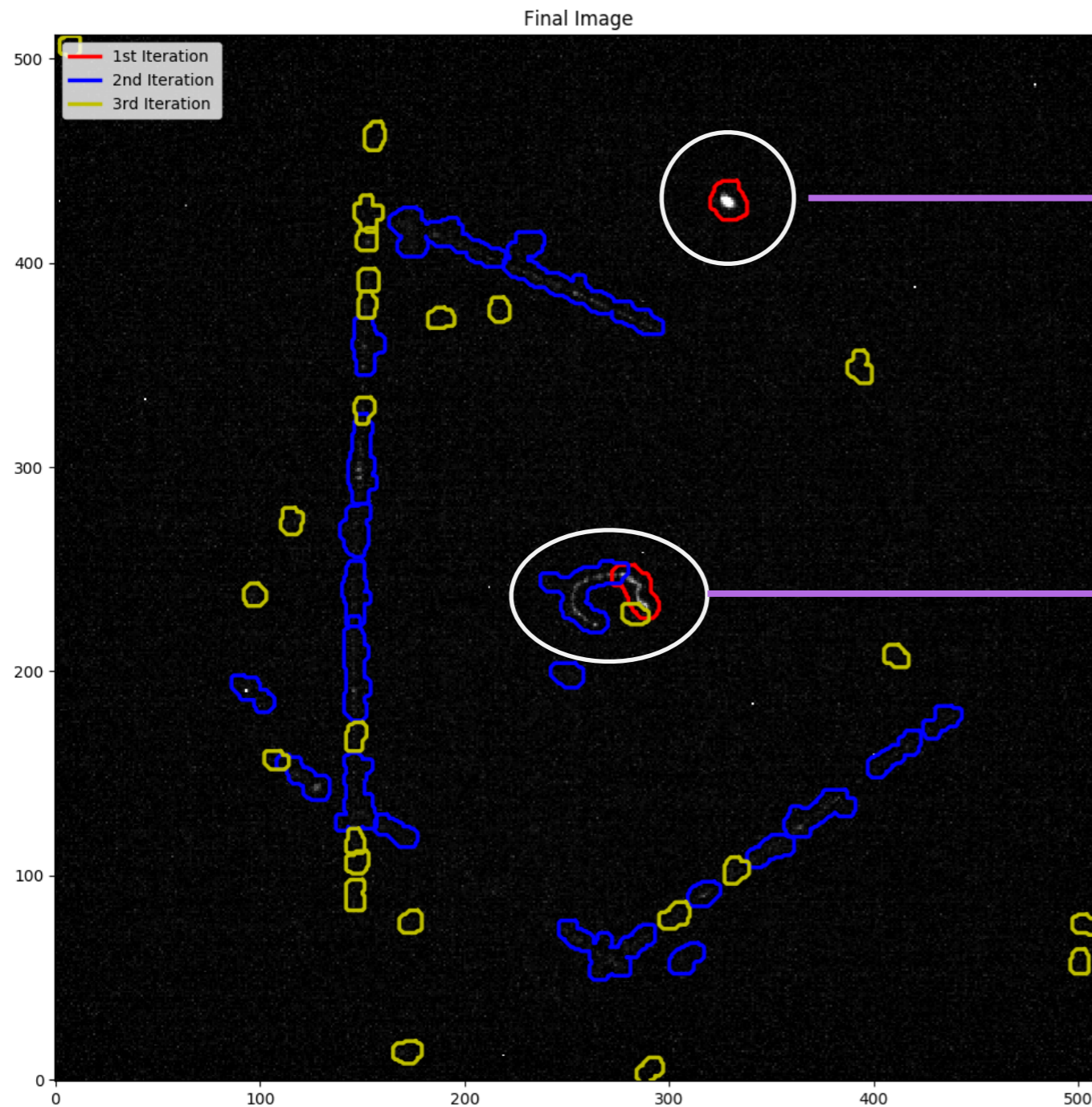


New data taking with AmBe and background rejection

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- Tuned the cluster reconstruction parameters to get efficiently α 's and what seem nuclear recoils in iteration 1, and the rest into iteration 2
 - reminder: **iteration 1 is made for "high-density" clusters**
 - **iteration 2 is made for medium/low-energy clusters** (e.g. 5.9 keV spots from Fe55, typical ambient radioactivity, cosmics products...)
- Achieved:
 - all the Fe55 is well separate (checked with the energy peak) => go to iter-2
- Problem:
 - there are clusters, pieces of the ubiquitous cosmics, that go into iteration 1
 - **=> iter-1 is not a pure sample of recoils**
- Possible solutions:
 - these are pieces of longer tracks, so they could be rejected by the supercluster length
 - at the moment not done, because the supercluster runs on separate classes (iterations) of clusters
 - **=> need to run the superclustering on the OR of it1 + it2 clusters**



Good recoil?

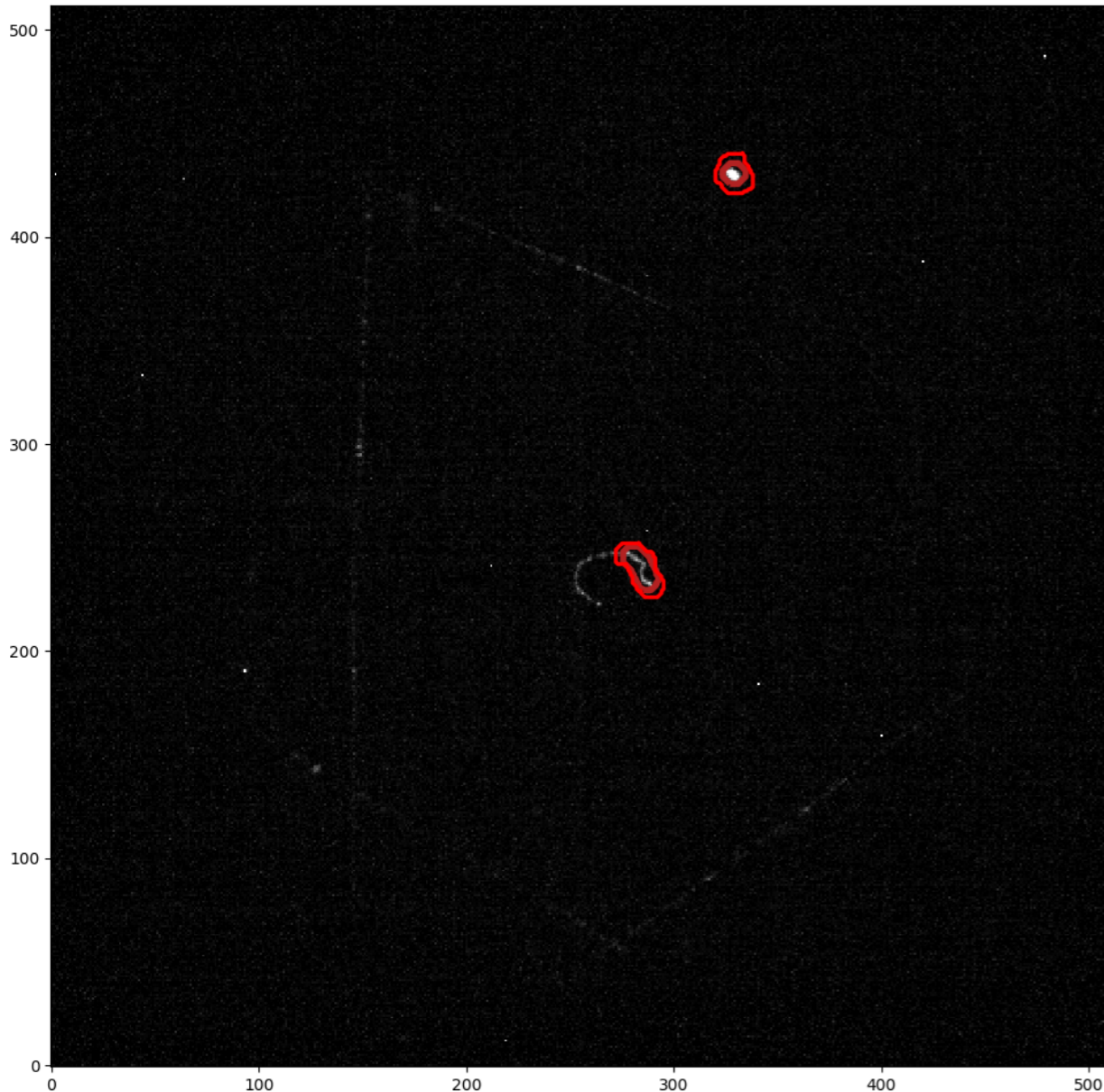
Two clusters with different iterations clearly from the same track

Clustering step

super-clusters in AmBe run

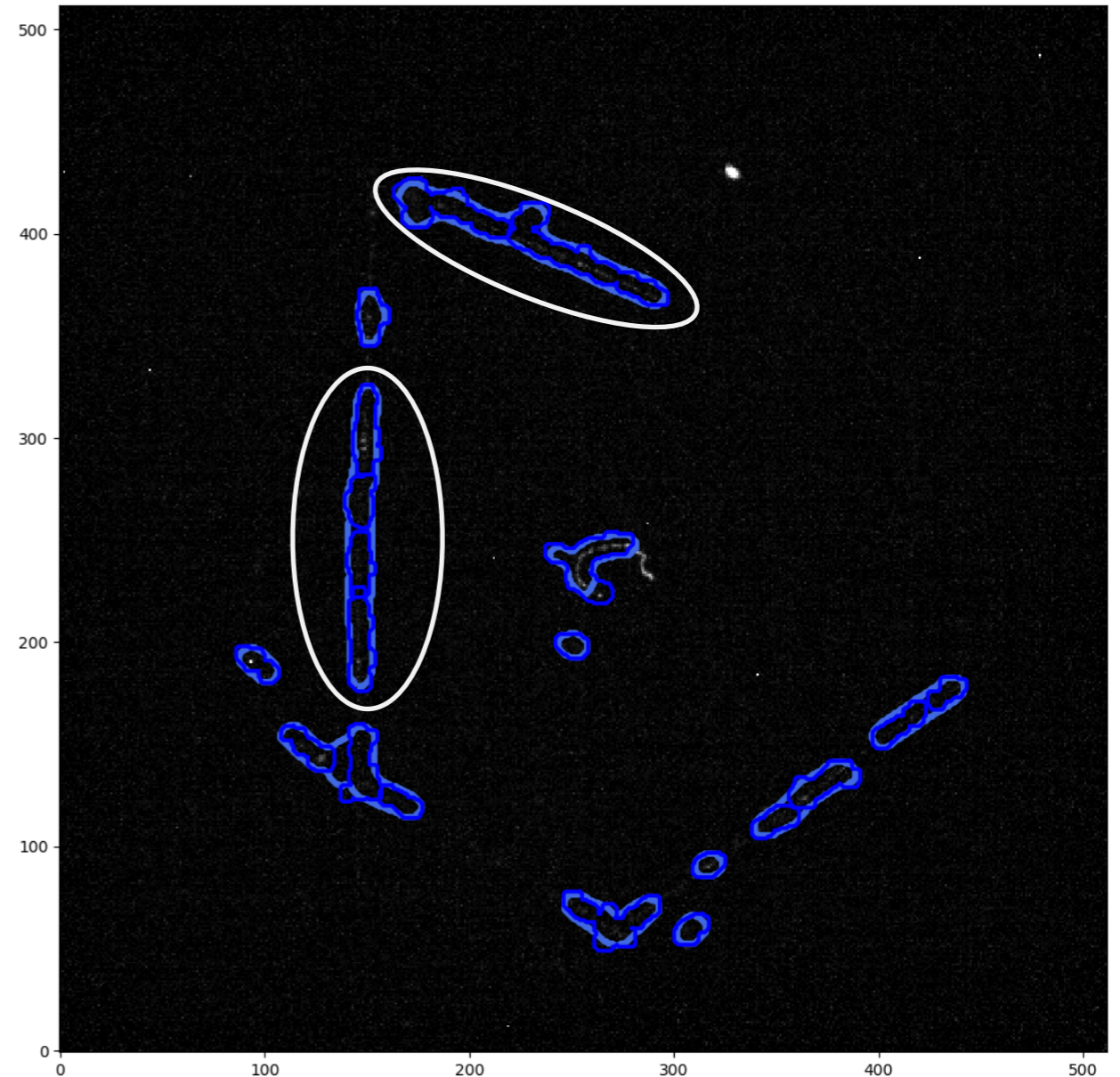
iteration 1: high density

Clusters found in iteration 1



iteration 2: medium density

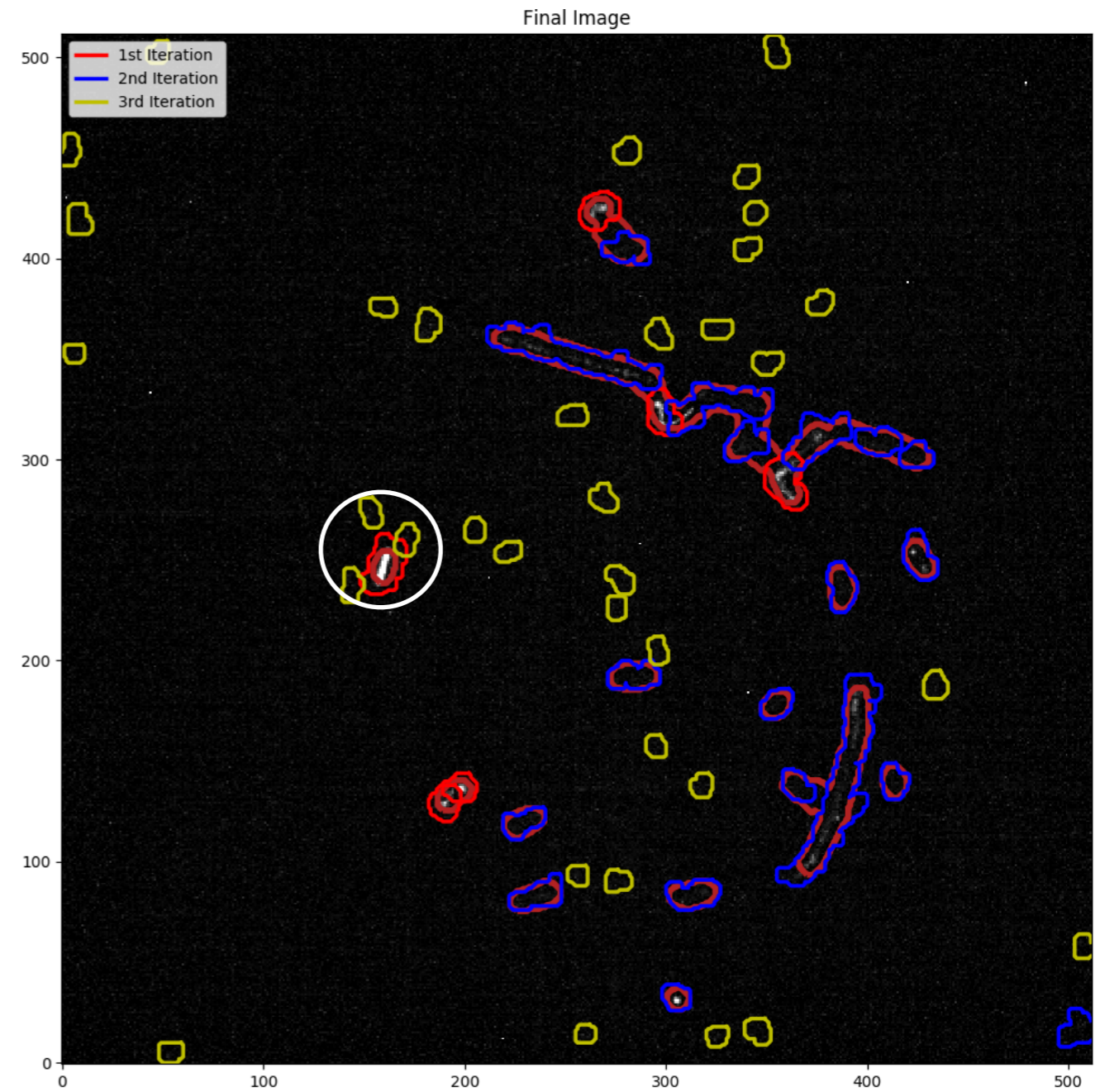
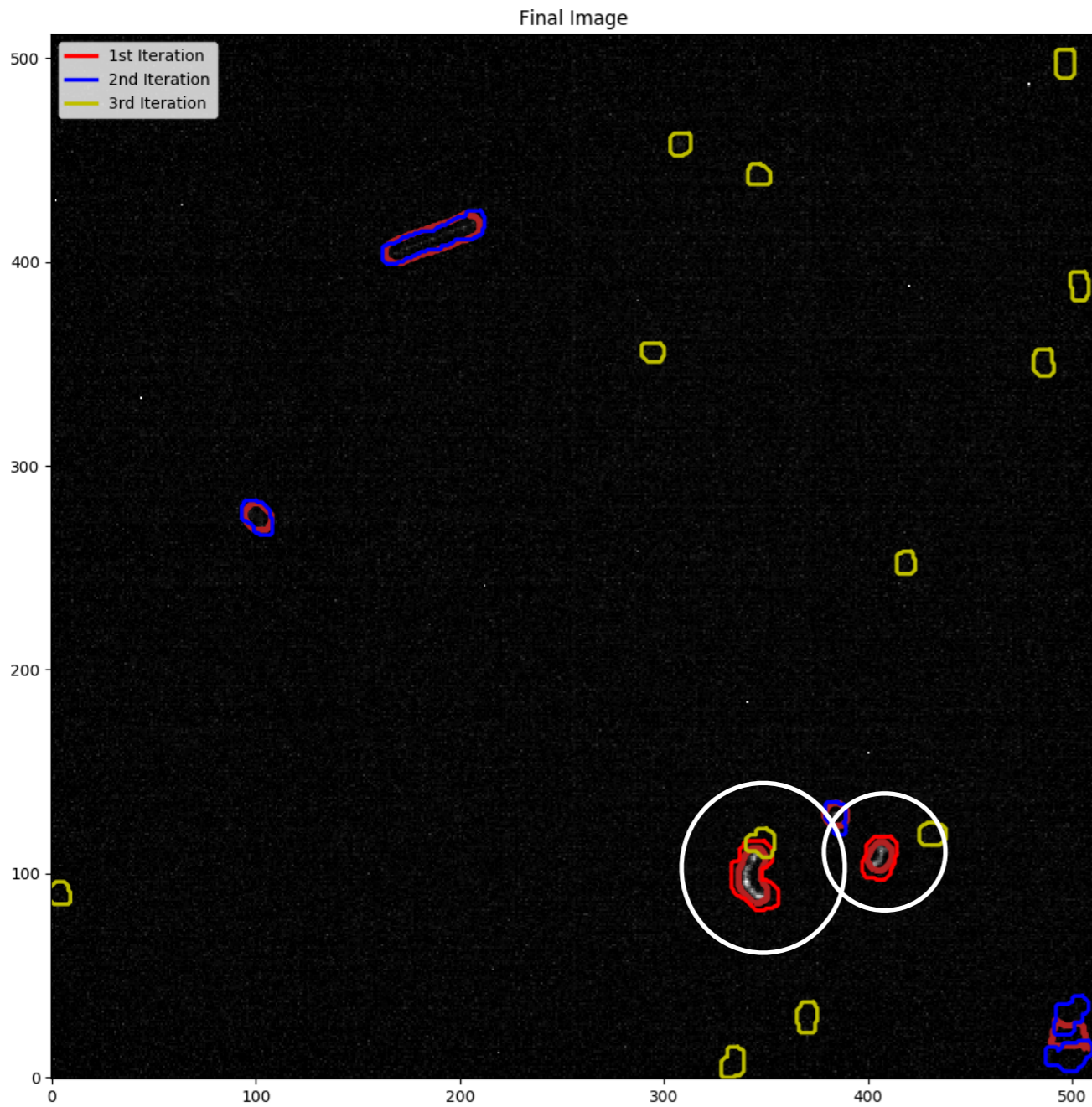
Clusters found in iteration 2



superclusters in it2 join pieces of tracks belonging to the same iteration, but by construction cannot join it1 + it2

Will change that to help the discrimination against “cosmics”
side note: not a big problem underground...

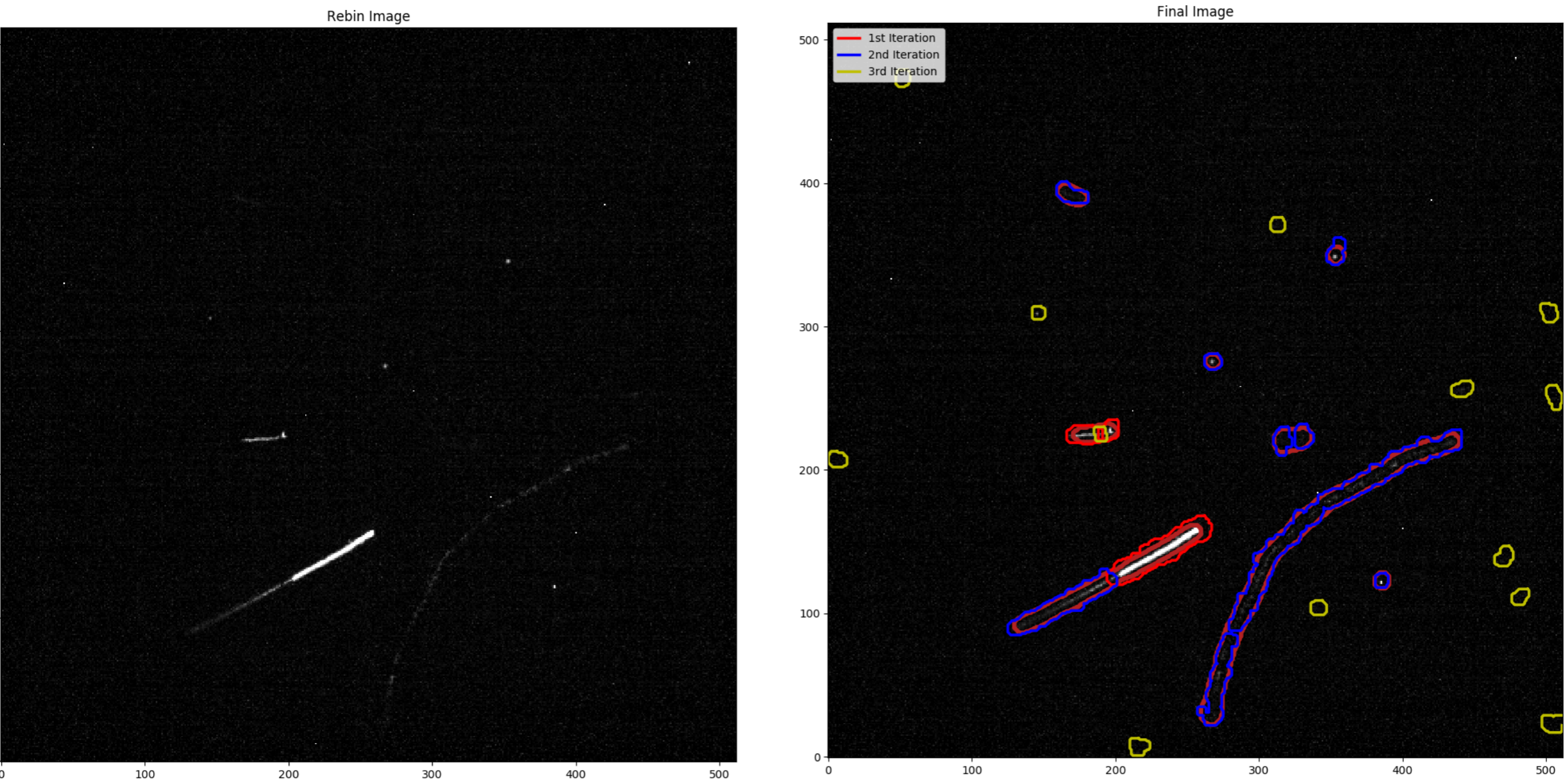
typical candidates



e.g. of 1 selected cluster

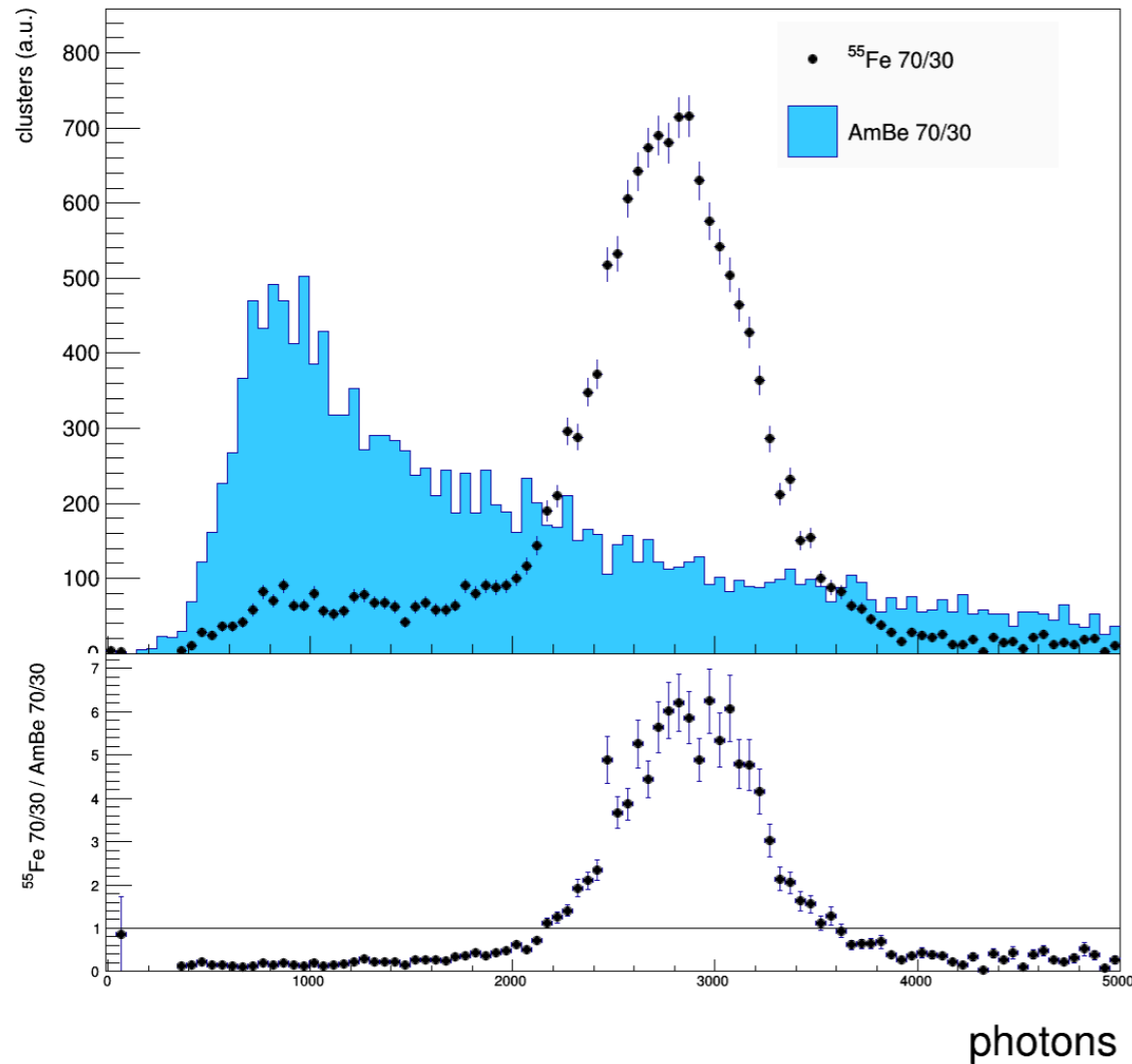
- One strange example with one **probably proton** track with two different “light-yield zones”

P.S. already implemented the supercluster inclusive wrt iterations



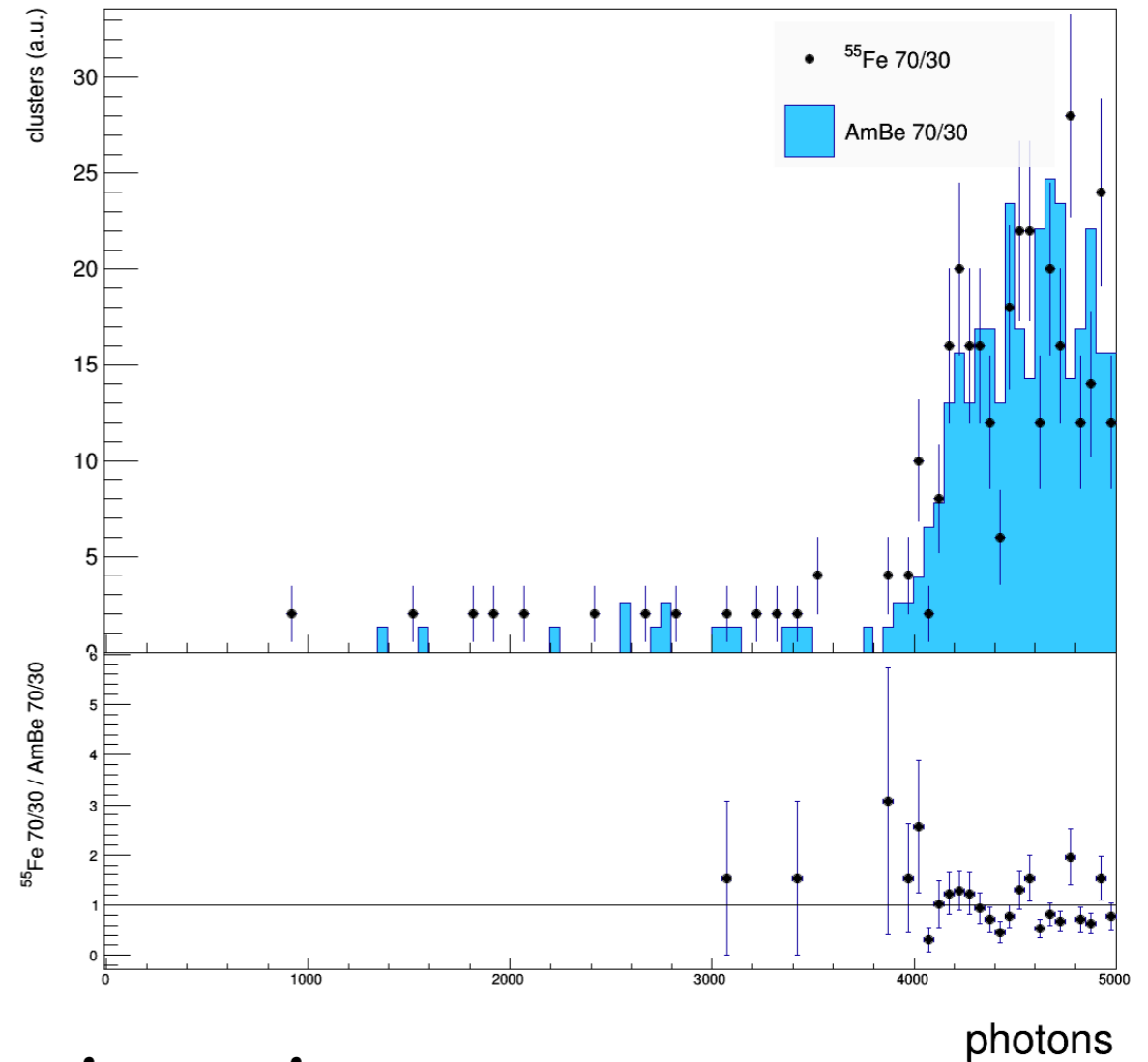
^{55}Fe vs AmBe: energy of the 2 classes

- Reconstructed Fe55 and AmBe with the same reconstruction parameters (runs of the same day)



iteration 2:

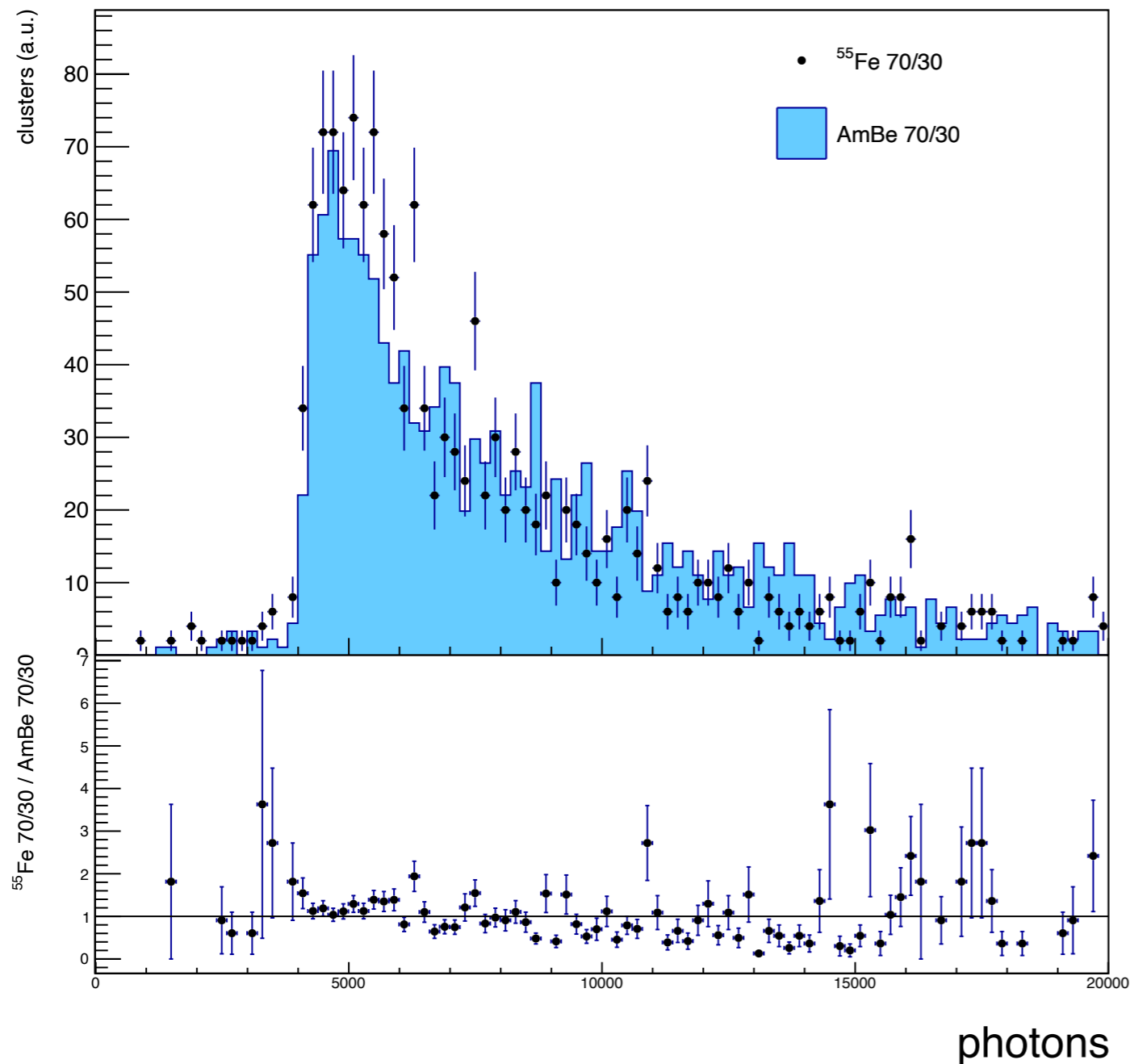
- gets the Fe55 peak (~2.8k photons)
- gets the bkg (note: same shape in Fe and AmBe)



iteration 1:

- do not get any part of the Fe55 peak (i.e. it's a decent discriminator)
- gets higher energy (full spectrum later)

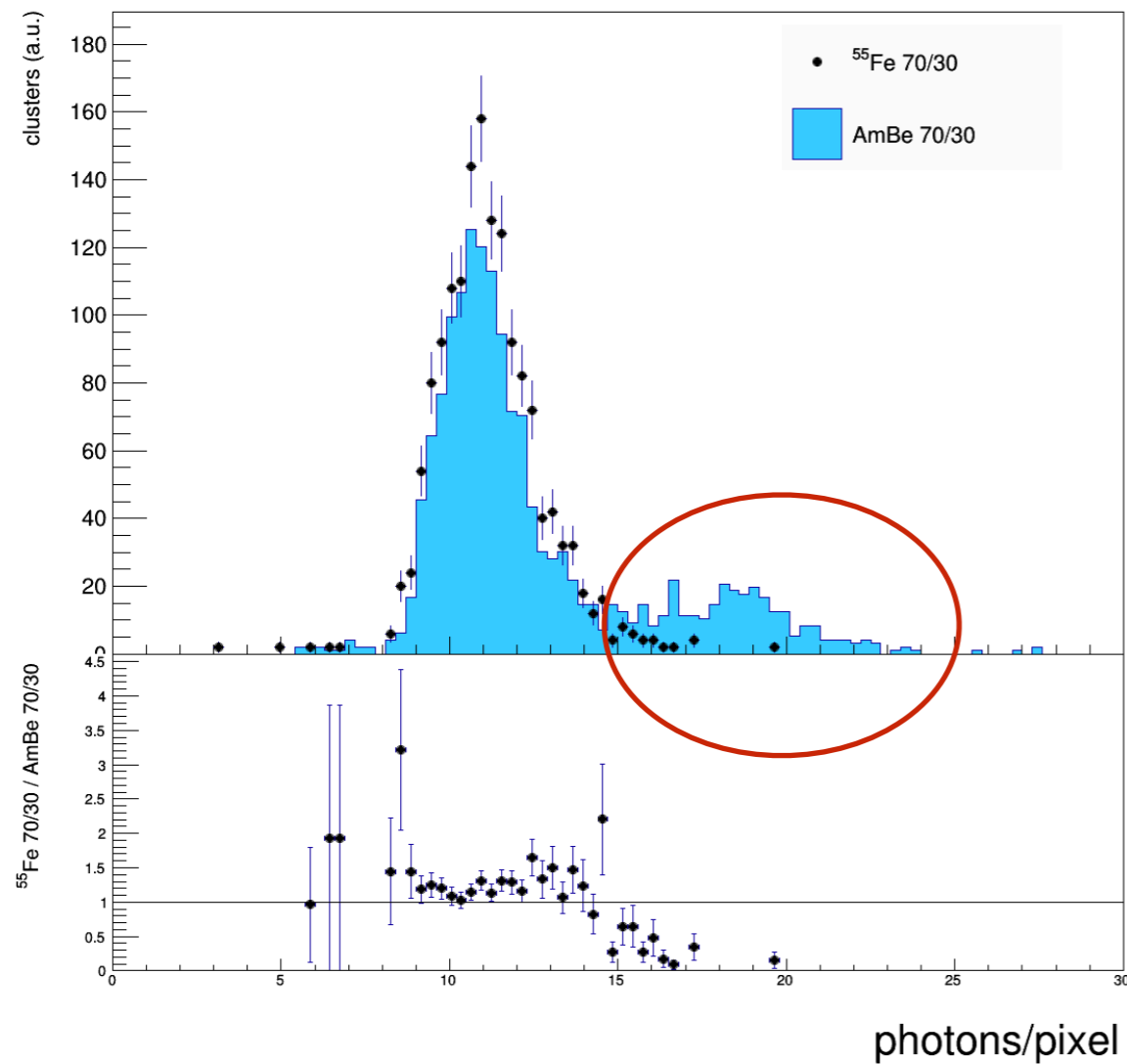
^{55}Fe vs AmBe: energy of iter-1



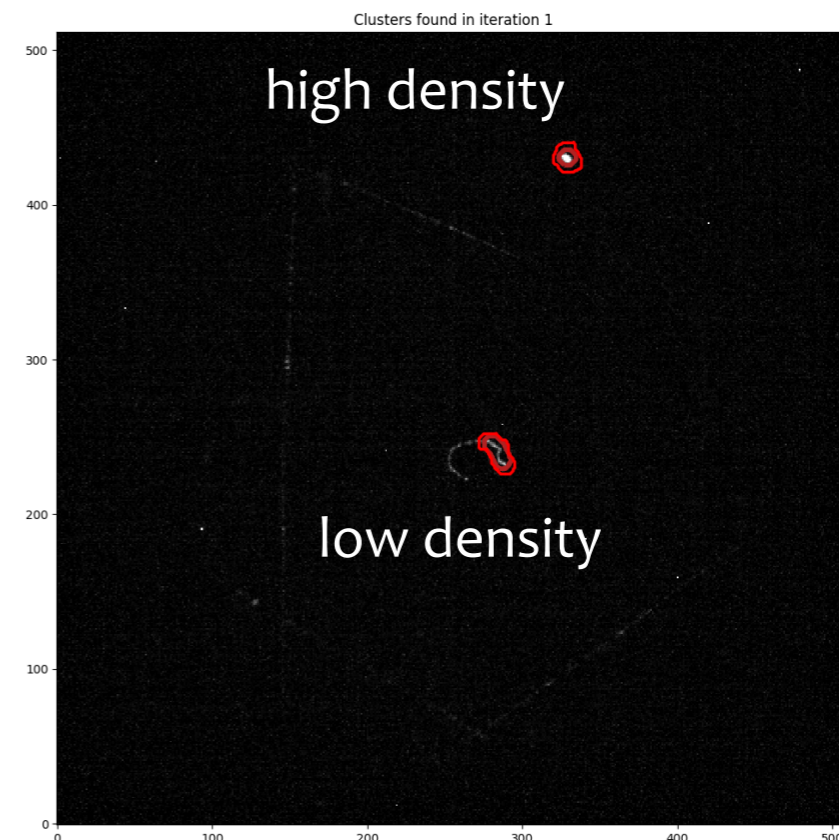
- The spectrum **in the mid-range** (calibrated energy in [10.3 - ~40 keV]) is very similar between the AmBe and Fe55 run
- **indeed these are MOSTLY pieces of cosmic-induced background, which are present both in Fe55 and AmBe runs**
- Also, the AmBe source was screened a lot, so it's a small contribution to this spectrum (if any, in the tails also, see later...)

do we see anything?

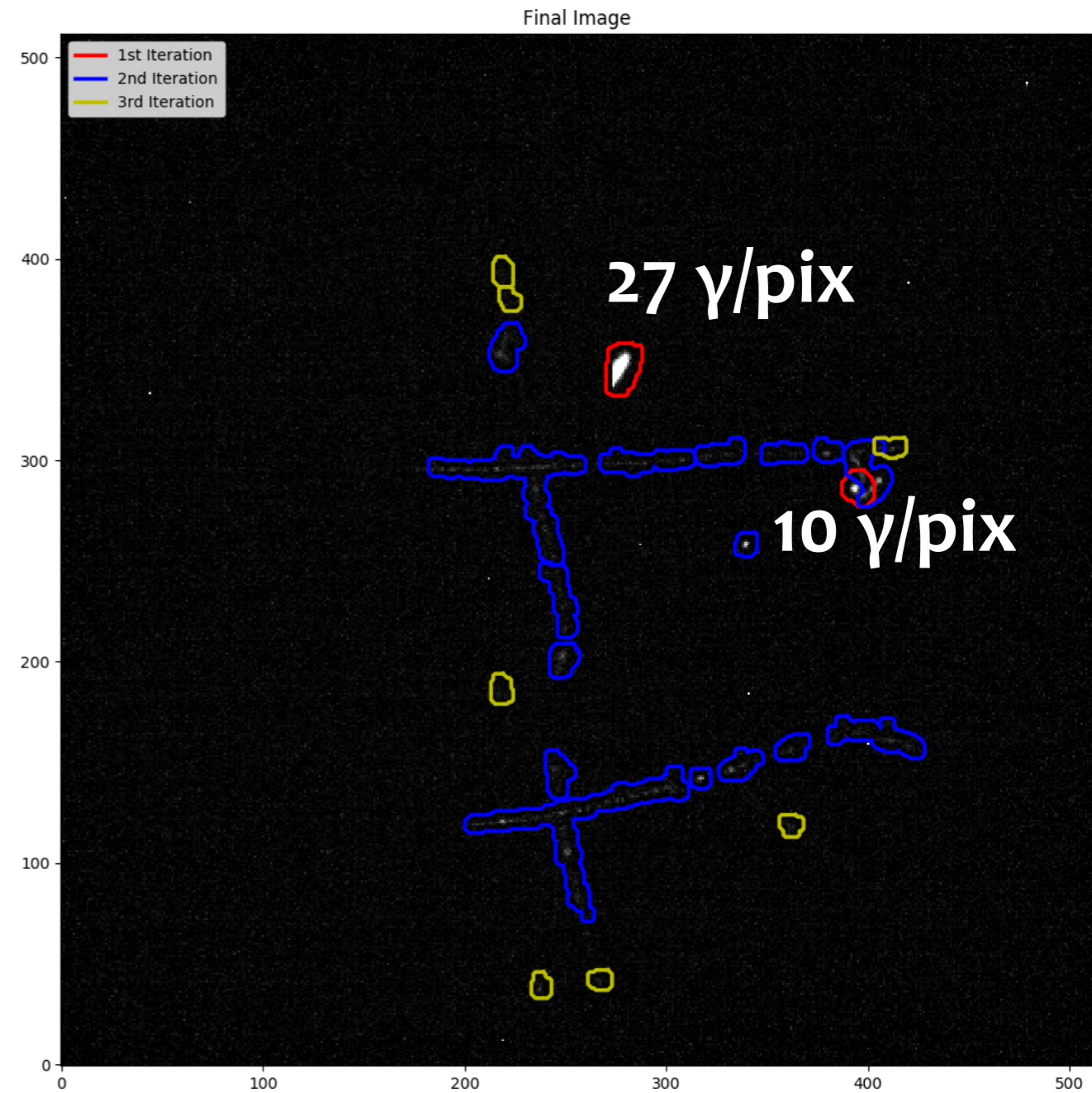
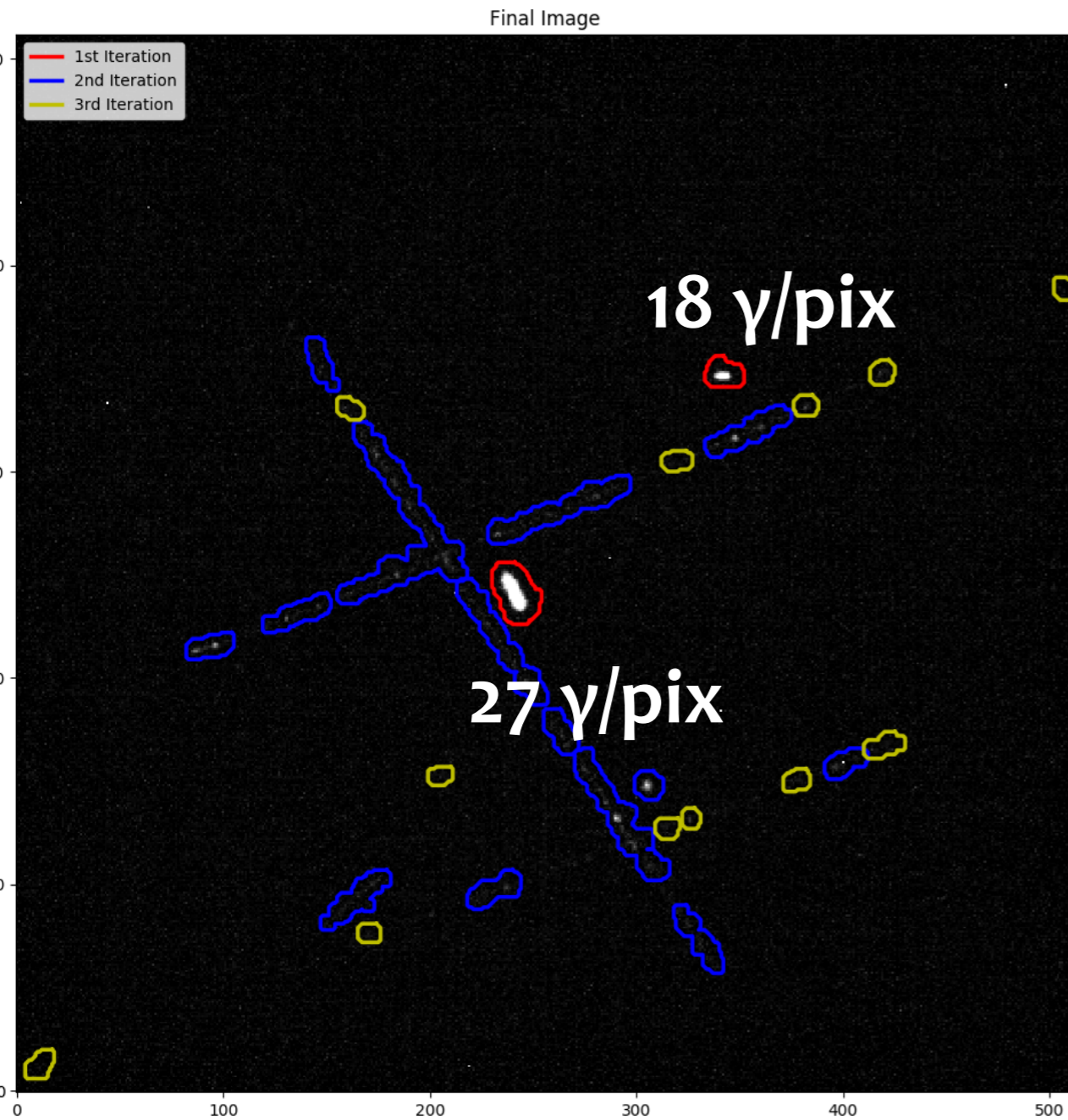
- Length, width, slimness, etc. are all very similar (should be solved as said by “superclustering” the OR of it1 and it2)
- apart the density: #photons/pixels in the cluster: the bump > 15 is only in AmBe runs**



can try to select those
e.g. the two it1 clusters of the
previous image were 1 high/1low density

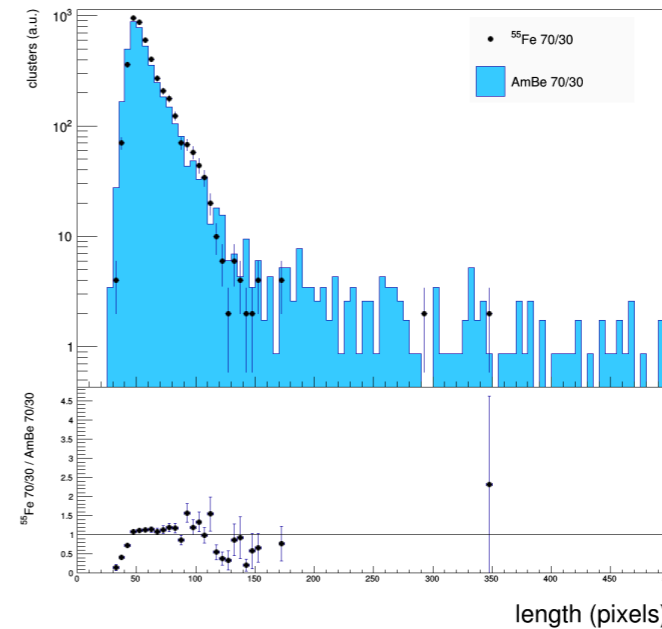
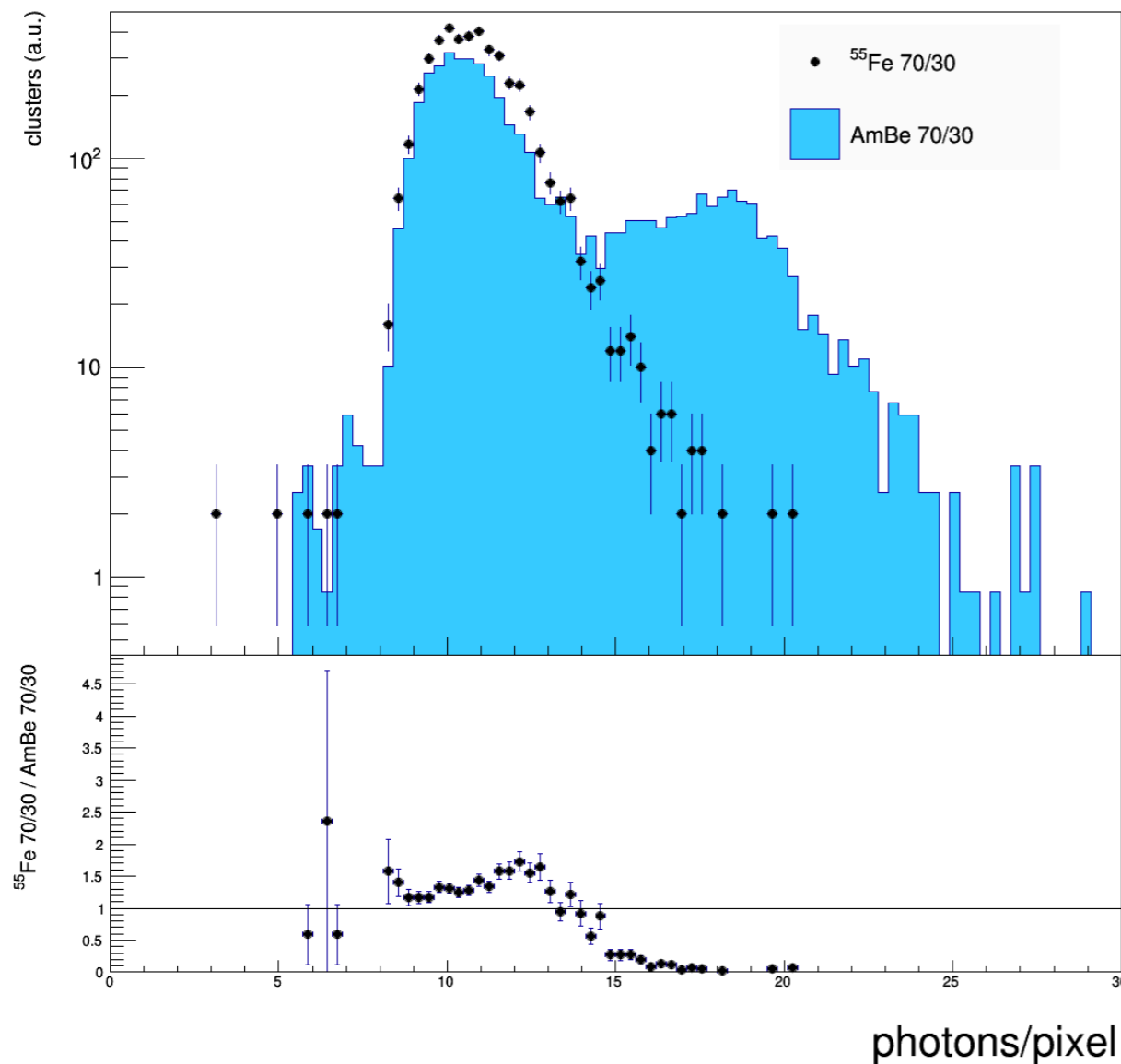


- looking at O(10) images, it seems that density > 15 selects good recoil candidates

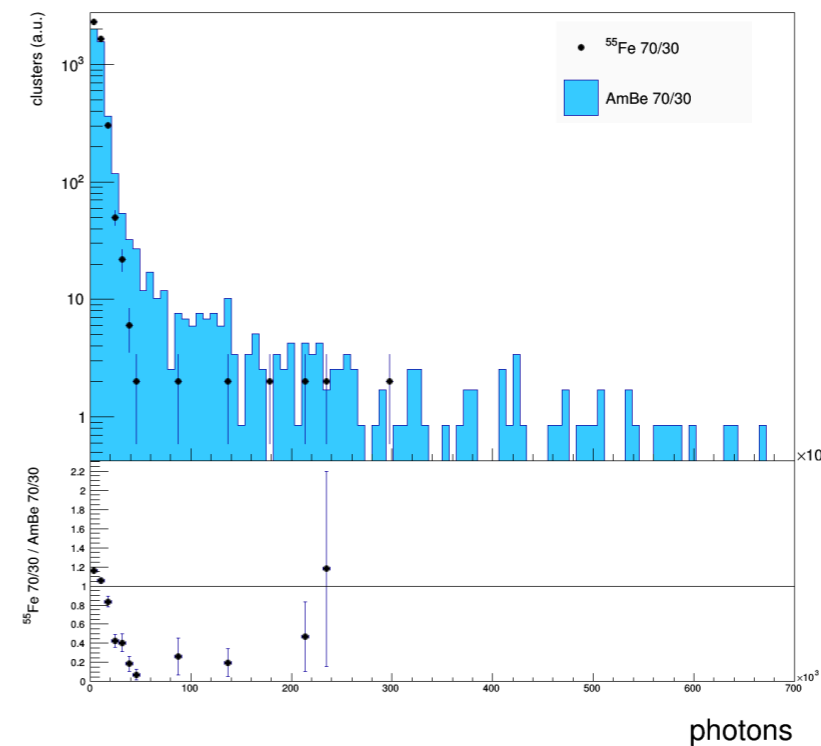


Very similar to α 's

- N.B. The previous plots are cutting in the center of the FC ellipse ($\sim 1/2$ of the area), to remove the protons coming from the interactions with the FC.
- Looking at the full volume the number of high-density clusters increases, as it should, because very long α tracks

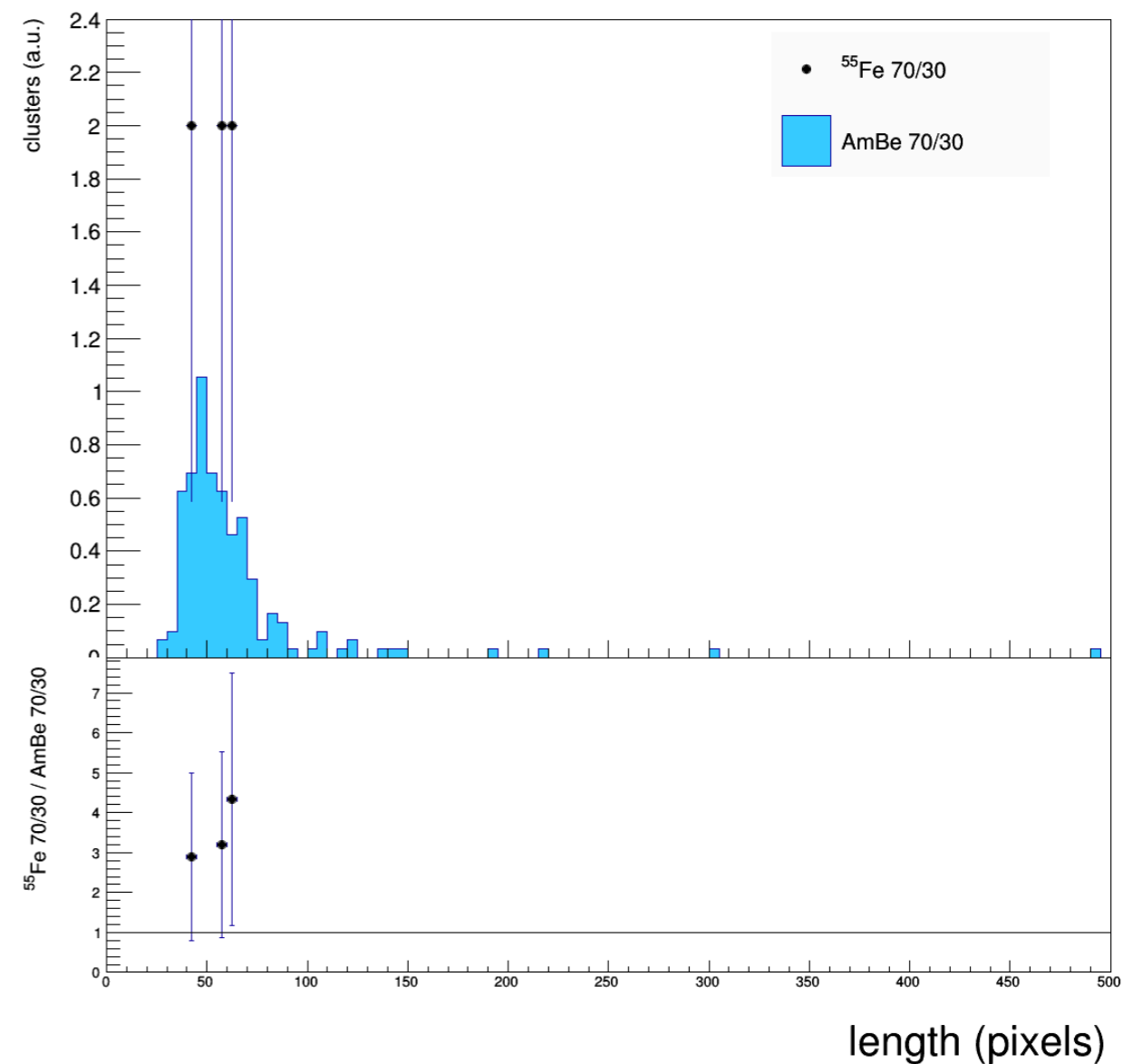
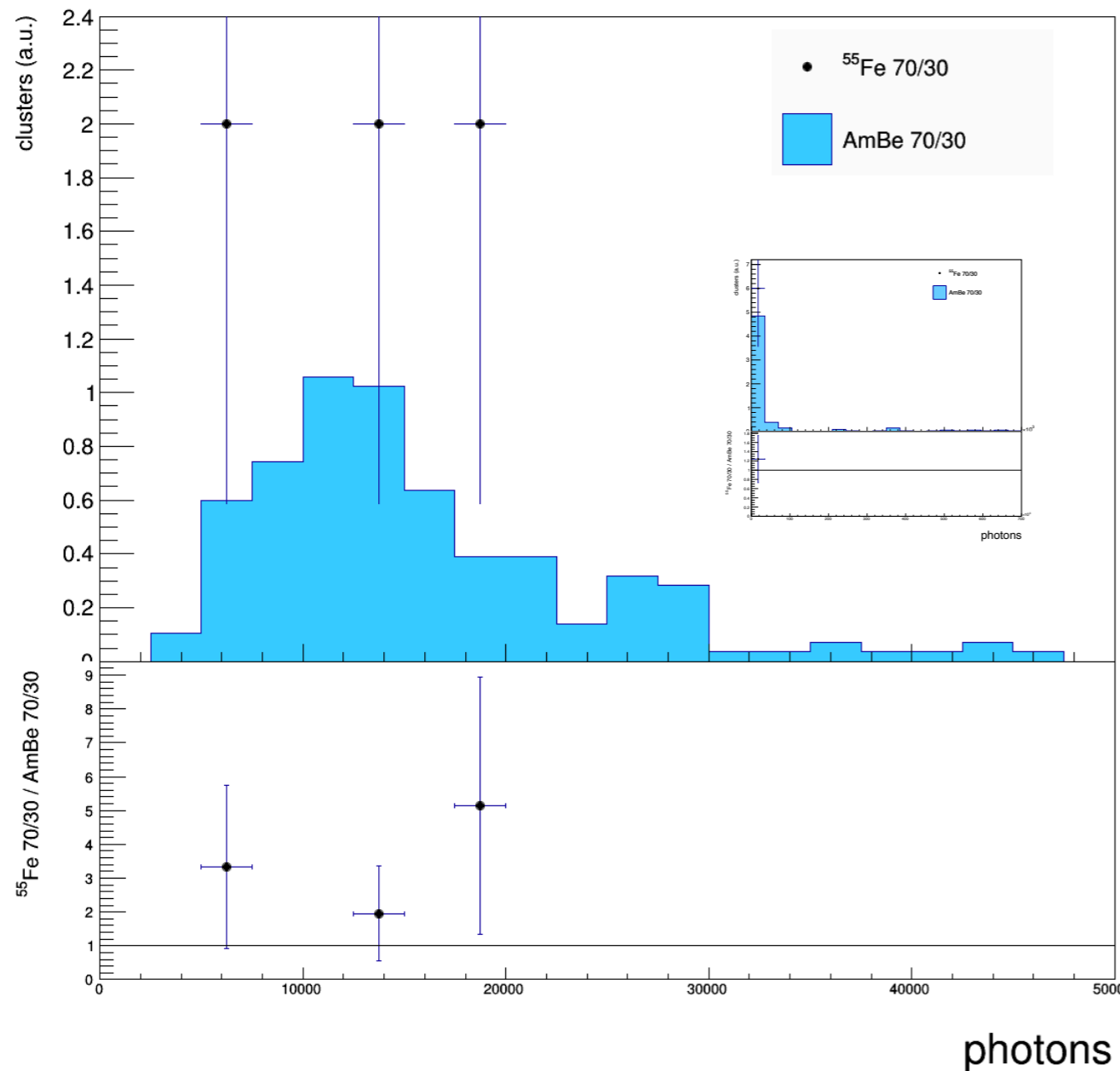


up to 10 cm long tracks



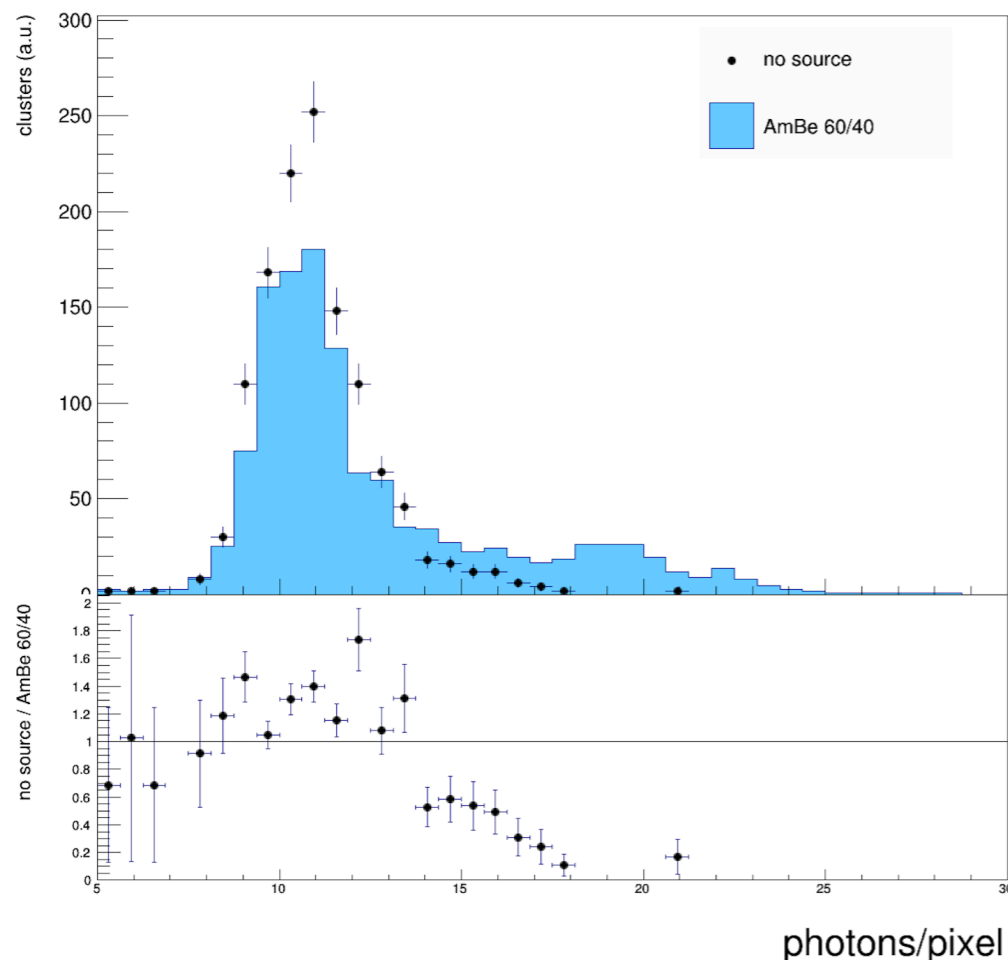
up to 1.5 MeV

- I.e. iter1 superclusters with density >15 γ /pixel inside the center of the FC
- what remains are not many candidates, with mode energy ~ 20 keV, but with a tail up to 1.5 MeV
- length up to ~ 1 cm (average 0.5 cm). ^{55}Fe “spots” are ~ 2.5 mm wide



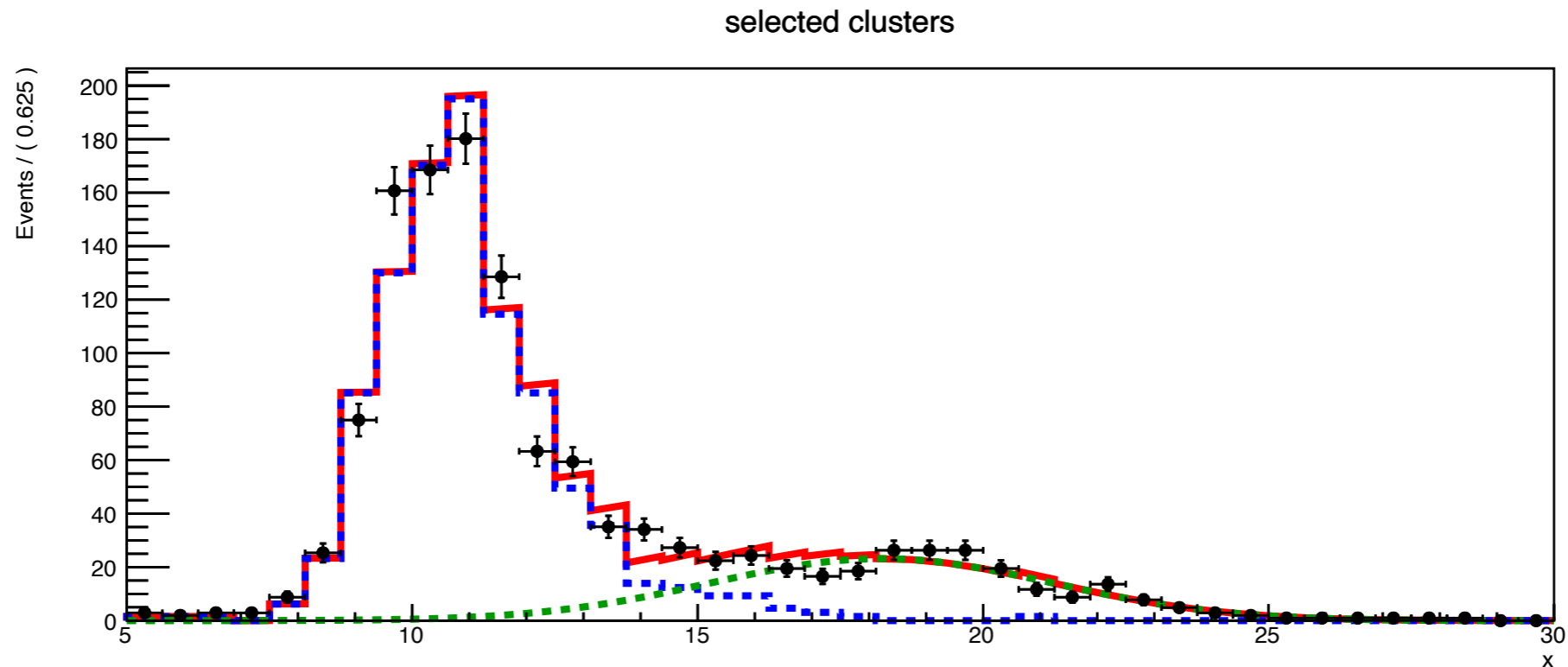
Signal yield and efficiency

- Choose the “density” as sig/bkg separation variable
- Need to subtract the background statistically to get pure signal yield and shape
 - from the signal shape one can get an efficiency for a density $> XX$ cut
- Make a simple likelihood fit
 - a run with no source (#2109 with 60/40 mixture) makes the PDF for the background
 - the “bump” around 19 ph/pix is modeled with a Gaussian (free mean/sigma)



The no-source data seems to model well the peak around 10 ph/pixel from pieces of cosmic-induced bkg

- 2 components fit
 - bkg from the binned template from cosmic data
 - signal as a completely floating Gaussian (not a rock-solid choice)

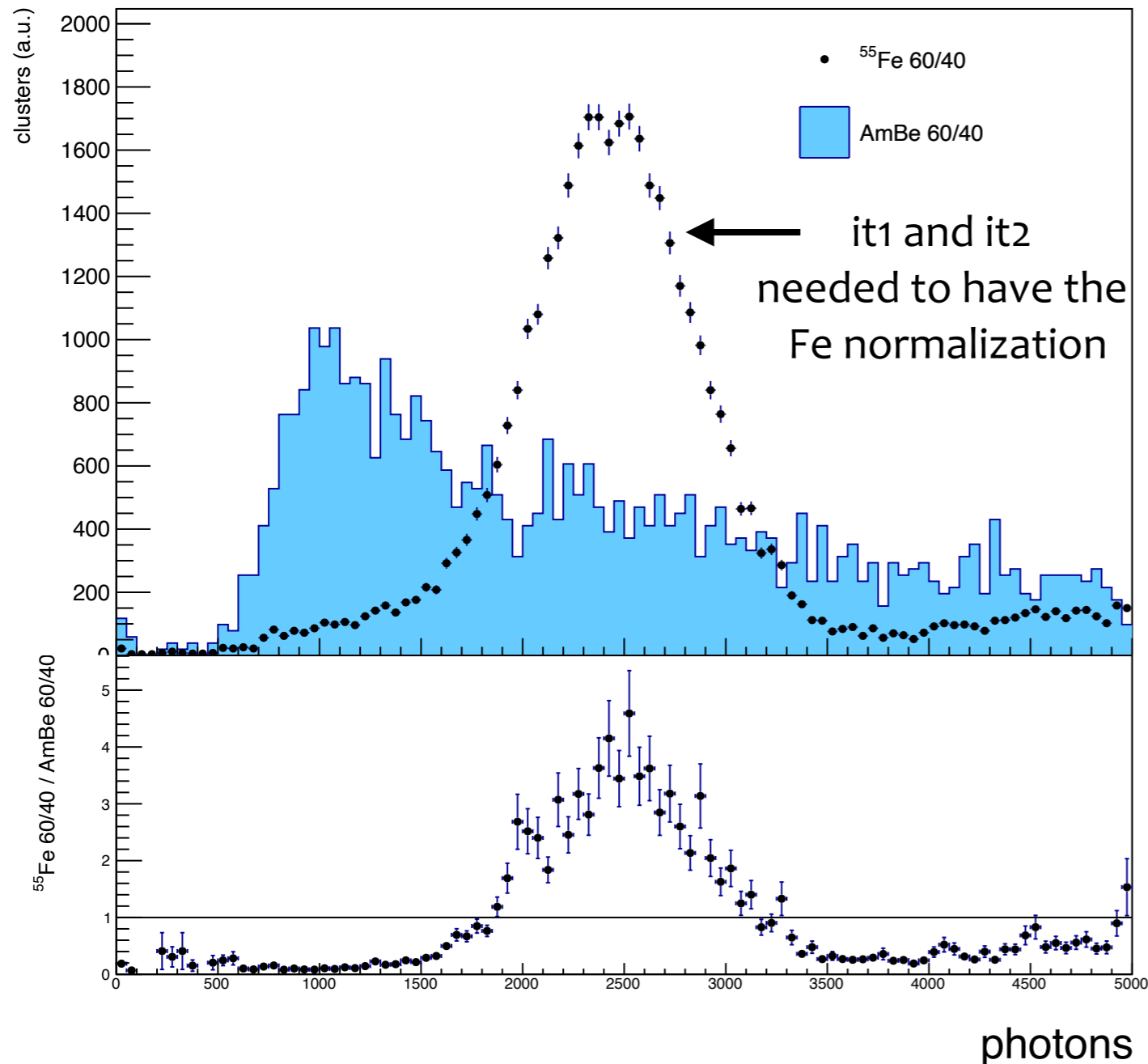


| threshold | Signal efficiency | Bkg efficiency |
|-----------|-------------------|----------------|
| >16 | 77% | 1.5% |
| >17 | 66% | 5.8e-3 |
| >18 | 53% | 1.9e-3 |
| >19 | 40% | 1.6e-3 |

signal/bkg efficiency
obtained
integrating the PDFs

Get bkg rejection for Fe at 6 keV

- not valid for recoils if there is saturation (because energy is shifted in non-linear way)



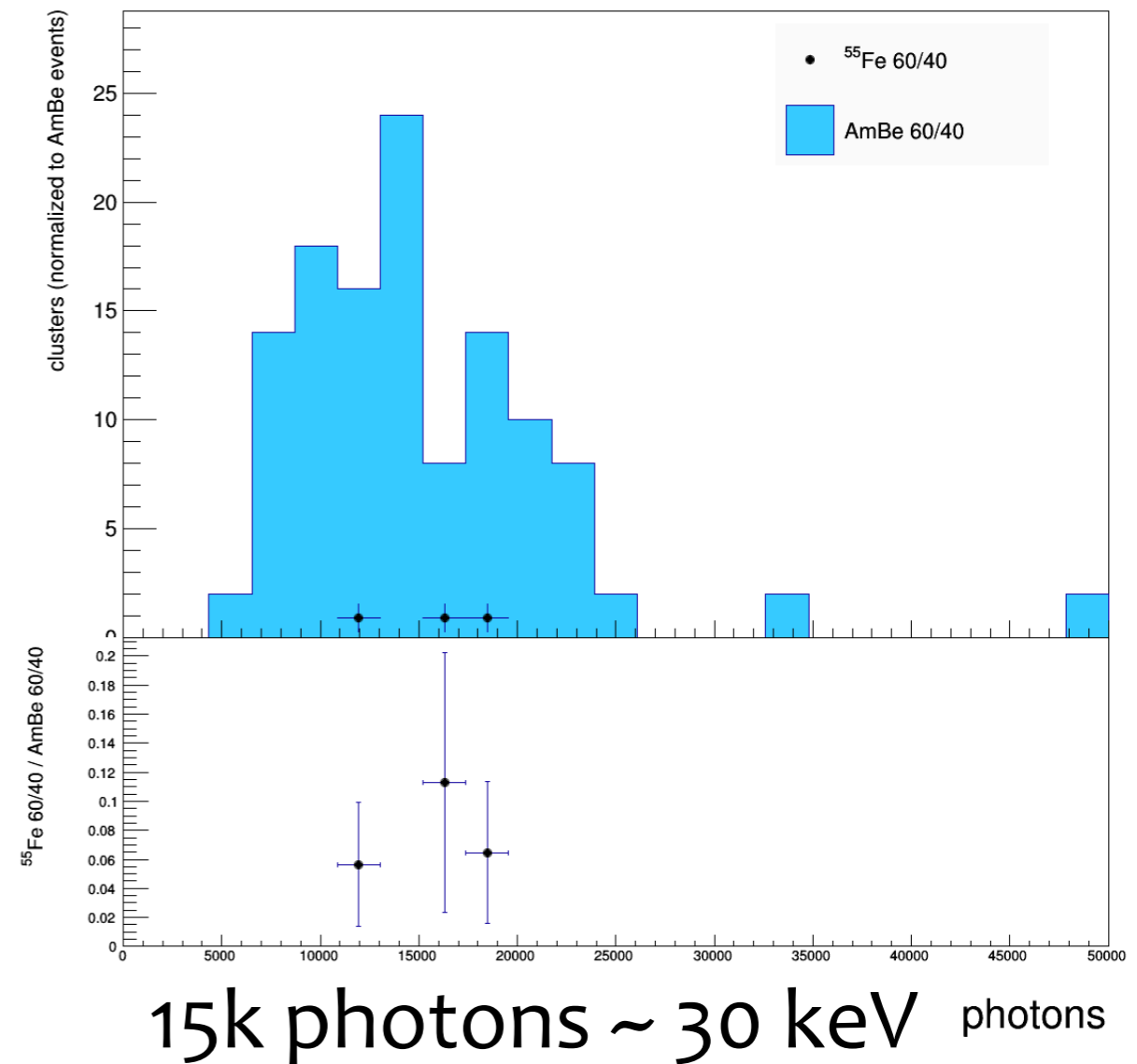
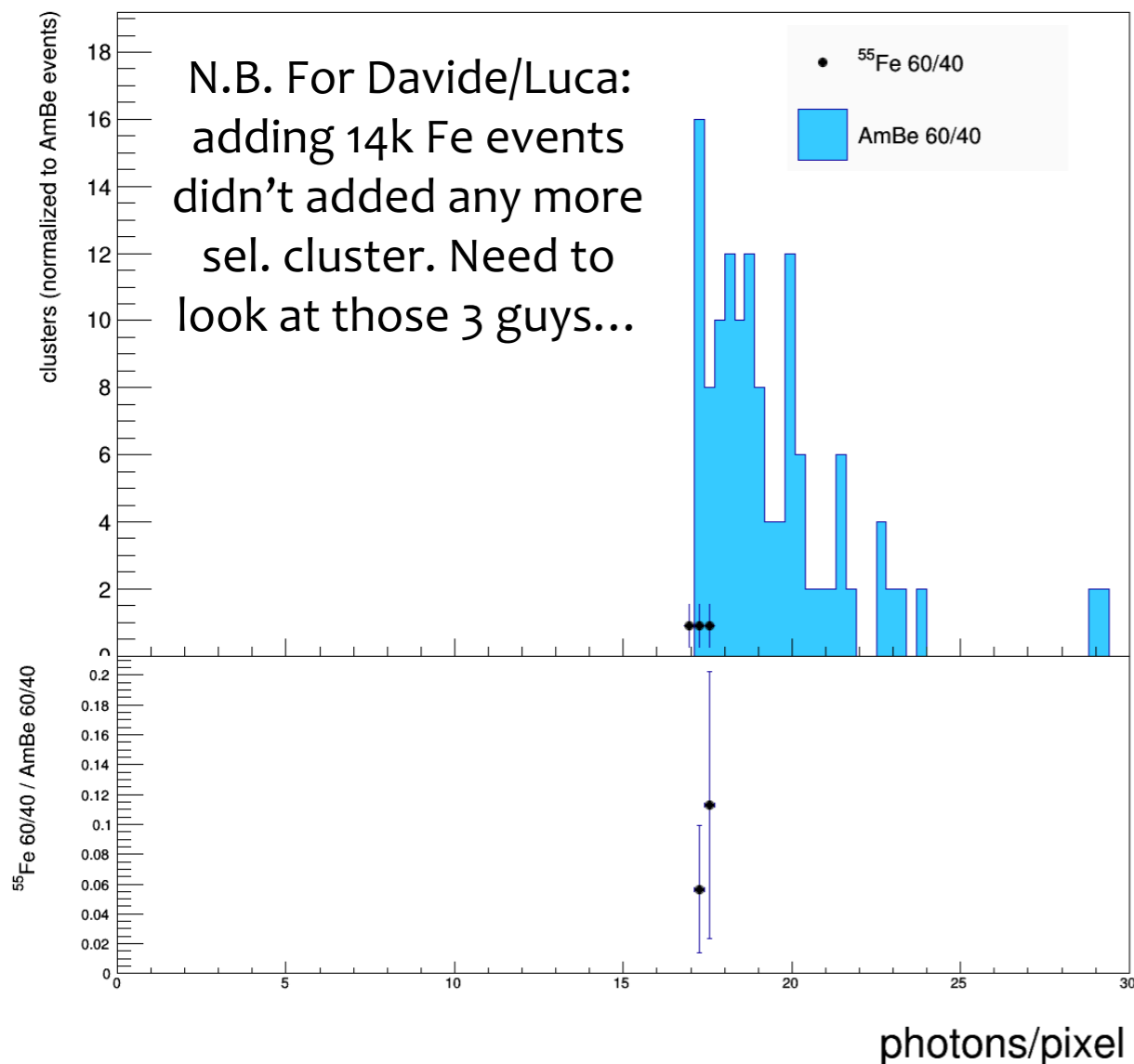
~3.5 Fe clusters/image
within the strict fiducial
region (reconstructed with
iter1 OR iter2 to get ~100%
efficiency on Fe)

$$\text{bkg } \varepsilon = \frac{n(^{55}\text{Fe})^{sel}}{n(^{55}\text{Fe})^{fiducial}}$$

$n(^{55}\text{Fe})^{sel}$ is the number of
clusters selected with a cut
on density > threshold

background rejection @ 6 keV

- Reconstructed 20k Fe images and 2900 AmBe images (...reconstructing more Fe...)
- applied the above selection with “density” cut > 17 photons/pixel.
 - 3 clusters surviving close to the boundary
 - efficiency = $3 \text{ clu} / (2e4 \text{ ev} * 3.5 \text{ clu/ev}) = 4.3e-5$
 - with a stat uncertainty of $\sim 3.8e-3$ (reconstruction of ALL Fe runs ongoing)

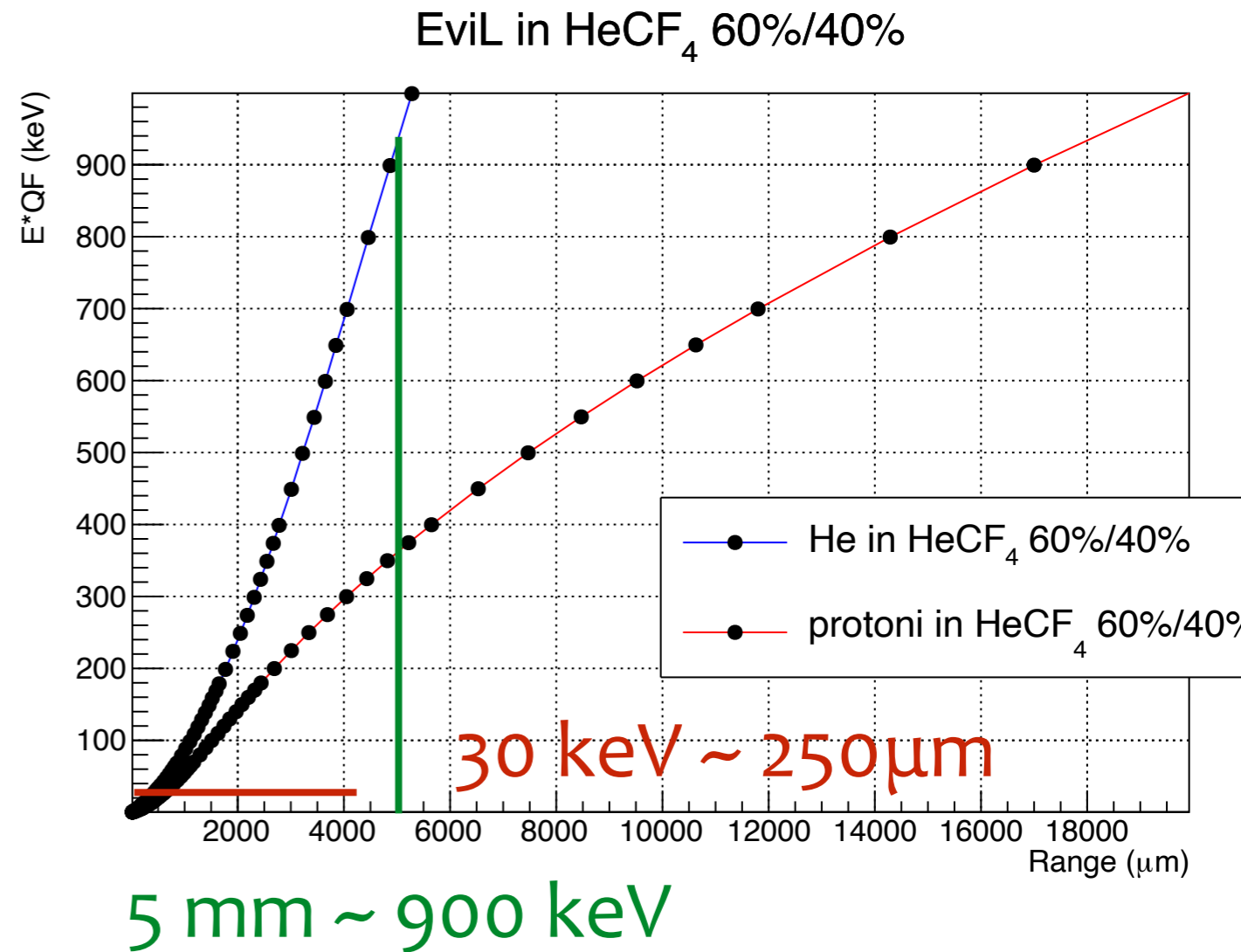
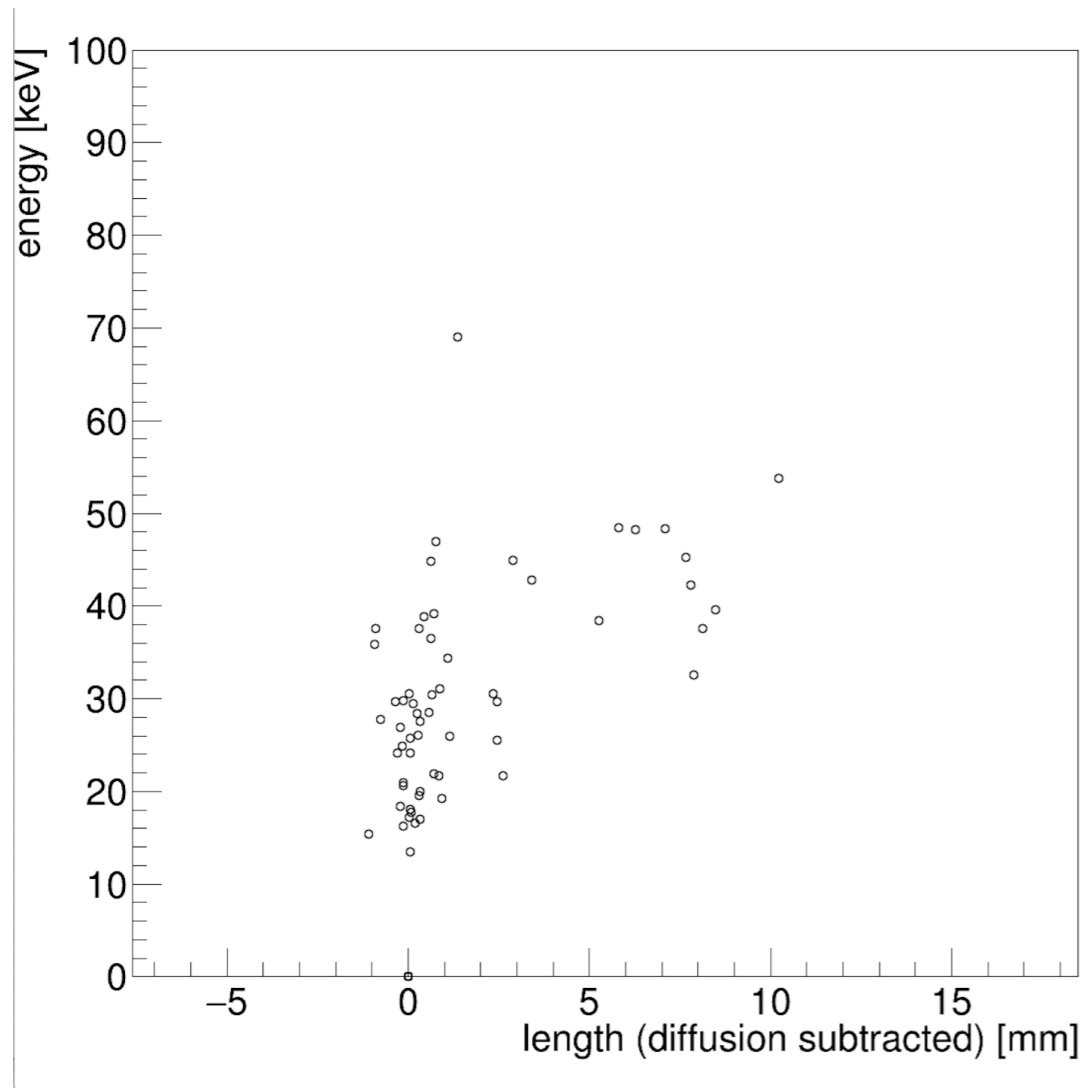


background rejection @ 6 keV

- Need to evaluate the total efficiency on signal (reconstruction, rest of the selection), not only the one from the density cut.
 - But presumably this is very high, looking at the images
 - so let's assume that $\text{Eff}(\text{Sig}) \sim \text{Eff}(\text{density cut})$

| threshold | Signal efficiency | Bkg efficiency of density cut | Bkg rejection |
|-----------|-------------------|-------------------------------|---------------|
| >16 | 77% | 1.5% | 1.1e-4 |
| >17 | 66% | 5.8e-3 | 4.3e-5 |
| >18 | 53% | 1.9e-3 | 1.4e-5 |
| >19 | 40% | 1.6e-3 | 1.2e-5 |

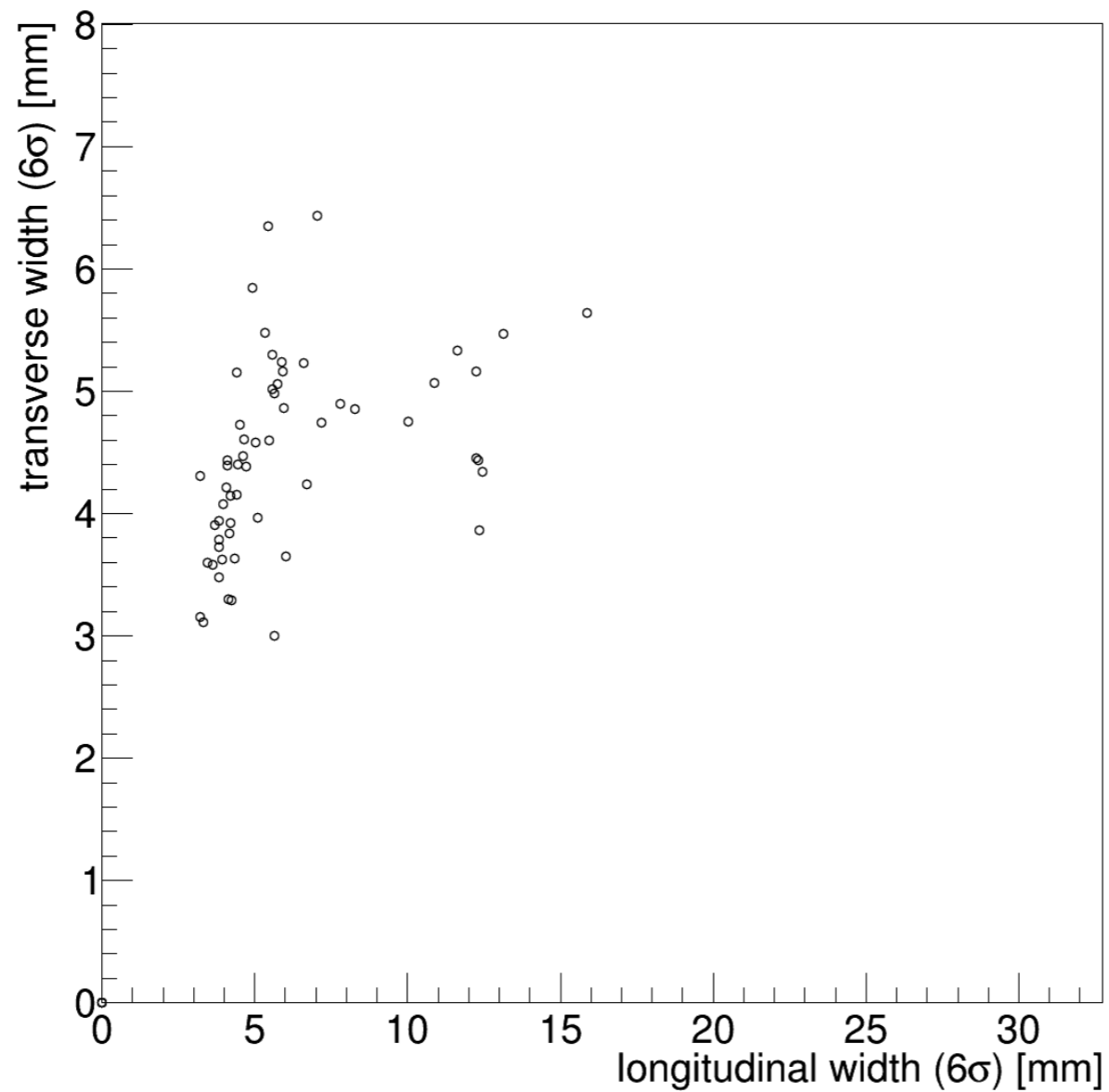
Length “diffusion subtracted” = length - width (cluster-by-cluster)
since in the transverse direction it should be a delta, while it is $\sim \langle 3 \text{ mm} \rangle$



5 mm tracks from SIM should correspond to $\sim 1 \text{ MeV}$
energy calibration @ 6keV is robust

may GEM saturation give x10 Energy underestimate?

- transverse size of a track is 3 to 6 mm. This needs to be subtracted to the width...



The End