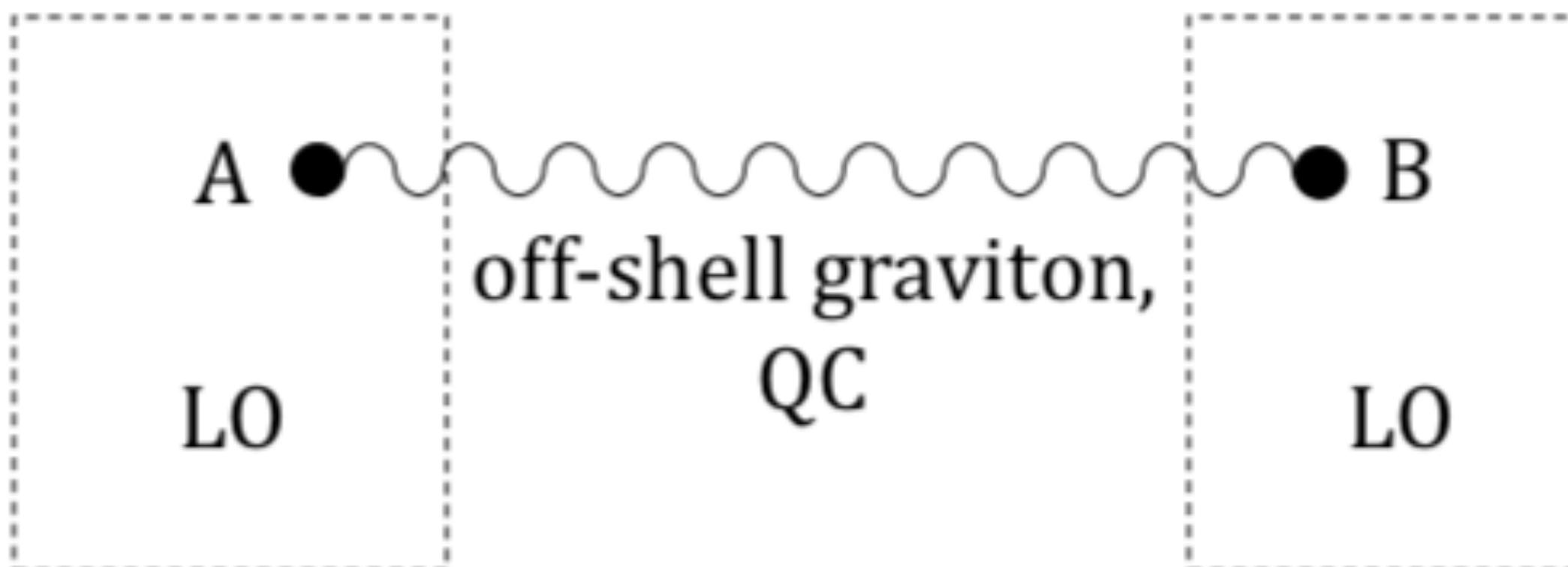
LOQC: Local operation & Quantum Communication

$$g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu} \quad S_{EH} = \frac{1}{4} \int d^4x h_{\mu\nu} \mathcal{O}^{\mu\nu\rho\sigma} h_{\rho\sigma} + \mathcal{O}(\kappa h^3)$$

(1) Non-Classicality

Minimal Coupling : $\kappa h_{\mu\nu} T^{\mu\nu}$

(2) Non-Commuting Operators

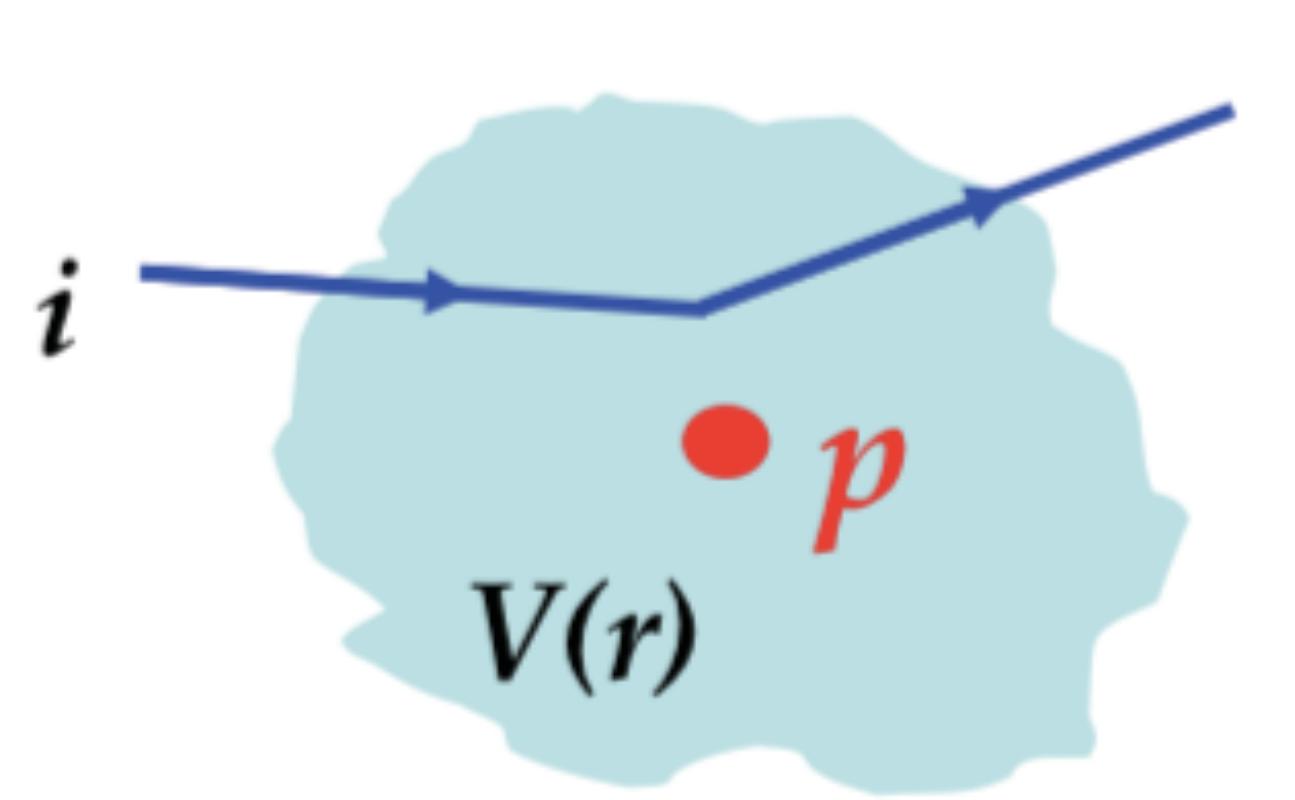


$$\begin{aligned} \mathcal{O}^{\mu\nu\rho\sigma} := & \frac{1}{4} (\eta^{\mu\rho}\eta^{\nu\sigma} + \eta^{\mu\sigma}\eta^{\nu\rho}) \square - \frac{1}{2} \eta^{\mu\nu}\eta^{\rho\sigma} \square \\ & + \frac{1}{2} (\eta^{\mu\nu}\partial^\rho\partial^\sigma + \eta^{\rho\sigma}\partial^\mu\partial^\nu - \eta^{\mu\rho}\partial^\nu\partial^\sigma - \eta^{\mu\sigma}\partial^\nu\partial^\rho) \end{aligned}$$

$$\Pi_{\mu\nu\rho\sigma}(k) = \left(\frac{\mathcal{P}_{\mu\nu\rho\sigma}^2}{k^2} - \frac{\mathcal{P}_{s, \mu\nu\rho\sigma}^0}{2k^2} \right)$$

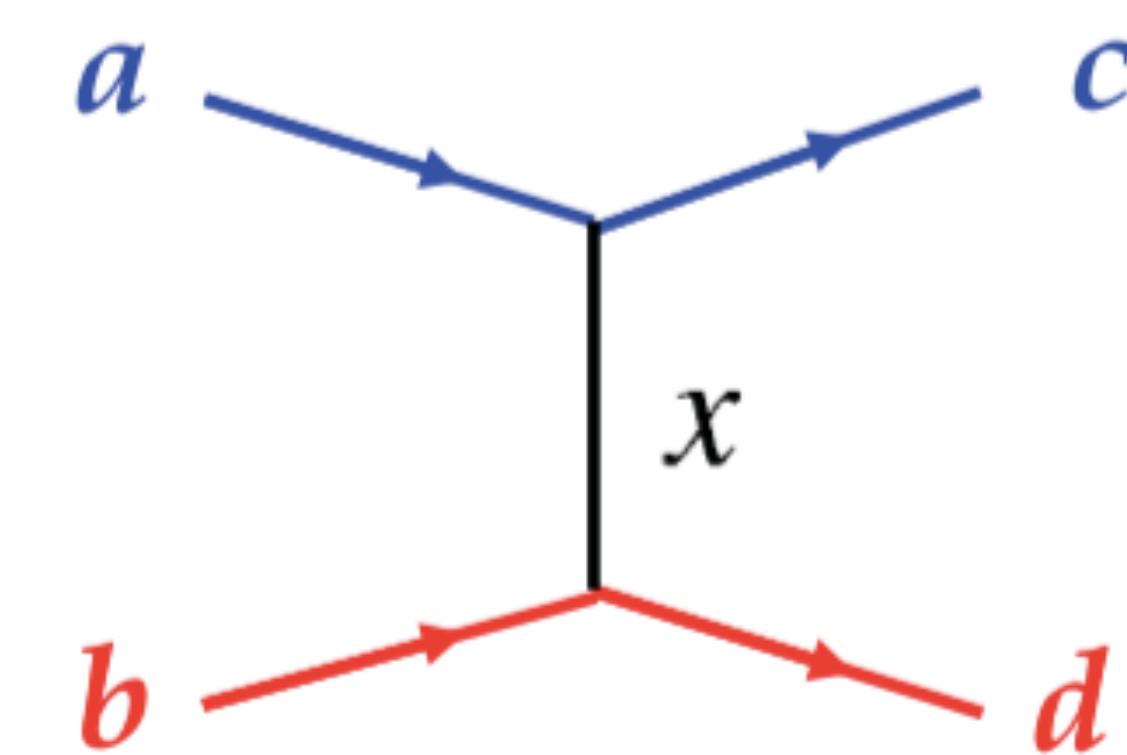
$$\begin{aligned} \Phi(r) &= -\kappa^2 \int \frac{d^3|\vec{k}|}{(2\pi)^3} T_1^{00}(k) \Pi_{0000}(k) T_2^{00}(-k) e^{i\vec{k}\cdot(\vec{r})} \\ &= -\frac{\kappa^2 m}{2} \int \frac{d^3|\vec{k}|}{(2\pi)^3} \frac{1}{\vec{k}^2} e^{i\vec{k}\cdot(\vec{r})} = -\frac{Gm}{r}, \end{aligned}$$

What Aspects of Quantum Gravity Can We test in the Lab ?



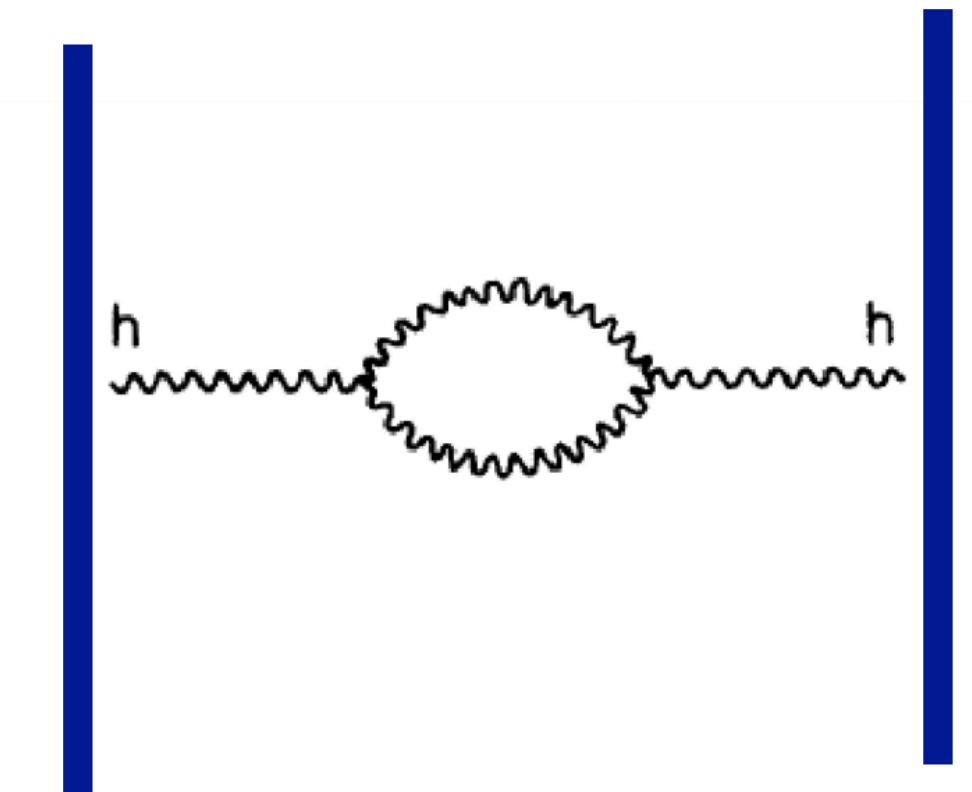
$$\Delta\phi \sim G/\hbar$$

Classical



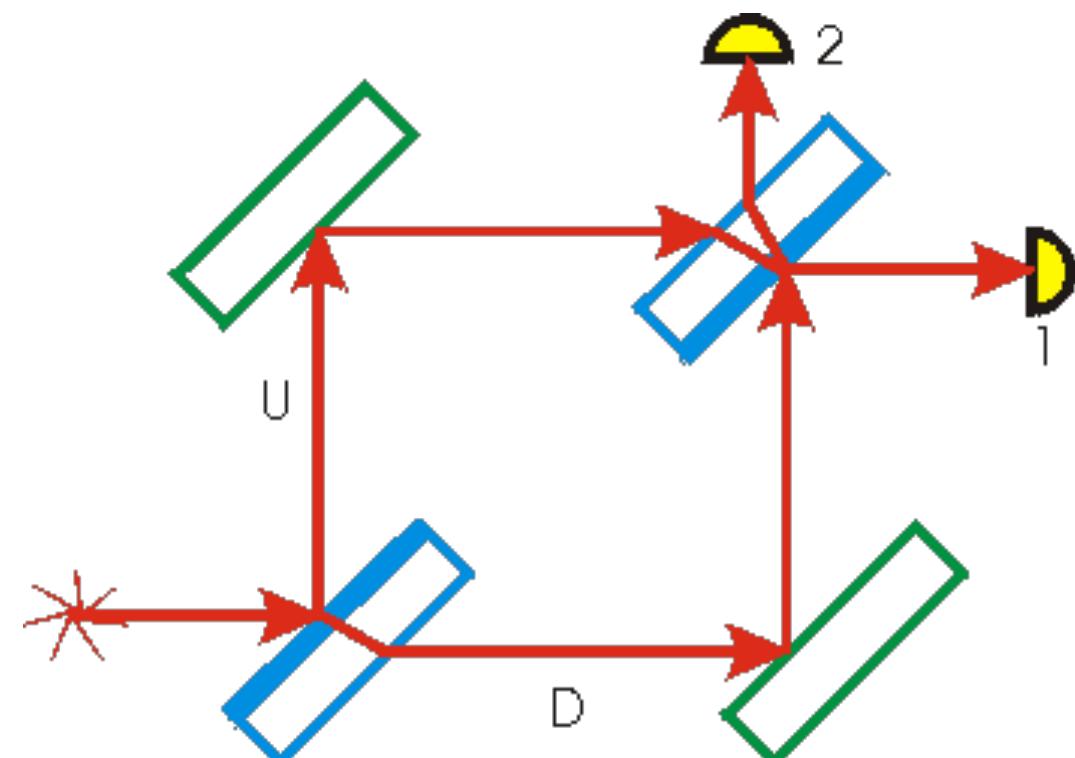
$$\Delta\phi \sim G/\hbar$$

Quantum
2nd order in perturbation



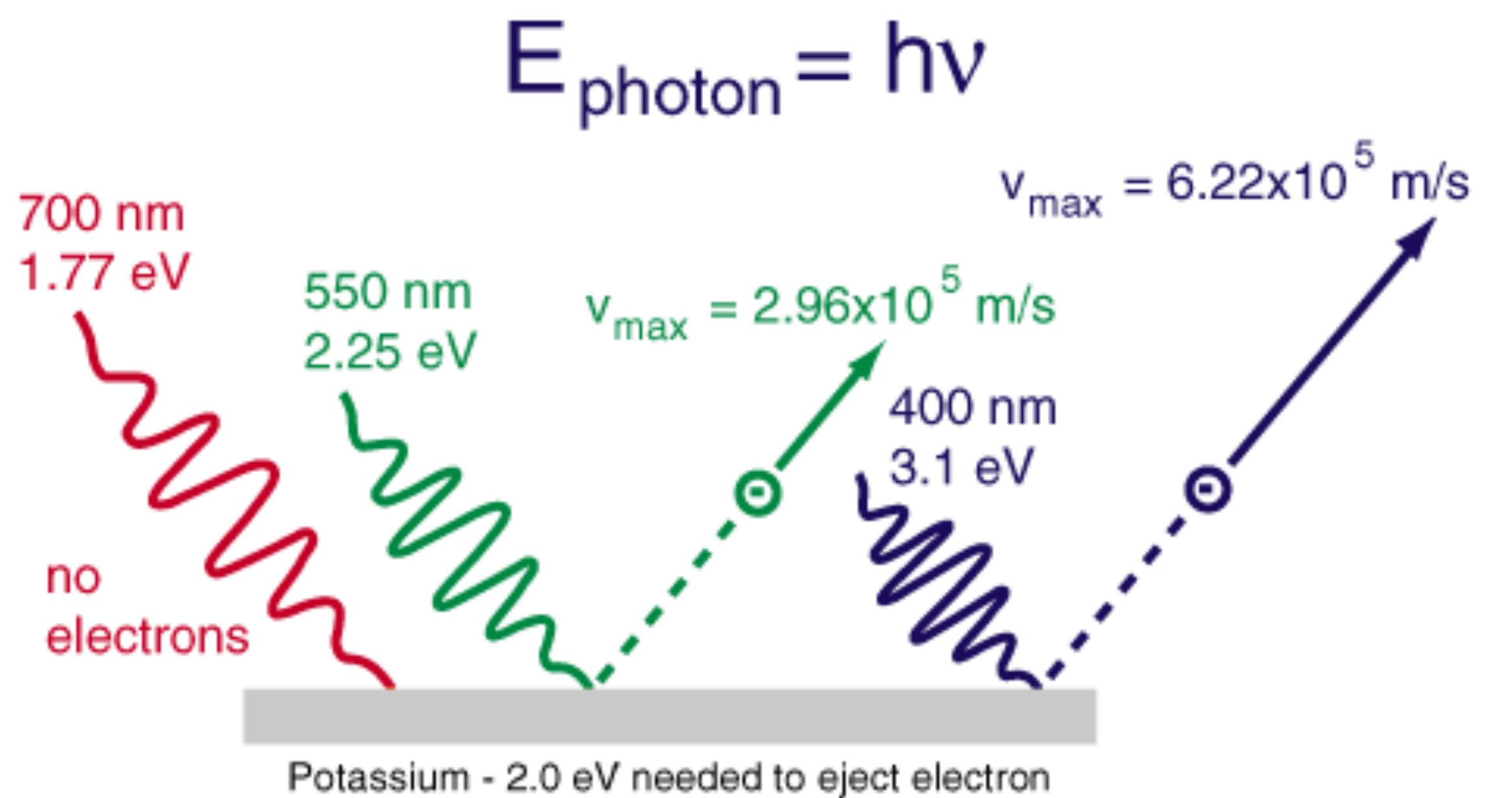
$$\Delta\phi \sim G^2 \hbar/\hbar$$

Quantum Loop
Effect

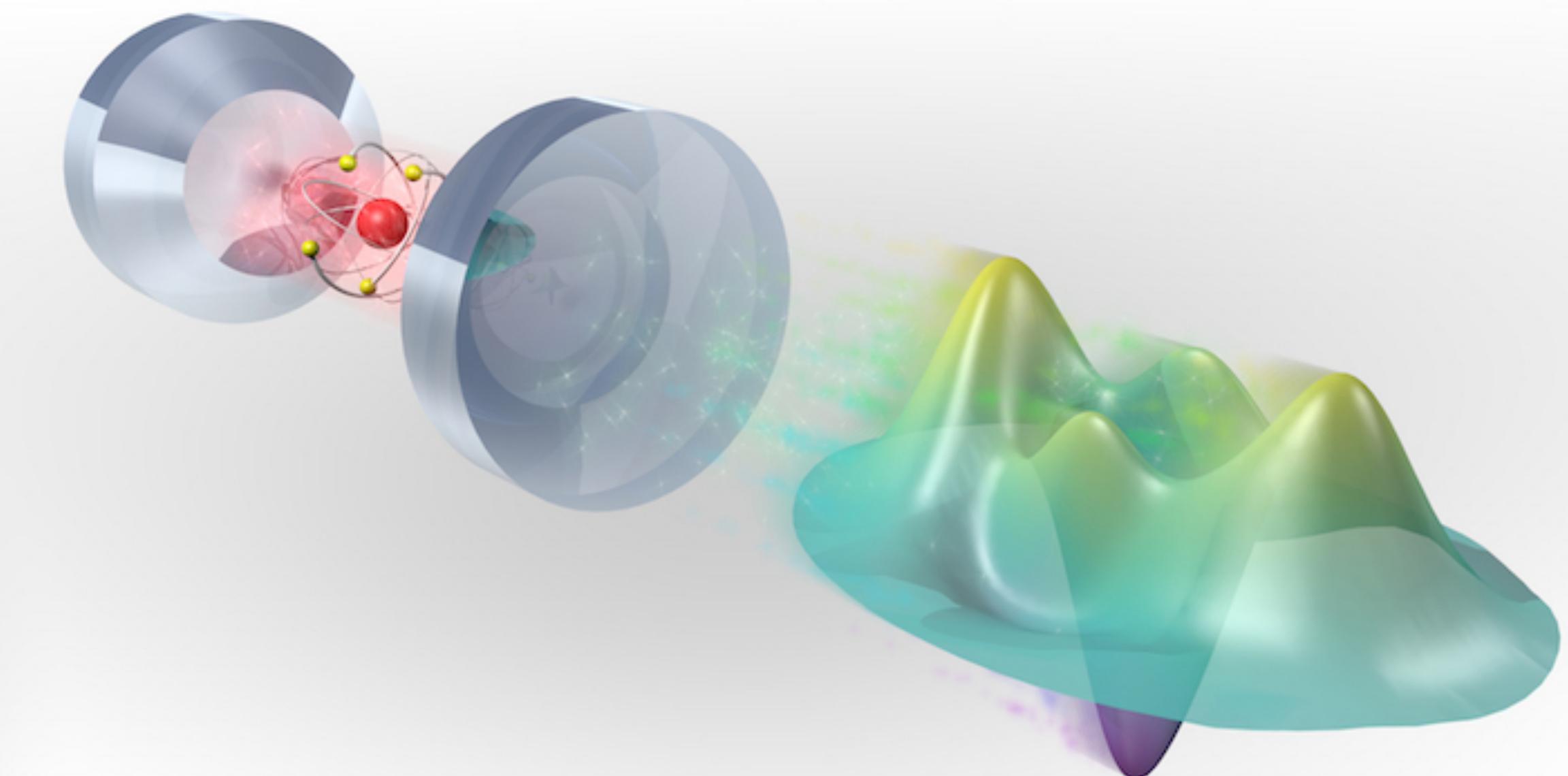


Fine Print
Quantum experiments are full of Off-shell processes
Beam splitter

From Photoelectric Effect to Schrödinger Cat States of Photon



Photoelectric effect



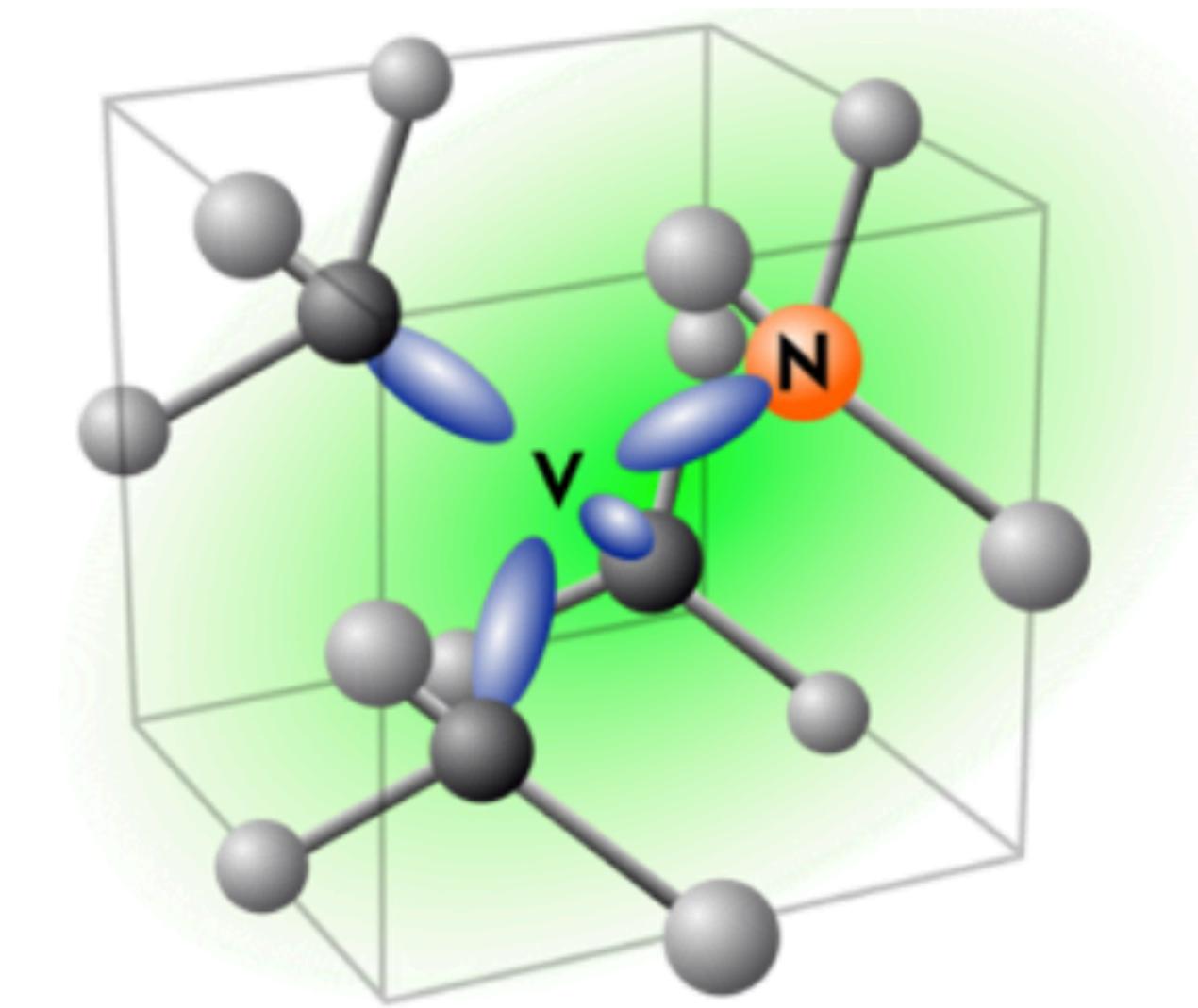
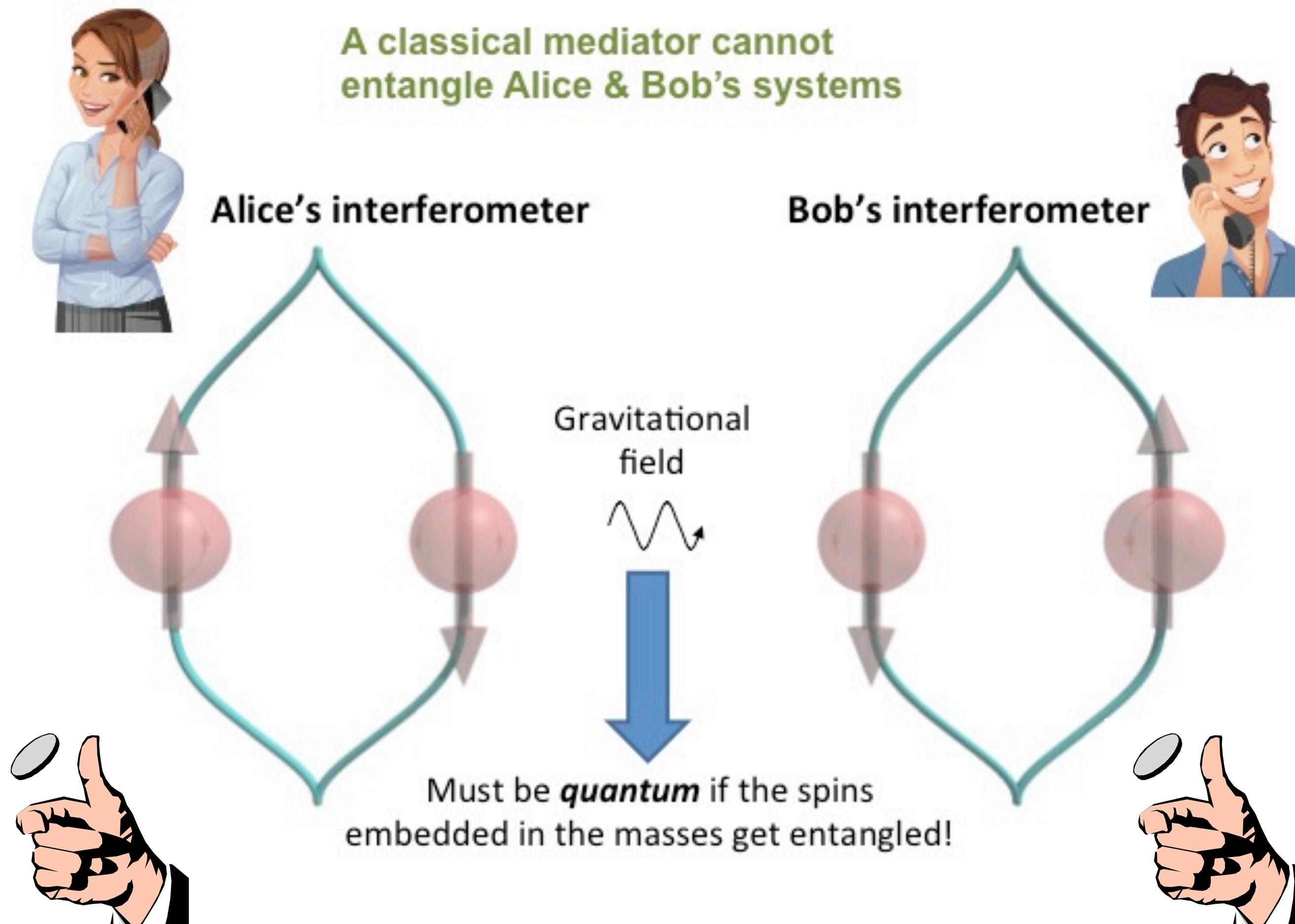
M. Brune, F. Schmidt-Kaler, A. Maali, J. Dreyer, E. Hagley, J.M. Raimond, S. Haroche,
Quantum Rabi Oscillation: A Direct Test of Field Quantization in a Cavity, Physical Review Letters, 76 (1996) 1800-1803.

Creating Schrödinger cat (kitten) states with photons and observing their decoherence:

M. Brune, E. Hagley, J. Dreyer, X. Maitre, A. Maali, C. Wunderlich, J.M. Raimond, S. Haroche,
Observing the progressive decoherence of the "meter" in a quantum measurement, Physical Review Letters, 77 (1996) 4887-4890.

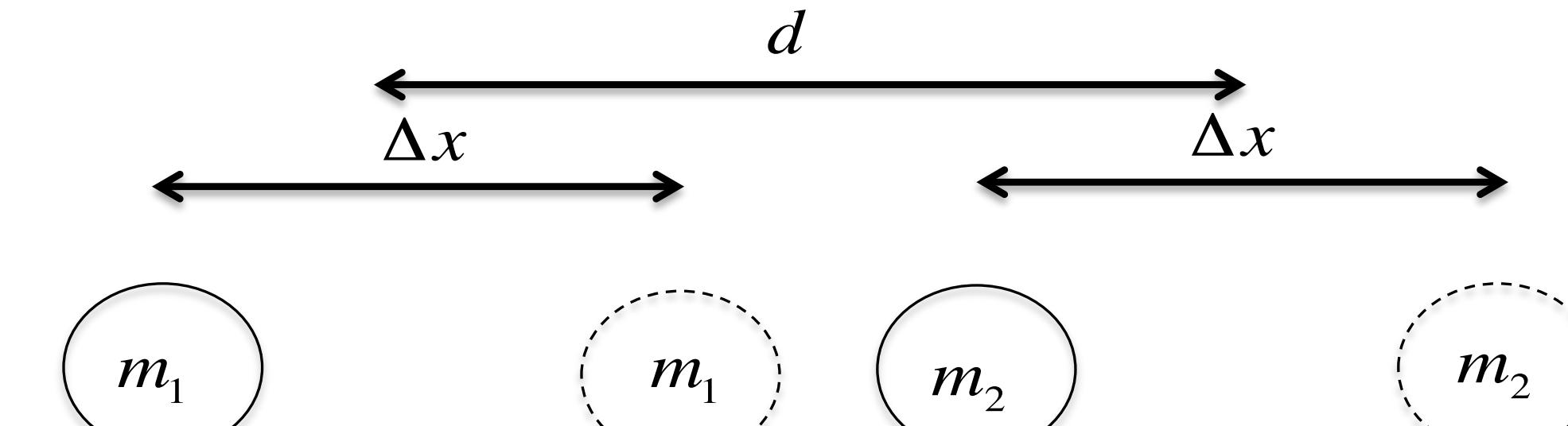
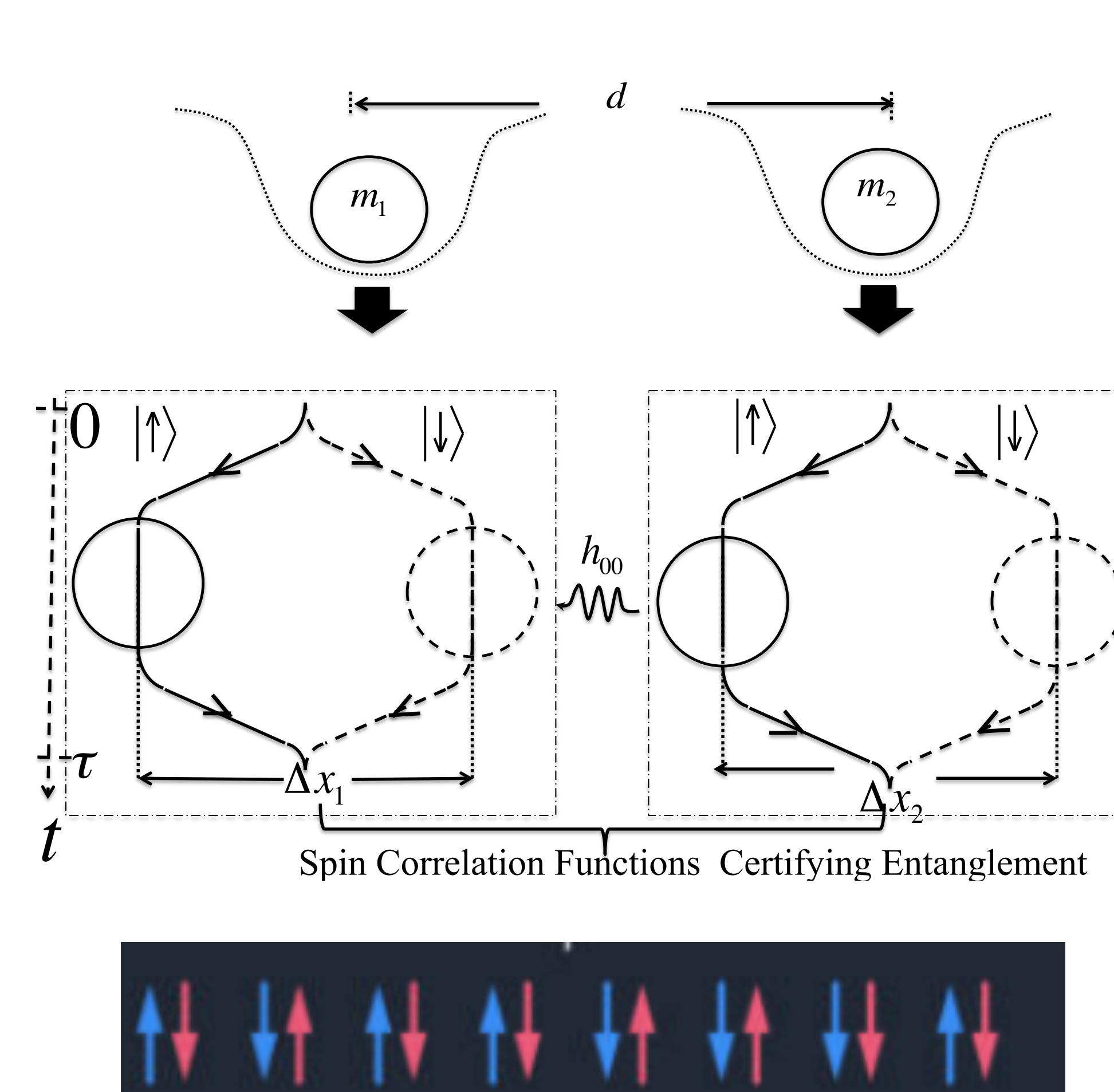
QGEM protocol: 2 Freely Falling Superposed masses

Macroscopic superposition with Spin Embedded Inside



Use very strong atom chip gradients to push and split massive objects using just one spin!

QGEM protocol: 2 Freely Falling Superposed masses



If they interact *only* through the gravitational force

$$|L\rangle_1 \quad |R\rangle_1 \quad |L\rangle_2 \quad |R\rangle_2$$

$$|\Psi(t=0)\rangle_{12} = \frac{1}{\sqrt{2}}(|L\rangle_1 + |R\rangle_1) \frac{1}{\sqrt{2}}(|L\rangle_2 + |R\rangle_2)$$

$$= \frac{1}{2}(|L\rangle_1|L\rangle_2 + |L\rangle_1|R\rangle_2 + |R\rangle_1|L\rangle_2 + |R\rangle_1|R\rangle_2)$$

$$\rightarrow |\Psi(t=\tau)\rangle_{12} = \frac{1}{2}(e^{i\phi_{LL}}|L\rangle_1|L\rangle_2 + e^{i\phi_{LR}}|L\rangle_1|R\rangle_2 + e^{i\phi_{RL}}|R\rangle_1|L\rangle_2 + e^{i\phi_{RR}}|R\rangle_1|R\rangle_2),$$

where

$$\phi_{RL} \sim \frac{Gm_1m_2\tau}{\hbar(d - \Delta x)}, \phi_{LR} \sim \frac{Gm_1m_2\tau}{\hbar(d + \Delta x)},$$

$$\phi_{LL} = \phi_{RR} \sim \frac{Gm_1m_2\tau}{\hbar d}$$

Maximum Entanglement Phase

Step 4: Witness spin entangled state:

$$\begin{aligned} |\Psi(t = t_{\text{End}})\rangle_{12} &= \frac{1}{\sqrt{2}} \left\{ |\uparrow\rangle_1 \frac{1}{\sqrt{2}} (|\uparrow\rangle_2 + e^{i\Delta\phi_{LR}} |\downarrow\rangle_2) \right. \\ &\quad \left. + |\downarrow\rangle_1 \frac{1}{\sqrt{2}} (e^{i\Delta\phi_{RL}} |\uparrow\rangle_2 + |\downarrow\rangle_2) \right\} |C\rangle_1 |C\rangle_2 \end{aligned}$$

through the correlations:

$$\mathcal{W} = |\langle \sigma_x^{(1)} \otimes \sigma_z^{(2)} \rangle - \langle \sigma_y^{(1)} \otimes \sigma_z^{(2)} \rangle|$$

we have

If $\mathcal{W} > 1 \implies \text{Graviton is quantum}$

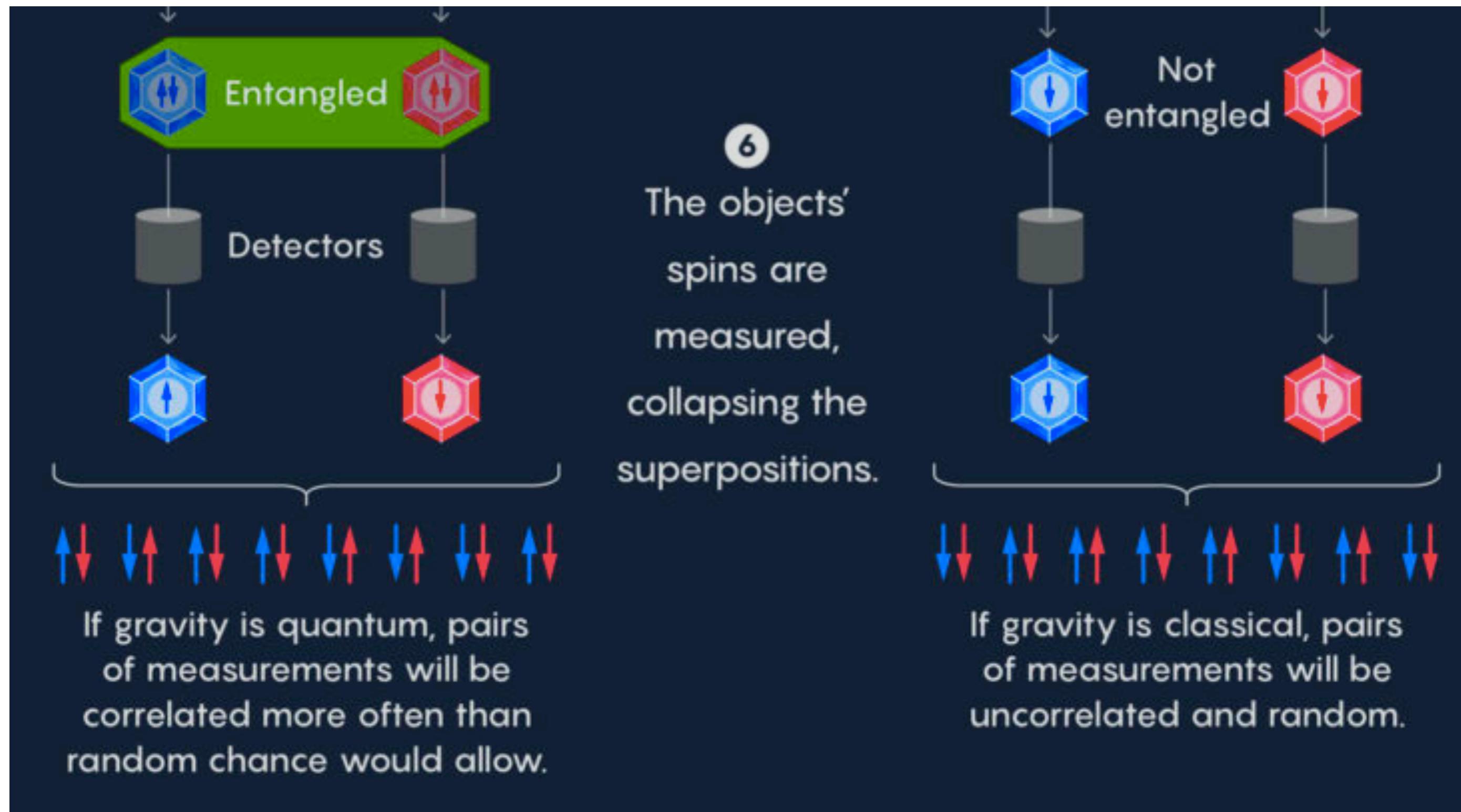
$$\Delta\phi_{LR} + \Delta\phi_{RL} \sim \mathcal{O}(1)$$

$$\Delta\phi_{RL} \sim \frac{Gm_1m_2\tau}{\hbar(d - \Delta x)} \gg \Delta\phi_{LR}, \Delta\phi_{LL}, \Delta\phi_{RR}$$

For mass $\sim 10^{-14}$ kg (microspheres), separation at closest approach of the masses ~ 200 microns (to prevent Casimir interaction), **time ~ 1 seconds**, gives:
Scale of superposition ~ 100 microns, **Delta phi_{RL} ~ 1**

Planck's Constant fights Newton's Constant!

Entanglement Witness: Spin Correlation



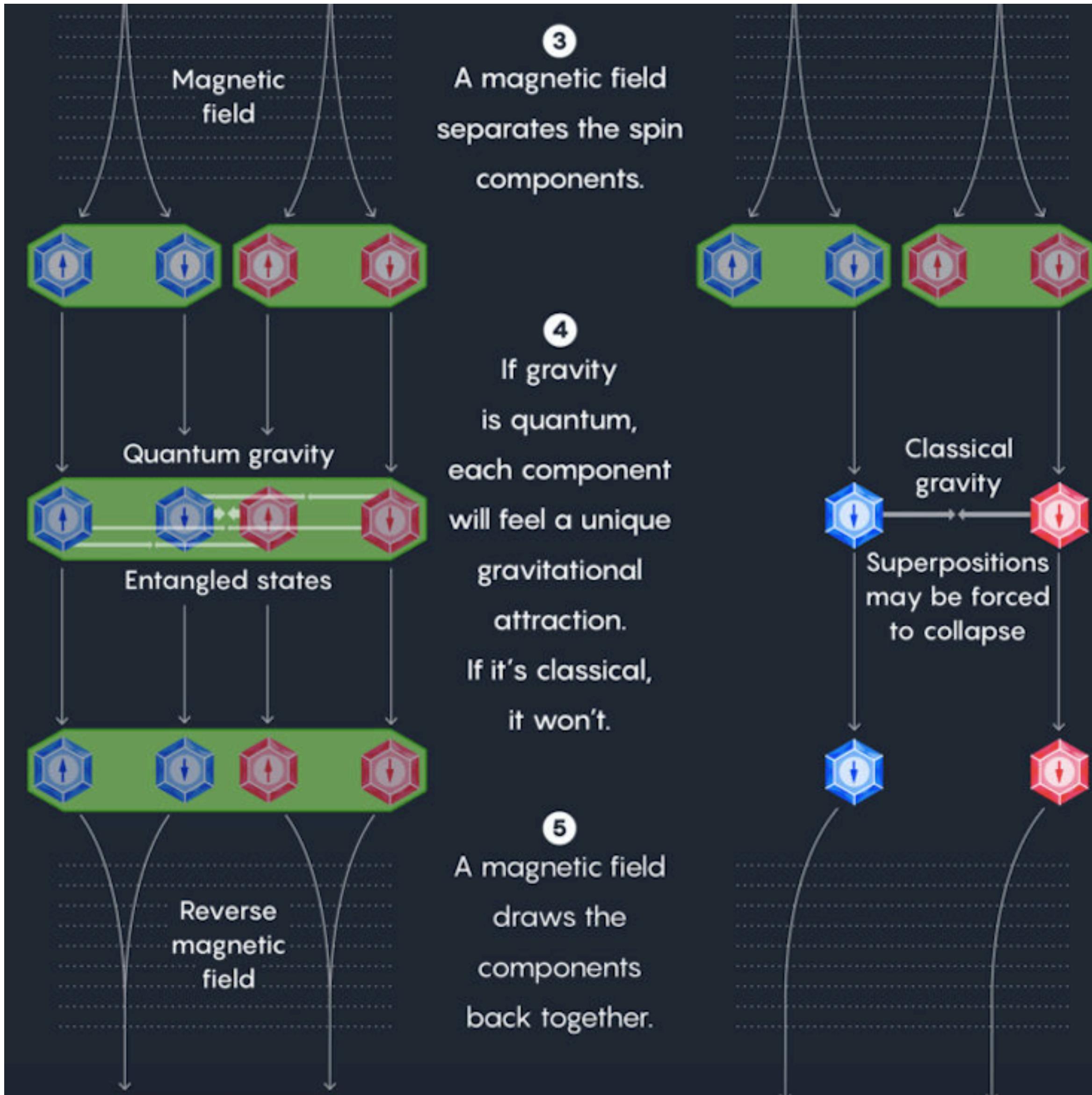
Basis Dependent Witness, similar to Bell's

$$\mathcal{W} = I^{(1)}I^{(2)} - \sigma_x^{(1)}\sigma_x^{(2)} - \sigma_y^{(1)}\sigma_z^{(2)} - \sigma_x^{(1)}\sigma_z^{(2)} \quad \text{Tr}(\mathcal{W}\rho) < 0$$

H. Chevalier, A. Paige, and M. Kim
arXiv:2005.13922 [quant-ph].

Basis Independent Witness: $S_A = -\text{Tr}_A\rho_A \log \rho_A = -S_B$ Entanglement Entropy!

Challenges for QGEM



10^{-14} Kg

Radius : 100nm

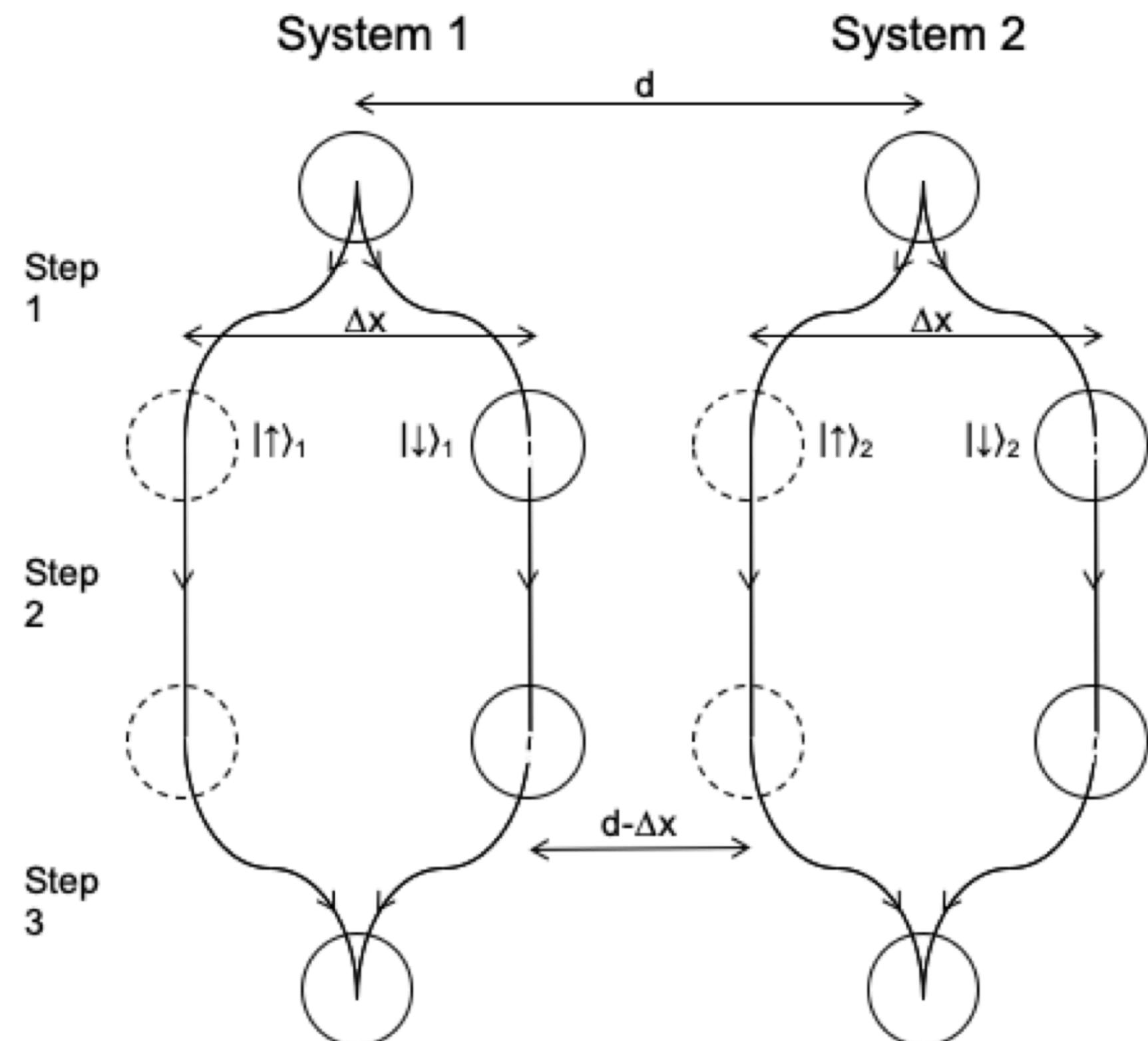
Neutralising e.m. charges

A magnetic field gradient of $\sim 10^6 \text{ T/m}$ and a time $\tau_{\text{acc}} \sim 500 \text{ m/s}^2$, $\Delta x \sim 250\mu\text{m}$, $d - \Delta x \sim 200\mu\text{m}$

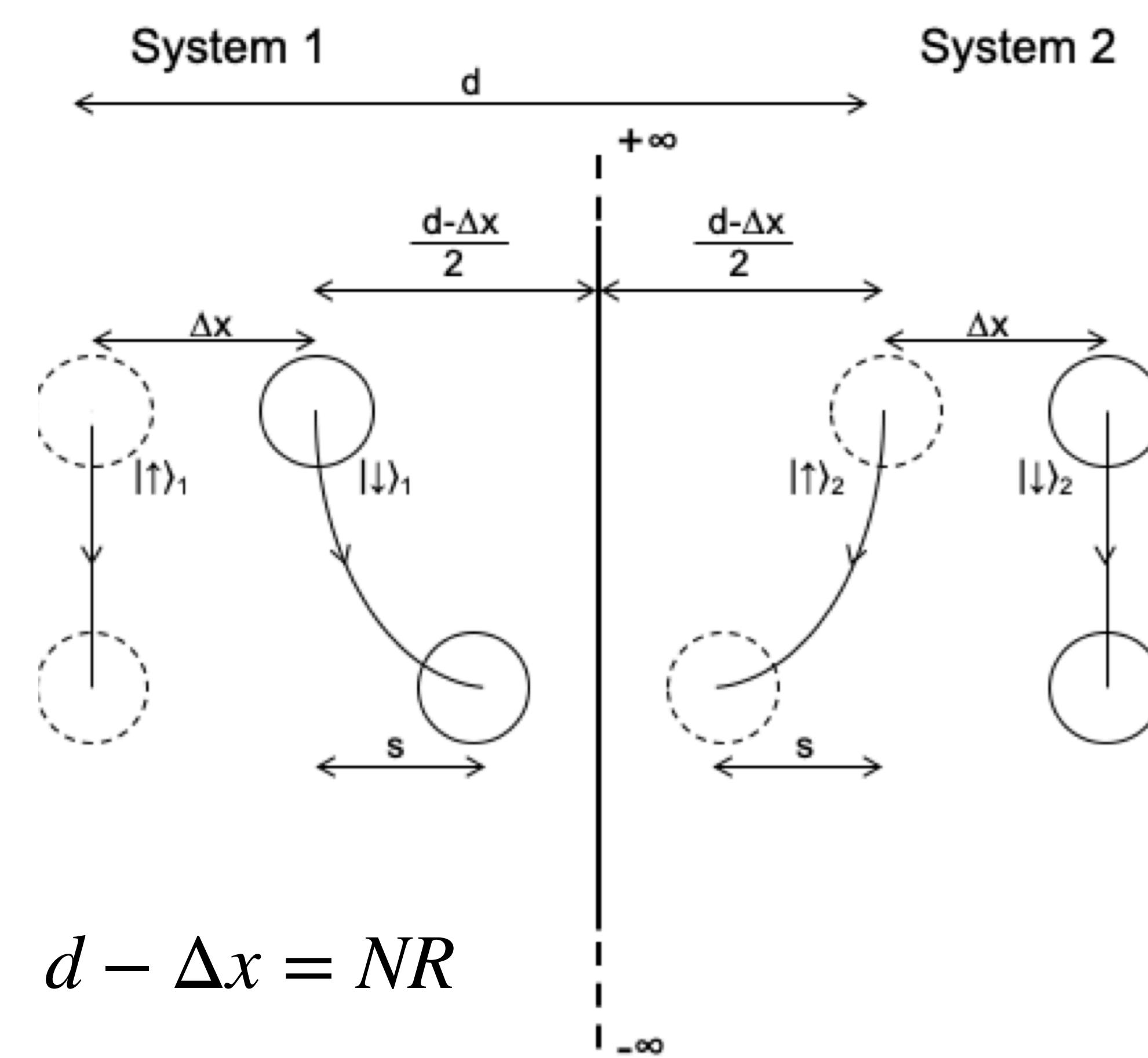
Electronic spins coherent for 1s, which should be possible for macro-diamond below 77 K

To estimate collisional and thermal decoherence times of the orbital degree of freedom we consider the pressure $P = 10^{-15} \text{ Pa}$ and the temperature 0.15 K. the collisional decoherence time for a superposition size of $\Delta x \sim 250\mu\text{m}$ is the same order of magnitude as the total microsphere's fall time $\tau + 2\tau_{\text{acc}} \sim 3.5 \text{ s}$

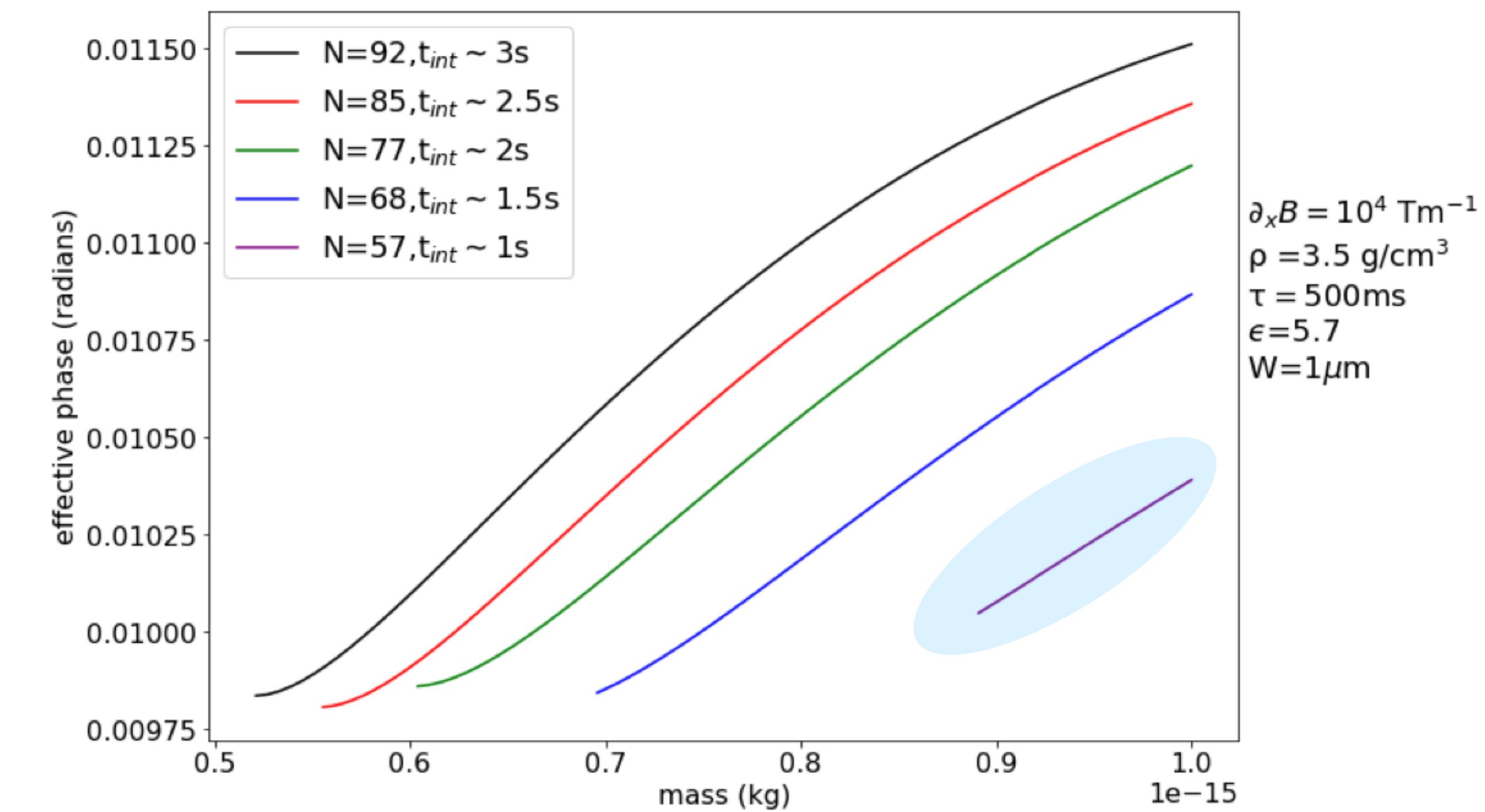
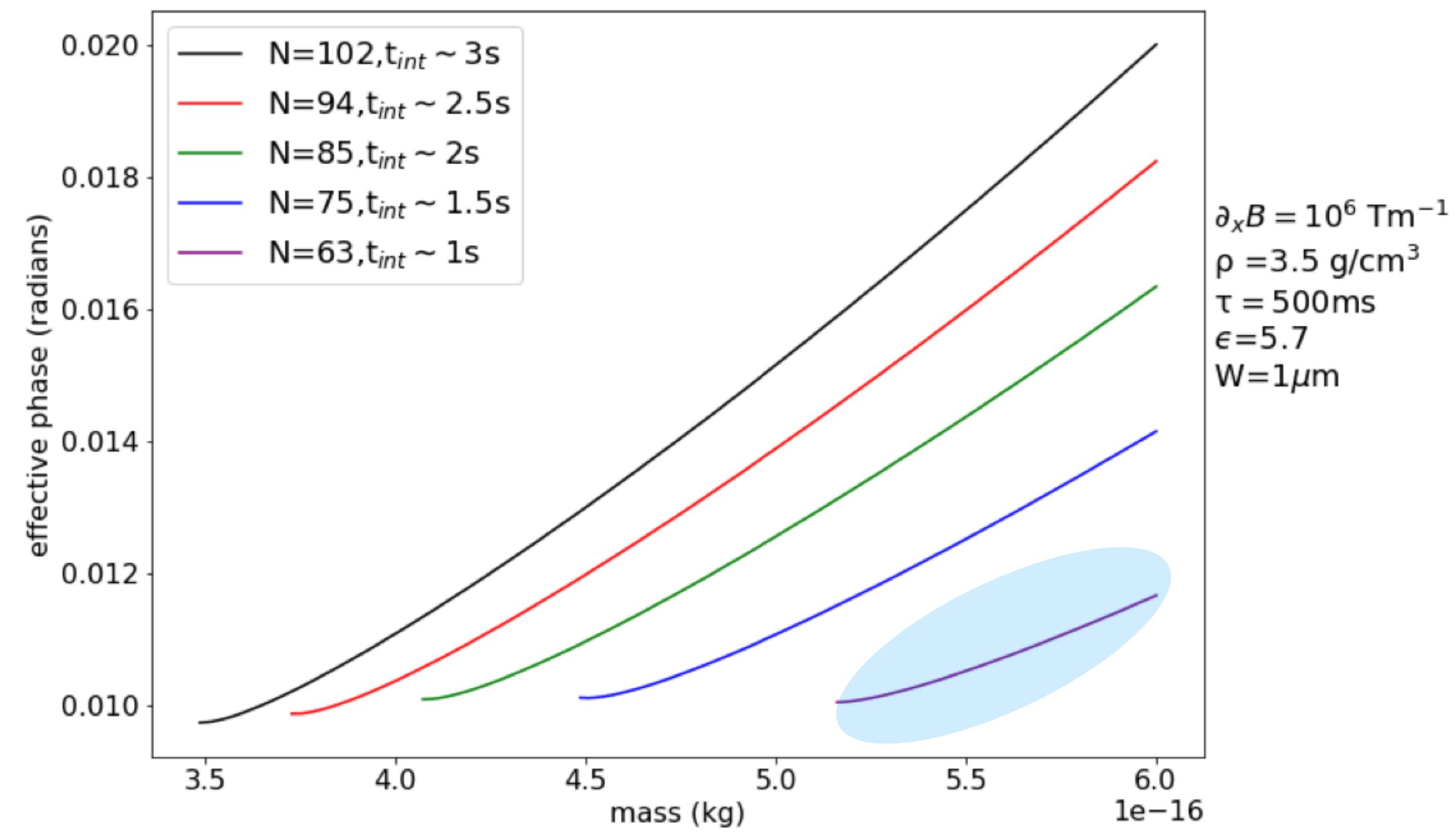
QGEM Protocol : Improved Design



$$V_{CP} \sim -\frac{23\hbar c}{4\pi} \frac{R^6}{r^7} \left(\frac{\epsilon - 1}{\epsilon + 2} \right)^2$$



QGEM Protocol (Phase vs Mass)



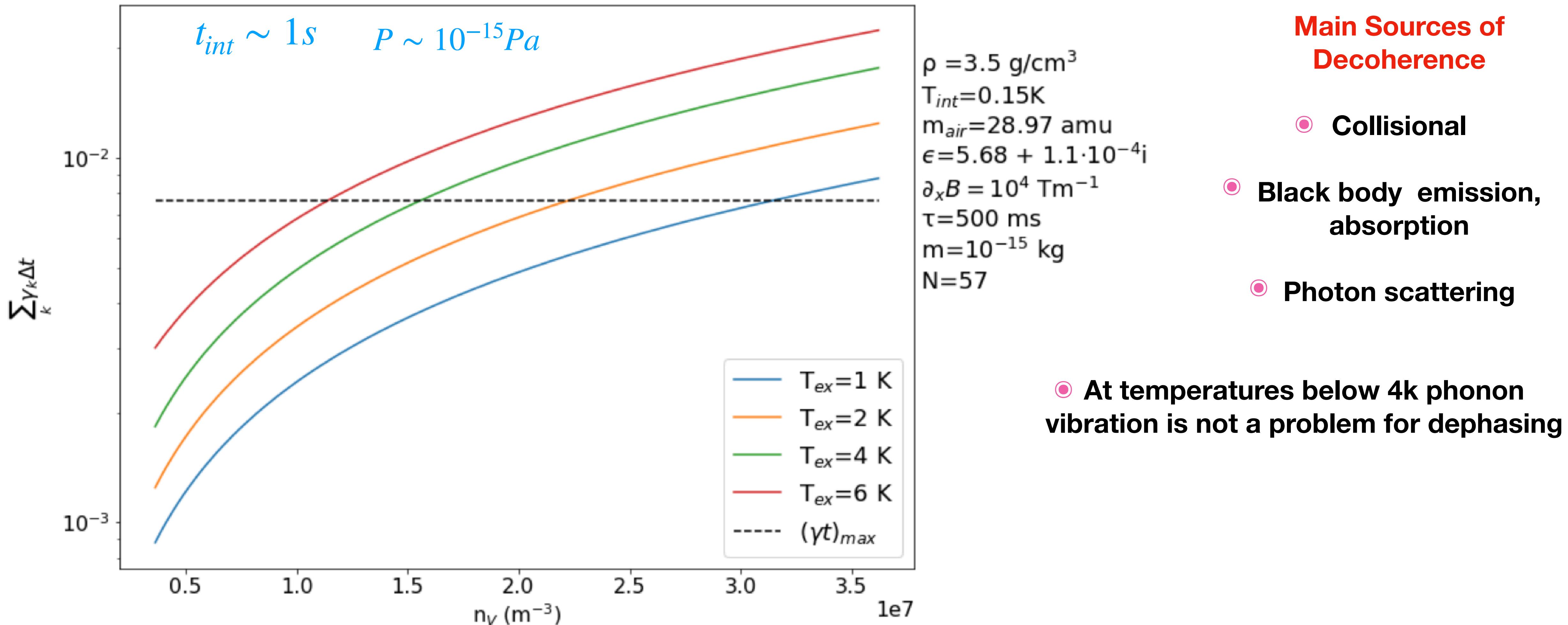
$$10^{-14}\text{Kg} \implies 10^{-16} - 10^{-15}\text{Kg}$$

$$\Delta x \sim 200\mu\text{m} \implies 20\mu\text{m}$$

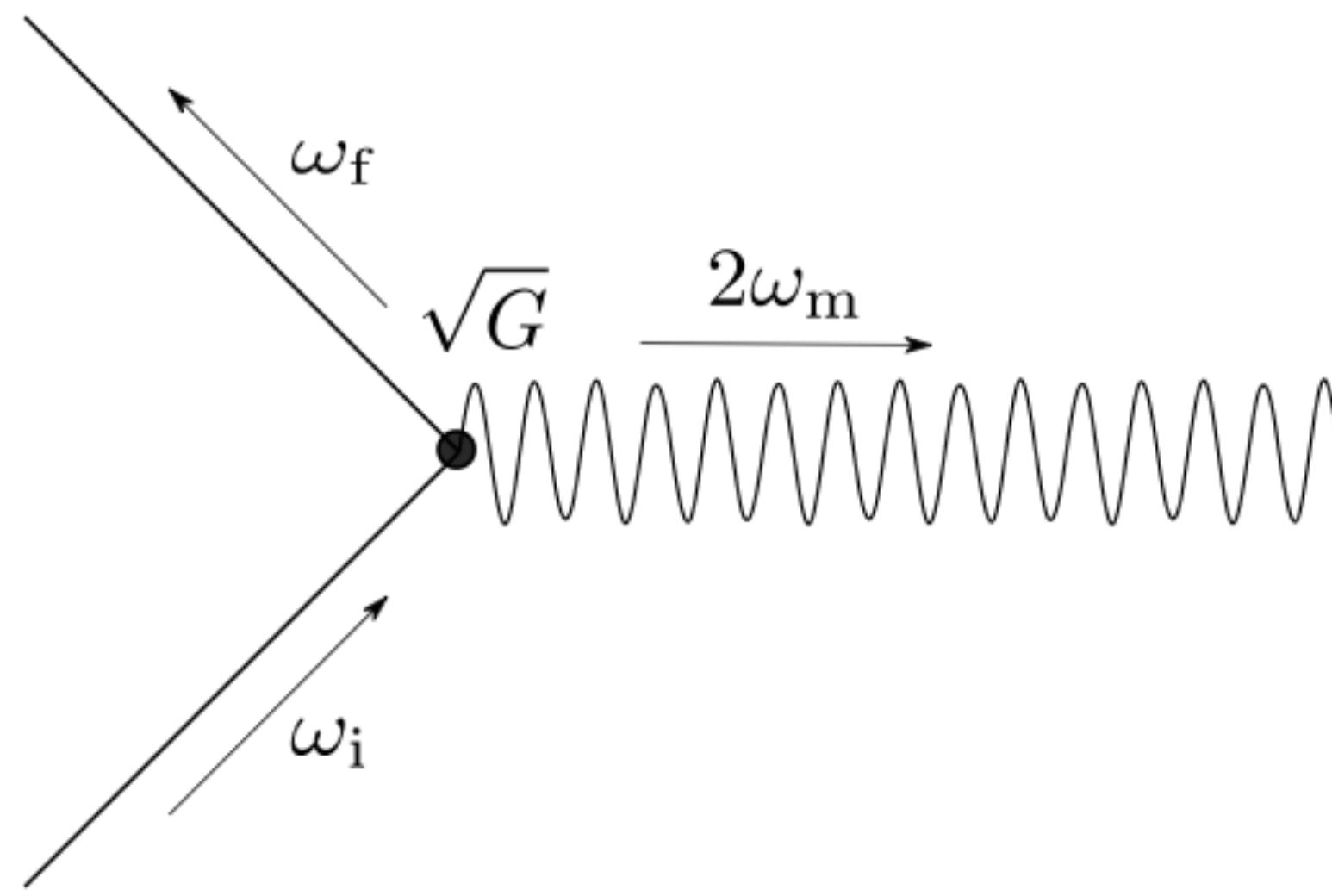
$$P \sim 10^{-16} \implies 10^{-15}\text{Pa}$$

Sources of Dephasing/Decoherence

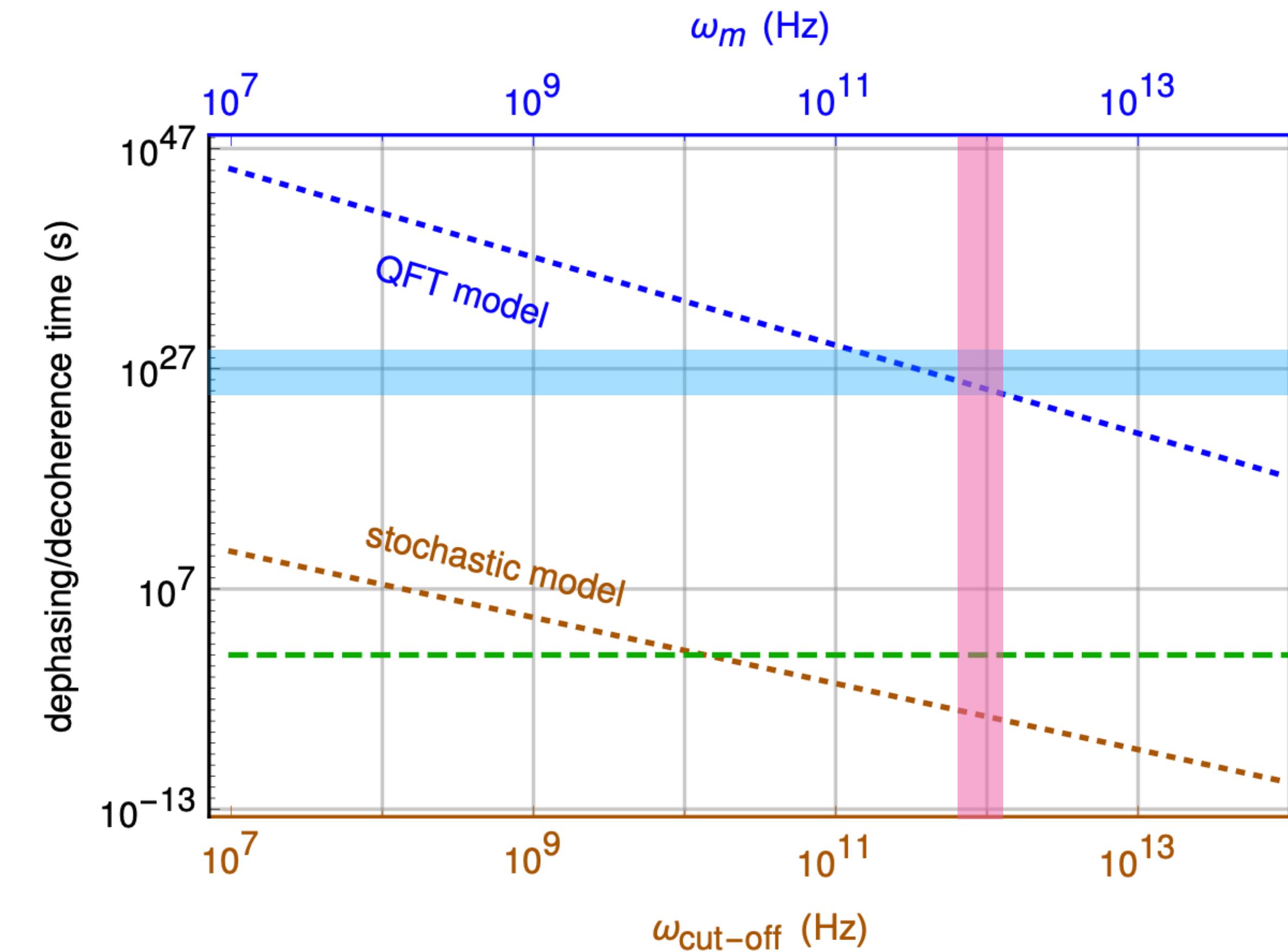
$$\mathcal{W} = I^{(1)}I^{(2)} - \sigma_x^{(1)}\sigma_x^{(2)} - \sigma_y^{(1)}\sigma_z^{(2)} - \sigma_x^{(1)}\sigma_z^{(2)} \quad \text{Tr}(\mathcal{W}\rho) < 0 \quad \Rightarrow \gamma_{dec}t_{int} < \Phi/2 \sim 0.0075$$



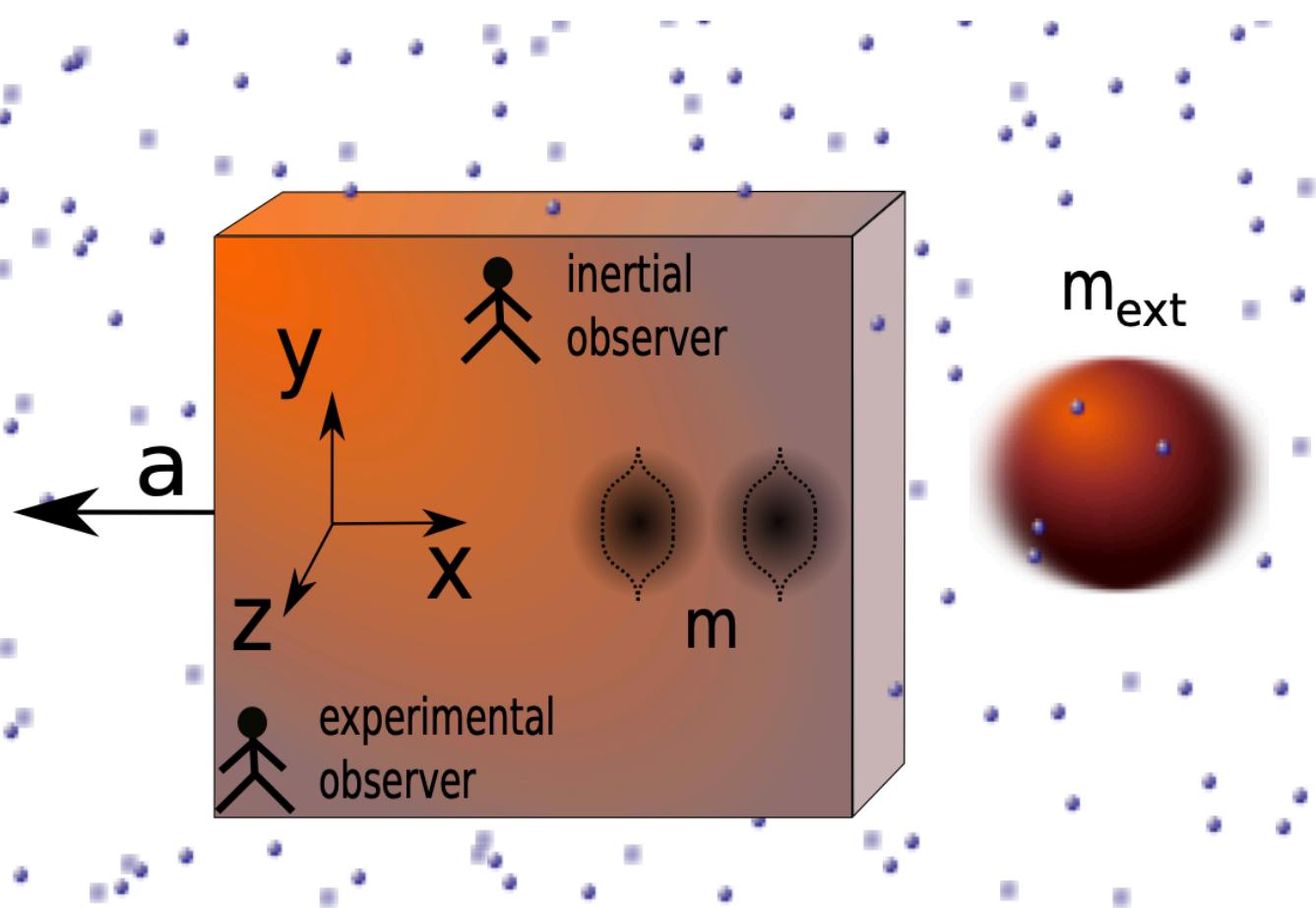
Dephasing due to Graviton is extremely tiny



$$\gamma_{grav} = \frac{32G\hbar\omega_m^3}{15c^5} = \frac{32}{15}t_{pl}^2\omega_m^3$$

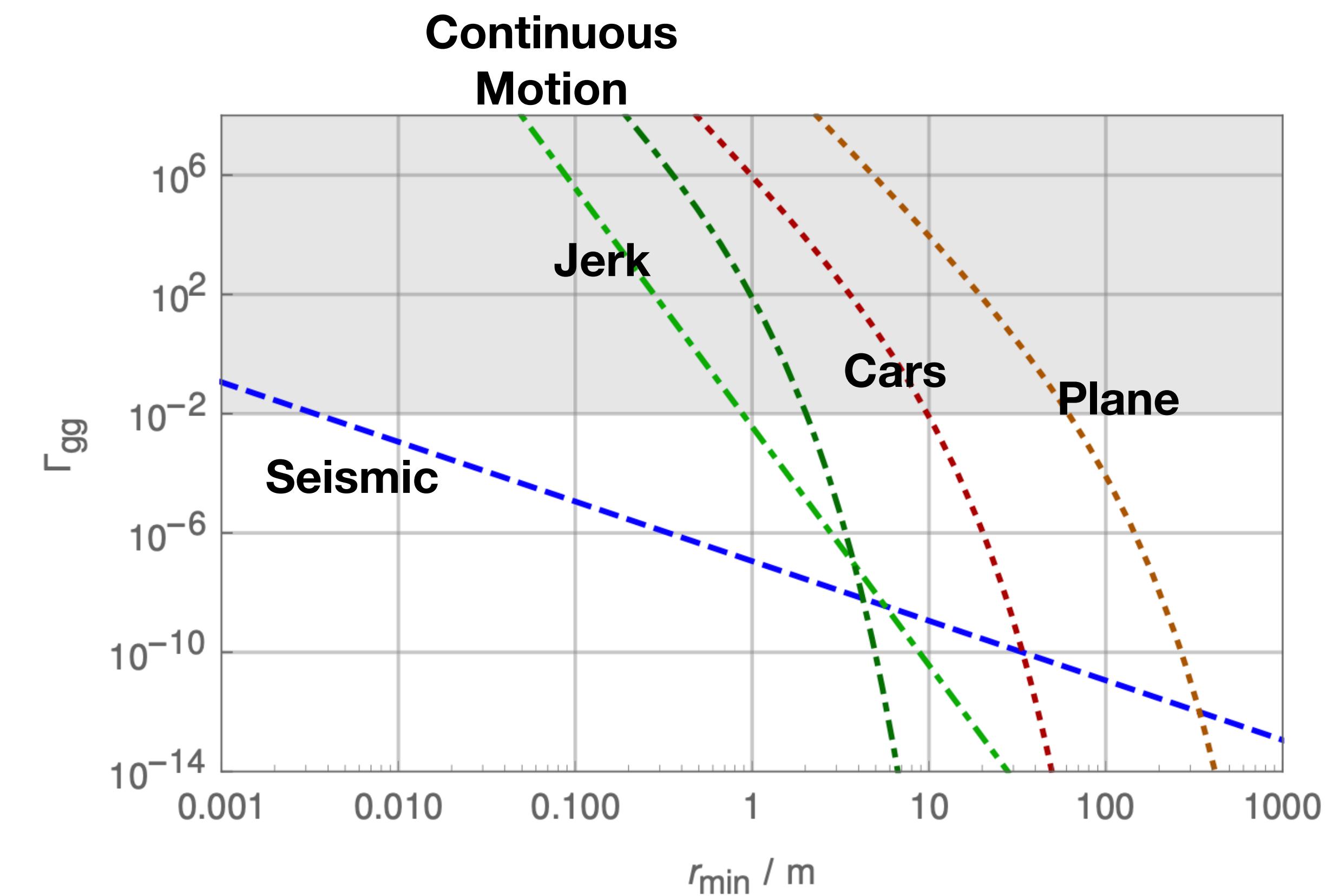
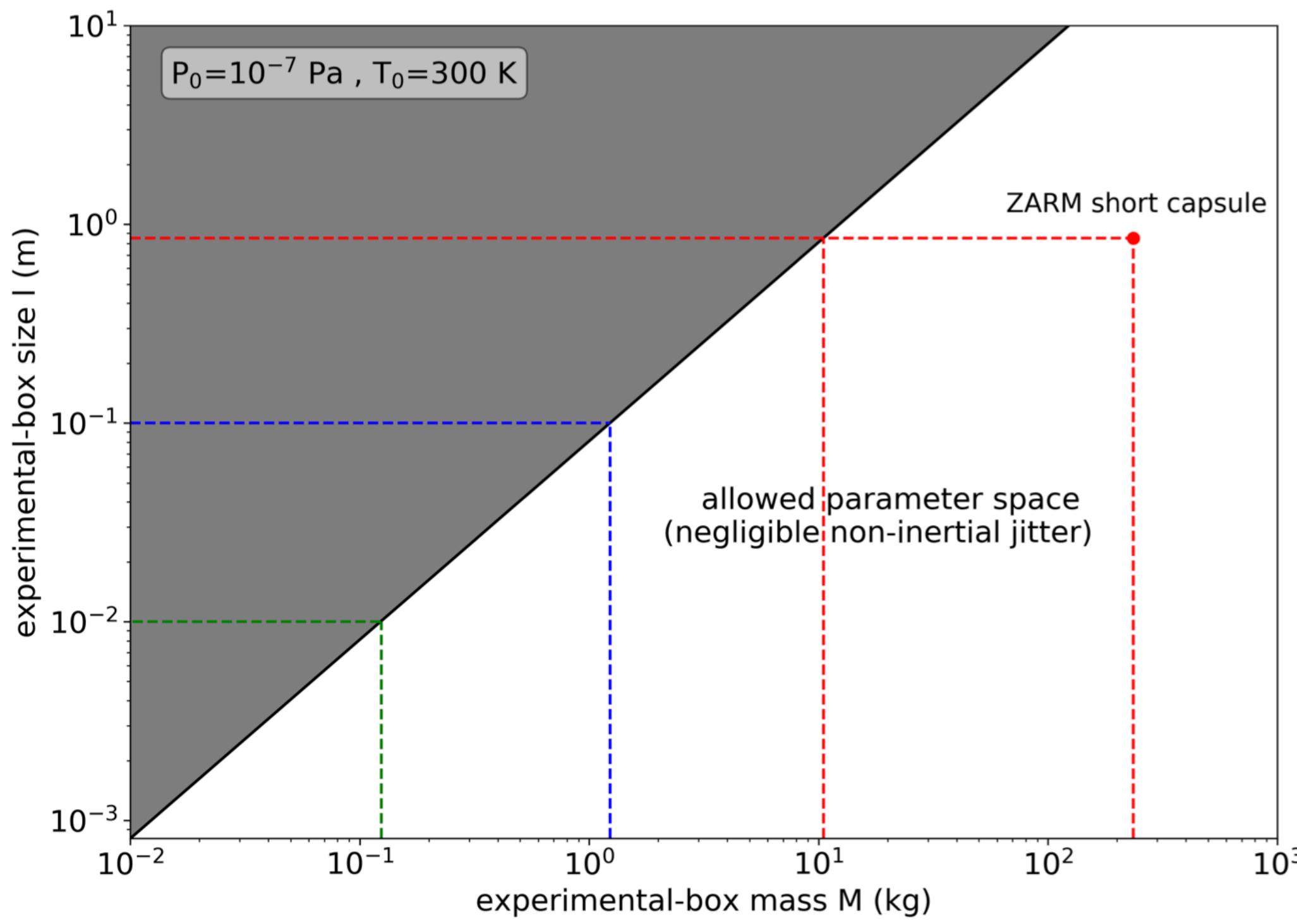


Relative Acceleration+Gravity Gradient Noise



$$\Phi_{eff} > \Gamma_d + \Gamma_{non-inertial} + \Gamma_{jitter} + \Gamma_{ggn}$$

Non-inertial, GGN
(Atmospheric+Seismic+ Constant
movement+ Jerks+Cars+Planes)

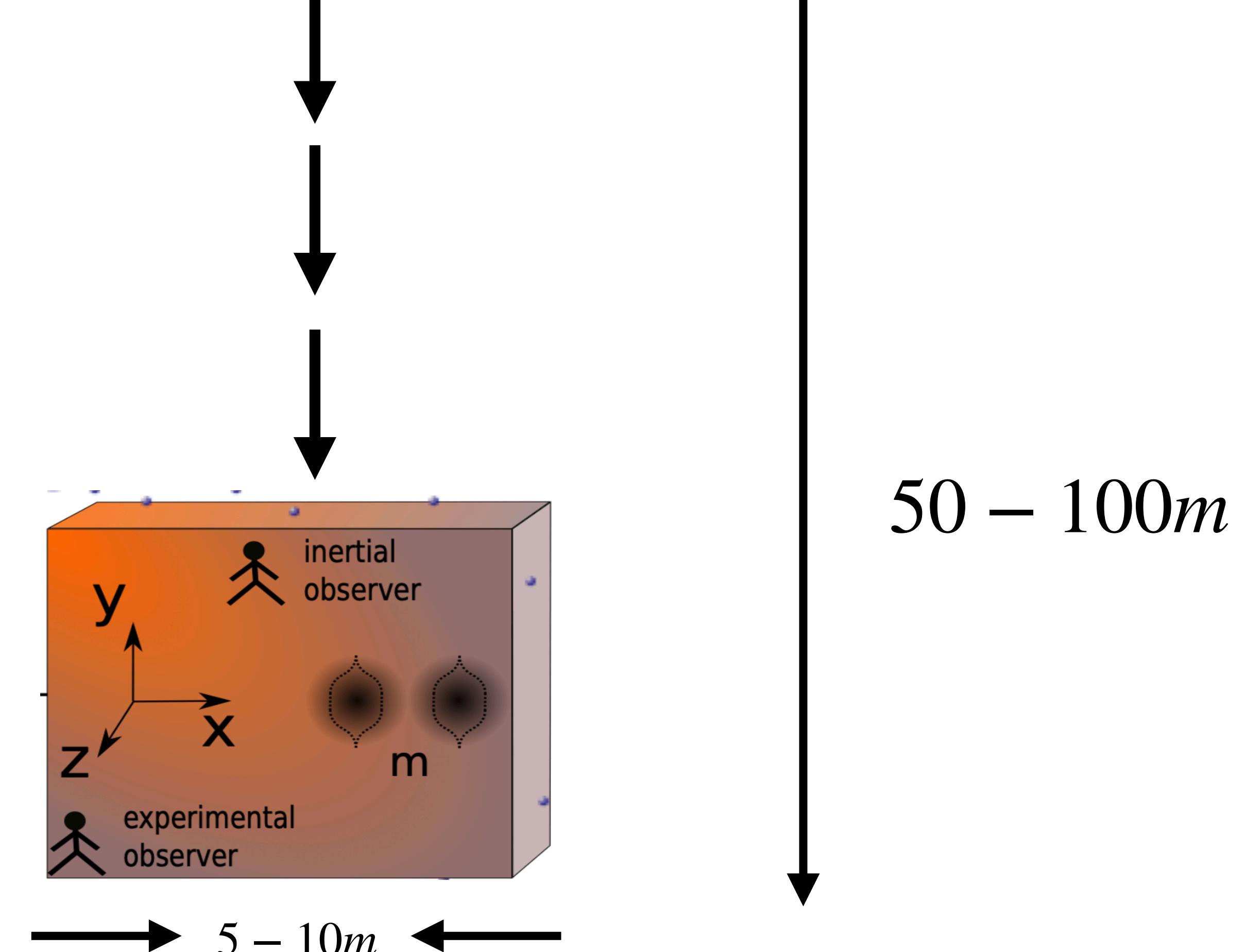


Underground Laboratory

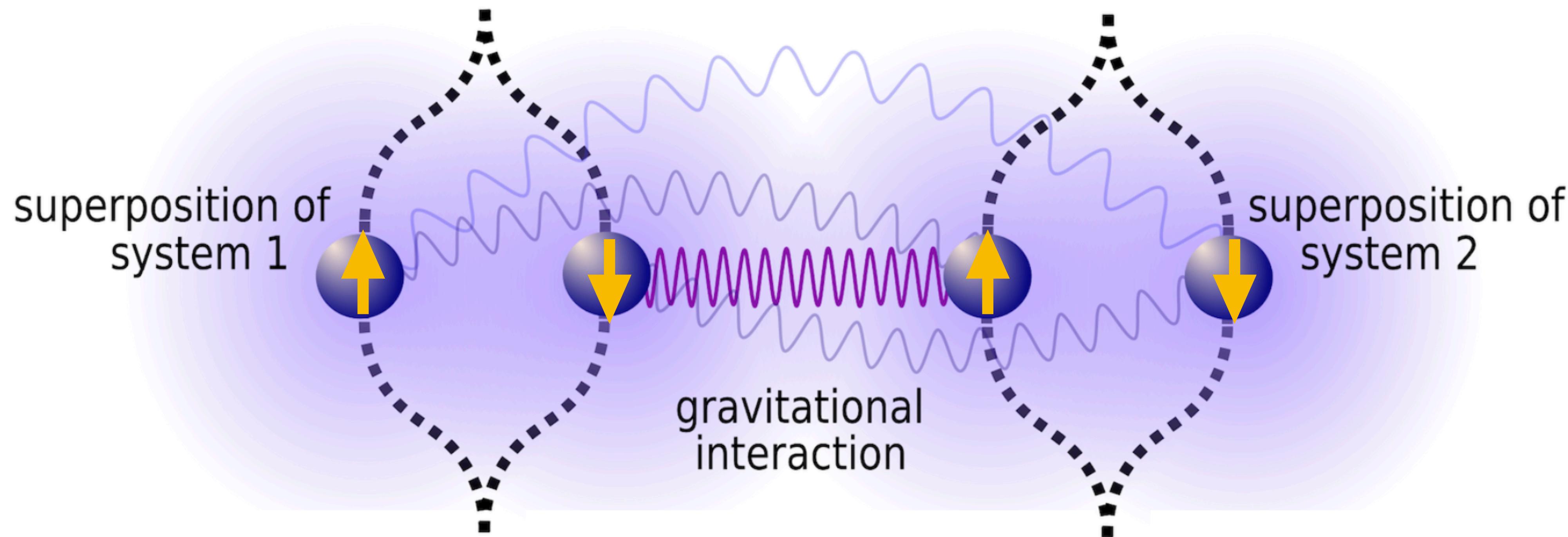


Drop Tower

Seismic Noise is Low



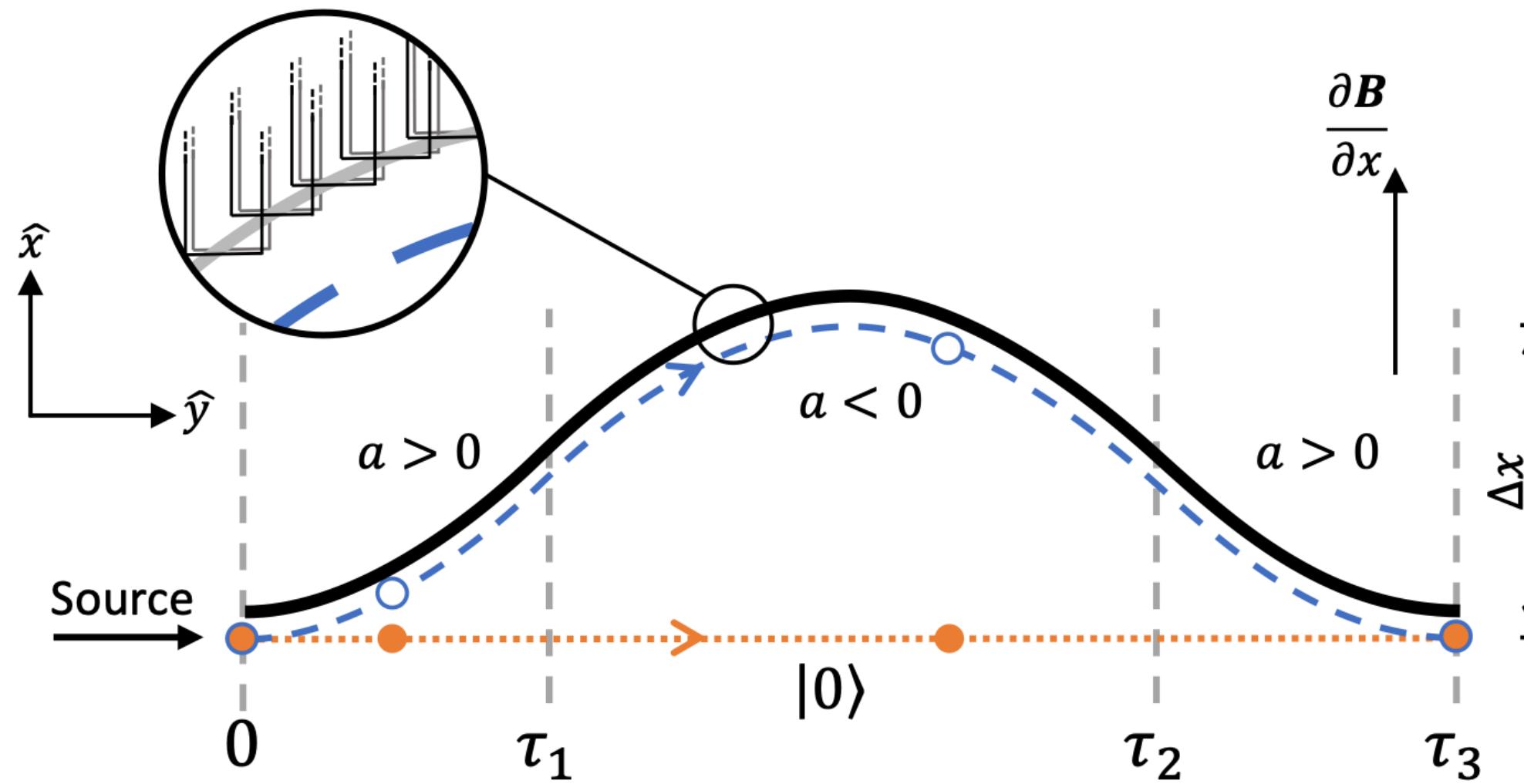
Formidable Challenges



$$H = \frac{1}{2m} \vec{p}^2 + mg\hat{z} + \hbar D \hat{S}_z^2 - \frac{\chi_\nu V}{2\mu_0} \vec{B}^2 - \hbar \gamma_e \vec{S} \cdot \vec{B}$$

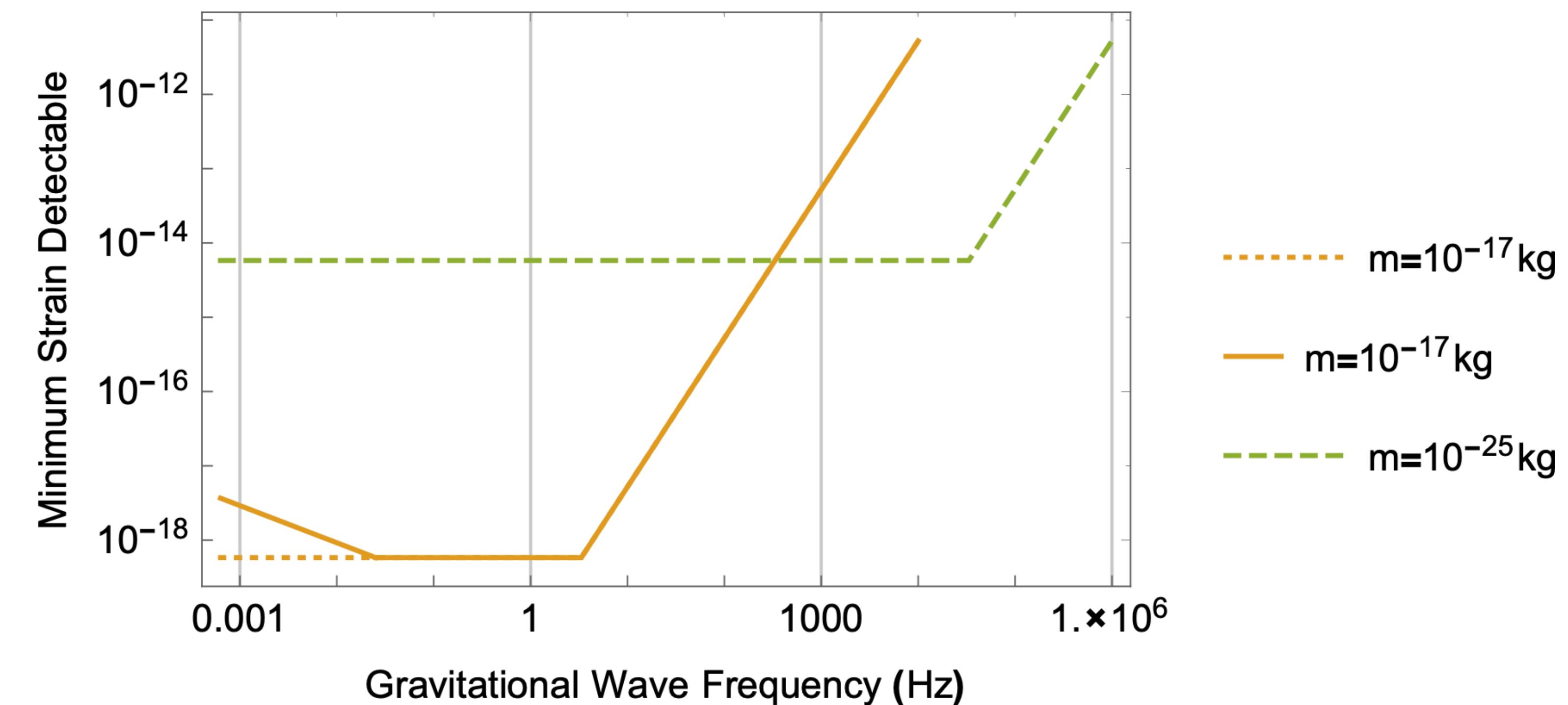
Zero-field Splitting Diamagnetic part NV-part

MIMAC: Mesoscopic Interferometer for Metric and Curvature



Sub-Hz Gravitational Wave Detector

We need to Improve upon the Strain Sensitivity



Theoretical Aspects

Alice, Bob and Eve

We are
all
Entangled

:

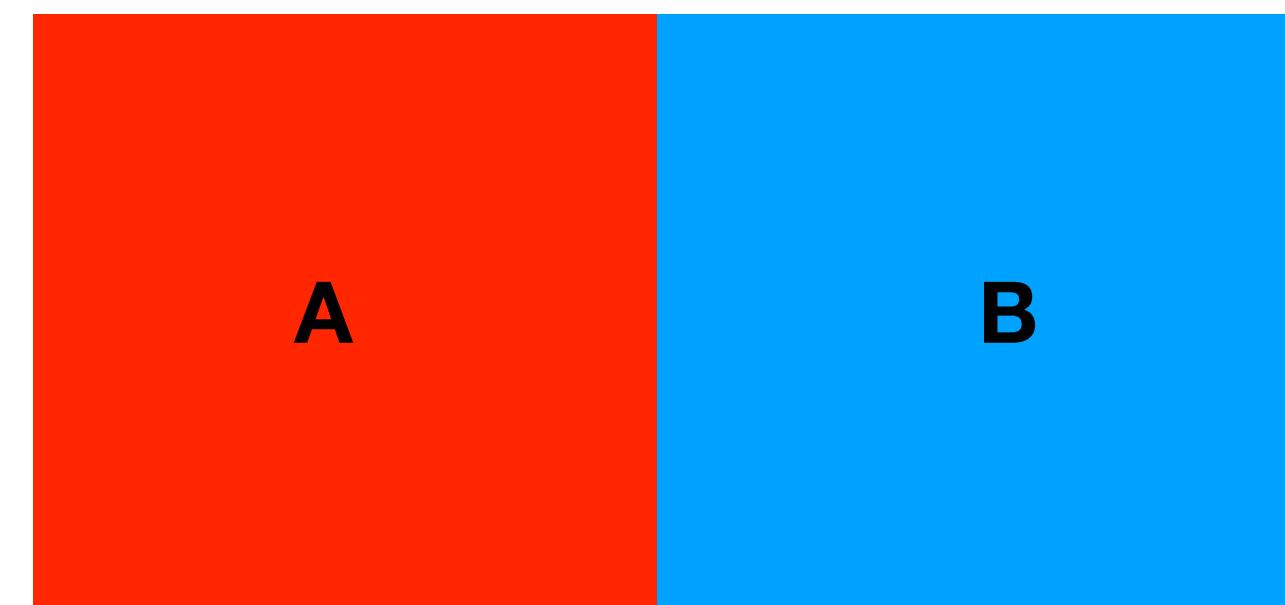
If Gravity is
QUANTUM !

We can test it !



Only Entropy is Entanglement Entropy

$$|\Psi\rangle = \sum_{i=1}^N C_i \psi_i^A \otimes \psi_i^B$$

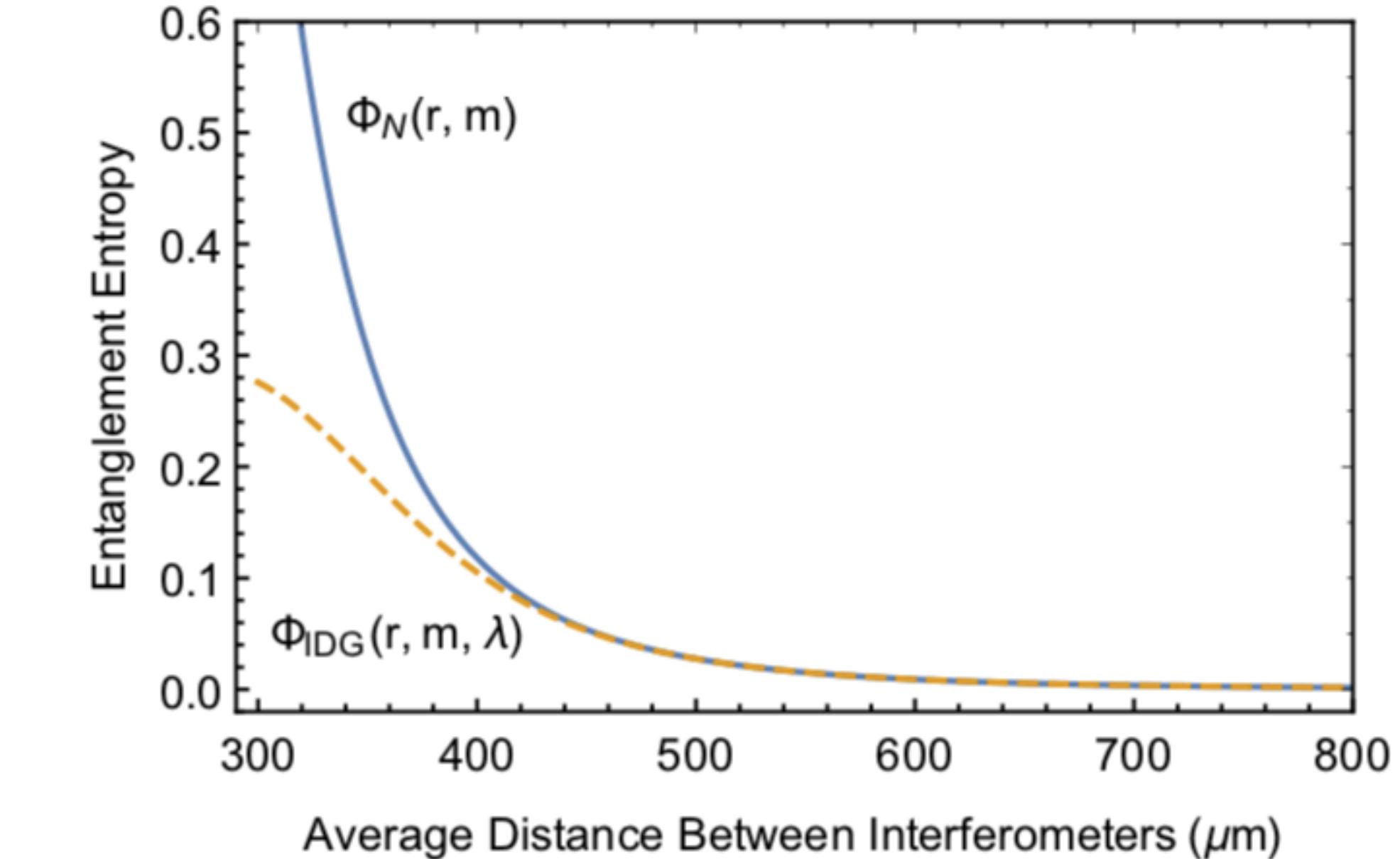
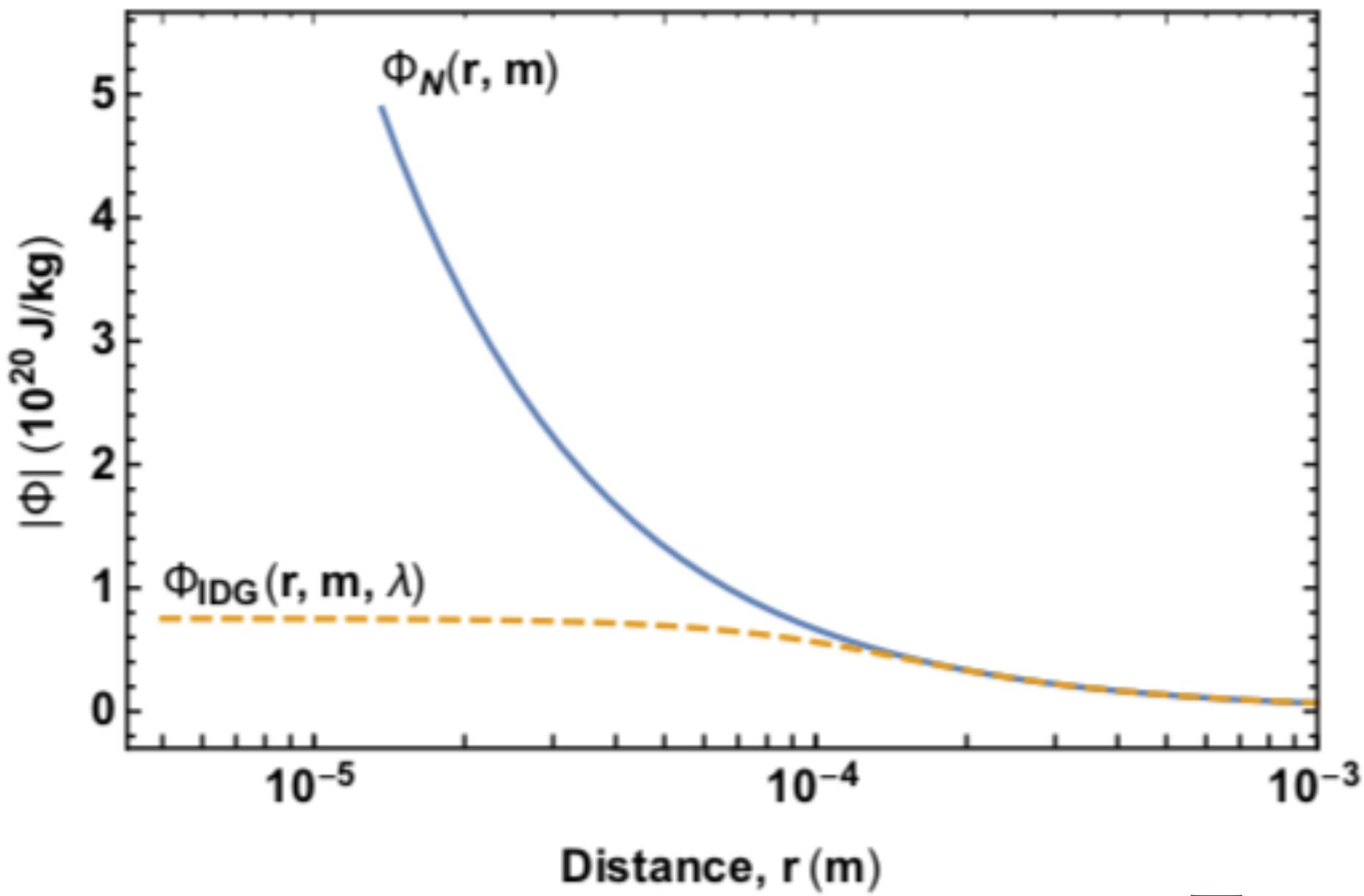


We can trace out the degrees of freedom of 'B' to get a reduced density matrix describing 'A'

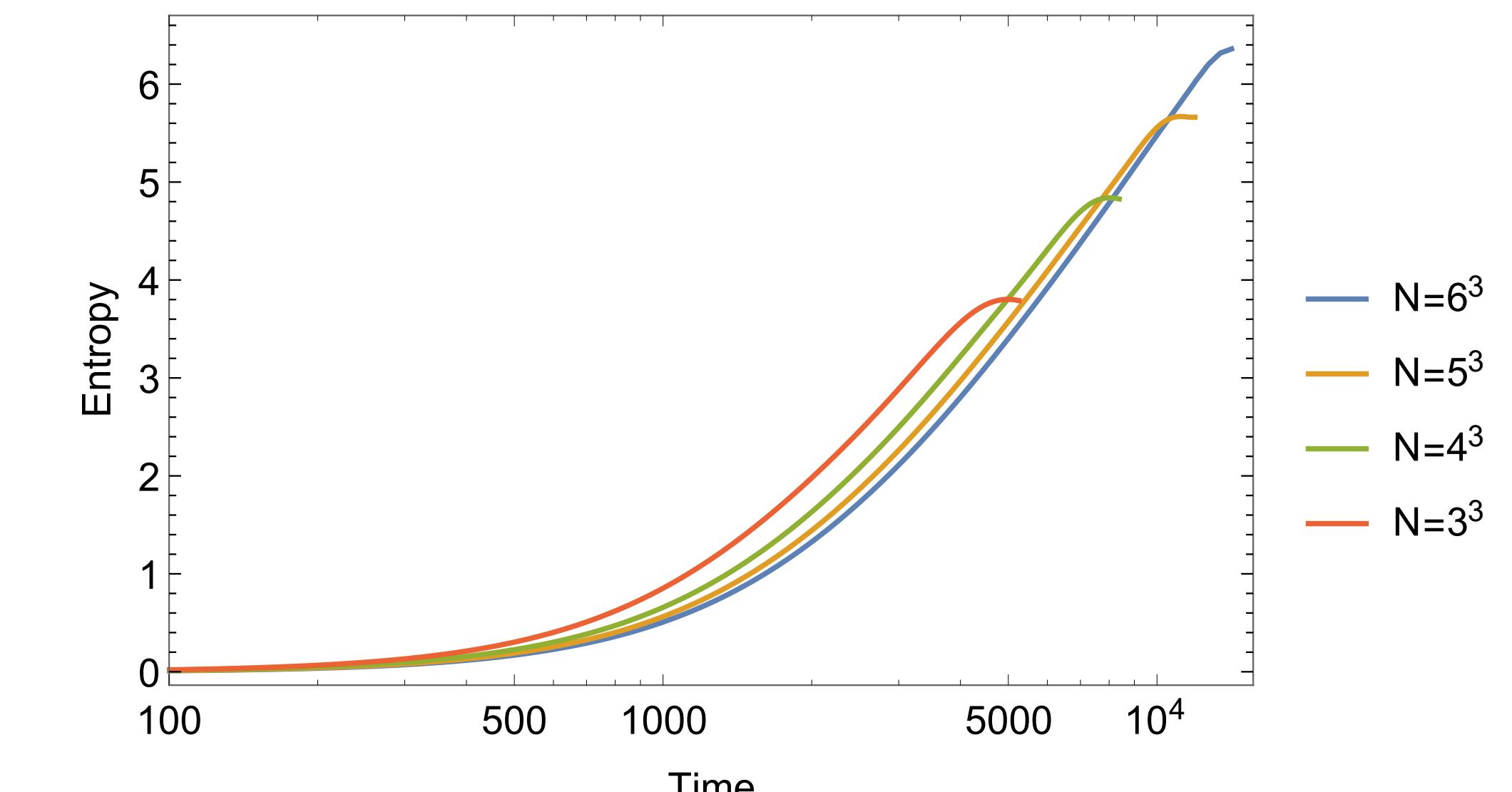
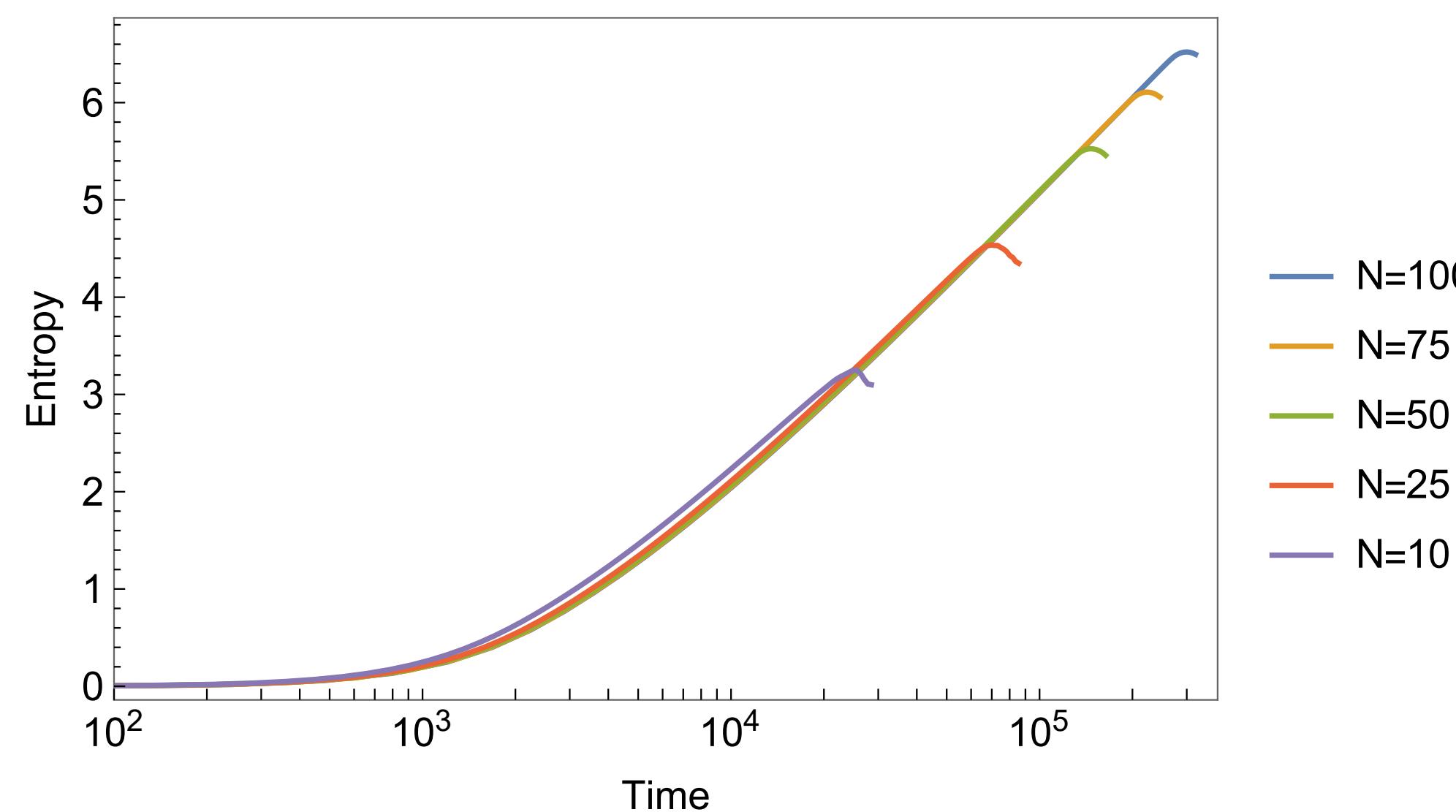
$$\rho_A = \sum_{i=1}^N |C_i|^2 \psi_i^A \rangle \langle \psi_i^A| \quad S(A) = -Tr[\rho_A \ln \rho_A] = \sum_{i=1}^N |C_i|^2 \ln |C_i|^2$$

$$S(A) + S(B) = 0$$

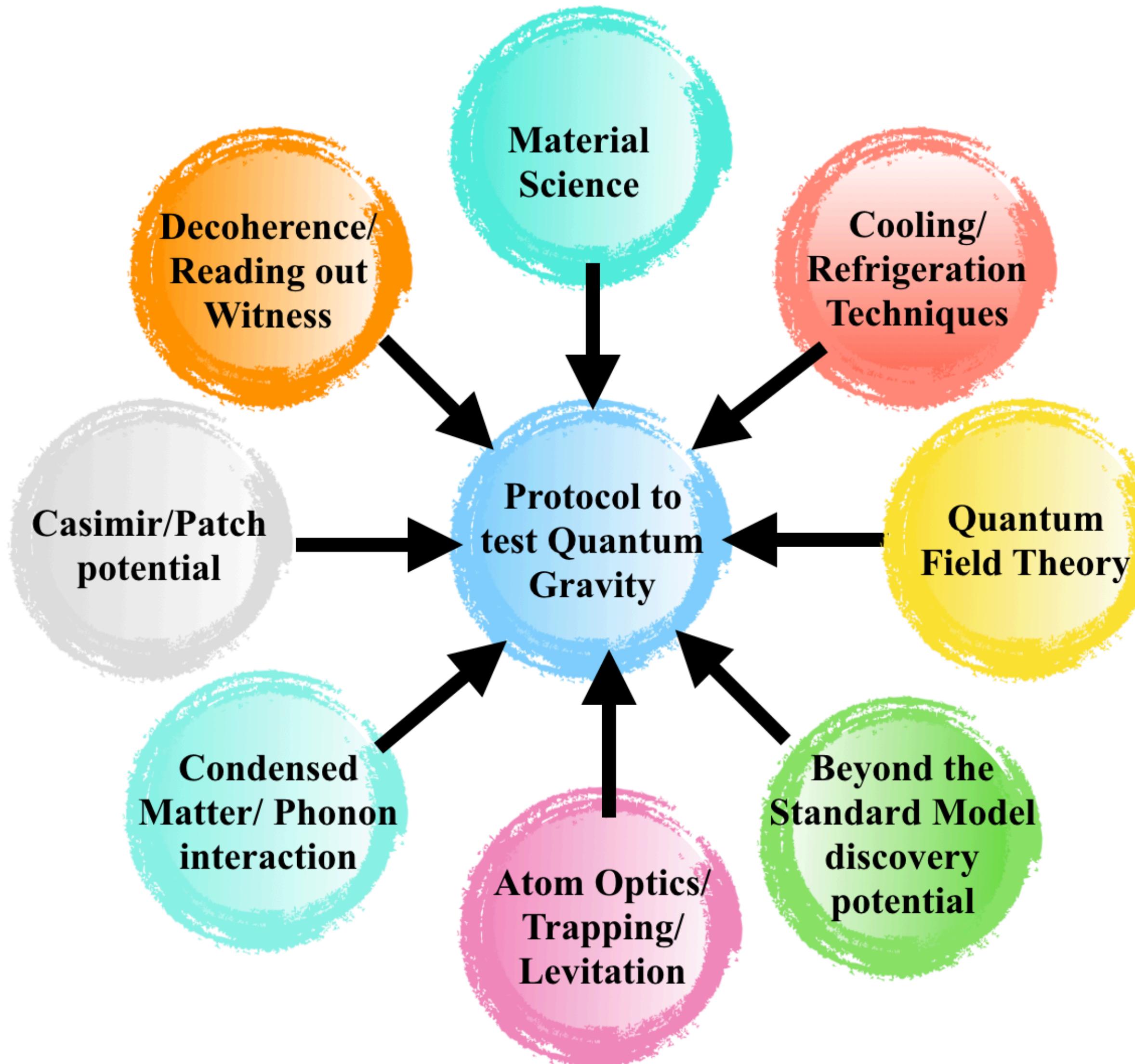
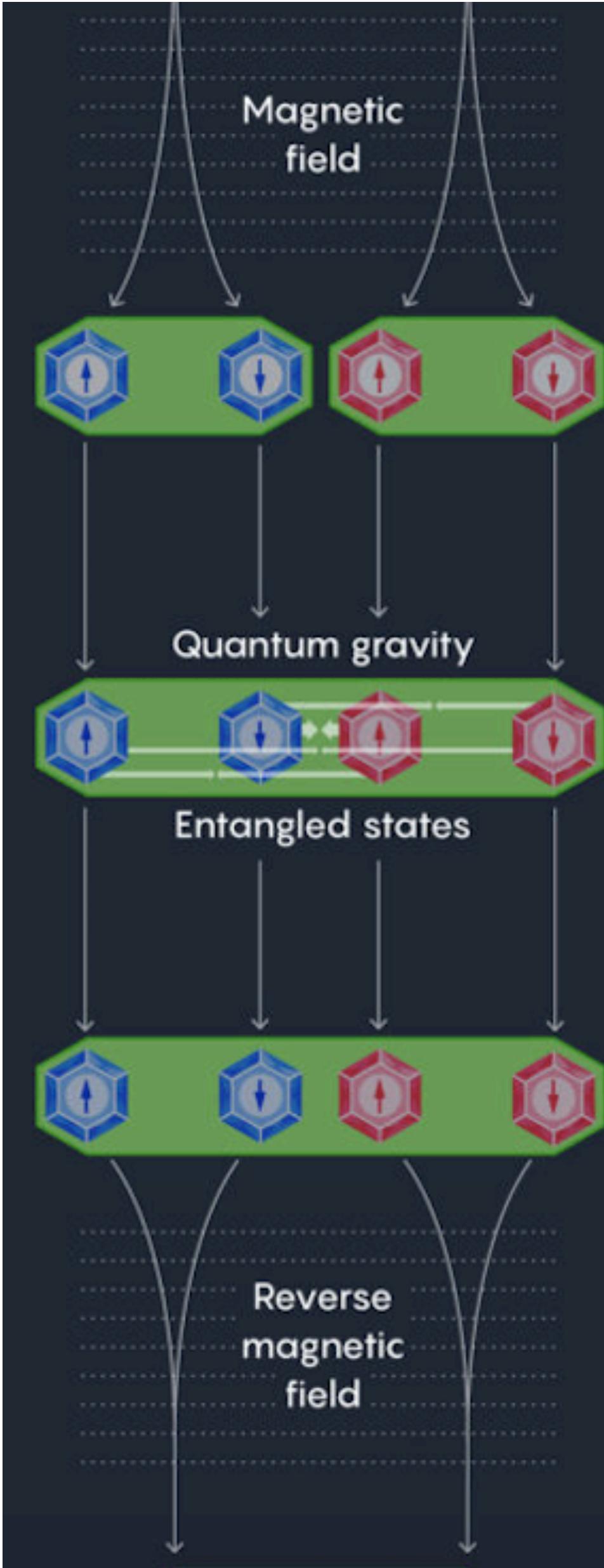
Entanglement Entropy & it's Universal Growth



Two Self-Gravitating systems



QGEM Protocol Unites All



- Foundations Of QM
- Foundations of QFT
- Foundations of Gravity
- Penrose's Collapse Model
- Gravitational Waves
- Axion/Hidden sector

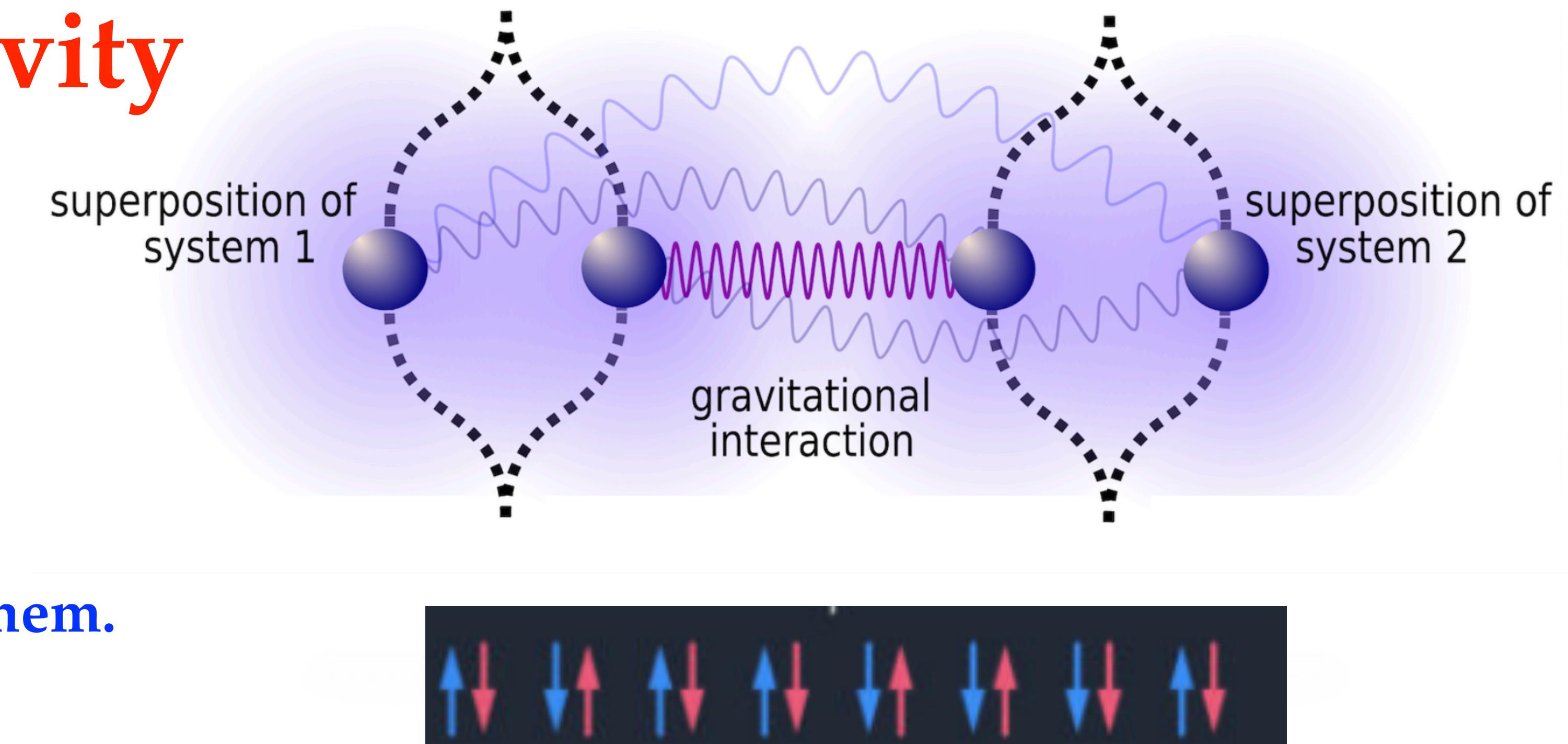
Conclusion & Future Outlook

Testing Quantum Gravity

The Final Frontier !

**One Protocol to rule them all,
One Protocol to find them,
One Protocol to bring them all,
and Quantum Gravity which binds them.**

We can test QGEM Protocol



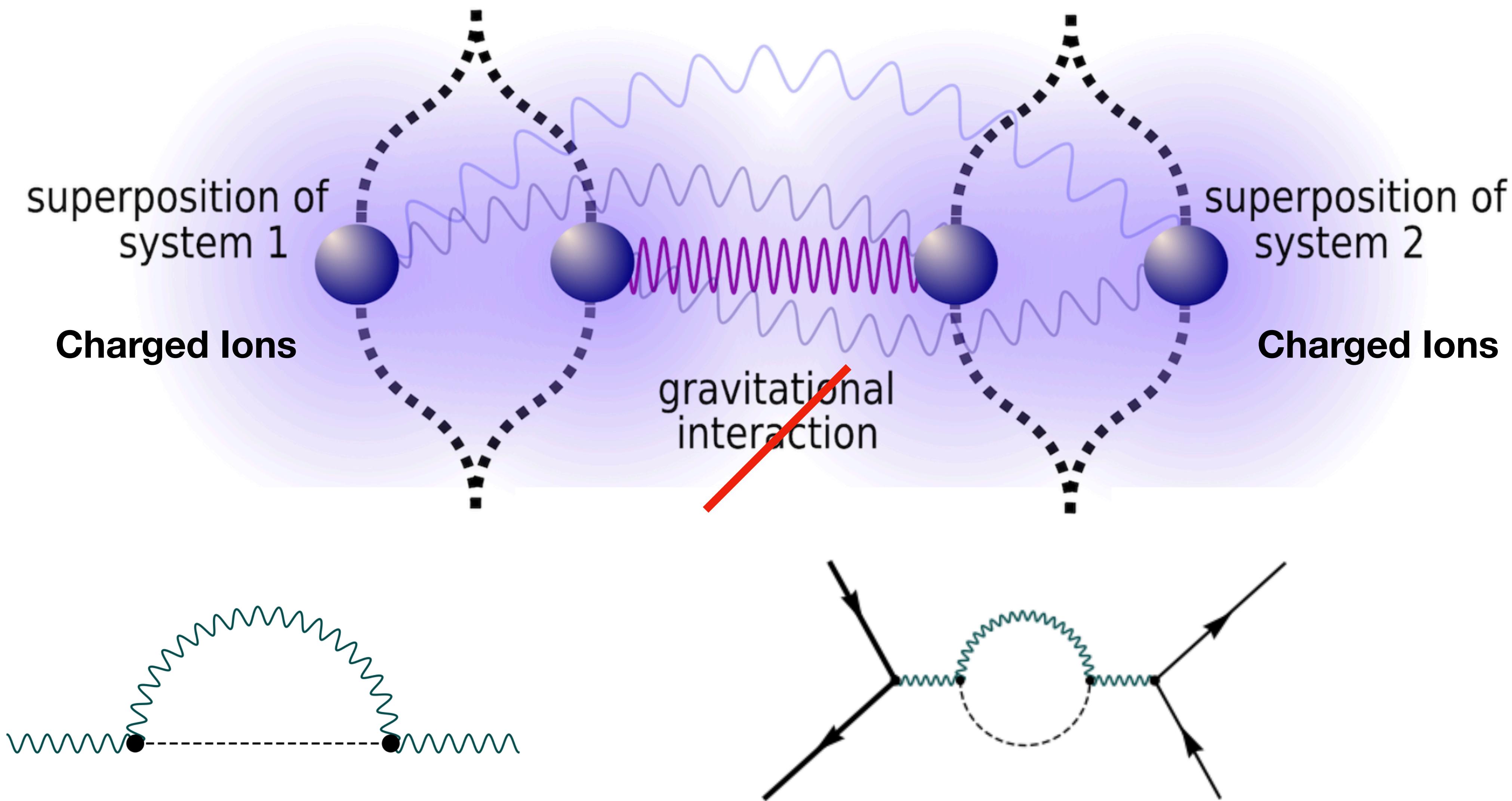
**Bose+AM+Morley+Ulbricht+Toros+Paternostro+Geraci+Barker+Kim+
Milburn, PRL (2017) [1707.06050]**

Marshman+AM+Bose, PRA (2020) [1907.01568]

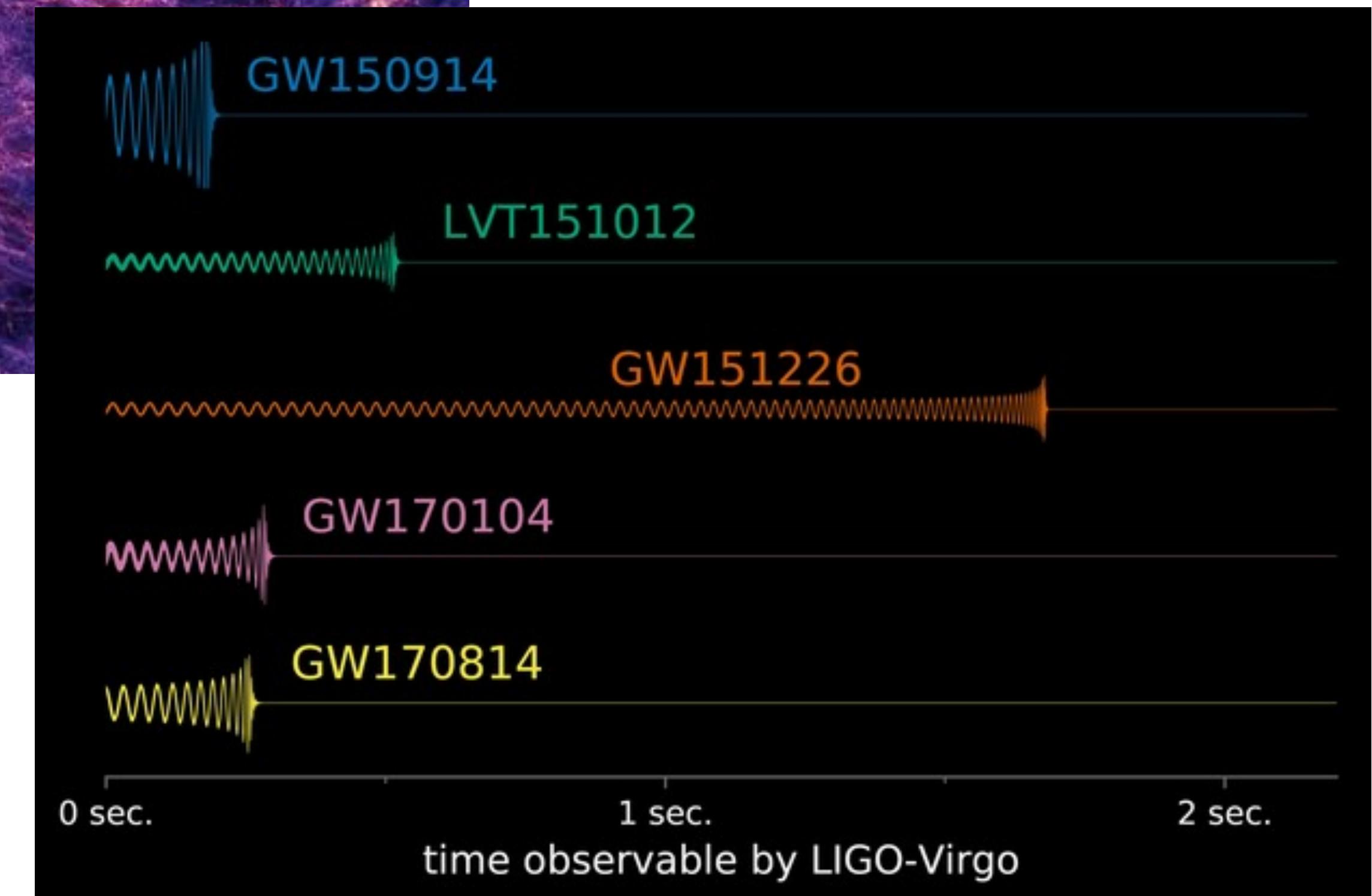
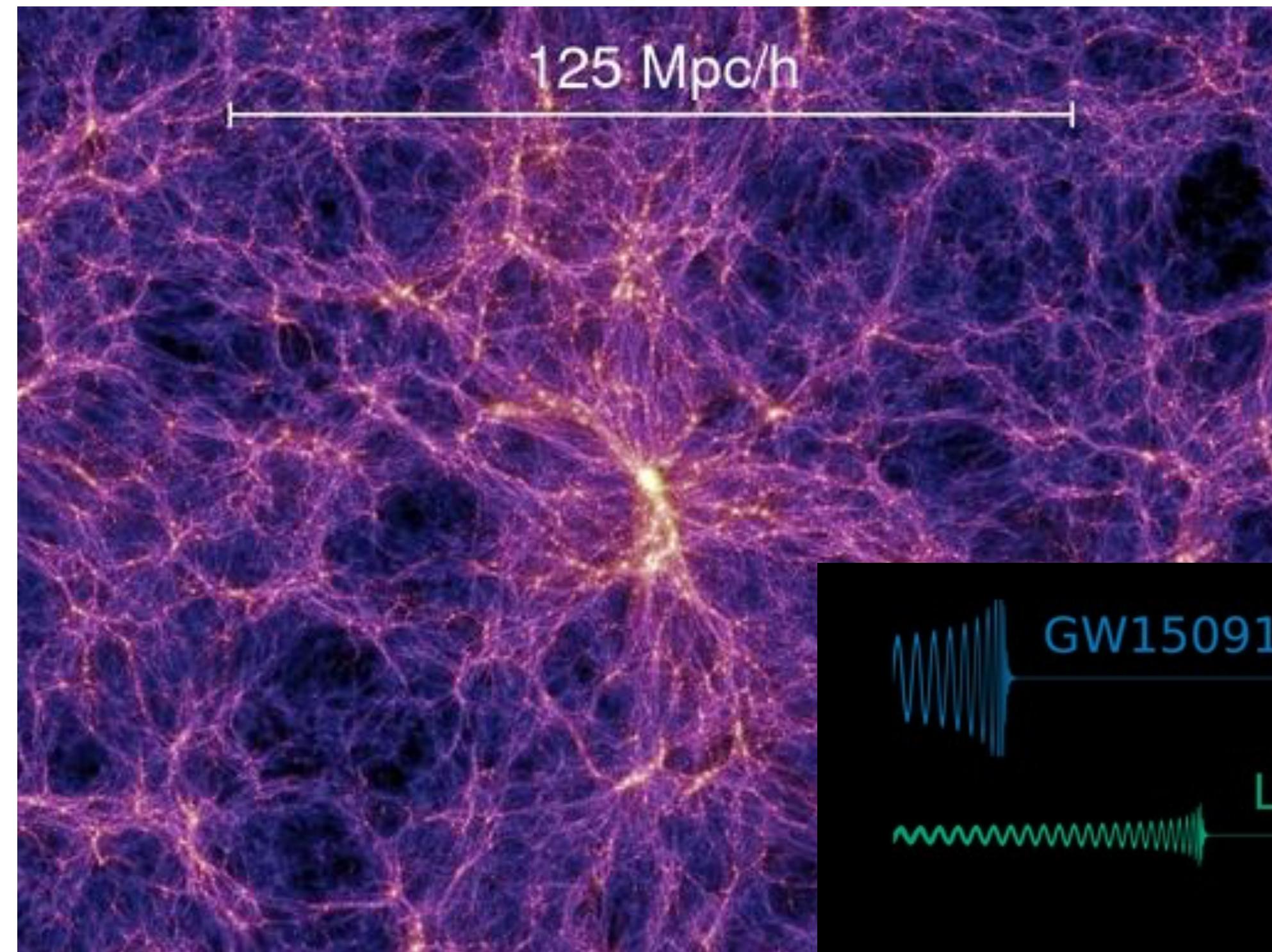
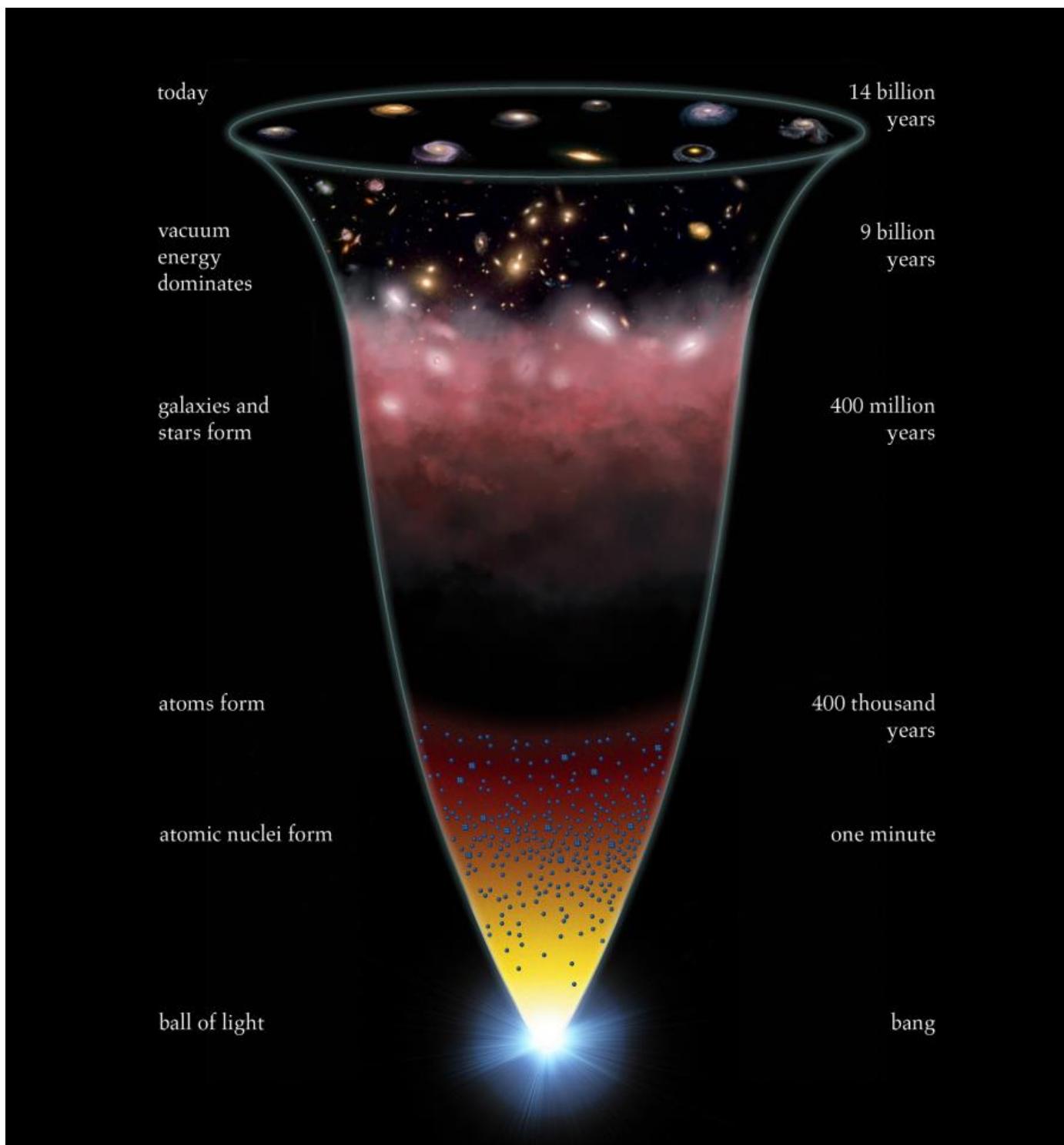
Extra slides

Probing Beyond the Standard Model Physics

Axion or Hidden Photon



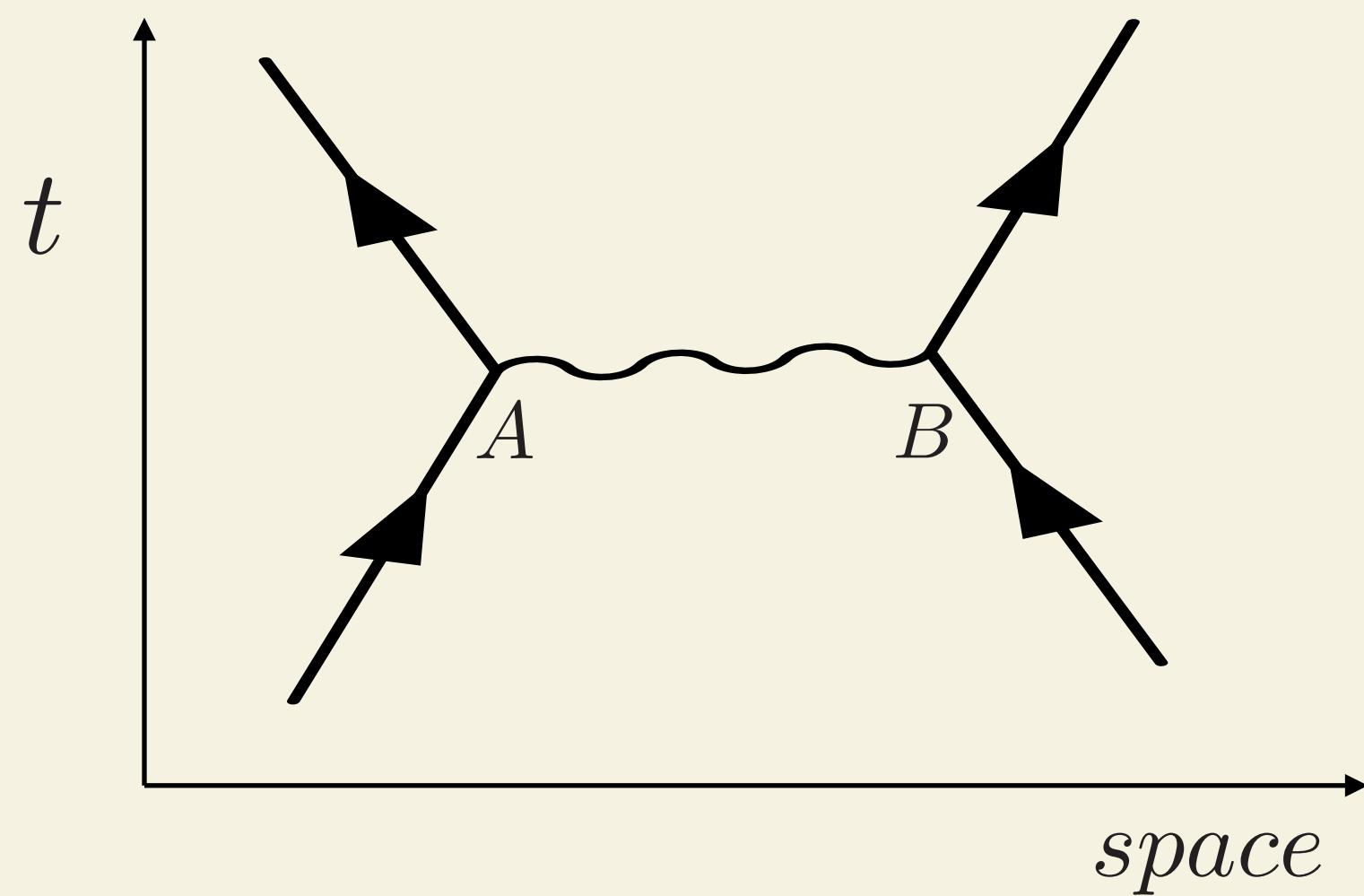
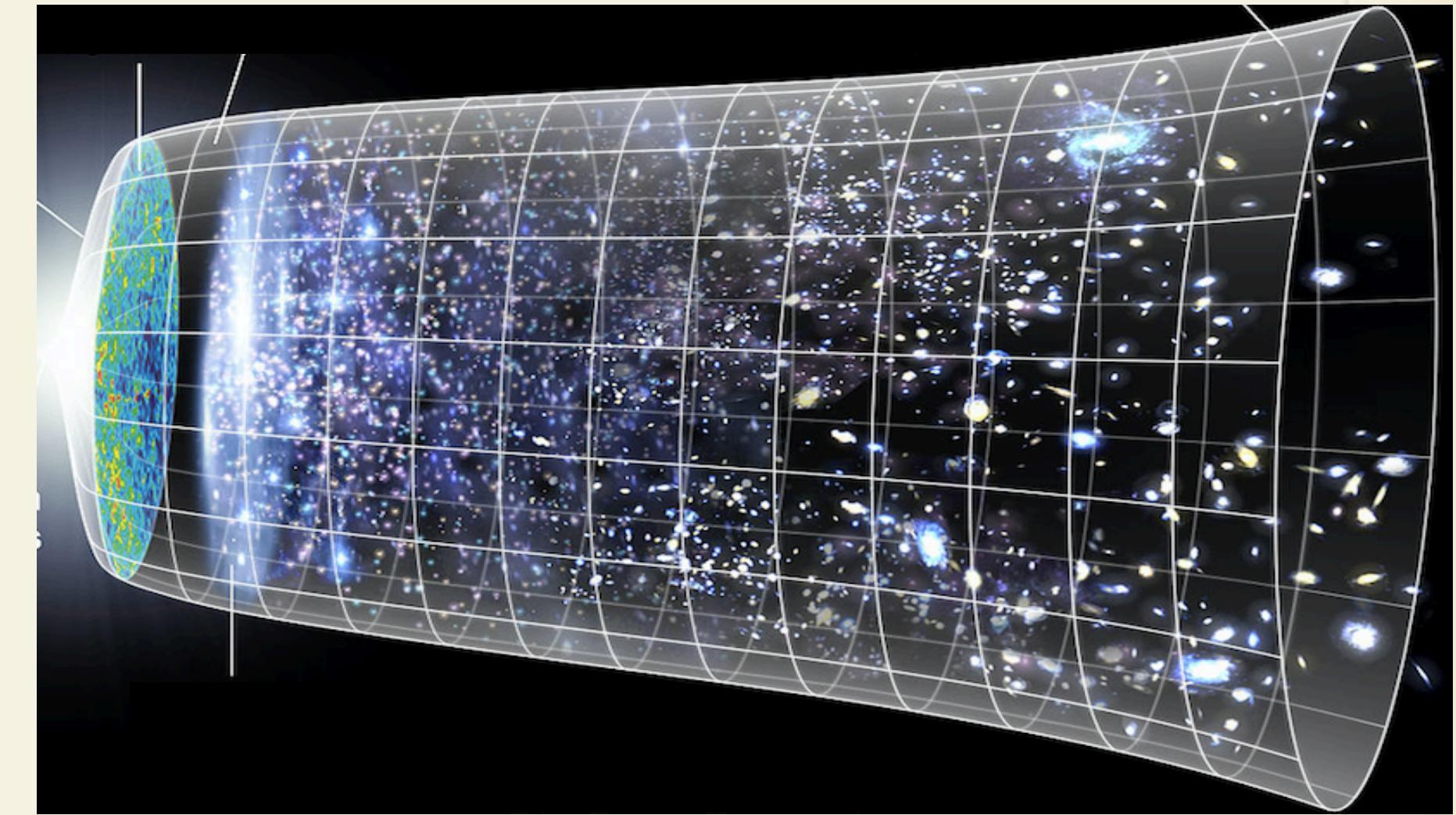
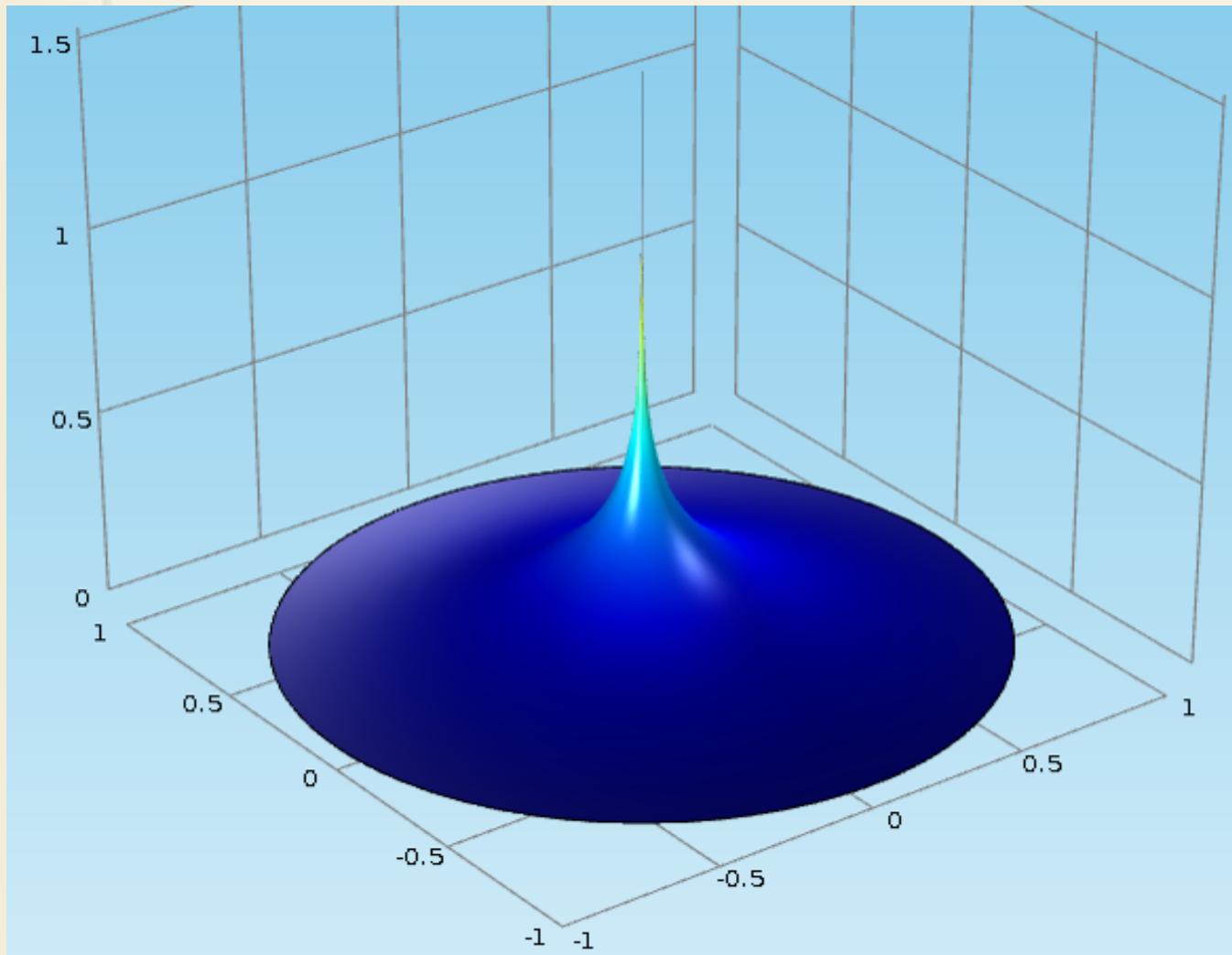
Gravity is Universal & Successful



In the simplest case: Minimal Coupling with matter

$$\kappa h^{\mu\nu} T_{\mu\nu}$$

Cosmological & Black hole Singularities



$$V \sim \frac{1}{r}$$

**Graviton or Photon
(mediator is massless)**

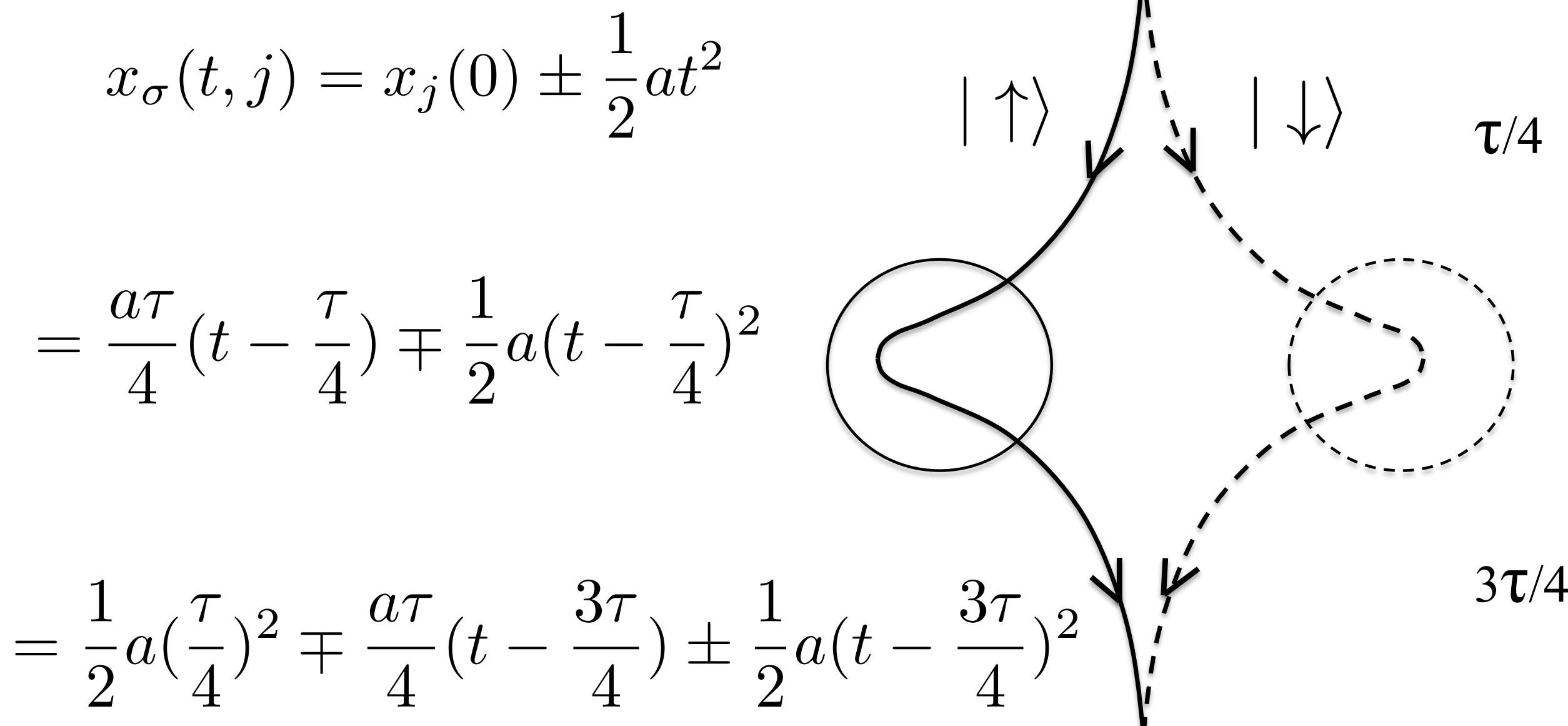
Conservation of Probability will not allow singularities to develop,

Non-local effects will forbid singularities to form

Macroscopic Spatial Superposition

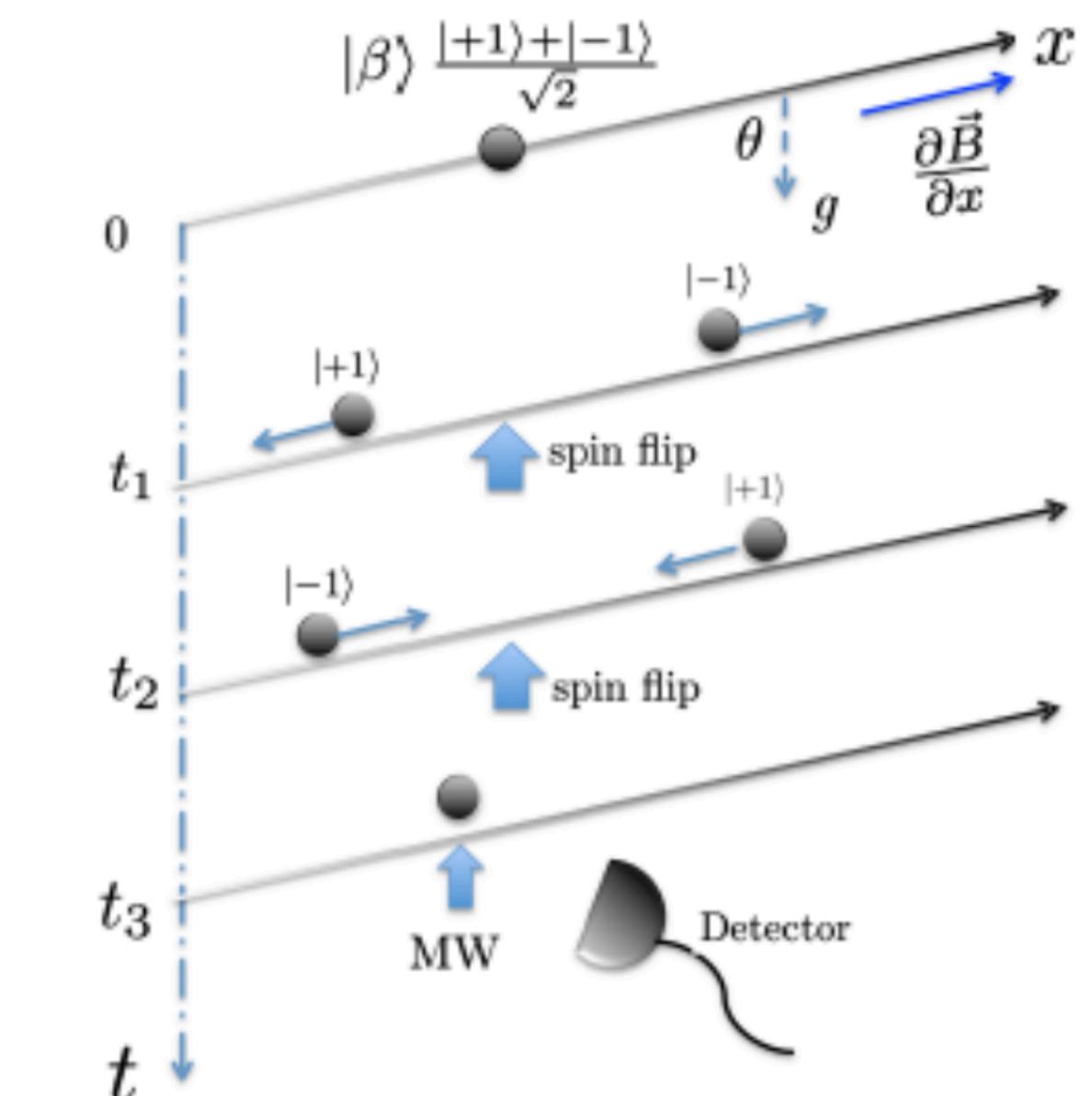
How can we increase the scale of the superposition?

Free particle in an inhomogeneous magnetic field (acceleration $+a$ or $-a$)



100 micron separation for 1 sec

$$\Delta x = 2s = 2\frac{g\mu_B}{m\hbar} S_x \partial_x B (\tau/2)^2 = \frac{g\mu_B \partial_x B}{2m} \tau^2$$



M. Scala, M. S. Kim, G. W. Morley, P. F. Barker, S. Bose, PRL. **111**, 180403 (2013)

J. S. Pedernales, G. W. Morley, and M. B. Plenio:1906.00835 (2019)

Van de Kamp+Marshman +Bose+Mazumdar [2006.06931 [quant-ph]]

Probing UV Gravity

(1) GR: $\lim_{k^2 \rightarrow 0} \Pi = (\mathcal{P}^2/k^2) - (\mathcal{P}_s^0/2k^2) \equiv \Pi_{GR}$

(2) F(R) Gravity:

$$\mathcal{L}(R) = \mathcal{L}(0) + \mathcal{L}'(0)R + \frac{1}{2}\mathcal{L}''(0)R^2 + \dots$$

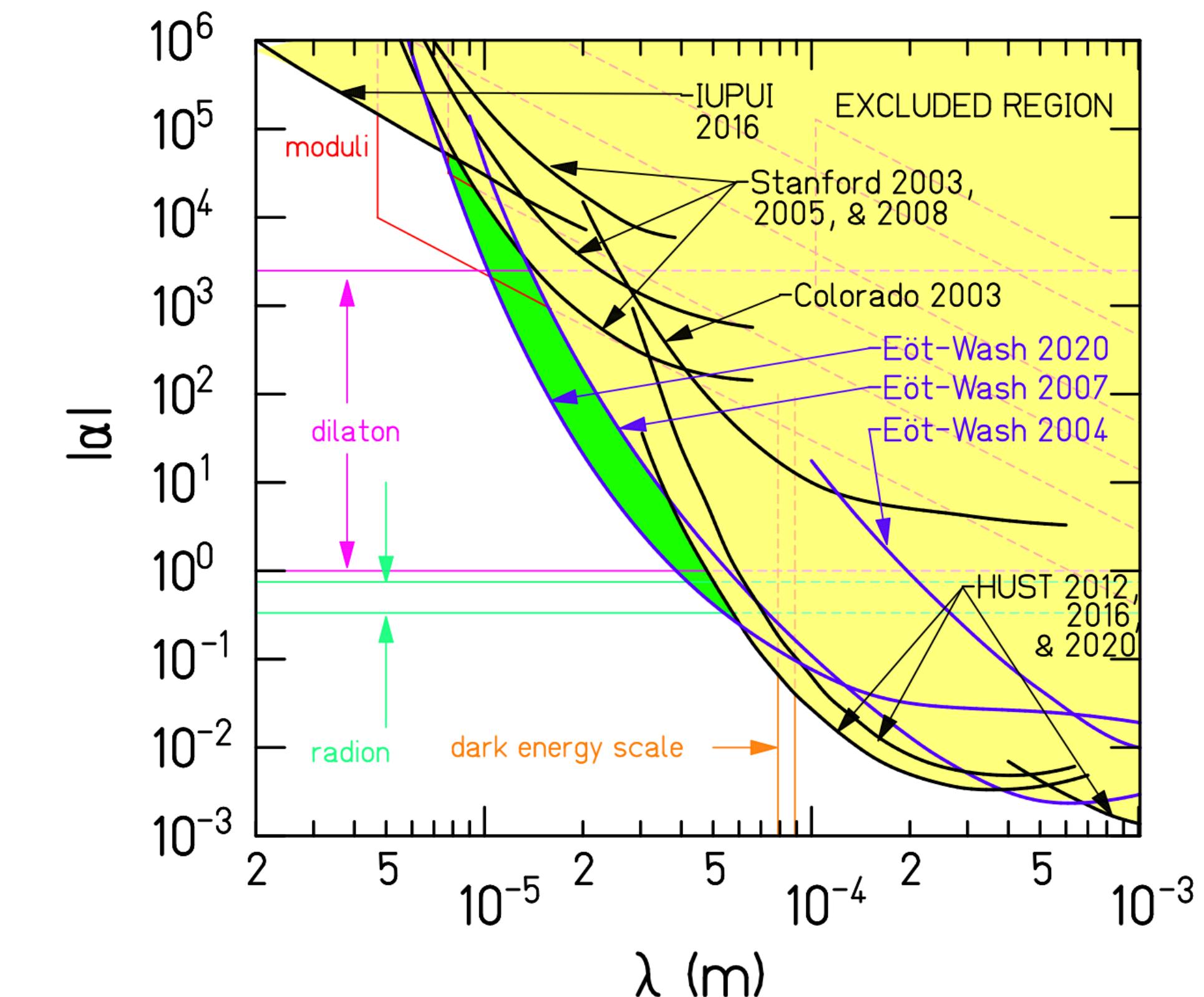
$$\Pi = \Pi_{GR} + \frac{1}{2} \frac{\mathcal{P}_s^0}{k^2 + m^2}, \quad m^2 = \frac{1}{3\mathcal{L}''(0)}$$

(3) Weyl Gravity: $\mathcal{L} = R - \frac{1}{m^2}C^2$

$$C^2 = R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} - 2R_{\mu\nu}R^{\mu\nu} + \frac{1}{3}R^2$$

$$\Pi = \frac{\mathcal{P}^2}{k^2(1-(k/m)^2)} - \frac{\mathcal{P}_s^0}{2k^2} = \Pi_{GR} - \frac{\mathcal{P}^2}{k^2+m^2}$$

P. Van Nieuwenhuizen, Nucl. Phys. B60, 478-492 (1973)
 T. Biswas, T. Koivisto and A. Mazumdar,
 "Nonlocal theories of gravity: the flat space propagator,"
 arXiv:1302.0532 [gr-qc]



$$V \sim \frac{1}{r} e^{-mr} + \dots$$

