

# INTENSE ANNUAL GENERAL MEETING AT PISA

## Muography of Vesuvius

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on behalf of the MURAVES Team, Italy

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# The Muography Team

Involved in the MURAVES experiment on Vesuvius

## People

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## Institutes

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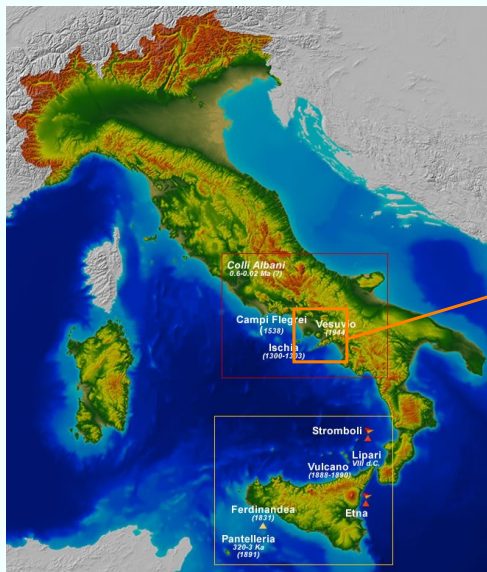
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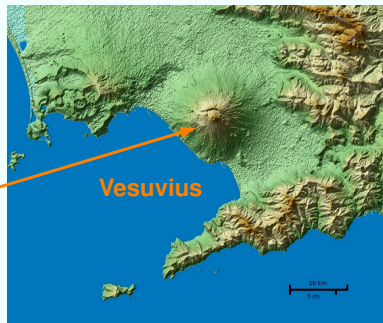
<sup>7</sup> Centro Siciliano di Fisica Nucleare e Struttura della Materia, Catania, Italy

<sup>8</sup> Earthquake Research Institute, The University of Tokyo

# The structure of Vesuvius



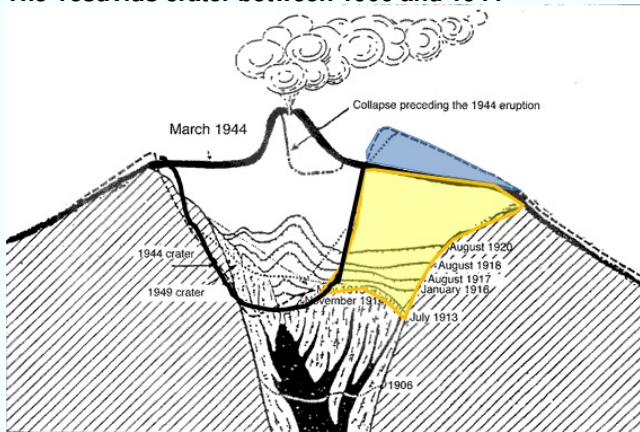
**Vesuvius is an active volcano**



- Last eruption occurred in 1944
- Now the conduit is closed
- Monitored 24/7 by INGV

# Observations of the upper part of Vesuvius

## The Vesuvius crater between 1906 and 1944



(from Imbò, 1949)

## Vesuvius crater in 1944

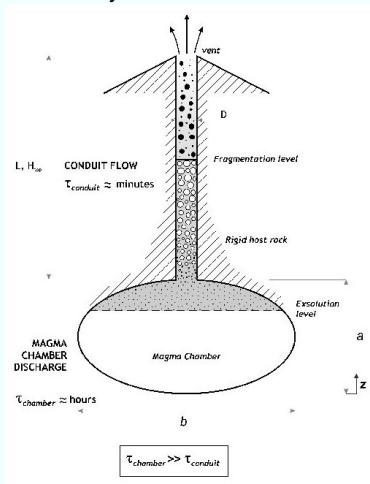


## Vesuvius crater today



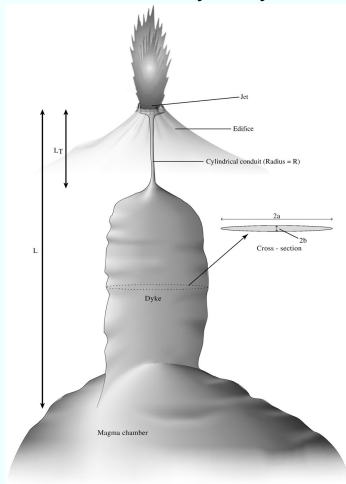
# Magma chamber-conduit model

## Cylindrical conduit



(Macedonio et al., 2005)

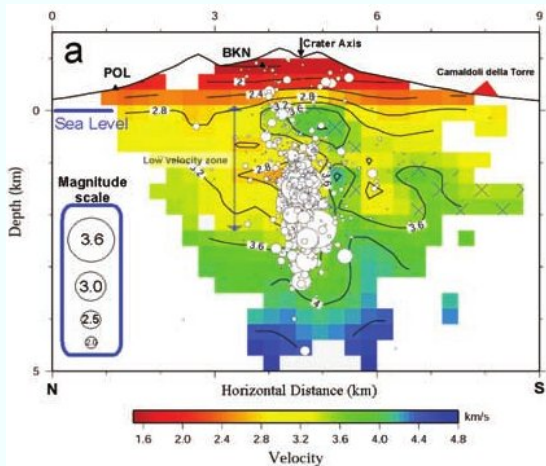
## More realistic: dyke+cylinder



(Costa et al., 2009)

# The internal structure of Vesuvius

High resolution **passive** seismic tomography (300-500 m)



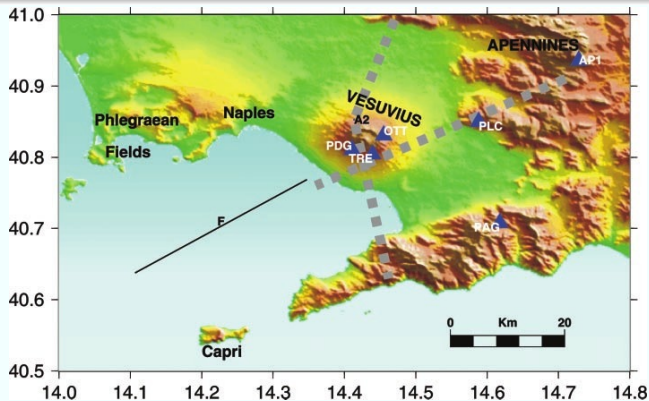
(modified after Scarpa et al., 2002)

## P-wave velocity structure

- Based on simultaneous inversion of the 3-D velocity structure and the earthquake location
- Data from 2139 earthquakes recorded from minimum 7, up to 19 stations (8600 P-waves, 1900 S-wave readings)
- Magmatic bodies inside the investigated volume were not found

# Active seismic tomography of Vesuvius

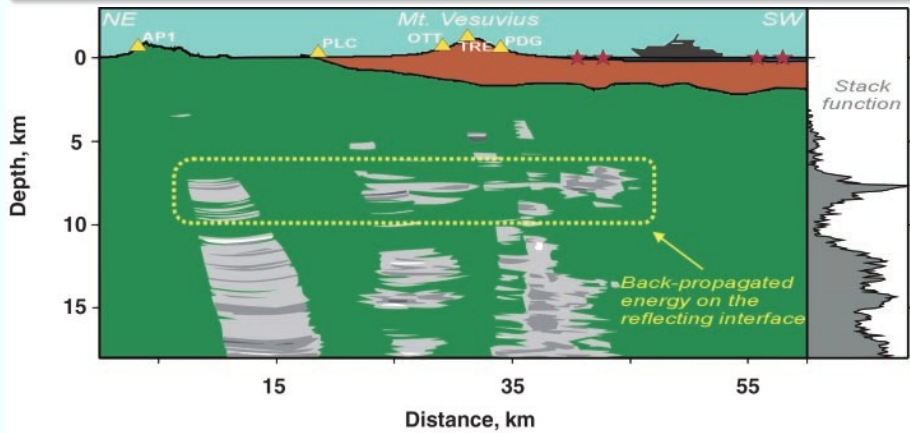
- Previous experiments Tomoves-1994, Tomoves-1996
- Mareves-1997: 1800 shots off-shore (along line F)
- 25 temporary seismic stations + permanent stations of the Osservatorio Vesuviano network



(from Auger et al., 2001)

# The MAREVES-1997 active tomography experiment

- Found the top of a magmatic body at 8 km depth, at least 400 km<sup>2</sup>
- Thickness of the magmatic body not constrained
- $\approx$  500-1000 m spatial resolution

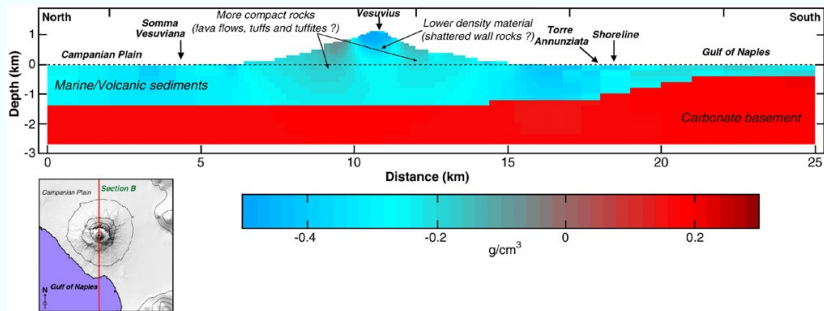


(from Auger et al., 2001)



# Gravimetric field inversion

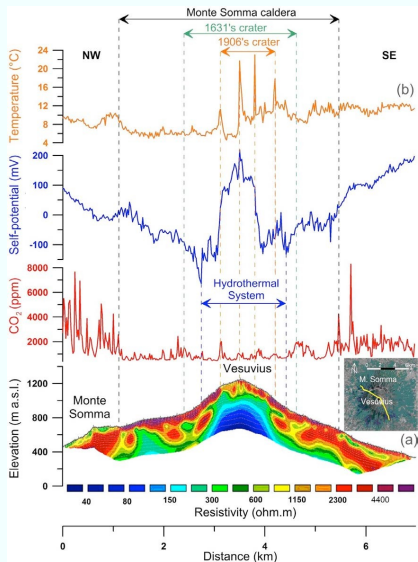
- About 300 m spatial resolution
- Model misfit  $\approx 5\%$



(from Cella et al., JVGR, 2007)

# The internal structure of Vesuvius

## Electric resistivity



- Total profile length 7 km
- Spacing between electrodes 40 cm
- Reach 500 m depth

(Finizola et al., 2014)

# First muography experiment at Vesuvius (2009-2010)

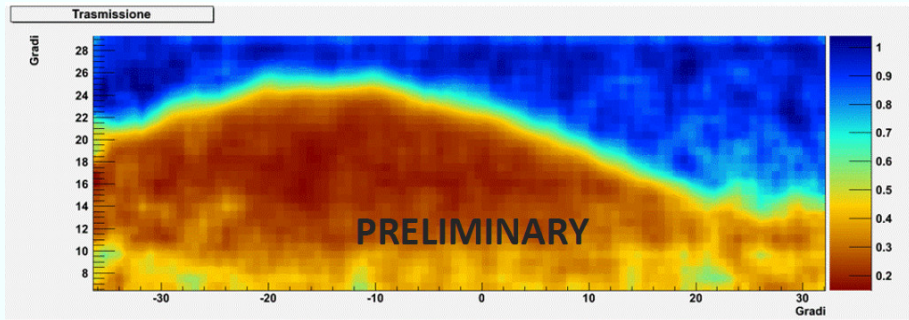
## Mounting the Japanese detector



# The MU-RAY prototype (2013)



# MU-RAY First results (April 2013)



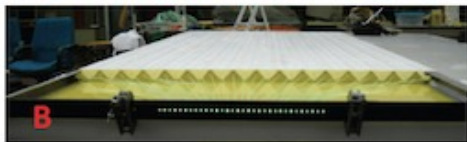
Vesuvius seen from  
the installation point



# The MURAVES telescope

INFN, INGV, Universities of Napoli and Florence

- Higher resolution, better background suppression
- The detector is based on **plastic scintillators** / SiPM
- Based on the **MU-RAY** technology
- At Vesuvius adopts **lead walls** for background suppression



# Selected site for the MURAVES detector

Inside the Vesuvius National Park

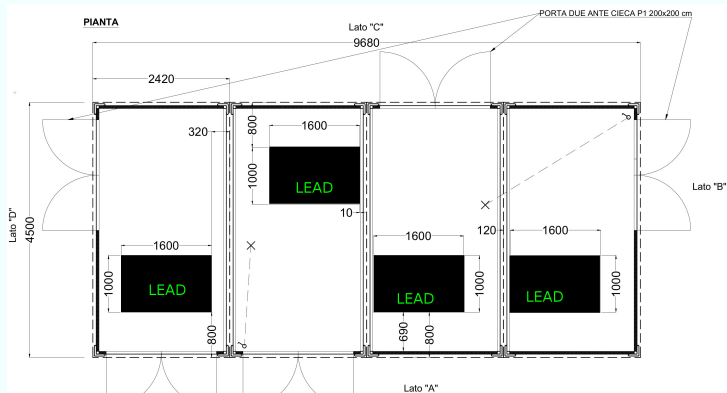


Position: E  $14^{\circ} 24' 41.76''$ ; N  $40^{\circ} 48' 36.75''$  (675 m a.s.l.)

# The Laboratory on Vesuvius



- Four aligned container modules
- Total internal: L=9.3 m, W=4.2 m, H=2.7 m
- Five doors
- Concrete basement (pre-built)
- Can host 3 detectors + 1 calibration





# Container assembly



# Lead from the OPERA experiment

Re-melt in bricks (Total: 35 tons)



# Power supply and data transmission

- Scintillators mounted in January-February 2019
- Acquisition tests already started for detector n.1
- Three years data acquisition

**Solar panels mounted on the roof**



## Solar panels

- 18 x 300 W solar panels
- Total power: 5400 W (peak)
- 32 lead batteries, Tot= 15.7 kWh
- Output voltage: 48 V

## Internet connection

- WiFi radio-bridge

# Conclusions

## MURAVES Project

- MURAVES project started in 2015 (INGV + INFN)
- Collaboration with UC-Louvain (Belgium) in progress
- End of laboratory setup on Vesuvius: January 2019
- Start of data acquisition (preliminary tests): February 2019
- Current status (November 2019): data acquisition

# Bibliography

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