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MADMAX

Towards a Dielectric Axion Haloscope

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Universität Hamburg

On behalf of the MADMAX Collaboration

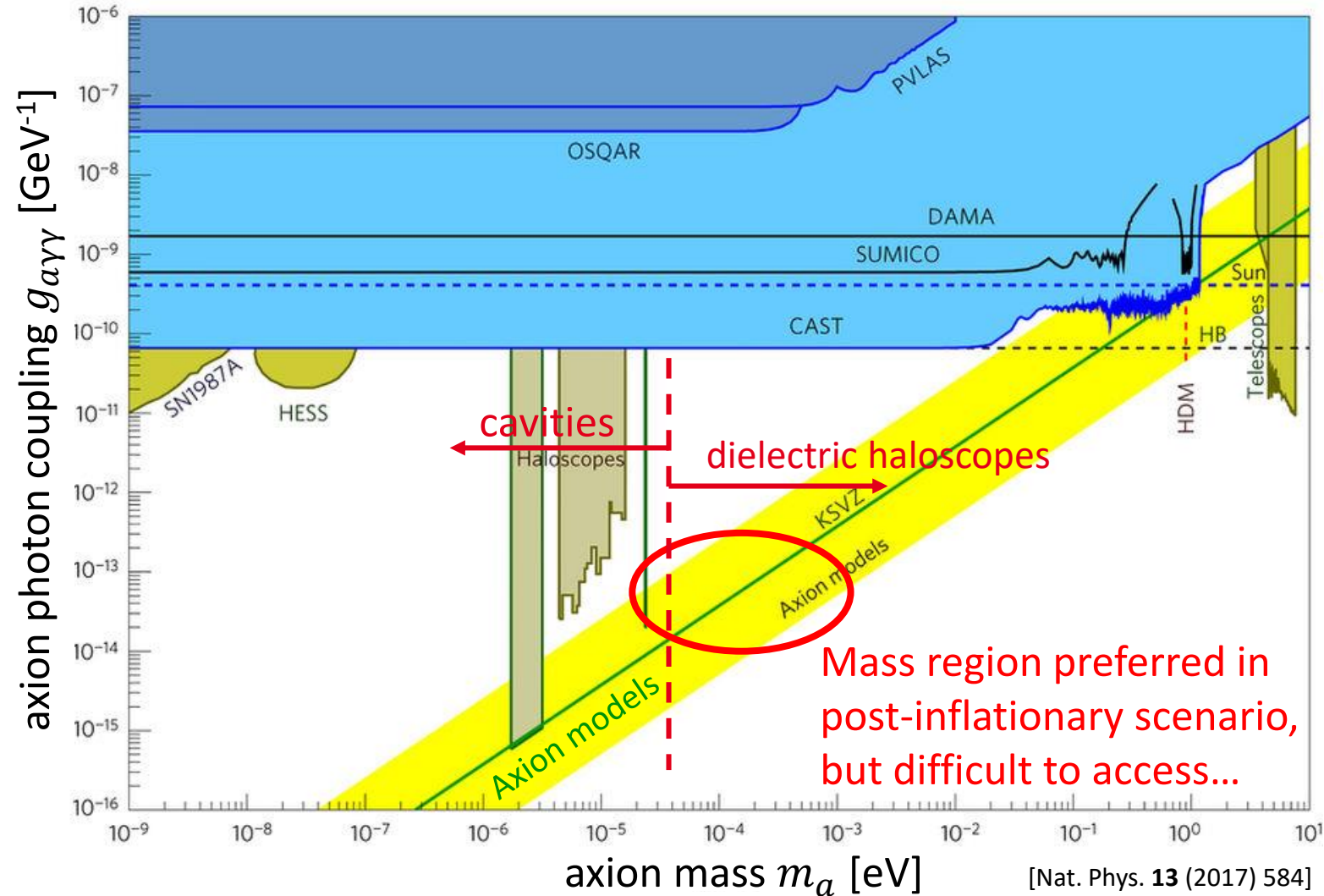
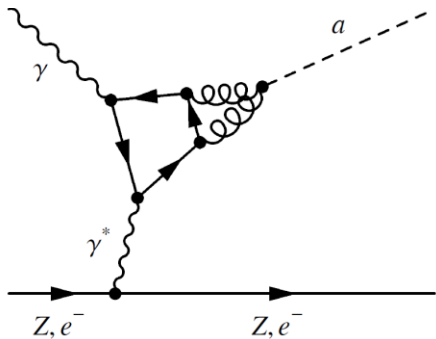
18th June 2021

16th Patras Workshop on Axions, WIMPs and WISPs

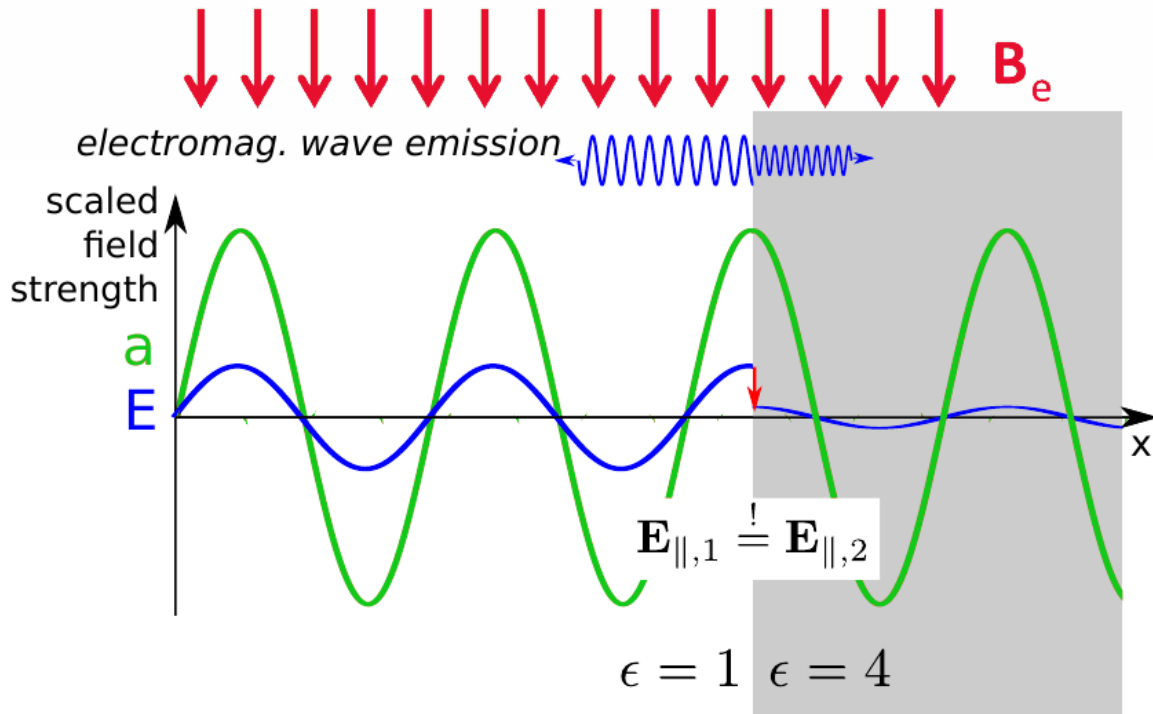
Trieste/Virtual

The Axion:

- Pseudo Nambu-Goldstone boson
- Small mass and small couplings
- Connected to solution of the strong CP problem
- Primakoff effect: Photon-Axion conversion in strong EM fields
- **Axion** can explain (part of) **Cold Dark Matter**



Dielectric Haloscope



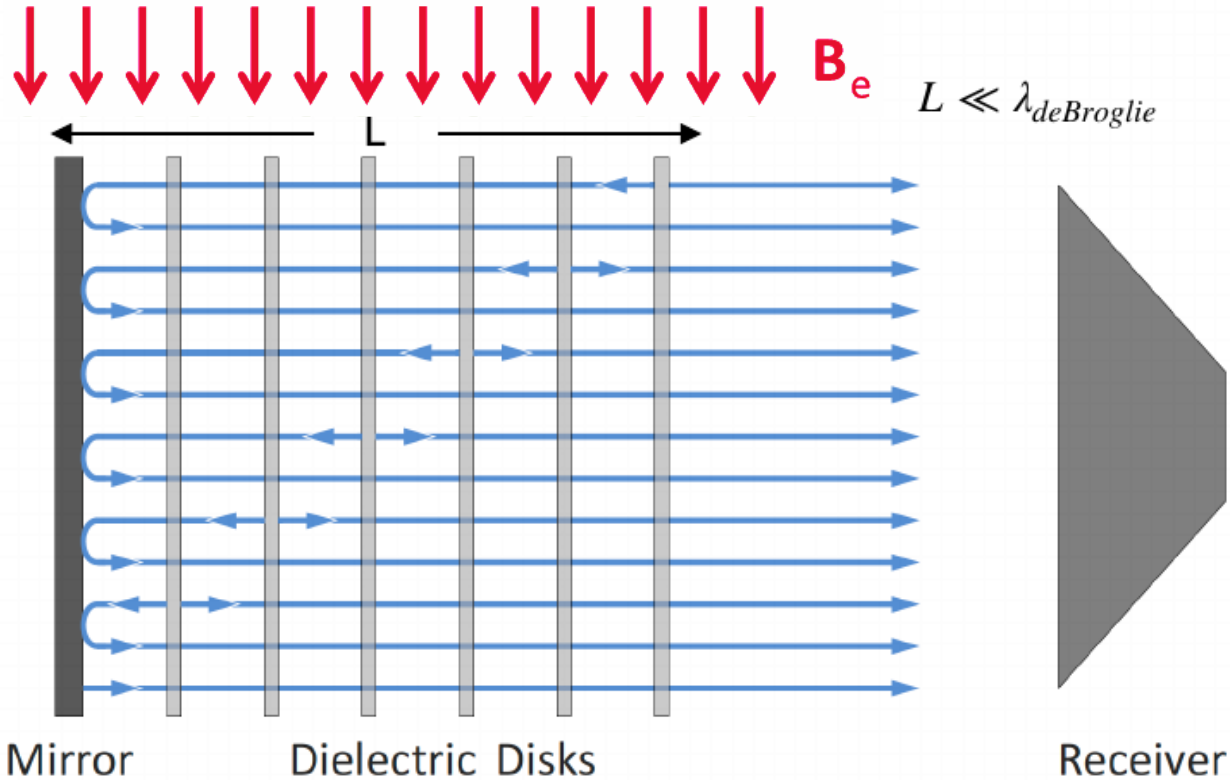
In an external magnetic field \mathbf{B}_e the axion field $a(t)$ sources an oscillating electric field \mathbf{E}_a

\mathbf{E}_a is different in materials with different ϵ

At the surface, E_{\parallel} must be continuous
→ Emission of electromagnetic waves

Power emitted from a single surface: $P/A = 2.2 \cdot 10^{-27} \frac{\text{W}}{\text{m}^2} C_{a\gamma} \left(\frac{B}{10 \text{ T}} \right)^2 \mathcal{O}(C_{a\gamma}) = 1$

Dielectric Haloscope



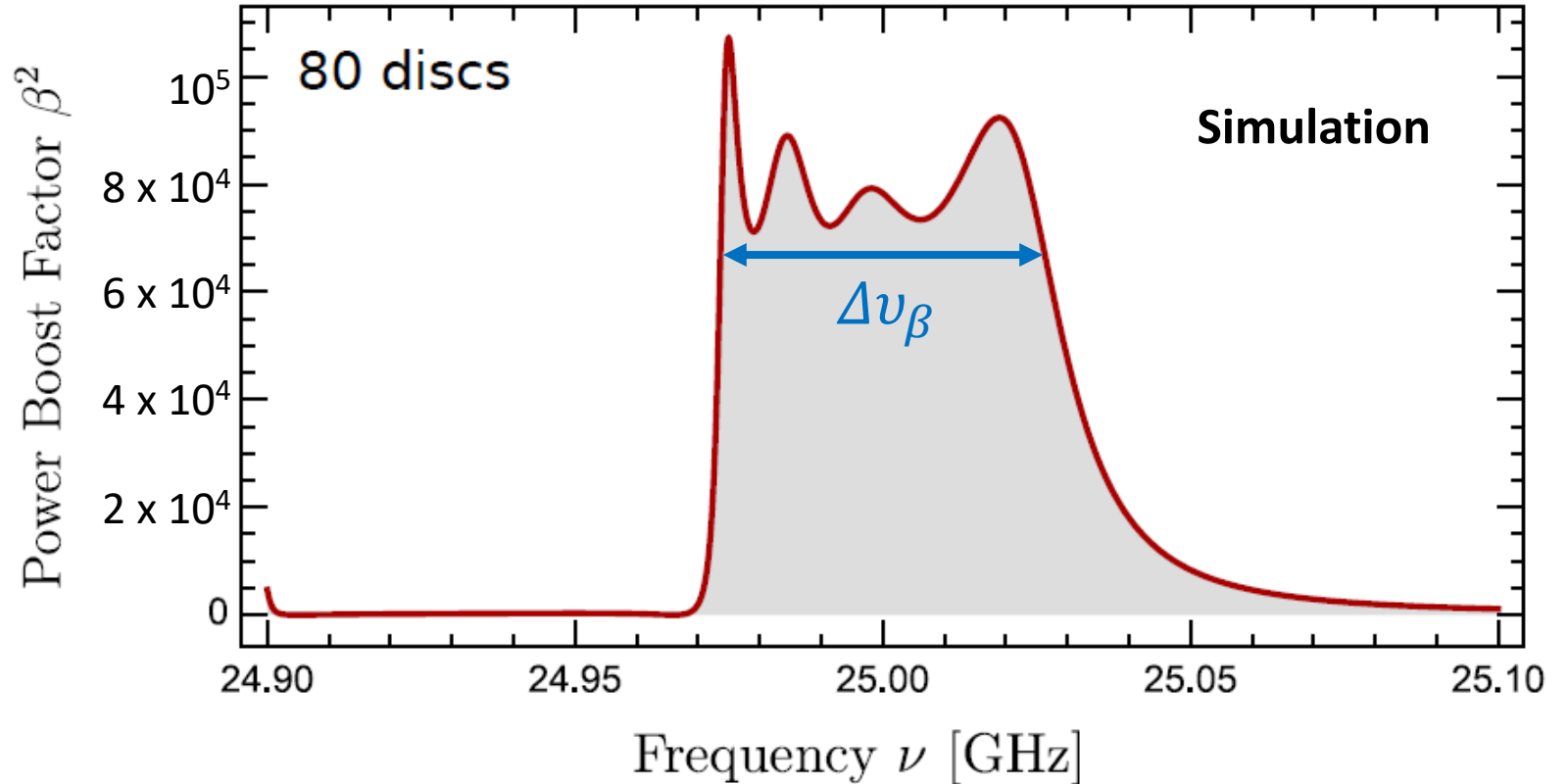
- Boost emitted power through:
- coherent emission from multiple interfaces
 - constructive interference effects

Power boost factor:

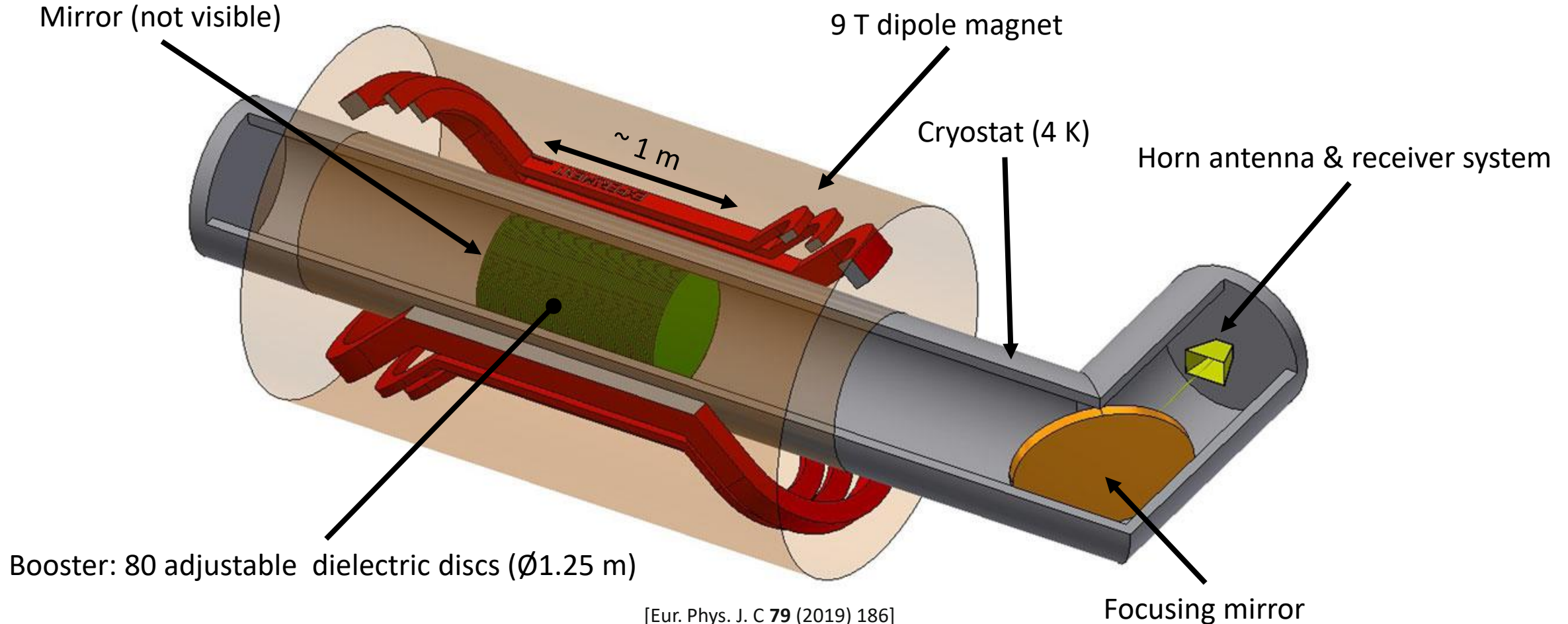
$$\beta^2 = P_{total} / P_{mirror}$$

Power emitted from all interfaces: $P/A = 2.2 \cdot 10^{-27} \frac{W}{m^2} C_{a\gamma} \left(\frac{B}{10 \text{ T}} \right)^2 |\beta^2|$

- $|\beta^2| > 10^4$ achievable with 80 discs and $\epsilon = 24$
- Non-uniform disk spacing of $\sim \lambda/2$ can achieve broadband response
- Tuning of sensitive frequency range by adjusting disc spacing
- Area law: $\beta^2 \Delta\nu_\beta \sim \text{const.}$



MAGnetized DISC and MIRROR AXION eXperiment





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The MADMAX Collaboration



Founded 18.10.2017
@ DESY/UHH



EBERHARD KARLS
UNIVERSITÄT
TÜBINGEN



Universidad
de Zaragoza

RWTH AACHEN
UNIVERSITY



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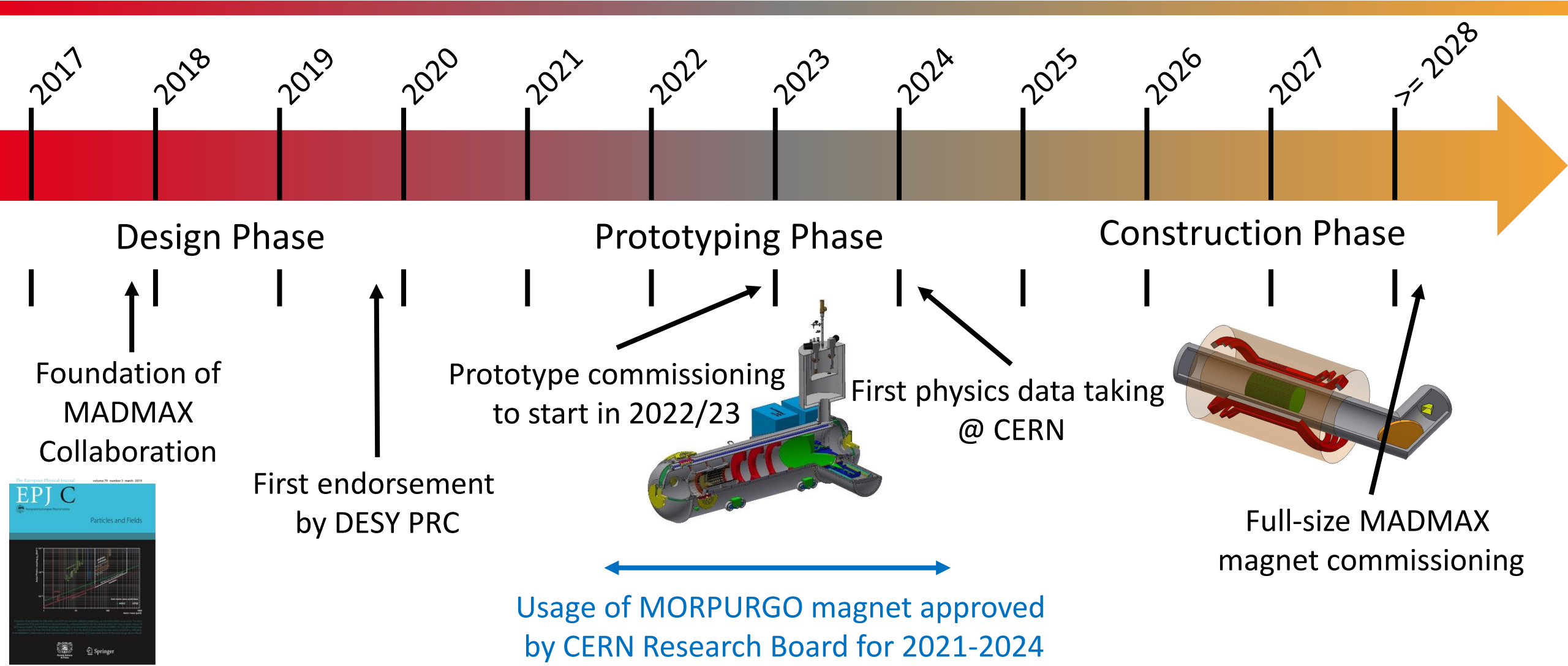
MAX-PLANCK-INSTITUT
FÜR PHYSIK



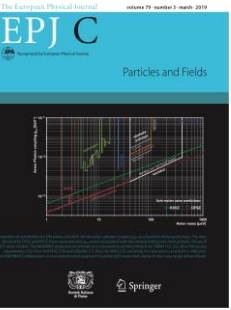
Max-Planck-Institut
für Radioastronomie



Time Scale



Foundation of MADMAX Collaboration



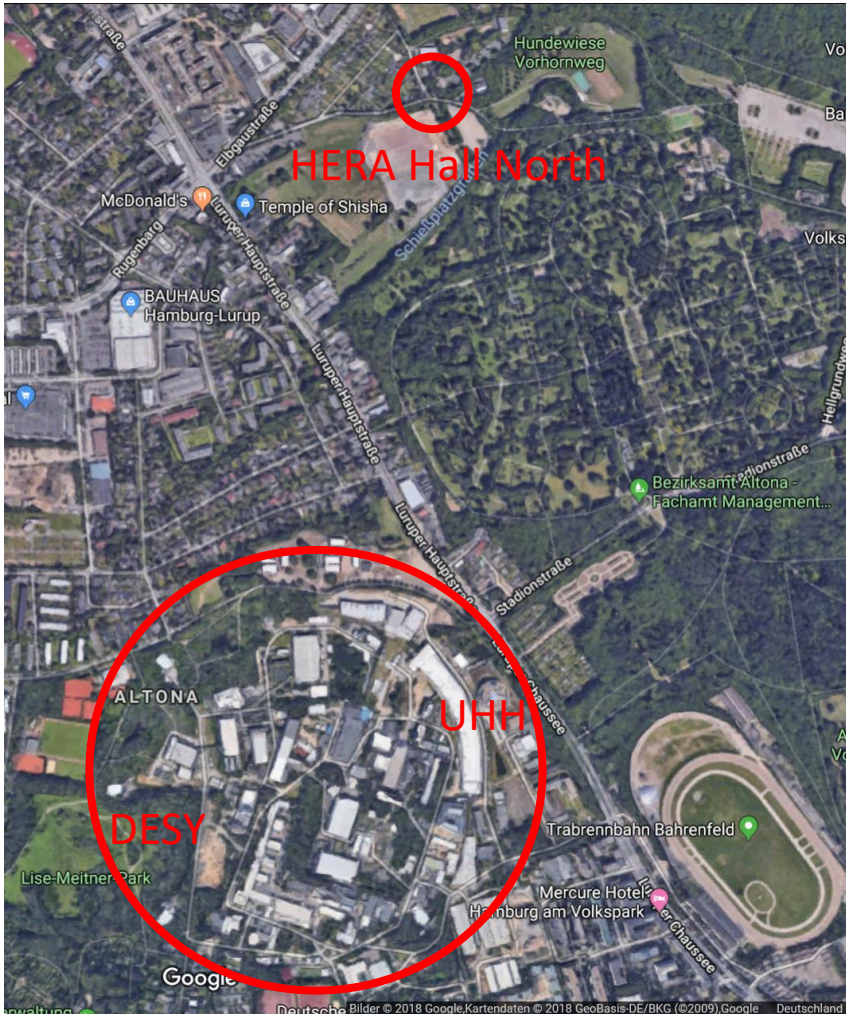
First endorsement by DESY PRC

Prototype commissioning to start in 2022/23

First physics data taking @ CERN

Full-size MADMAX magnet commissioning

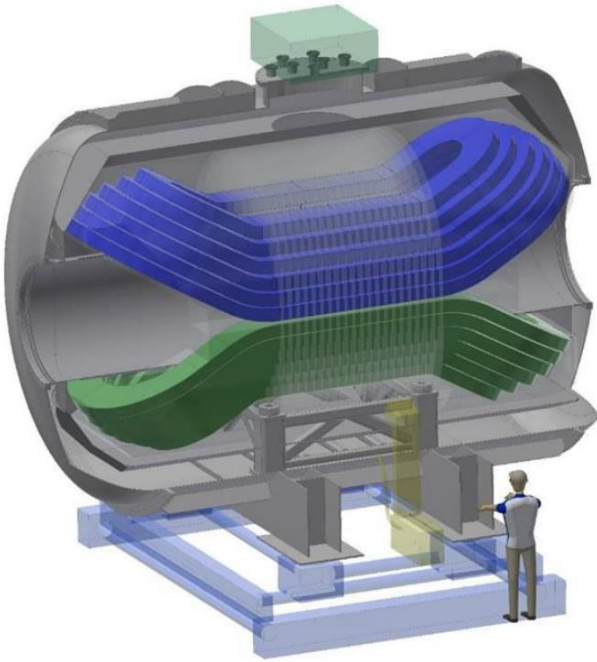
Usage of MORPURGO magnet approved by CERN Research Board for 2021-2024



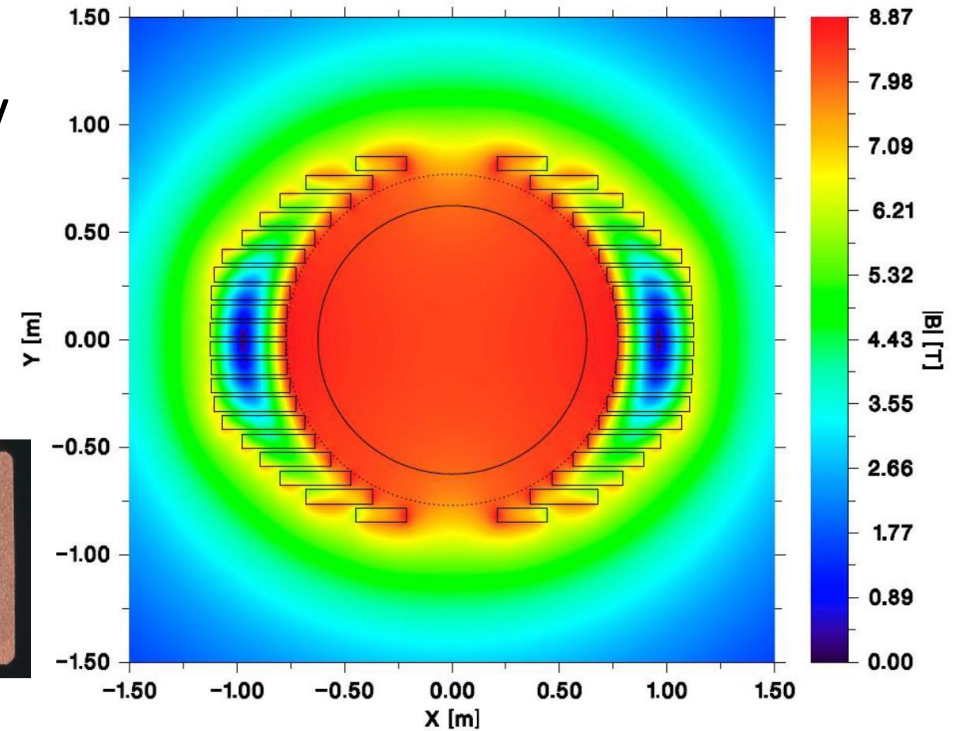
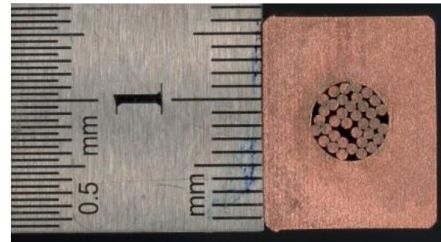
- MADMAX to be built at HERA Hall North
- Make use of DESY infrastructure
- Benefit: re-use H1 yoke as magnetic shielding to reduce fringe field



The Magnet



- Innovation Partnership with Bilfinger Noell and CEA Saclay
- Block Design with CICC superconducting cable: NbTi with Cu jacket

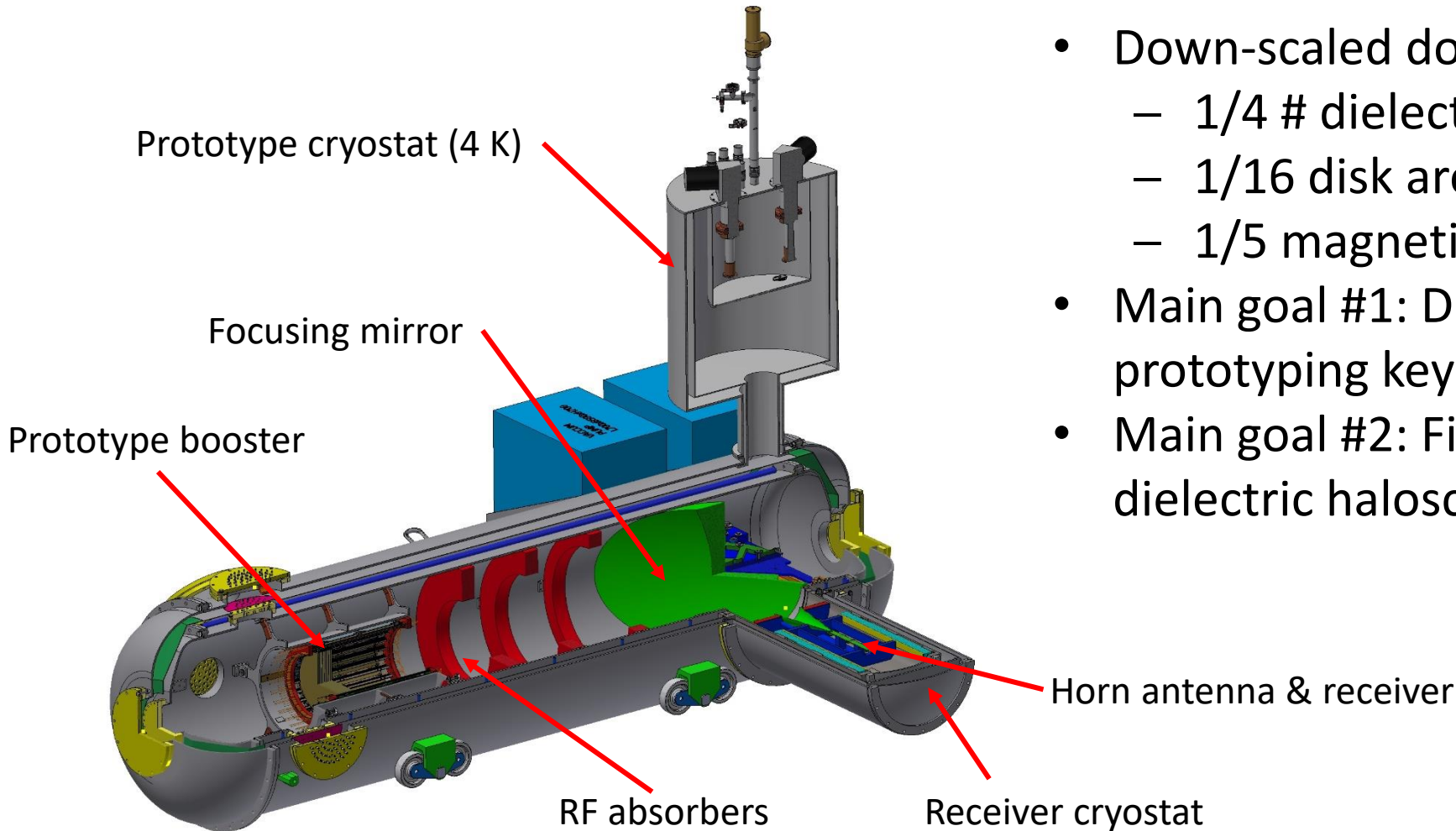


- Magnet design and construction drives the time scale of the project
- **Peak field 9 T**, homogeneity better than 20%
- Warm bore: \varnothing **1.35 m**

First of a kind!

$$\text{FoM} = B^2 A = 100 \text{ T}^2 \text{ m}^2$$

The MADMAX Prototype



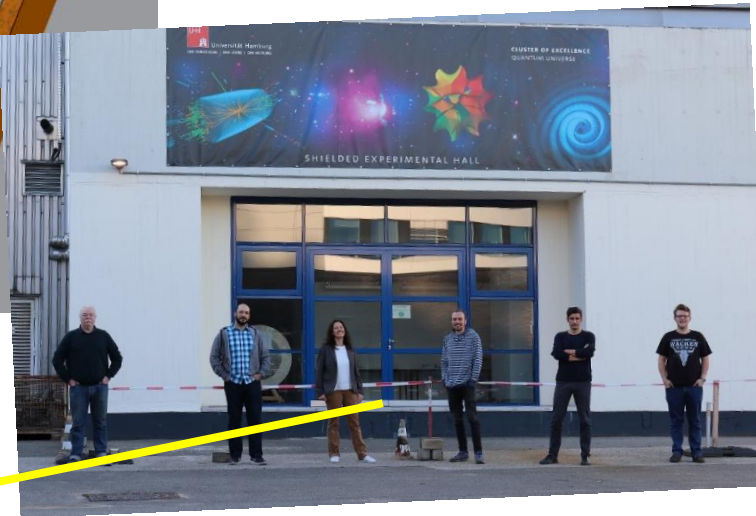
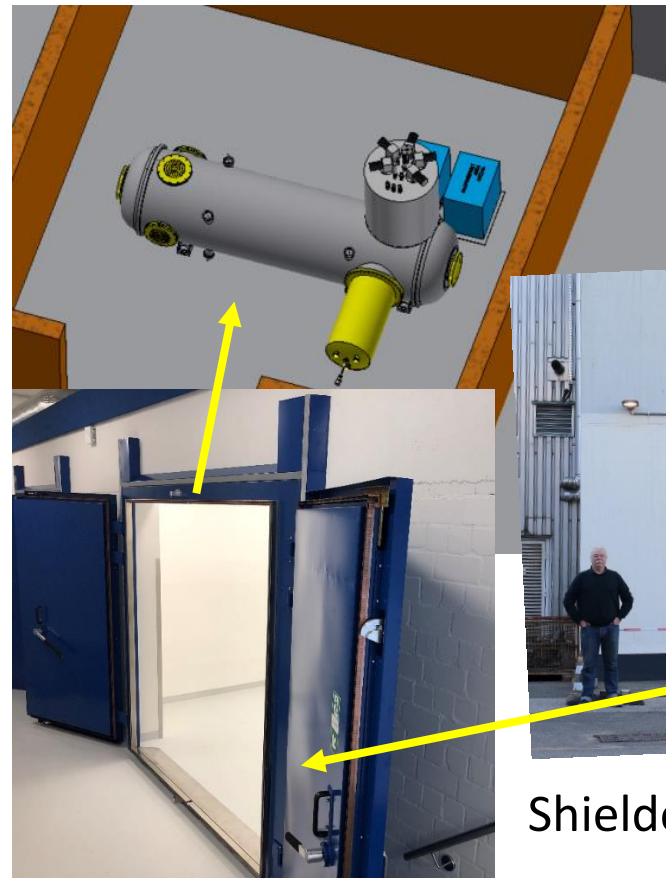
- Down-scaled down version of MADMAX:
 - 1/4 # dielectric disks
 - 1/16 disk area
 - 1/5 magnetic field
- Main goal #1: Demonstrating and prototyping key technologies for MADMAX
- Main goal #2: First axion search with a dielectric haloscope

The MADMAX Prototype

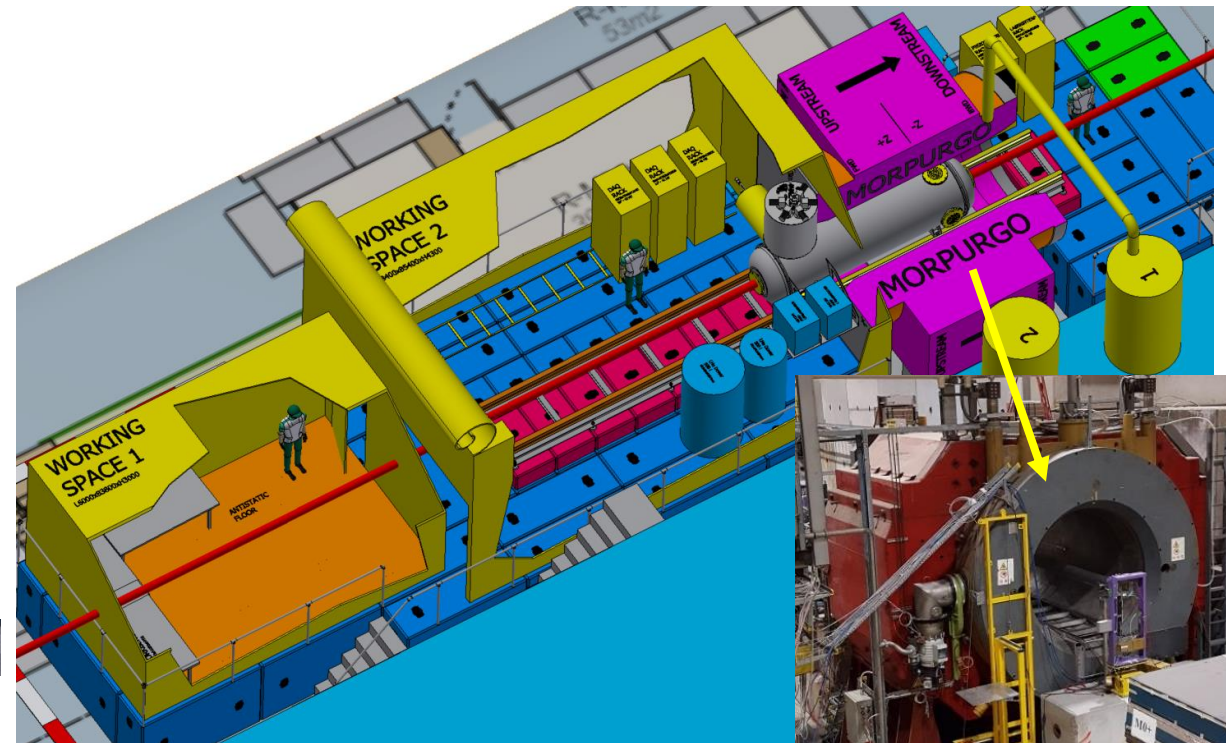


- Gain experience in operation and commissioning in SHELL @ UHH
- First physics measurements in MORPURGO magnet @ CERN

CLUSTER OF EXCELLENCE
 QUANTUM UNIVERSE

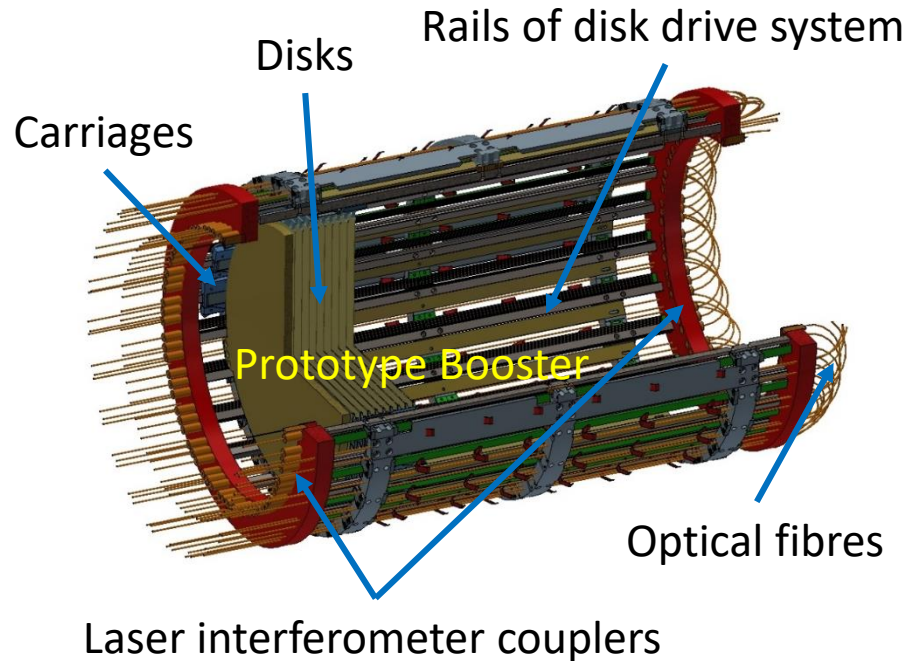


Shielded RF lab SHELL in Hamburg

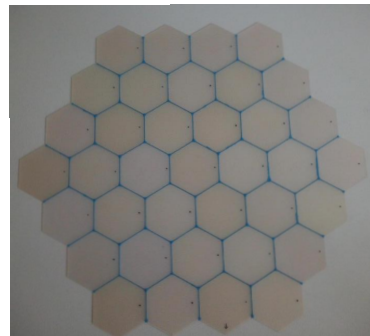
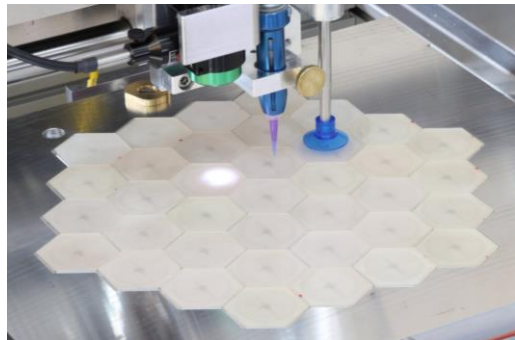


MORPURGO magnet @ CERN



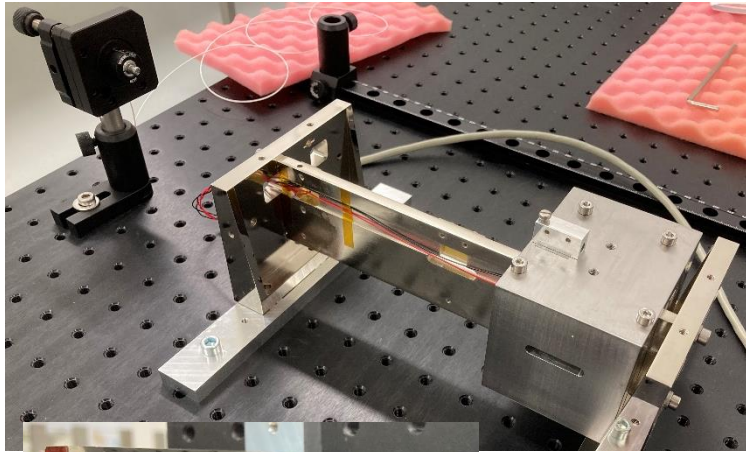


- Booster is the heart of MADMAX
- Need to manipulate many large area disks with precision $< 10 \mu\text{m}$
- Operating conditions:
 - Cryogenic temperatures: 4 K
 - High magnetic field: up to $\sim 10 \text{ T}$
 - Vacuum
- Long travel range
- Disk weight: 600 g for $\varnothing 300 \text{ mm}$
- **Piezo-driven actuator system with feedback from laser interferometer with absolute precision**



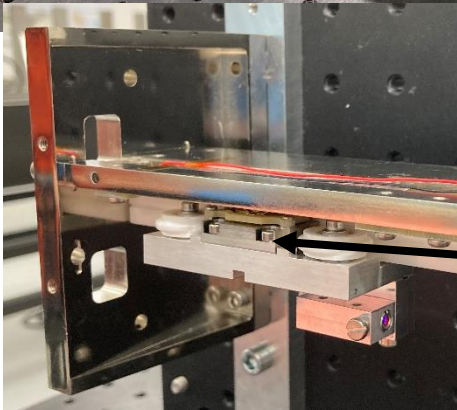
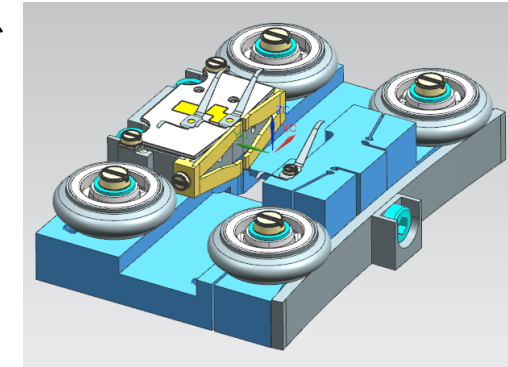
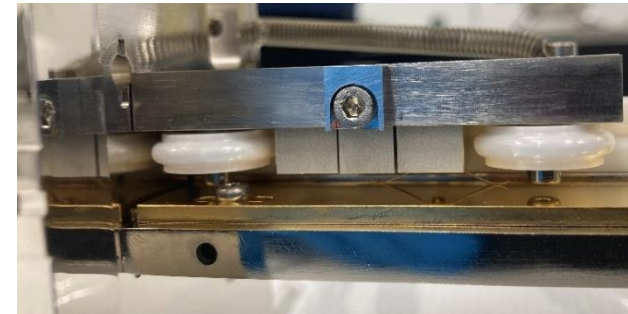
- Candidate disk materials:
 - **LaAlO_3** ($\epsilon \approx 24$, $\tan\delta \approx \text{a few } 10^{-5}$)
 - Sapphire ($\epsilon \approx 9$, $\tan\delta \approx 10^{-5}$)
- LaAlO_3 available as 3" wafers at maximum
- **Tiling necessary \rightarrow Semi-automatic gluing machine**

Prototype Disk Drive



- Carriage on rail driven by piezo actuator (stick-slip principle)
- Successfully tested by manufacturer @ 12 K

First disk drive unit developed by



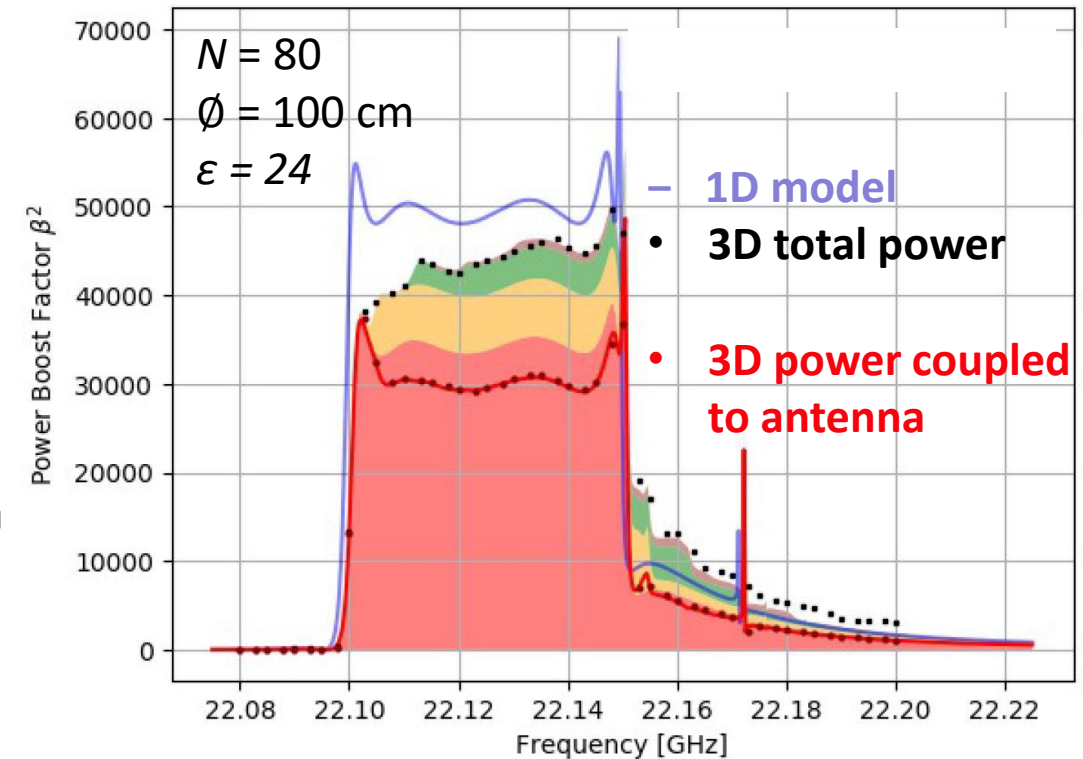
Disk drive movement @ 12 K



- Room temperature tests confirmed manufacturer results
- Cryogenic tests at MPP Munich currently in progress

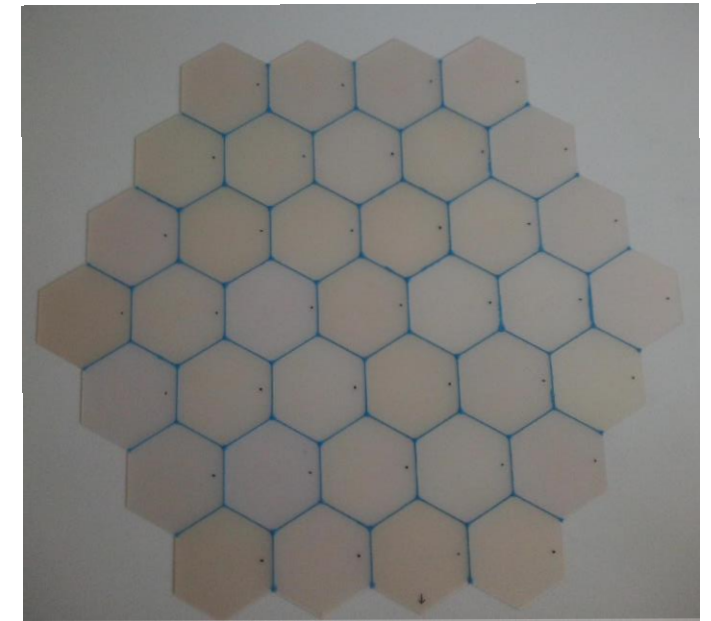
Detector feasibility study and design optimization using simulation of achievable boost factor

- **Geometry factor and coupling to antenna (beam shape)** → **~70 % compared to 1D**
- Dielectric losses → small losses for $\tan\delta < 10^{-4}$
- Inaccuracy (position, roughness, tilt, thickness,...) → positioning precision $< 10 \mu\text{m}$
roughness $< 10 \mu\text{m}$
tilt $< 0.1 \text{ mrad}$
thickness measured to $\pm 5 \mu\text{m}$
- DM velocity dispersion → no significant loss if $v < 10^{-2} c$
- Tiling of disks

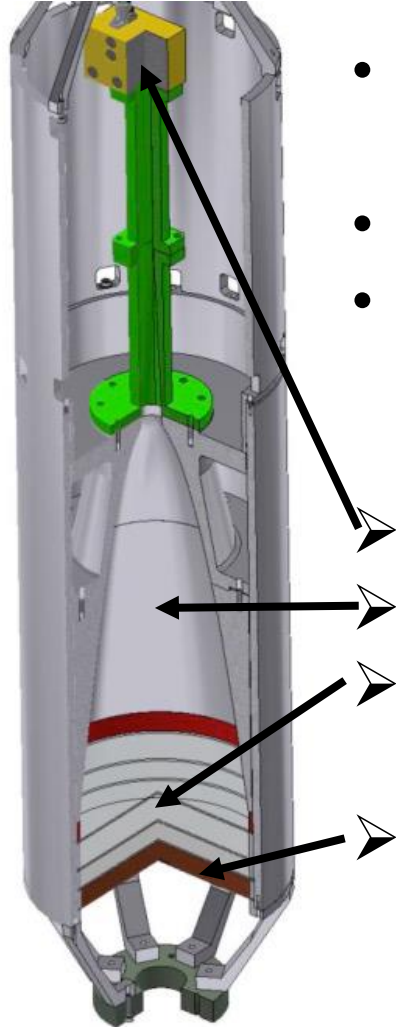


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- DM velocity dispersion → no significant loss if $v < 10^{-2} c$
- **Tiling of discs** → **possibly large impact**
(very preliminary)



Closed Booster System



- „Simpler“ closed system to understand behaviour and simulations
- Can be operated at cryogenic temperatures
- Planned operation at MORPURGO at CERN to understand influence of magnetic field on RF measurements

Receiver

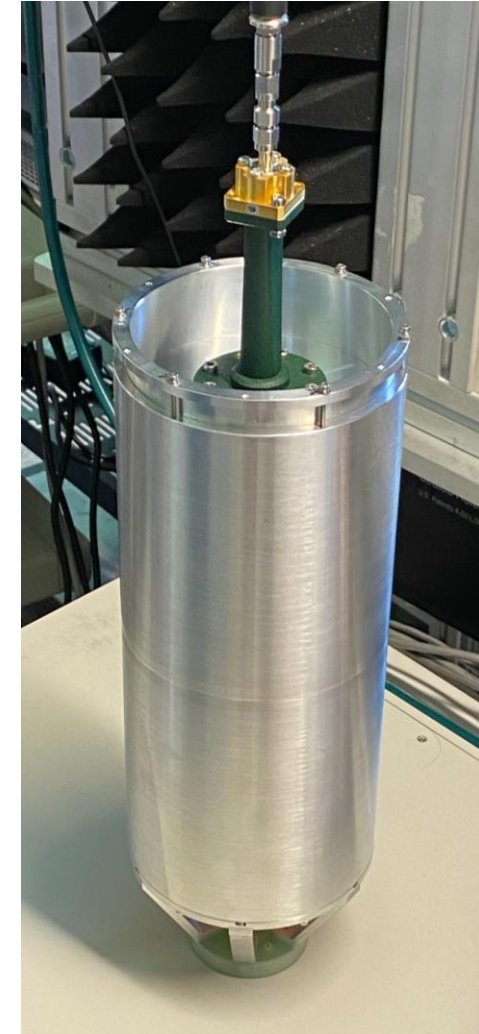
Parabolic taper

3x Ø100 mm sapphire disks
 (fixed distances)

Copper mirror

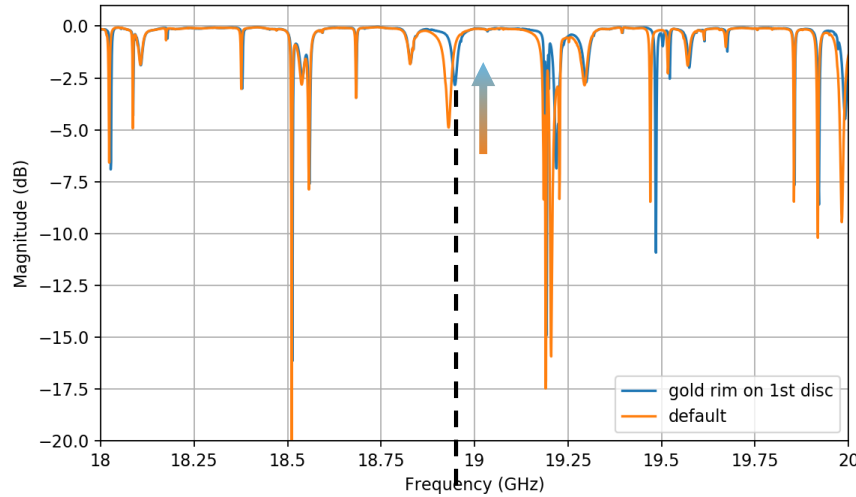


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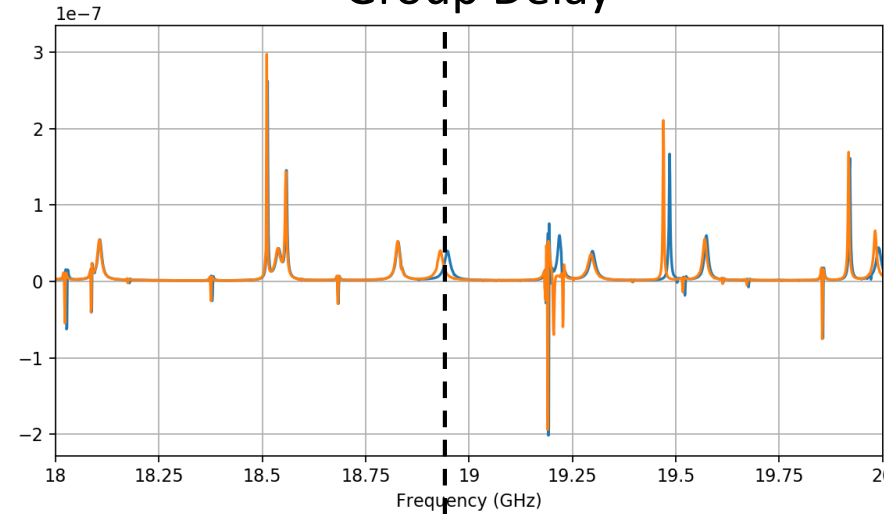


Data

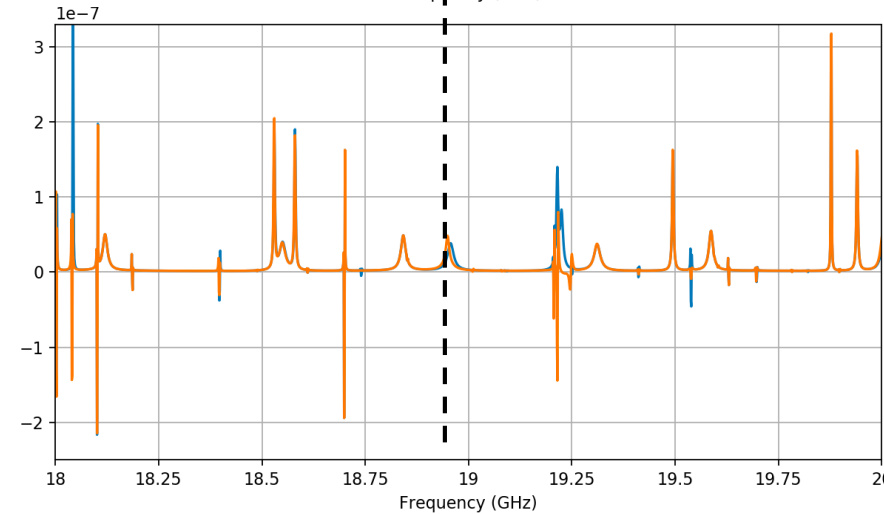
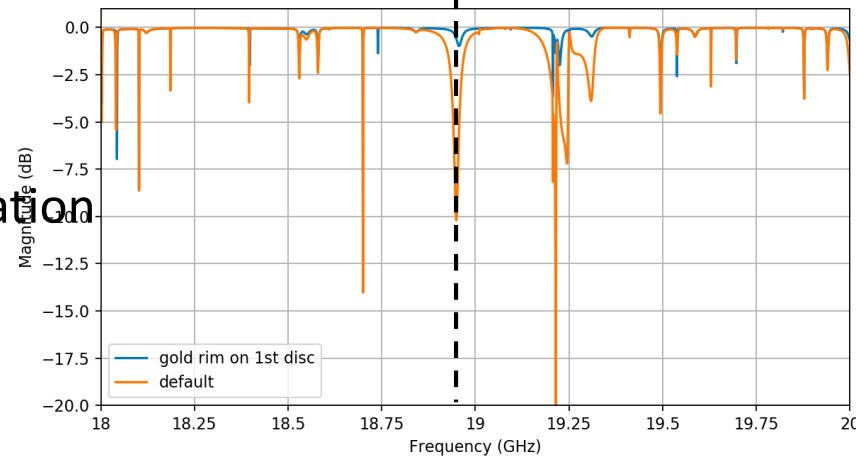
Reflectivity



Group Delay



Simulation



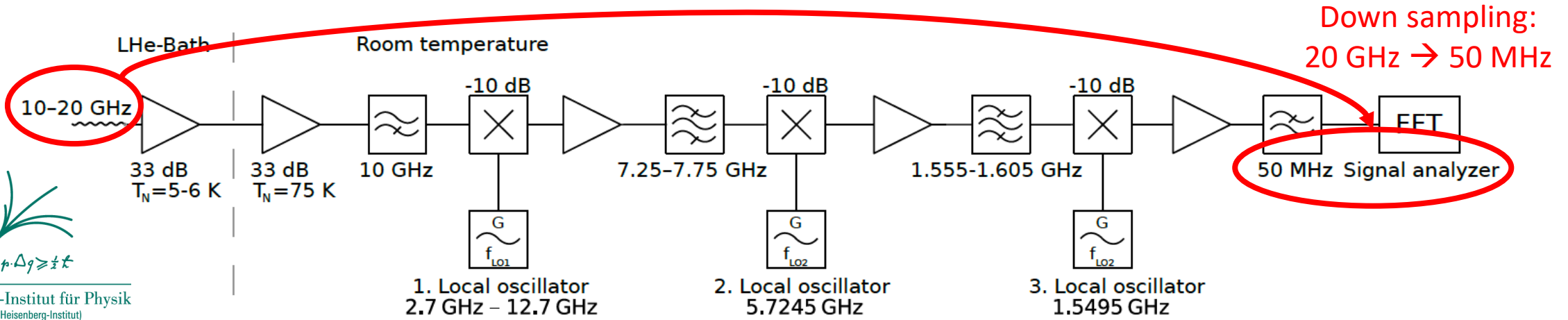
Hunting losses:

- Sapphire disks
- Disk with gold coated rim

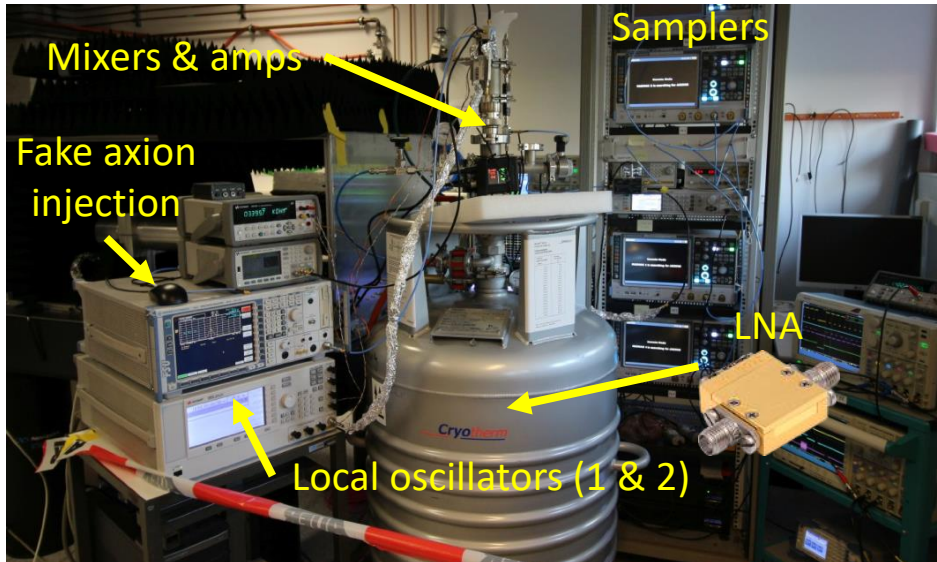


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Receiver Chain



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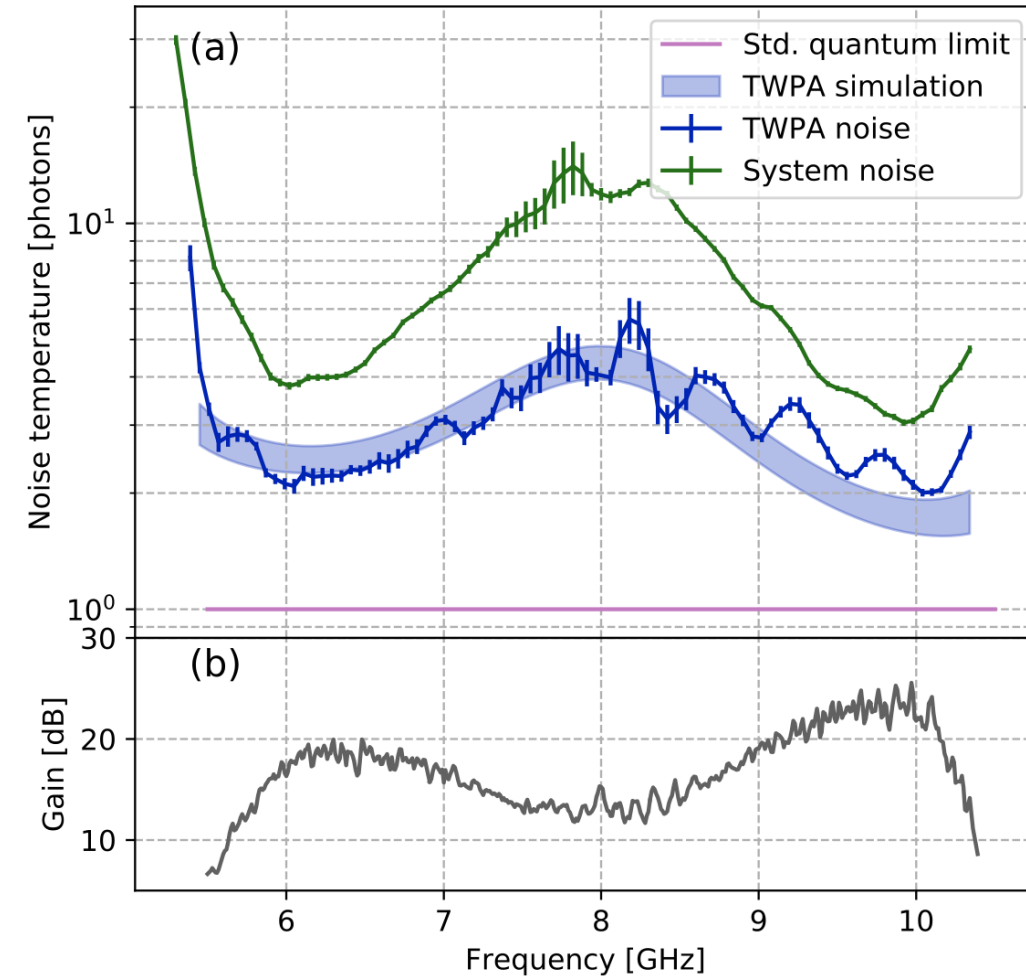
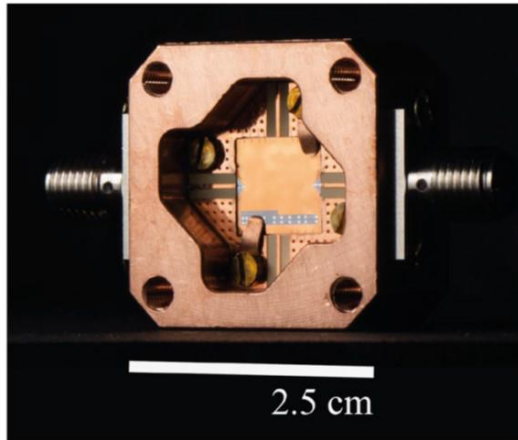
- Receiver chain with low-noise amplifier and three mixing stages
- Amplifiers for high frequencies (> 40 GHz) still have to be developed, e.g. TWPAs

Test setup at MPP with 4 samplers and fake axion injection:
Detection of 1.2×10^{-22} W signal within few days

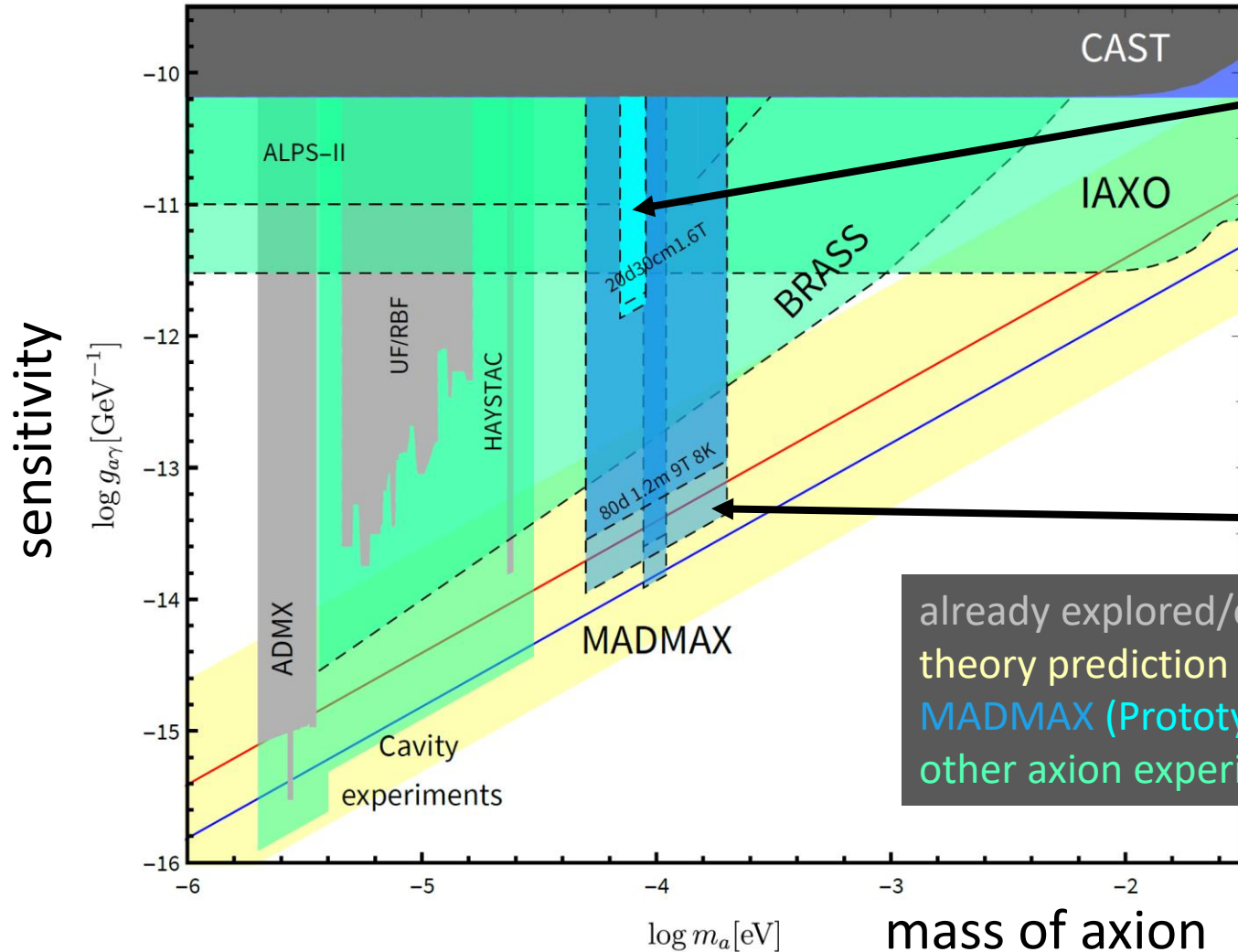


Low-noise cryogenic amplifier (noise temperature 5 to 6 K)

- Traveling wave parametric amplifier (TWPA)
- **First 10 GHz TWPA** produced (PRX 10, 021021)
- 1K noise temp, 20 dB gain @ 10 GHz
- Future development to 30 GHz



[Reversed Kerr TWPA arXiv:2101.05815]



MADMAX Prototype:

$$N_{\text{disk}} = 20$$

$$A_{\text{disk}} = 0.07 \text{ m}^2$$

$$B_{\parallel} = 1.6 \text{ T}$$

$$T_{\text{sys}} = 8 \text{ K}$$

MADMAX:

$$N_{\text{disk}} = 80$$

$$A_{\text{disk}} = 1.2 \text{ m}^2$$

$$B_{\parallel} = 9 \text{ T}$$

$$T_{\text{sys}} = 8 \text{ K}$$

already explored/excluded
theory prediction
MADMAX (Prototype)
other axion experiments

- **MA**gnetized **D**isk and **M**irror **A**xion **eX**periment: a dielectric haloscope to detect post-inflationary dark matter axions
- MADMAX is now in the beginning of the prototyping phase
- Key progress so far:
 - Booster disk drive system: piezomotor and interferometer successfully testes
 - Gained understanding of the Closed Booster System
 - Magnet R&D ongoing: so far no show stoppers!
- Prototype to be commissioned in SHELL @ UHH in 2022/23
- First physics measurements with prototype at CERN (MORPURGO magnet)
- Commissioning of full-size MADMAX at DESY in HERA Hall North starting ≥ 2028

