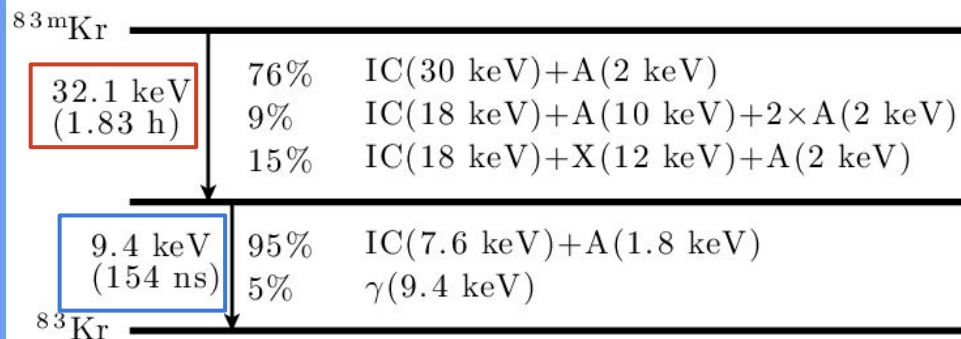
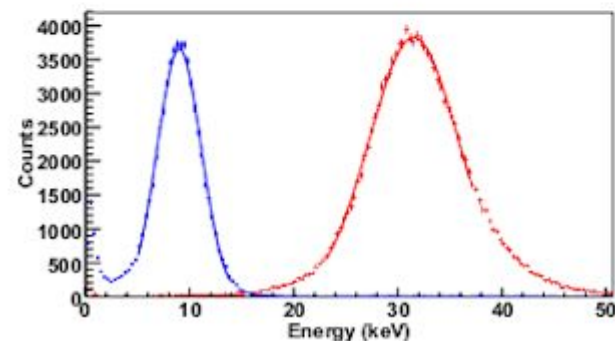
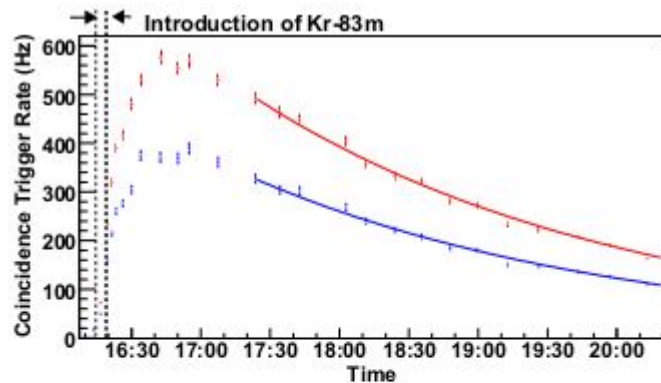


Calibration Sources for CYGNO

Flavio Di Clemente
Giulia D'Imperio
Gianluca Cavoto

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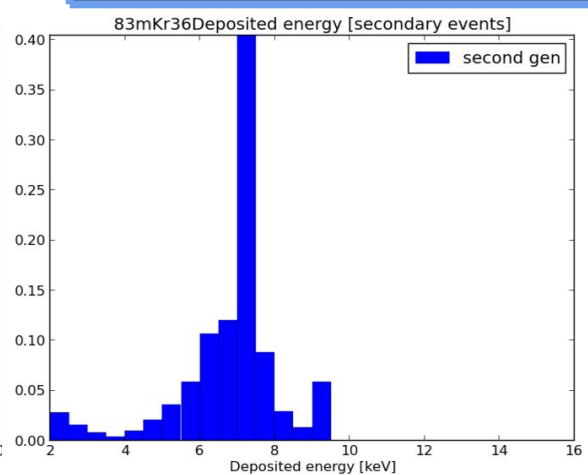
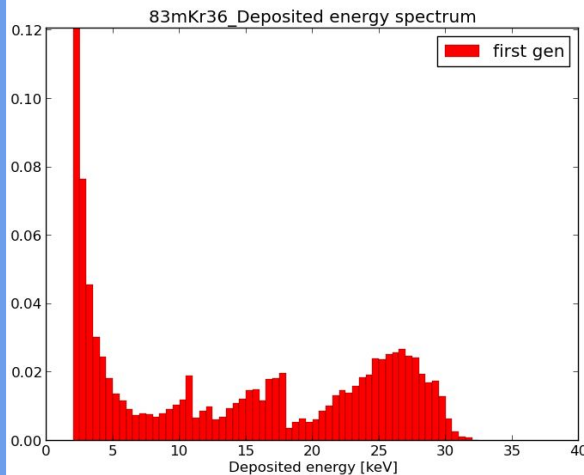
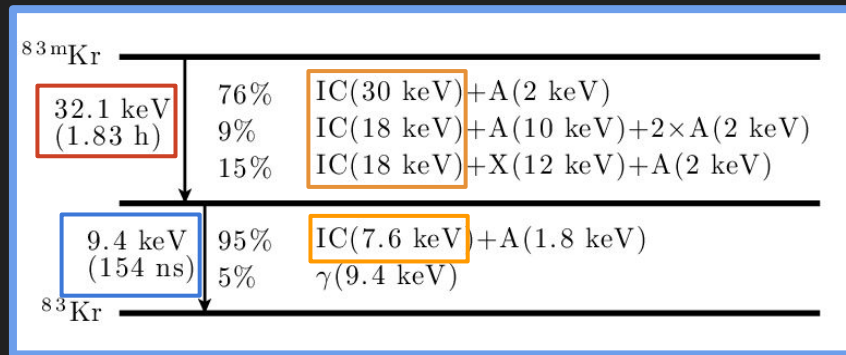
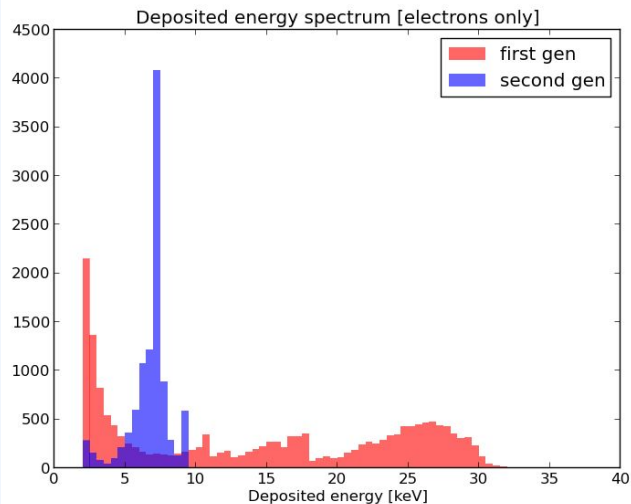
83m Krypton



<https://arxiv.org/abs/0905.1766v2>

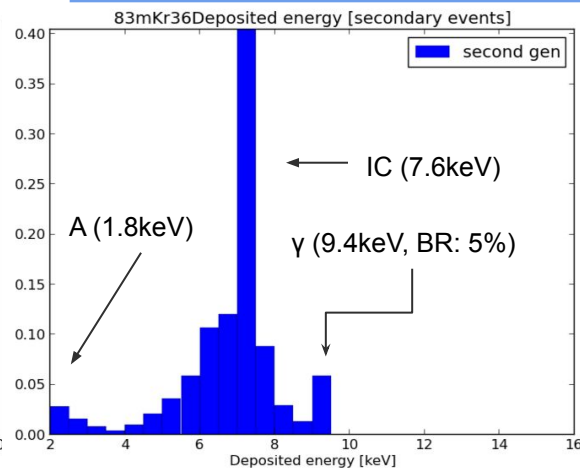
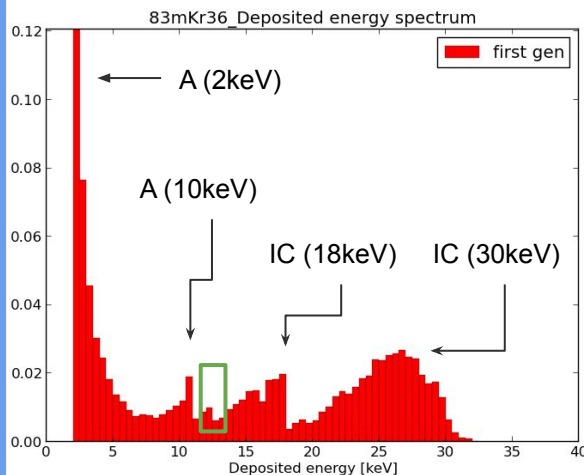
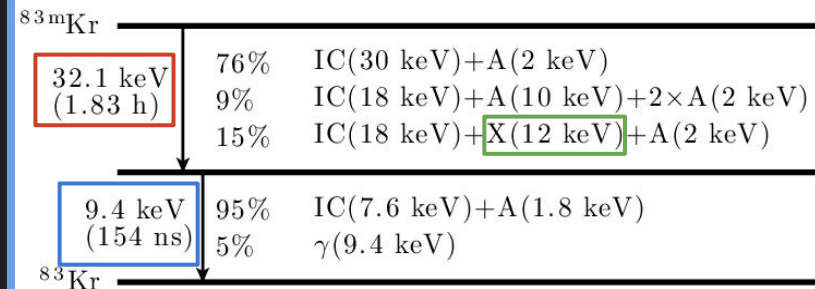
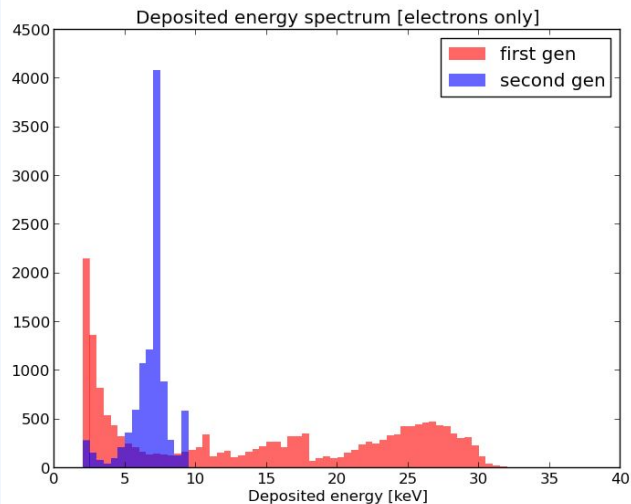
Calibration of a Liquid Xenon Detector with $^{83}\text{Kr}^m$

L. W. Kastens, S. B. Cahn, A. Manzur, and D. N. McKinsey
 Department of Physics, Yale University, P.O. Box 208120, New Haven, CT 06520



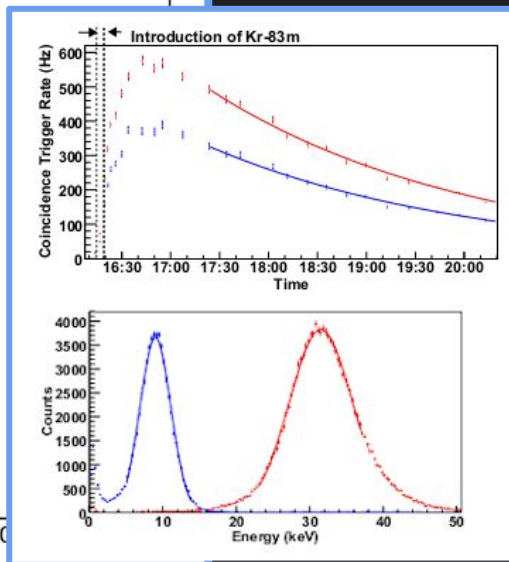
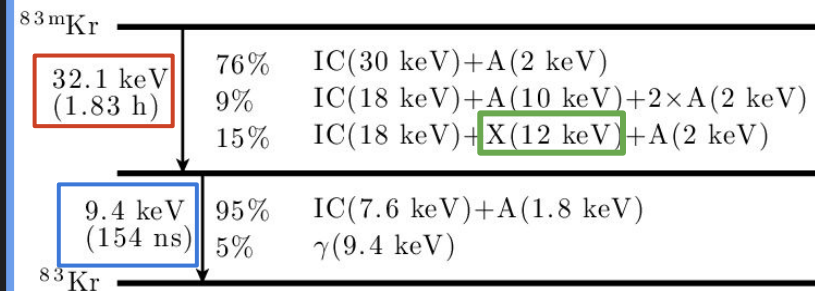
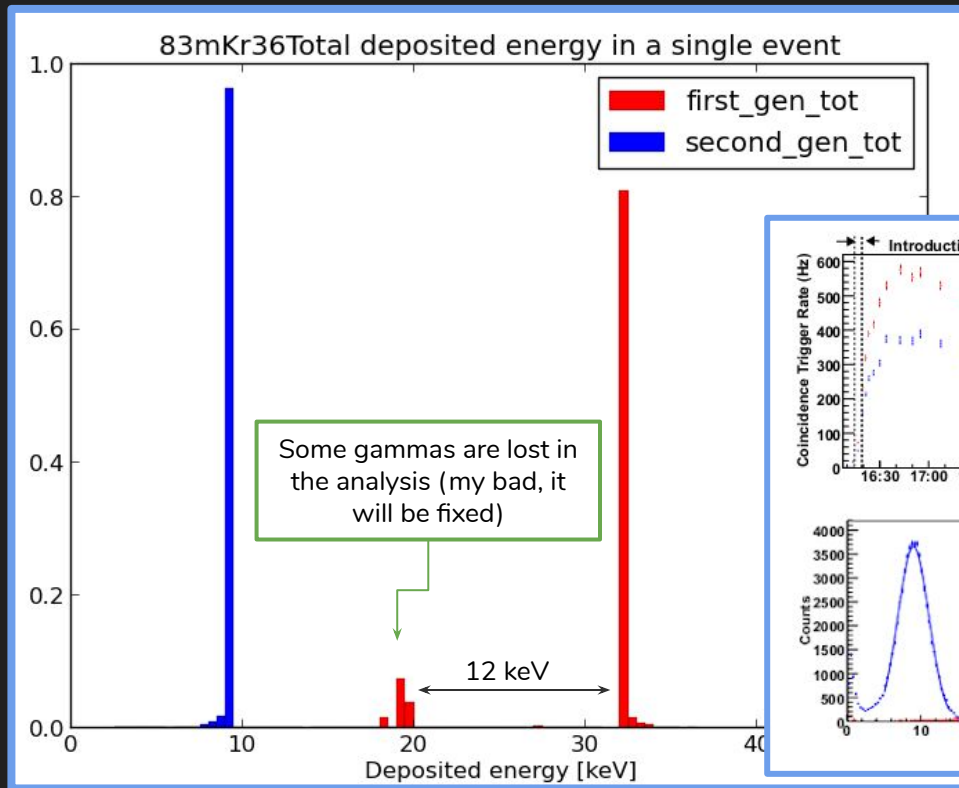
Internal conversion (IC) in the inner shells leaves a hole behind. This produces a chain of electrons filling lower levels consequently.

This chain reaction produces a number of **X-rays (X)** and **Auger Electrons (A)**.



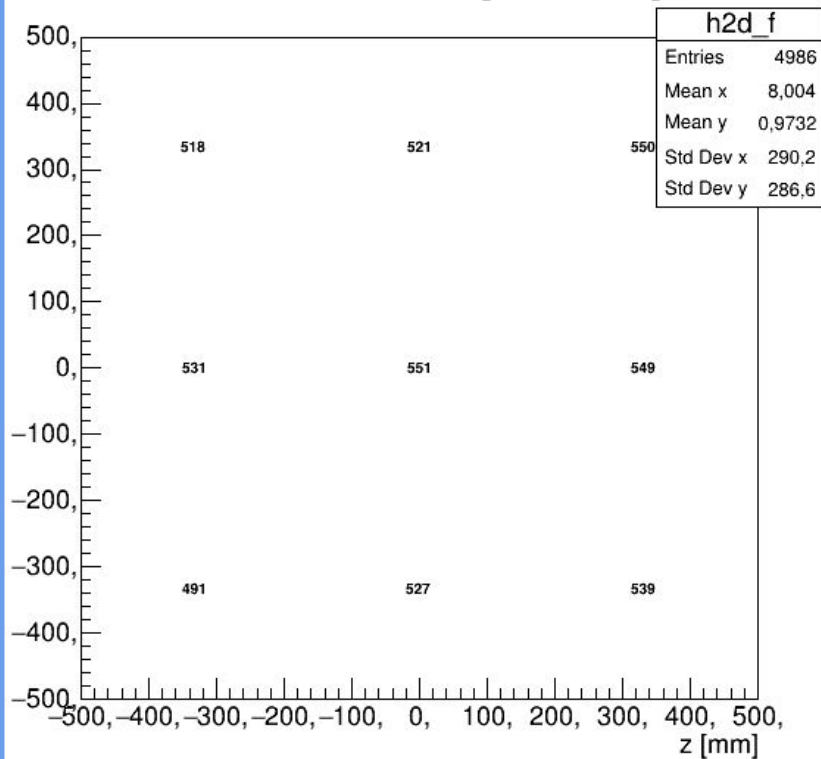
The two phases of the de-excitation can be seen as **two independent processes**, resulting in two different source of calibration at once.

The histogram is filled with the sum of the energies of all the particles resulting from the de-excitation of a single event (dividing the two steps)

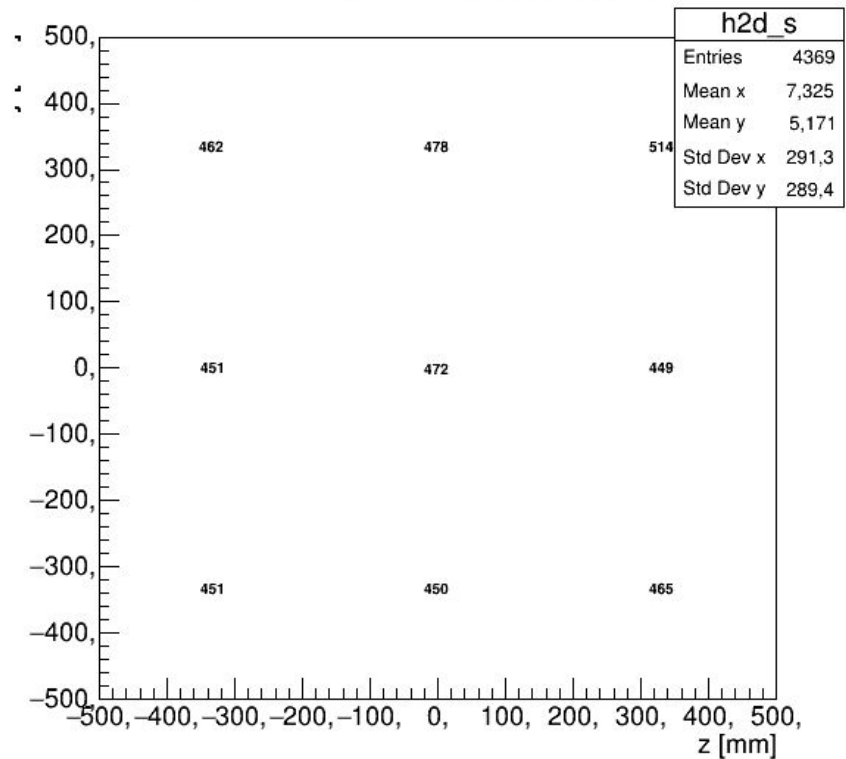


Since this is a MC-truth-level simulation, we don't see gaussians as presented in the article listed before (no digitization added, no smearing effect)

83mKr36_hist2d_first_gen_coverage



83mKr36_hist2d_second_gen_coverage



We are working on 10^4 events. Events which are fully contained in a single camera represent ~40-50% of the sample. The source is in a gaseous form, thus it can evenly cover all the cameras. Each camera contains ~5% of the entire sample.

Specific Activity of the 83mKr

$$A_{83\text{mKr}} = \frac{\ln(2) N_A}{T_{1/2} m_w} \left\{ \begin{array}{l} T_{1/2} = 1.83\text{h} = 6588\text{s} \\ m_w = 83.8 \text{ g mol}^{-1} \end{array} \right. \Rightarrow A_{83\text{mKr}} = 7.6 \times 10^{20} \text{ Bq/kg}$$

But 83mKr comes from 83Rb, which has $T_{1/2} = 86 \text{ days!}$ $\Rightarrow A_{83\text{Rb}} = 6.7 \times 10^{17} \text{ Bq/kg}$

83Rb, again, is not a natural nuclide...

Conclusions

^{83m}Kr would be a great calibration source for CYGNO, but it needs to be produced in some way.

A gaseous source will perfectly satisfy the need for a uniform coverage.

The ^{83m}Kr half-life is short enough to make the source easily removed from the sensitive region, but long enough to make a calibration run possible.

Also the energy spectrum has a well-resolved double peak at low energy.