# Axiogenesis

Raymond Co

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Online "Newton 1665" Seminars April 6th 2020

Collaborators:

arXiv: 1910.02080 Keisuke Harigaya Phys. Rev. Lett. 124, 111602 (2020) arXiv: 1910.14152 Lawrence Hall, Keisuke Harigaya



### 1998 Nobel Prize in Physics "for their discovery of a new form of quantum fluid with fractionally charged excitations".

Horst Ludwig Störmer

born on April 6th

1949

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https://en.wikipedia.org/wiki/Horst\_Ludwig\_St%C3%B6rmer

#### - QCD axion

(1)

(2)

#### (0) Misalignment mechanism

Preskill, Wise, Wilczek 1983, Abbott, Sikivie 1983, Dine, Fischler 1983

Possible initial <u>condit</u>ions

(1

 $(\circ)$ 

×

(2)

(1')

#### Peccei-Quinn symmetry

### • QCD axion

### (0) Misalignment mechanism Preskill, Wise, Wilczek 1983, Abbott, Sikivie 1983, Dine, Fischler 1983 (1) Parametric resonance RC, L. Hall, K. Harigaya 2017 K. Harigaya, J. Leedom 2019 (2) - Kinetic misalignment mechanism RC, L. Hall, K. Harigaya 2019 - Axiogenesis RC, K. Harigaya 2019

Possible initial conditions

(0)

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#### Peccei-Quinn symmetry

Dark Matter

Baryon Asymmetry

### • QCD axion



This talk

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(0)

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Peccei-Quinn symmetry

# Experimental Searches of QCD Axions

	CASPEr	$0 - 10^{-9} eV$	D. Budker et al. 1306.6089
	ABRACADABRA	$10^{-9} - 10^{-6} \text{ eV}$	Y. Kahn et al. 1602,01086
	ADMX	$10^{-6} - 10^{-3} \text{ eV}$	P. Sikivie 1983 N. Du et al. 1804.05750
	IAXO	$10^{-3} - 1 \text{ eV}$	J. K. Vogel et al. 1302,3273 E. Armengaud et al. 1401,3233
	ARIANDE	$10^{-6} - 10^{-2} \text{ eV}$	A. Arvanitaki et al. 1403,1290 A. A. Geraci et al. 1401,3233
	Orpheus	$10^{-5} - 10^{-3} eV$	G. Rybka et al. 1403.3121
	MADMAX	$10^{-5} - 10^{-4} \text{ eV}$	A. Caldwell et al. 1611.05865
	TASTE	$10^{-3} - 1 \text{ eV}$	V. Anastassopoulos et al. 1706.09378
Multilayer optical	haloscopes	0,1 - 10 eV	M. Baryakhtar, J. Huang, R. Lasenby 1803.11455
Resonant absorption	in molecule	s 0.2 - 20 eV	A. Arvanitaki, S. Dimopoulos, K. Van Tilburg 1709.05354

# Experimental Searches of QCD Axions

Axiogenesis favors 10-3 - 1 eV	CASPEr	$0 - 10^{-9} \text{ eV}$	D. Budker et al. 1306.6089
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# Neutron EDM



 $|d_n| \sim e \times d \sim 10^{-13} \text{ e} \cdot \text{cm}$ 

 $10^{-13}$  cm

# strong CP Problem



 $\mathcal{L} \supset \bar{\theta} \frac{g_s^2}{32\pi^2} G_b^{\mu\nu} \widetilde{G}_{b\mu\nu}$  $|d_n| \simeq 3 \times 10^{-16} \bar{\theta} \,\mathrm{e} \cdot \mathrm{cm}$ 

 $|d_n| \sim e \times d \sim 10^{-13} \text{ e} \cdot \text{cm}$ 

# strong CP Problem



 $|d_n| \sim e \times d \sim 10^{-13} \text{ e} \cdot \text{cm}$  $\mathcal{L} \supset \bar{\theta} \frac{g_s^2}{32\pi^2} G_b^{\mu\nu} \widetilde{G}_{b\mu\nu}$  $|d_n| \simeq 3 \times 10^{-16} \overline{\theta} \,\mathrm{e} \cdot \mathrm{cm}$  $|d_n| \lesssim 3 \times 10^{-26} \text{ e} \cdot \text{cm}$ Experiments:  $|\bar{\theta} < 10^{-10}|$ 

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Pendlebury et al. 2015, Bake et al. 2016, Abel et al. 2020

# strong CP Problem

 $\mathcal{L} \supset \bar{\theta} \frac{g_s^2}{32\pi^2} G_b^{\mu\nu} \widetilde{G}_{b\mu\nu}$ 

 $\mathcal{L} \supset \bar{\theta} \frac{g_s^2}{32\pi^2} G_b^{\mu\nu} \widetilde{G}_{b\mu\nu} + \frac{a}{f_a} \frac{g_s^2}{32\pi^2} G_b^{\mu\nu} \widetilde{G}_{b\mu\nu}$ 

Peccei-Quinn 1977

# Axions



# Axions



### • QCD axion



This talk

(0) Misalignment mechanism Preskill, Wise, Wilczek 1983, Abbott, Sikivie 1983, Dine, Fischler 1983
(1) Parametric resonance RC, L. Hall, K. Harigaya 2017 K. Harigaya, J. Leedom 2019
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} Baryon Asymmetry

Dark Matter

Possible initial conditions

(0)

(2)

Peccei-Quinn symmetry

# StoryLine

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### QCD axion



Misalignment mechanism
Preskill, Wise, Wilczek 1983, Abbott, Sikivie 1983, Dine, Fischler 19
Parametric resonance
RC, L. Hall, K. Harigaya 2017 K. Harigaya, J. Leedom 2019
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RC, L. Hall, K. Harigaya 2019
- Axiogenesis R
RC, K. Harigaya 2019
This talk

Possible initial conditions

(0)

×

(2)

#### Peccei-Quinn symmetry

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# Misalignment Mechanism: Scalars

$$\left(\partial_t^2 + 3H\partial_t + m_\phi^2\right)\phi = 0$$

Early time  $H \gg m_{\phi}$ 

Hubble friction dominates

 $\rho_{\phi} = m_{\phi}^2 \phi^2$ 

Energy density

Except for long inflation: P. Graham et al. 1805,07362 F. Takahashi et al. 1805,08763

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 $\phi = \text{constant}$ 

Field value is "stuck"

 $ho_{\phi} = ext{constant}$ 

is also "stuck"

Late time

 $m_{\phi} \gg H$ 

Oscillations begin

$$\rho_{\phi} = m_{\phi}^2 \phi^2$$

Energy density

 $\phi \propto a^{-rac{3}{2}}$ 

Field value redshifts

$$ho_\phi \propto a^{-3}$$

scales like matter

Preskill, Wise, Wilczek 1983 Abbolt, Sikivie 1983 Dine, Fischler 1983

### Misalignment Mechanism: Axions



# Axion Misalignment Mechanism

H.

 $\Omega_a h^2 \approx 0.11 \left(\frac{f_a}{5 \times 10^{11} \text{GeV}}\right)^{1.184} F \bar{\Theta}_i^2$ 

# Axion Misalignment Mechanism



# New Production Mechanisms

### Misalignment Mechanism

- QCD axions: misalignment driven to the hilltop/bottom arXiv:1812.11186, 1812.11192 RC, E. Gonzalez, K. Harigaya - QCD axions: kinetic misalignment mechanism

> arXiv:1910.14152 RC, L. Hall, K. Harigaya arXiv:2004.00629 RC, L. Hall, K. Harigaya, K. Olive, S. Verner

### Exponential Particle Production

- QCD axions: parametric resonance

arXiv:1711.10486 RC, L. Hall, K. Harigaya

- Dark photons: tachyonic instability

arXiv:1810.07196 RC, A. Pierce, Z. Zhang, Y. Zhao

# New Production Mechanisms

### Misalignment Mechanism

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This talk

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# Rotation



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# Why Robation?

Large field value : Inflaton coupling  $V(|P|) \sim -H_I^2 |P|^2 + \frac{|P|^{2d}}{M^{2d-4}}$ 

Angular motion : Explicit PQ breaking

 $V(P) \sim \frac{P^n}{M^{n-4}} + \text{h.c.}$ 

expected from quantum gravity or PQ as an accidental symmetry

S. Giddings et al. 1988, S. Coleman 1988, G. Gilbert 1988, D. Harlow et al. 2019 R. Holman 1992, S. Barr 1992, M. Kamiokowski 1992, D. Dine 1992

Dynamics analogous to that in Affleck-Dine baryogenesis

I. Affleck and M. Dine 1991

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PRL 124, 111602 (2020) RC and K. Harigaya

Initial condition

 $P = \frac{S + f_a}{\sqrt{2}} e^{i\frac{a}{f_a}}$ 

# Asymmetry of PQ Charge

Noether charge associated with the shift symmetry

 $n_{\rm PQ} = iP\dot{P^*} - iP^*\dot{P}$  $n_{\rm PQ} = S^2\dot{\theta}$ 

PQ asymmetry PQ charge density = Rotation of PQ field

PQ charge is conserved soon after the onset.

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PRL 124, 111602 (2020) RC and K. Harigaya

 $P = \frac{S + f_a}{\sqrt{2}} e^{i\frac{a}{f_a}}$ 

# PQ Field Evolution

# Redshift $n_{\rm PQ} = S^2 \dot{\theta}$ Thermalization

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} Baryon Asymmetry

Dark Matter

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Peccei-Quinn symmetry

### QCD axion



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(2) - Kinetic misalignment mechanism

RC, L. Hall, K. Harigaya 2019

- Axiogenesis

RC, K. Harigaya 2019

This talk

Possible initial conditions

(0)

X

(2)

#### Peccei-Quinn symmetry

Dark Matter

Baryon Asymmetry



# Kinetic Misalignment Mechanism

(Misalignment + non-zero kinetic energy)

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# Kinetic Misalignment Mechanism

(Misalignment + non-zero kinetic energy)

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# V(0) *Hisalignment Mechanism*

A

Reason:

Why such a large angular speed?

 $n_{\rm PQ} = S^2 \dot{\theta}$   $n_{\rm PQ} R^3 = \text{conserved charge}$ 

Conventional:

 $S^2 = f_a^2$   $\dot{ heta} \propto R^{-3}$ matrix  $S^2 \propto R^{-3}$   $\dot{ heta} = {
m constant}$  f since  $S^2 \propto R^{-2}$   $\dot{ heta} \propto R^{-1}$  f since  $S^2 \propto R^{-2}$   $\dot{ heta} \propto R^{-1}$ 

Slower redshift!

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### QCD axion



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RC, K. Harigaya 2019

This talk

Possible initial conditions

(0)

X

(2)

#### Peccei-Quinn symmetry

Dark Matter

Baryon Asymmetry

# Axiogenesis

(QCD axion + baryogenesis)

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#Космос, #Наука

For T > TEW,



$$\mathcal{L} = \frac{a}{f_a} \frac{g_s^2}{32\pi^2} G\widetilde{G}$$

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For T > TEW,



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For T > TEW,



 $n_B = c_B \dot{\theta} T^2$ 

 $c_B \simeq 0.1 - 0.15 \, c_W$ 

PRL 124, 111602 (2020) RC and K. Harigaya

For T < TEW, the baryon asymmetry is frozen



 $n_B = c_B \dot{\theta} T^2$ 

 $c_B \simeq 0.1 - 0.15 \, c_W$ 

PRL 124, 111602 (2020) RC and K. Harigaya

At T < TEW, the baryon asymmetry is frozen

# $Y_B \equiv \frac{n_B}{s} \simeq 2 \times 10^{-3} \left(\frac{c_B}{0.1}\right) \frac{\theta(T_{\rm EW})}{T_{\rm EW}}$

Baryon asymmetry fixes rotational speed

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### QCD axion



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RC, K. Harigaya 2019

Dark Matter

Baryon Asymmetry Possible initial conditions

(0)

X

(2)

Peccei-Quinn symmetry



arXiv:1910.02080 RC and K. Harigaya arXiv:1910.14152 RC, L. Hall, and K. Harigaya



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arXiv:1910.02080 RC and K. Harigaya arXiv:1910.14152 RC, L. Hall, and K. Harigaya

see talk

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fa ~

 $\frac{\Omega_a h^2}{\Omega_{\rm DM} h^2} \simeq \Pr_{\rm talk}^{\rm see} \left(\frac{f_a}{10^8 \ {\rm GeV}}\right) \left(\frac{130 \ {\rm GeV}}{T_{\rm EW}}\right)^2 \left(\frac{0.1}{c_B}\right)$ 





see talk

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 $\frac{\Omega_a h^2}{\Omega_{\rm DM} h^2} \simeq \frac{\sec}{\tan k} \left( \frac{f_a}{10^8 \text{ GeV}} \right) \left( \frac{130 \text{ GeV}}{T_{\rm EW}} \right)^2 \left( \frac{0.1}{c_B} \right)$ 

CB ~ see talk SO CW ~ see talk ?



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### Conclusions

Axiogenesis allows the QCD axion to simultaneously explain

- ~ the strong CP problem
- V the dark matter abundance

v the baryon asymmetry

- ✓ Possible signatures:
  - ✓ QCD axion searches
  - ✓ TeV-scale new particles
  - ✓ dark matter ultracompact minihalos

New model building opportunities

