High Precision X-ray Measurements 2021

Probing Electron Properties in ECR Plasmas using X-ray Bremsstrahlung and Fluorescence Emission



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Brief Overview



O.Tarvainen, I. Izotov, D. Mansfeld, V. Skalgya, S. Golubev, T. Kalvas, H. Koivisto, J. Comppula, R. Cronholm, J. Laulainen and V. Toivanen, Plasma Sources Sci. Technol. 23, 025020 (2014)
 D. Mascali, G. Torrisi, L. Neri, G. Sorbello, G. Castro, L. Celona and S. Gammino, Eur. Phys. J. D 69, 27 (2015).

[3] L. Celona, G. Ciavola, F. Consoli, S. Gammino, F. Maimone, P. Spaedtke, K. Tinschert, R. Lang, J. Mader, J. Rosbach, S. Barbarino, R. S. Catalano and D. Mascali, Rev. Sci. Instrum. 79, 023305 (2008).

ECRIS: Operation and Properties



[4] D. Mascali, A. Musumarra, F. Leone, F.P. Romano, A. Galatá, S. Gammino and C. Massimi, Eur. Phys. J. A 53, 145 (2017).
[5] S. Gammino, G. Ciavola, L. Celona, D. Mascali and F. Maimone, IEEE Transactions on Plasma Science 36, 4 (2008)

ECRIS: Operation and Properties (contd.)



12.84 GHz

3D structure of plasma from simulations – the spatial anisotropy and non-homogeneity with respect to density and energy is visible [6]

Energetic electrons (~ keV) and radiation interact with the ions through multiple processes (ionization, excitation etc.) to produce high charge states (NLTE population kinetics solved with CR model)





Net particle density distribution (positive charge in blue, negative in red [7]

[6] R. Racz, D. Mascali, S. Biri, C. Caliri, G. Castro, A. Galata, S. Gammino, L. Neri, J. Palinkas, F.P. Romano and G. Torrisi, Plasma Sources Sci. Technol. **26**, 075011 (2017).

[7] D. Mascali, L. Neri, L. Celona, G. Castro, G. Torrisi, S. Gammino, G. Sorbello and G. Ciavola, Rev. Sci. Instrum. 85, 02A511 (2014)

PANDORA: Plasmas for Astrophysics, Nuclear Decays Observation and Radiation for Archaeometry A new ECR Ion Trap for in-plasma β-decay measurements









[4] D. Mascali, A. Musumarra, F. Leone, F.P. Romano, A. Galatá, S. Gammino and C. Massimi, Eur. Phys. J. A 53, 145 (2017).
[8] K. Takahashi and K. Yokoi, Nucl. Phys. A (1963)

Emission spectroscopy excellent indirect diagnostic tool

Different energy ranges of electrons can be probed by analysing different photon ranges

Optical spectra for $T_{\rho} \sim 10 \text{ eV}$ [9]

Fluorescence and low energy bremsstrahlung for $T_a \sim 10$ keV

ECE and bremsstrahlung for $T_e > 50$ keV [10]

Novel plasma diagnostic technique developed by ATOMKI, Debrecen and LNS, Catania [6,11,12] – overall objective to study ECR plasma properties through multiple perspectives: space-resolved X-ray images, volumetric X-ray emission and ion current extraction



Schematic of experimental setup showing various detectors used and distances [11]

[6] R. Racz, D. Mascali, S. Biri, C. Caliri, G. Castro, A. Galata, S. Gammino, L. Neri, J. Palinkas, F.P. Romano and G. Torrisi, Plasma Sources Sci. Technol. 26, 075011 (2017). [11] D. Mascali, G. Castro, S. Biri, R. Racz, J. Palinkas, C. Caliri, L. Celona, L. Neri, F. P. Romano, G, Torrisi and S. Gammino, Rev. Sci. Instrum. 87, 02A510 (2016). [12] R. Racz, S. Biri, J. Palinkas, D. Mascali, G. Castro, C. Caliri, F.P. Romano and S. Gammino, Rev. Sci. Instrum. 87, 02A741 (2016).

2D Space-Resolved X-Ray Imaging – Experiment Details

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2D Space-Resolved X-Ray Imaging – Theoretical Model: General Outline

and EEDF



[6] R. Racz, D. Mascali, S. Biri, C. Caliri, G. Castro, A. Galata, S. Gammino, L. Neri, J. Palinkas, F.P. Romano and G. Torrisi, Plasma Sources Sci. Technol. **26**, 075011 (2017).

2D Space-Resolved X-Ray Imaging – Theoretical Model: Electron Kinetics





4 choices of EEDF – mix of Maxwell and Druyvesteyn distribution functions

Ref. case name	Type
EEDF1	Low- $E f_M$ + High- $E f_M$
EEDF2	Low- $E f_M$ + High- $E f_D$
EEDF3	Low- $E f_M$ + Medium- $E f_M$ + High- $E f_M$
EEDF4	Low- $E f_M$ + Medium- $E f_D$ + High- $E f_M$

$$f_M(E;k_BT_e) = \frac{2}{\sqrt{\pi}} \frac{\sqrt{E}}{(\sqrt{k_BT_e})^3} e^{-E/k_BT_e}$$

$$f_D(E;k_BT_e) = 1.04 \frac{\sqrt{E}}{(\sqrt{k_BT_e})^3} e^{-0.55E^2/(k_BT_e)^2}.$$

Each EEDF tested in each ROI – better and more physical analysis

MSE and r² calculated for each cell of the ROI, then mean and SD of both quantities evaluated

Mean – average value of the statistic in the ROI SD – variation of actual value from the mean *within* the ROI

Thus, low mean MSE, high mean r², and low SD for both implies best performance



EEDF2 = Cold Maxwell + Hot Druyvesteyn works in each ROI [14] But simulations give relative electron density distribution, absolute n_e unknown still

[14] B. Mishra, accepted to be published, Il Nuovo Cimeno 2021

^[13] A. Galata, C.S. Gallo, D. Mascali and G. Torrisi, eprint arXiv:1912.01988

2D Space-Resolved X-Ray Imaging – Theoretical Model: Cross-Section and LGE

Lotz formula for K-shell ionisation cross-section



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Spatially-resolved geometrical efficiency for anisotropic and non-homogeneous emitting source



Schematic of the detection system



Isometric view of the chamber, pinhole and CCD for ray-tracing

Ray-Tracing Monte Carlo code to evaluate a 3D map of geometrical efficiency, a value for each plasma cell

> [15] B. Mishra, to be submitted to Phys. Of Plasmas

2D Space-Resolved X-Ray Imaging – Theoretical Model: LGE (contd.)



2D Space-Resolved X-Ray Imaging – Comparison and Conclusion



Romano and G. Torrisi, Plasma Sources Sci. Technol. 26, 075011 (2017).

Volumetric X-Ray Spectroscopy – Experiment Details



Bremsstrahlung Emissivity Density

$$J_{theo,brem}(h
u) =
ho_e
ho_i \int_{h
u}^{\infty} rac{\mathrm{d}\sigma_K(h
u)}{\mathrm{d}h
u} v_e(E) f(E) \mathrm{d}E$$



$$\frac{\mathrm{d}\sigma_{K}(h\nu)}{\mathrm{d}h\nu} = \frac{16\pi}{3\sqrt{3}}\alpha^{3} \left(\frac{\hbar}{m_{e}c}\right)^{2} \left(\frac{c}{v_{e}}\right)^{2} \frac{Z^{2}}{h\nu}$$

Single-component Maxwell EEDF

$$f_M(E;k_BT_e) = \frac{2}{\sqrt{\pi}} \frac{\sqrt{E}}{(\sqrt{k_BT_e})^3} e^{-E/k_BT_e}$$

$$J_{theo,brem}(h
u)=
ho_e
ho_i(Z\hbar)^2igg(rac{4lpha}{\sqrt{6m_e}}igg)^3igg(rac{\pi}{k_BT_e}igg)^{1/2}{
m e}^{h
u/k_BT_e}$$

Fluorescence Emissivity Density (Ar ions)

$$J_{theo,2.96} = \frac{h\nu_{2.96}}{\Delta E} \rho_e \rho_i \omega_{2.96} \int_{3.205}^{\infty} \sigma_{K,ion}(E) v_e(E) f(E) dE$$
$$J_{theo,3.19} = \frac{h\nu_{3.19}}{\Delta E} \rho_e \rho_i \omega_{3.19} \int_{3.205}^{\infty} \sigma_{K,ion}(E) v_e(E) f(E) dE$$

Lotz formula for K-shell ionisation cross-section

$$\sigma_{K,ion} = a_K q_K rac{\ln arepsilon/I}{arepsilon I} \{1 - b_K \exp[-c_K(arepsilon/I - 1)]\}$$



Pseudo-Voigt profile for line-broadening

$$D_{PV}(x-x_0,f) = \eta L(x-x_0, au_L) + (1-\eta)G(x-x_0,\sigma_G)$$





Volumetric X-Ray Spectroscopy – Theoretical Model: Cross-Section

Fluorescence from plasma interior – Ar ions and confined electrons

Lotz formula for K-shell ionisation cross-section

Fluorescence from extraction plate – Cr/Fe atoms and escaping electrons

Deutsch-Mark formalism for K-shell ionisation cross-section



Volumetric X-Ray Spectroscopy – Comparison and Conclusion



Chosen model fits experimental data well and gives valuable information about confined and escaping electrons

Conclusions

Space-resolved X-ray Imaging

Theoretical model developed based on selfconsistent numerical simulations on warm electrons + LGE evaluation

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Preliminary results show match with respect to overall structure (correlated with simulations/relative distribution) and counts (correlated with absolute particle density)

Uncertainties in both experimental and theoretical counts $-n_e$ used too high owing to small contribution from warm electrons, corroborated with volumetric analysis

UPDATE SIMULATIONS CONSIDER HOTTER ELECTRONS IMPROVE LGE + OVERALL MODEL

Volumetric X-Ray Spectroscopy

Theoretical model developed based on bremsstrahlung and fluorescence from trapped particles and fluorescence from escaping electrons impinging on extraction plate

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Excellent match between model and experiment, fit procedure furnishes particle density in expected range, predicts contribution from hotter electrons towards Kα emissions





Sandor Biri Richard Rácz



David Mascali Domenico Santonocito Angelo Pidatella Eugenia Naselli Maria Mazzaglia Giorgio Mauro Giuseppe Torrisi

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Alessio Galatà

THANK YOU FOR YOUR ATTENTION!