

Gravitational Laboratories for Nuclear Physics

Prospects for Binary Neutron Star Observations
and their Impact on the Equation of State

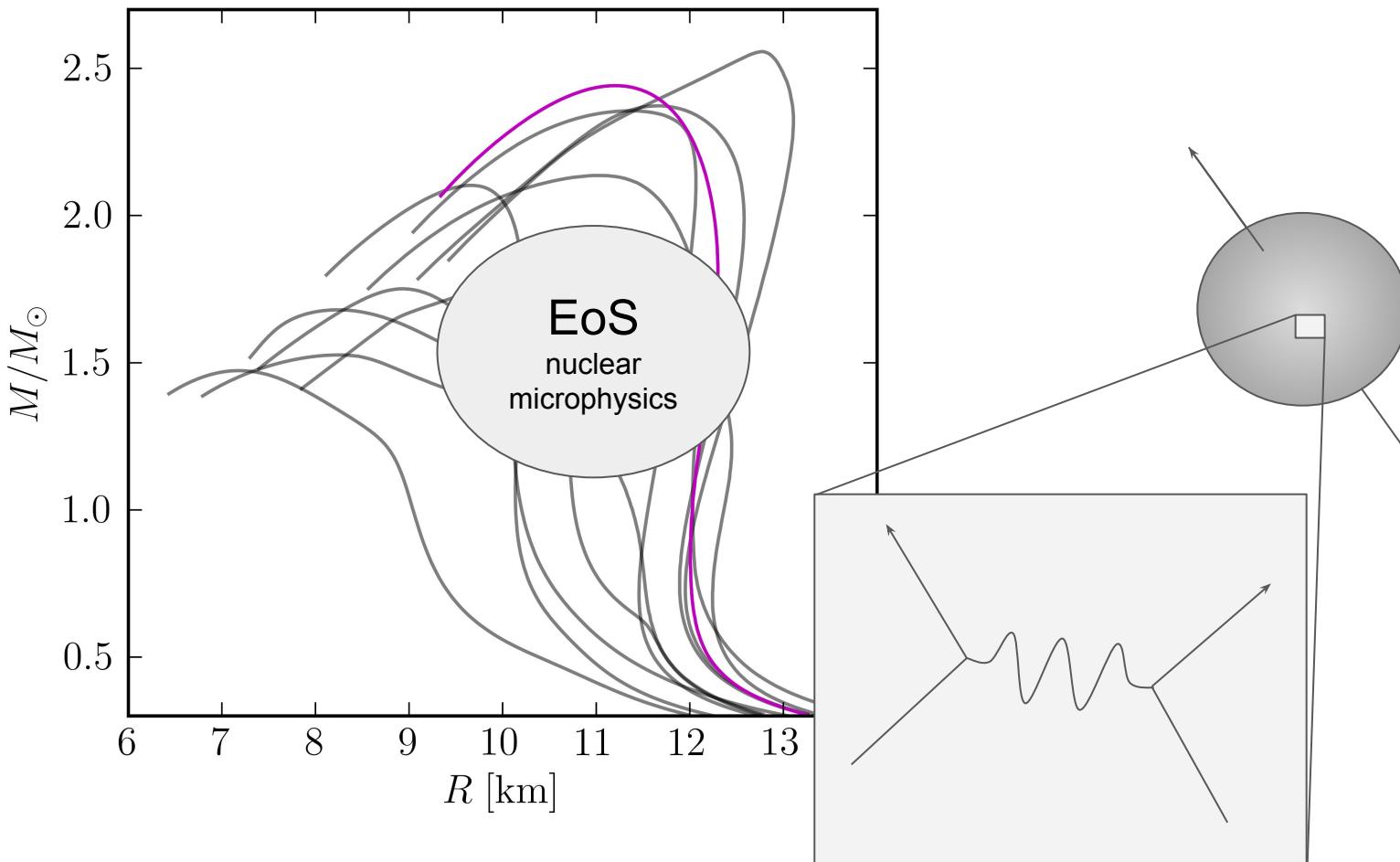
Reed Essick

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Perimeter Institute for Theoretical Physics

What Do We Observe?

What do we observe?

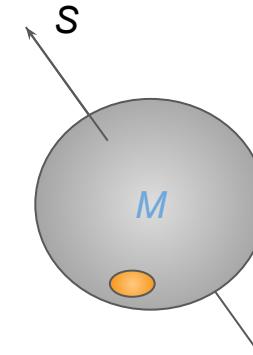
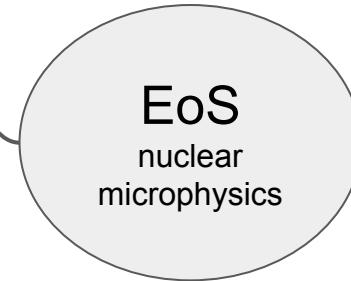


What do we observe?

M_{max}

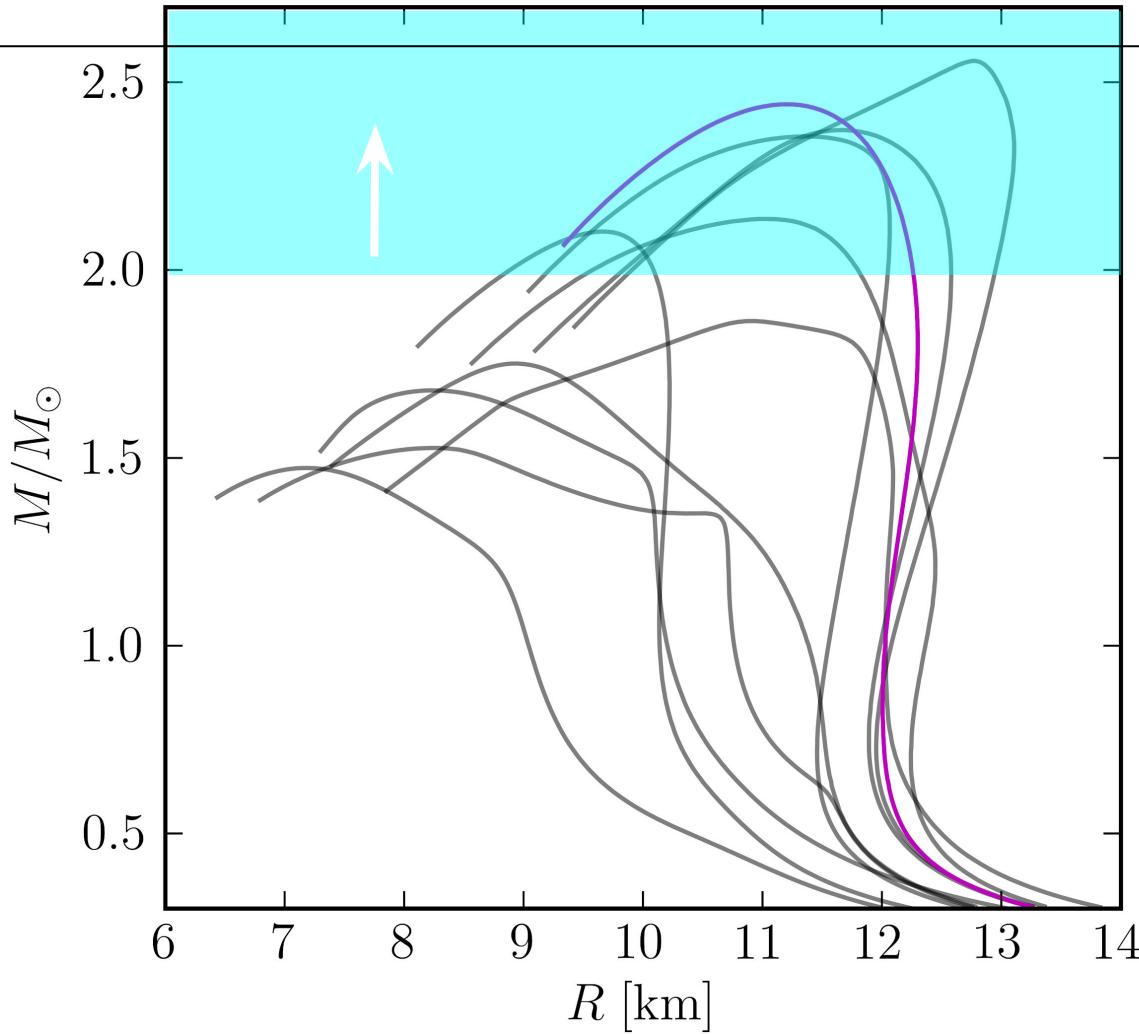
Massive pulsars

J0740+6620 Cromartie+(2019)
Fonseca+(2021)

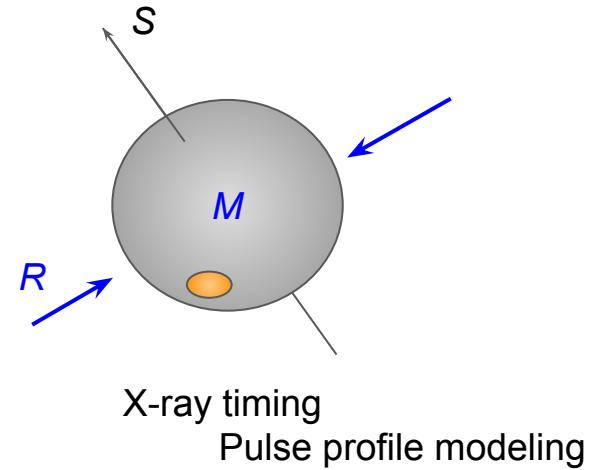
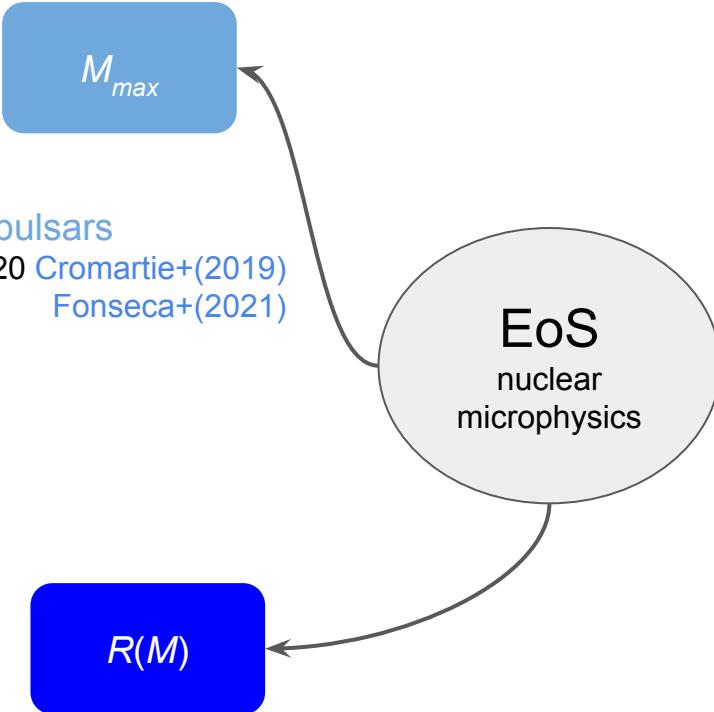


pulsar observations
radial velocity
Shapiro delay

What do we observe?

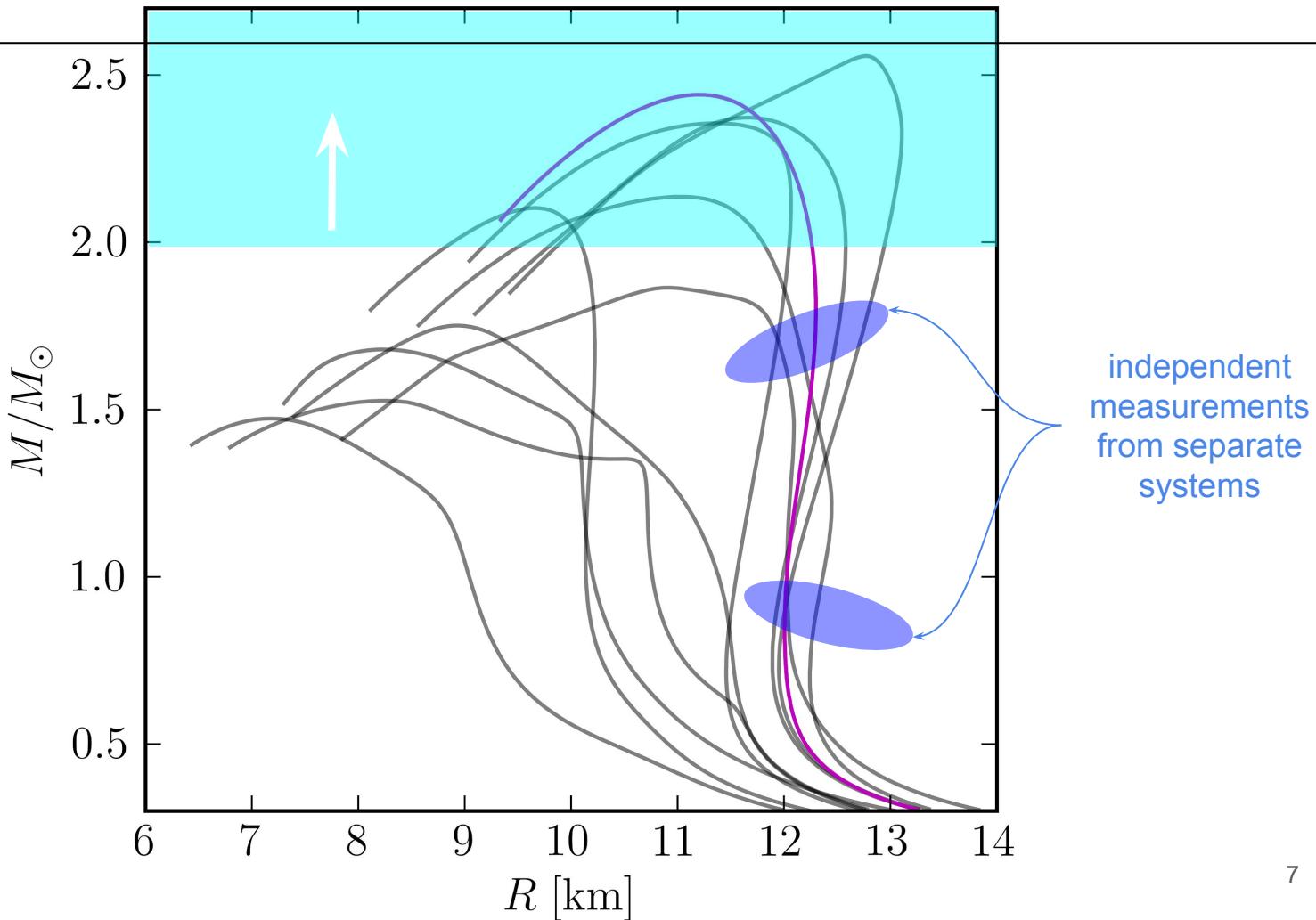


What do we observe?

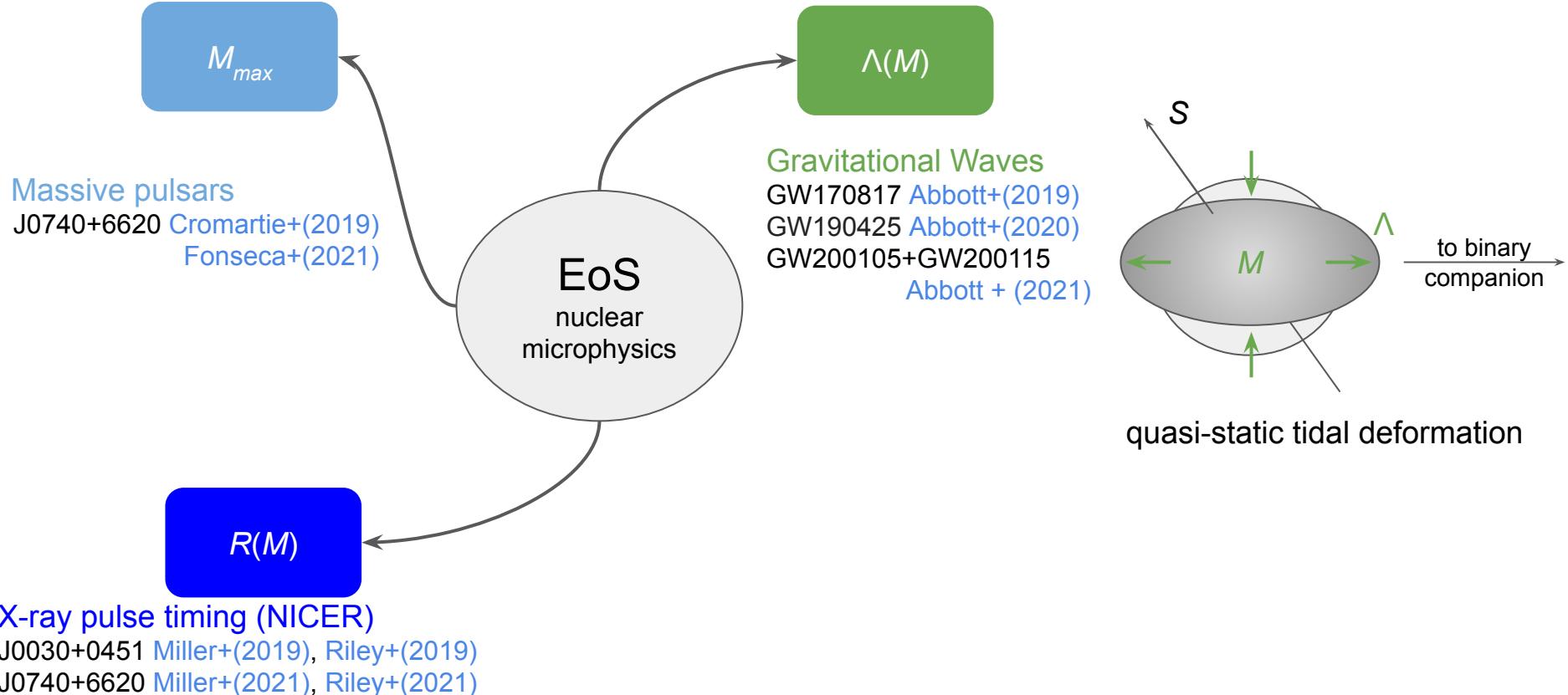


X-ray pulse timing (NICER)
J0030+0451 Miller+(2019), Riley+(2019)
J0740+6620 Miller+(2021), Riley+(2021)

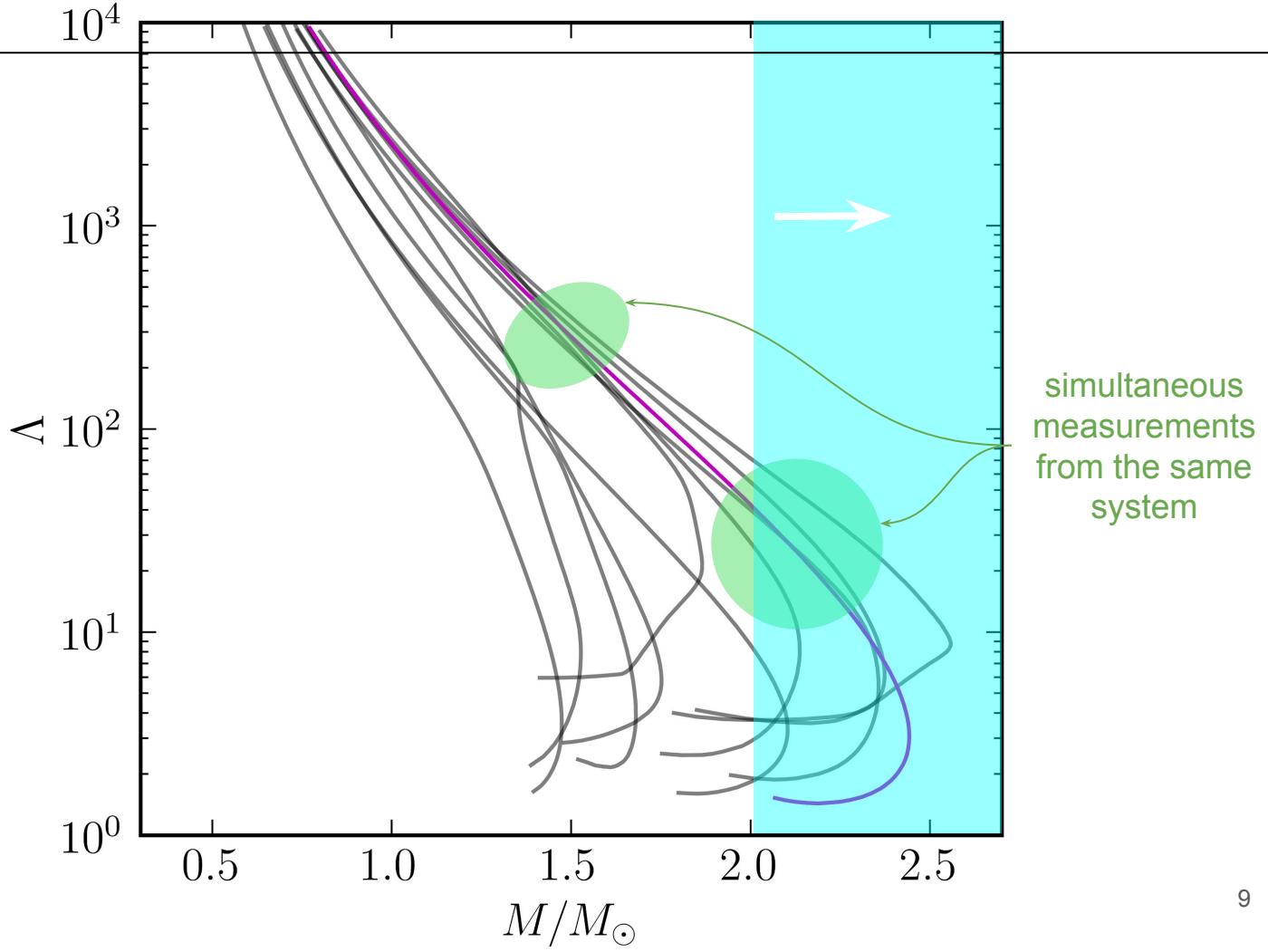
What do we observe?



What do we observe?



What do we observe?



Current Astrophysical Constraints on the High-Density, Cold Equation of State

P. Landry and R. Essick, *Nonparametric Inference of the Neutron Star Equation of State from gravitational Wave Observations*, PRD 99, 084049 (2019)

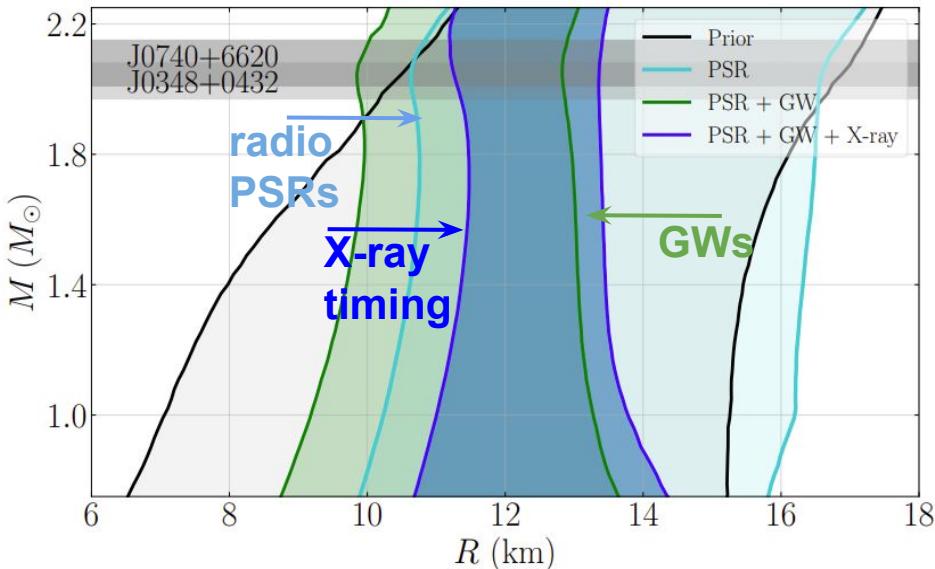
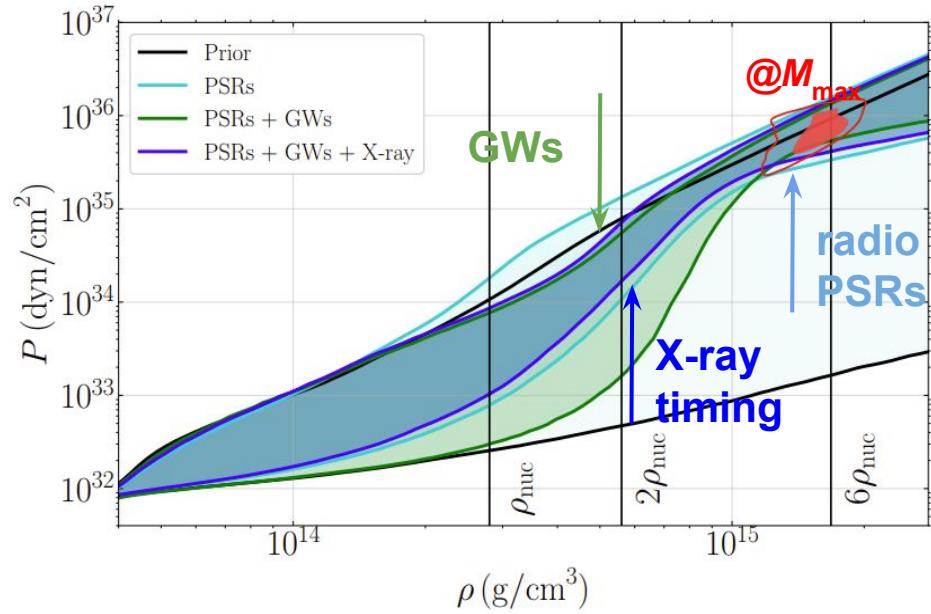
R. Essick, P. Landry, and D. E. Holz, *Nonparametric Inference of Neutron Star Composition, Equation of State, and Maximum Mass with GW170817*, PRD 101, 063007 (2020)

P. Landry, R. Essick, and K. Chatzioannou, *Nonparametric Constraints on Neutron Star Matter with Existing and Upcoming Gravitational Wave and Pulsar Observations*, PRD 101, 123007 (2020)

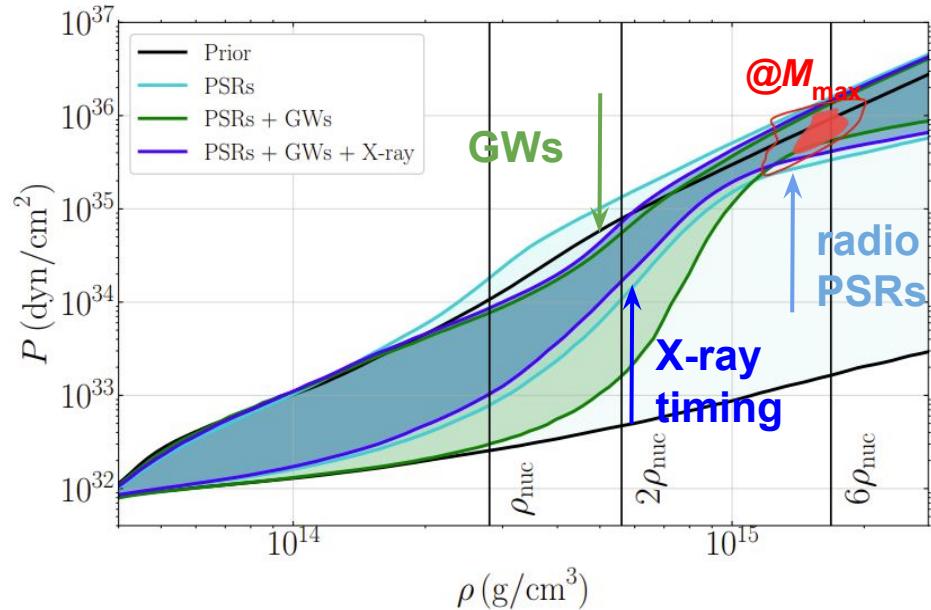
I. Legred, K. Chatzioannou, R. Essick, S. Han, and P. Landry, *Impact of the PSR J0740+6620 Radius Constraint on the Properties of High-Density Matter*, PRD 104, 063003 (2021)

I. Legred, K. Chatzioannou, R. Essick, and P. Landry, *Implicit correlations within phenomenological parametric models of the neutron star equation of state*, PRD 105, 043016 (2022)

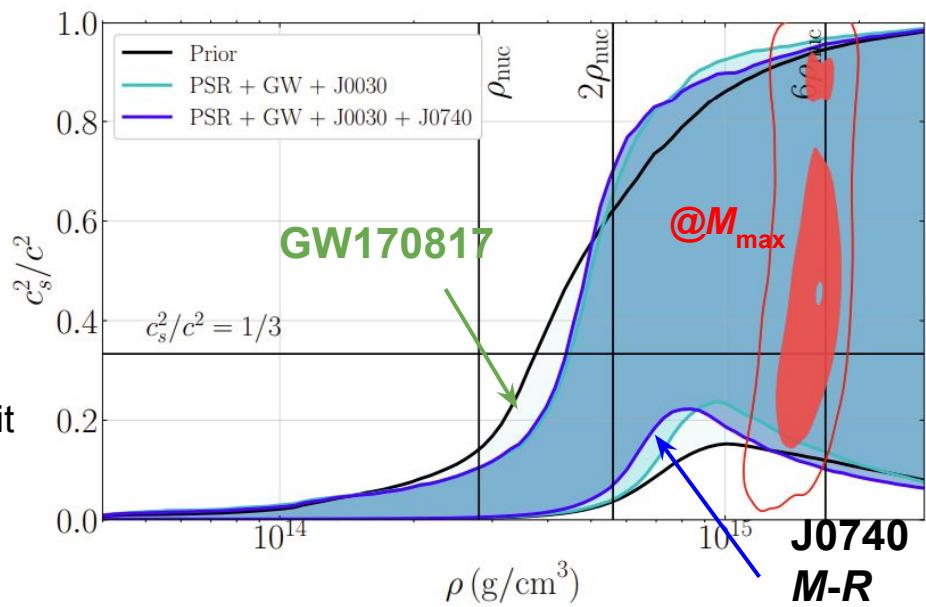
Current Astrophysical Constraints on the High-Density, Cold Equation of State



Current Astrophysical Constraints on the High-Density, Cold Equation of State

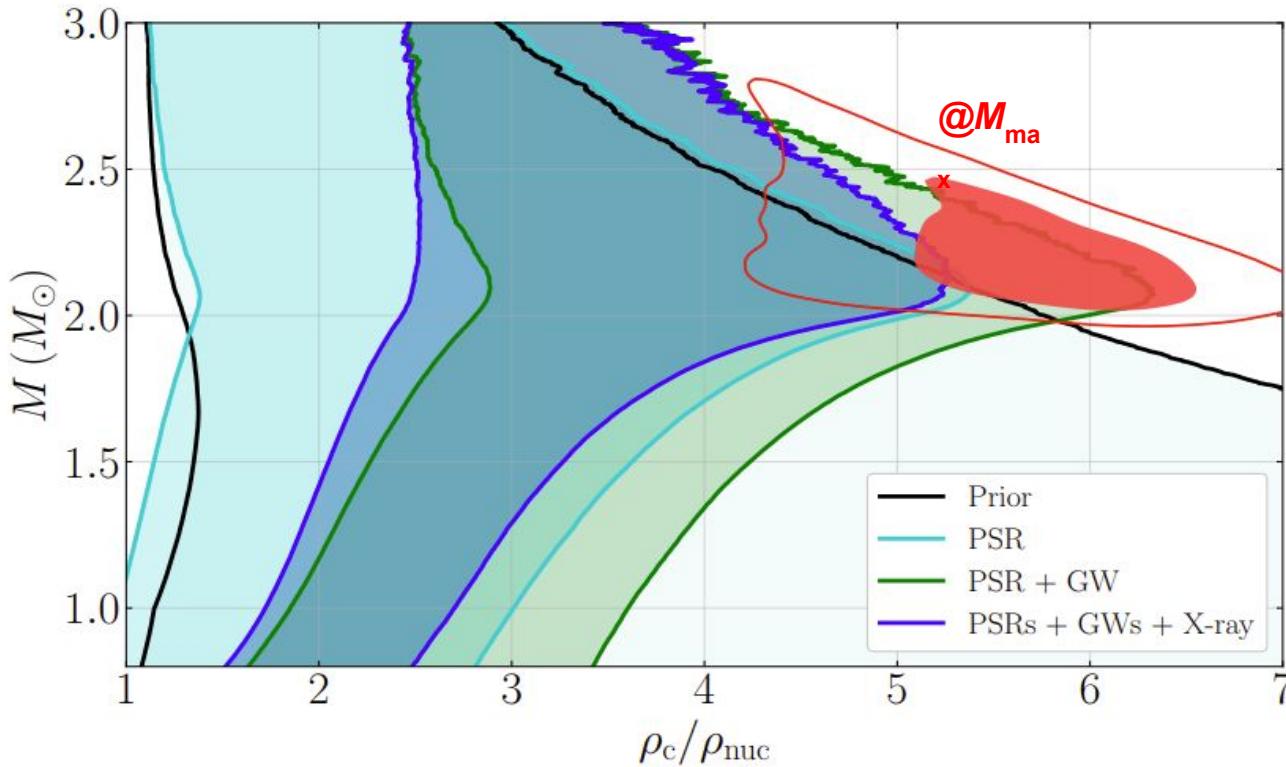


sound speed almost certainly exceeds the conformal limit
(suggests strongly-coupled interactions)

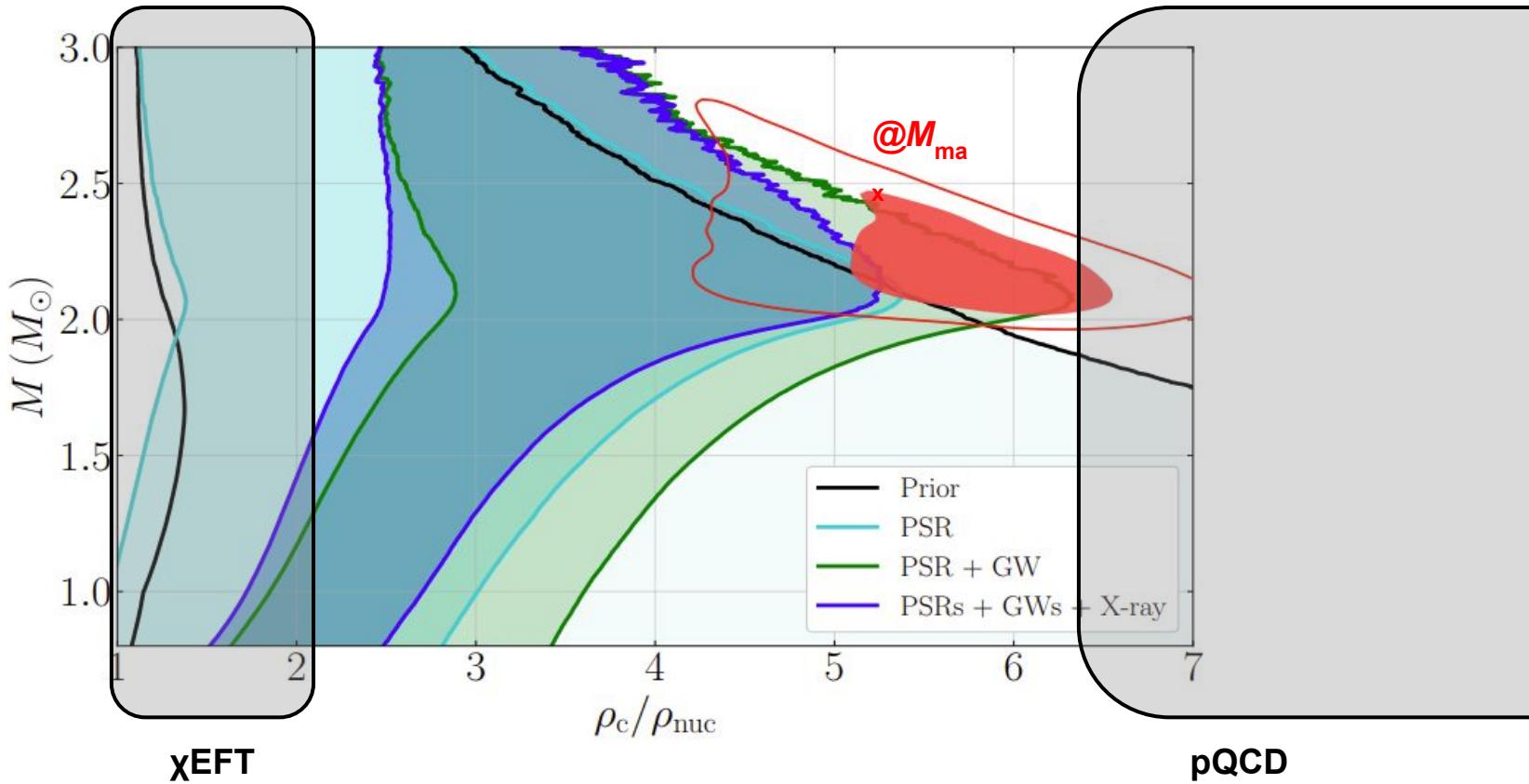


Current Astrophysical Constraints on the High-Density, Cold Equation of State

There is a limit to the densities we can probe within NSs



Current Astrophysical Constraints on the High-Density, Cold Equation of State



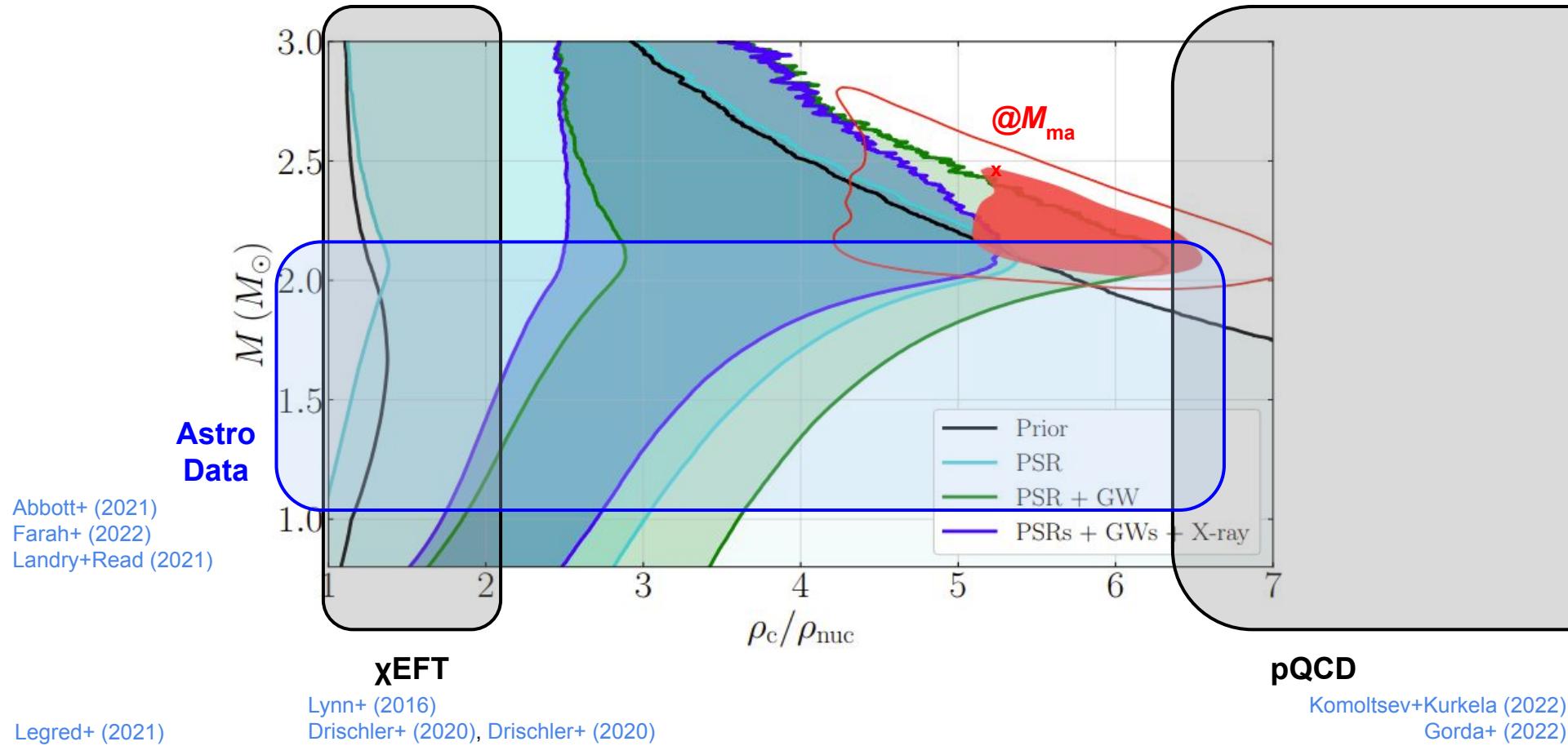
Lynn+ (2016)

Drischler+ (2020), Drischler+ (2020)

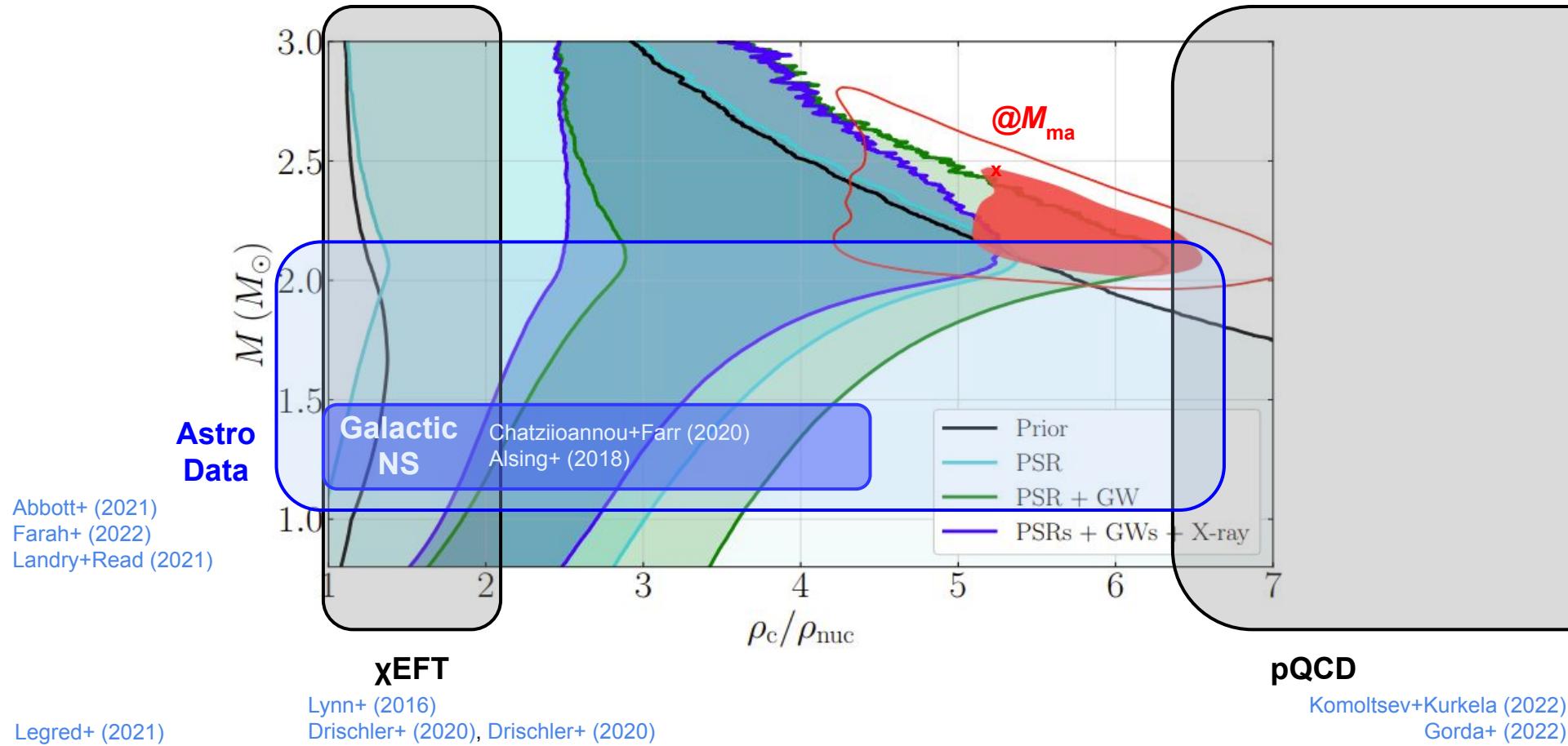
Komoltsev+Kurkela (2022)

Gorda+ (2022)

Current Astrophysical Constraints on the High-Density, Cold Equation of State



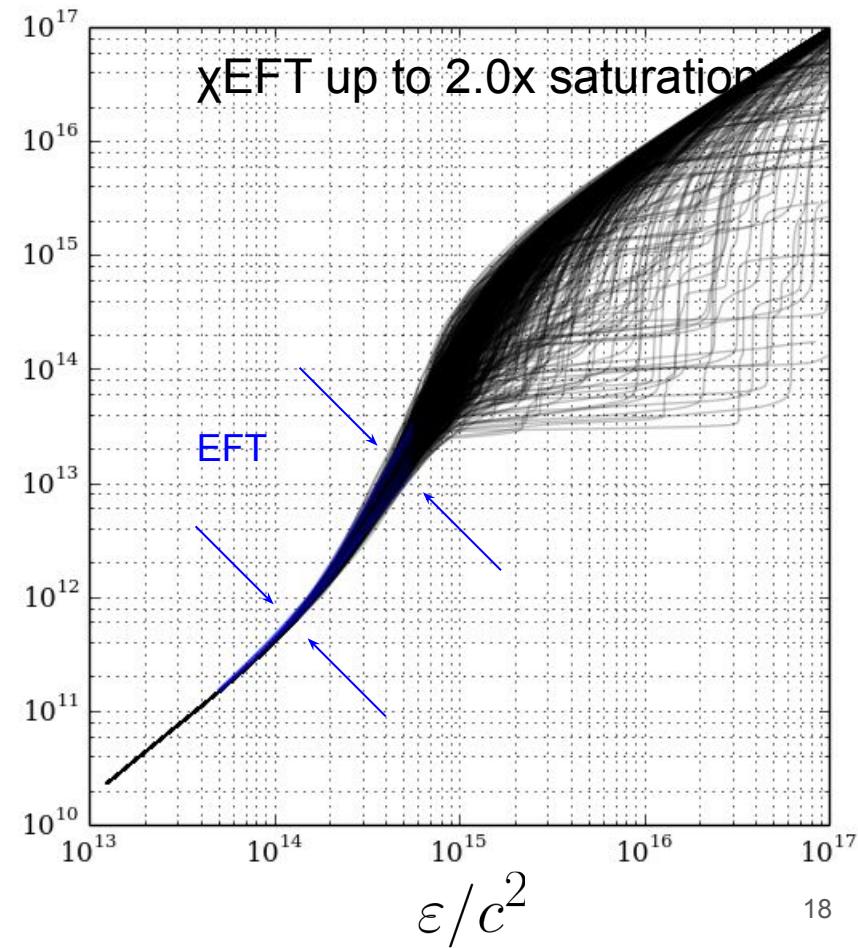
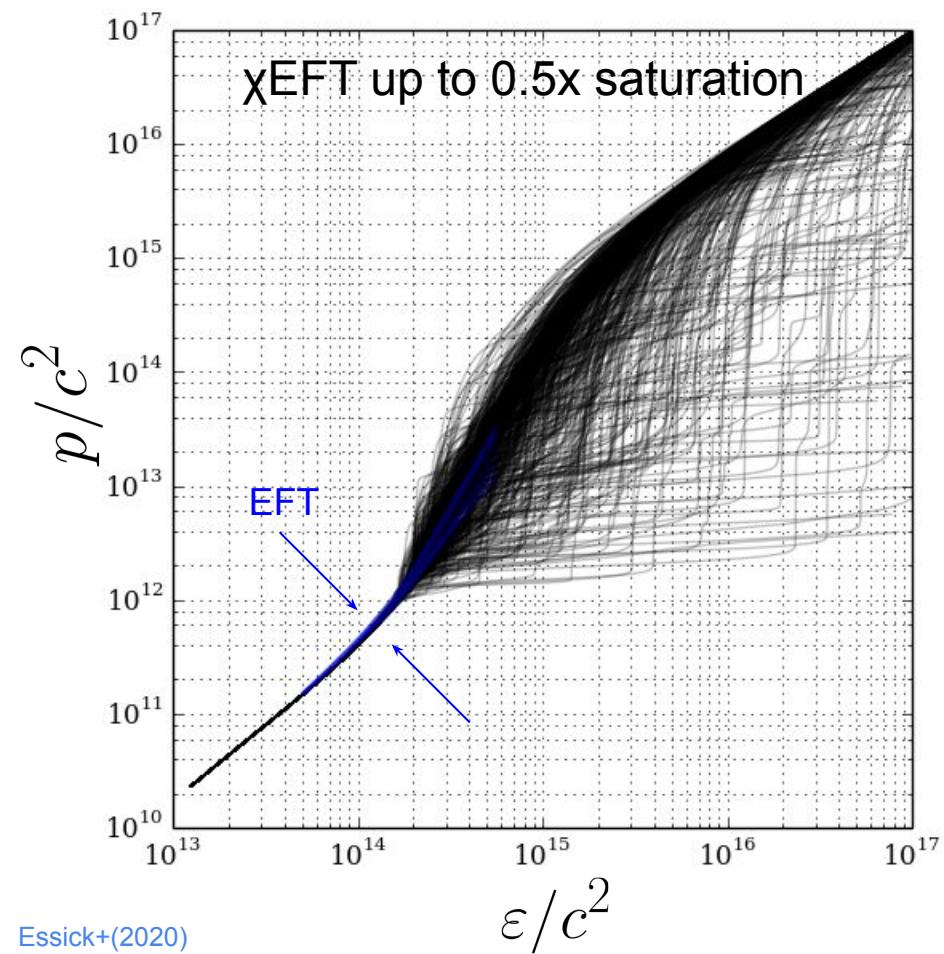
Current Astrophysical Constraints on the High-Density, Cold Equation of State



Comparisons with Terrestrial Nuclear Experiments and *ab initio* Theory

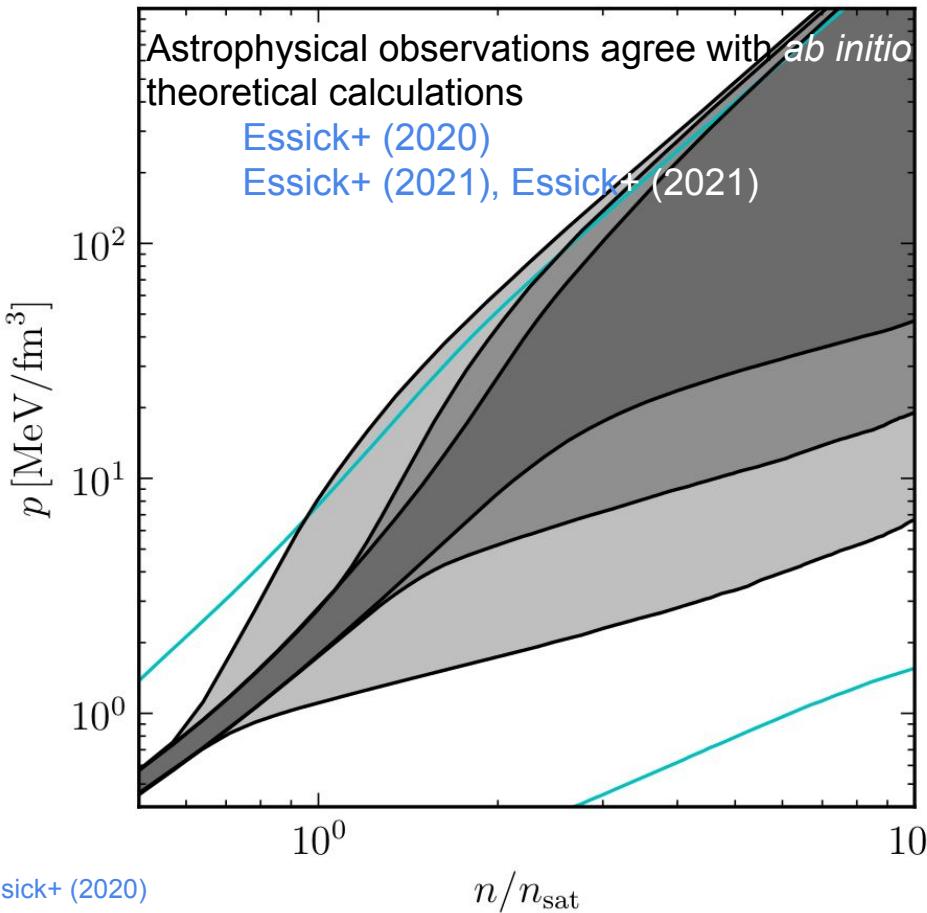
- R. Essick, I. Tews, P. Landry, S. Reddy, D. E. Holz, *Direct Astrophysical Tests of Chiral Effective Field Theory at Supranuclear Densities*, PRC 102, 055803 (2020)
- R. Essick, I. Tews, P. Landry, A. Schwenk, *Astrophysical Constraints on the Symmetry Energy and the Neutron Skin of ^{208}Pb with Minimal Modeling Assumptions*, PRL 127, 192701 (2021).
- R. Essick, I. Tews, P. Landry, A. Schwenk, A Detailed Examination of Astrophysical Constraints on the Symmetry Energy and the Neutron Skin of ^{208}Pb with Minimal Modeling Assumptions**, PRC 104, 065804 (2021)

Comparisons with Terrestrial Nuclear Experiments and *ab initio* Theory

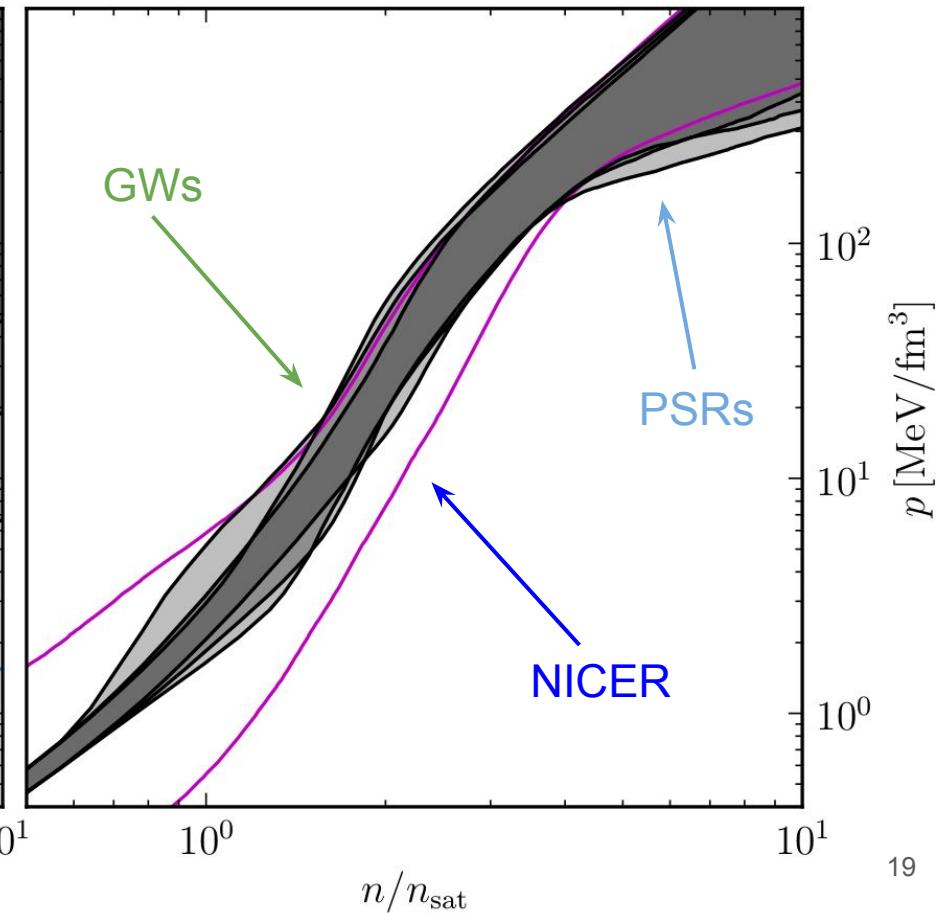


Comparisons with Terrestrial Nuclear Experiments and *ab initio* Theory

Priors

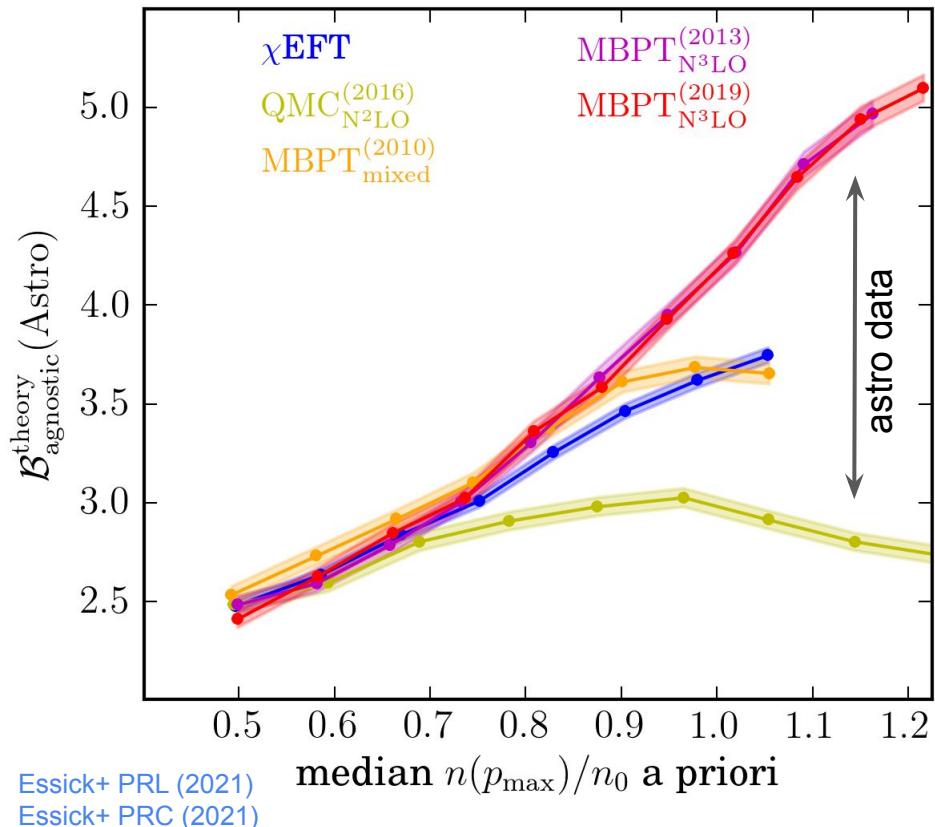


Posteriors with PSRs+GWs+NICER

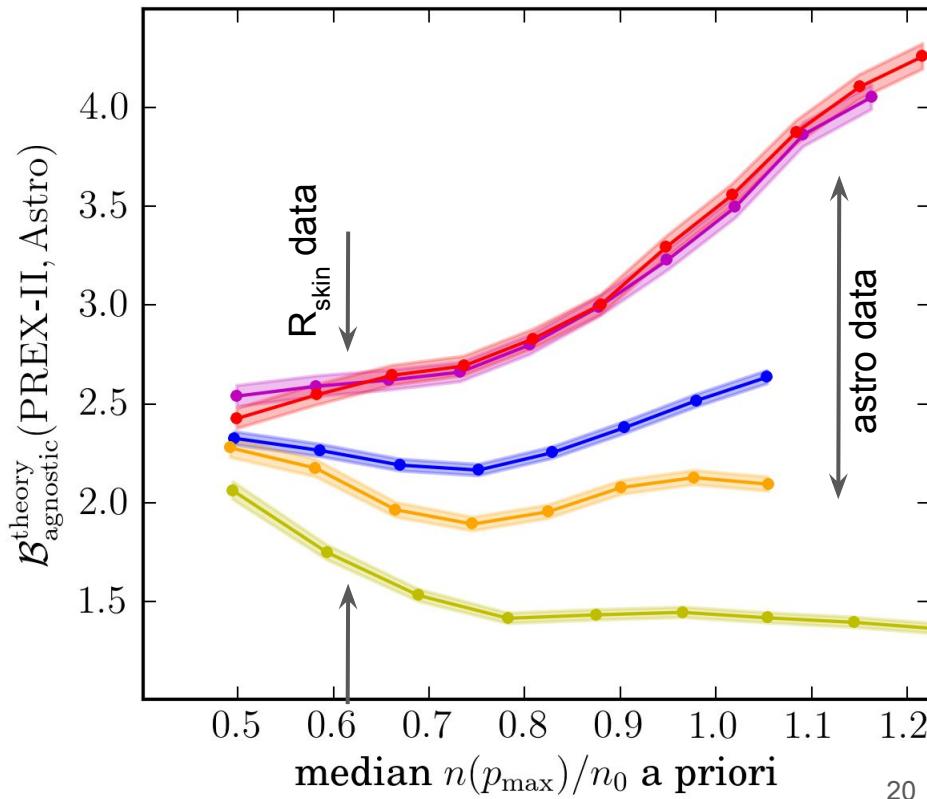


Comparisons with Terrestrial Nuclear Experiments and *ab initio* Theory

astro data can distinguish between nuclear theories at high densities



nuclear experiments probe lower densities



Comparisons with Terrestrial Nuclear Experiments and *ab initio* Theory

Nonparametric Prior

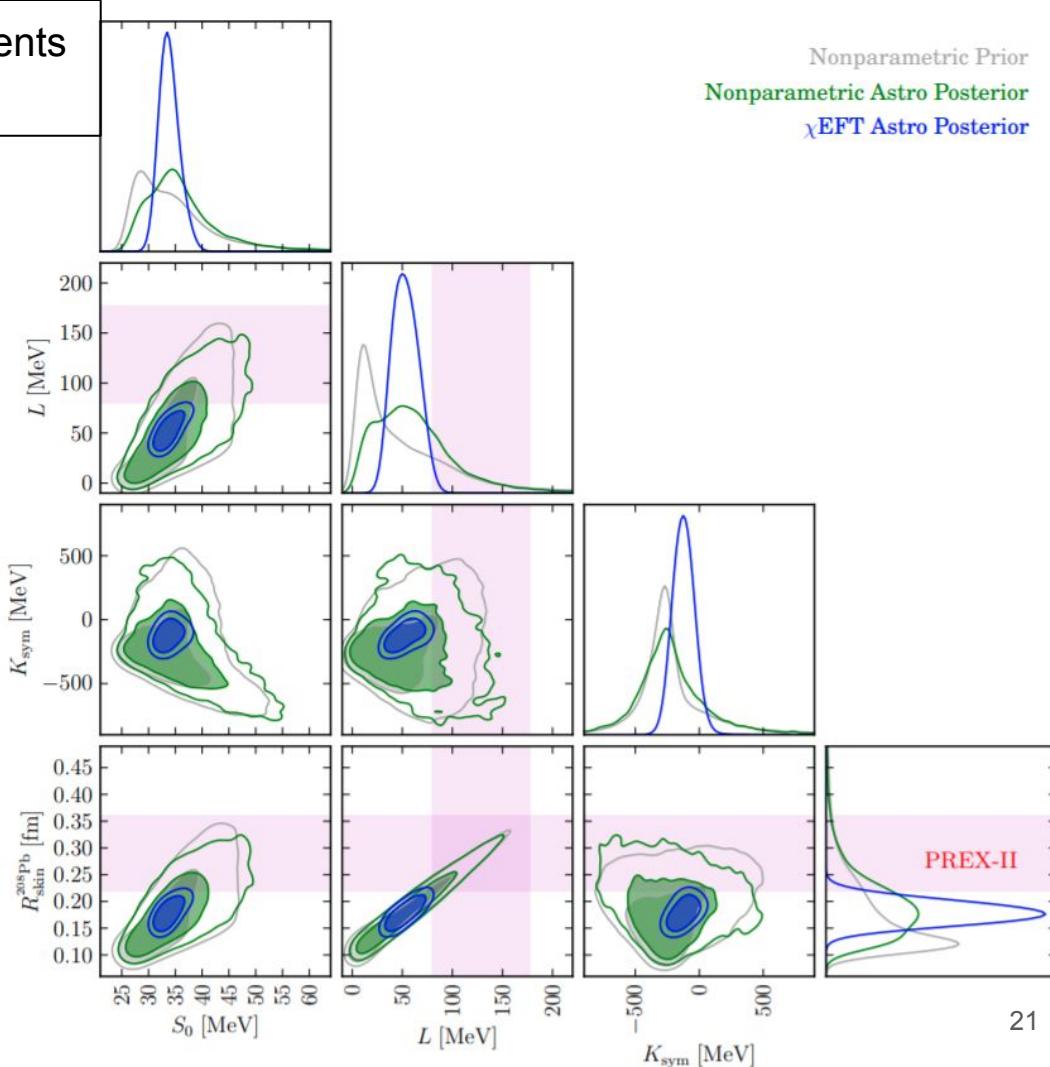
Nonparametric Astro Posterior

χ EFT Astro Posterior

Comparison to “theory marginalized” χ EFT calculations and (preliminary) PREX-II measurement

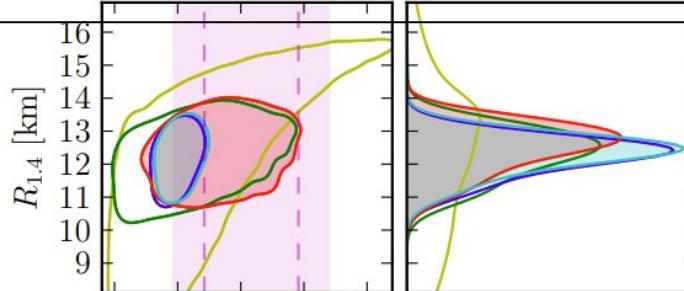
- χ EFT calculations fall near the maximum *a posteriori* of our agnostic nonparametric analysis
- Incorporating PREX (Gaussian) likelihood shifts posteriors towards larger R_{n-p} , L .
- **No significant tension between PREX, χ EFT, and/or agnostic astrophysical posterior**

Reed+(2021) suggest there might be some tension. We find this is because they use a single model class (RMF) to high densities, which may not be warranted.



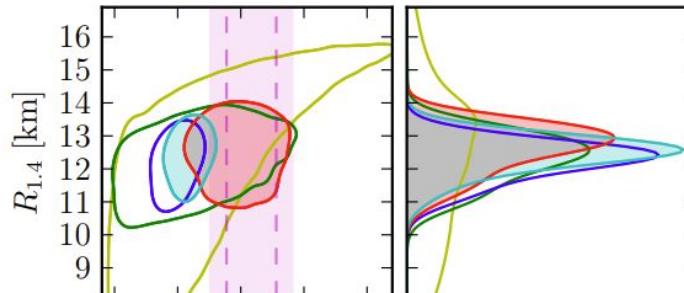
Comparisons with Terrestrial Nuclear Experiments and *ab initio* Theory

current R_{skin} uncertainty

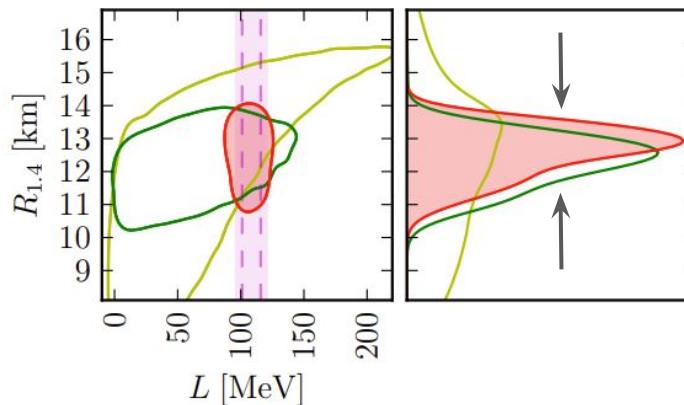


nonparametric prior
nonparametric astro-only posterior
XEFT+astro posterior
nonparametric astro+ R_{skin} posterior
XEFT+astro+ R_{skin} posterior

R_{skin} uncertainty improved by a factor of 2



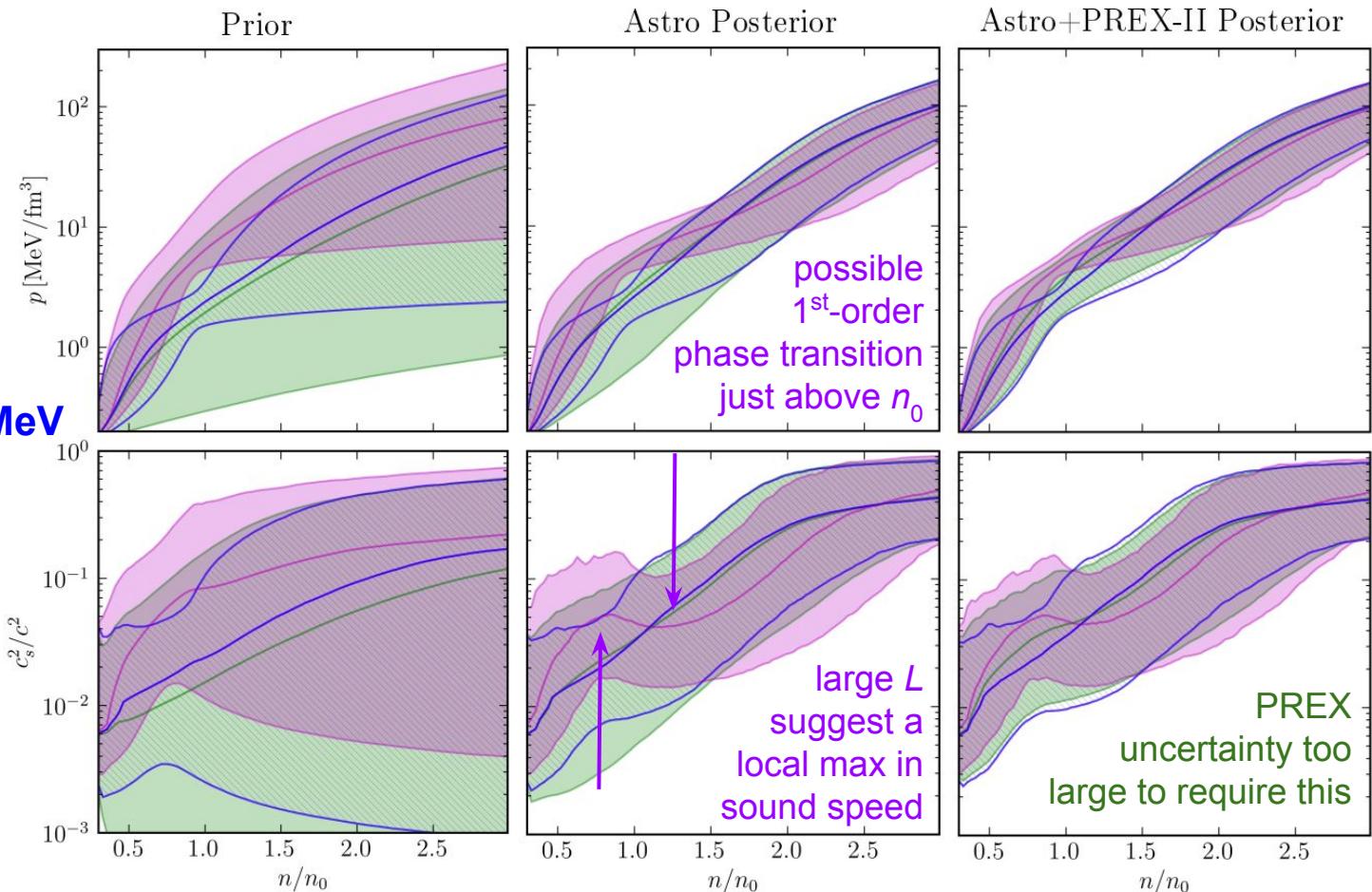
hypothetical perfect R_{skin} measurement



improved precision in nuclear experiments is unlikely to affect our knowledge of NS radii without improved theoretical calculations

Comparisons with Terrestrial Nuclear Experiments and *ab initio* Theory

100 MeV $< L$
30 MeV $< L < 70$ MeV
All L



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