# Weyl-Invariant Gravity and the Nature of Dark Matter

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# **Basic Assumptions**

- Weyl-invariant (WI) version of general relativity (GR). WI: freedom to locally rescale our fundamental units
- The standard model (SM) of particle physics
- A distinction between active & passive gravitational masses
- G & M<sub>ac</sub> weakly vary over galactic and supergalactic scales

#### Weyl-Transformation

$$g_{\mu\nu} \to \Omega^2(x)g_{\mu\nu}$$
  
 $\mathcal{L}_M \to \Omega^{-4}(x)\mathcal{L}_M$   
 $\phi \to \Omega^{-1}(x)\phi$ 

Note that  $\Omega(x)$  is an arbitrary function

Can be used in the WI extension of GR to generate new solutions.

# Weyl-Invariant Scalar-Tensor (WIST) Theory

$$\mathcal{I}_{EH} = \int \left[ R/(16\pi G) + \mathcal{L}_M \right] \sqrt{-g} d^4 x = \int \left( \frac{1}{2} m_P^2 R + \mathcal{L}_M \right) \sqrt{-g} d^4 x$$
$$\mathcal{I}_{WIST} = \int \left( \frac{1}{6} |\phi|^2 R + \phi_{,\mu} \phi^{*,\mu} + \mathcal{L}_M(|\phi|, \{\psi\}) \right) \sqrt{-g} d^4 x$$

Deser 1970

# **Field Equations**

$$\frac{\delta \mathcal{I}_{WIST}}{\delta g^{\mu\nu}} = 0 \Rightarrow \qquad \frac{|\phi|^2}{3} G_{\mu\nu} = T_{M,\mu\nu} + \Theta_{\mu\nu}$$

$$3\Theta_{\mu\nu} \equiv \phi^*_{\mu;\nu}\phi - 2\phi^*_{\mu}\phi_{\nu} - g_{\mu\nu}(\phi^*\Box\phi - \frac{1}{2}\phi^*_{\alpha}\phi^{\alpha}) + c.c.$$

$$\phi^* \frac{\partial \mathcal{L}_M}{\partial \phi^*} + \phi \frac{\partial \mathcal{L}_M}{\partial \phi} = T_M$$

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Linearly perturbed metric assuming baryons only (G is a constant):

 $ds^{2} = -(1 - 2\Phi_{b})dt^{2} - dtdx^{i}v_{i} + (1 + 2\Phi_{b})dx_{i}dx^{i}$ 

Adding DM potential:

$$\Omega(x) = 1 + \Phi_{DM}(x)$$

$$ds^{2} \to \Omega^{2} ds^{2} = -[1 - 2(\Phi_{b} + \Phi_{DM})]dt^{2} - dt dx^{i}v_{i} + [1 + 2(\Phi_{b} + \Phi_{DM})]dx_{i}dx^{i}$$

# Summary

- DM phenomena on galactic and galaxy cluster scales could be explained by a WI version of GR, i.e. a `new' symmetry of gravitation
- This entails weak spatial (and possibly temporal as well) variation of  $GM_{ac}$  of  $O(\varphi_{DM})$  that act as an effective  $\rho_{DM}$

#### **Backup Material**

Machianity of Brans-Dicke Theory

$$(3+2\omega_{BD})\Box\Phi_{BD}=8\pi T_M$$

But there is a (reasonable) `cost':

$$GM_{ac} \propto m_P^{-2} M_{ac} \propto \phi^{-2} \phi = \phi^{-1} \propto 1 + \Phi_{DM}(x)$$

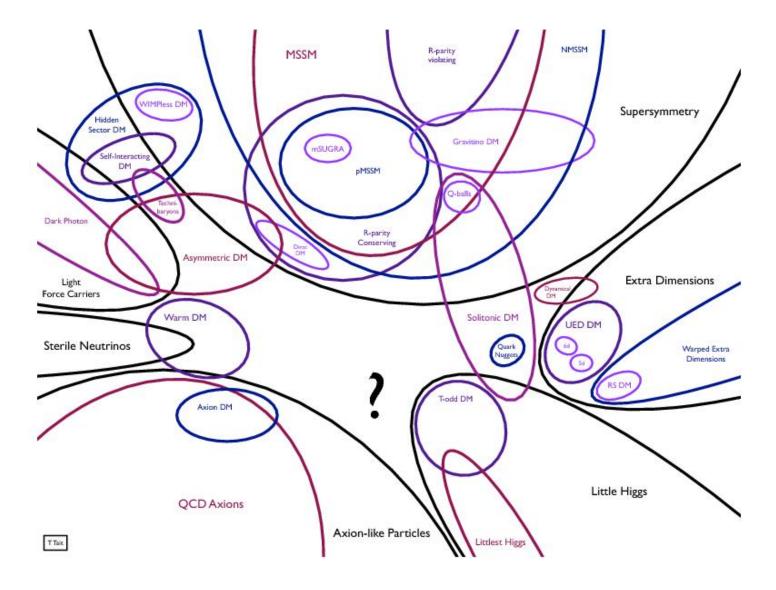
e.g 
$$\Phi_{DM} = \Phi_{NFW}(r) \propto r^{-1} \ln(1 + r/r_s)$$

Why is the gravitational potential linear whereas the overdensity is non-linear ?

$$\nabla^2 \Phi = 4\pi G \delta \rho = 4\pi G \rho \frac{\delta \rho}{\rho} = \frac{3H^2}{2} \frac{\delta \rho}{\rho}$$
$$\Rightarrow \qquad \left(\frac{l_H}{l_{gal}}\right)^2 \Phi = \frac{\delta \rho}{\rho}$$

# Slightly modified baryon density $\rho_b \to \rho_b/(1 + \Phi_{DM})^4 \approx \rho_b(1 - 4\Phi_{DM})$

#### Dark Matter Venn Diagram



#### Mannheim-Kazanas Metric

$$ds_{MK}^{2} = -(1 - \frac{\beta}{r} + \gamma r - \kappa r^{2})dt^{2} + \frac{dr^{2}}{1 - \frac{\beta}{r} + \gamma r - \kappa r^{2}} + r^{2}(d\theta^{2} + \sin^{2}\theta d\varphi^{2})$$

Mannheim & Kazanas, 1989 MS , arXiv:1702.08472 Li & Modesto, 2020