

Light dark matter, dark sector EFT and long-lived states

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Based on 1807.10314 and 1912.xxxx

Outline

Introduction: sub-GeV dark matter and dark sectors

Long-lived states in and EFT description

Production of light dark sector at the intensity frontier and EFT

Signatures of light dark sectors and limits

Introduction

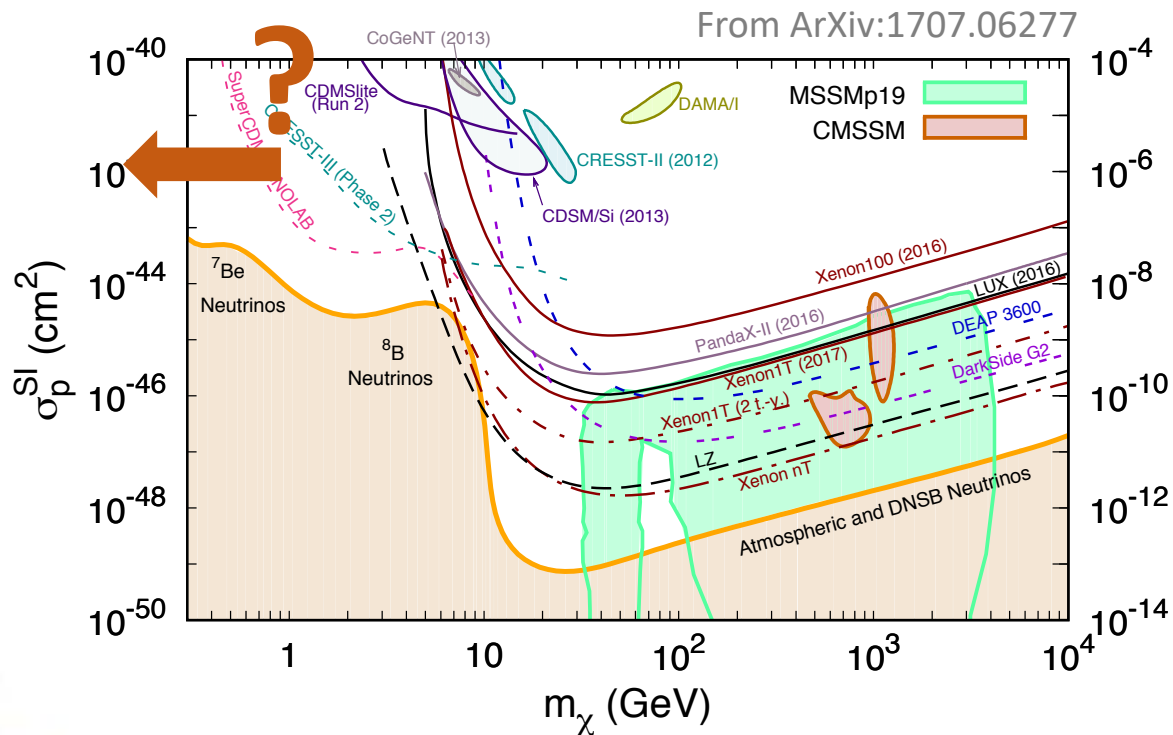
sub-GeV DM and dark sector: an example

Light dark sectors and dark matter

- WIMP Dark matter is a mature area of research
- what about lighter (sub-GeV) DM ?

- Strong experimental effort in the intensity frontier

→ $\gtrsim 10$ relevant experiment in next 2 years



Information References (339) Citations (182) Files Plots

US Cosmic Visions: New Ideas in Dark Matter 2017: Community Report

Marco Battaglieri (SAC co-chair) (INFN, Genoa), Alberto Belloni (Coordinator) (Maryland U.), Aaron Chou (WG2 Convener) (Fermilab), Priscilla Cushman (Coordinator) (Minnesota U.), Bertrand Echenard (WG3 Convener) (Caltech), Rouven Essig (WG1 Convener) (SUNY, Stony Brook), Juan Estrada (WG1 Convener) (Fermilab), Jonathan L. Feng (WG4 Convener) (UC, Irvine), Brenna Flaugher (Coordinator), Patrick J. Fox (WG4 Convener) (Fermilab) *et al.* [Show all 251 authors](#)

Jul 14, 2017 - 113 pages

Information References (355) Citations (19) Files Plots

Physics Beyond Colliders at CERN: Beyond the Standard Model Working Group Report

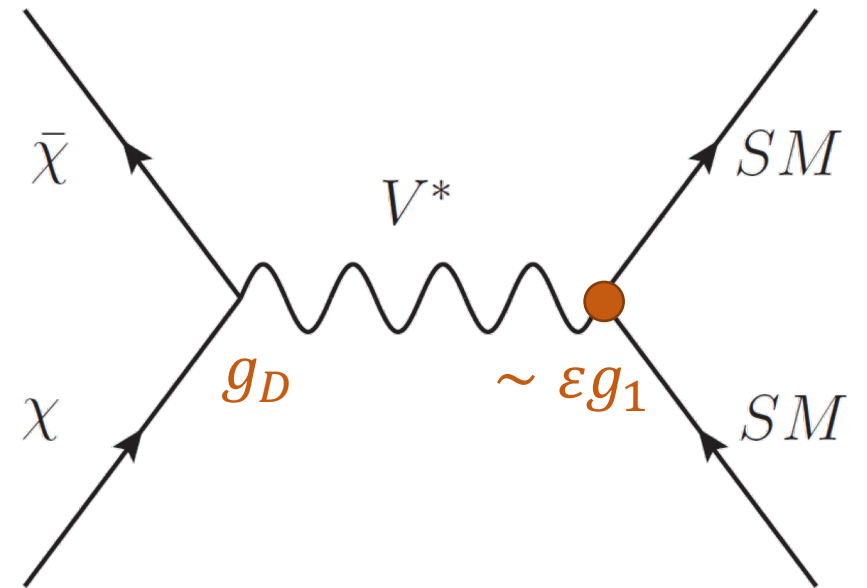
J. Beacham (Ohio State U., Columbus (main)), C. Burrage (U. Nottingham), D. Curtin (Toronto U.), A. De Roeck (CERN), J. Evans (Cincinnati U.), J.L. Feng (UC, Irvine), C. Gatto (INFN, Naples & NIU, DeKalb), S. Gninenko (Moscow, INR), A. Hartin (U. Coll. London), I. Irastorza (U. Zaragoza, LFNAE) *et al.* [Show all 33 authors](#)

Jan 20, 2019 - 150 pages

- Not all DM-motivated, yet most can be used to constrained sub-GeV thermal DM and dark sectors

Sub-GeV DM, a step-by-step example

- Sub-GeV DM typically require an additional “dark sector”
- Let’s try to build a simple, self-consistent dark sector model with sub-GeV dark matter → presence of long-lived state
- Suppose a vector mediator (dark photon), and (mostly in this talk) fermion dark matter
- Try to keep model building SM/WIMP-like and maintain top-down consistency



Astrophysics of sub-GeV DM

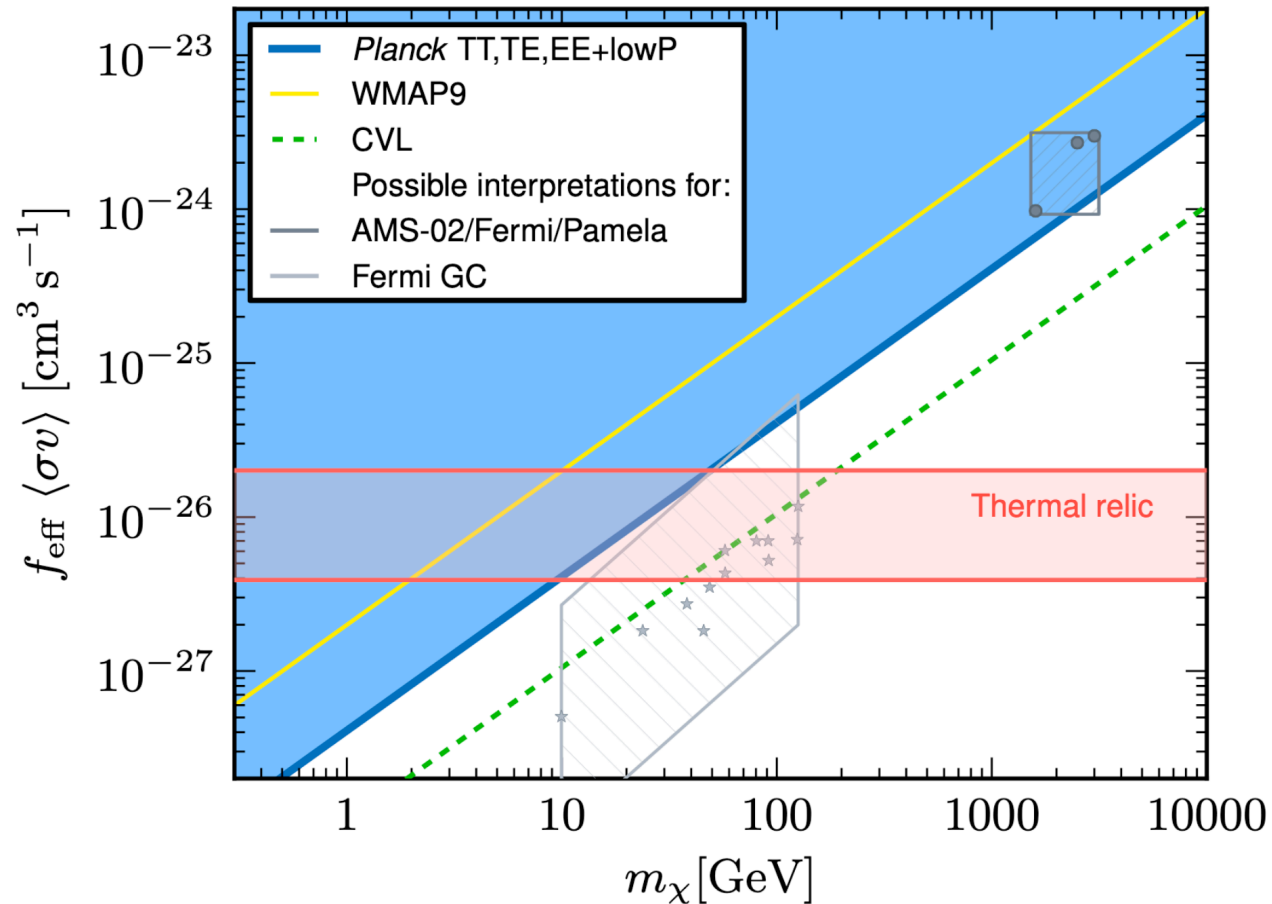
- **Relic density** -> sub-GeV particles,
 $\varepsilon \sim 10^{-3}$ suppression

$$\Omega h^2 \sim 0.1 \times \left(\frac{10^{-3}}{\varepsilon}\right)^2 \left(\frac{0.1}{\alpha_D}\right) \times \left(\frac{25 \text{ MeV}}{M_\chi}\right)^2 \left(\frac{M_V}{75 \text{ MeV}}\right)^4$$

- **CMB limits**

- No active annihilation process by the time of CMB
- Exclude s-wave annihilation, requires additional mechanism (p-wave, co-annihilation ...)

From Klasen, Pohl and Sigl 1507.03800



Kinetic mixing and dark Higgs mechanism

- Coupling to SM obtained through “kinetic mixing” term

$$\mathcal{L}_{A'} = -\frac{1}{4} F'^{\mu\nu} F'_{\mu\nu} - \frac{1}{2} \frac{\epsilon}{\cos \theta_w} B_{\mu\nu} F'^{\mu\nu} + (D^\mu S)^* (D_\mu S) + \mu_S^2 |S|^2 - \frac{\lambda_S}{2} |S|^4$$

Kinetic mixing term

Dark Higgs potential, with SSB

- Anomaly cancellation -> Introduce a Dirac fermion dark matter
 $\chi = (\chi_L, \bar{\chi}_R)$

Fermion dark matter example

$$\mathcal{L}_{pDF}^{\text{DM}} = \bar{\chi} (i\not{D} - m_\chi) \chi + y_{SL} S \bar{\chi}^c P_L \chi + y_{SR} S \bar{\chi}^c P_R \chi + \text{h.c.}$$

- Yukawa couplings to the dark Higgs S

- Avoid Dirac DM (CMB exclusion)

- After $U(1)_D$ symmetry breaking, the dark matter acquires a Majorana mass

$$M_\chi = \begin{pmatrix} \sqrt{2}v_S y_{SL} & m_\chi \\ m_\chi & \sqrt{2}v_S y_{SR} \end{pmatrix}$$

$$M_V = g_{\alpha_D} q_S v_S \quad \uparrow \quad V$$

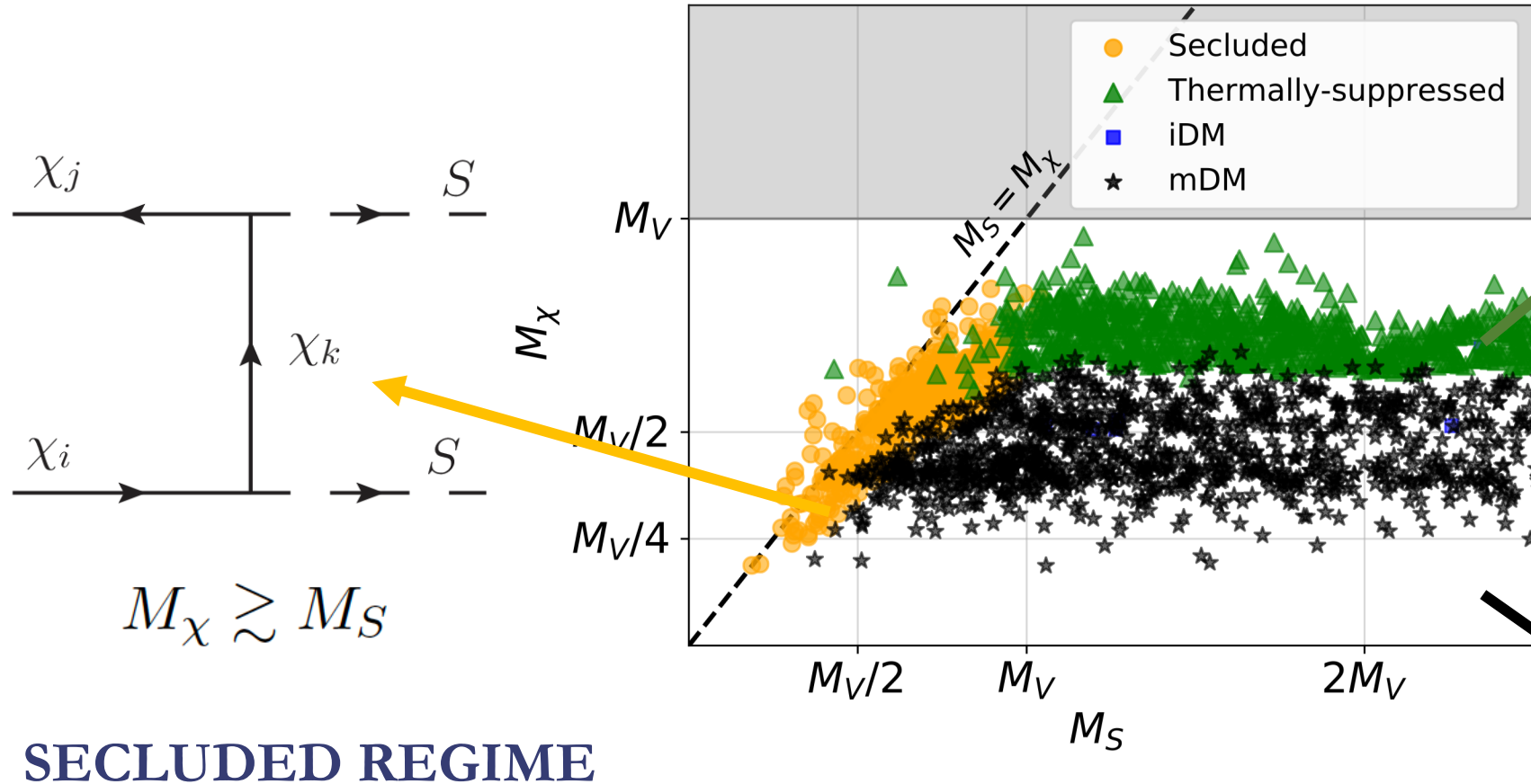
$$M_S = \sqrt{2\lambda_S} v_S \quad \uparrow \quad S$$

$$M_{\chi_2} - M_{\chi_1} = \sqrt{2}v_S (y_{SR} + y_{SL}) \quad \updownarrow \quad \begin{matrix} \chi_2 \\ \chi_1 \end{matrix}$$

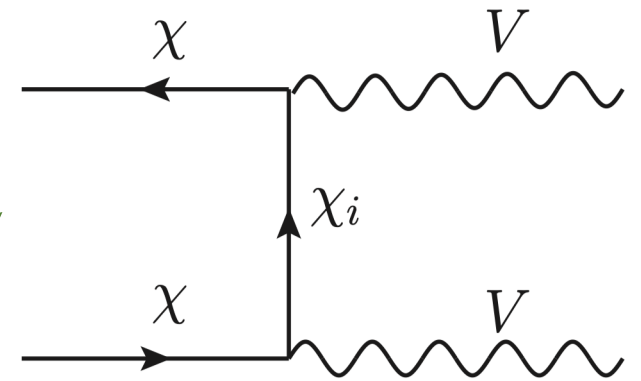
- After diagonalization → two Majorana fermions

Typical regimes with correct relic density

LD, Rao, Roszkowski, 1710.00971

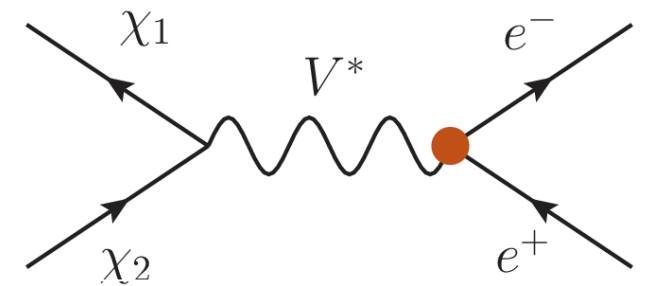


FORBIDDEN REGIME



$$M_\chi \lesssim M_V \lesssim M_S$$

iDM/mDM REGIME



Long-lived states and EFT

Dark sectors, their long-lived states and how to describe them

Complete models of light thermal DM

- Strong theoretical developments toward building models of **thermal sub-GeV DM**
- **Dark matter is bundled with a dark sector,** with potentially many particles in it
 - Required to obtain the proper relic density (while avoiding CMB limits)
 - Implied from top-down approach (e.g anomaly cancellations, Higgs mechanism for dark photon mass, etc...)

iDM
hep-ph/0101138, ...

Secluded DM
0711.4866, ...

Semi-annihilating DM
1003.5912, ...

Boosted DM
1405.7370, 1503.02669...

Selfish DM
1504.00361,...

Forbidden DM
Griest-Seckel, 1505.07107, ...

Co-decaying DM
1607.03110, ...

Impeded DM
1609.02147,...

...and many more recent

A key consequence: long-lived particles

- Decays involving SM particles are often the only option for unstable dark sector states

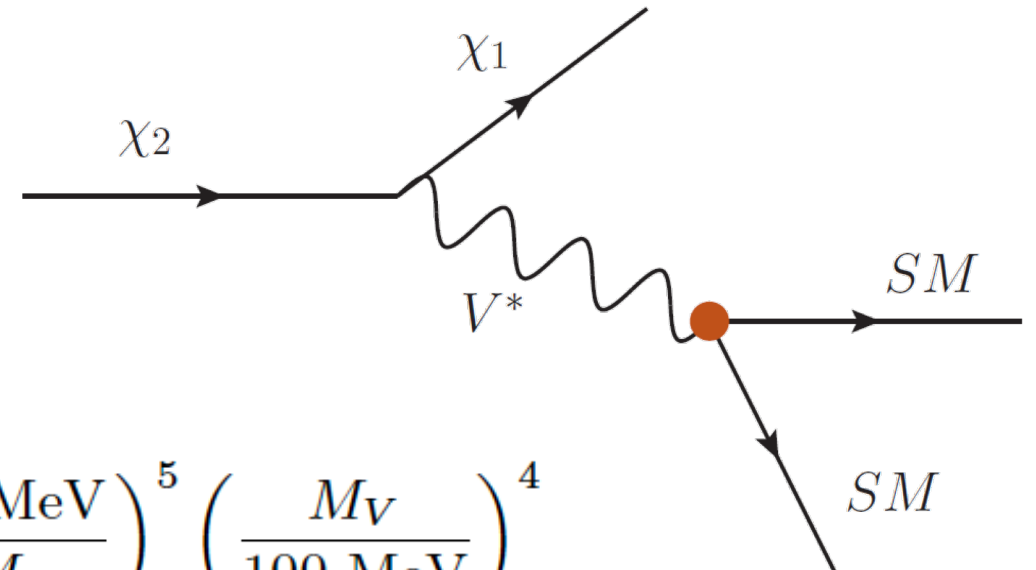
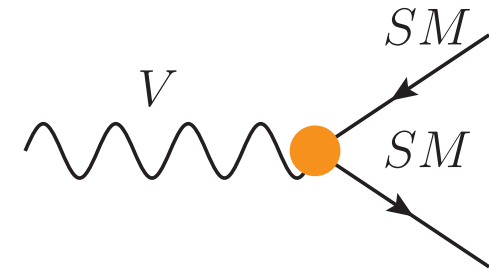
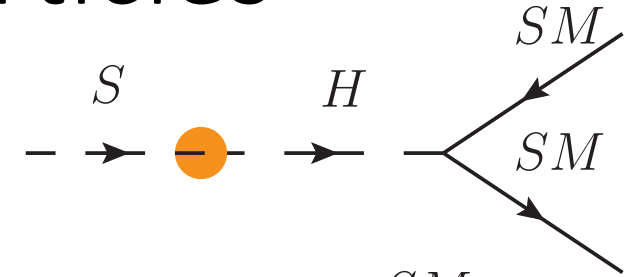
- Through the portal -> e.g. dark Higgs boson, dark photon

$$c\tau_{V \rightarrow e^+e^-} \sim 1 \text{ cm} \times \left(\frac{10^{-5}}{\epsilon} \right)^2 \left(\frac{M_V}{100 \text{ MeV}} \right)$$

- Mixed visible/dark decays are also often relevant.

- Here: dark sector decays which proceed through off-shell mediator \rightarrow e.g. semi-visible 3-body decays (iDM, certain sterile neutrino models, etc....)

$$c\tau_{\chi_2} \propto 100 \text{ m} \times \left(\frac{0.1}{\alpha_D} \right) \left(\frac{10^{-3}}{\epsilon} \right)^2 \left(\frac{0.2 M_\chi}{\Delta_\chi} \right)^5 \left(\frac{25 \text{ MeV}}{M_\chi} \right)^5 \left(\frac{M_V}{100 \text{ MeV}} \right)^4$$



The off-shell effective approach

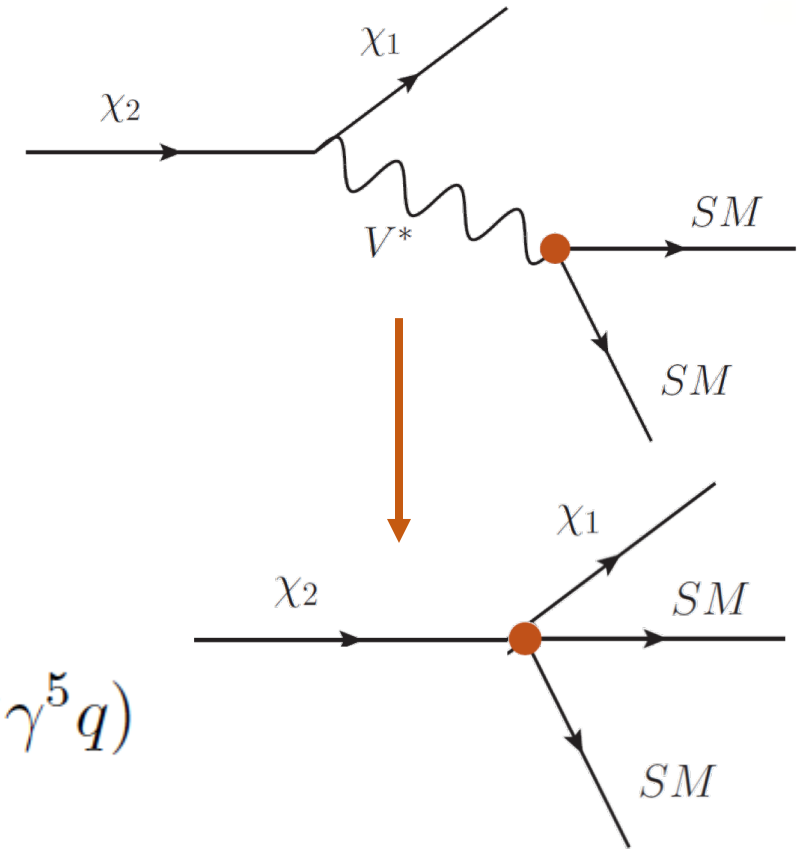
- Relevant for decay mediated by an “off-shell” mediator
 - Effective theory description (cf Fermi theory) can be seen as a “fermion portal”

Vector operator

$$\sum_q \frac{g_q}{\Lambda^2} (\bar{\chi}_1 \gamma_\mu \chi_2) (\bar{q} \gamma^\mu q)$$

Axial vector operator

$$\sum_q \frac{\tilde{g}_q}{\Lambda^2} (\bar{\chi}_1 \gamma_\mu \gamma^5 \chi_2) (\bar{q} \gamma^\mu \gamma^5 q)$$



- Basic equivalence with a dark photon model with kinetic mixing ε and coupling g_D :

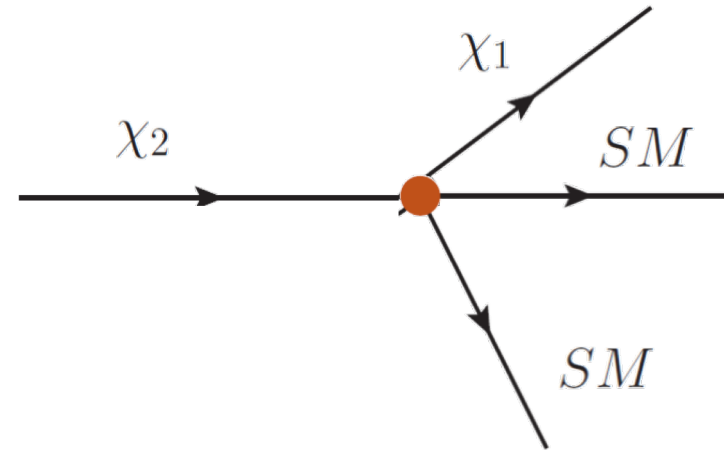
$$\frac{\Lambda}{\sqrt{g}} \sim \frac{M_V}{\sqrt{\varepsilon g_D e}}$$



Could probe scale 2 to 3 orders of magnitudes larger than χ

Signatures and decay

- Typical decay length of order meter
 - Decay into pair of electrons (no background from neutrino scattering)
 - In optimum region, large portion of the heavy dark states decay in the detector



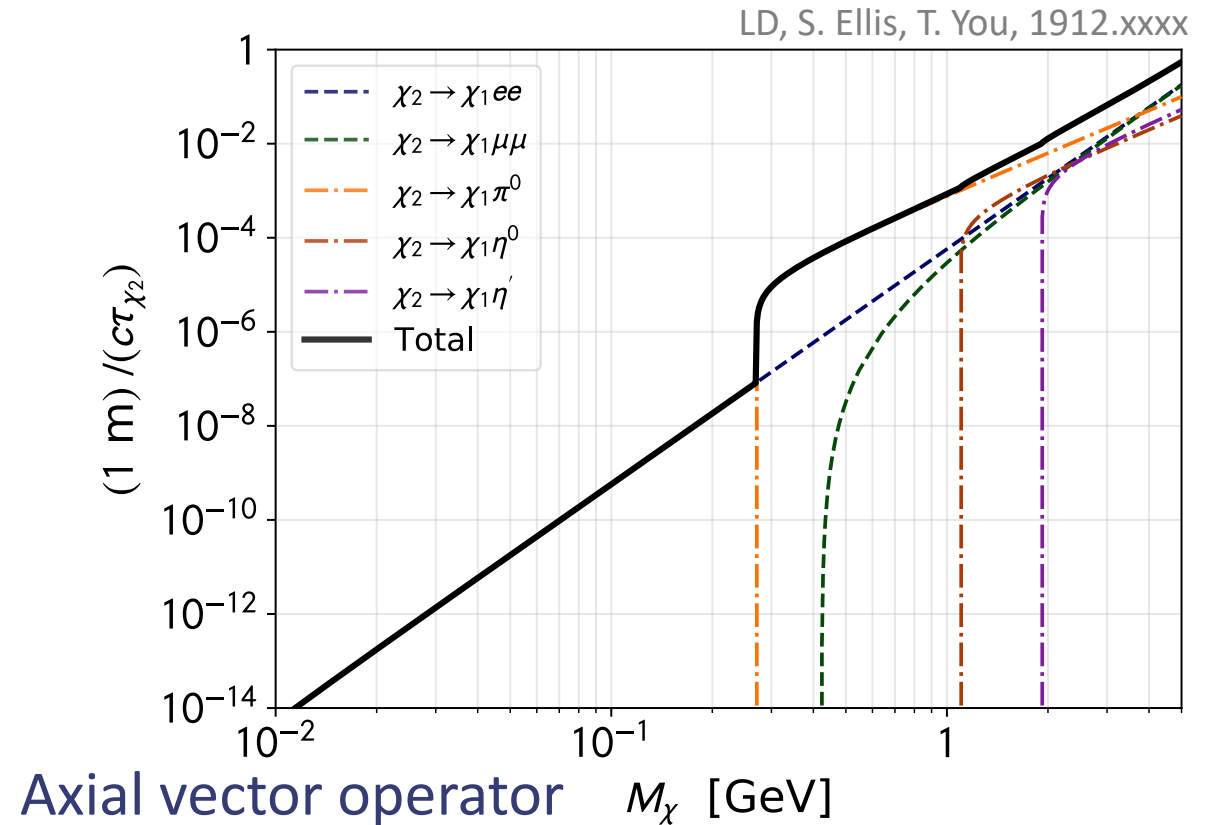
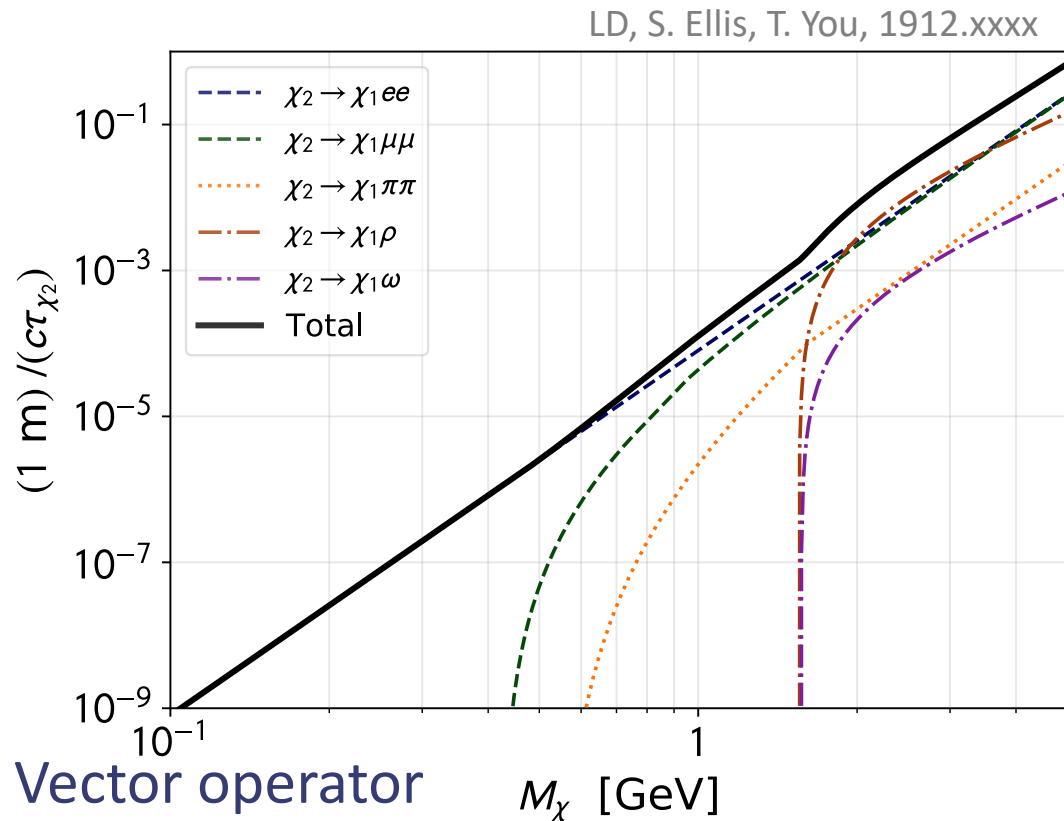
$$c\tau^{\text{PD}} \sim 375 \text{ m} \times \left(\frac{100 \text{ GeV}}{\Lambda}\right)^4 \left(\frac{1 \text{ GeV}}{M_{\chi_1}}\right)^5 \left(\frac{0.25}{\Delta_\chi}\right)^5 \left(\frac{0.01}{g}\right)^2 \quad (\text{Vector operator})$$

- Leptonic channel almost **always dominant/significant BR** (exception AV case)
- Reach also for $M_2 \gg M_1$

$$c\tau^{\text{sat}} \sim 2 \text{ m} \times \left(\frac{\Lambda/\sqrt{g}}{1 \text{ TeV}}\right)^4 \left(\frac{1 \text{ GeV}}{M_{\chi_2}}\right)^5$$

Decay rate of heavy state

- The decay rate depends on the possible decay channel \rightarrow depends on the operator type ($M_2 \gg M_1$ and $\Lambda = 5$ TeV)



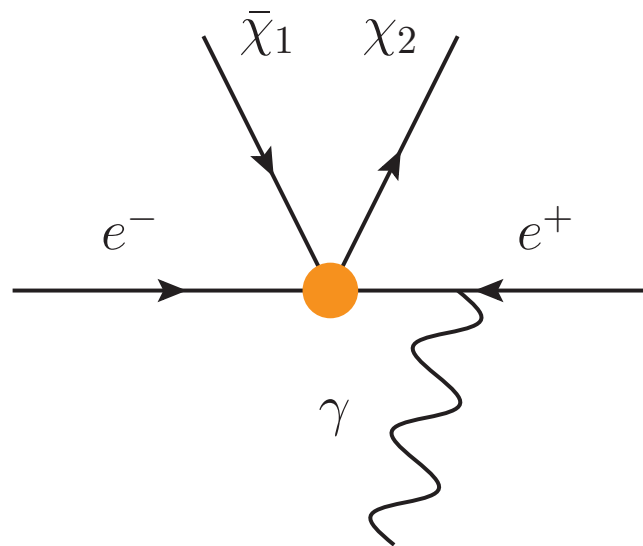
Light dark sector EFT and the intensity frontier

Production at accelerator-based experiments

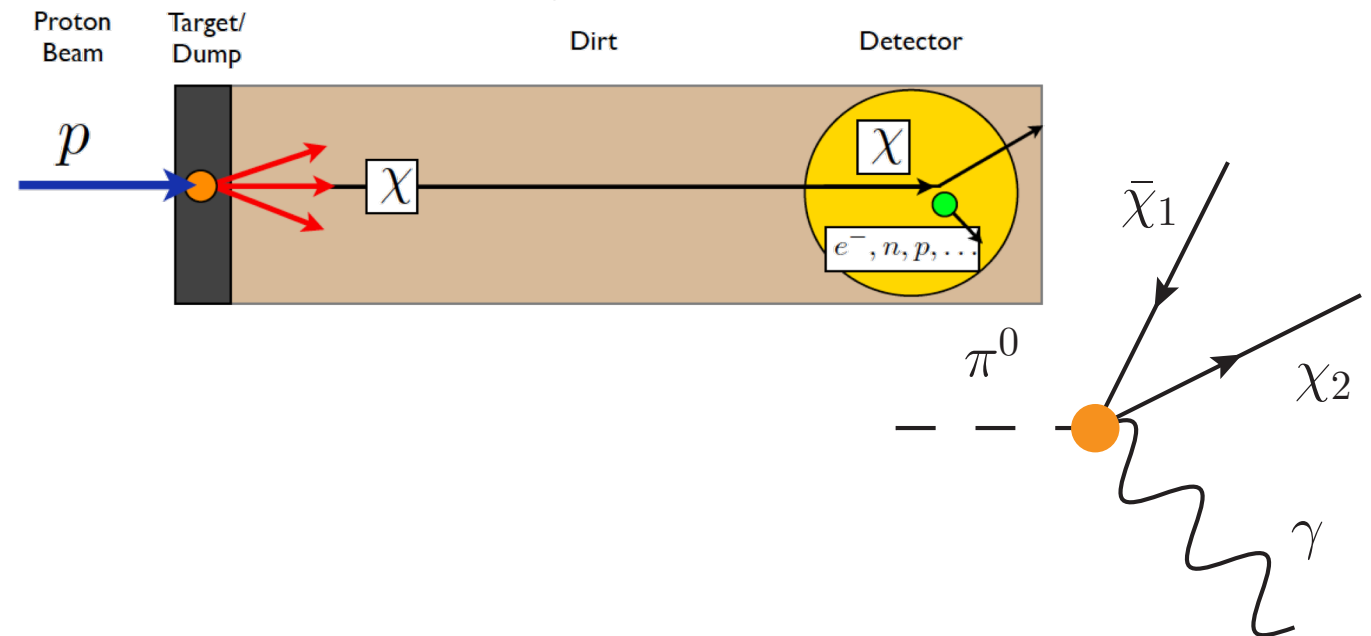
Dark Sector searches - production

- Light dark sector particles may be accessible at the *intensity frontier* (GeV energy / large intensity / good background rejection)
 - Since typically $\Lambda > E_{cm}$, EFT description of off-shell production well-defined

Precision experiments at collider (e.g BaBar, BELLE...)



Beam-dump/fixed-target types of experiments (LSND, CHARM...)



Production in the light dark sector EFT

Production is strongly modified w.r.t the on-shell mediator production

- Off-shell nature of the process -> Strong suppression of low energy production mechanism.

→ For meson decay, BR typically suppressed $\propto \frac{M_m^4}{\Lambda^4}$

- On-shell mediator bremsstrahlung $e^- N \rightarrow e^- N V$ or $p N \rightarrow p N V$ not available

→ Electron beam-dump production suppressed

- When available, direct production more relevant since higher c.o.m energy compared to Λ

Depending on the nature of the operators, different production channels from meson decay

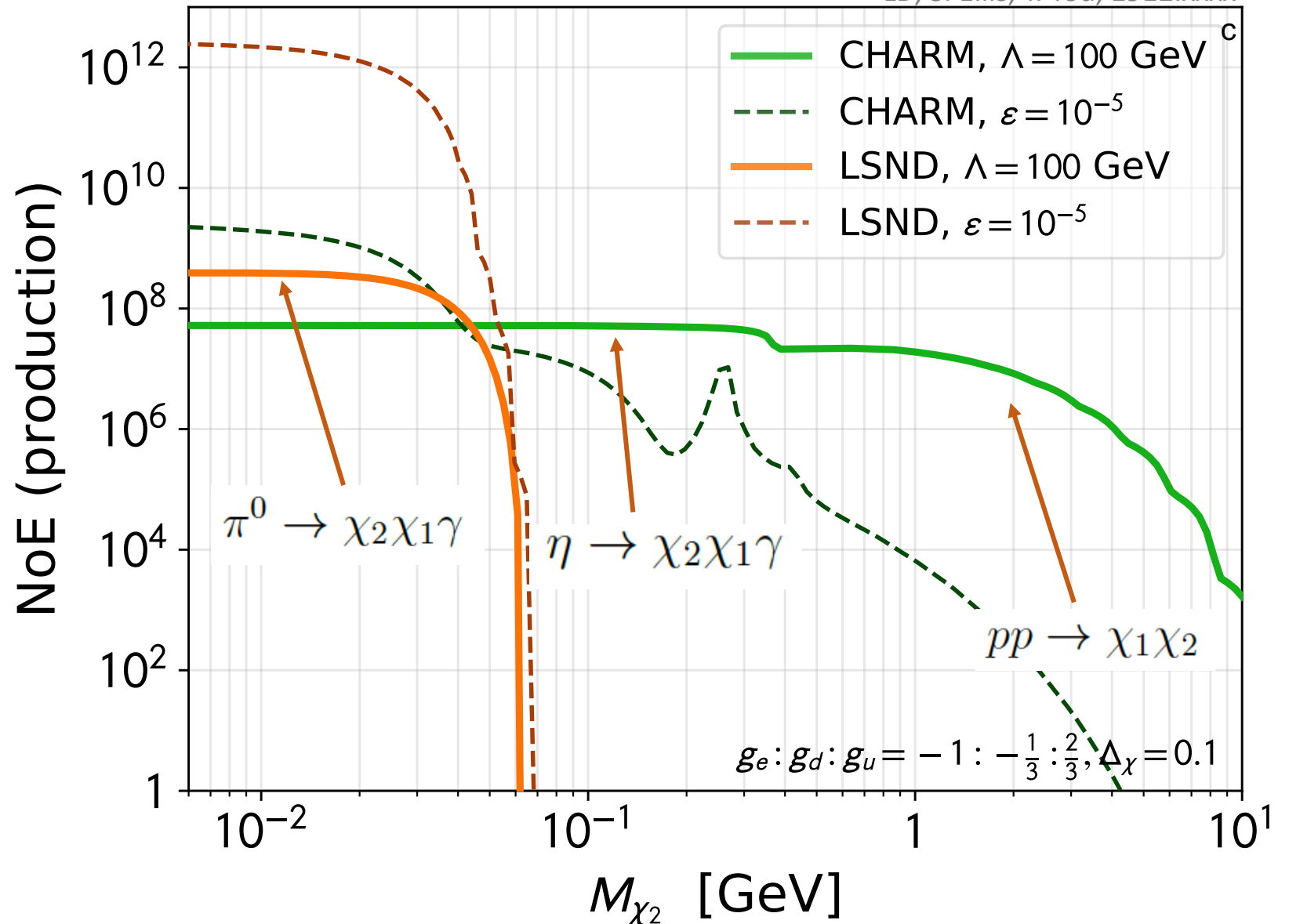
Full production – vector coupling

LD, S. Ellis, T. You, 1912.xxxx

- Main exp. properties
 - LSND: $\sim 10^{23}$ PoT and 0.8 GeV beam
 - CHARM: $\sim 10^{18}$ PoT and 400 GeV beam
- Strong differences with dark photon case
- Meson decay allowed for VV operator:

$$\pi^0, \eta, \eta' \rightarrow \gamma \chi \chi$$

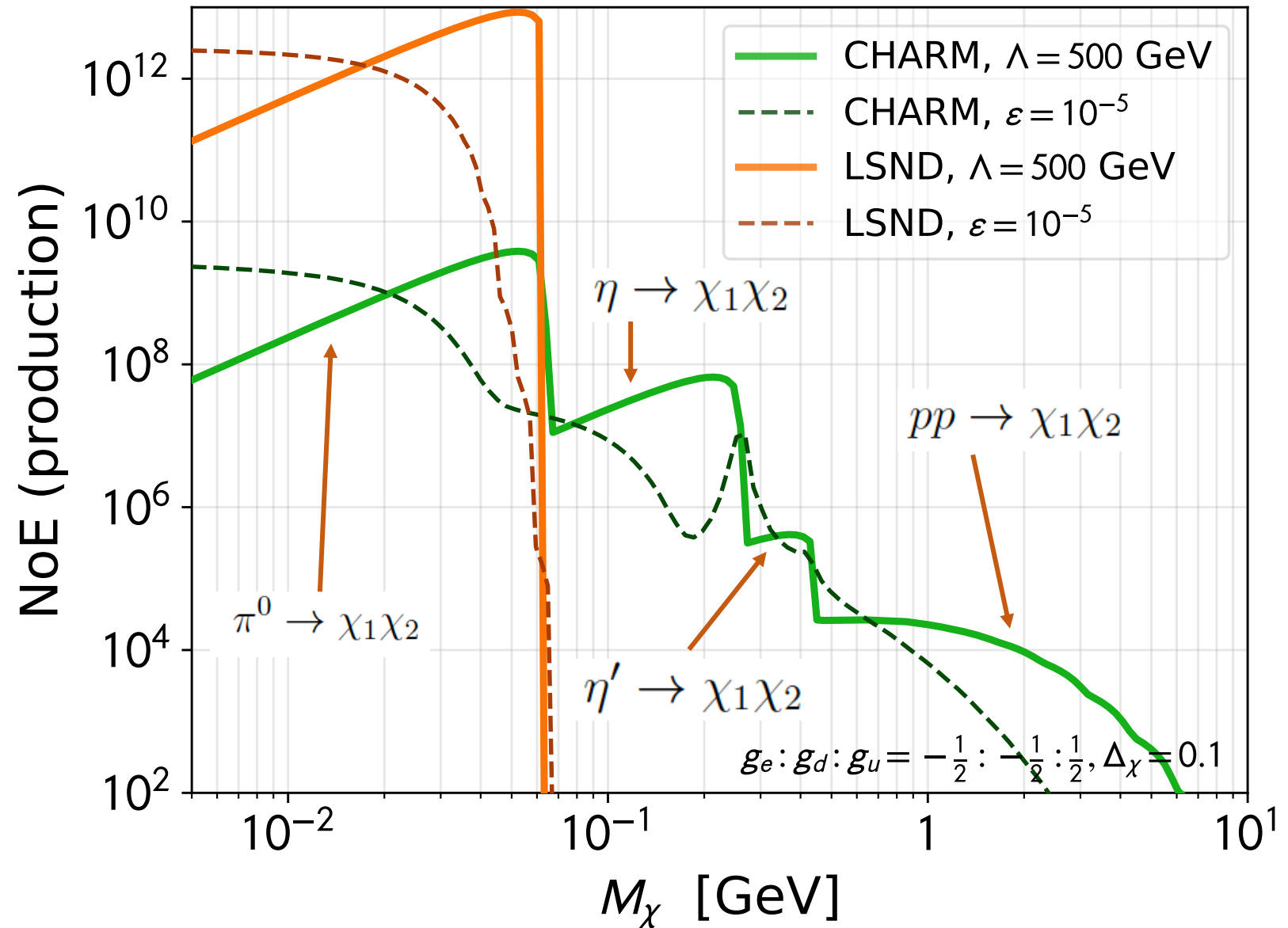
$$\rho, \omega \rightarrow \chi \chi$$



Full production – axial vector

LD, S. Ellis, T. You, 1912.xxxx

- Meson production strongly enhanced
 - 2-body decays dominant
- $$\pi^0, \eta, \eta' \rightarrow \chi\chi$$
- Pion decay contribute significantly

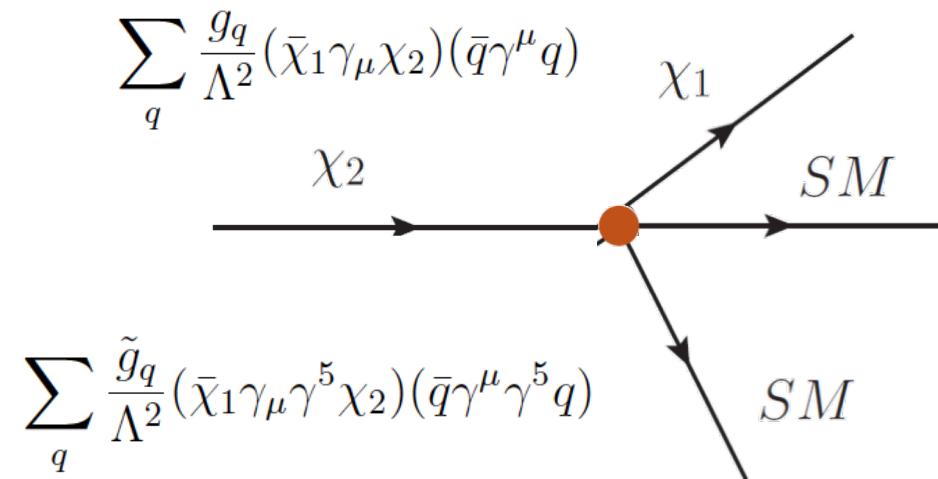
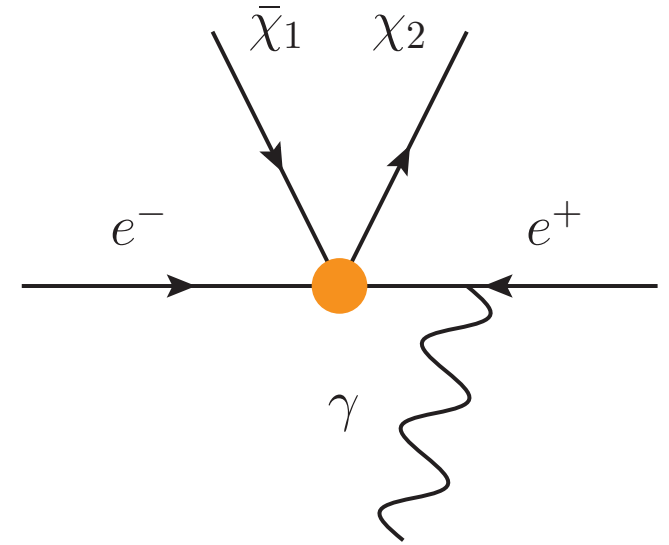


Dark sector searches and constraints

Recasting and limits using the EFT approach

Dark Sector searches in the lab

- **Missing energy/ Invisible decay:** Mono-photon/mono-jet searches missing energy signature @ BaBar, Belle, NA64, LEP, LHC.
- **Invisible meson decay:** for instance $\pi^0 \rightarrow \bar{\chi}\chi$ (NA62)
 - Important for flavour-violating operators
- **Dark sector beam production and detection**
 - **Scattering:** Searching for DM via scattering (E137, LSND, miniBooNE ...)
 - **Dark sector visible decay:** (LSND, CHARM, Seaquest, FASER, etc...)



Recasting in the light EFT approach

Most existing limits are obtained for vanilla cases (e.g iDM, pure dark photon ...) → need to recast these searches as function of the EFT

- Mono—photon searches are also weakened since no “bump-search” can be performed in $\chi\chi$ invariant mass
- Different approaches for each search strategies → Decay limits are particularly challenging
 - need to rescale for production rates

$$\Lambda_{\text{lim}} = 410 \text{ GeV} \times \sqrt{g_{\text{eff}}} \left(\frac{0.001}{\varepsilon} \right)^{1/2} \left(\frac{\mathcal{N}_{\text{prod}}^{\text{eff}}}{\mathcal{N}_{\text{prod}}^{\text{DP}}} \right)^{1/8}$$

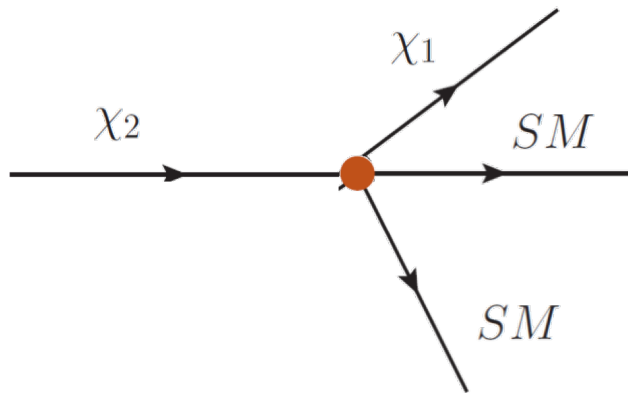
→ For different splitting, detection probability is modified (also rescale for decay rates)

Limits in the vector case

Include limits/projections:

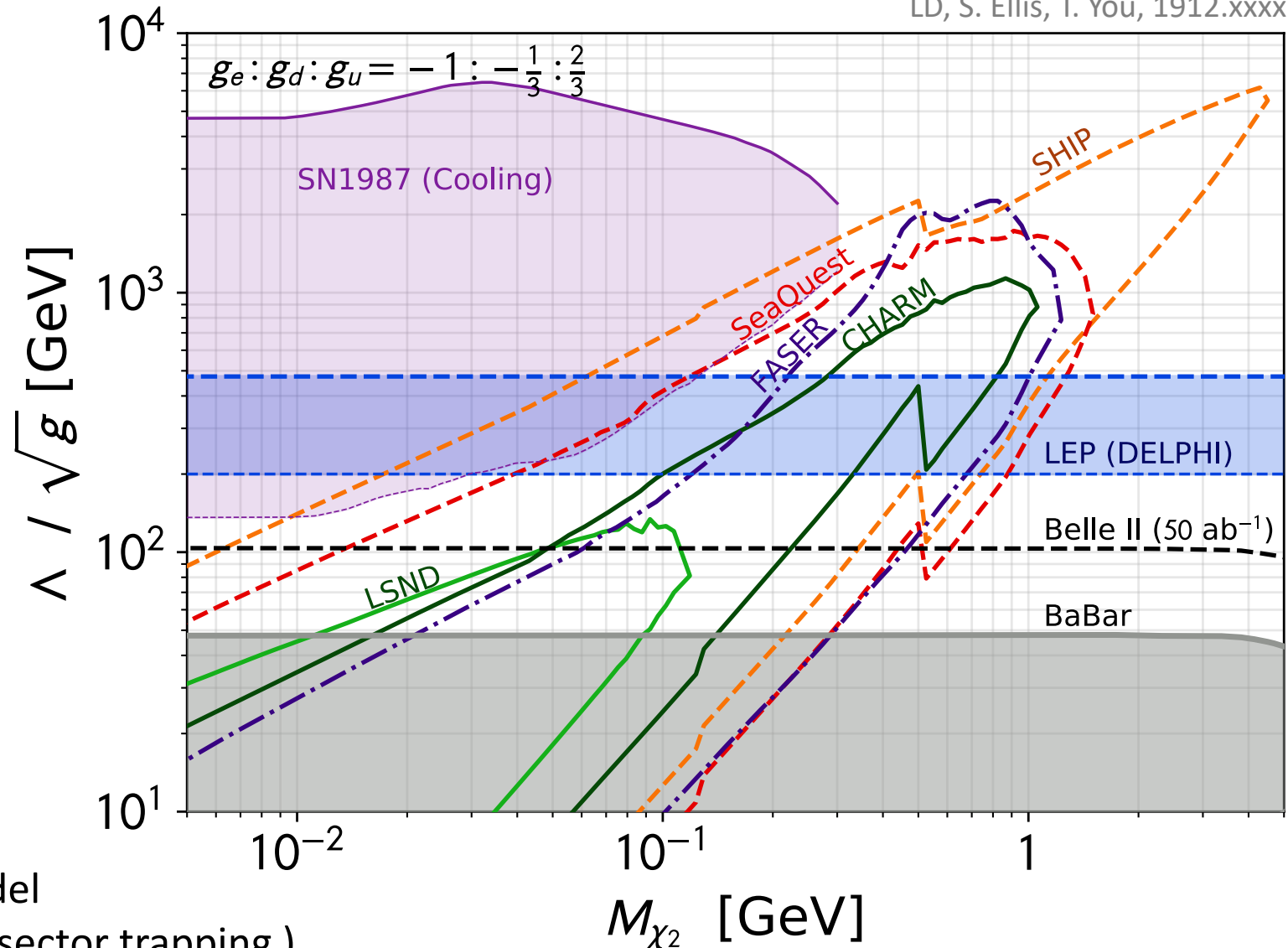
→ Mono-photon: LEP, BaBar and Belle II

→ Decay searches at saturation ($M_2 \gg M_1$) at LSND, CHARM, SeaQuest (hypothetical Phase 2 with $\sim 10^{18}$ PoT) and SHIP



→ SN1987 cooling limits, but strong model dependence in the lower bounds (dark sector trapping)

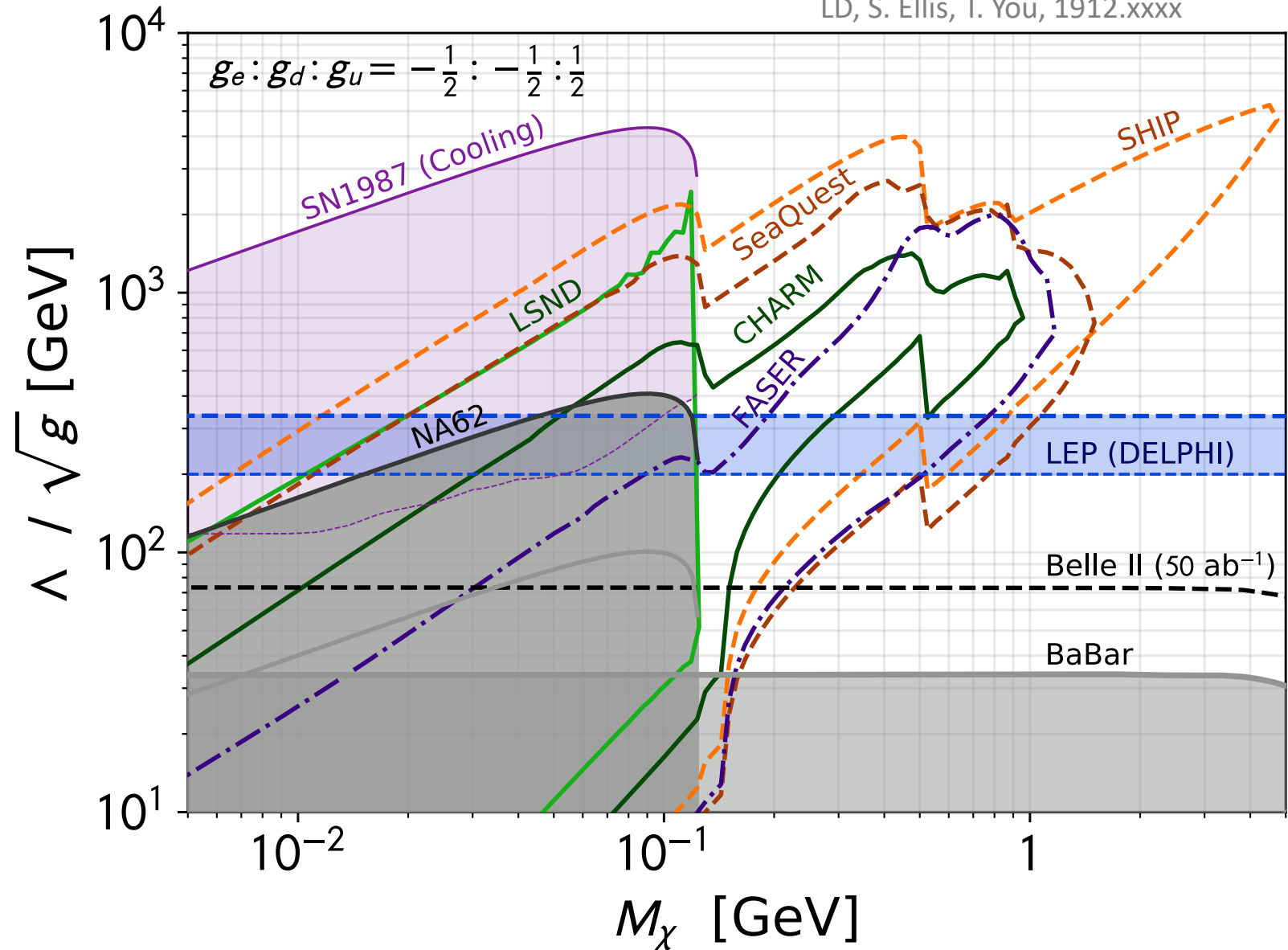
LD, S. Ellis, T. You, 1912.xxxx



Limits in the axial-vector case

LD, S. Ellis, T. You, 1912.xxxx

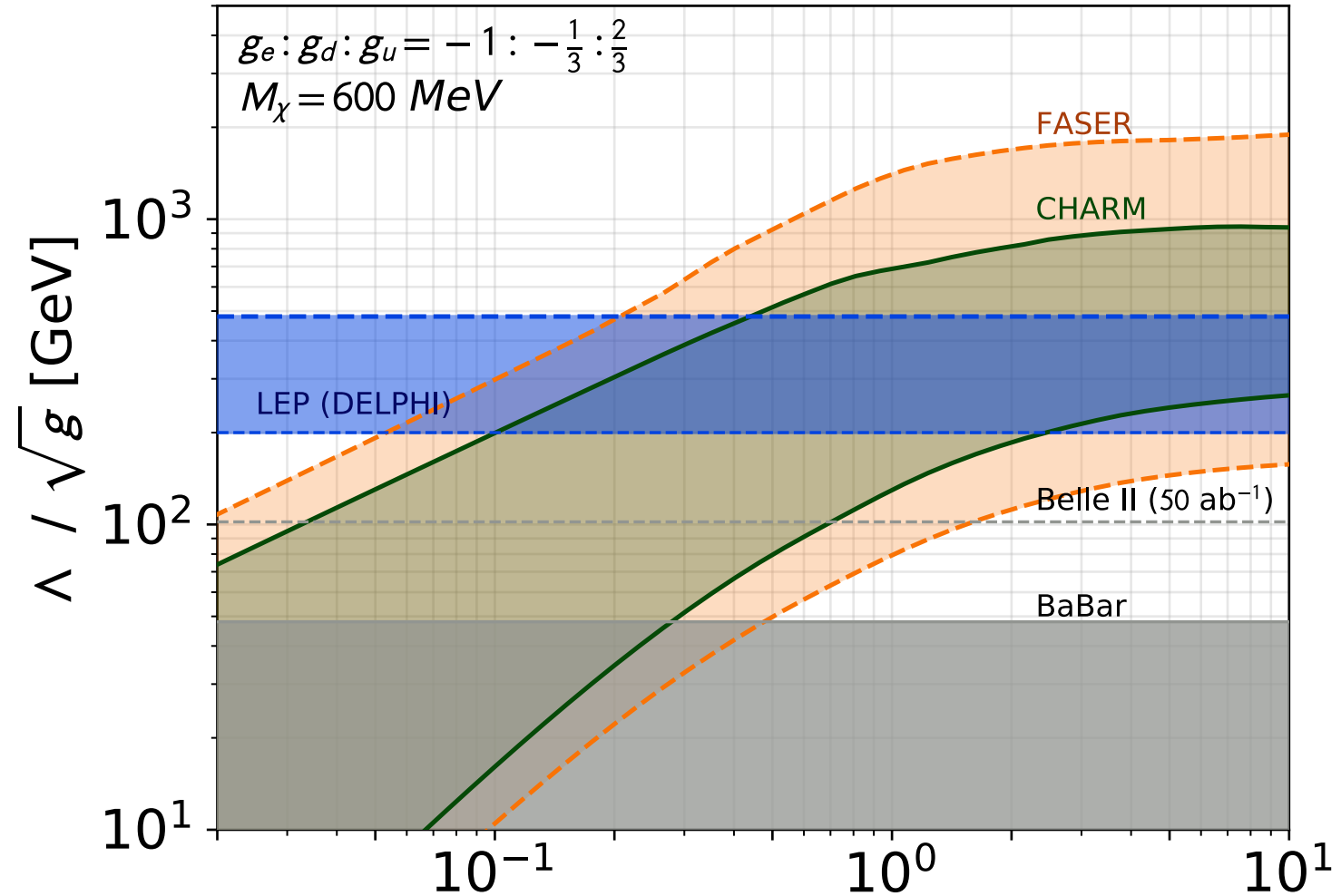
- Mesons production strongly enhanced
- Better low-mass limits
 - LSND (0.8 GeV beam) probes up to 1 TeV
- SN1987 based on invisible π^0 decay
- All limits based only on first generation couplings



Varying the splitting

LD, S. Ellis, T. You, 1912.xxxx

- Decay signatures depends strongly on splitting $M_2 - M_1$
 - Lifetime scales as $(M_2 - M_1)^{-5}$
 - Then reach saturation for $M_2 \gg M_1$
- Both upper limits and lower limits are modified
 - Long-lived limit -> linear suppression
 - Short-lived limit -> exponential dependence

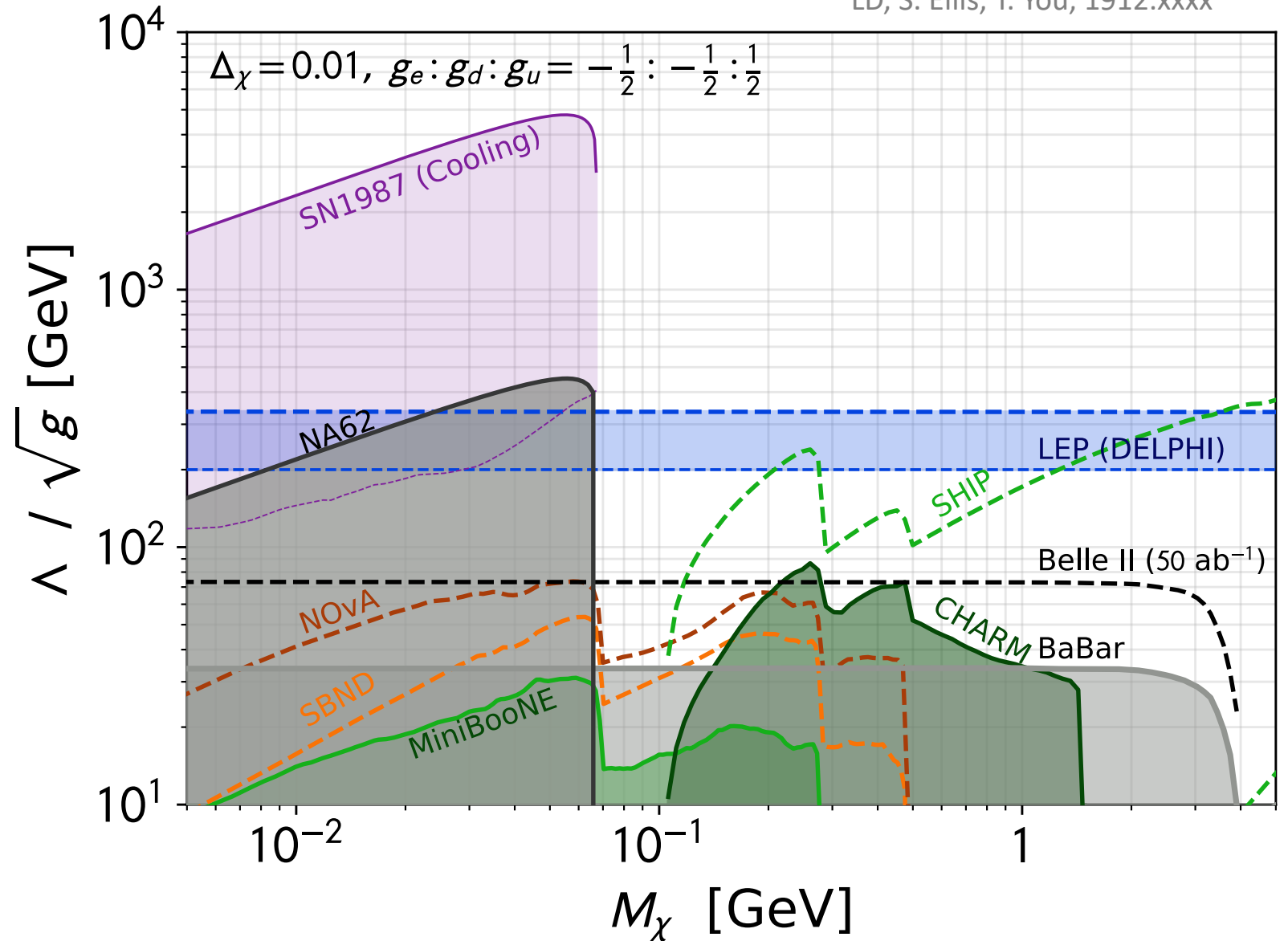


$$c\tau^{\text{PD}} \sim 375 \text{ m} \times \left(\frac{100 \text{ GeV}}{\Lambda}\right)^4 \left(\frac{1 \text{ GeV}}{M_{\chi 1}}\right)^5 \left(\frac{0.25}{\Delta_\chi}\right)^5 \left(\frac{0.01}{g}\right)^2 \quad \Delta_\chi \equiv (M_2 - M_1) / M_1$$

Small-splitting limits

LD, S. Ellis, T. You, 1912.xxxx

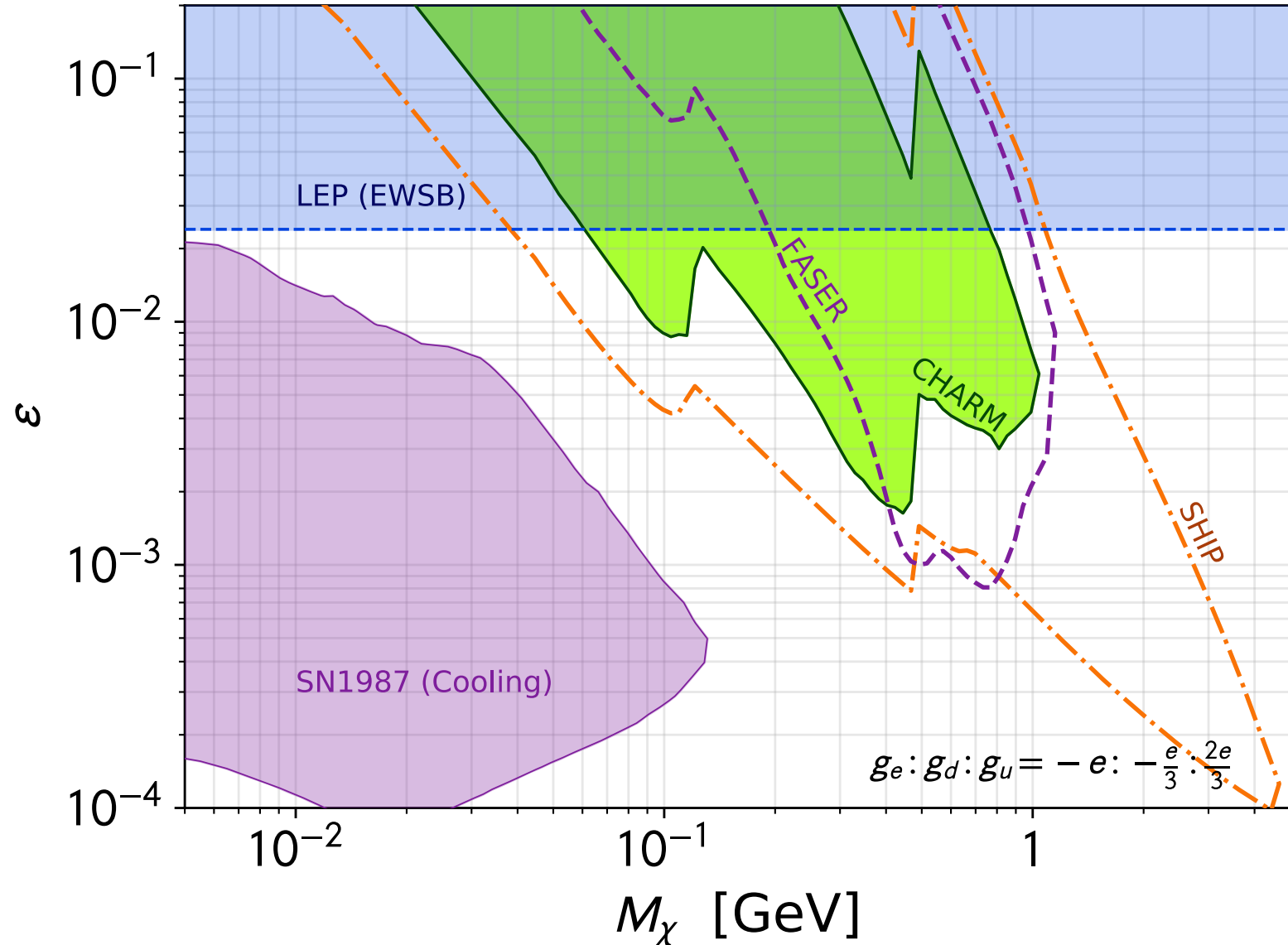
- Invisible meson decay
 - Recent NA62 on $\pi^0 \rightarrow inv$
- Scattering limits in future neutrinos experiment can play a role
- Decay limits shifted to higher mass ($2m_e$ threshold)



Practical example: off-shell dark photon

LD, S. Ellis, T. You, 1912.xxxx

- Standard iDM scenario with a heavy dark photon ($M_V \sim 30$ GeV, with $M_V \gg M_\chi$)
- Very weak limits from Babar (no resonance search available)
- Relic density through e.g. $\chi_1 \bar{\chi}_1 \rightarrow SS$ dark Higgs boson



Conclusion

Looking forward ...

- Many upcoming relevant experiments:
 - **Neutrino** experiments -> the near detectors can search for dark sector particles
 - **Dark sector-oriented** -> looking for decays/ missing energy
 - **Flavour/ Rare mesons** decay -> Missing energy searches, invisible meson decay, etc...

Belle II, **BDX@Jlab**

PADME

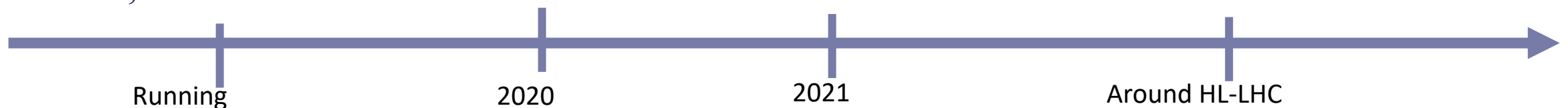
SBN, SeaQuest,

NA62, KOTO

LHCb Run 3,
FASER, NA62+

MAGIX,
BDX @MESA

MATHUSLA, SHIP,
CODEX-b, KLEVER
LBNF/DUNE



(Many missing, not all of them are funded yet...)

Conclusion

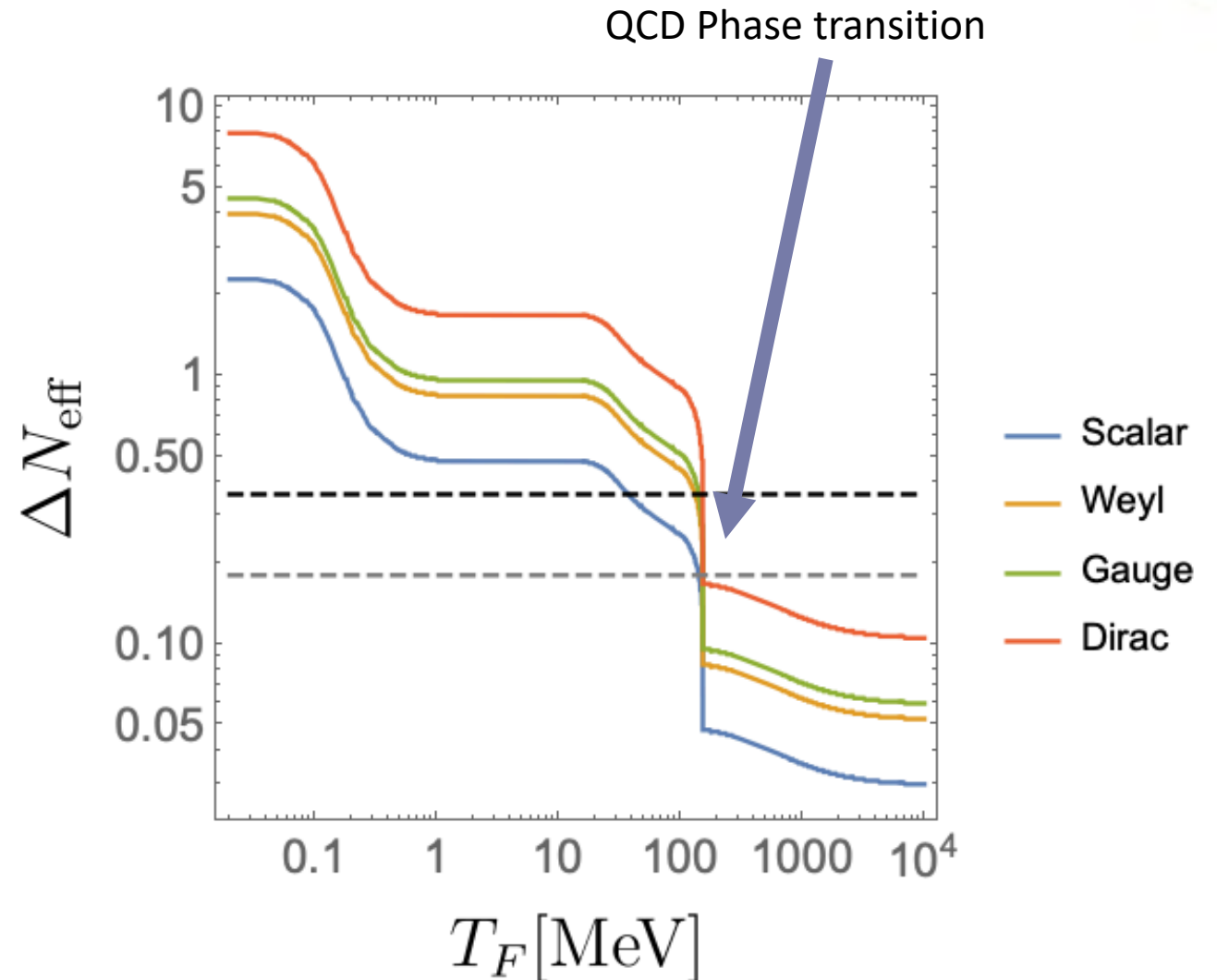
- Light thermal dark matter models typically include a dark sector with long-lived particles -> Important search targets for intensity frontier experiments
 - When the mediator is too heavy to be produced directly, describe the phenomenology as an “off-shell” fermion portal -> EFT description
 - Lead to rich phenomenology in intensity frontier experiments, with different prospects than standard “on-shell” portals
- Will be release in a python package, to provide recasted limits for any effective coupling

Backup slides

Astrophysical limits

- For $M_2 \gg M_1$, the lightest dark sector can be relativistic relic
- One can still obtain dark matter candidate for iDM setup for masses around the GeV

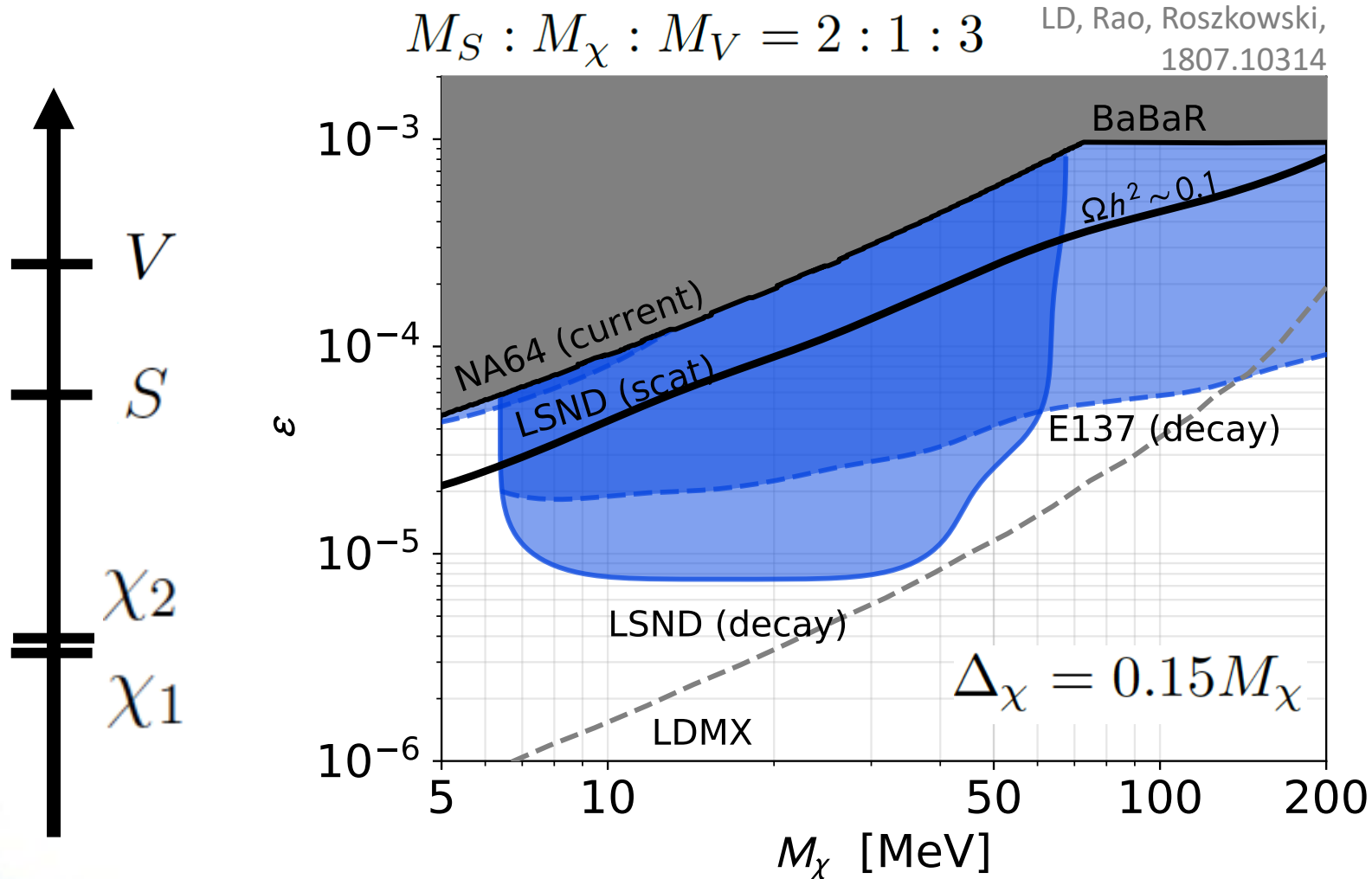
$$\Omega h^2 \sim 0.3 \times \left(\frac{2 \text{ GeV}}{M_\chi} \right)^2 \left(\frac{\Lambda/\sqrt{g}}{500 \text{ GeV}} \right)^4$$



- Additional dynamics in the hidden sector may fix the relic density, e.g. $\chi_1 \bar{\chi}_1 \rightarrow SS$ of iDM with a dark Higgs boson

Inelastic DM regime

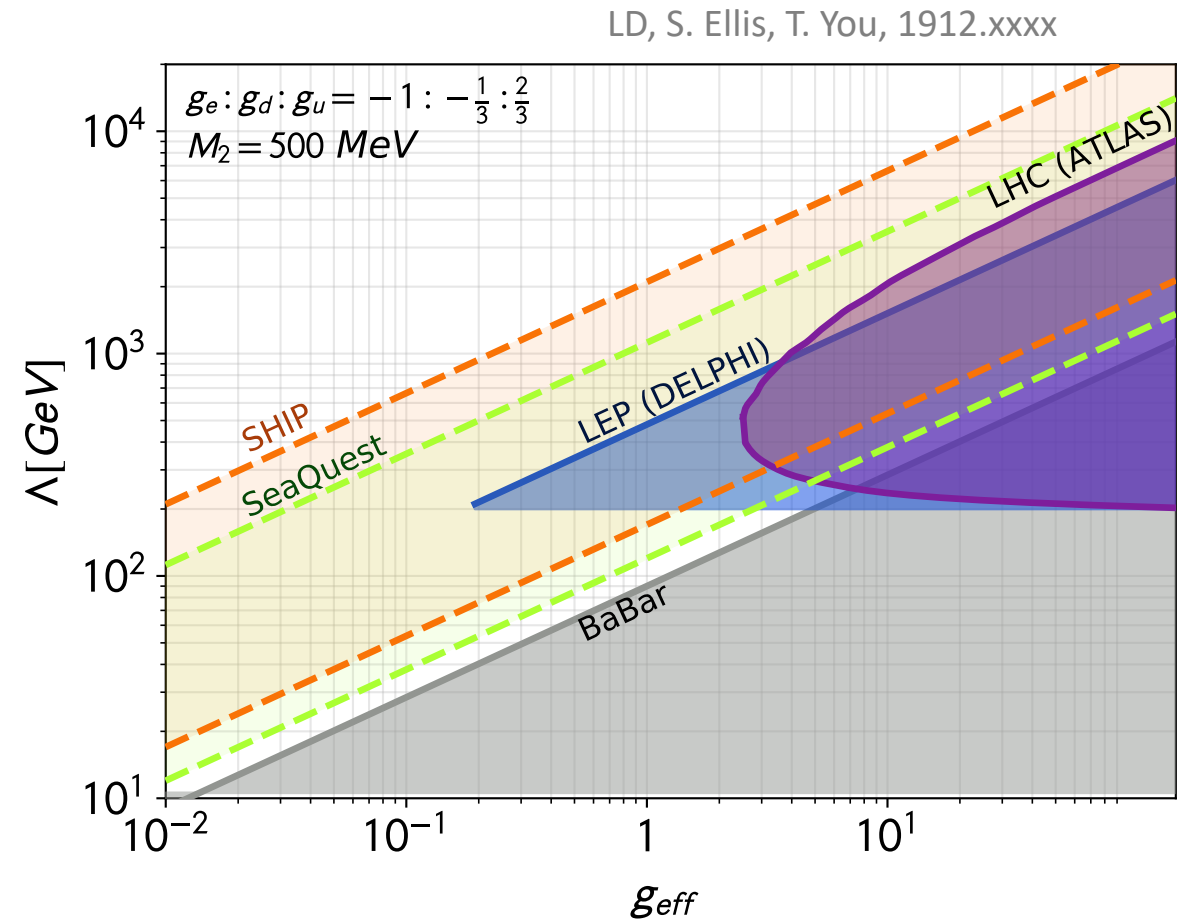
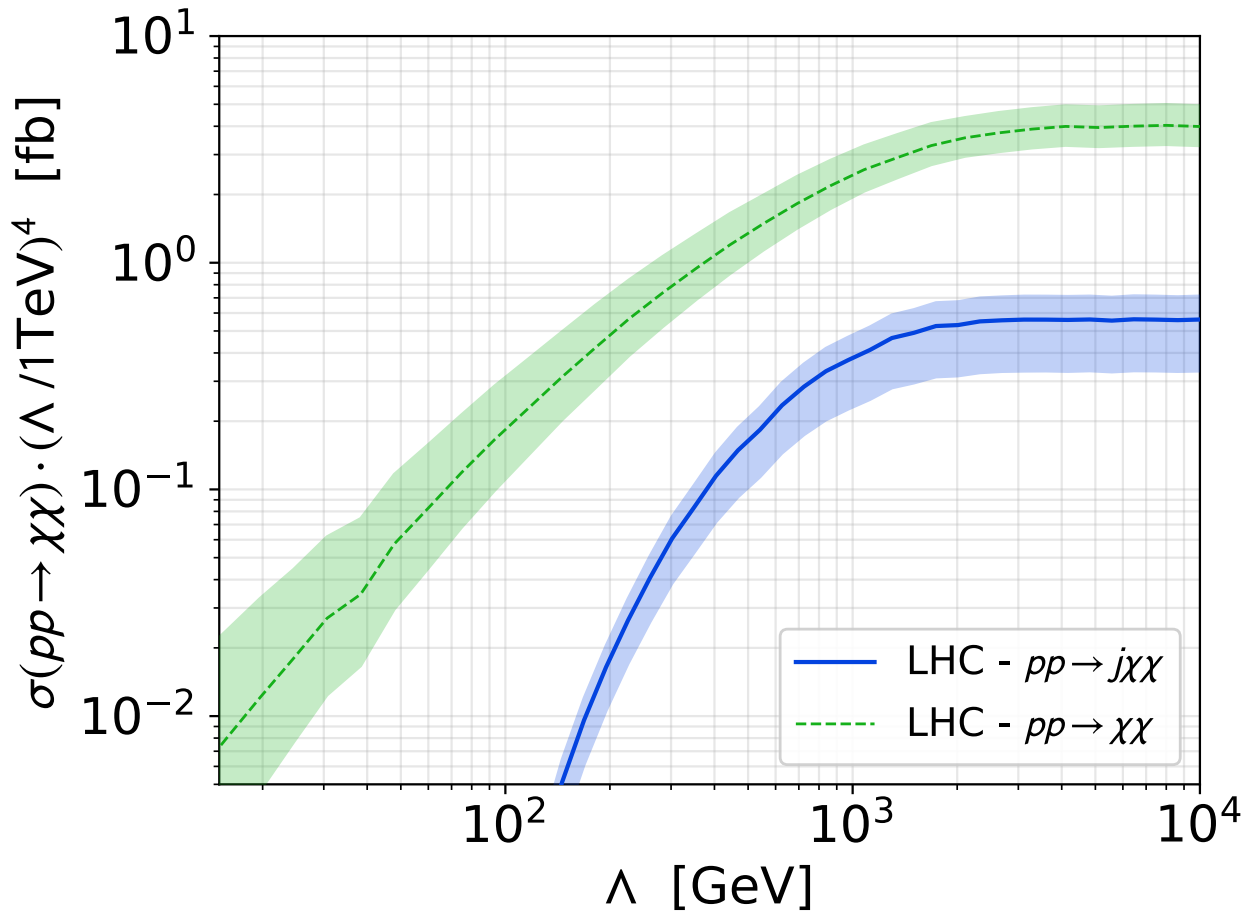
- Relic density fixed by s-channel, co-annihilation process: $\chi_1\chi_2 \rightarrow e^+e^-$



- Main signatures:
 - Missing energy searches
 - $\chi_2 \rightarrow \chi_1 e^+ e^-$ decay
 - χ_1 scattering
- When consider dark sector decays, decades-old experiment are still strongly ahead of current mono-photon searches!

EFT limitation at LEP and LHC

- EFT not applicable if roughly the c.o.m energy of the process higher than the scale \rightarrow significantly discussed for dark matter at LHC



SN 1987A bounds

- Typical bounds arise when DM do not scatter enough and escape the SN core and escape the SN core

$$\alpha_D \epsilon^2 < O(\text{few}) \times 10^{-14}$$

- Not relevant for pseudo-Dirac case/Majorana case at the thermal target
- Dark Higgs bounds may be relevant at $m_S < M_{\chi_1}$ or $m_S < M_{\chi_2} - M_{\chi_1}$

→ But scattering with DM halo inside the SN should be enough to trap it

