Total Ionizing Dose effects on CMOS Image Sensor for the ULTRASAT space mission

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The ULTRASAT mission

ULTRASAT is a **wide-angle space telescope** that will perform a deep time-resolved all-sky survey in the **near-ultraviolet** (NUV) [1]. The science objectives are the detection of counterparts to gravitational wave sources and supernovae.

The mission is led by the Weizmann Institute of Science [2]. Launch is expected in early 2025 and 3 years of orbit operations are planned as a minimum. DESY will provide the **UV camera**, composed by the detector assembly located in the telescope focal plane and the remote electronics unit.

First results on pixel Total Ionizing Dose (TID) effects

We present **TID** results on Tower **test structures** with similar pixel to the final flight ULTRASAT sensor. Both sensors are Back Side Illuminated (**BSI**), **4T CMOS** Imaging Sensors made in the **Tower 180 nm** architecture. Due to the expected **similar photo-diodes** and technology node, we can use the conclusions from testing these test structures, to **optimize** the flight **sensor** performance. However, the final sensor is designed to be rad hard up to at least 20 krad TID and 60 MeV*cm²/mg Linear Energy Transfer Single Event Effects. Nevertheless, due to the similar photo-diodes and technology node, we expect that analysing the test structures' response to radiation will help us to gain a better understanding of the flight sensors.

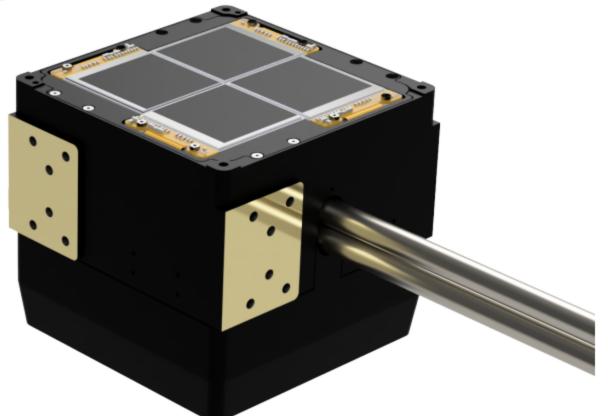


Figure 1: Rendering of the ULTRASAT camera components that will be developed at DESY Zeuthen. The detector assembly (DA) is based on four 180 nm back-side illuminated (BSI) CMOS sensor arrays with a total of 90 Megapixels that are operated at -70°C

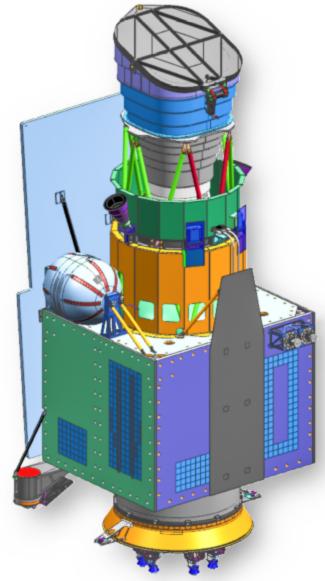


Figure 2: Rendering of the ULTRASAT space mission [3]





Figure 3: Left: Scout test structure board set-up. Right: Experimental setup at the CC-60 irradiation facility at CERN

Test structure (Scouts) experimental setup

- The Scouts readout chain is comprised out of a mainboard that acts as both a bias board and a read out board for the scout sensors and a daughterboard, in which Scouts can be socketed.
- Irradiations have been performed with a Co-60 source at CC-60, CERN and at HZB in Berlin. The dark signal has been measured before and after irradiation at the same temperature.
- During irradiations, the sensors were **biased** at the highest bias configuration, at room temperature.

Dark Signal vs TID first results

Sensor noise

- An increase in temporal pixel noise is observed with elevated accumulated TID.
- The noise is elevated both in terms of distribution mean, caused by the increase in-pixel shot noise and distribution tail, which hints towards a non-Gaussian noise component.

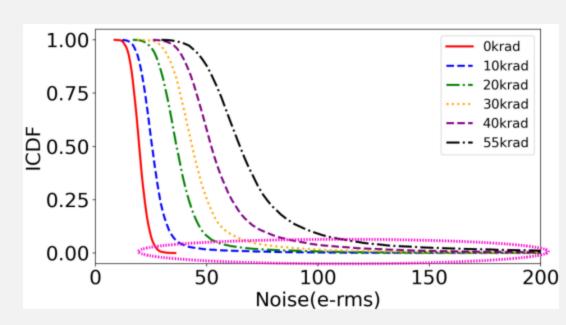


Figure 4: Inverse Cumulative Distribution of total inpixel temporal noise for multiple TID

Sensor Dark Current

 Elevated kurtosis and skewness of the dark counts distribution is observed for higher TID values.

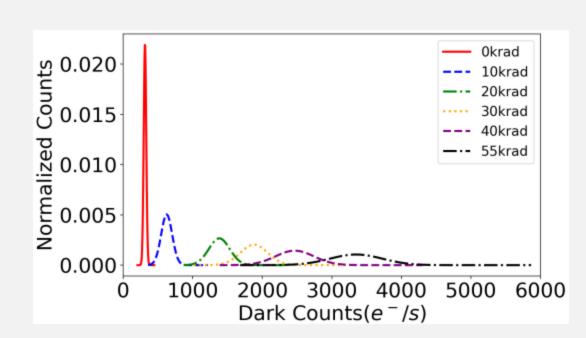


Figure 5: Probability Distribution Function of dark counts for multiple TID

These results hint towards an important **Dark Current** - **Random Telegraph Signal (DC-RTS)** contribution to the total post-irradiation noise.

References:

3094.

- [1] https://www.weizmann.ac.il/ultrasat/
- [2] https://www.weizmann.ac.il[3] https://www.iai.co.il/commercial
- [4] Virmontois, C., et all, (2011). IEEE
 Transactions on Nuclear Science, 58(6), 3085-

- Random Telegraph Signal detection
- In order to detect RTS jumps in signal, a Real-time Auto-detection
 Method for RTS is employed [4].
- The signal is first filtered with the following filter:

$$H(z) = \sum_{i=0}^{L/2-1} \left(\frac{2}{L}z^{-i}\right) - \sum_{j=L/2}^{L-1} \left(\frac{2}{L}z^{-j}\right)$$

 Triangles are formed at the leading edge, then are used to detect RTS jumps, based on a dynamic threshold for each noise segment.

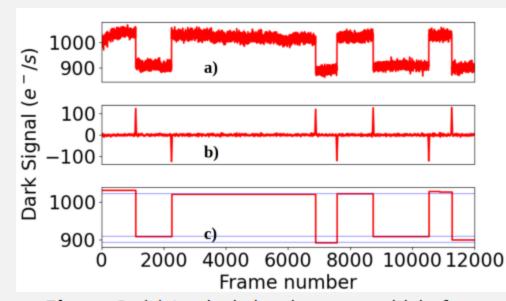


Figure 6: (a) In pixel signal across multiple frames; (b) Filtered signal; c) Reconstructed RTS levels

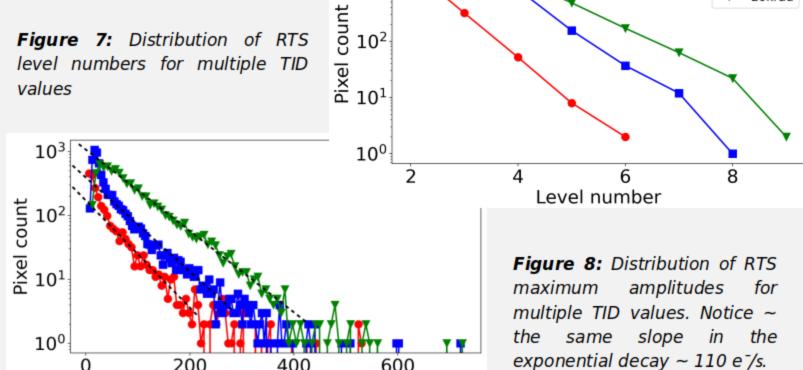
1krad

→ 5krad

→ 20krad

Random Telegraph Signal results

- Significant RTS is observed for relatively low doses
 ~1-5 krad. This is most probably due to the non-accumulated operation mode of the Transfer Gate.
- No RTS was detected for non-irradiated samples.
- Distribution of RTS levels and amplitudes for multiple TID levels show an increase of TID induced defects in the photo-diode.
- The distribution of the maximum RTS amplitudes follow the same exponential decay, showing that the same RTS mechanism takes place at all TID values.



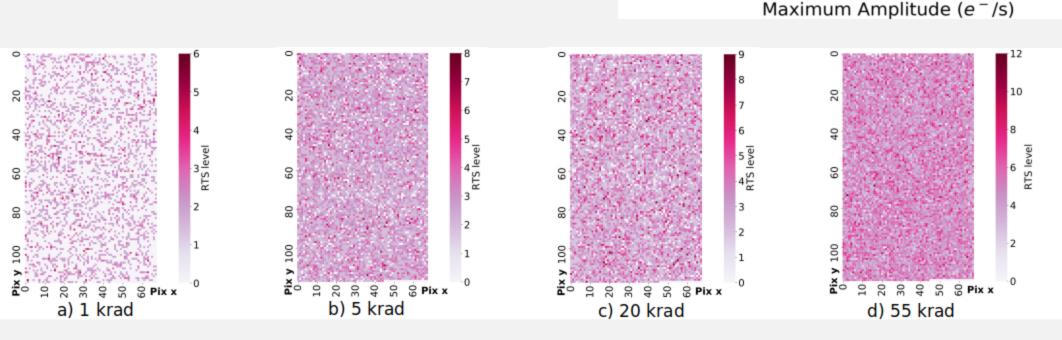


Figure 9: Levels of RTS projected over the whole pixel array at: a) 1 krad TID; b) 5 krad TID; c) 20 krad TID; d) 55 krad TID.

Summary

- The preliminary investigation on TID induced defects in the Scout test structures provided us with information on the expected response of the ULTRASAT flight grade sensors to radiation.
- The **lack of RTS in non-irradiated test structures**, shows a good isolation between the Space Charge Region of the photo-diode and the Silicon-Oxide interface.
- Due to TID and the **non-accumulated mode** in which the **Transfer Gate** operates, sensor performance degradation is observed.
- Further studies are planned in characterizing the RTS properties against the Low Transfer Gate Voltage and **accelerated annealing**, in order to showcase the source of these defects.
- This study will improve the understanding of the final operation of the flight sensor for the ULTRASAT mission and will help to streamline the radiation characterization of the flight sensors.



