

MAKING THE INVISIBLE VISIBLE

Search for dark photons using a dielectric multilayer haloscope

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16th Patras Workshop on Axions, WIMPs and WISPs

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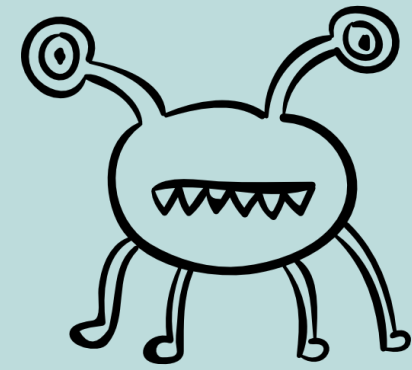
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INTRODUCTION



DARK MATTER

While the existence of dark matter is motivated by its gravitational coupling to the Standard Model, we (mostly) search for it on its presumed weak interaction.



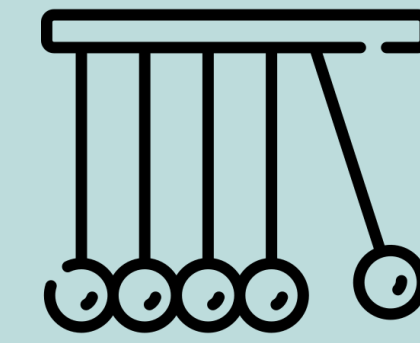
WIMPS

Weakly Interacting Massive Particles have been highly favourable candidates. Several experiments have been searching for them in last couple of decades.



DARK PHOTON

So far... nothing has been detected yet. Interest in a dark sector—dark because completely neutral under SM interactions—has increased over the past few years. The dark photon is a hypothetical new boson belonging to the dark sector.

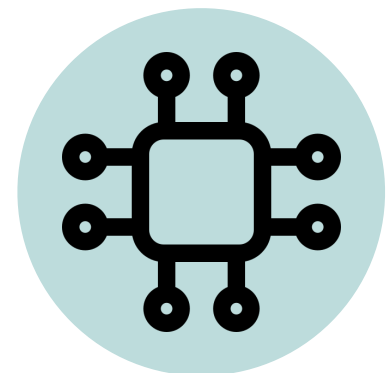
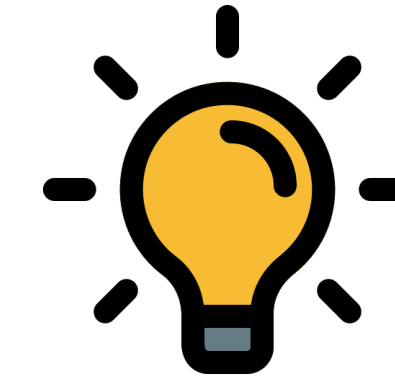
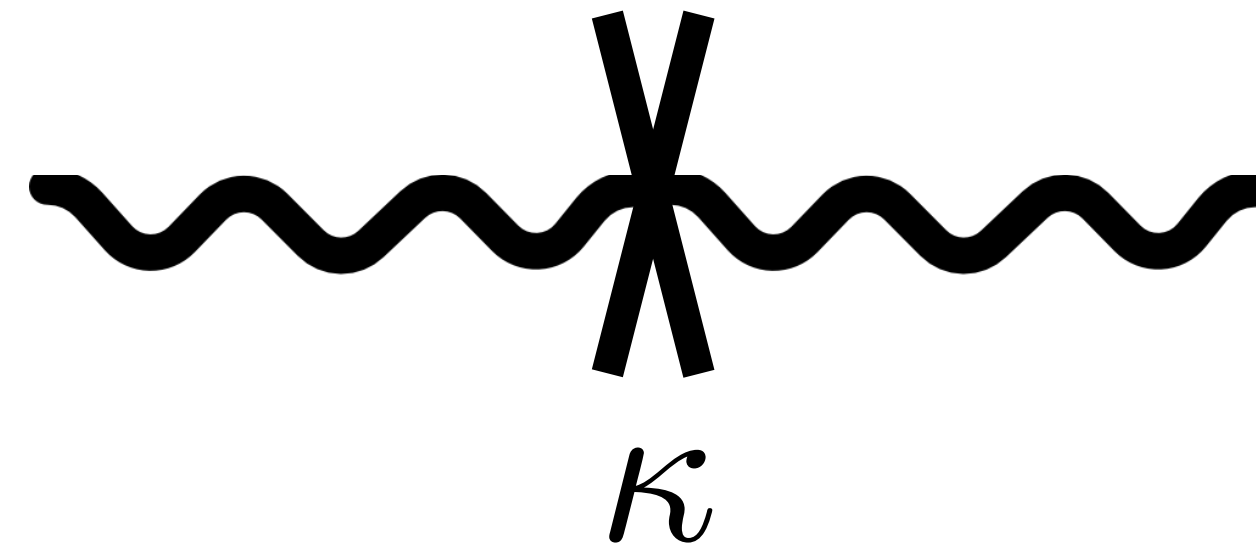
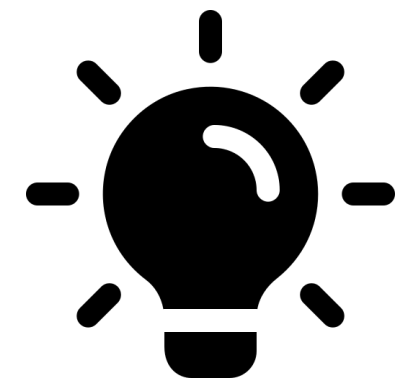


KINETIC MIXING

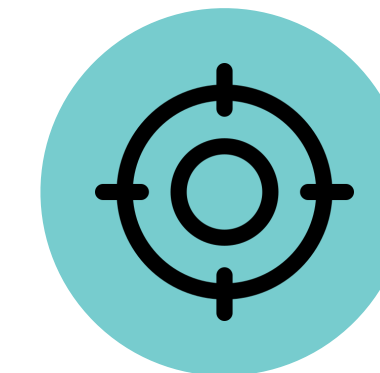
The dark photon can be detected because of its kinetic mixing with the SM photon (vector portal).

$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{4}(F'_{\mu\nu})^2 + \frac{\epsilon}{2}F'_{\mu\nu}F_{\mu\nu} + \frac{1}{2}m_{A'}^2(A'_\mu)^2$$

DIELECTRIC MULTILAYER HALOSCOPE



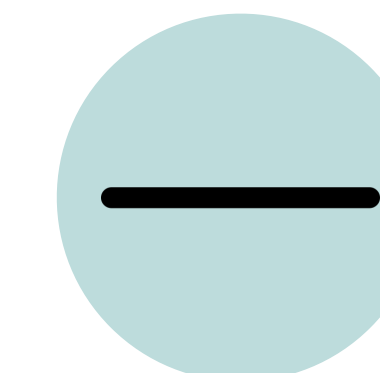
Single-photon sensor.



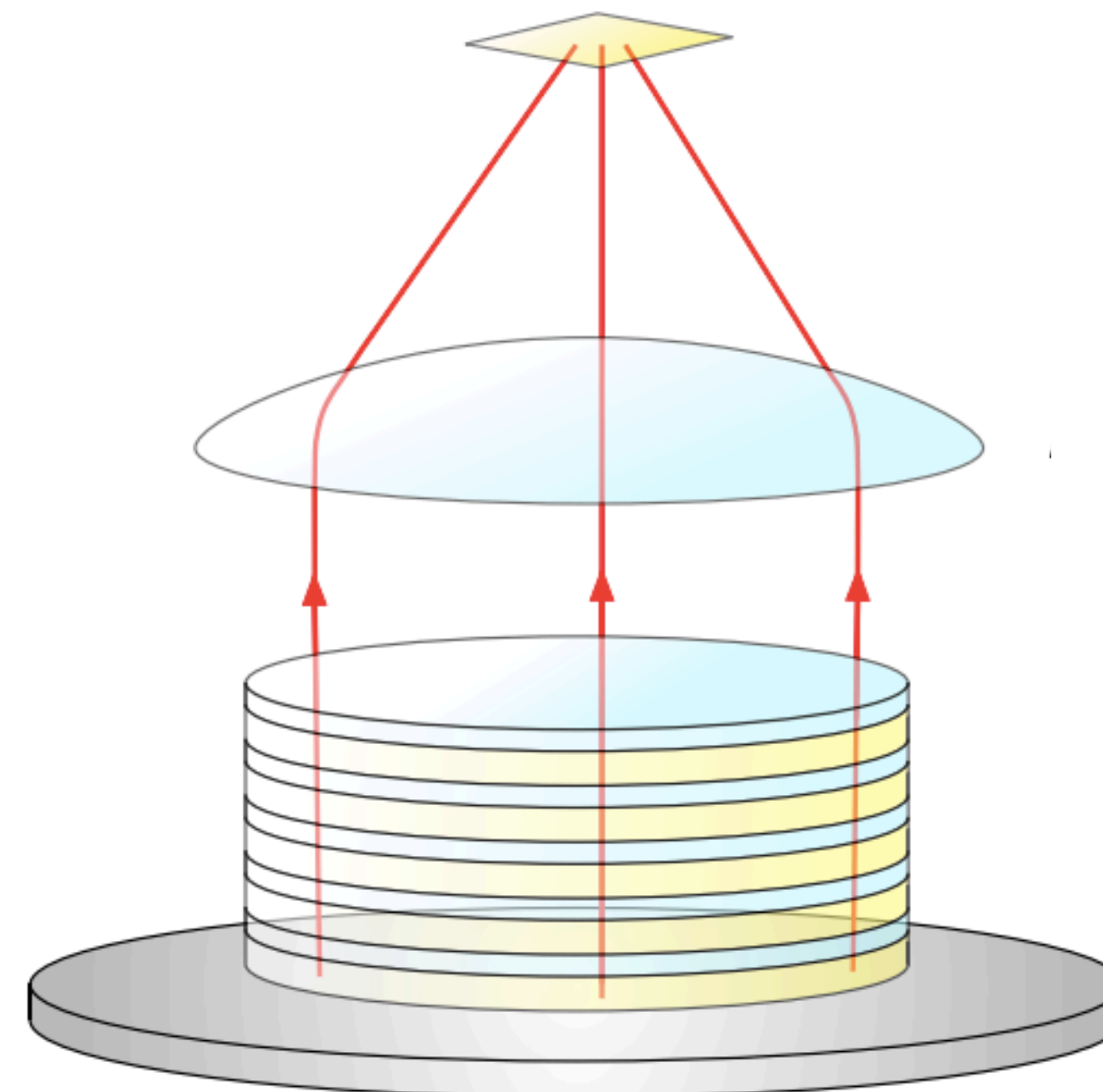
Lens for focusing the converted dark photons onto the photosensor.

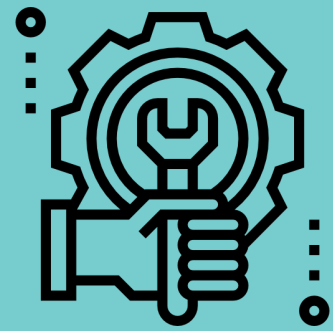


Half-wave stack of dielectric bilayers with indices of refraction $n_1 \gg n_2$ enable the conversion dark photon to SM photon.

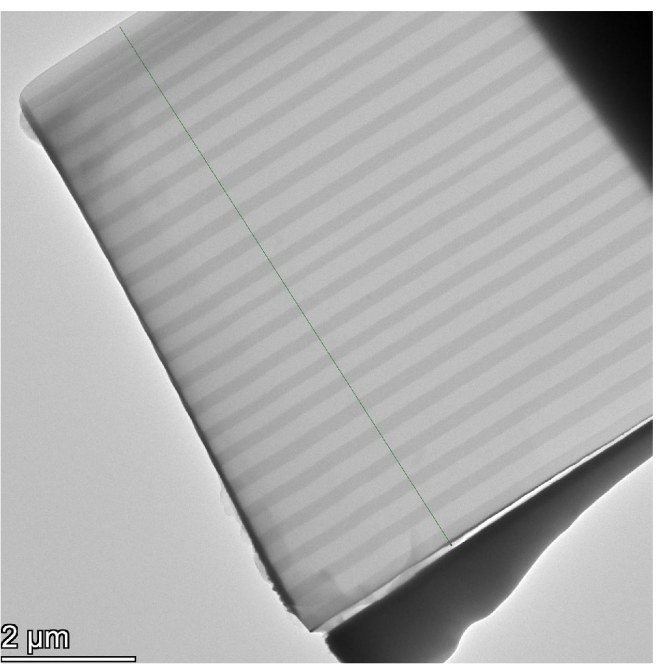
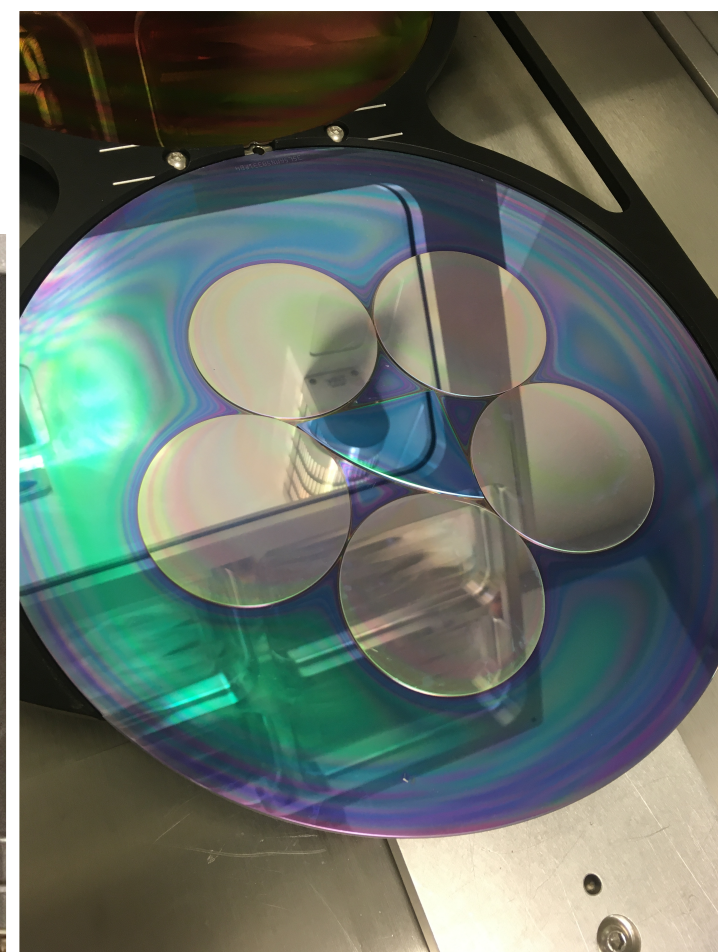
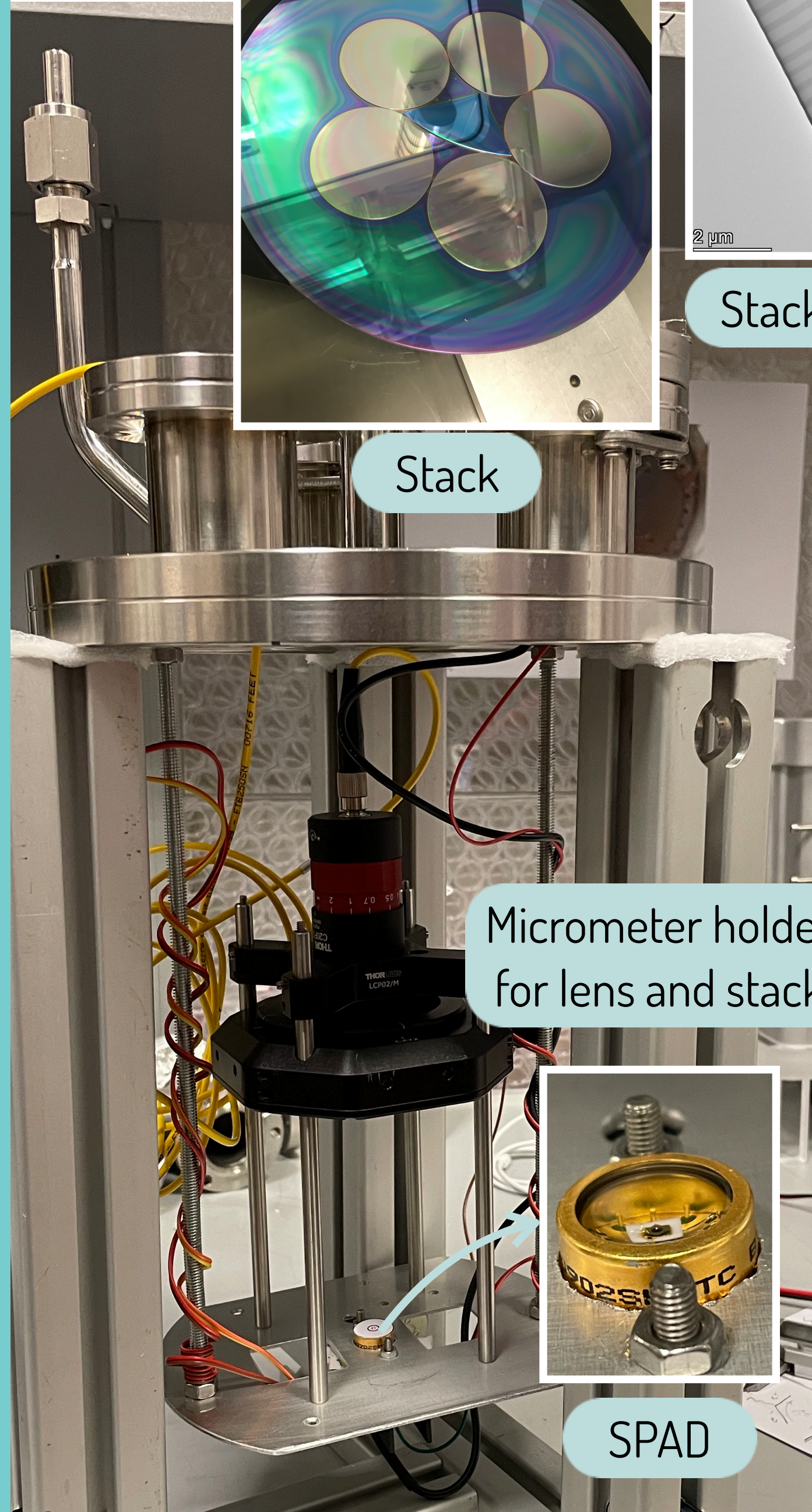


Mirrored substrate to enhance DP to SM photon conversion.





THE SETUP



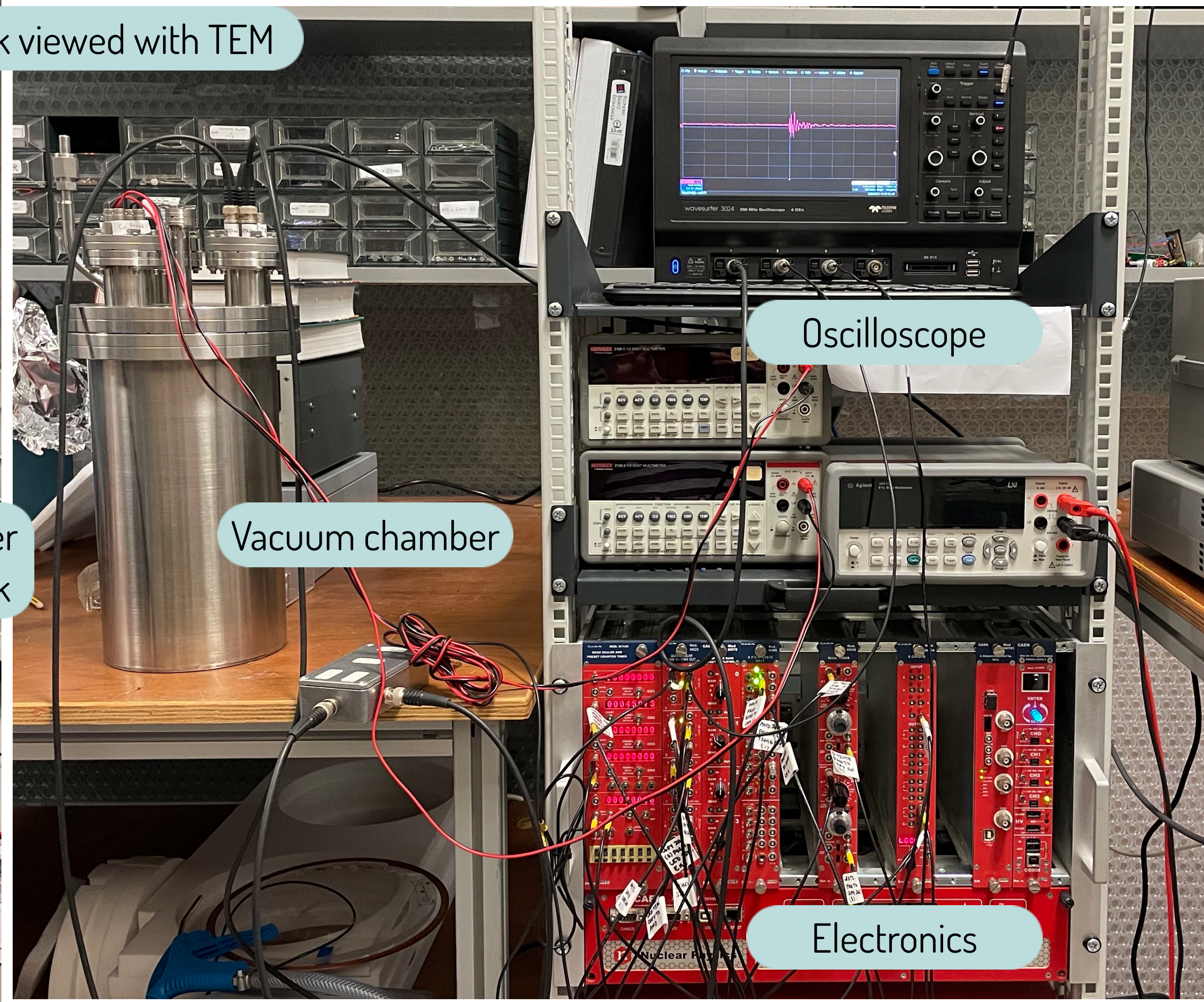
Stack

Micrometer holder for lens and stack



SPAD

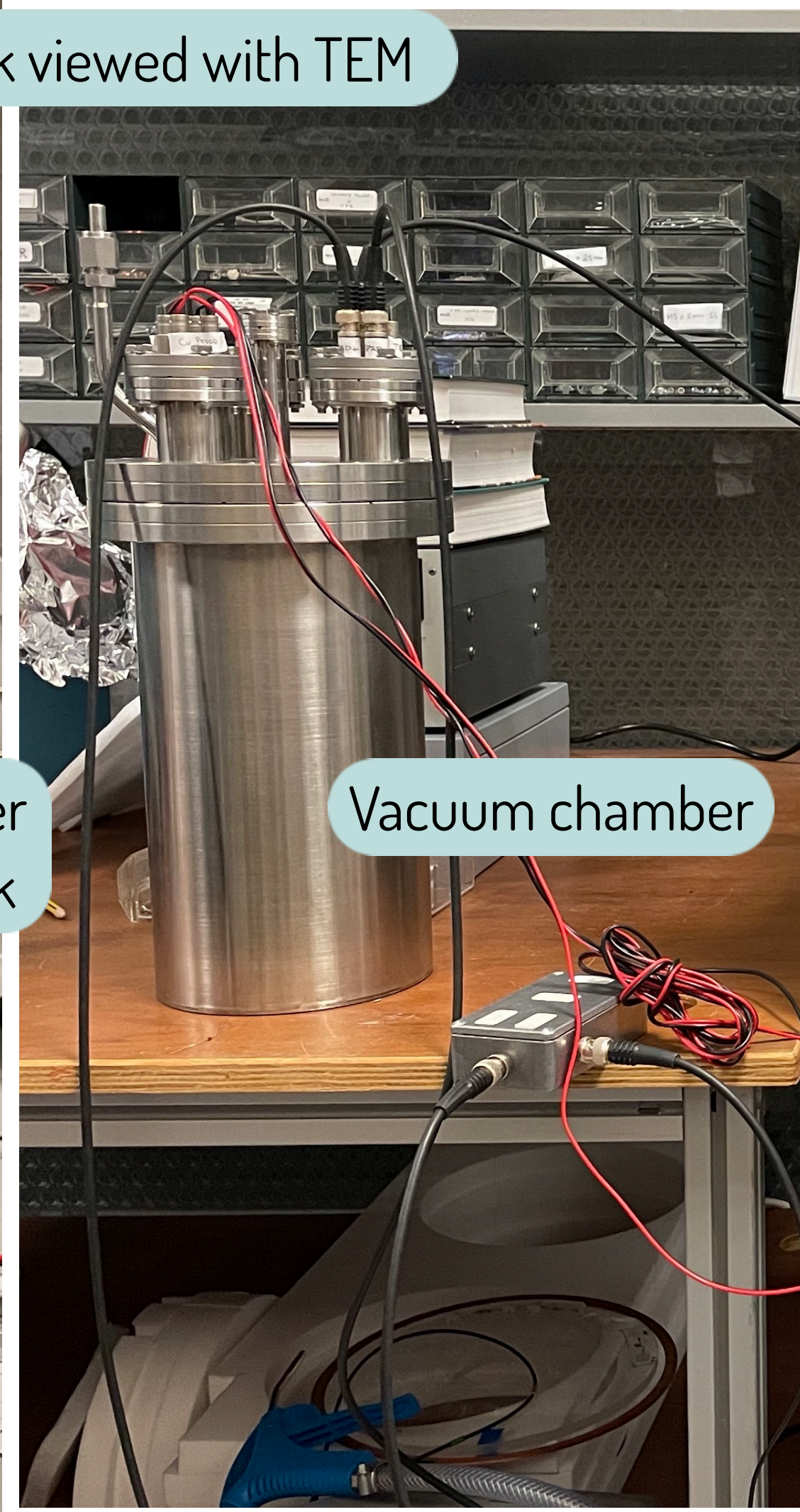
Stack viewed with TEM

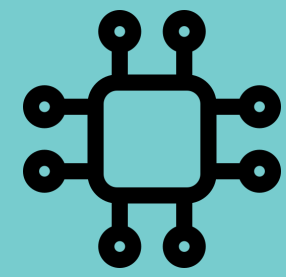


Vacuum chamber

Oscilloscope

Electronics





THE EXPERIMENT

**“OFF”
MEASUREMENT**



1. No stack present gives measurement of background.

**“ON”
MEASUREMENT**

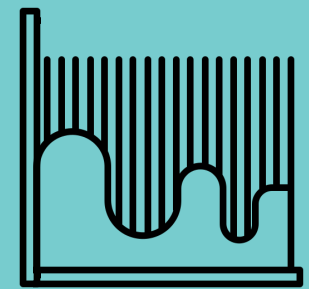


2. Stack present gives measurement of signal+background.

PLACE LIMITS



3. Compute the one-sided upper limit at 90% confidence level.



PROJECTED EXCLUSION LIMITS

“OFF” MEASUREMENT



1. No stack present gives measurement of background.

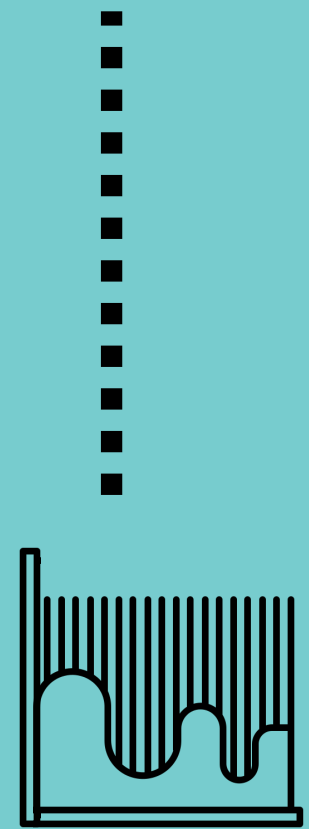
N/A

2. Assume no signal, i.e. $n_{on} = n_{off}$

PROJECTED LIMITS

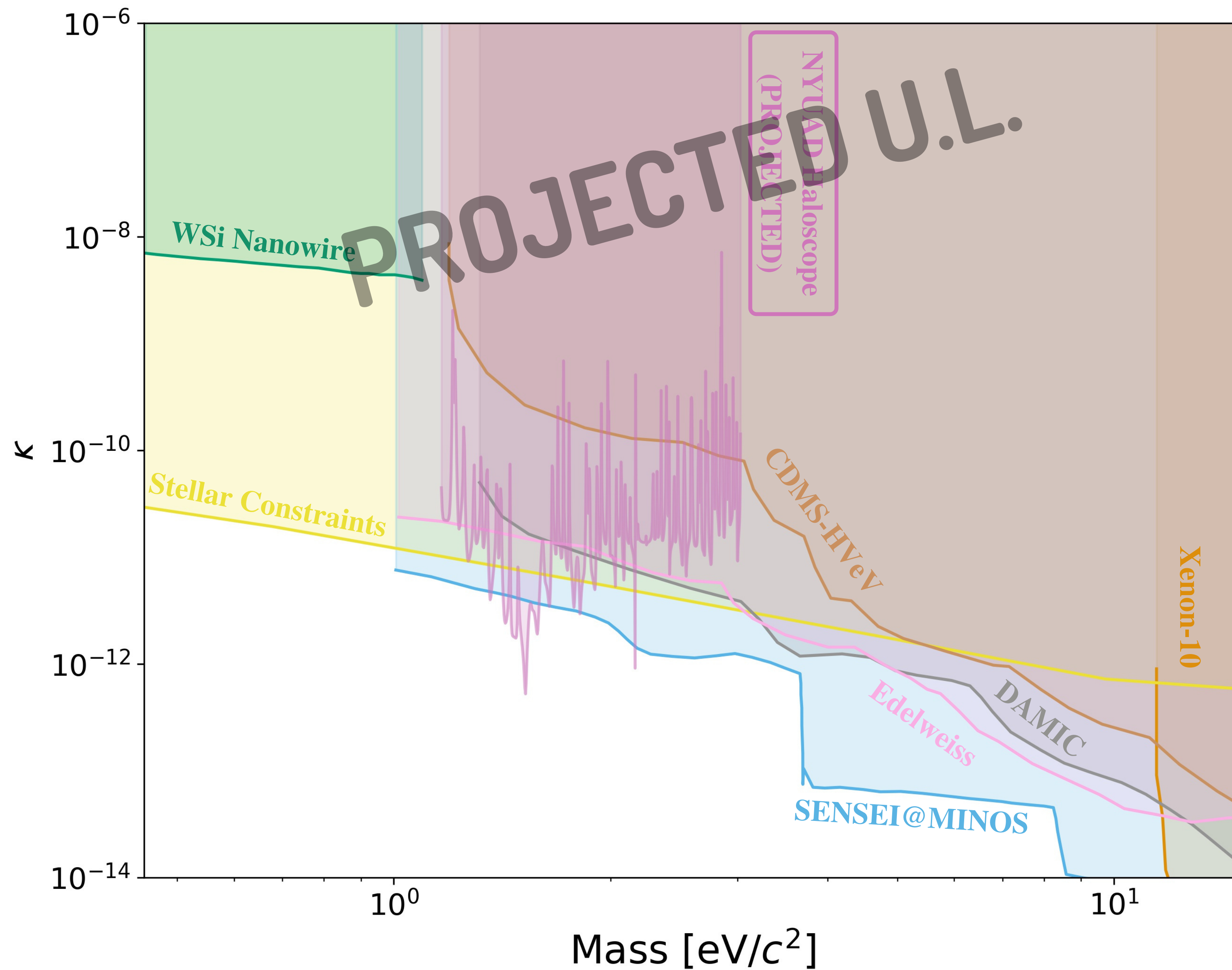


3. Compute the one-sided upper limit at 90% confidence level for a no signal measurement.



PROJECTED EXCLUSION LIMITS

N.B. Colorblind-
friendly plot thanks
to “color blind”
palette from the
seaborn library in
Python





CONCLUSIONS AND OUTLOOK



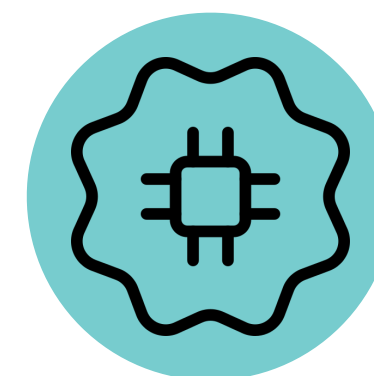
First-ever prototype of a multilayer dielectric haloscope in the 800 nm range

Economic prototype (~\$20,000) , that can be run in a standard physics lab with no special requirements.



Mechanical assembly developed and tested. The same setup will be used for the second stage of the experiment.

Complete the “on measurement” to place experimental upper limits on the mixing constant.



We are working with collaborators at INRiM (Turin, Italy) to develop a **Transition Edge Sensor (TES)** that will have dark count rates of the order of 10^{-4} Hz.