

By the end of 2024, the Advanced Virgo+ gravitational wave interferometer (AdVirgo+) will undergo a major upgrade, called Phase II and aimed at a reduction of thermal noise and an extension of the observing horizon to more than 150 Mpc.

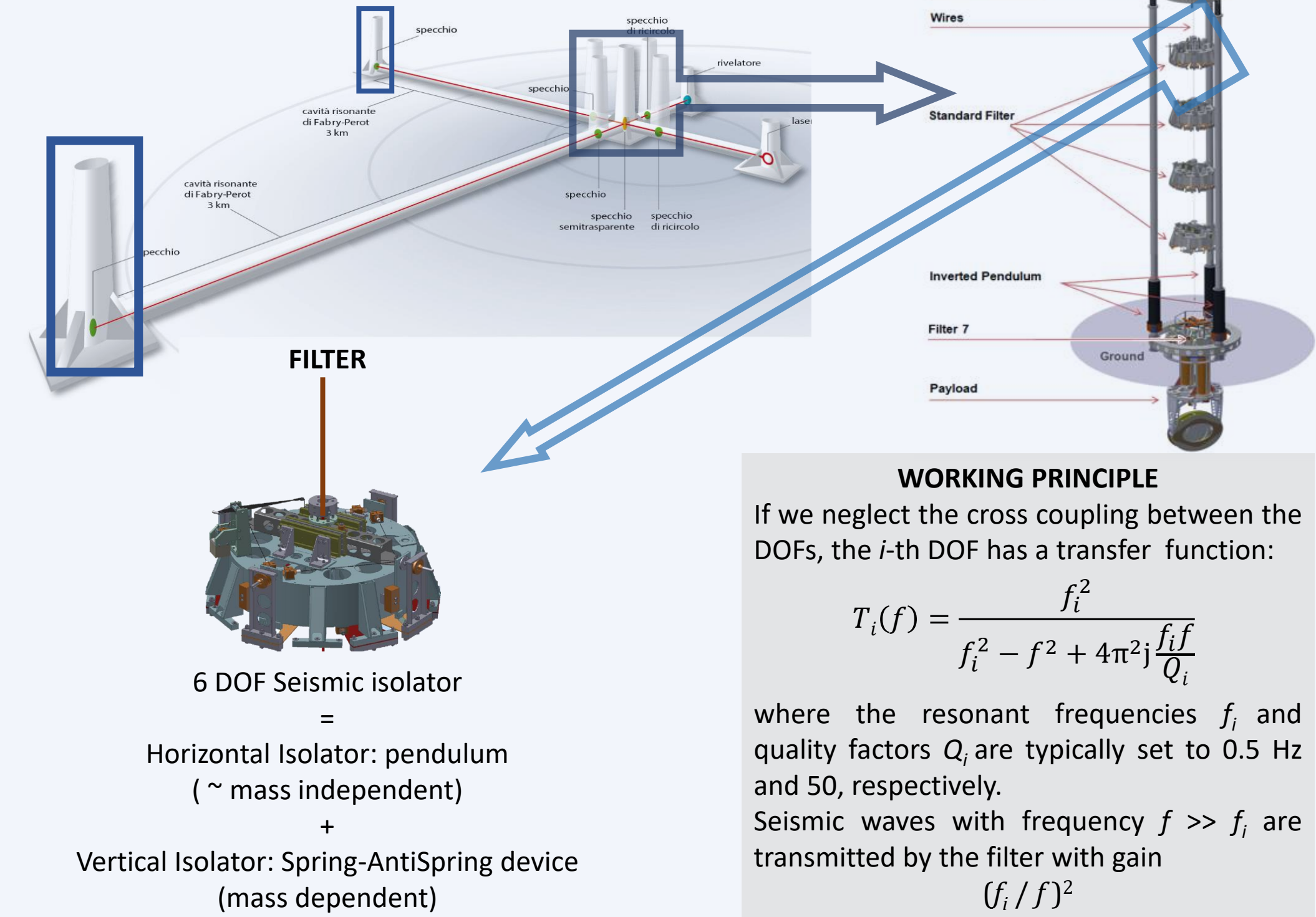
The laser beam size will be enlarged on end test masses and better coatings will be implemented on mirrors to lower mechanical losses. In particular, end mirrors will be larger (55 cm in diameter) and heavier (104 kg) to deal with the larger beam size.

The seismic isolation of AdVirgo+ mirrors and suspended benches will be provided by an upgraded version of the SuperAttenuator (SA), a passive attenuation system capable of reducing the seismic noise by more than 10 orders of magnitude in all six degrees of freedom above a few Hz.

The new end mirrors will require a re-scaling of the payload and, consequently, a re-design of all the elastic elements of the end SAs (blade springs, suspension wires, magnetic AntiSprings and Inverted Pendulum flex joints), in order to sustain the new loads without significant changes of the resonant frequencies. Several studies are being performed in order to design and validate the required mechanical updates.

Such studies are also providing useful insights on the design of seismic isolation systems for the third generation detectors.

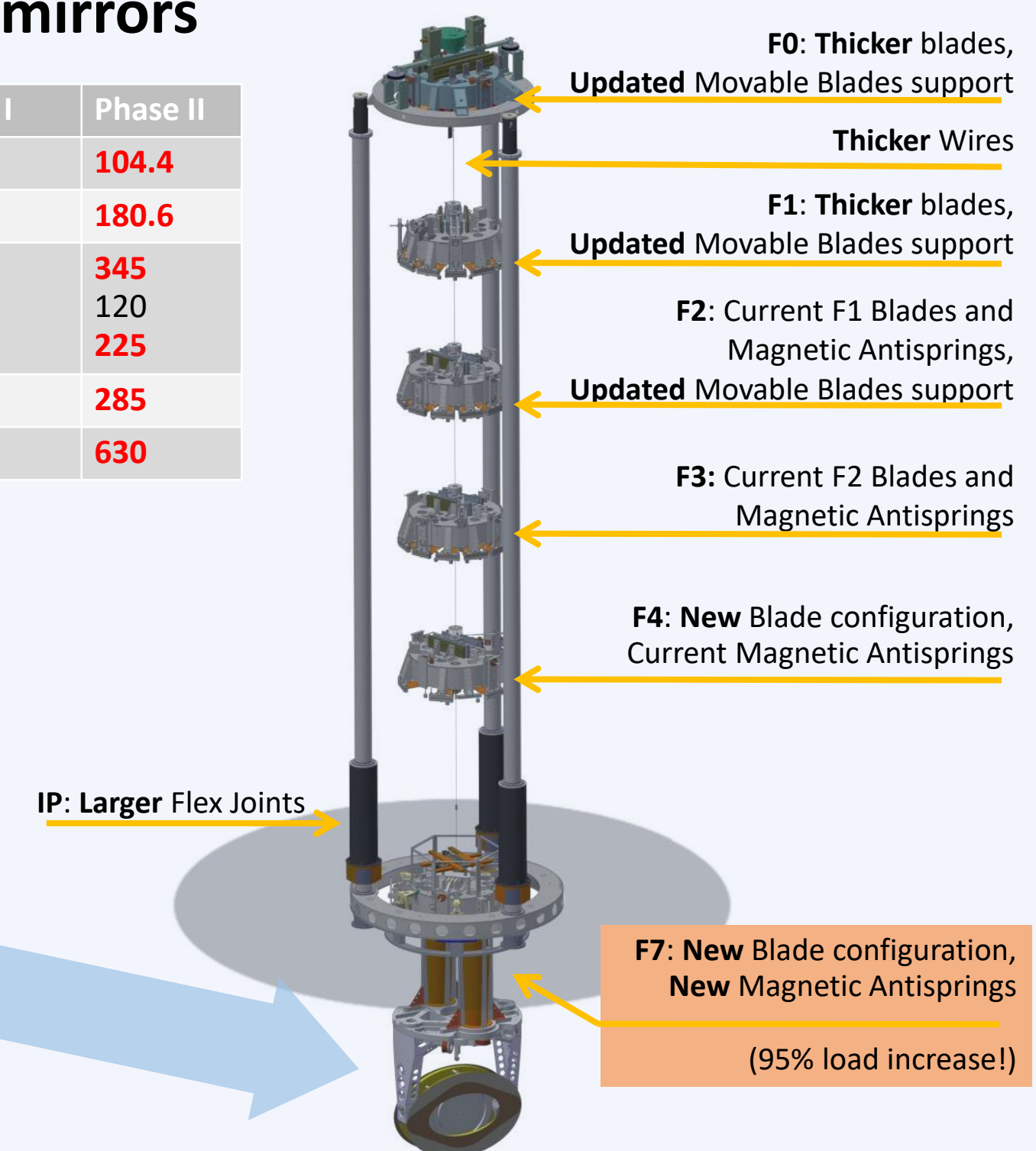
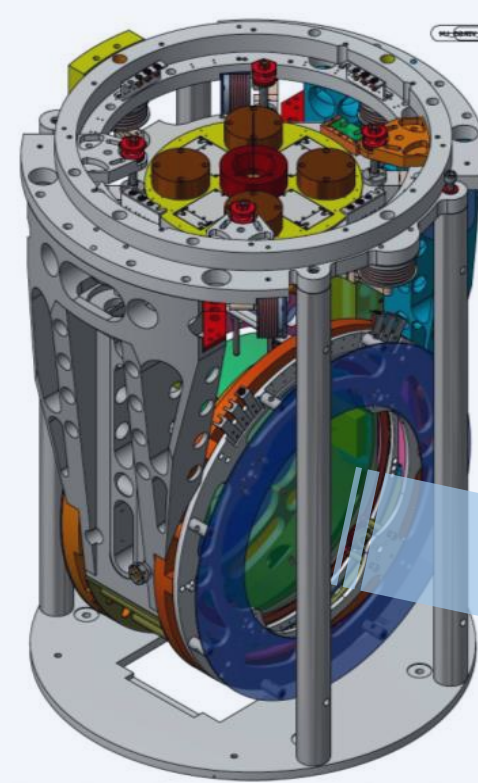
SuperAttenuator recap



New large mass end mirrors

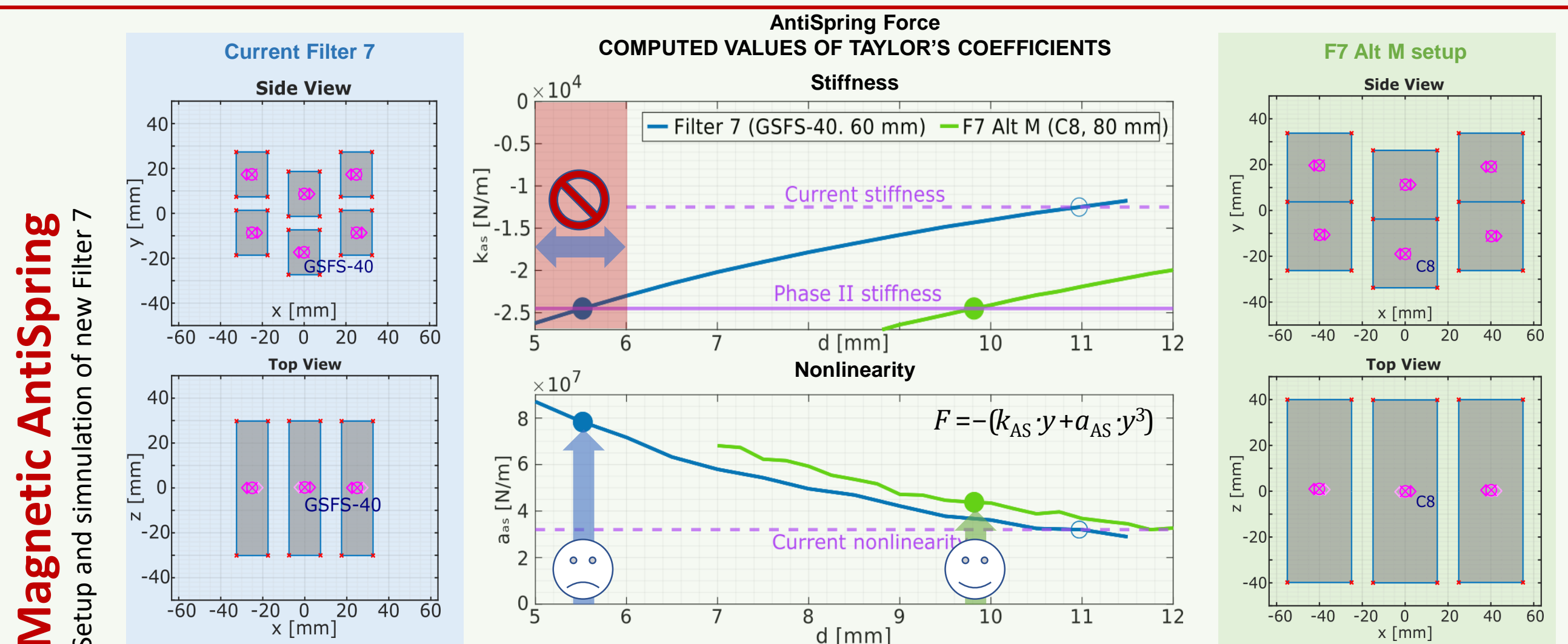
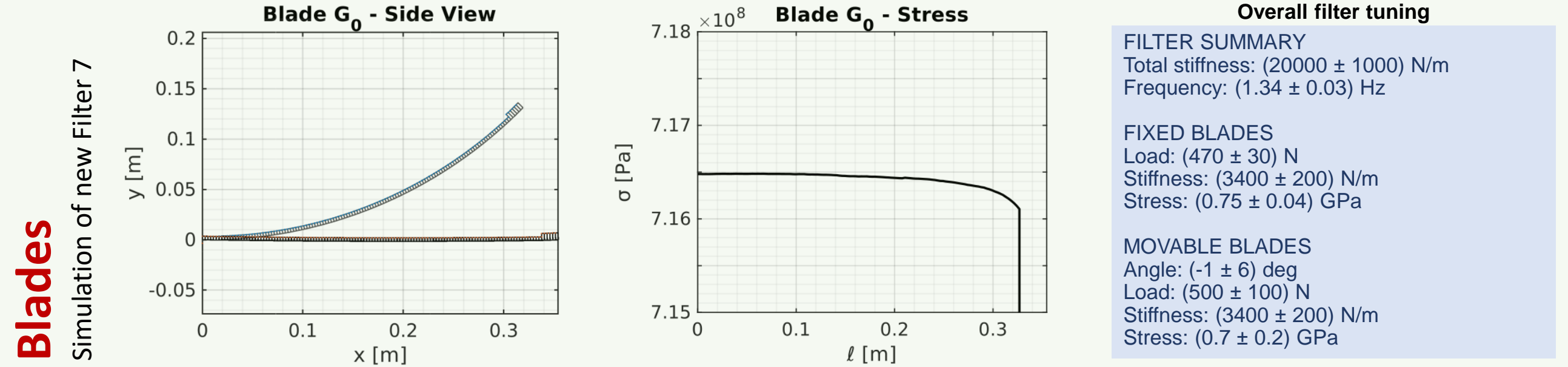
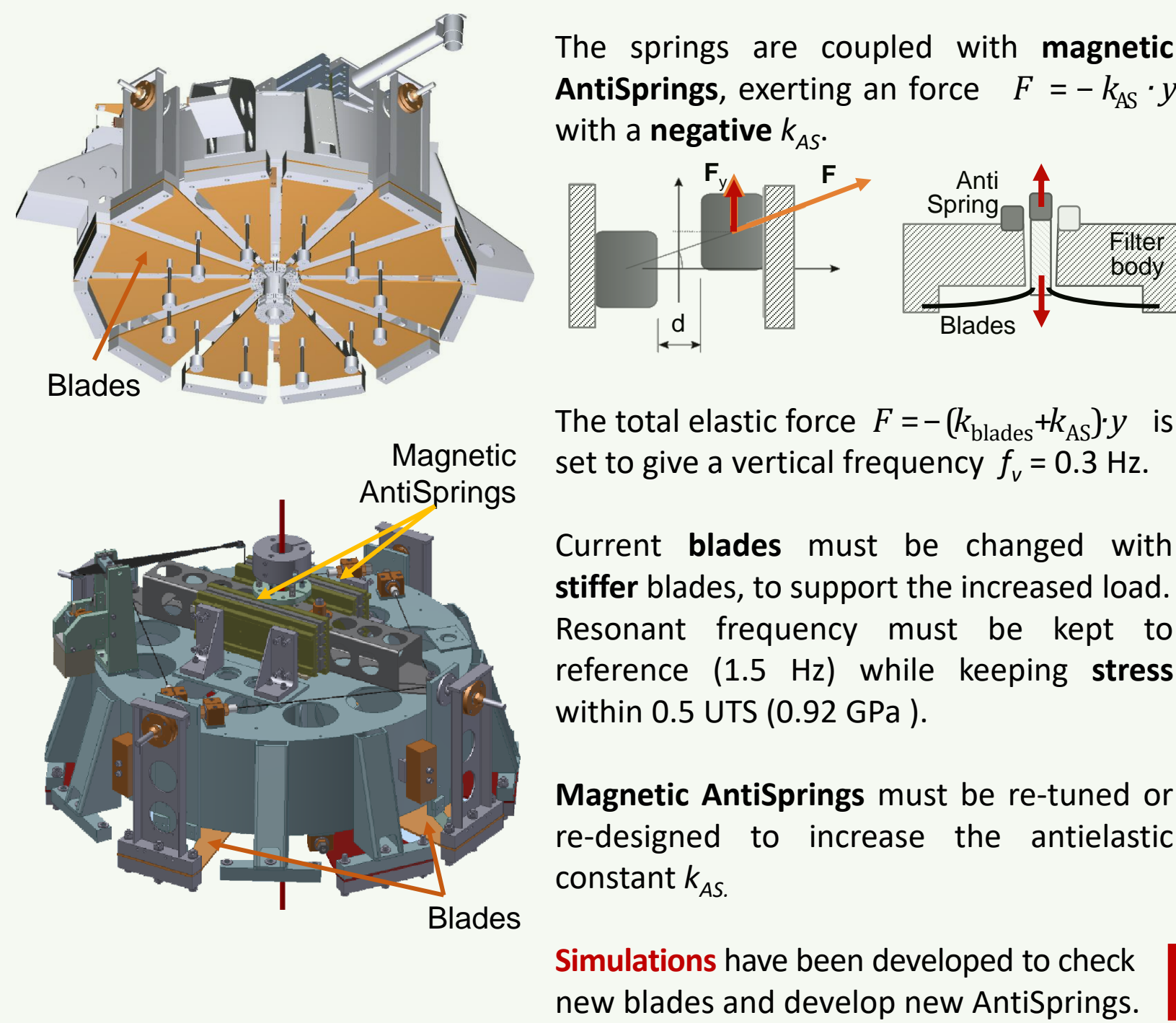
Physical quantity (SI units)	Phase I	Phase II
Mirror mass	42	104.4
Marionette mass	103.7	180.6
F7 mass (total)	315	345
• Body+interface	120	120
• Actuation cage	195	225
Total mass below F7 blades	145.7	285
Total mass below F4 blades	460.7	630

LARGE MASS PAYLOAD



Filters: vertical resonant frequency

Each filter supports the weight of the subsequent chain of filters by means of pre-bended **elastic maraging blades**. The blades must be stiff enough to support load. Such stiff **springs** have a high vertical resonant frequency (~ 1.5 Hz).



New blade equipments

Filter 0 and **Filter 1** have filled out blade slots (with the stiffest preset blades available)
⇒ Stiffer blades have to be designed.

Simple laws apply to the **flattening load** of pre-bended triangular blades and the **stress** of the loaded blades:

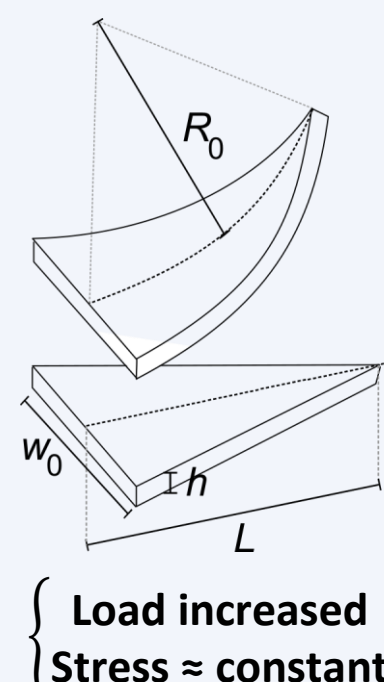
$$L = E \frac{h^3 w_0}{12 L R_C}, \quad \sigma = E \frac{h}{2 R_C} \quad (E = \text{Young modulus}).$$

Thickness h is **increased** from 3.5 mm to 3.7 mm ⇒

UPGRADED BLADE SERIES
10 kinds of **preset blades**:

Filter #	Load [kg]	Blades	Nominal Load [kg]	Load Mismatch [kg]	Total Stiffness [N/m]	Res. Freq. [Hz]
0	1223	8·A ₁ + 4·B ₁	1275	52	108644	1.53
1	1050	8·C ₁ + 4·B ₁	1113	63	93277	1.54
2	885	8·C ₀ + 4·B ₀	942	57	78620	1.55
3	745	6·D ₀ + 4·B ₀	771	26	66185	1.53
4	630	4·D ₀ + 4·B ₀	625	-2	55703	1.50
7	285	6·G ₀	290	3	25493	1.51

A₀, B₀, ..., G₀ ($h = 3.5$ mm)
+ A₁, B₁, C₁ ($h = 3.7$ mm)

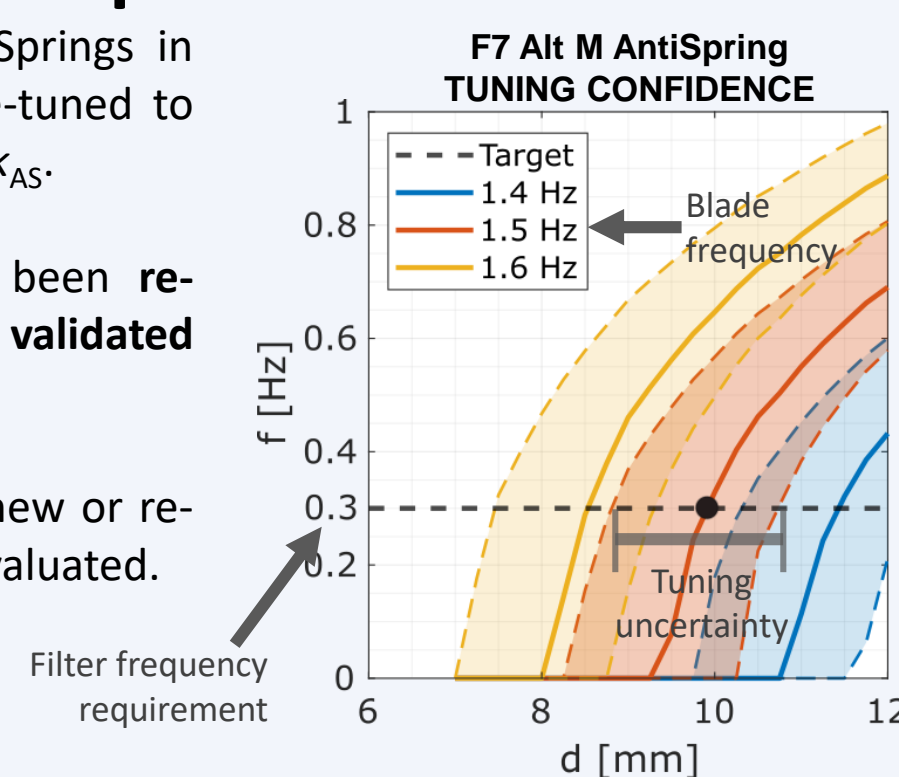


New AntiSpring setups

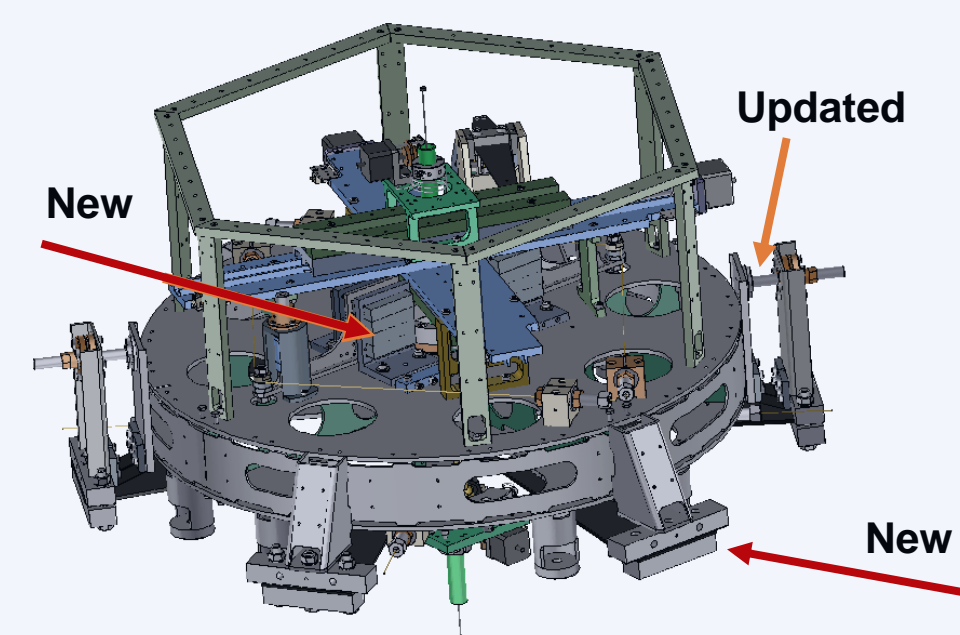
Simulations indicate that AntiSprings in filters 0 to 4 can be simply re-tuned to the required negative stiffness k_{AS} .

AntiSprings of Filter 7 have been **re-designed** and **experimentally validated** at INFN Pisa.

The tuning confidence of the new or re-tuned AntiSprings have been evaluated.



Filter 7 design upgraded



Perspectives

Phase II upgrade is designed and a prototype of new Filter 7 is being assembled and studied. AdVirgo+ will run with the new configuration in the fifth observation period O5 expected to start in Dec 2024.

Based also on these results, new generation attenuators are currently being studied.