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Update from the  
**XENON dark matter project**  
**OBSERVATION OF EXCESS  
ELECTRONIC-RECOIL EVENTS  
IN XENON1T**  
**PRD 102, 072004 (2020)**

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PATRAS WORKSHOP  
16th June 2021

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On behalf of:  
XENON Collaboration





Columbia



KIT



Nikhef



Muenster



Stockholm



Mainz



MPIK, Heidelberg



Freiburg



University of Zurich

Zurich



Chicago



UCSD



Rice



Purdue



Subatech



Coimbra



LPNHE



Torino



Bologna



L'Aquila



LNGS



Napoli



Weizmann



NYUAD



Tsinghua



Tokyo

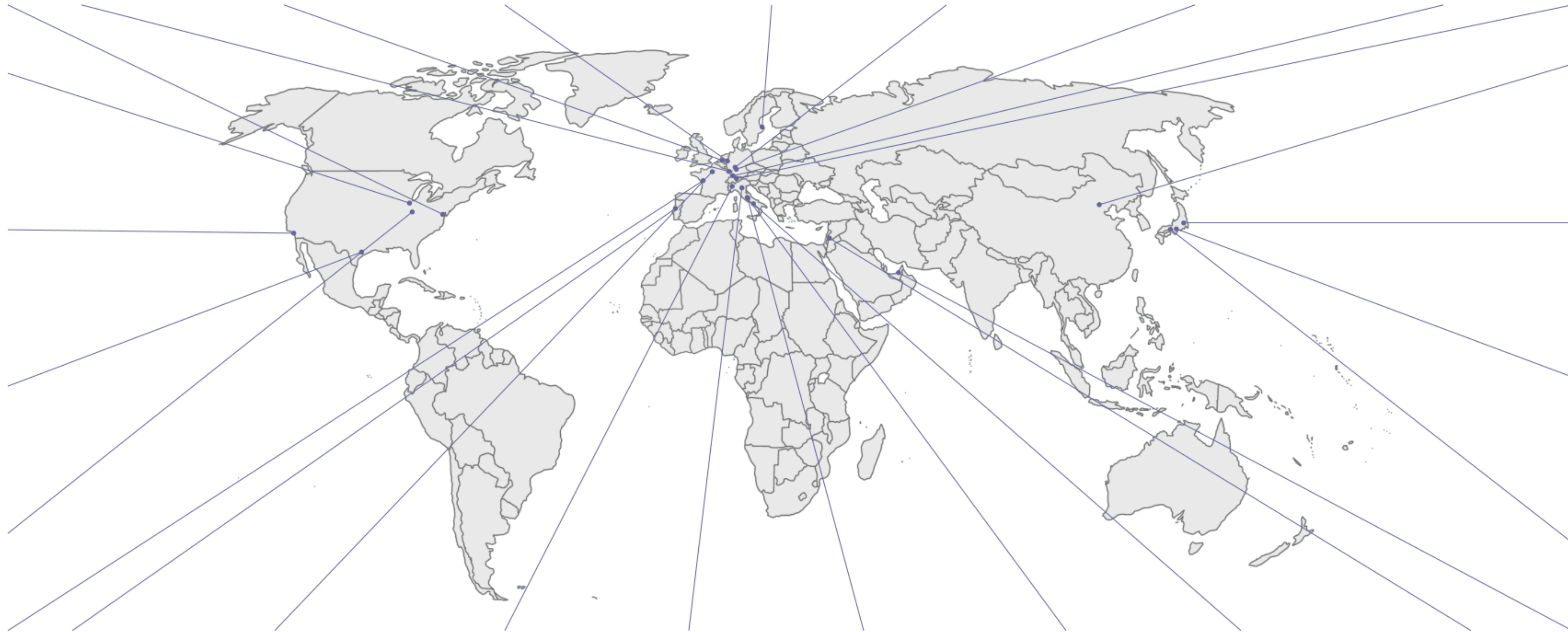


NAGOYA UNIVERSITY

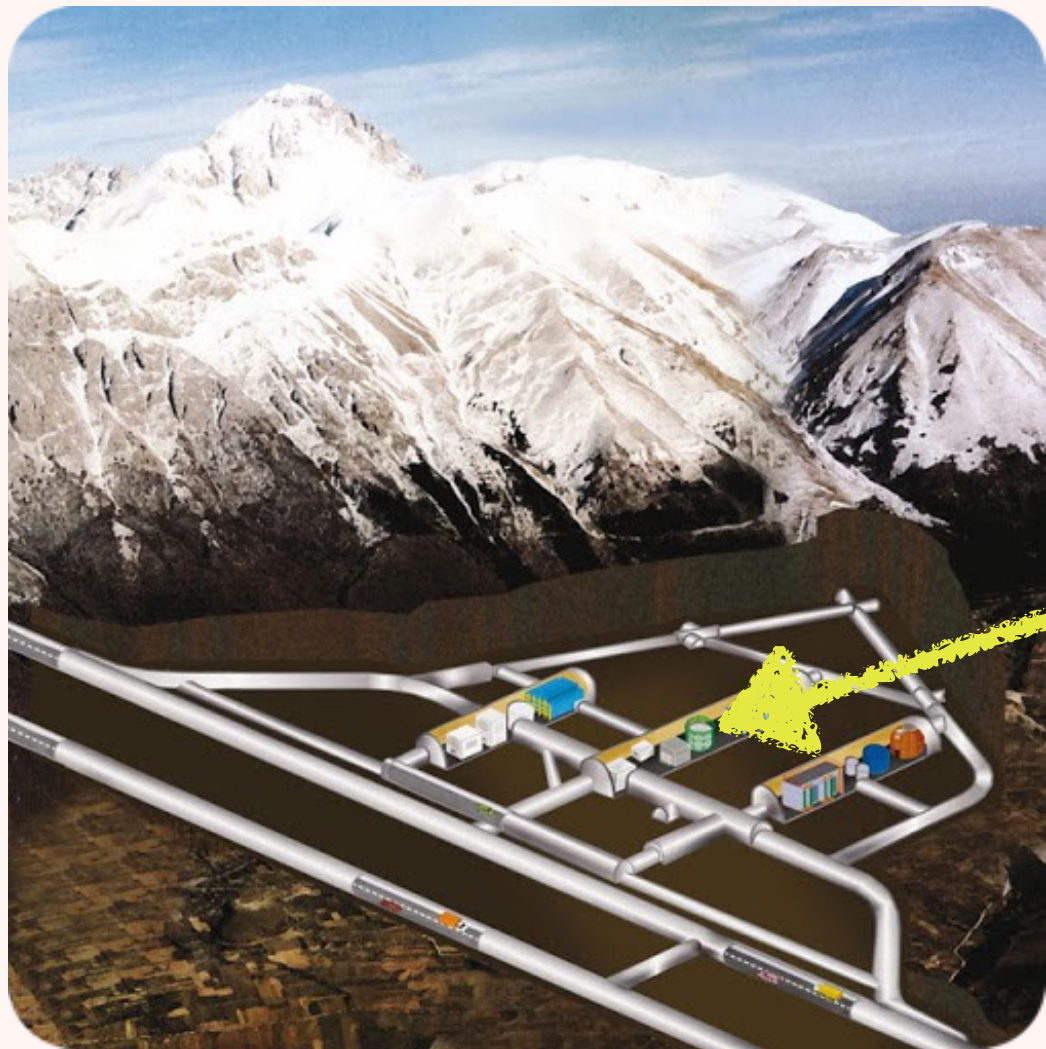
Nagoya



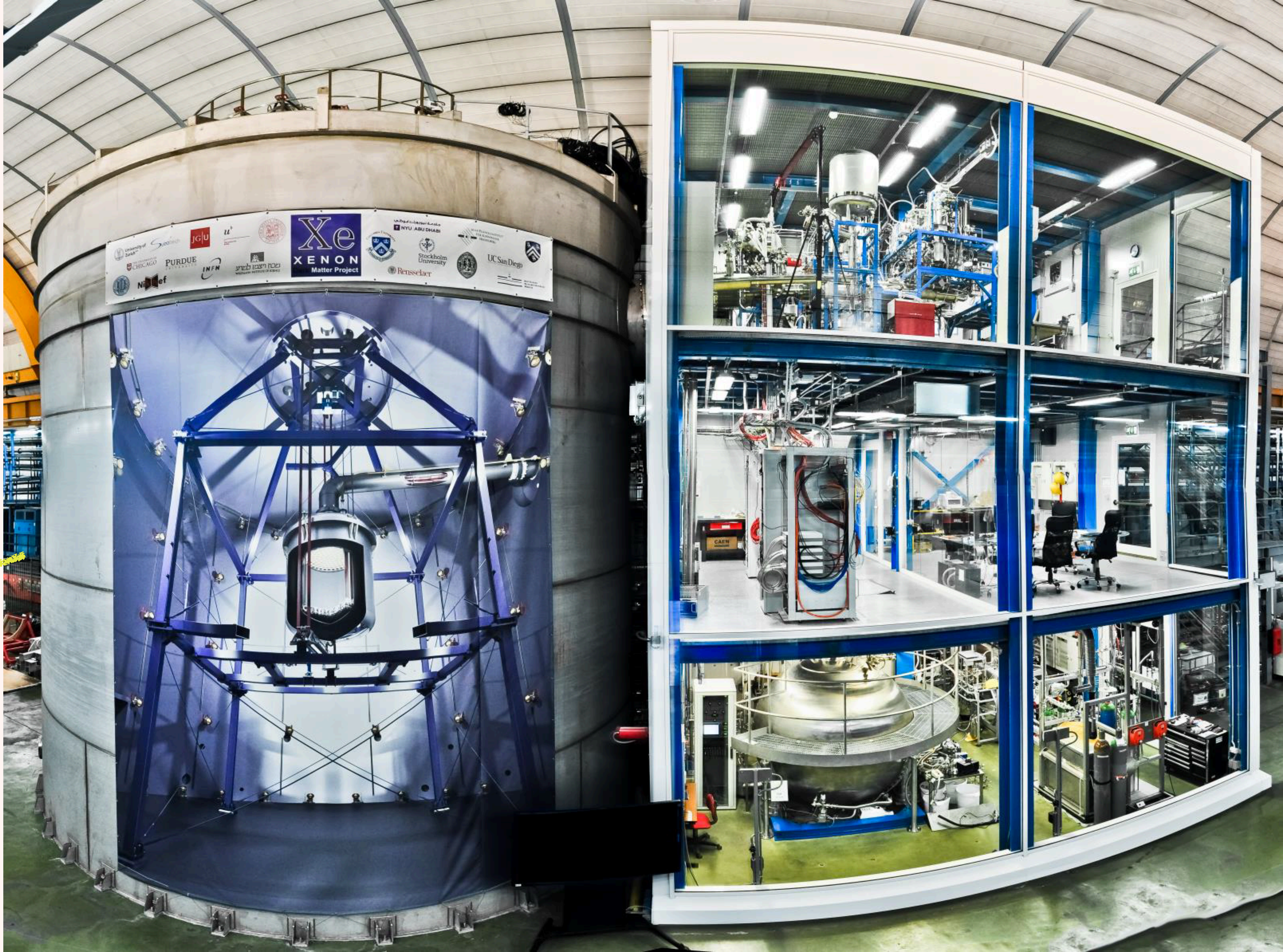
Kobe



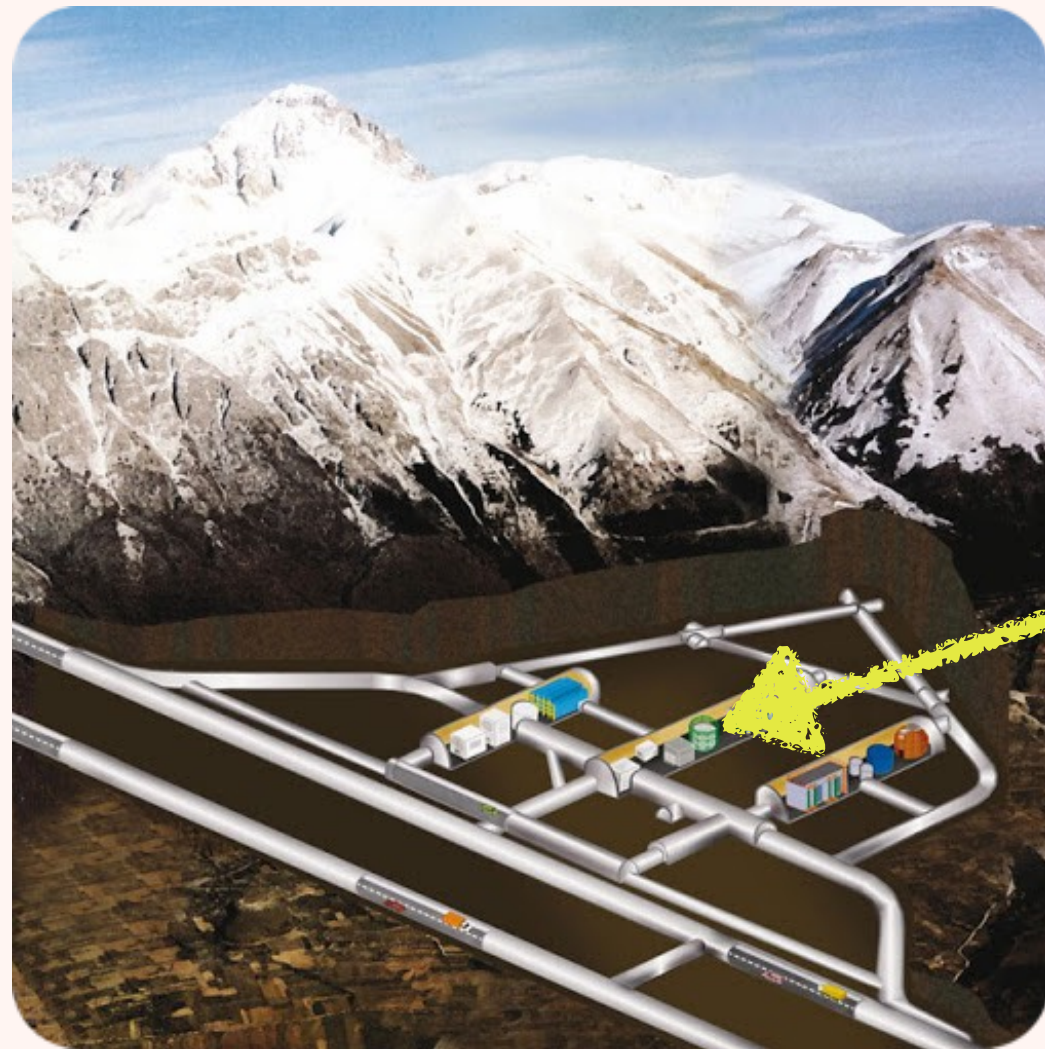




Laboratori Nazionali  
del Gran Sasso  
Italy





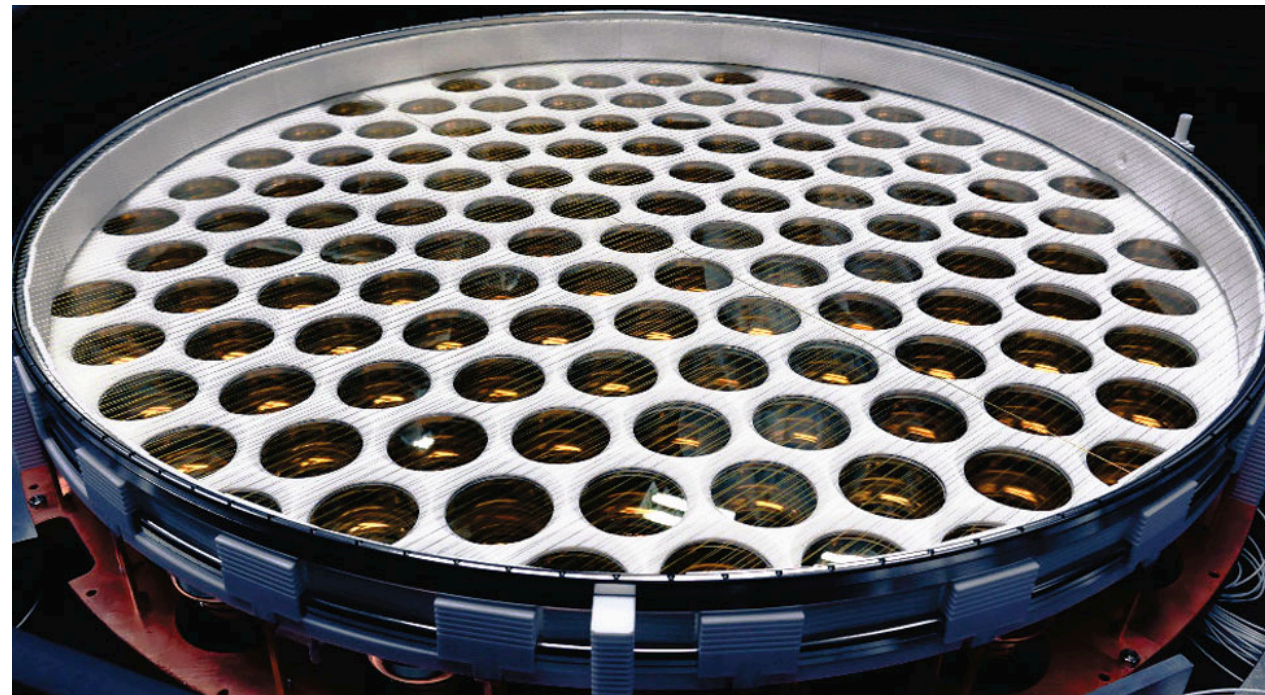
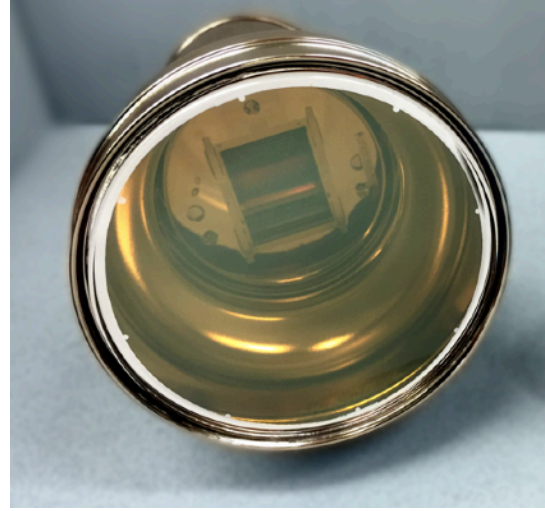


Laboratori Nazionali  
del Gran Sasso  
Italy



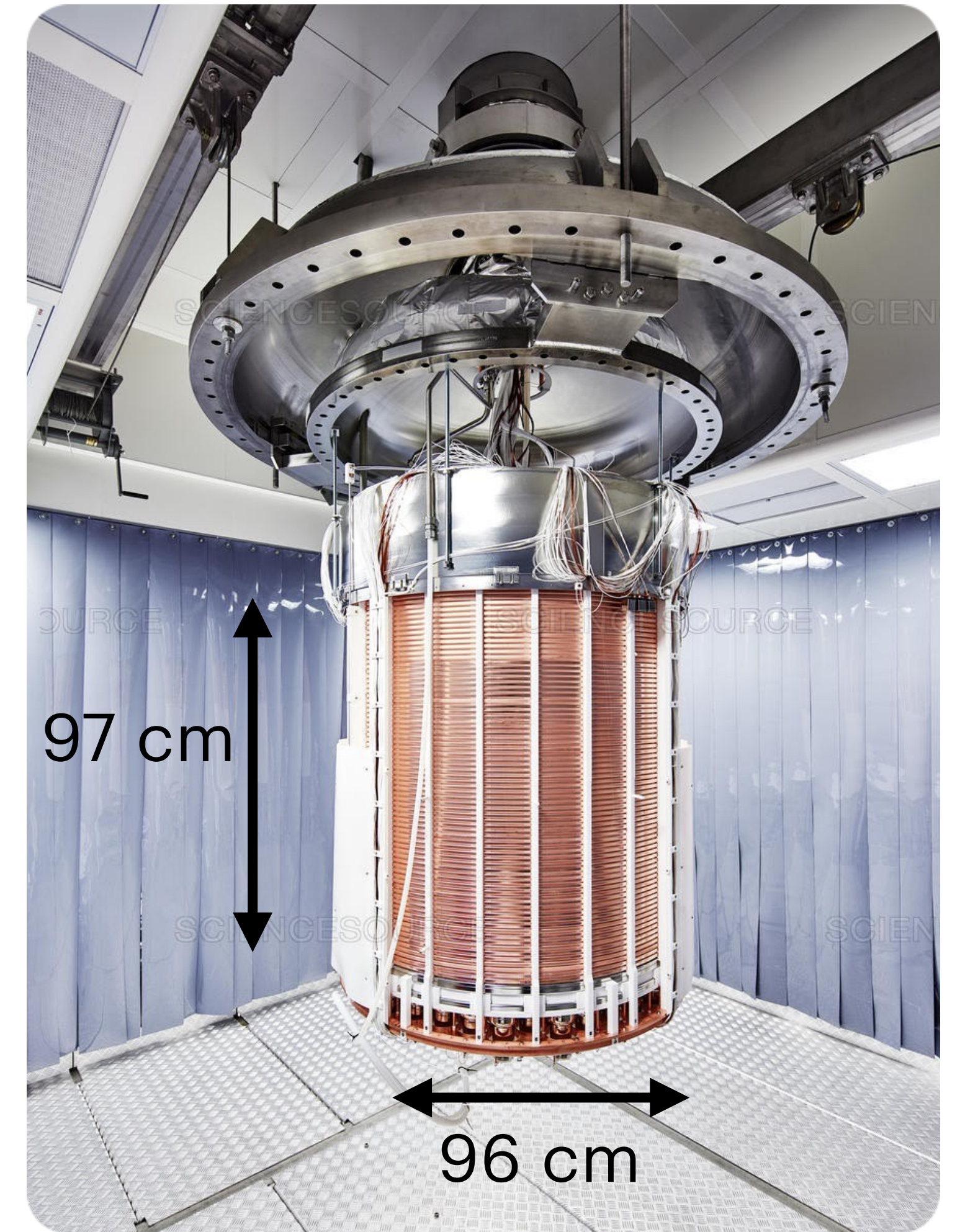
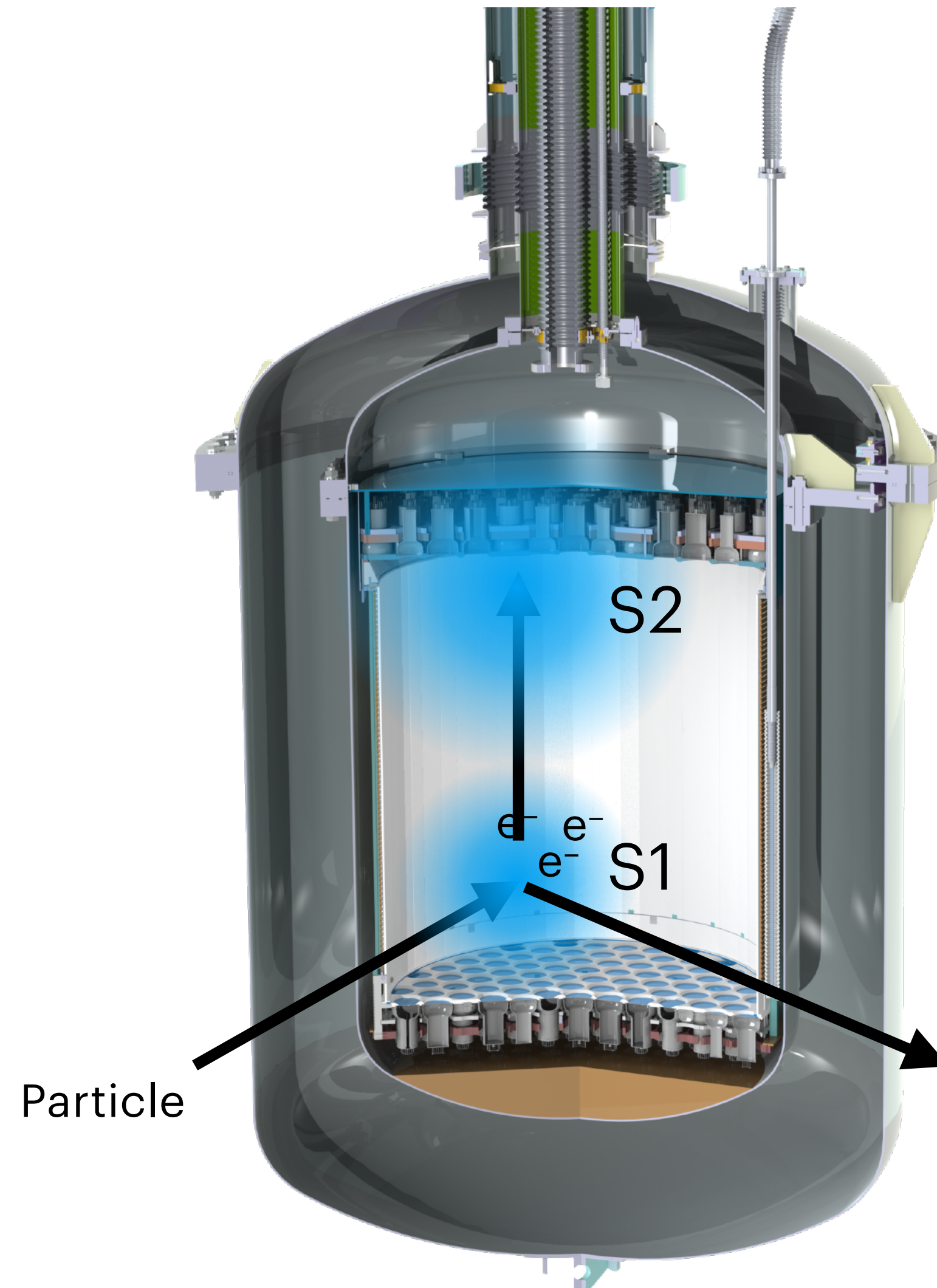


# The TPC



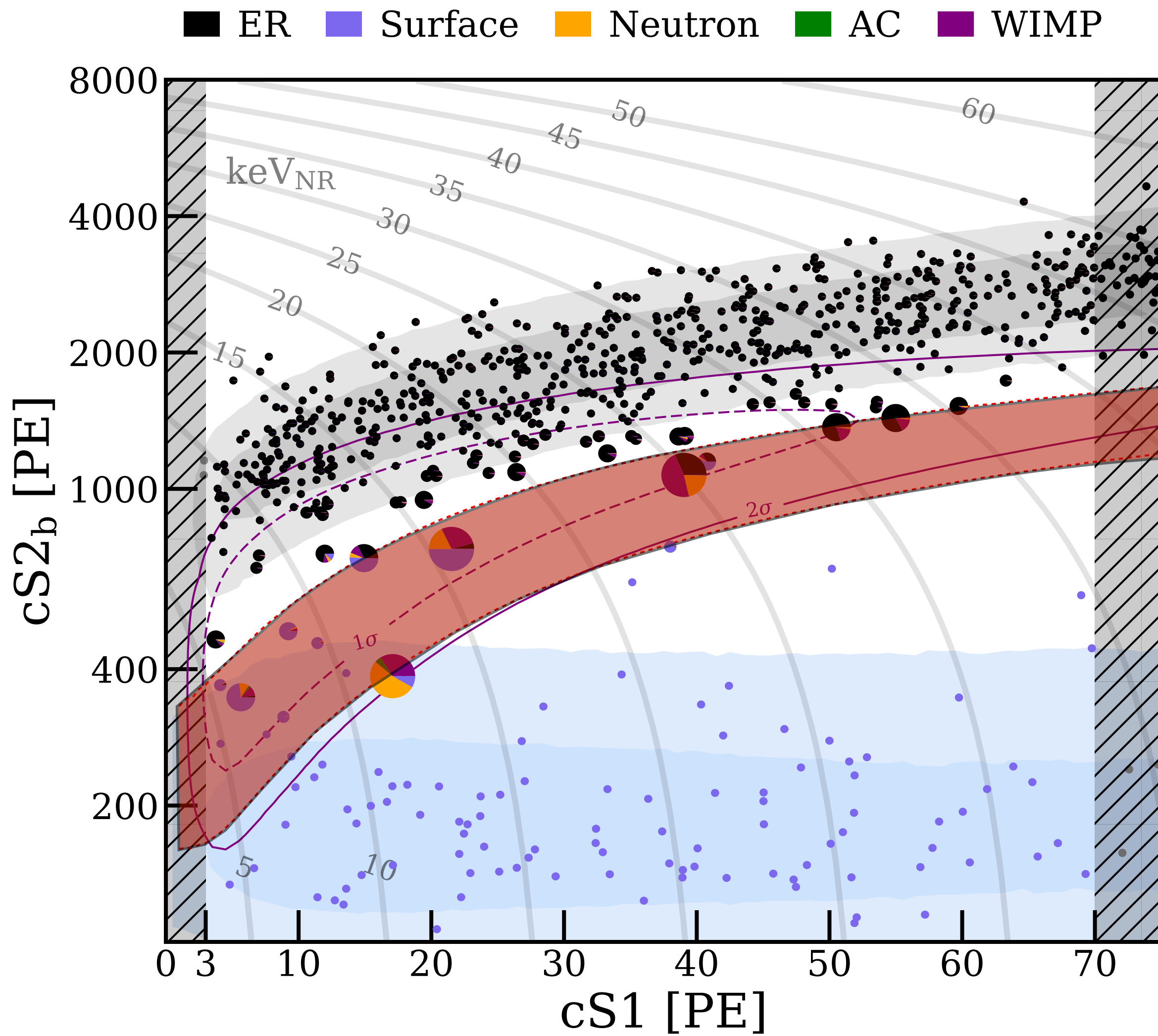
2 t Liquid xenon in TPC  
3.2 t in total

248 PMTs

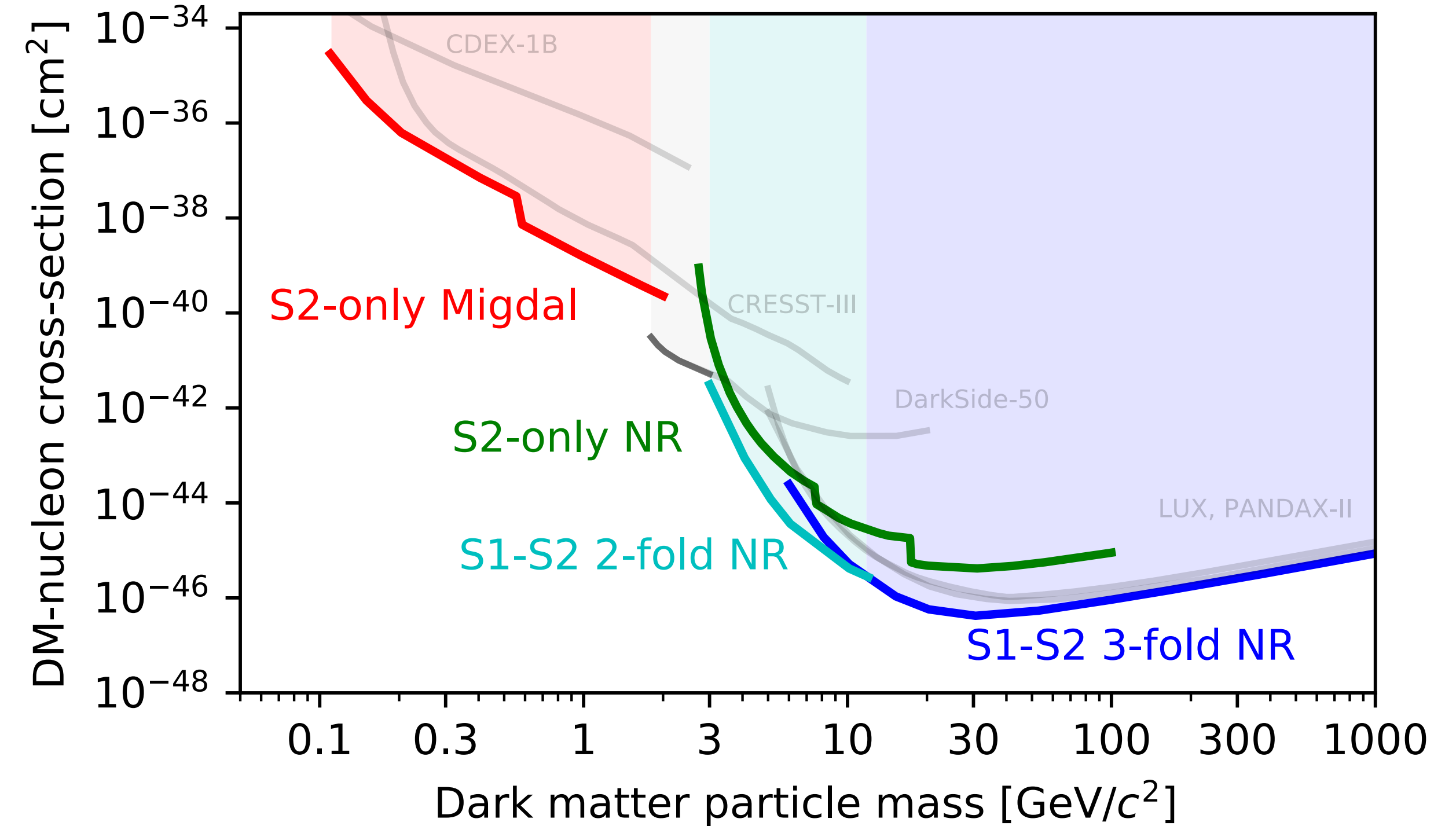




# Nuclear recoil searches



No excess of NR events found



PRL 123, 241803 (2019)

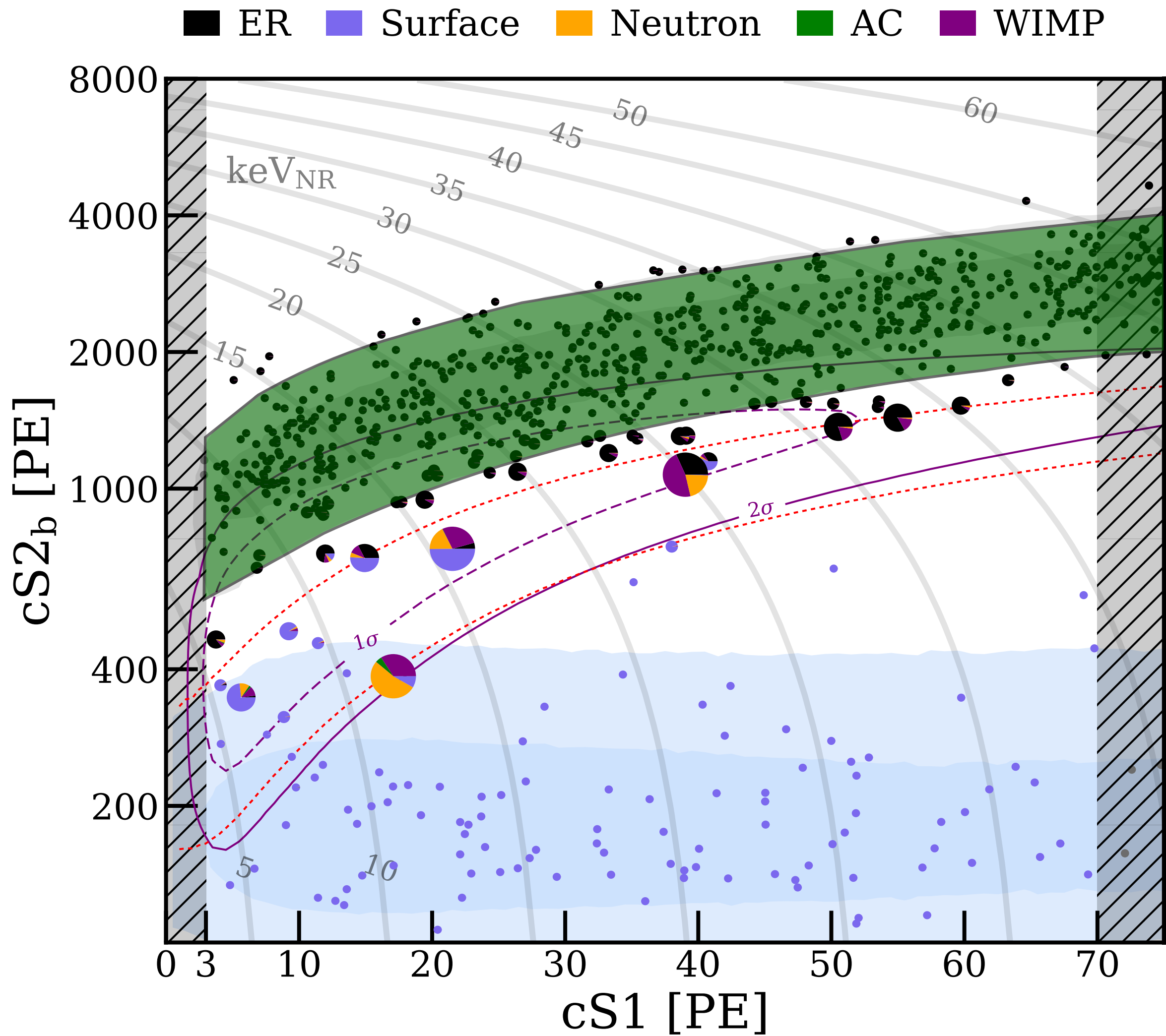
PRL 126, 091301 (2021)

PRL 123, 251801 (2019)

PRL 121, 111302 (2018)



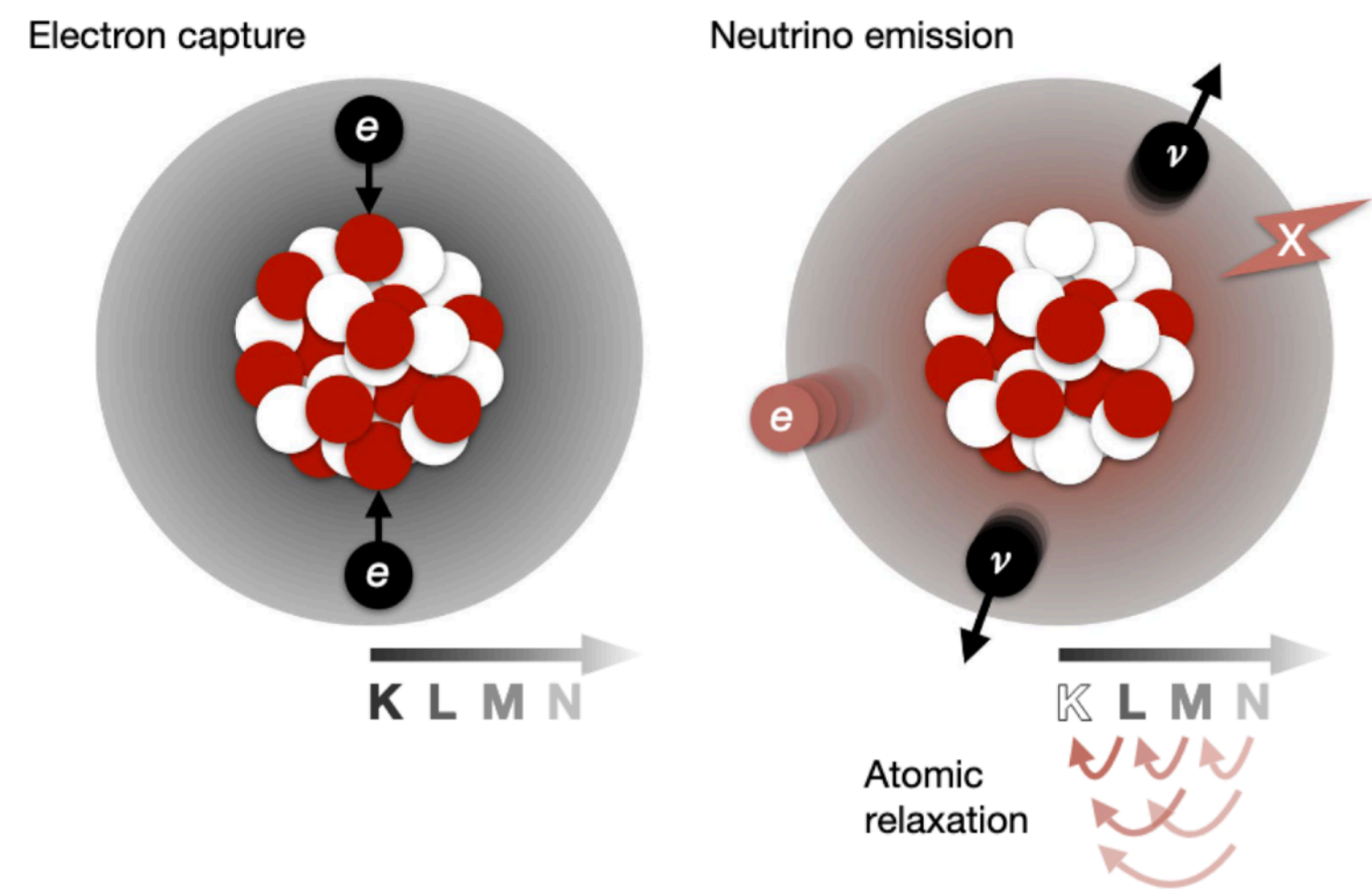
# What about the electronic recoils?



**This talk**

Search for excess above known BGs

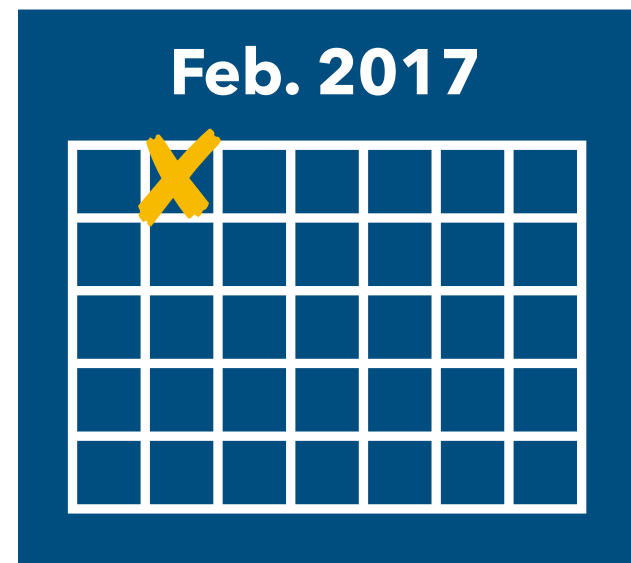
Nature 568, 532 (2019)  
— Double electron capture  $^{124}\text{Xe}$



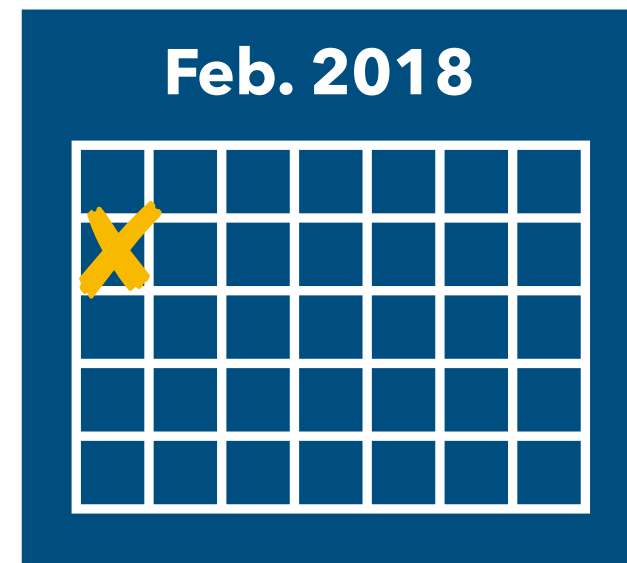
**This talk: low energies (up to 10 keV)**



# The low energy excess



Science Run I  
(SRI)



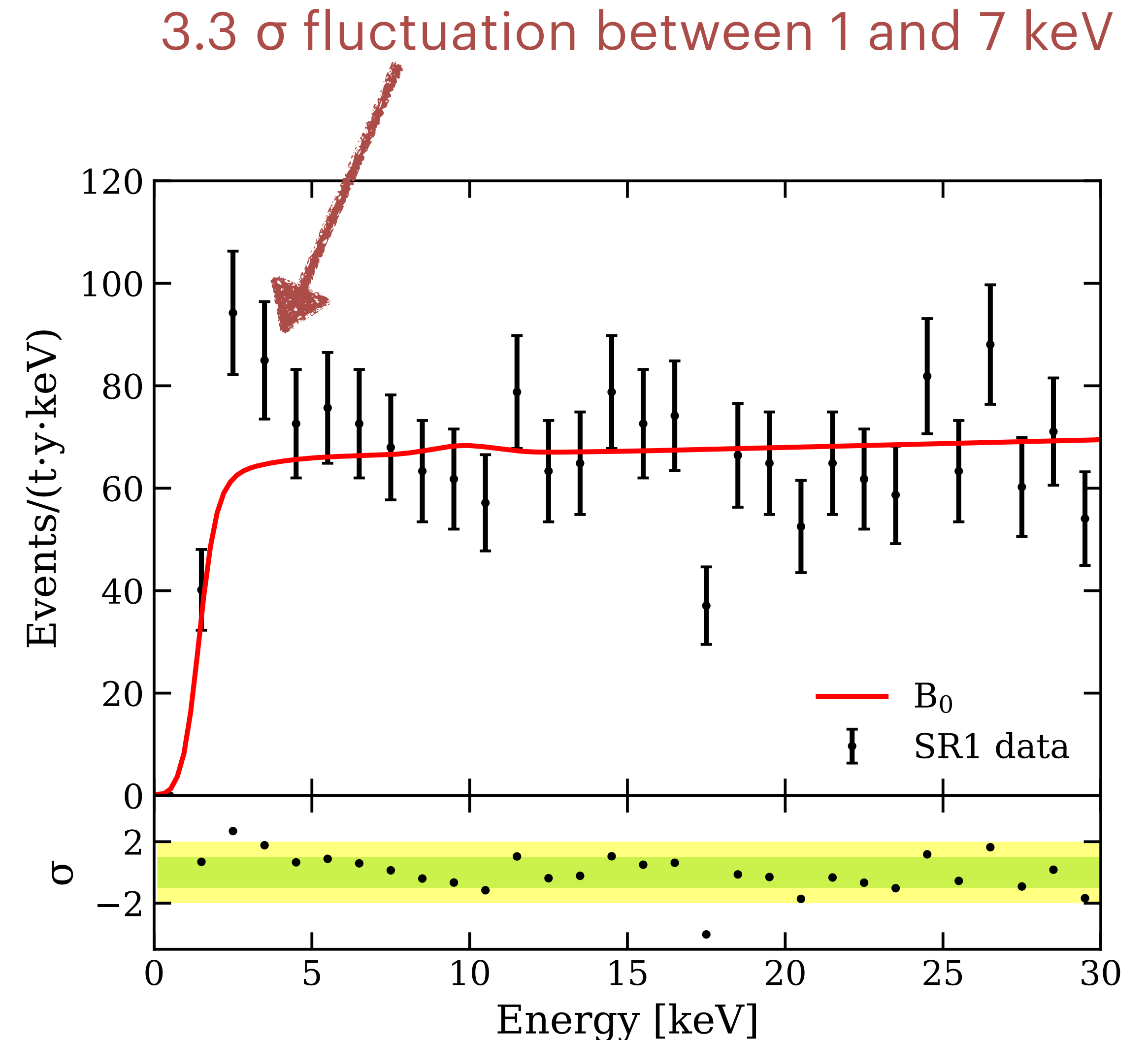
226.9 live days

1 tonne fiducial volume

[1, 210] keV energy range

Consider efficiencies of reconstruction and data quality cuts

Threshold at 10% detection efficiency



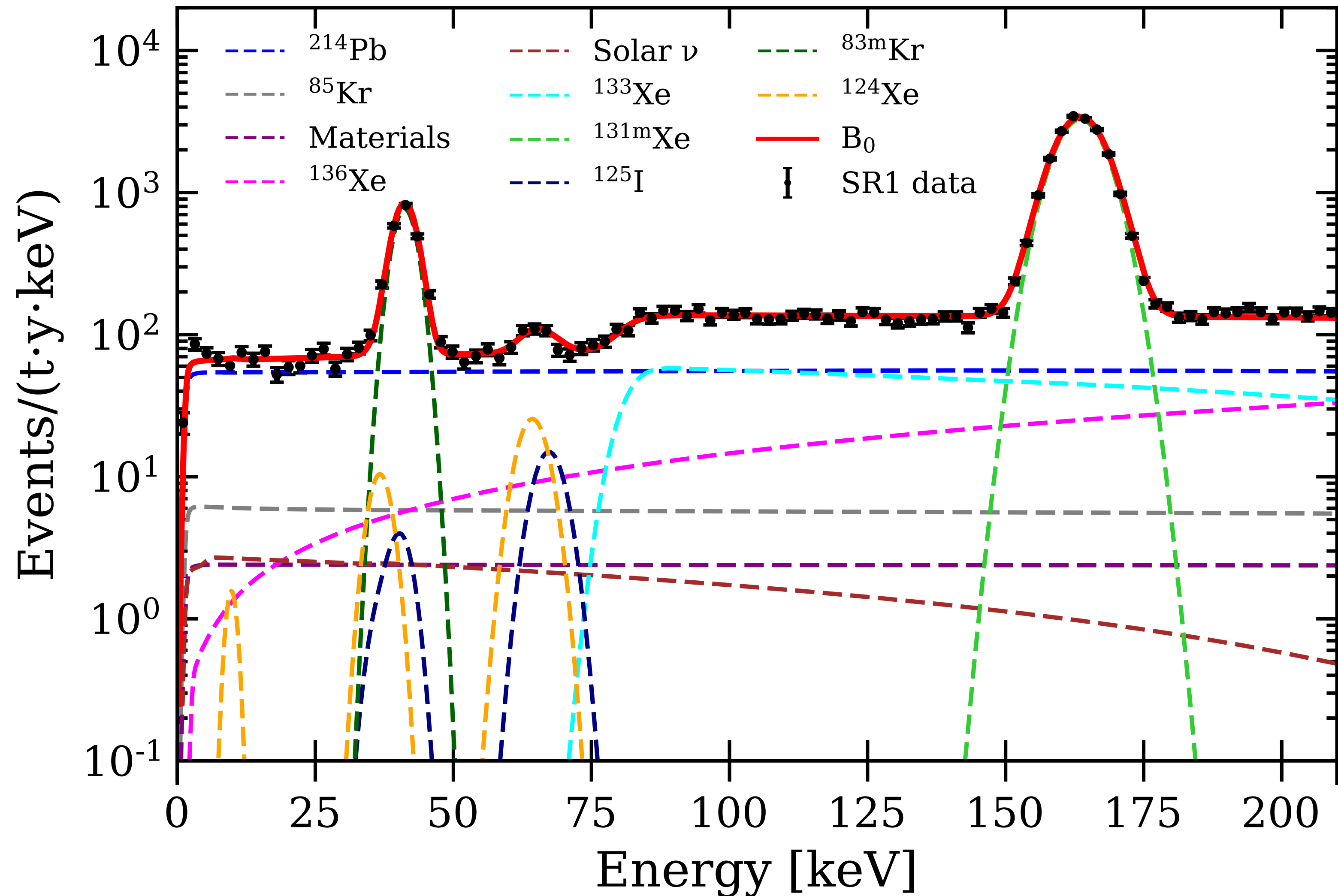


# Background modelling

Background sources modelled with Geant4

Most rates constrained by other measurements or time dependence

Search for excess over known backgrounds between 1 and 210 keV





# Efficiency and energy reconstruction



Mistake in energy reconstruction?

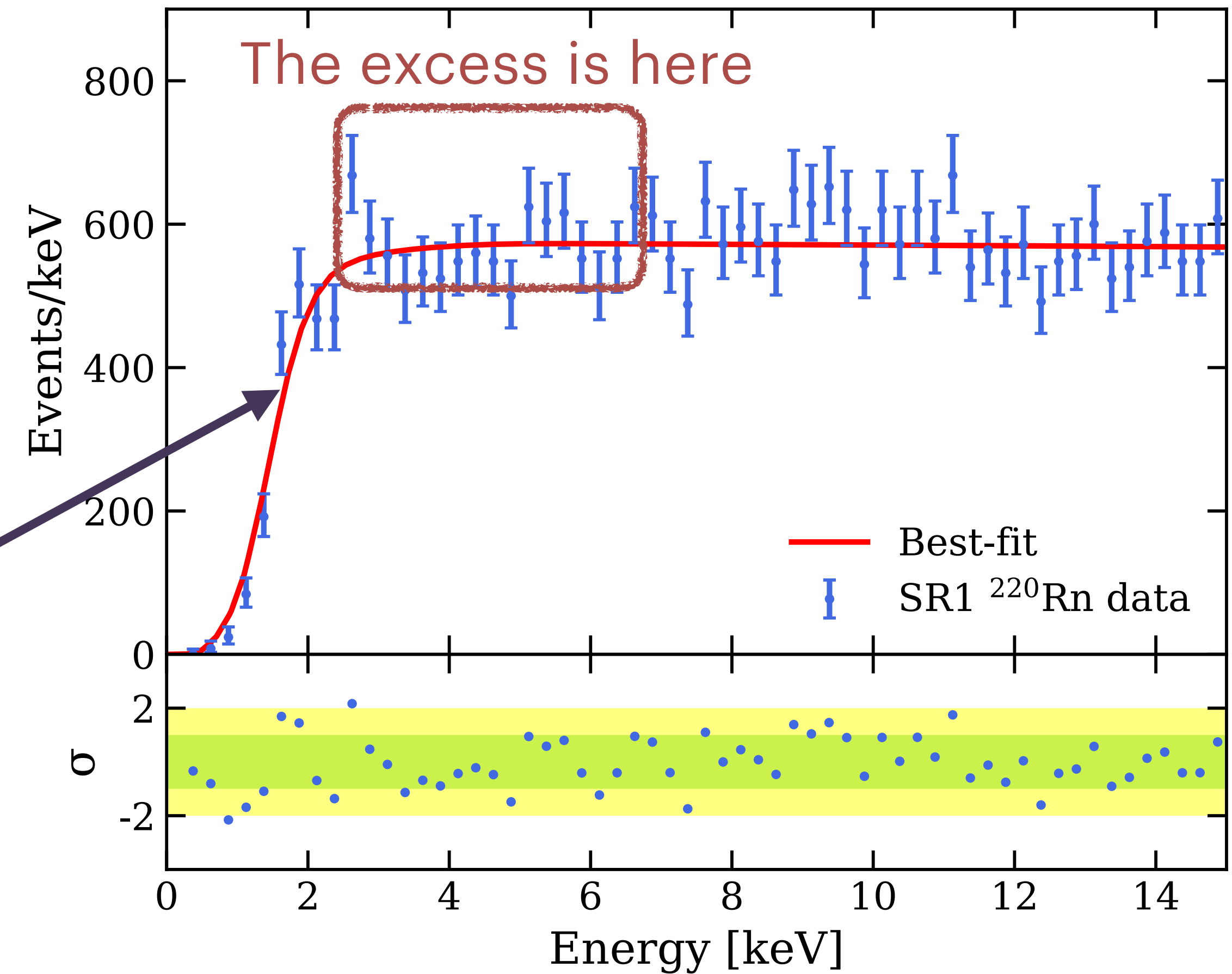
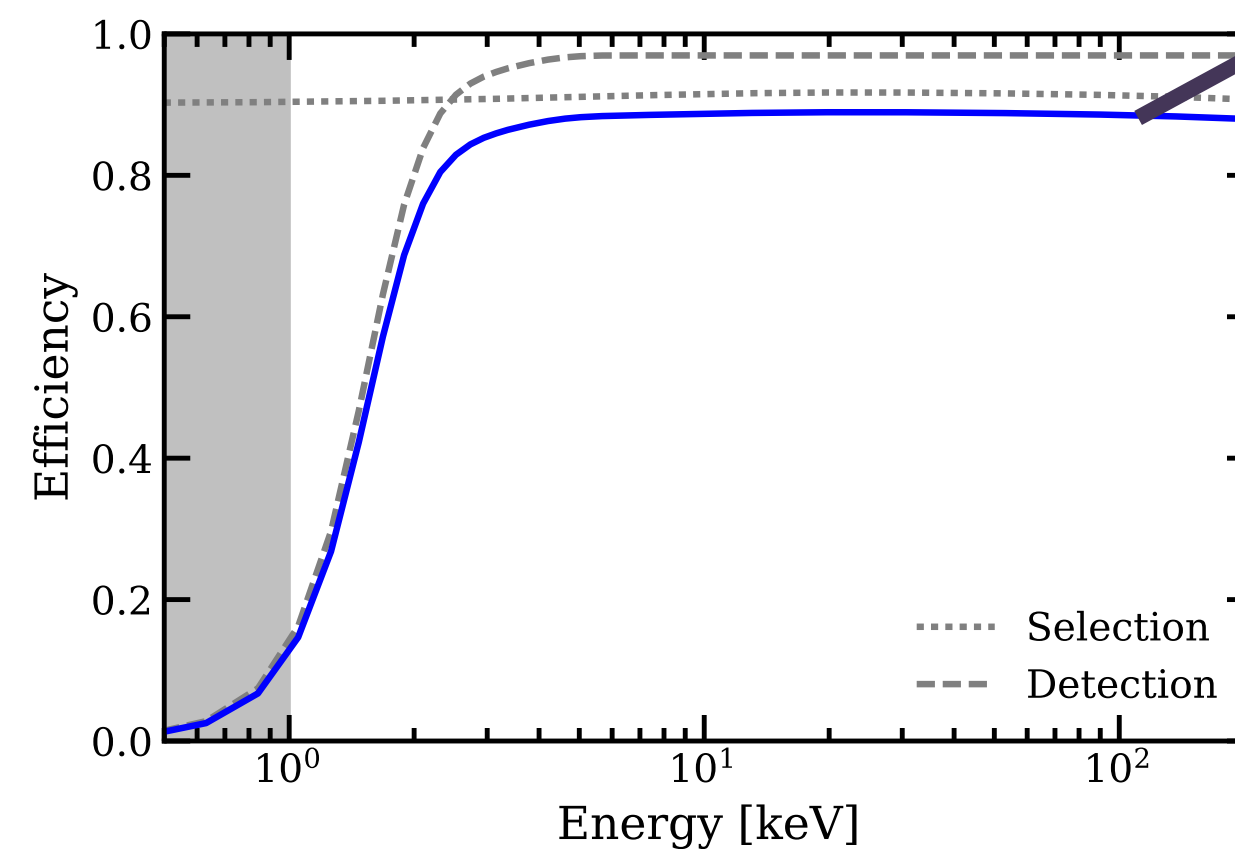
Mis-modelled efficiency?

Look at Rn-220 calibration data

Beta-decay just like dominant background

p-value 0.58

Cannot explain  
the excess





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**IT COULD BE A NEW  
BACKGROUND**

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# Tritium

**$3.2\sigma$  over background**

**$(159 \pm 51)$  ev/keV/t/yr**

**$< 3$  atoms  $^3\text{H}$  / kg Xe**

Beta decay

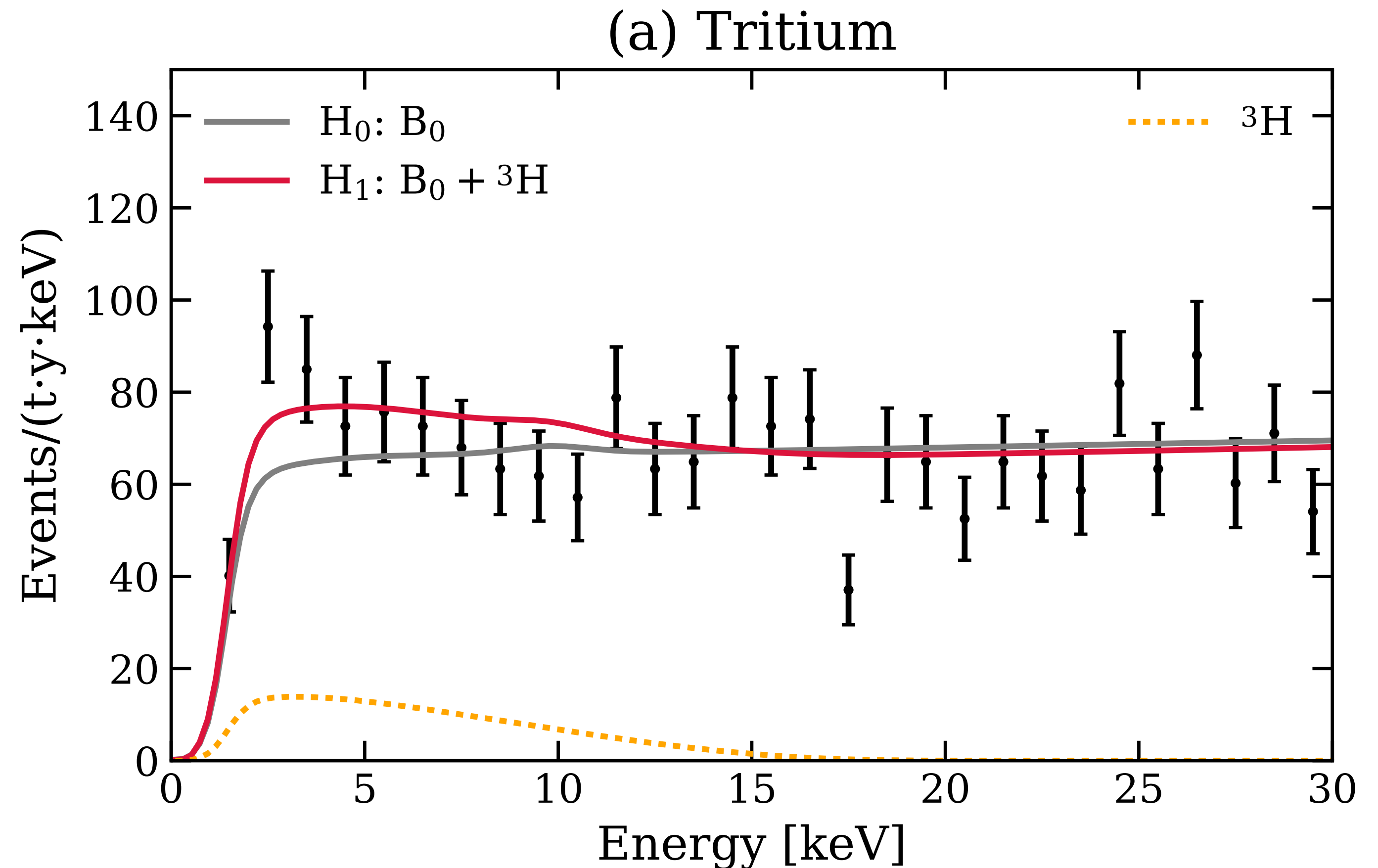
Q value 18.6 keV

Half life 12.3 years



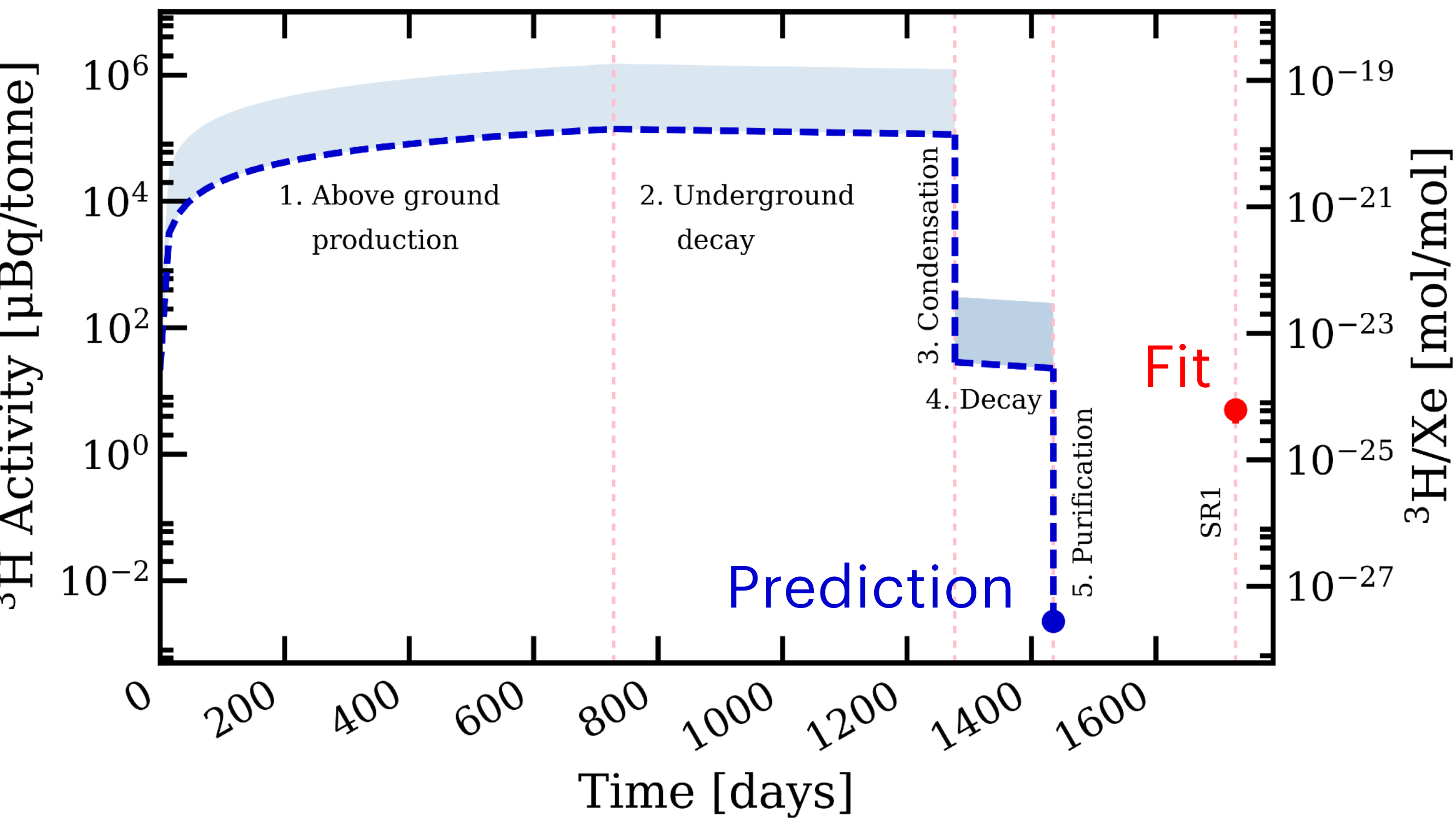
Where from?

Emanation from detector materials neither confirmed nor ruled out





# Tritium — possible origins



Where from?

Cosmogenic activation?

Xe spallation

31.58/kg/d at sea level

(Zhang et al., Astropart. Phys 84, 62 (2016))

Seems unlikely



# Tritium — possible origins



Where from?

Emanation from detector materials?

Atmospheric abundance  $(5-10) \times 10^{-18}$  HTO/H<sub>2</sub>O

Best fit  $\Rightarrow$  60–120 ppb H<sub>2</sub>O+H<sub>2</sub>

Can neither confirm nor rule out tritium

All other significances reported  
both with and without  
tritium in BG mode

HTO

Light yield  $\Rightarrow$  O(1) ppb H<sub>2</sub>O

HT

Electron lifetime (xenon purity)  
 $\Rightarrow$  < ppb O<sub>2</sub>-equivalent impurities

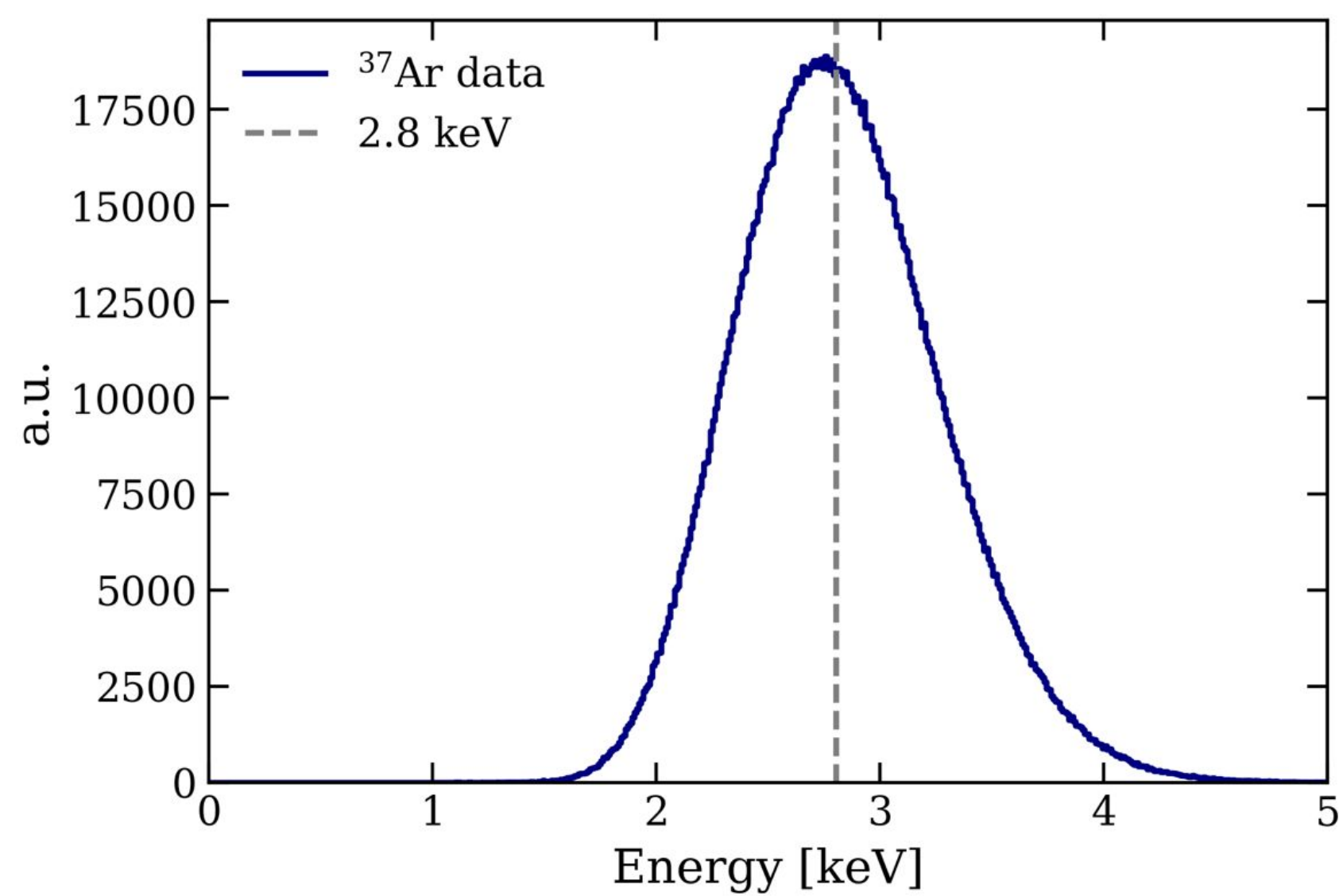


# Argon-37

2.8 keV energy released after EC

X-rays / Auger electrons

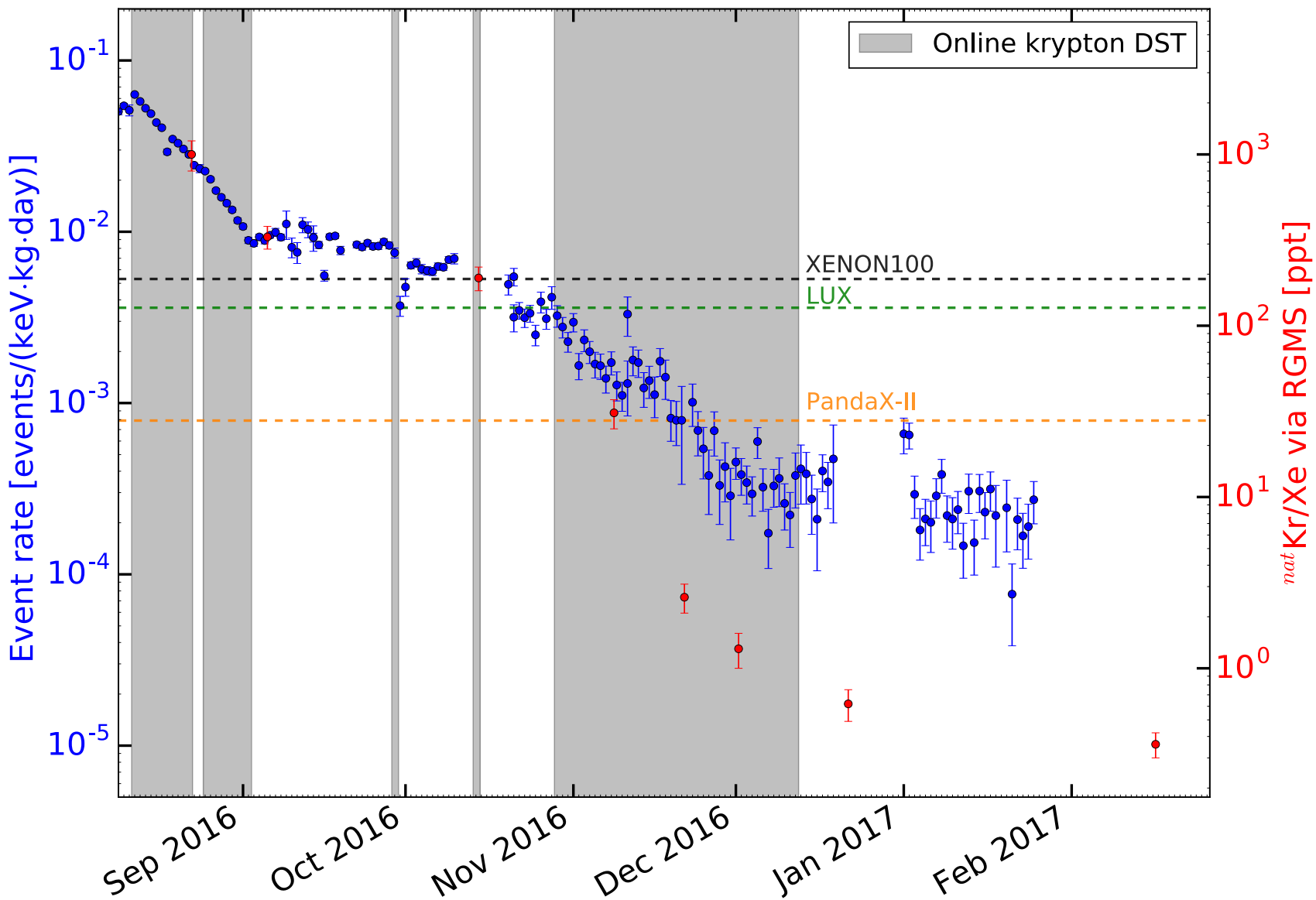
Tested as calibration source



Where from?

Always present? No, removed by distillation

Air leak? Would also introduce Kr



Ruled out



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**COULD IT BE  
NEW PHYSICS?**

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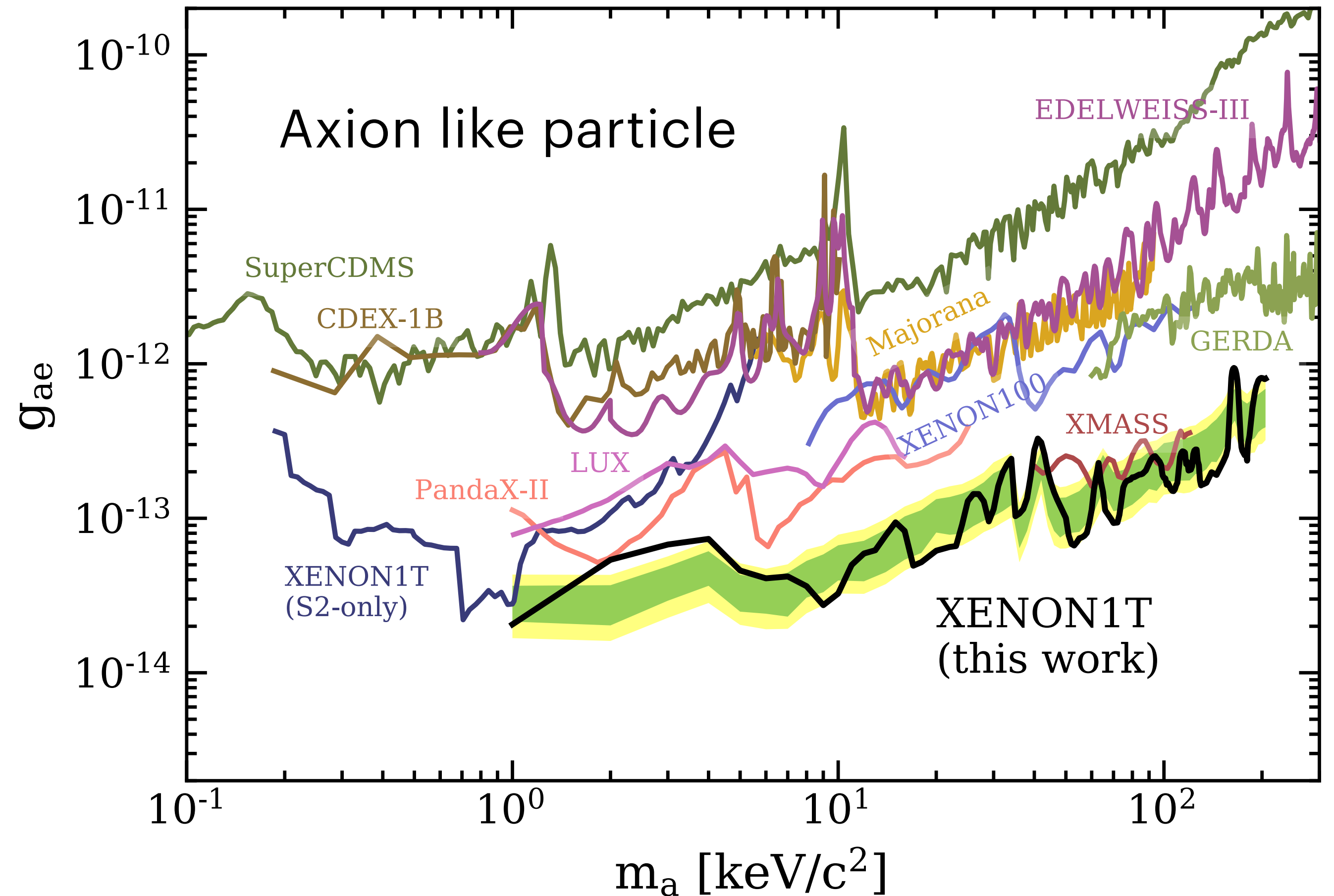
# Bosonic dark matter

Search for a mono-energetic peak

Could be dark matter,  
e.g. axion-like particle or dark photon

Most significant at  $2.3 \pm 0.2$  keV

No  $> 3\sigma$  excess  $\Rightarrow$  only report limits



**Mono-energetic peak:  $3.0\sigma$  over background (global)**



# Neutrino magnetic moment

Mag. Moment:  $3.2\sigma$  over background

with  $^3\text{H}$ :  $0.9\sigma$  over background +  $^3\text{H}$

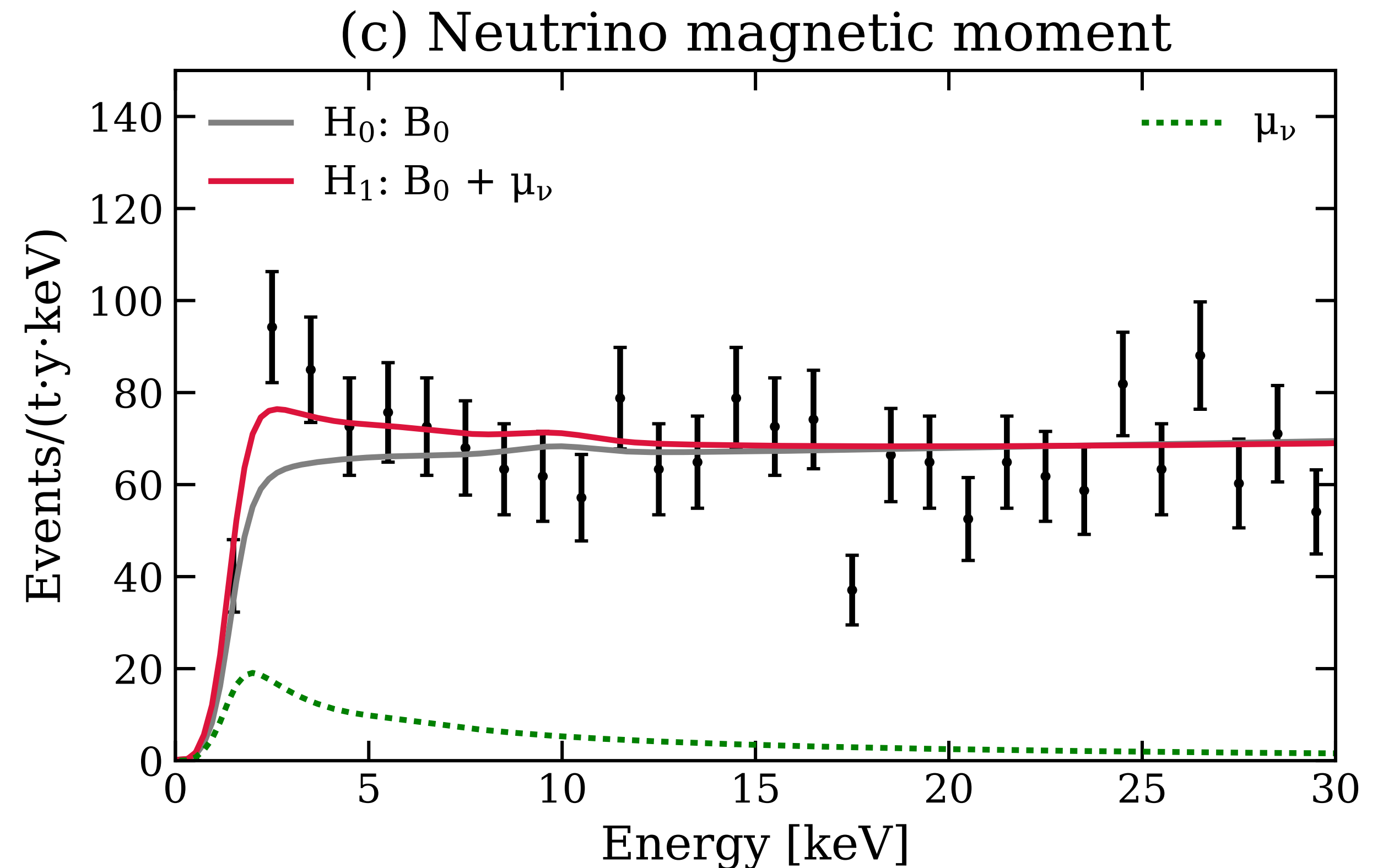
$$\mu_\nu : (1.4 - 2.9) \times 10^{-11} \mu_B$$

$$\gtrsim 10^{-15} \mu_B \implies \text{Majorana neutrinos}$$

Compatible with other experiments

In tension with astrophysical constraints

arXiv 1910.10568, 1907.00115





# Solar axions

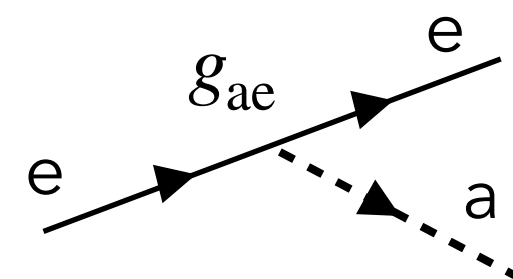
## Production

Solar physics

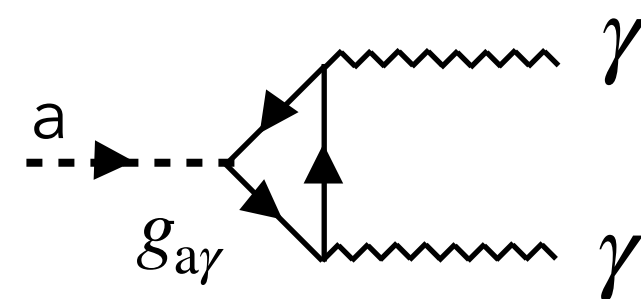


1. ABC

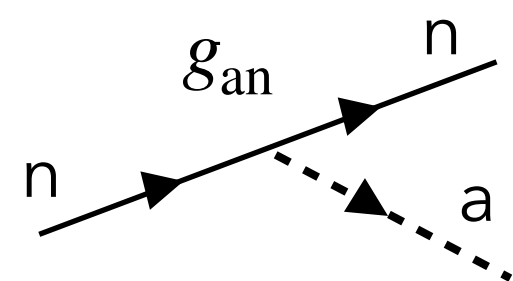
(atomic recombination and de-excitation, bremsstrahlung and Compton)



2. Primakoff



3.  $^{57}\text{Fe}$



## Detection

Axioelectric effect

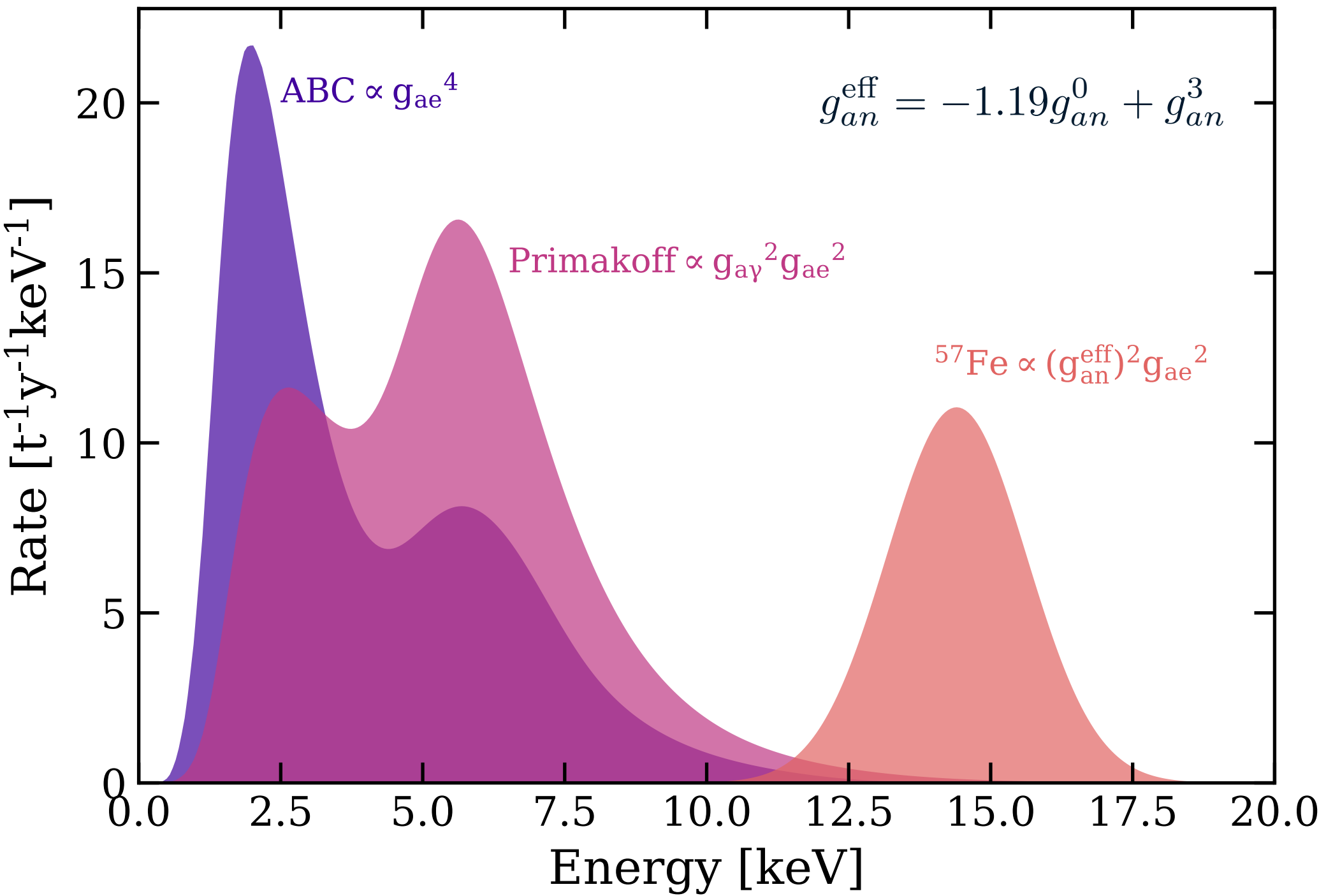


$$\sigma_{ae} = \sigma_{pe} \frac{g_{ae}^2}{\beta} \frac{3E_a^2}{16\pi\alpha m_e^2} \left(1 - \frac{\beta^{2/3}}{3}\right)$$

We ignore the inverse Primakoff effect

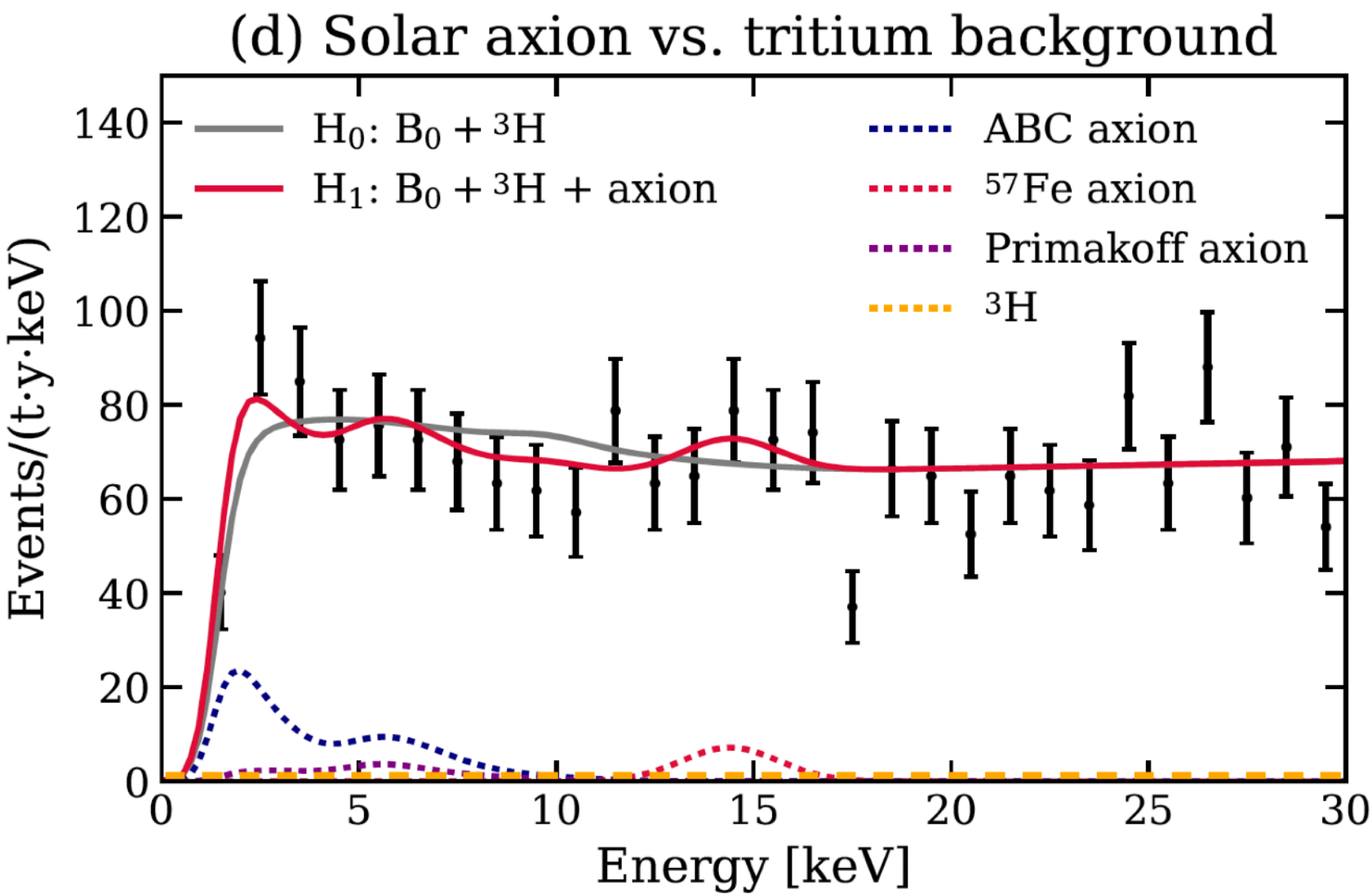
## Reconstruction

XENONIT resolution, efficiency

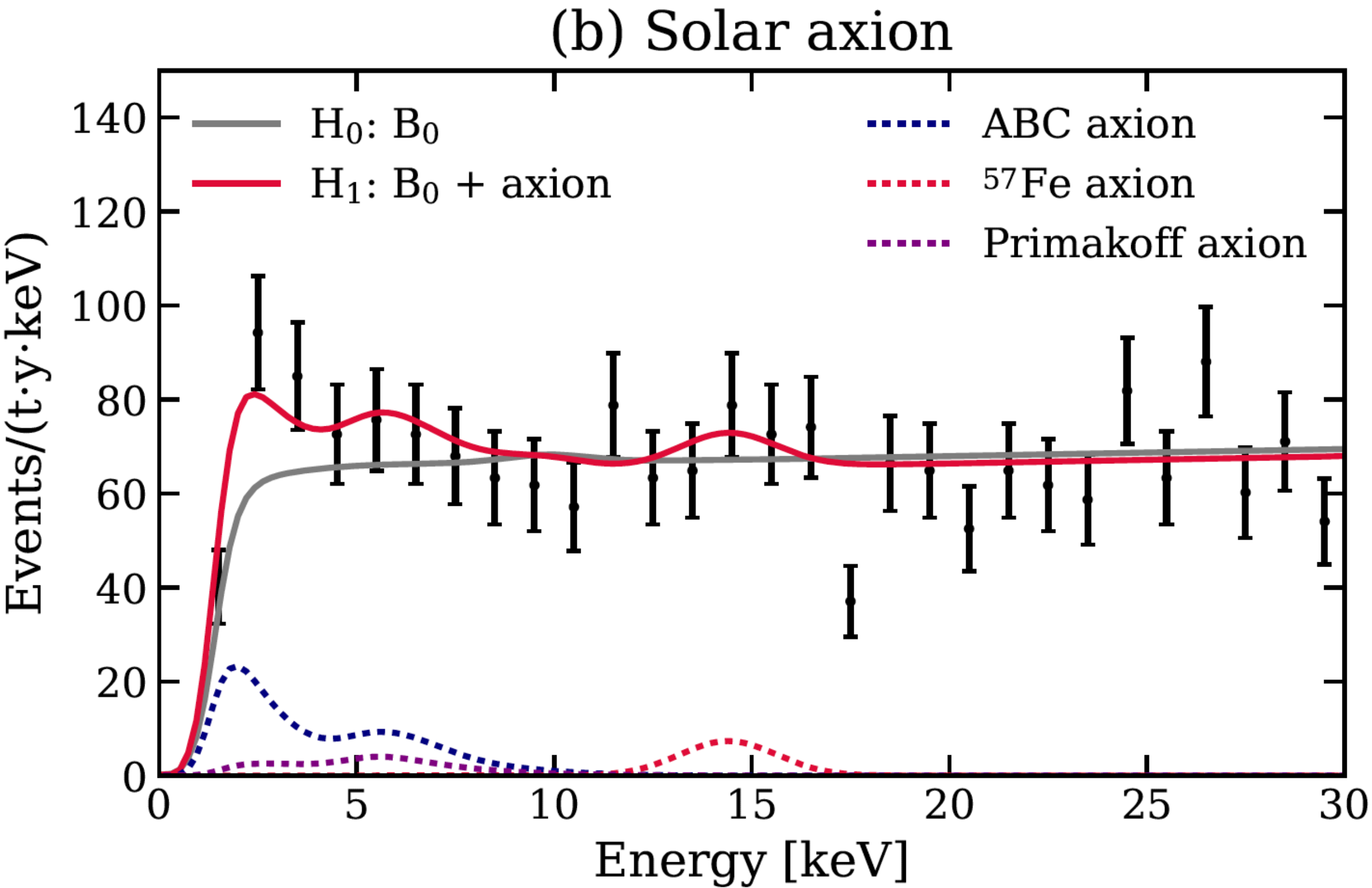




# Solar axions



Solar axions only:  $3.4\sigma$  over background  
Axions +  ${}^3\text{H}$ :  $2.0\sigma$  over background +  ${}^3\text{H}$



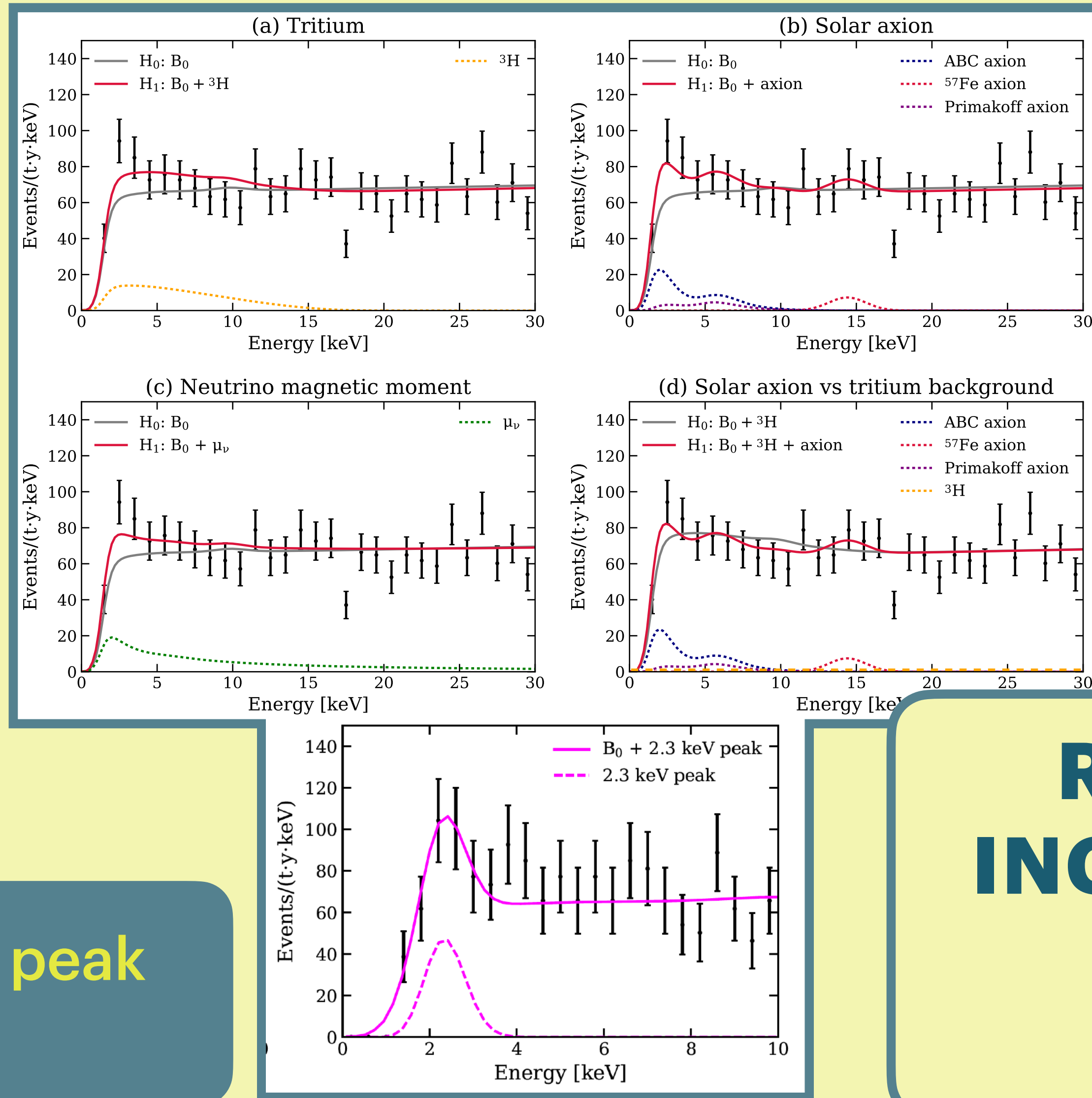
In tension with astrophysical constraints  
e.g. from stellar cooling  
(arXiv 1708.02111)



Tritium  
 $3.2\sigma$

$\mu_\nu$   
 $3.2\sigma$

Mono-energetic peak  
 $3.0\sigma$



Solar axions  
 $3.4\sigma$

Axions +  ${}^3\text{H}$   
 $2.0\sigma$

**RESULTS ARE  
INCONCLUSIVE...**

**...FOR NOW**



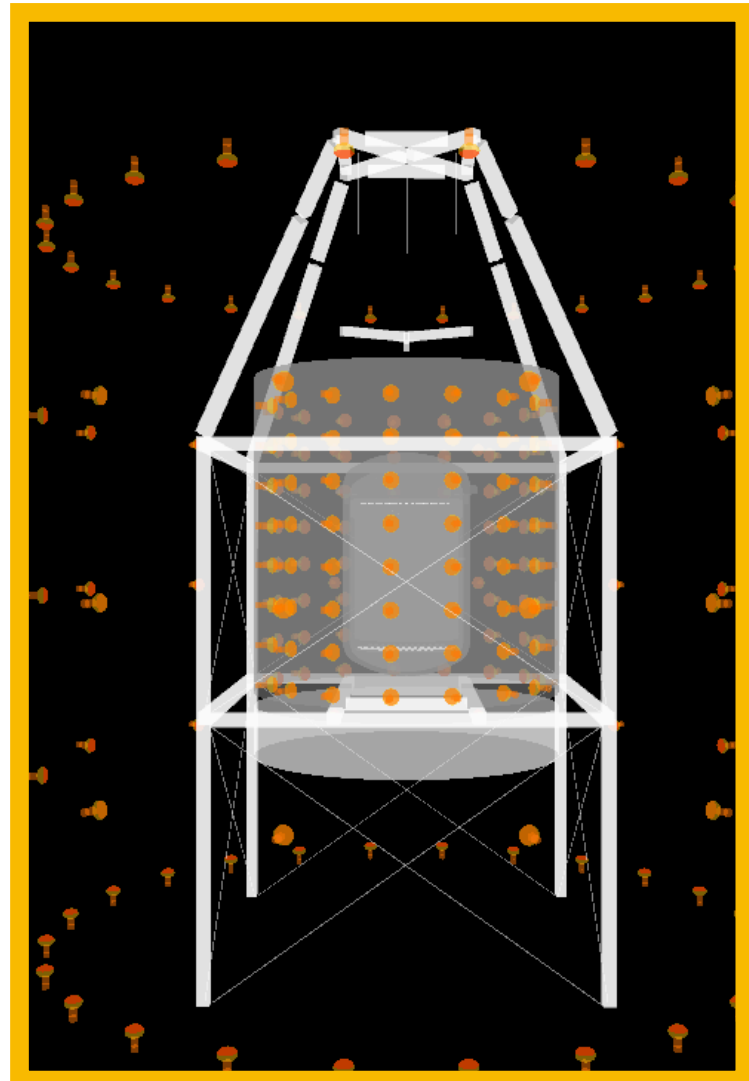
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# **XENON<sub>n</sub>T**

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# Some of what's new in XENONnT



## Neutron veto

- Separated region of existing muon veto
- 120 additional PMTs
- Gd in the water tank
- 0.5 %  $\text{Gd}_2(\text{SO}_4)_3$

⇒ Lower BG



## $^{222}\text{Rn}$ distillation

- Reduce Rn ( $^{214}\text{Pb}$ ) from pipes, cables, cryogenic system
- New

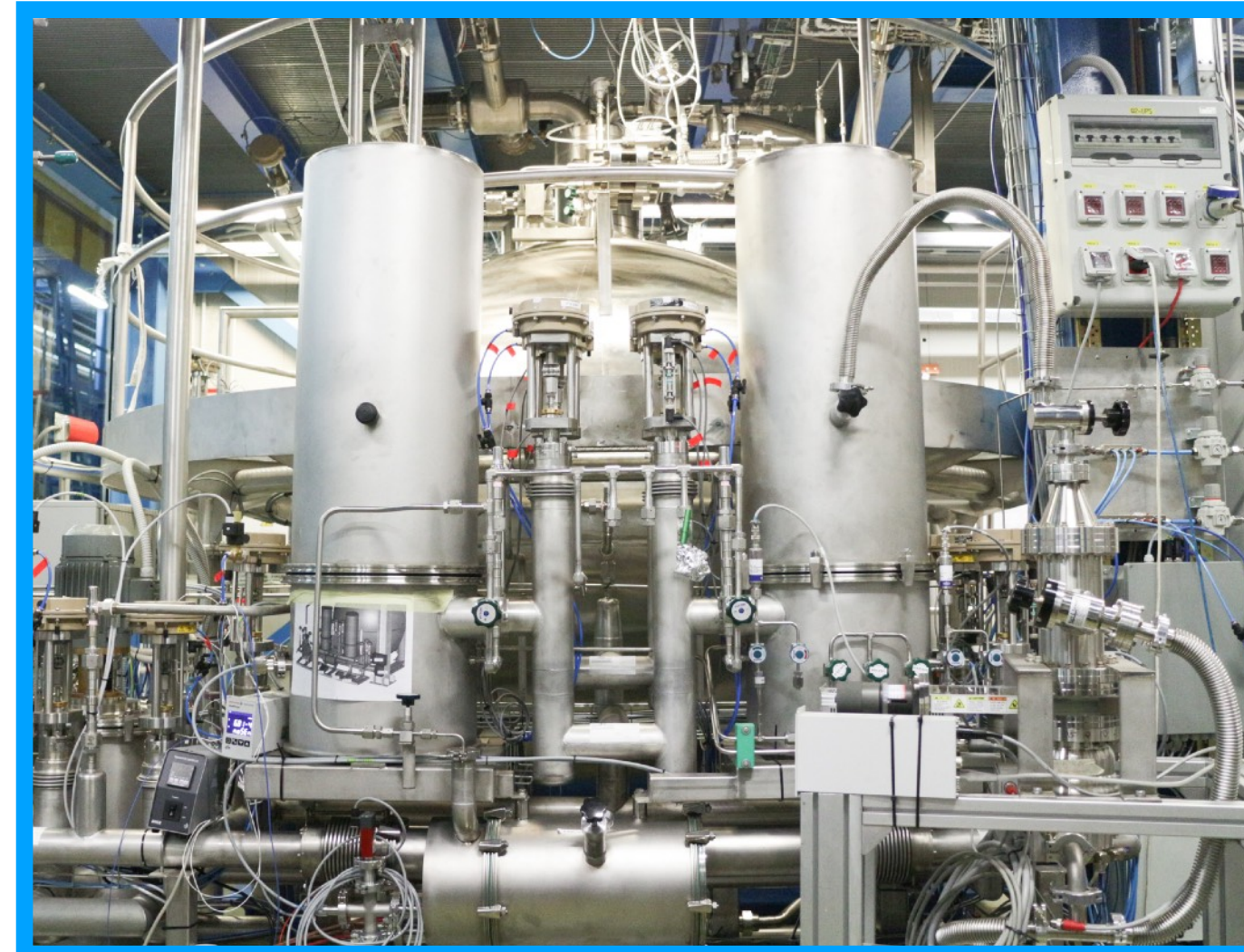
⇒ Lower BG



## Larger TPC

- Total 8.4 t LXe
- 5.9 t in TPC
- ~ 4 t fiducial
- 248 → 494 PMTs

⇒ Higher exposure  
and lower BG

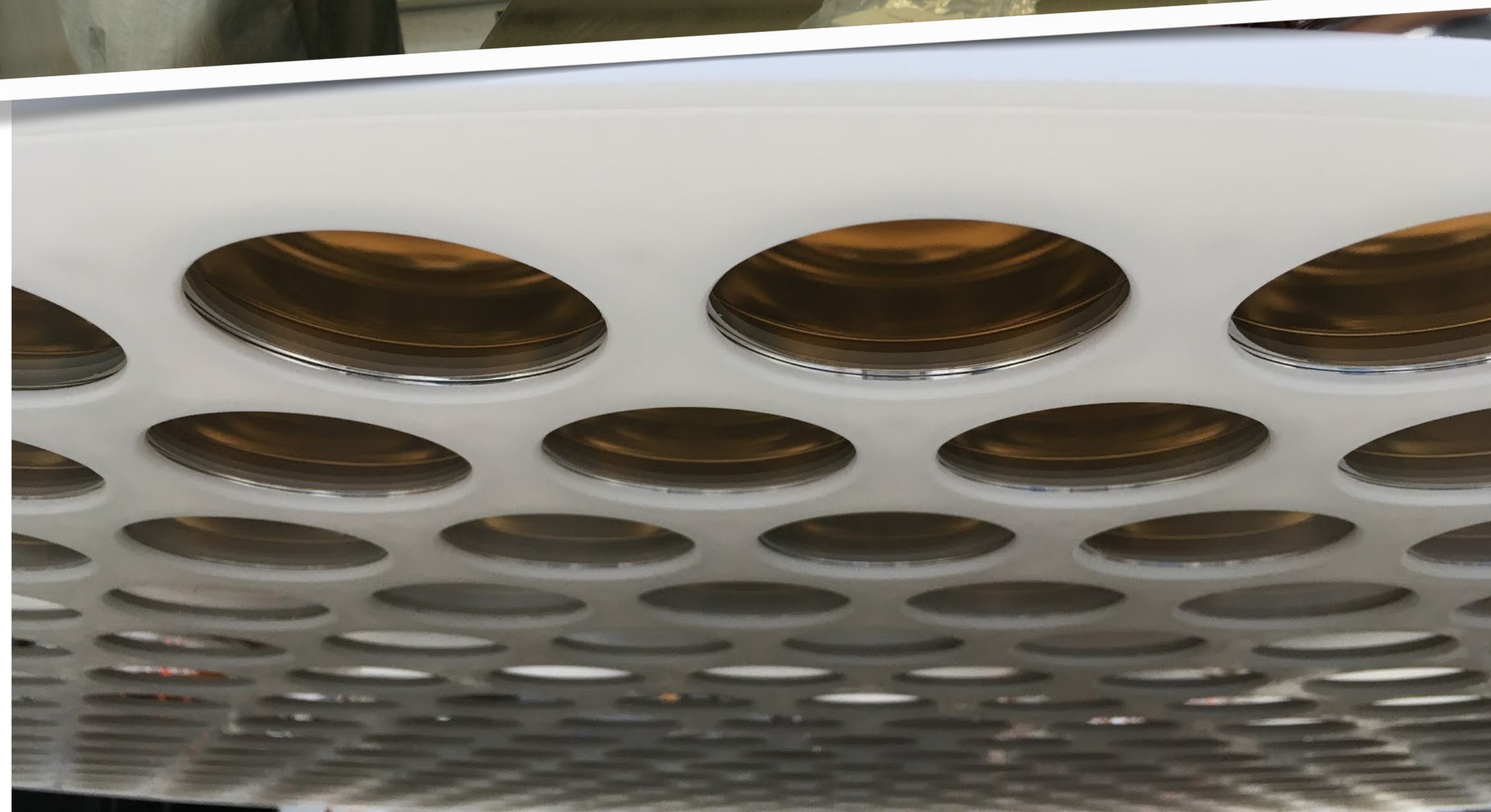


## Liquid Xe purification

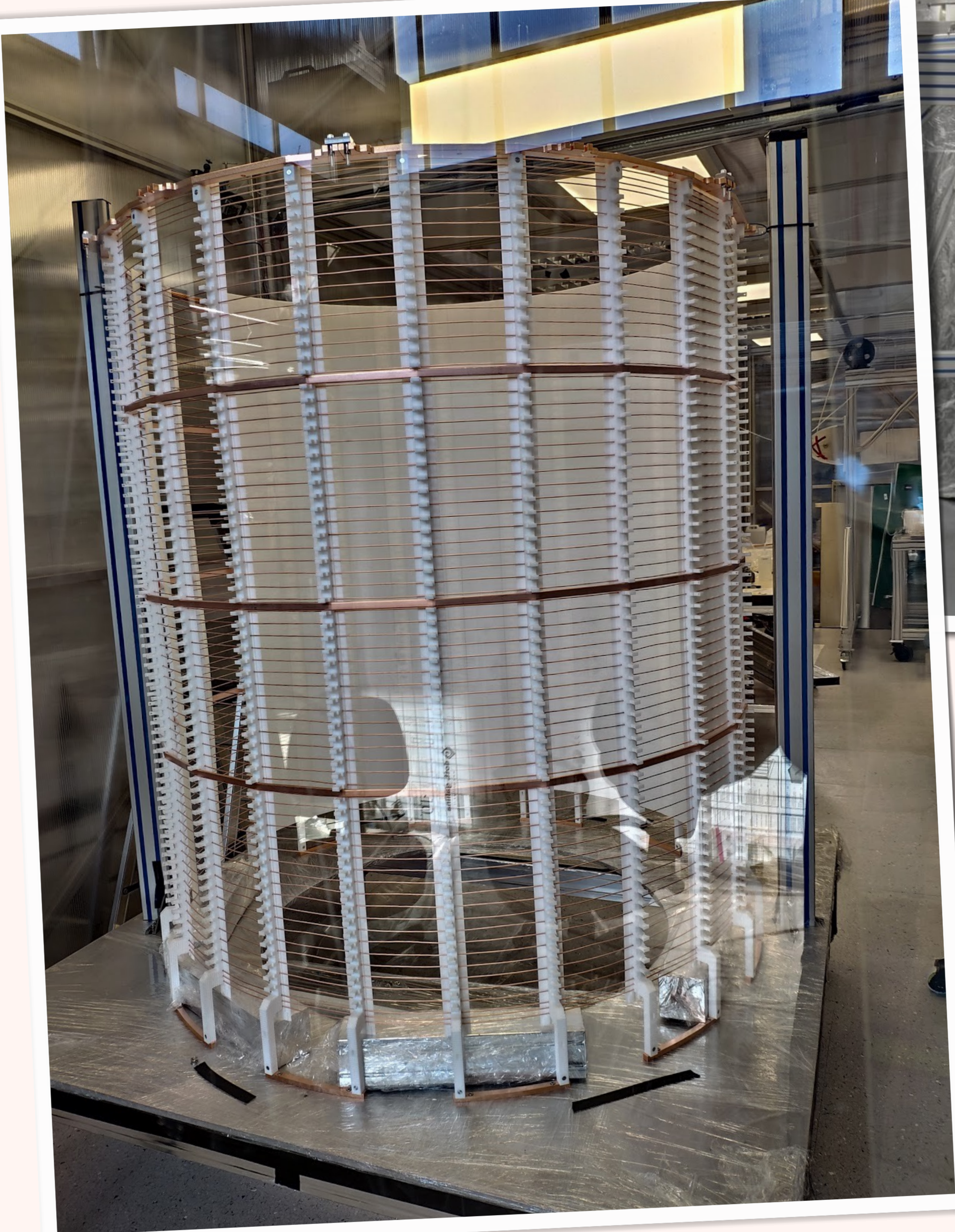
- Faster xenon cleaning
- 5 L/min LXe (2500 slpm)
- XENON1T ~ 100 slpm

⇒ Better resolution  
Lower thresholds

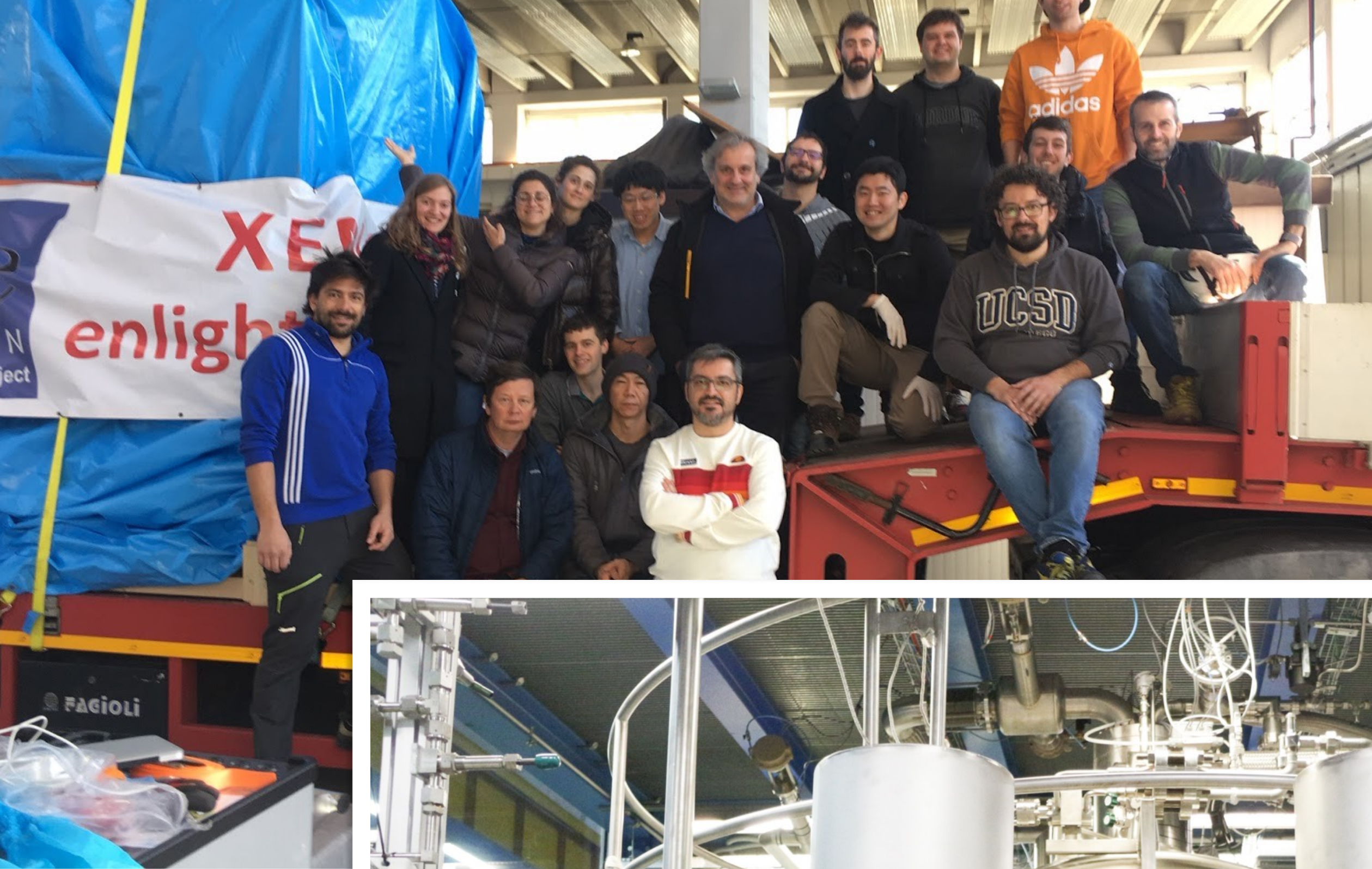






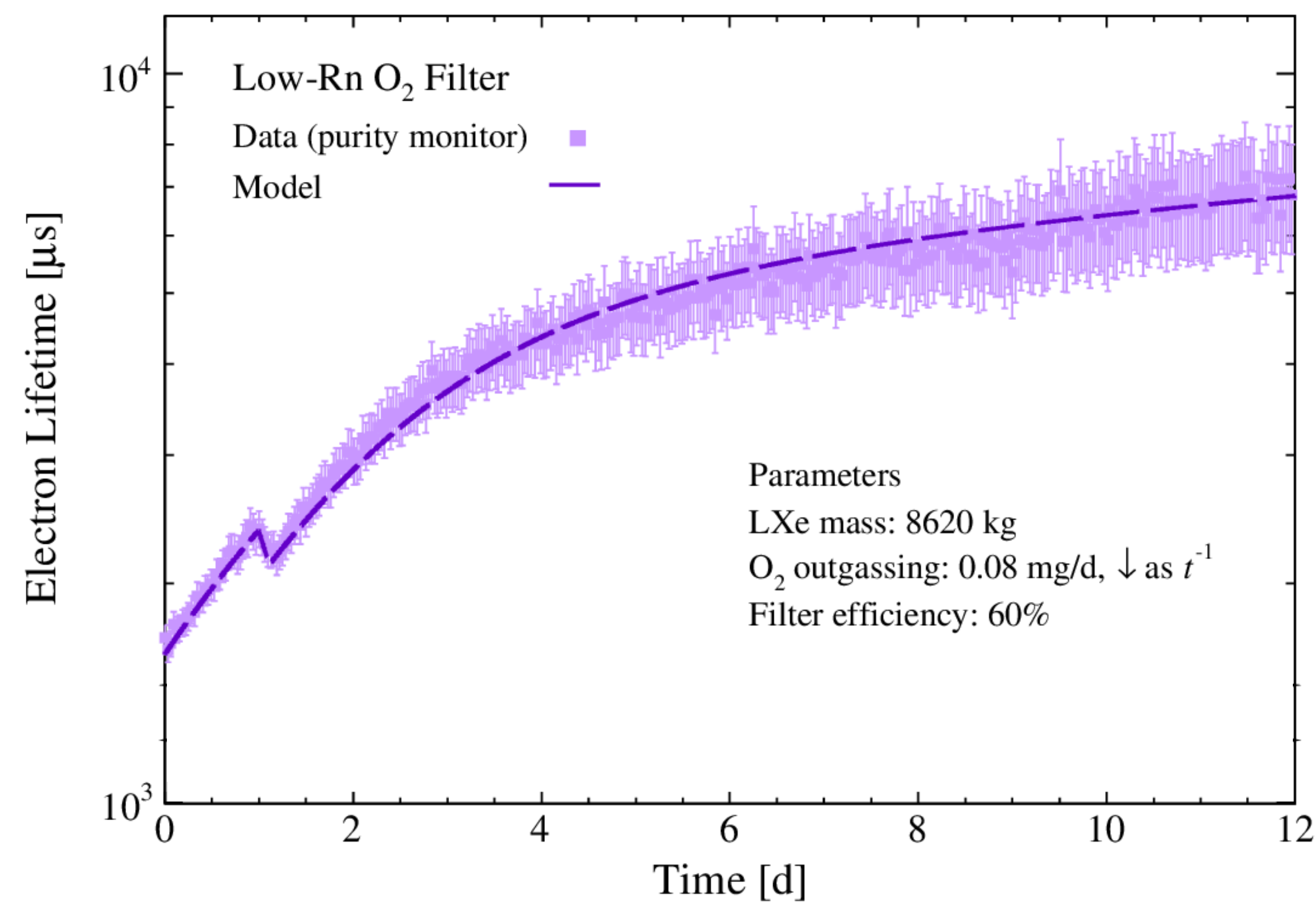




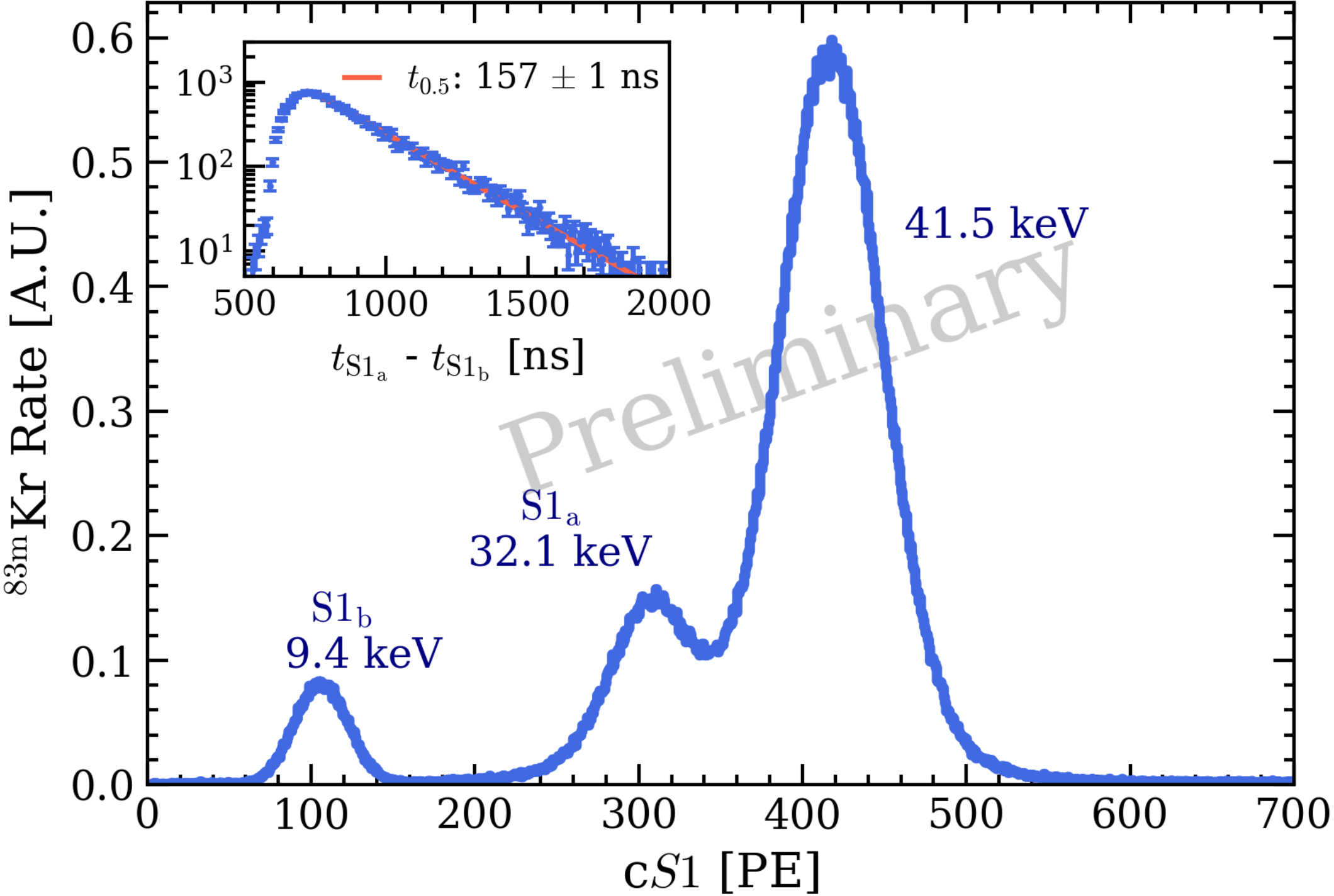




# XENOnT — watch this space



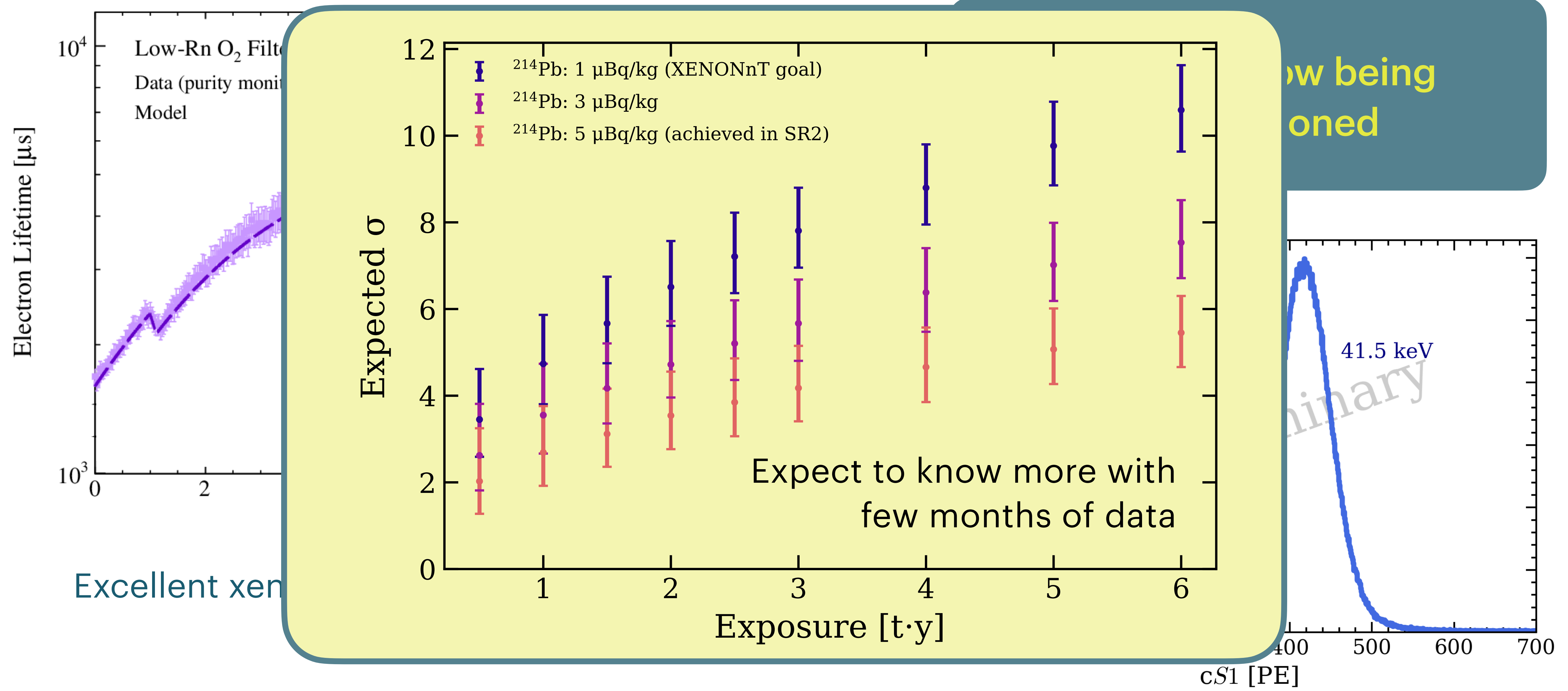
XENONnT now being commissioned



Excellent xenon purity achieved



# XENOnT — watch this space







**THANK YOU  
FOR LISTENING**

[xe-pr@lngs.infn.it](mailto:xe-pr@lngs.infn.it)

[www.xenonexperiment.org](http://www.xenonexperiment.org)

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