

# The ANDROMeDa Project

Searching for Dark Matter with Vertically-Aligned Carbon Nanotubes

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**ANDROMeDa**

Aligned Nanotube Detector for Research On MeV Darkmatter



**ROMA  
TRE**  
UNIVERSITÀ DEGLI STUDI

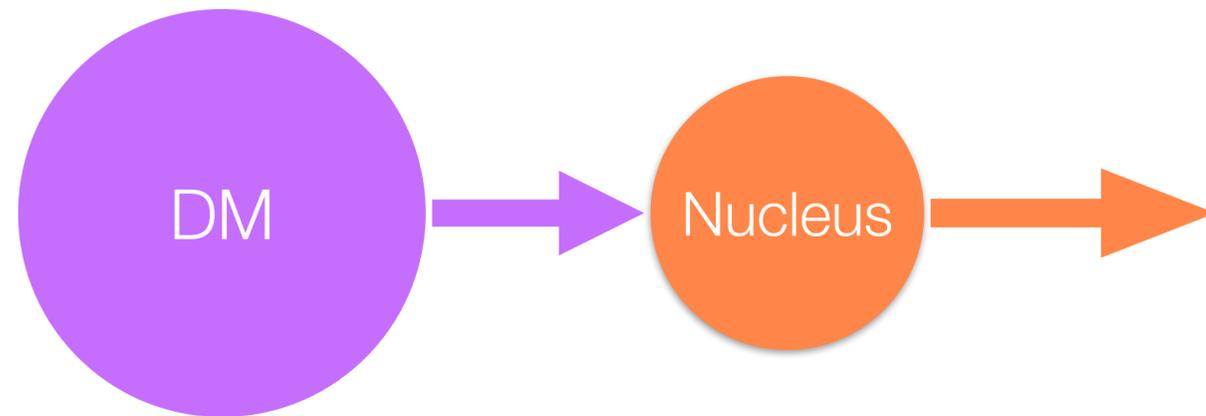


Istituto Nazionale di Fisica Nucleare



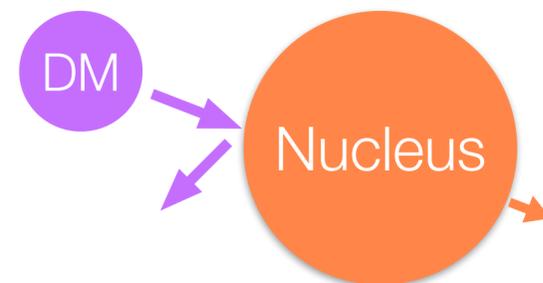
# Poor Sensitivity for Dark Matter in Sub-GeV Range

▶ Current experiments mainly based on **nuclear recoil**

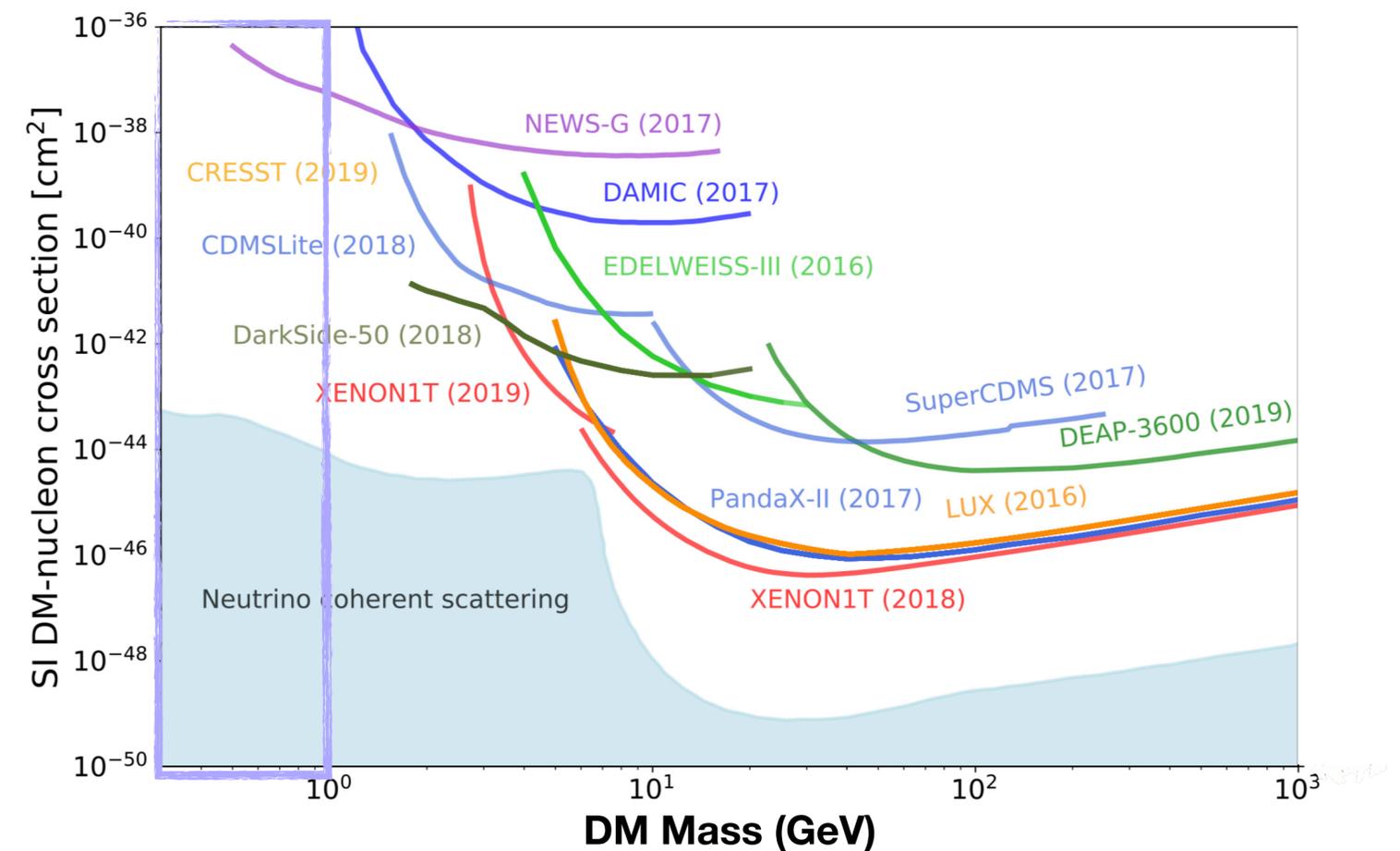


▶ If Dark Matter (DM) mass  $< \text{GeV}$

- no visible nuclear recoil



- **lighter target needed**





# Less Strict Limits Using Electron Recoil

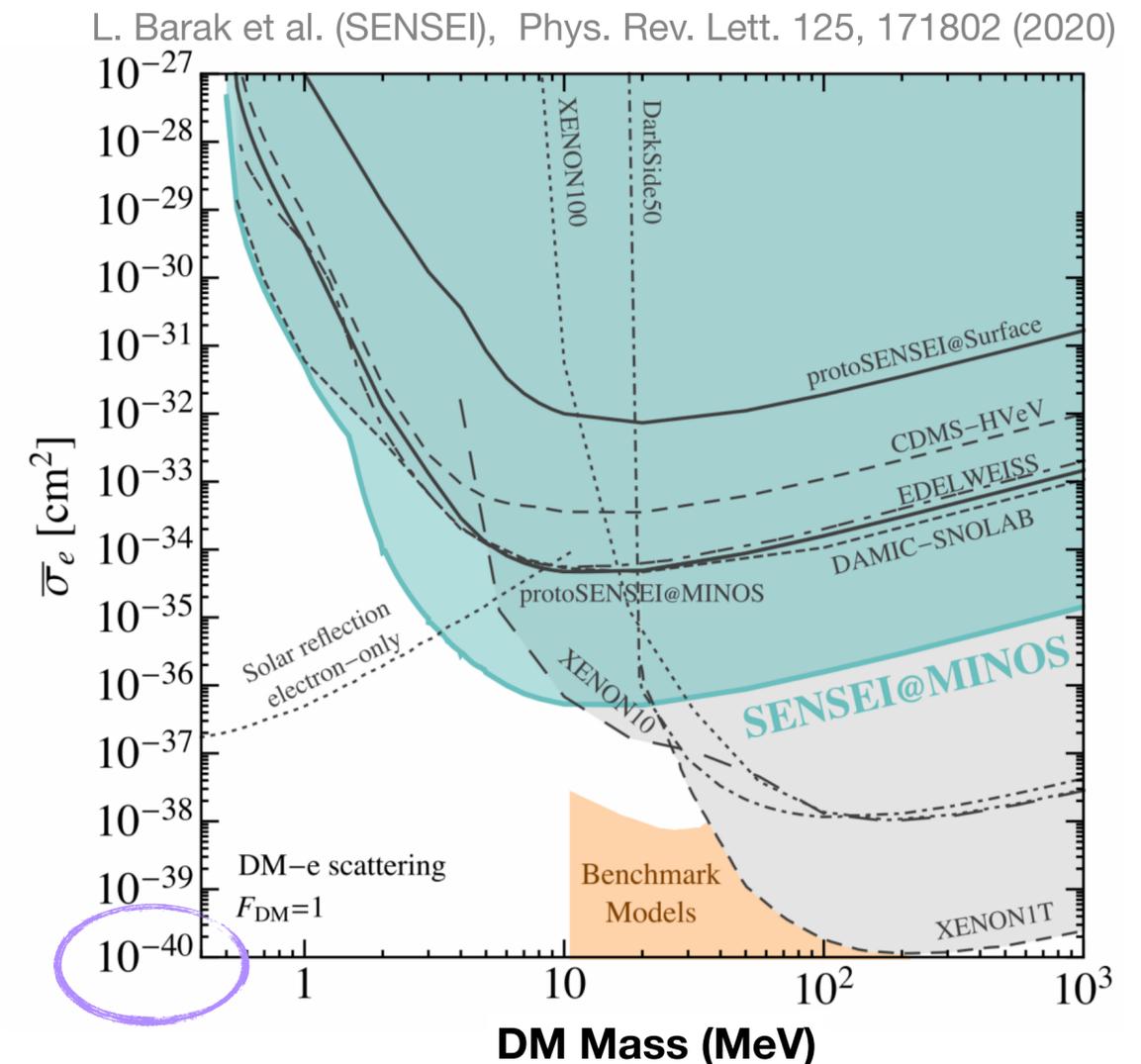
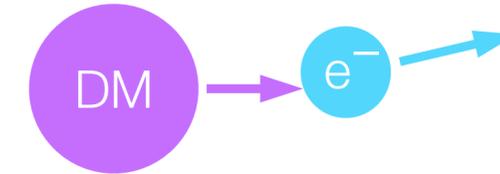
▶ For Sub-GeV DM better look for **DM-electron scattering**

- visible also with few MeV DM mass

▶ Few experiments sensitive to electron recoil

- cross section limits  $10^6$  times weaker
- sensitivity drop for mass  $< 100$  MeV

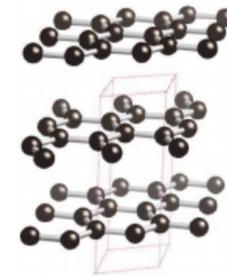
▶ **New detection strategy** needed to go below 100 MeV



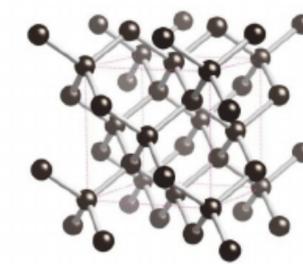


# Electrons Directly Into Vacuum: 2D Materials

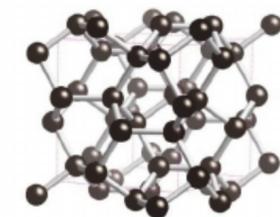
▶ Assuming  $\begin{cases} v_{DM} = 300 \frac{km}{s} \\ m_{DM} = 10 - 100 \text{ MeV} \end{cases} \rightarrow K_{DM} = 5 - 50 \text{ eV}$



graphite



diamond



BC8

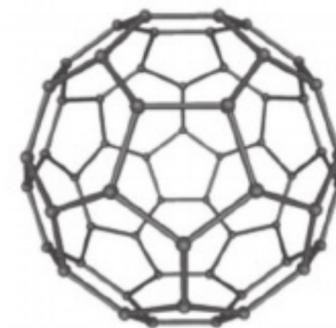
▶ Able to **extract electrons** from carbon ( $\Phi_C = 4.7 \text{ eV}$ )

▶ Low energy electrons = extremely **short range** in matter

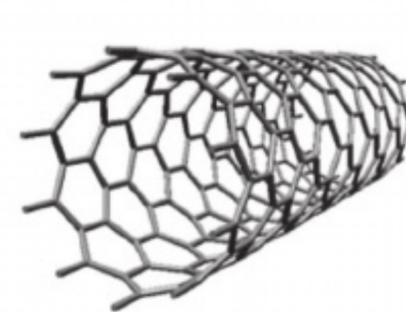
▶ Problem solved using **2D materials**:

- electrons directly ejected into vacuum
- no additional energy loss

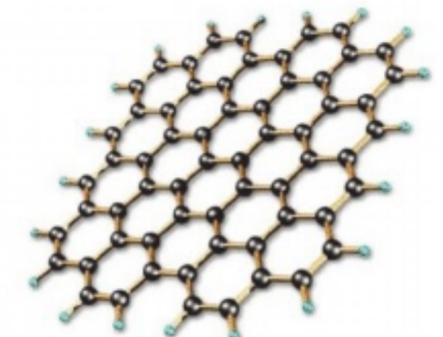
↓ Low dimension



fullerene



nanotube

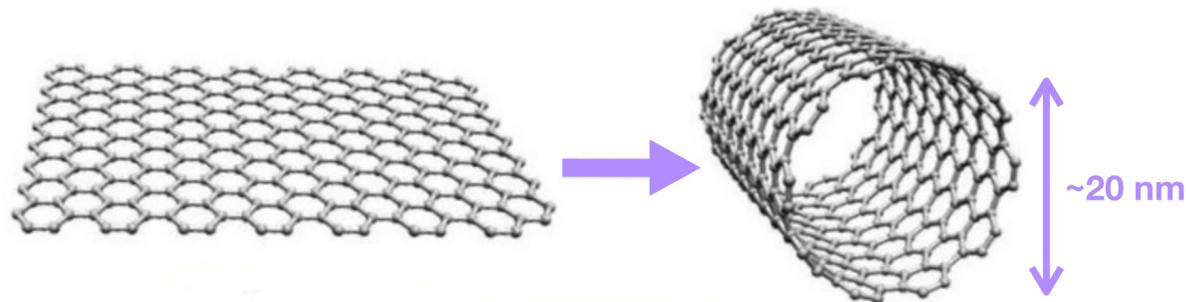


graphene



# Growing Carbon Nanotubes Forests

## ▶ Idea:

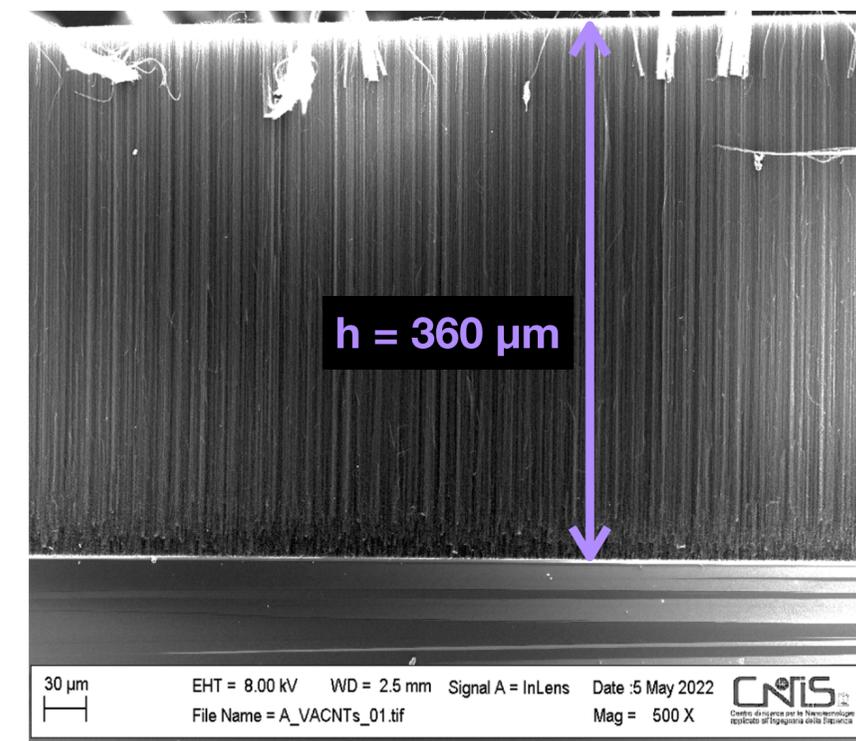
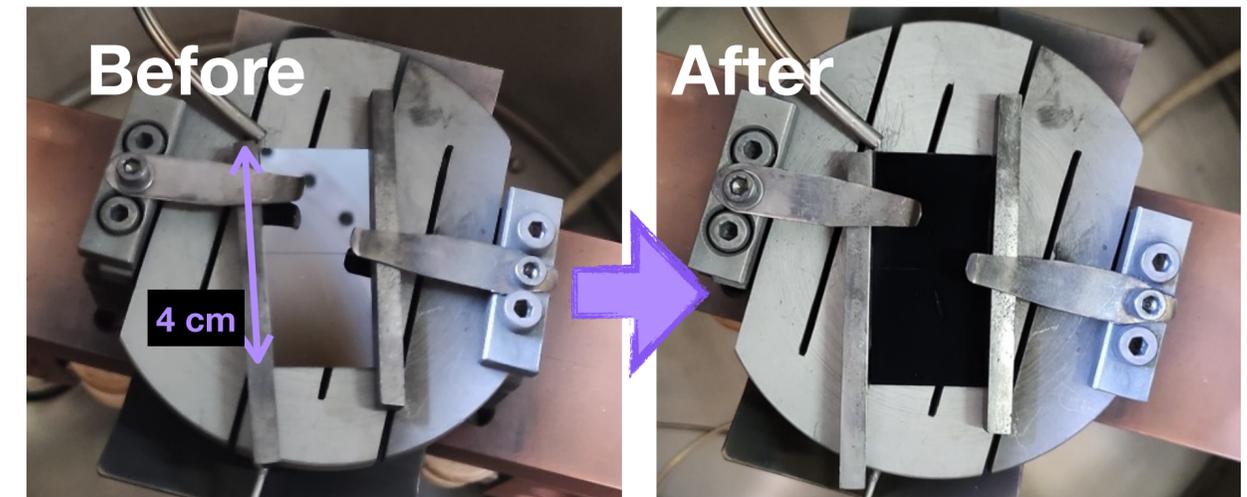


## ▶ State-of-the-art nanotube facility in Rome Sapienza

- Chemical Vapour Deposition (CVD) technique
- Up to 8 cm<sup>2</sup> extension on various substrates
- Diameter ~20 nm, length up to 400 μm

## ▶ Result: vertically-aligned nanotubes forests

- Ideal target for DM search?



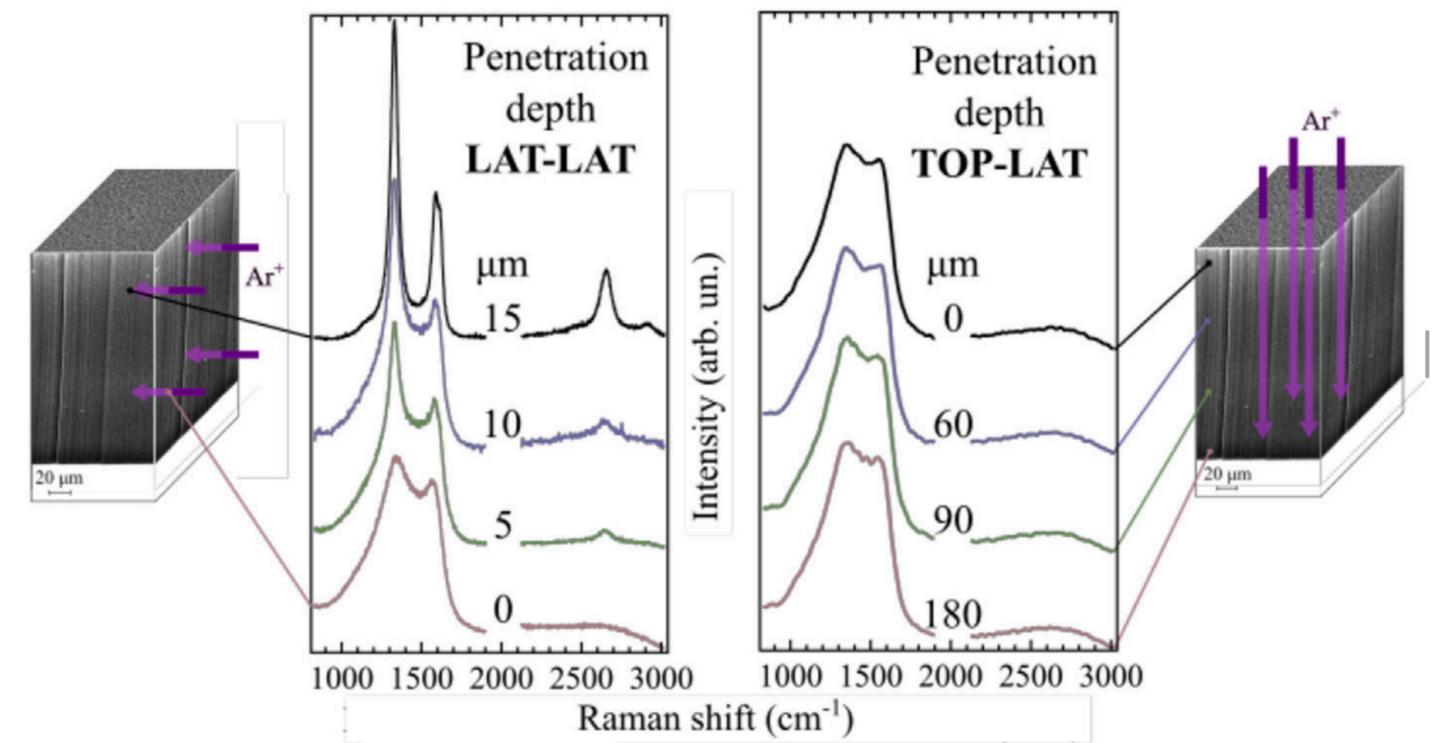


# Directional Sensitivity with Carbon Nanotubes

G. D'Acunto, et al., Carbon 139 (2018) 768

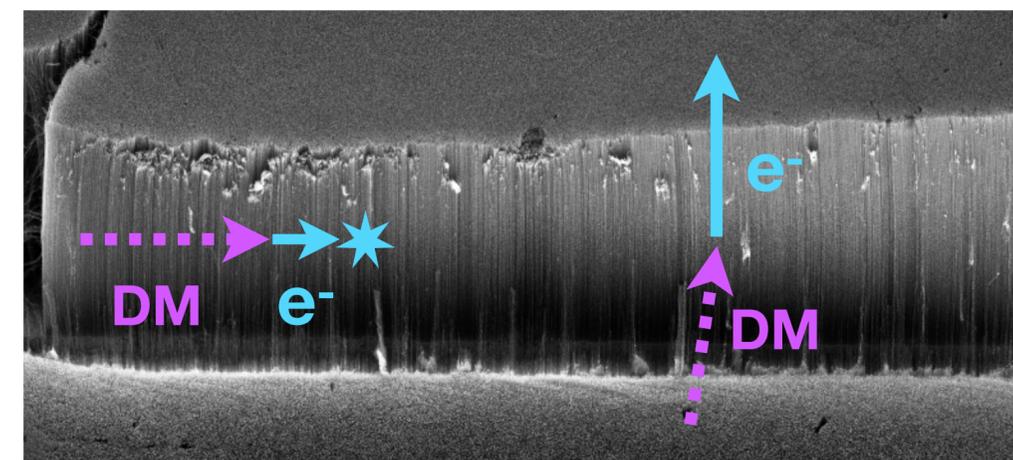
## ▶ Raman analysis after Ar<sup>+</sup> bombardment

- Lateral penetration < 15 μm
- Longitudinal damage along full length (180 μm)
- **Highly anisotropic density**



## ▶ CNTs as target for DM-electron scattering

- ~vanishing density in tube axis direction
  - electron ejected **only if parallel to tubes**





# A New Detector Concept: The Dark-PMT

## ▶ Working principle:

- DM-electron scattering on a target of VA-CNTs
- Electrons out if tubes parallel to the DM wind
- Acceleration up to keV
- Detection by silicon sensor

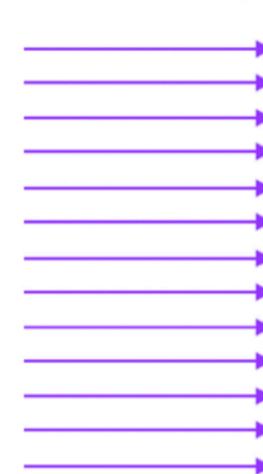
## ▶ Key features:

- ✓ Directional Sensitivity
- ✓ Sensitive to few eV electrons
- ✓ ~Unaffected by thermal noise ( $\Phi_C = 4.7$  eV) even at room temperature

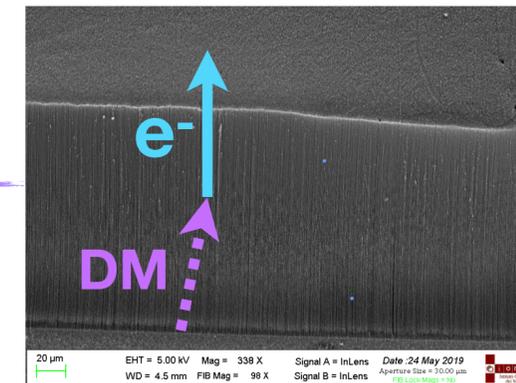
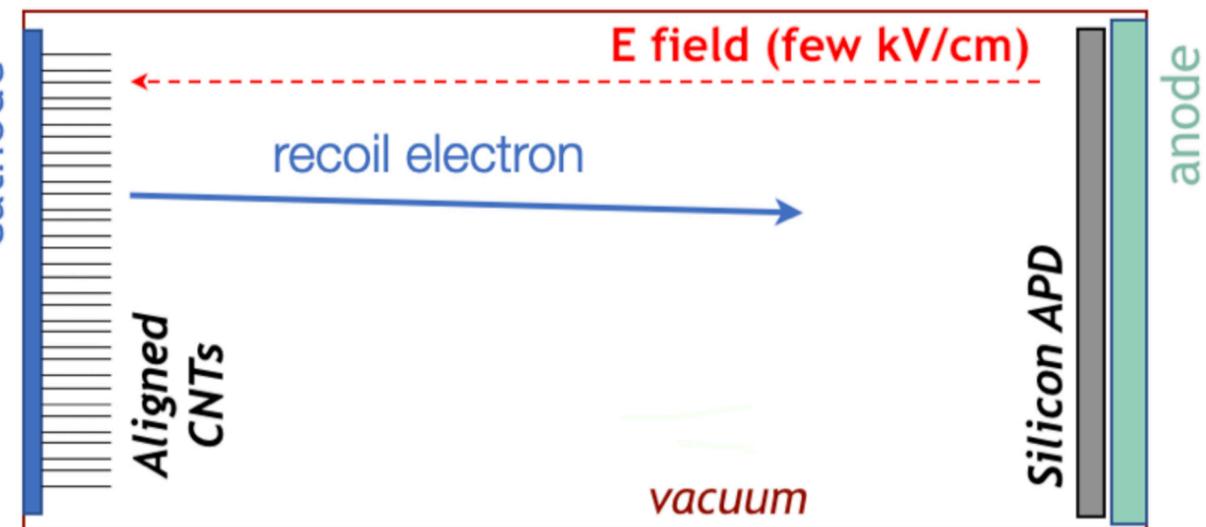


Cygnus

DM Wind



cathode





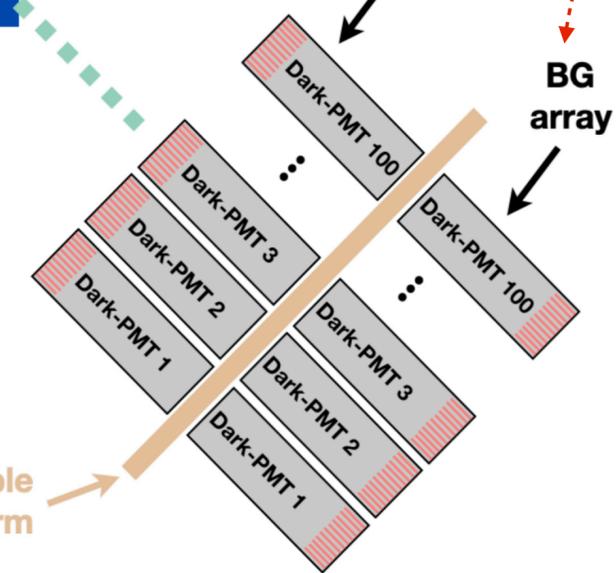
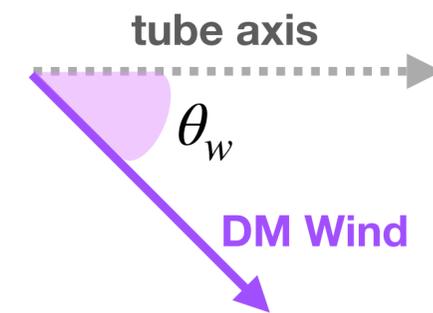
# Dark Matter Search with 2 Dark-PMT Arrays

## ▶ Looking at expected rate of electrons ejected from VA-CNTs

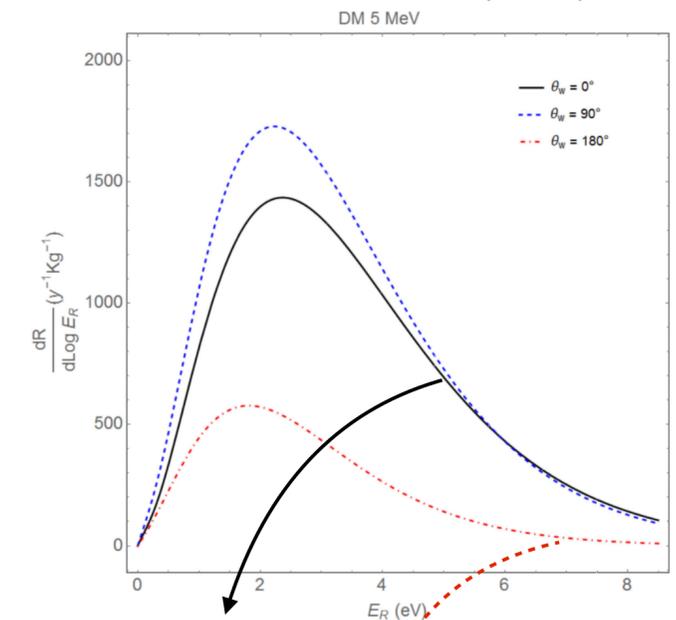
- rate for  $\theta_w = 0^\circ \gg$  rate for  $\theta_w = 180^\circ$
- counts excess if dark-PMT pointed in DM wind direction

## ▶ two arrays of dark-PMTs on a moving platform

- 1st pointed towards Cygnus  $\rightarrow$  **DM signal**
- 2nd in opposite direction  $\rightarrow$  **backgrounds**
- $\geq 1$  g mass for array so  $\sim 100$  units with  $10 \text{ cm}^2$  cathode area



G. Cavoto, et al., PLB 776 (2018) 338





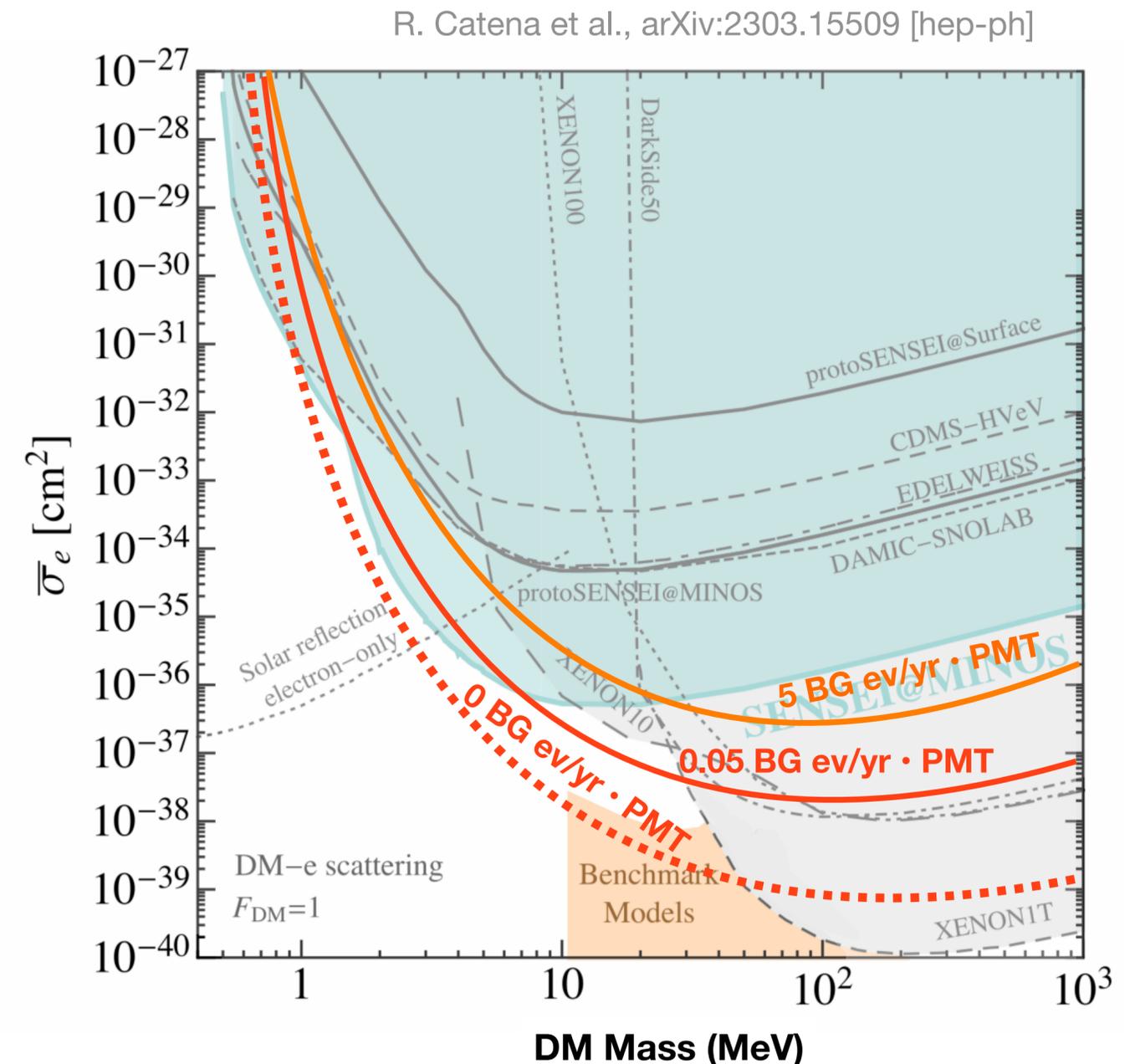
# Background Minimisation Will Be Essential

## ▶ Expected exclusion limits

- For just 1 g x 1 year exposure
- Using 2 arrays of 100 Dark-PMTs

## ▶ Performance **strongly depends** on BG event rate

- BG rate < 0.05 events/year x dark-PMT needed  
to **extend current limits**





# The ANDROMEaDa Project

▶ Awarded PRIN2020 Grant (1M€)



- 3 years project started in 2022
- 3 units: 

}	INFN (F. Pandolfi - P.I.)
	Sapienza (G. Cavoto)
	Roma 3 (A. Ruocco)

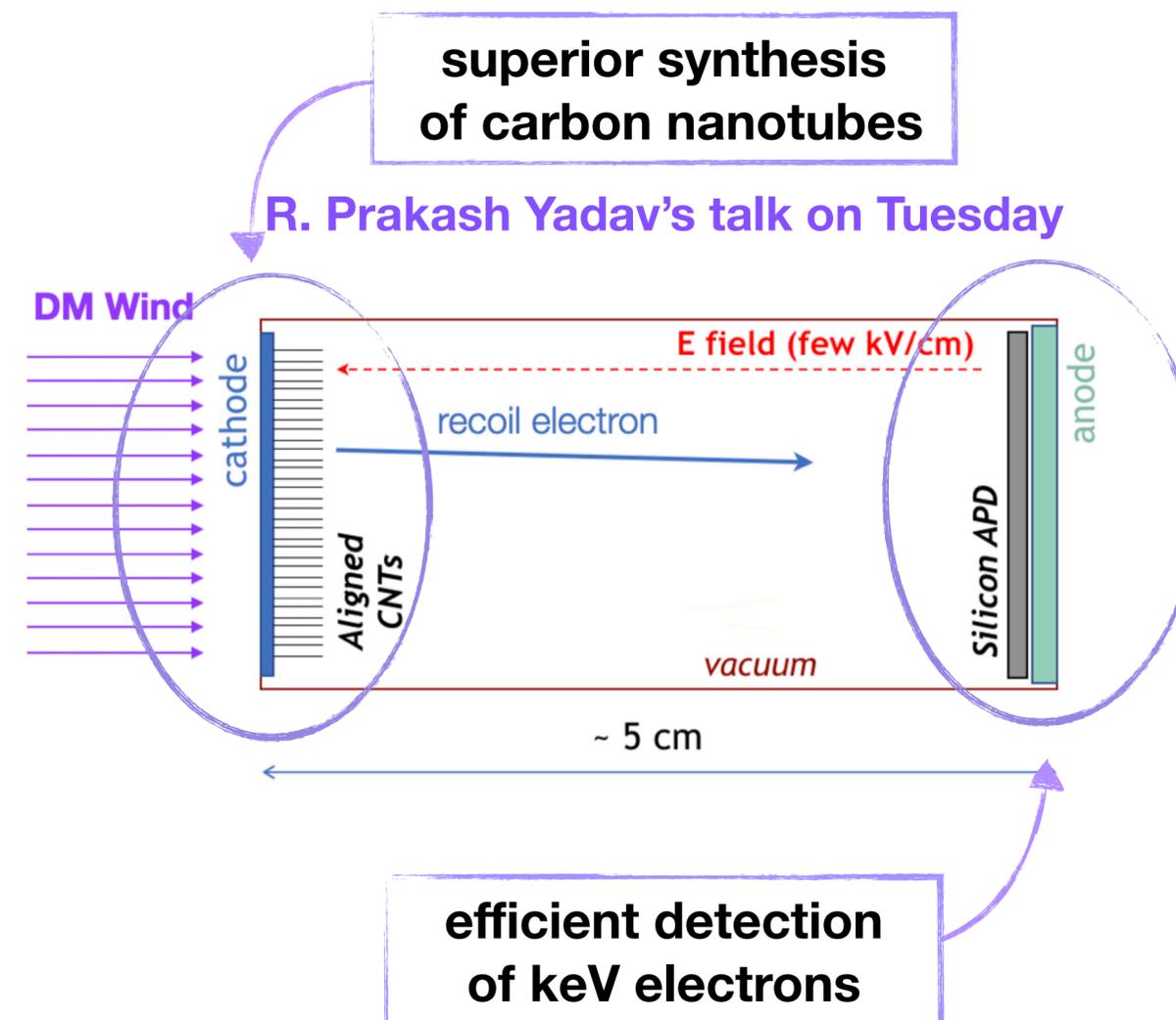
▶ **Main objective:** development & construction of a dark-PMT prototype

- **challenges on both sides** of detector



**ANDROMEDA**

Aligned Nanotube Detector for Research On MeV Darkmatter

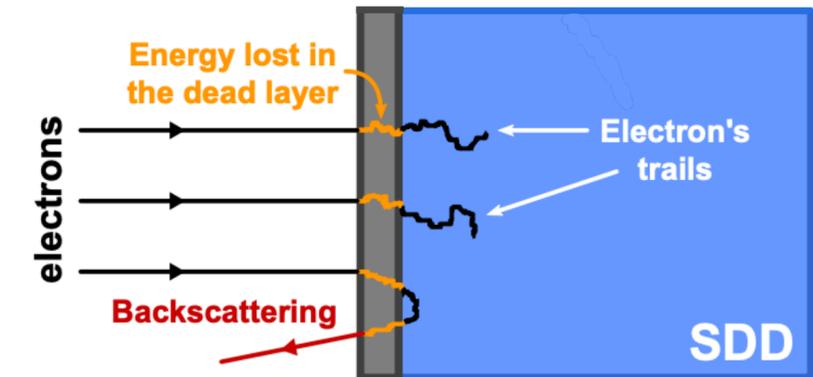




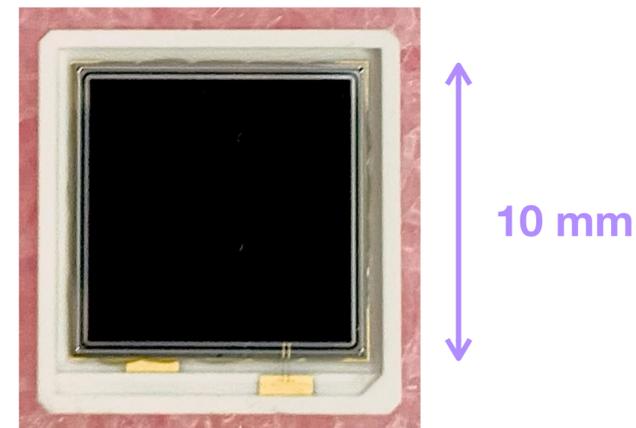
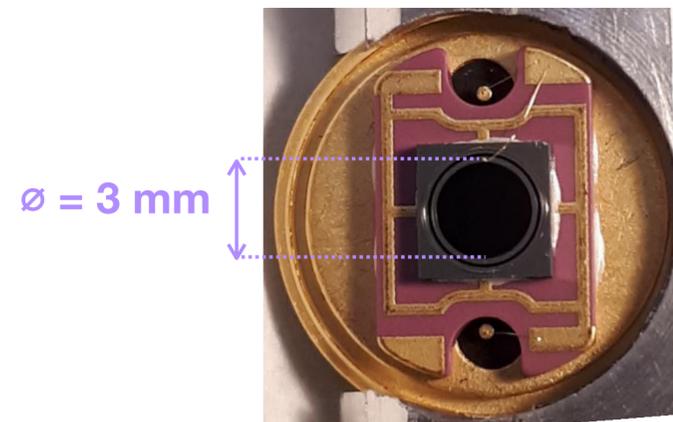
# Detecting keV Electrons with Silicon Detectors

## ▶ SDDs & APDs born as photon detectors

- with thick dead layer (Si oxidation) → able to **detect electrons**

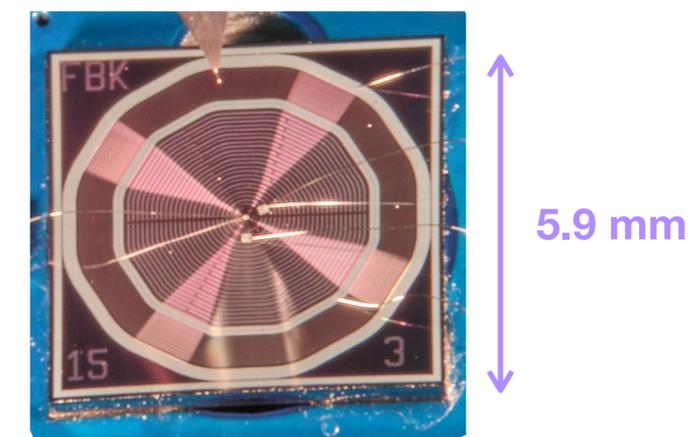


## ▶ Benchmark: **Windowless Avalanche Photodiodes**



- simple, cost-effective
- produced by Hamamatsu

## ▶ Backup: **Silicon Drift Detectors**



- ultimate energy resolution
- produced by FBK + electronics by PoliMi

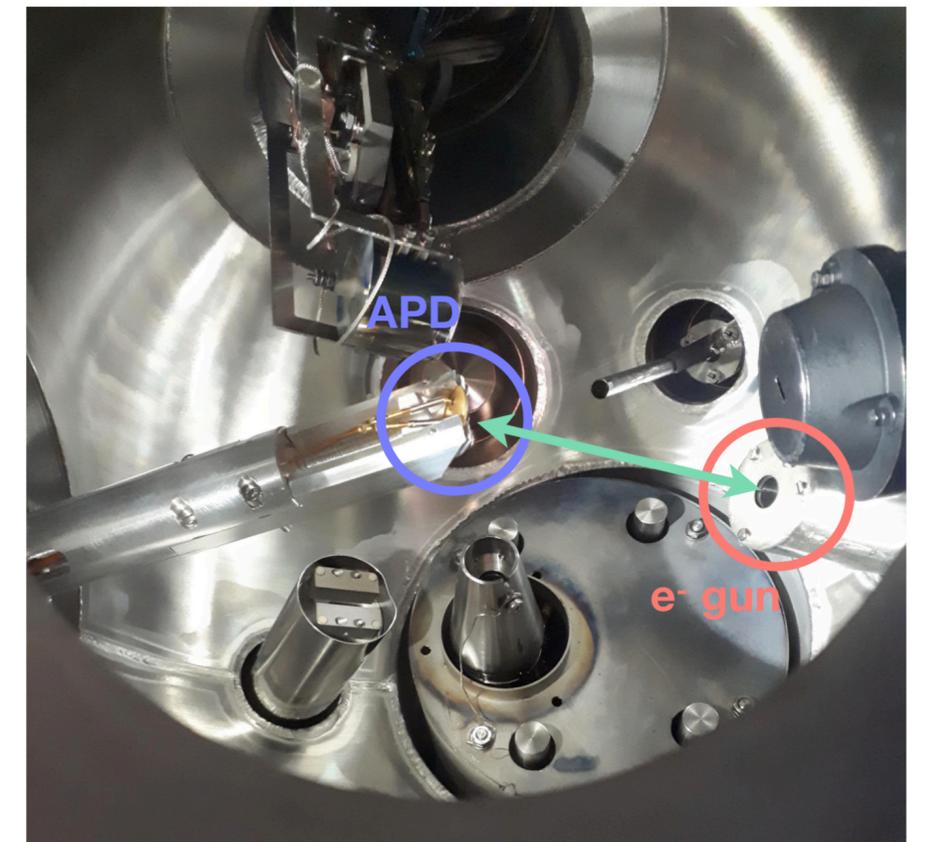


# APD Characterisation @ LASEC Labs (Roma Tre)

▶ Hot tungsten filament + electrostatic lenses

▶ **Key features:**

- Electron energy:  $30 < E < 1000$  eV
  - Energy uncertainty  $< 0.05$  eV
  - Beam spot  $\sim 0.5$  mm
  - Current as low as a few fA
- ☑ can probe **single electron regime**





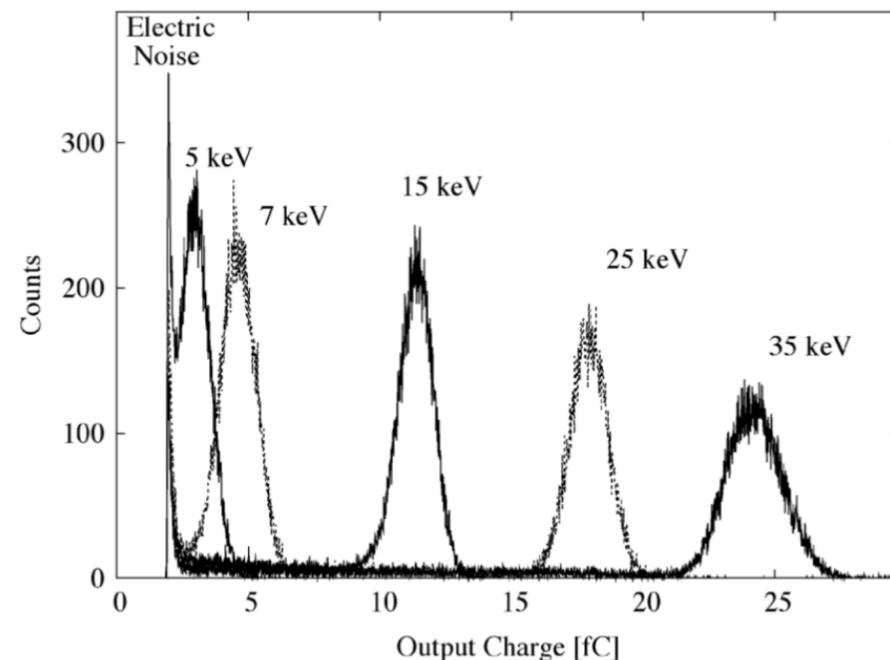
# Linearity Between Gun and APD Currents

## ▶ Reading APD bias current when shooting gun on it

- Clear **linear correlation** with gun current

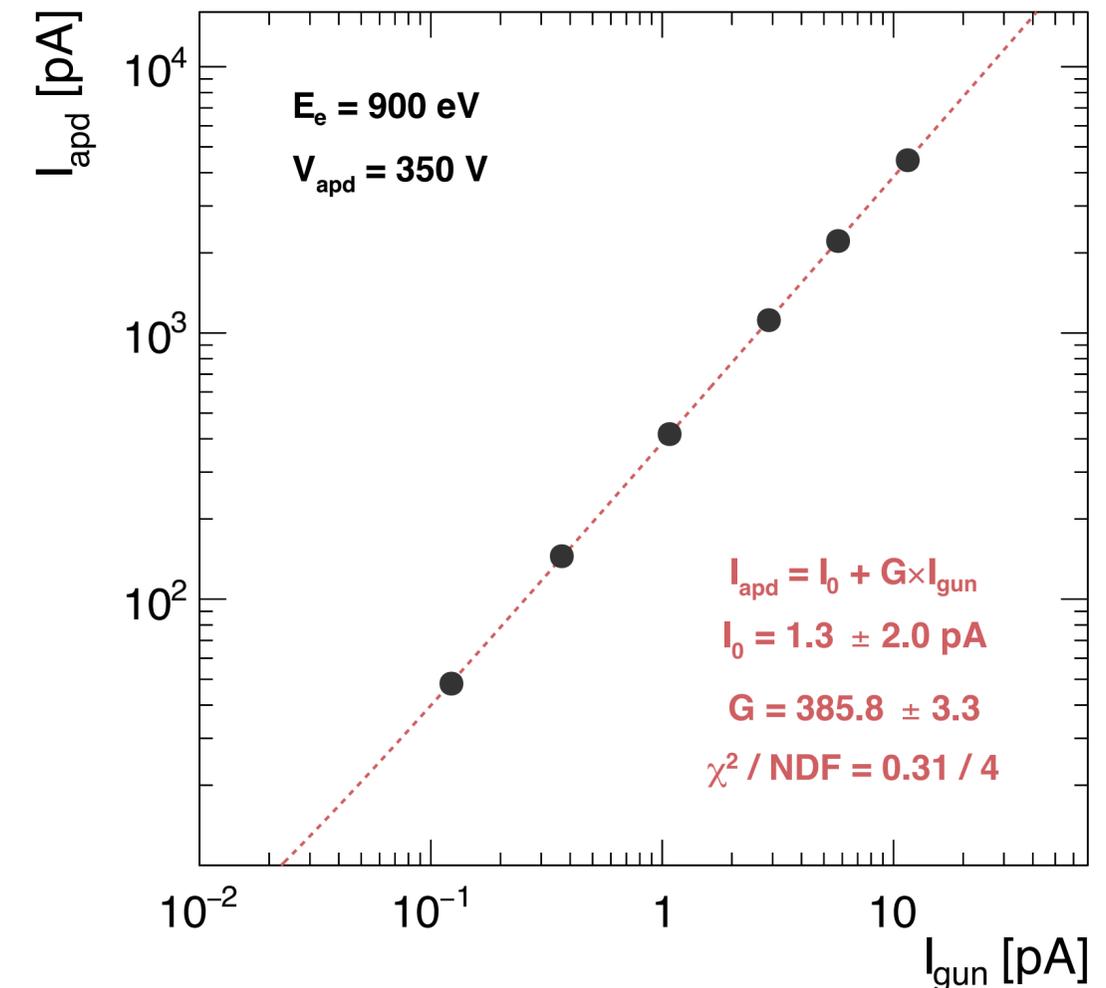
## ▶ Not able to see single-electron pulses

- from literature:  
 $E \geq 5$  keV to see single-electron signals with APDs



S. Kasahara, et al.,  
IEEE Trans. Nucl. Sci. 57 (2010) 1549

A. Apponi et al 2020 JINST **15** P11015



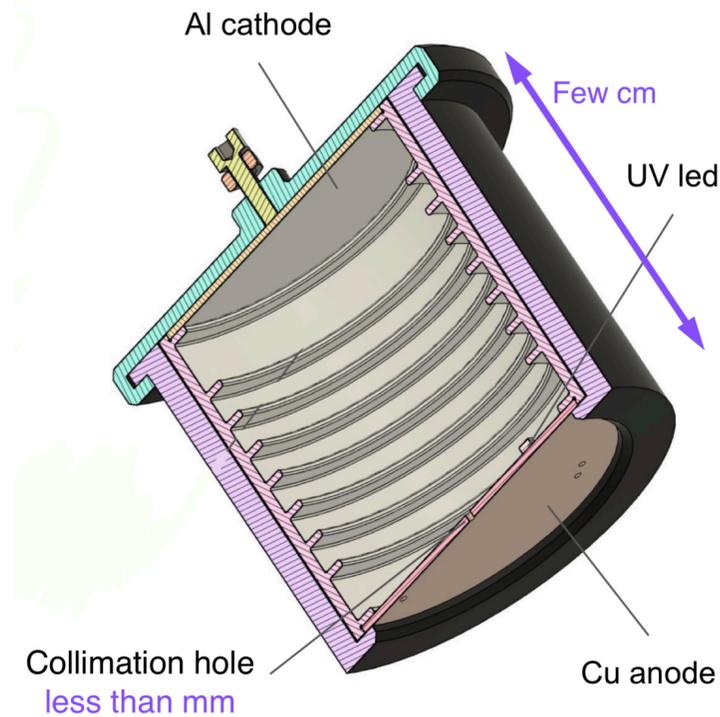


# Electron Gun @ Milano Bicocca

▶ UV led + metallic electrodes + electric field

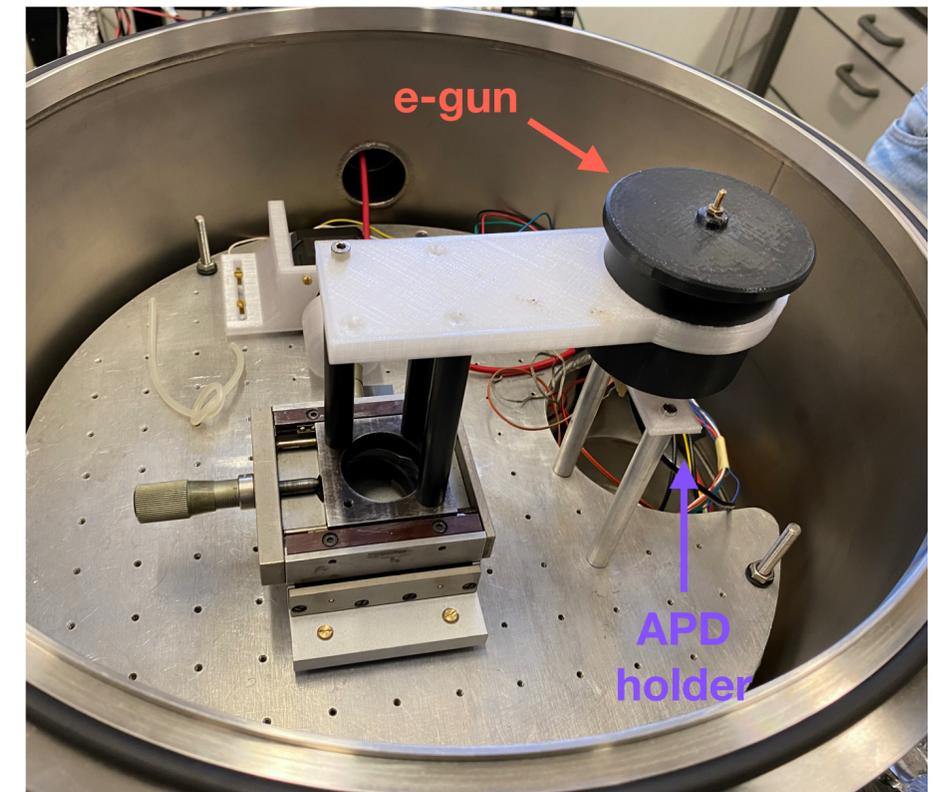
▶ **Key features:**

- Electron energy:  $0 < E < 30$  keV
- Energy uncertainty  $< 2$  eV
- Beam spot  $< 1$  mm
- Current as low as a few fA
- Compact & easy to move



▶ First measurements on APDs in Jan 2023 →

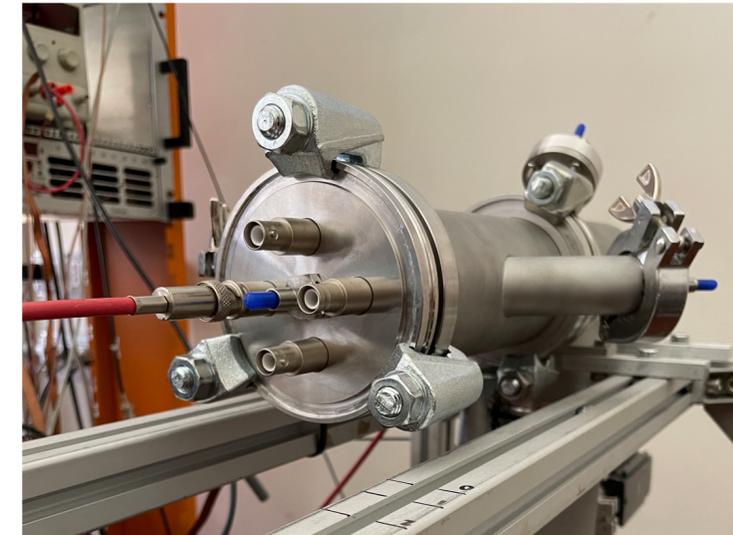
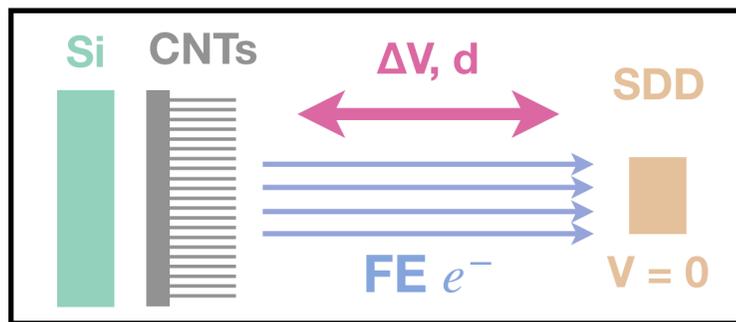
**ongoing analysis**  
(my thesis work)



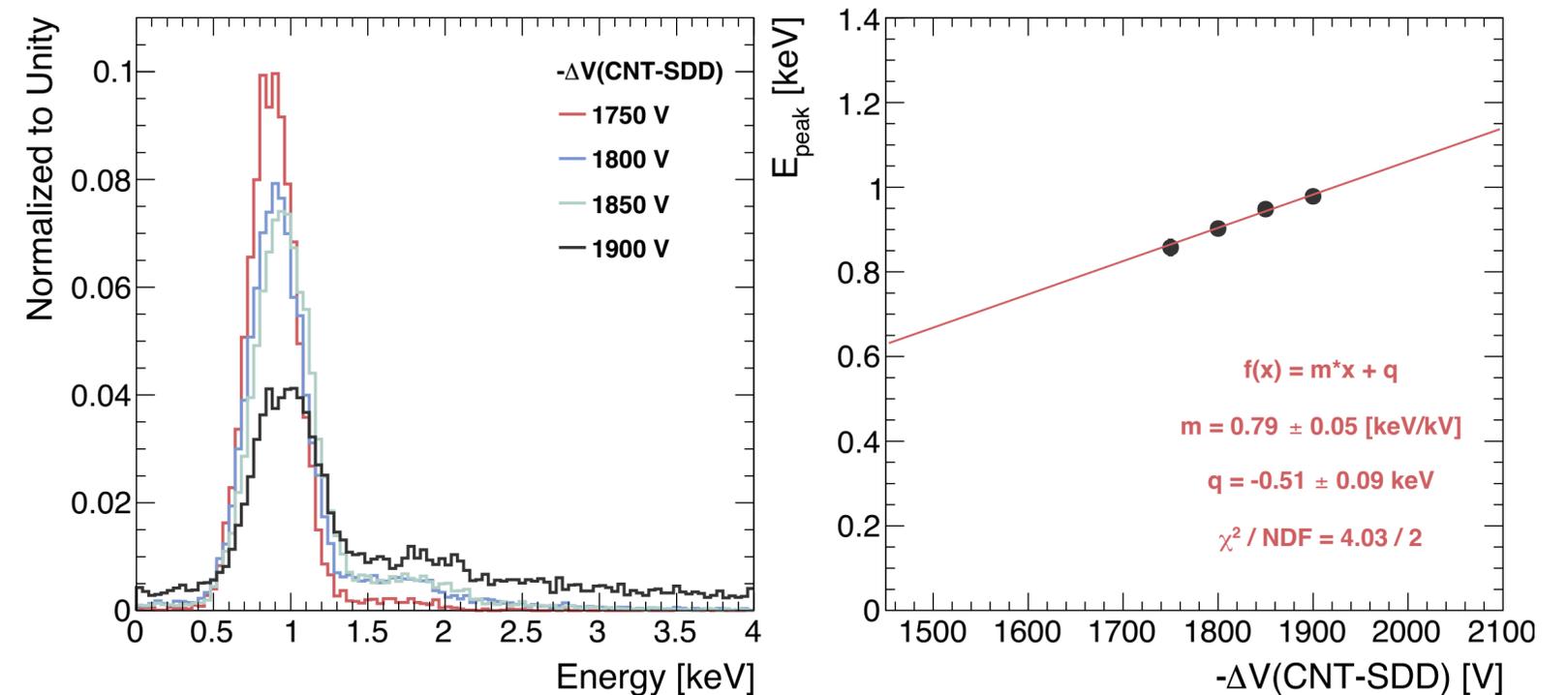


# First Dark-PMT Prototype: Hyperion II

- ▶ Prototype-0 already taking data in Rome Sapienza
- ▶ Observed **field electron emission** from CNTs



- Measurements with SDD
- For high  $\Delta V$  / small  $d(\text{CNT-SDD})$
- $\sim 2$  keV electrons emitted by CNTs detected

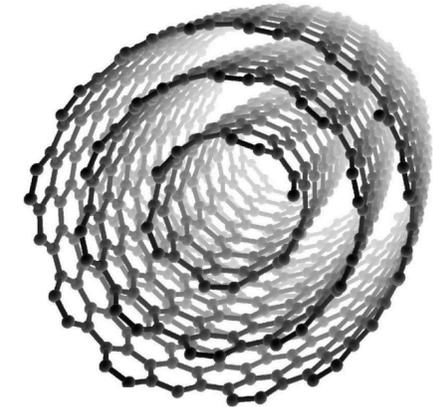




# Conclusions

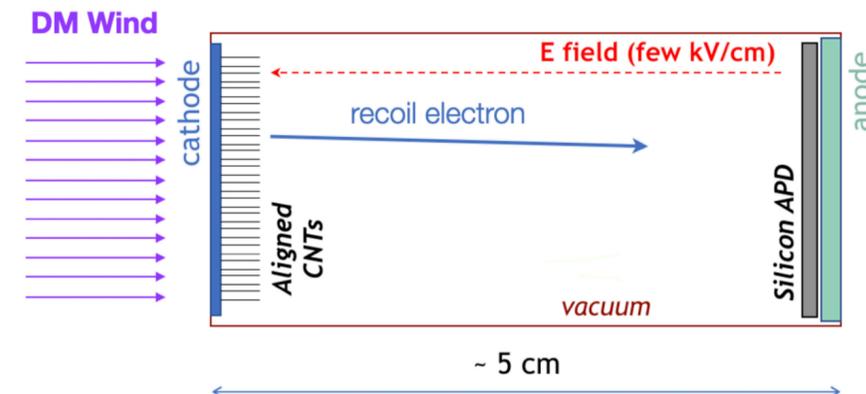
▶ **Carbon nanotubes:** exciting material for dark matter detectors

- 2D material: electrons ejected directly into vacuum
- density anisotropy



▶ **Dark-PMT:** a new dark matter detector concept

- Portable, no thermal noise, directional sensitivity
- Competitive with just 1 g target



▶ **ANDROMeDa:** a young and ambitious program in Rome

- Aiming to build first working Dark-PMT prototype by 2025



**ANDROMeDa**  
Aligned Nanotube Detector for Research On MeV Darkmatter



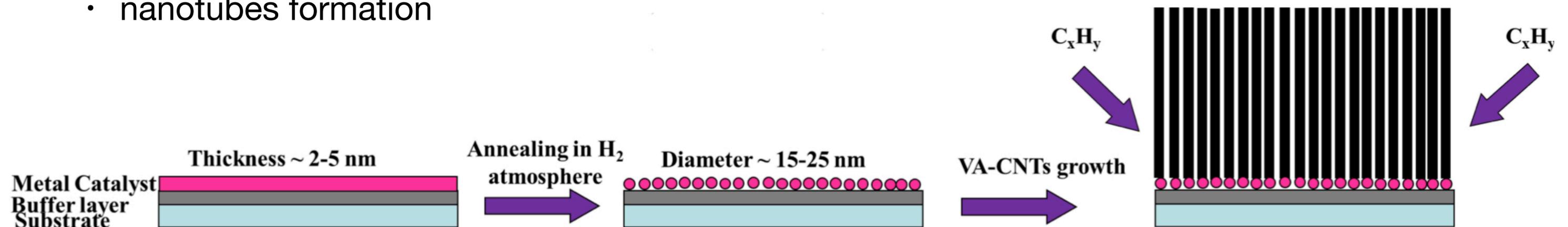
# Backup Slides



# CVD: How Does It Works?

## ▶ Main Steps:

1. metallic nanolayer (e.g. iron) deposited on the substrate
2. annealing at high temperature
  - nanolayer forms nanoparticles = catalyst seeds during synthesis
3. Carbon precursor gas (e.g. acetylene) oriented on nanoparticles at high temperature
  - nanotubes formation

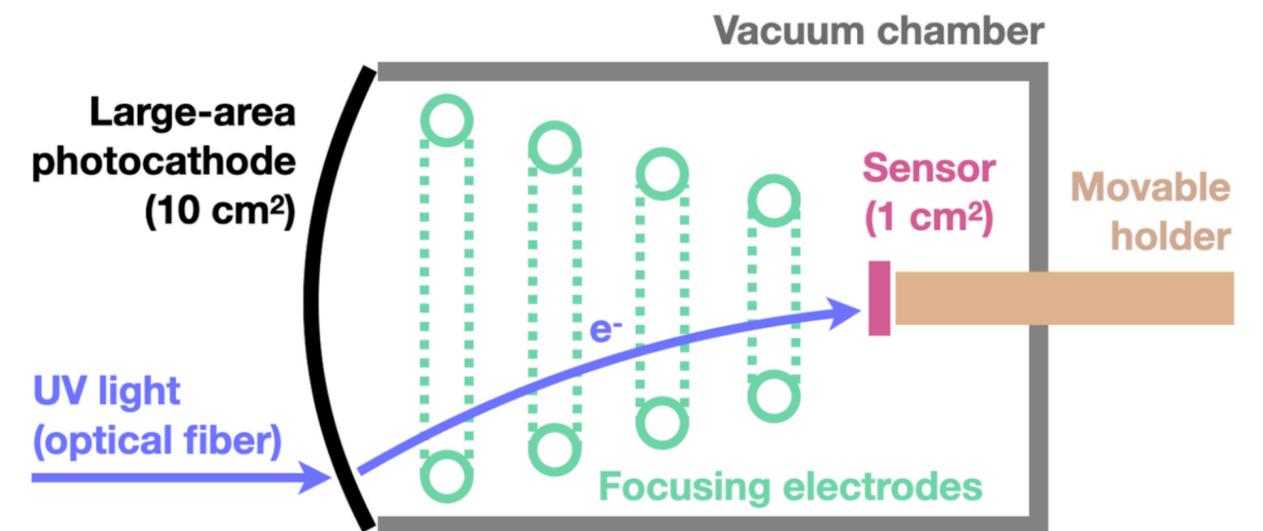




# From 10 cm<sup>2</sup> Cathode to 1 cm<sup>2</sup> Sensor

- To reach 1 g target with 100 dark-PMTs → 10 cm<sup>2</sup> cathodes
- Large Area Silicon Detectors → 1 cm<sup>2</sup> sensors
- ▶ 10:1 **electron focusing system** needed
- ▶ Key parameter: **focusing efficiency**

- computed as 
$$\frac{\# e^- \text{ detected}}{\# e^- \text{ from cathode}}$$
- aim for efficiency > **90%**
- can be optimised using UV light + standard photocathode

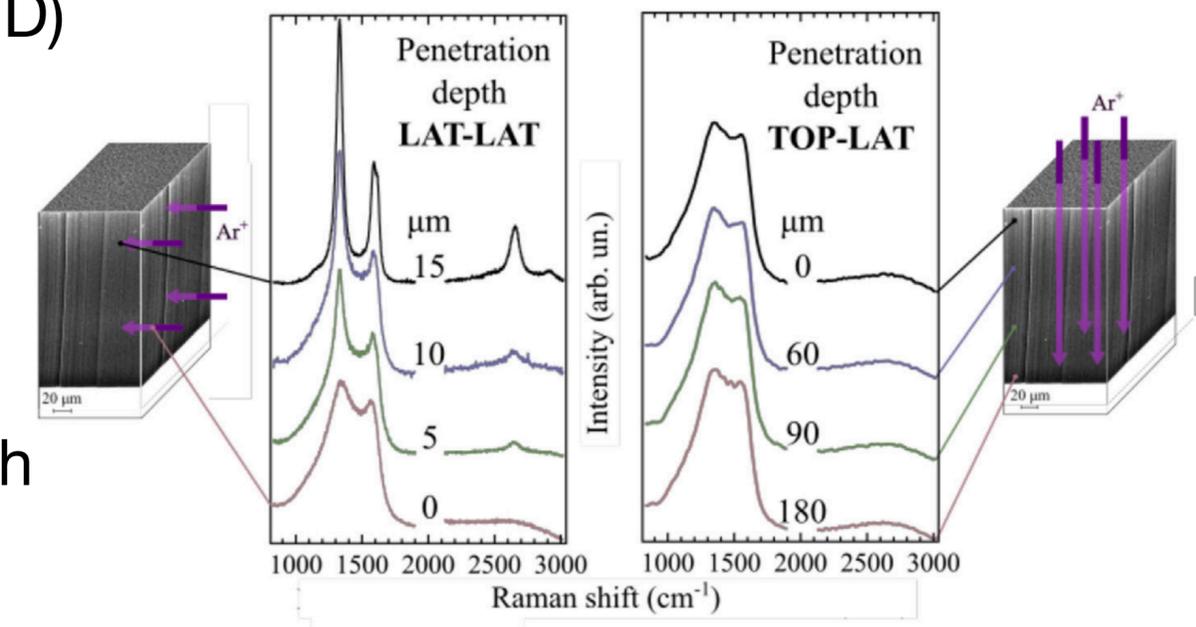
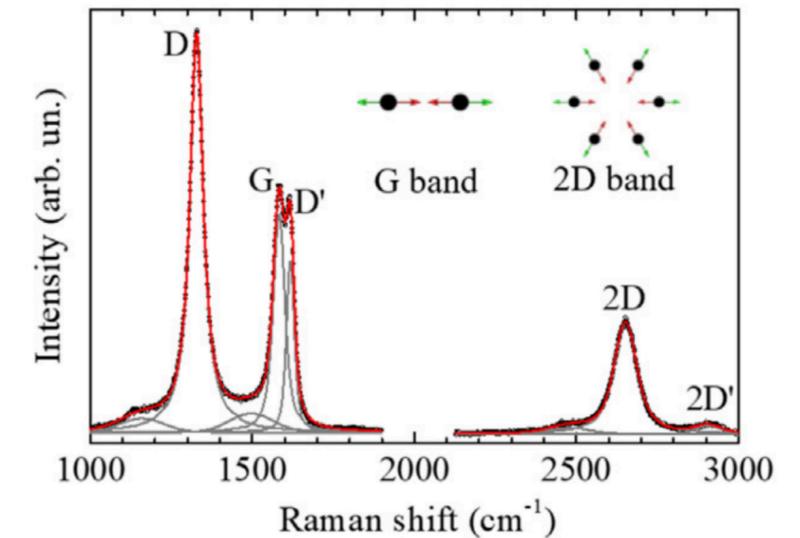


**ONGOING:**  
Electron trajectories simulations  
with SIMION software



# Going Into Detail on Raman Spectroscopy

- ▶ **Raman Spectroscopy** used to determine vibrational modes
  - can assess quality of chemical bondings
  - can be focused at various depth and coordinates of the sample
- ▶ on pristine sample:
  - excitation of vibrational modes of the carbon lattice (G and 2D)
  - prominent 'defect' (D) peak
- ▶ after Ar ions bombardment ( 5 keV at  $1.5 \times 10^{17}$  ions/cm<sup>2</sup> )
  - LATERAL: Raman spectrum unchanged at 15  $\mu\text{m}$  lateral depth
  - TOP: amorphization from at any height



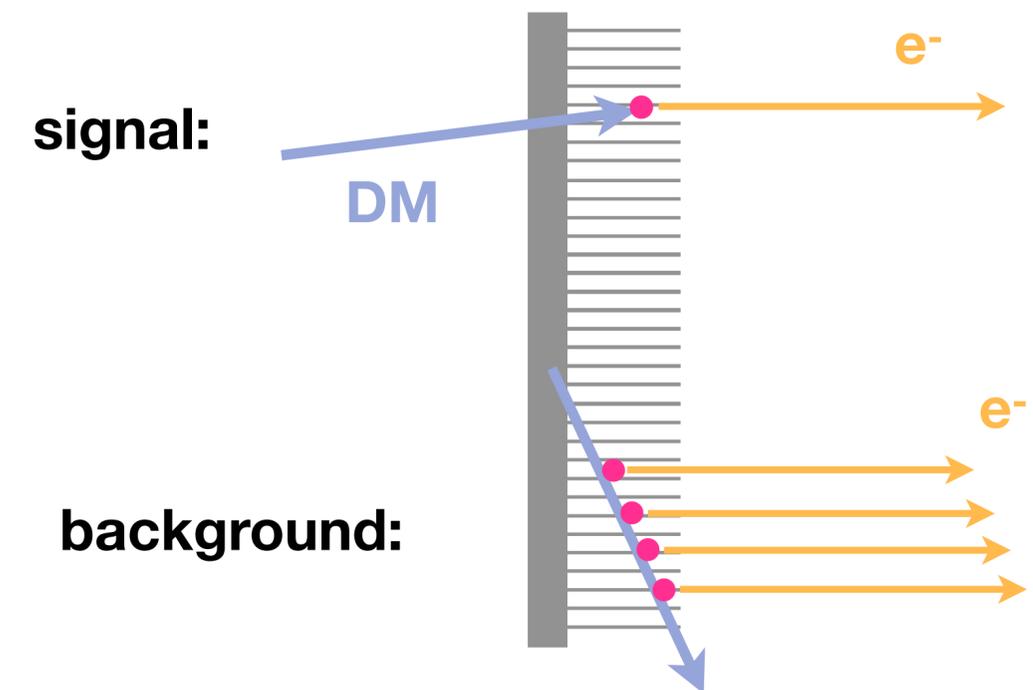
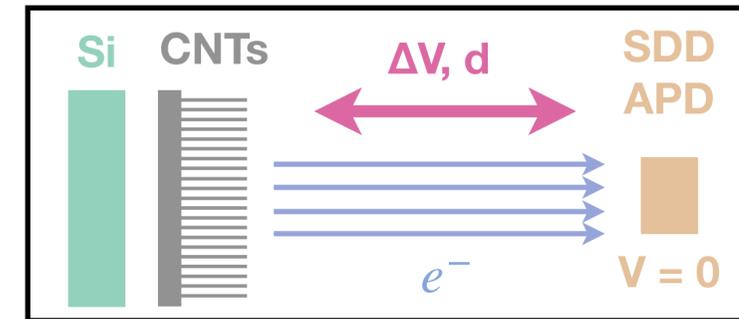


# Electron Detection: Main Challenges

▶ Electron energy =  $\Delta V$ (anode-cathode) to keep  $< 10\text{kV}$

▶ **Key parameters :**

1. **compactness** to have a portable dark-PMT detector;
2. high ( $>90\%$ ) **efficiency** on single  $e^-$  detection in keV energy range;
3. percent-level **discrimination** between  $1e^-$  and  $2e^-$  events
4. **suppression** at permil level of fake single  $e^-$  signals due to noise





# Unwanted Features of Carbon Nanotubes

▶ Two **problems** with as-grown nanotubes forests:

1. non-aligned **top crust** layer

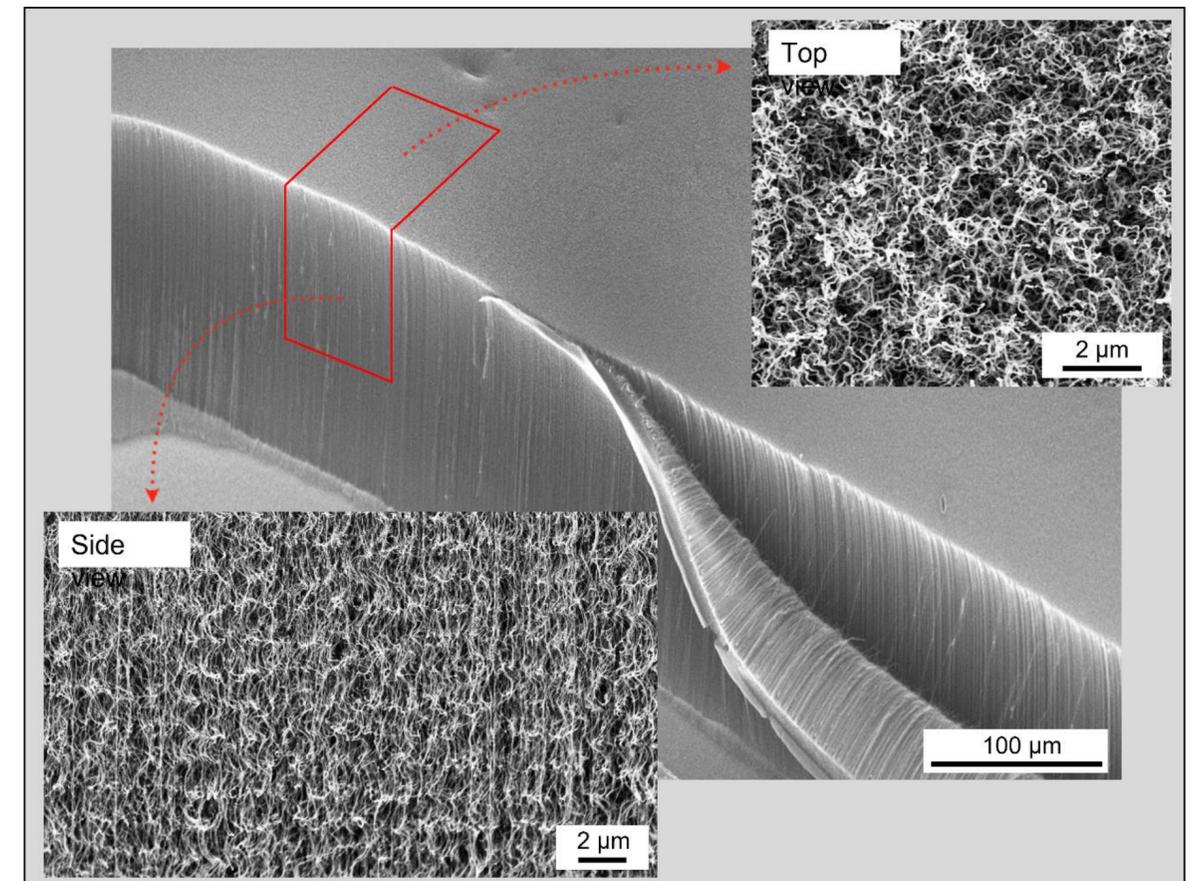
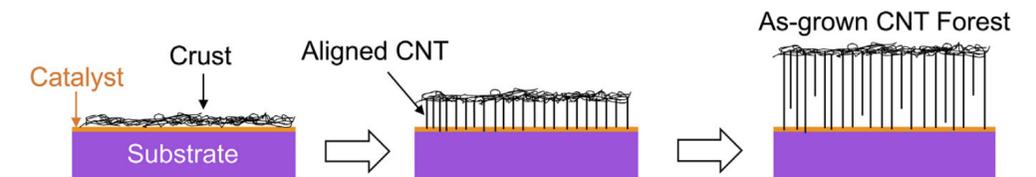
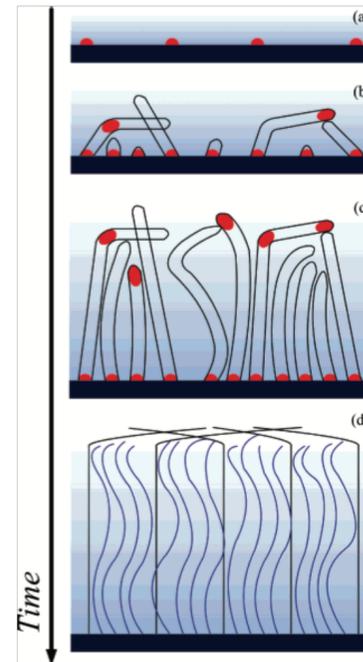
- due to initial growth instabilities

2. side **waviness** at the nanoscale

- due to different growth rates

▶ Both hamper electron transmission

- **minimisation** needed for ideal DM target

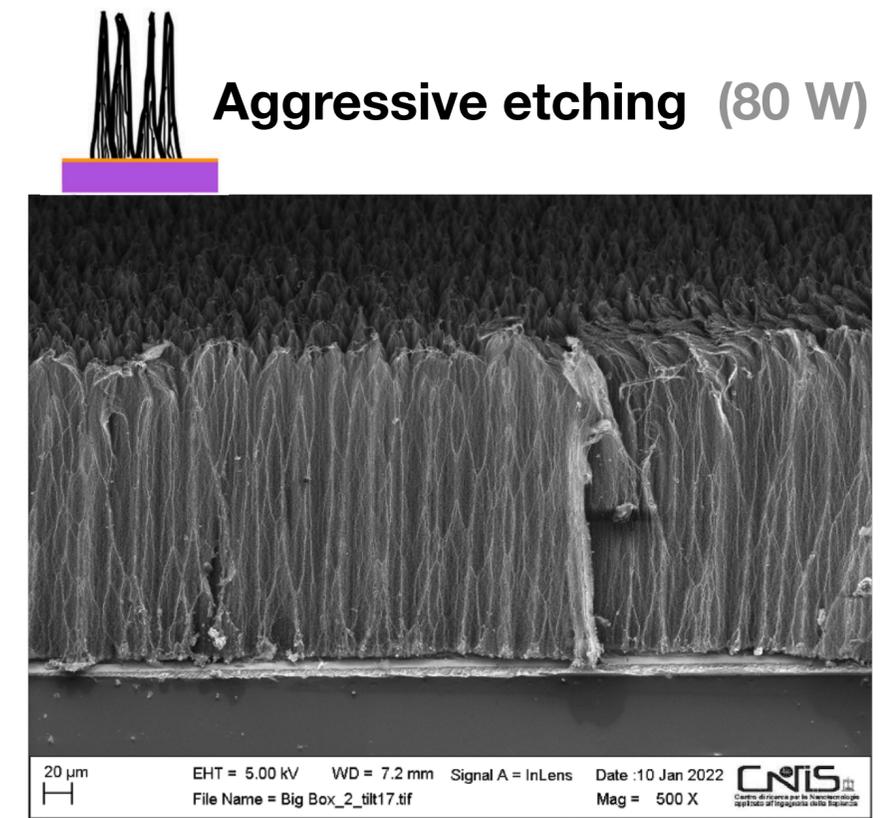
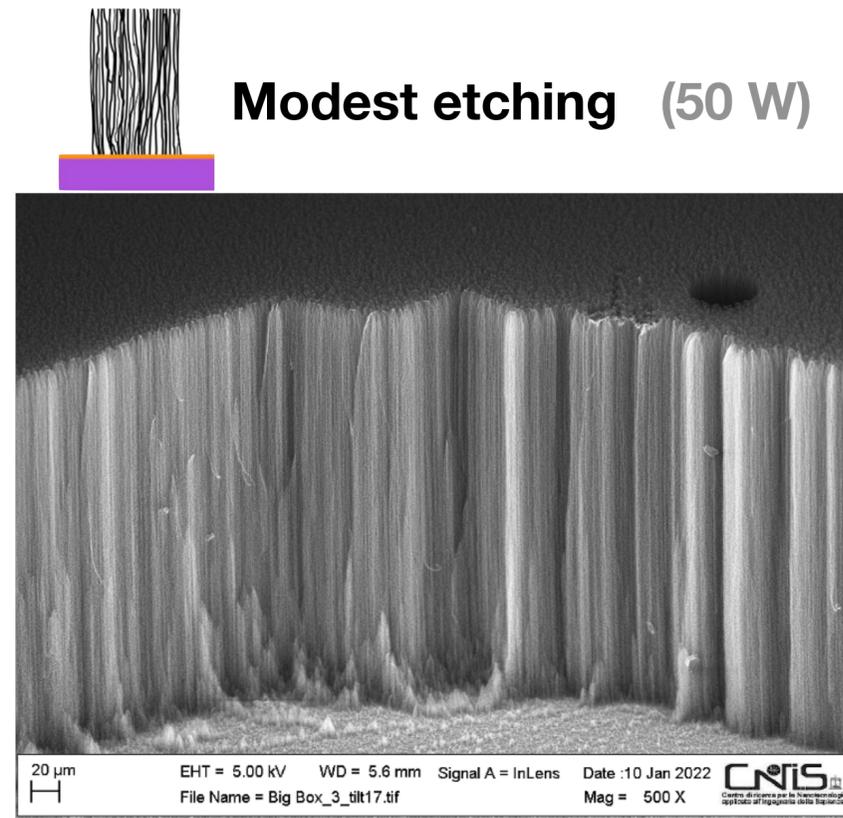
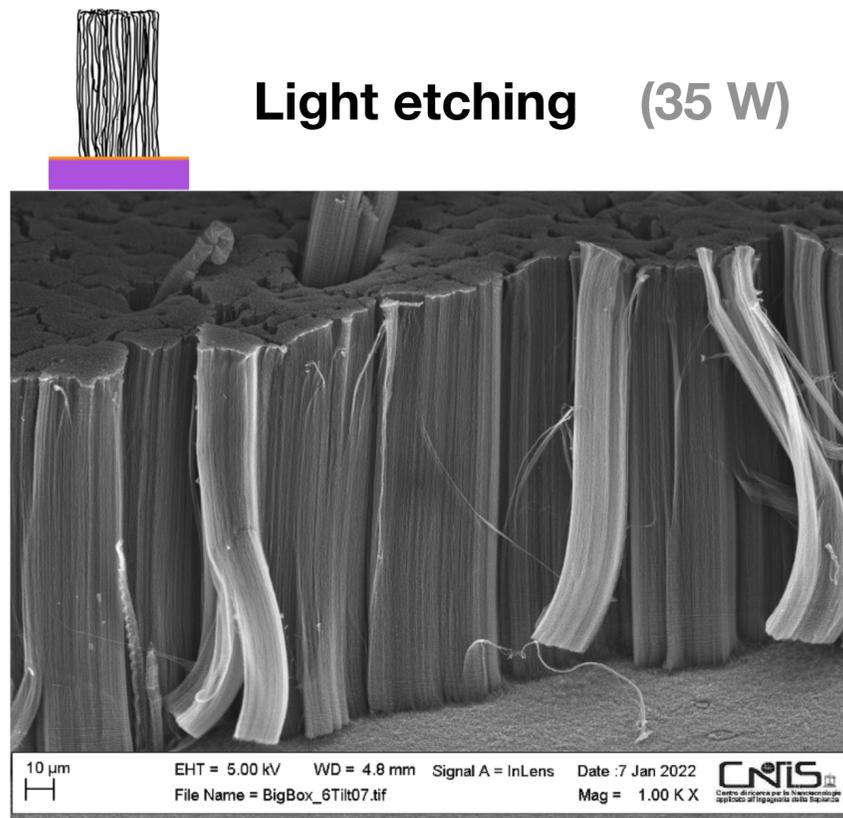




# Plasma Etching to Remove Crust

► **Optimisation** of  $Ar_2/O_2$ -plasma etching parameters (e.g. time, plasma power, frequency, pressure)

- measuring morphology (SEM), roughness (AFM) and electron emission

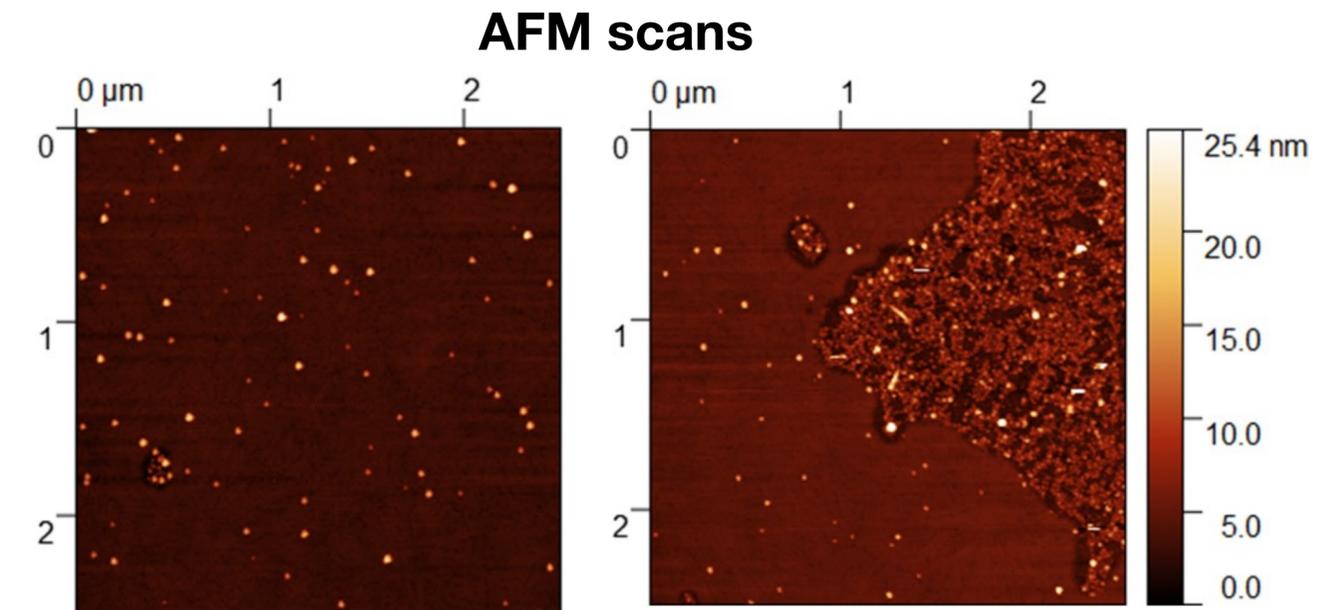




# Aiming for Ultimate Parallelism at the Nanoscale

## ▶ Parallelism strongly influenced by iron **catalyst seeds**

1. **non-uniformity** in seeds size
  - leads to different growth rates
2. **density** of seeds
  - farer seeds = weaker interaction between tubes



## ▶ Evaporation chamber being built in Rome

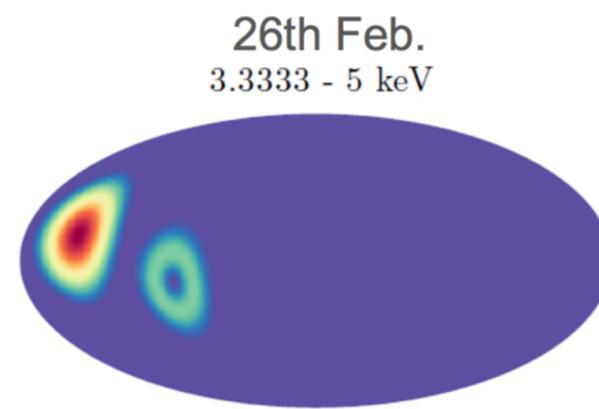
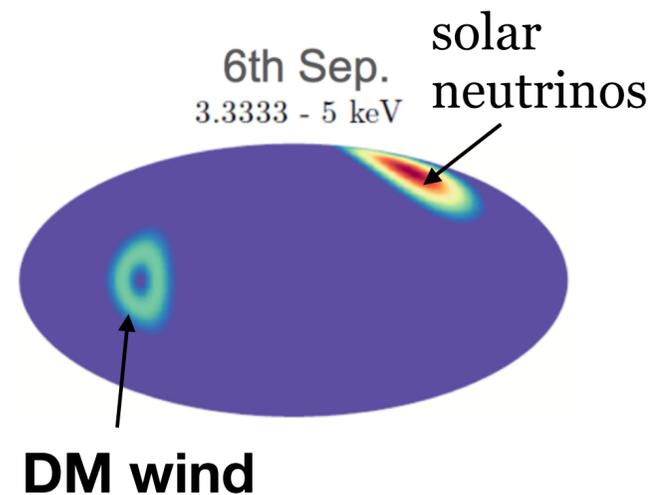
- aim for seed density  $> 10^{12} \text{ cm}^{-2}$
- **AFM** to check seeds size, density and distribution
- **iterative optimisation** of nucleation parameters

	Now	Goal
seeds density [cm <sup>-2</sup> ]	$10^{10}$ - $10^{11}$	$> 10^{12}$
seeds size [nm]	15-30	5 ( $\pm 20\%$ )

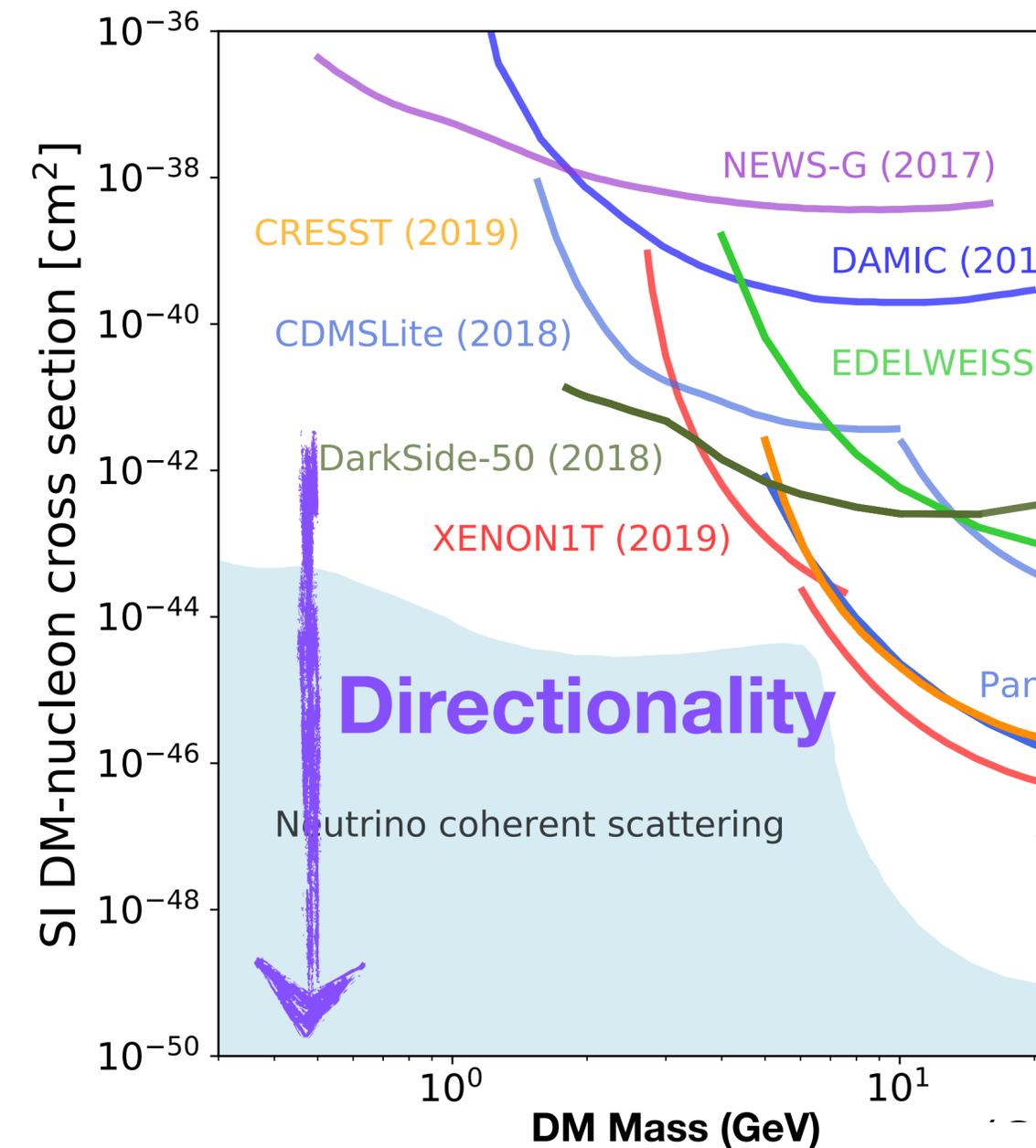


# Avoiding Neutrino Background with Directionality

- ▶ **Directionality:** link a signal with region of the sky
  - DM 'wind' expected to come from Cygnus constellation
- ▶ But also to be **insensitive to neutrino floor**
  - Low mass neutrino floor mostly from solar neutrinos
  - Cygnus **never overlaps** with Sun



O'Hare et al, Phys. Rev. D 92, 063518 (2015)





# World-Leading Sensitivity Below 30 MeV?

## ▶ **Competitive** with other light DM searches

- with just 1 year exposure
- using 1 kg target

## ▶ In principle **sensitive to few MeV DM**

- **extend search** below 30 MeV
- using just 1 g target

G. Cavoto, et al., PLB 776 (2018) 338

